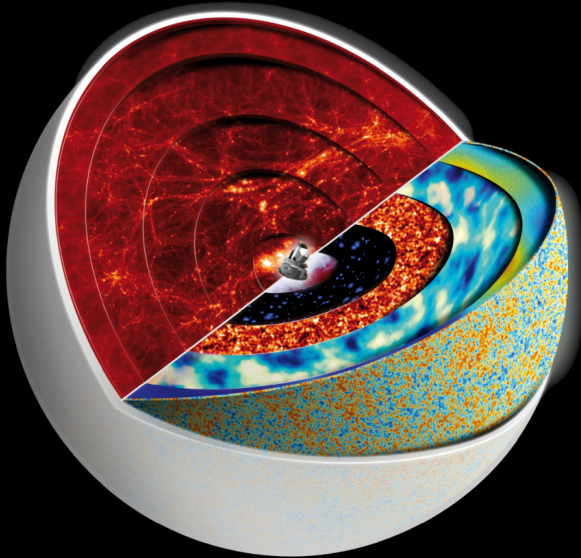


Planck polarization at low ℓ and reionization

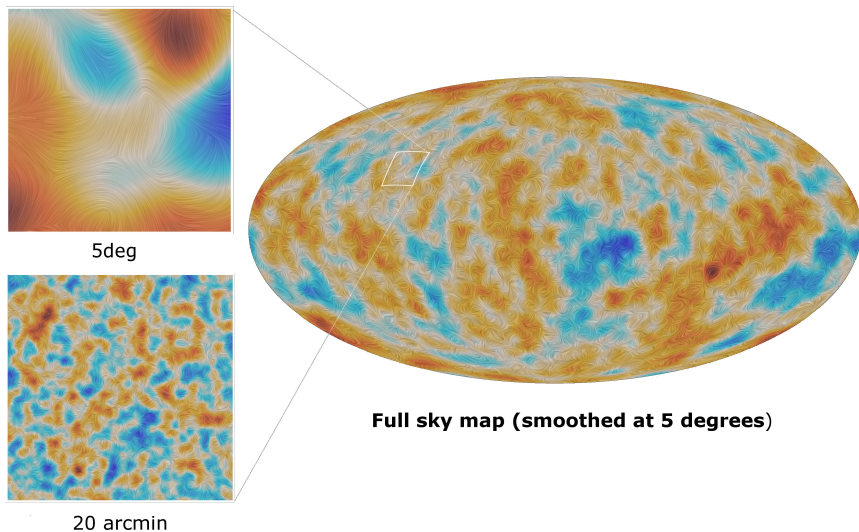


Anne Ducout, on behalf of the Planck collaboration

Outline

- 1 The Planck polarization
- 2 Reionization : Introduction
- 3 Reionization : Results and Planck low- ℓ measurements
- 4 Conclusion

The *Planck* 2015 CMB polarization

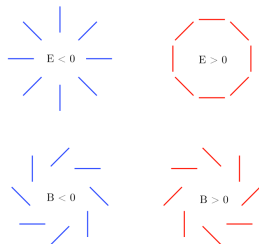


LFI : 30, 44, 70 GHz, **HFI : 100, 143, 217, 353, 545, 857 GHz**

New : HFI low-ell polarization

The CMB angular power spectra

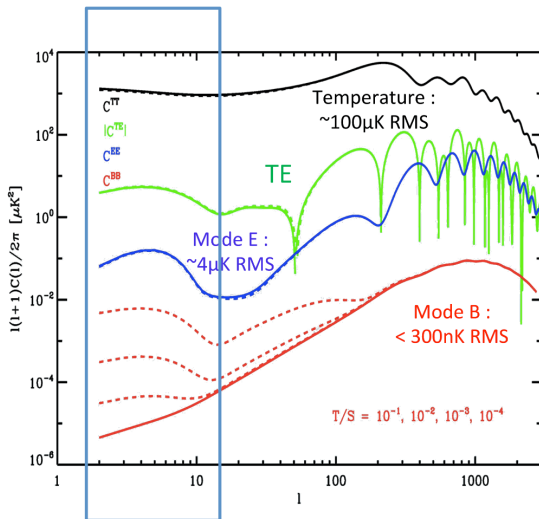
- 3 observables : T, E, B



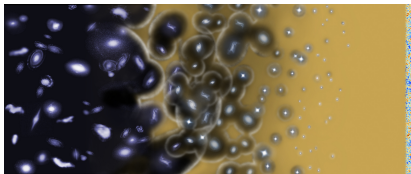
- Low-ell polarization :

Reionization "bump"
around $\ell \lesssim 10$

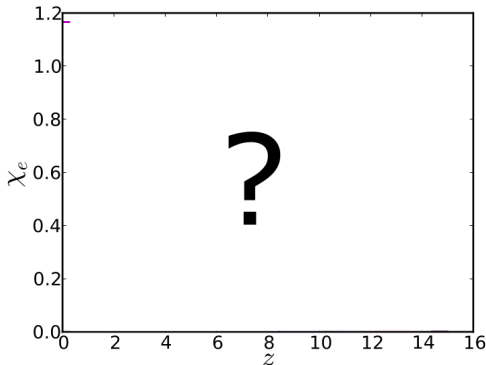
$$\theta \sim \frac{180}{\ell}$$



Reionization



The Epoch of Reionisation (EoR) describes the period during which the cosmic gas went from neutral to ionised at the onset of the first emitting sources.



χ_e is the ionisation fraction as a function of the redshift.

Reionisation optical depth is defined as

$$\tau = \int_0^{\eta_0} a n_e \sigma_T d\eta$$

Reionization history

- **symmetric** (standard tanh)

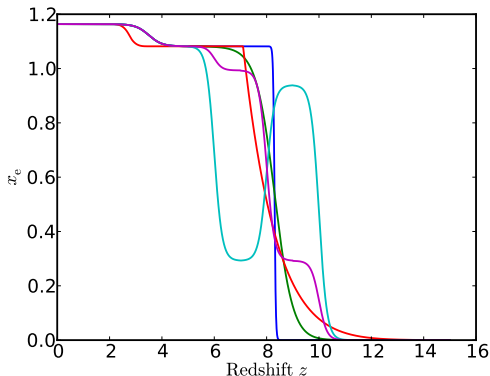
- ▶ 2 parameters : z_{re} , Δz (δz)
includes “instantaneous” reionization

- **asymmetric**

- ▶ 2 sources :
 1. “gentle” : stars & DGs
 2. “abrupt” : QSOs finish
- ▶ phenomenological description :
 z_{begin} , z_{end} , z_{re}

- **model independent**

- ▶ $x_e(z)$ in redshift bins
- ▶ Principal Component Analysis

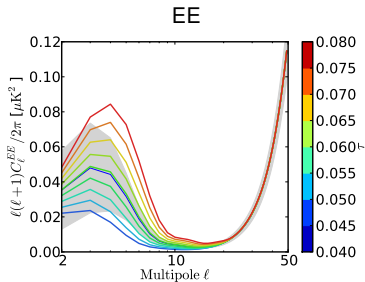
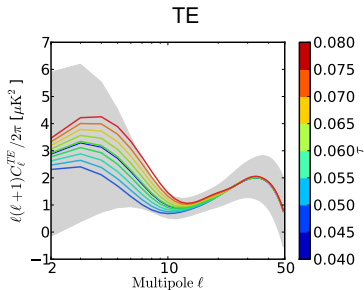


CMB & Reionization

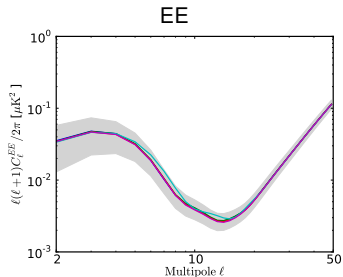
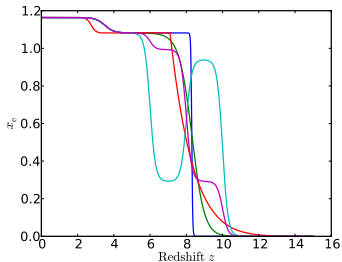
CMB can give information on EoR through :

- **temperature anisotropies :**
suppression of TT power at high multipole (very degenerate with other cosmological parameters and foregrounds)
- **large scale CMB anisotropies in polarisation :**
new polarisation anisotropy at large angular scale because the horizon has grown to a much larger size by that epoch
- **kinetic Sunyaev-Zel'dovich effect :**
Scattering of photons off newly liberated electrons [Sunyaev & Zel'dovich 1980]
(generates CMB temperature anisotropies)

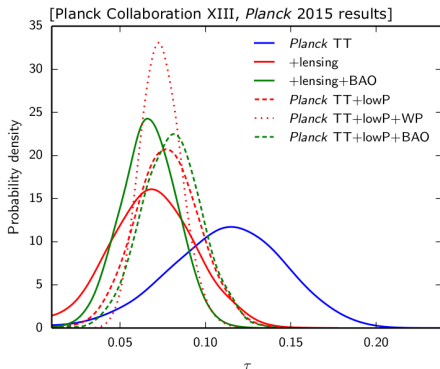
Effect on CMB polarisation at low- ℓ



CMB degeneracies :



Reionization optical depth



From CMB data :

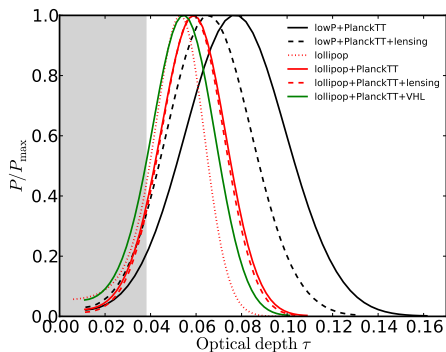
- **WMAP 9yr**
 $\tau = 0.089 \pm 0.014$
- **Planck 2013**
 $\tau = 0.089 \pm 0.014$ (TT with WP)
 $\tau = 0.075 \pm 0.013$ (TT with Planck dust)
- **Planck 2015**
 $\tau = 0.078 \pm 0.019$ (TT with lowP)
 $\tau = 0.066 \pm 0.016$ (TT + lowP + lensing)
 $\tau = 0.067 \pm 0.016$ (TT + lensing + BAO)
- **Planck HFI EE low- ℓ**
Decreasing trend continues...

Case of standard instantaneous reionization model

Planck-HFI low- ℓ data [Planck PIP XLVI, 2016]

- **In previous Planck data, the biggest systematic was ADC-NL**
still the case, but has been reduced by a factor almost 10, and produces very low levels of residuals in the maps.
- **Updated mapmaking procedure :**
significant reduction of systematic effects in the maps, including leakage from temperature to polarization, and very good calibration (better than 0.001%).
- **Results on E2E Monte-Carlo simulations including ADC-NL**
bias remaining : very low and accounted for in the likelihoods.
- **Component separation :** Main data at 100 & 143 GHz
use of 30 and 353 GHz for synchrotron and dust templates.
- **Likelihood based on cross-spectra between Planck frequency maps**
Lollipop : [Hamimeche & Lewis 2008] approximation modified for cross-spectra [Mangilli+2015]
Simbal : biases removed with simulations.

Results : Reionization optical depth and redshift



• Data used, likelihood :

- ▶ Planck Temperature (PlanckTT)
- ▶ Planck Lensing
- ▶ ACT and SPT (VHL)
- ▶ Planck HFI polarization EE at low- ℓ (lolipop)

• Optical depth

$$\tau = 0.058 \pm 0.012 \text{ (Planck TT + lolipop)}$$

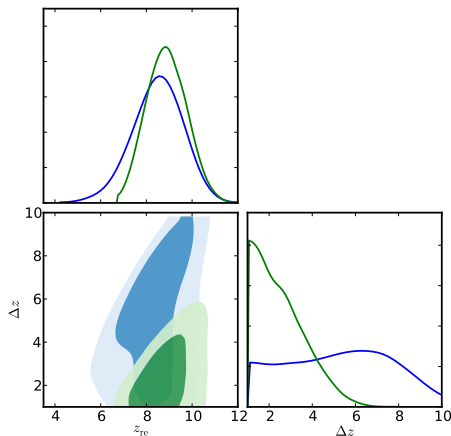
• Redshift

$$7.8 \leq z_{re} \leq 8.8 \text{ (All data)}$$

$$\Delta z < 2.8 \text{ (All data)}$$

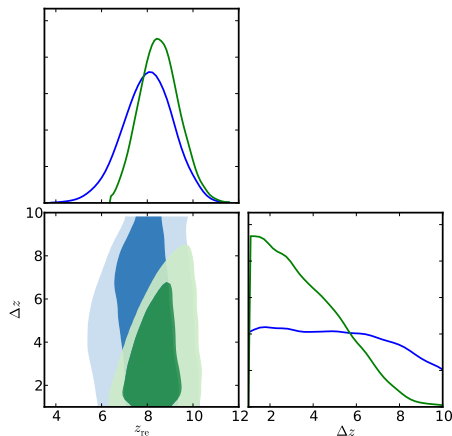
The Universe is ionized at less than 10% level at redshifts above $z \simeq 10$.

Symmetric / asymmetric model (Planck TT+lollipop)



Symmetric

$$z_{re} = 8.8 \pm 0.9 \text{ (with prior)}$$



Asymmetric

$$z_{re} = 8.5 \pm 0.9 \text{ (with prior)}$$

$$\Delta z < 6.8 \text{ (with prior)}$$

Green Contour : added prior $z_{end} > 6$

Using the kSZ effect (+VHL)

⇒ allows to break the degeneracy with Δz .

2 types of kSZ :

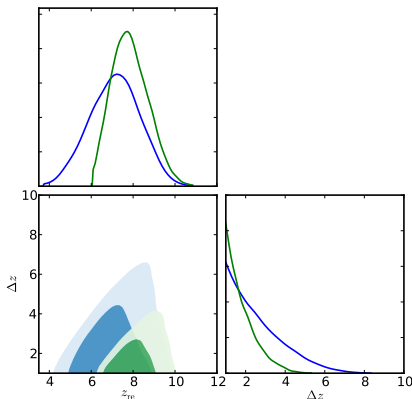
- Homogeneous (h) kSZ : scattering on electrons in bulk motion from clusters, when **reionization is complete**.
- Patchy (p) kSZ : **during the process of reionization**, from proper motion of ionized bubbles around emitting sources.

Different spectra for the different types of kSZ :

$$\mathcal{D}_\ell^{\text{kSZ}} = \mathcal{D}_\ell^{\text{h-kSZ}} + \mathcal{D}_\ell^{\text{p-kSZ}}$$

Amplitude of p-kSZ power primarily depends on duration of reionization.

But very model-dependent...



$\Delta z < 2.8$ (with prior $z_{\text{end}} > 6$)
(symmetric model)

Conclusion

- A lower value for τ as suggested by Planck data would be
 - ▶ **consistent with a fully reionised Universe at $z \sim 6$**
(Gunn-Peterson effect showing Universe is mostly ionized up to $z \sim 6$ [Fan et al. 2006a])
 - ▶ **in good agreement with recent constraints on reionisation in the direction of particular objects** (in particular distant GRB and Ly- α emitters)
- Constraints on the reionisation history with such a low optical depth would **disfavour large abundances of star-forming galaxies beyond $z = 15$**
- Maintaining a UV-luminosity density at the maximum level allowed by the luminosity density constraints at redshifts $z < 9$ and considering only the currently observed galaxy population at $M_{UV} < -17$ seems to be **sufficient to comply with all the observational constraints without the need for high redshift ($z = 10$ to 15) galaxies.**

The scientific results that we present today are a product of the Planck Collaboration, including individuals from more than 100 scientific institutes in Europe, the USA and Canada.



Planck is a project of the European Space Agency, with instruments provided by two scientific Consortia funded by ESA member states (in particular the lead countries: France and Italy) with contributions from NASA (USA), and telescope reflectors provided in a collaboration between ESA and a scientific Consortium led and funded by Denmark.