

Minicharged Dark Matter from Gauge Coupling Unification

Norimi Yokozaki (Tohoku U.)

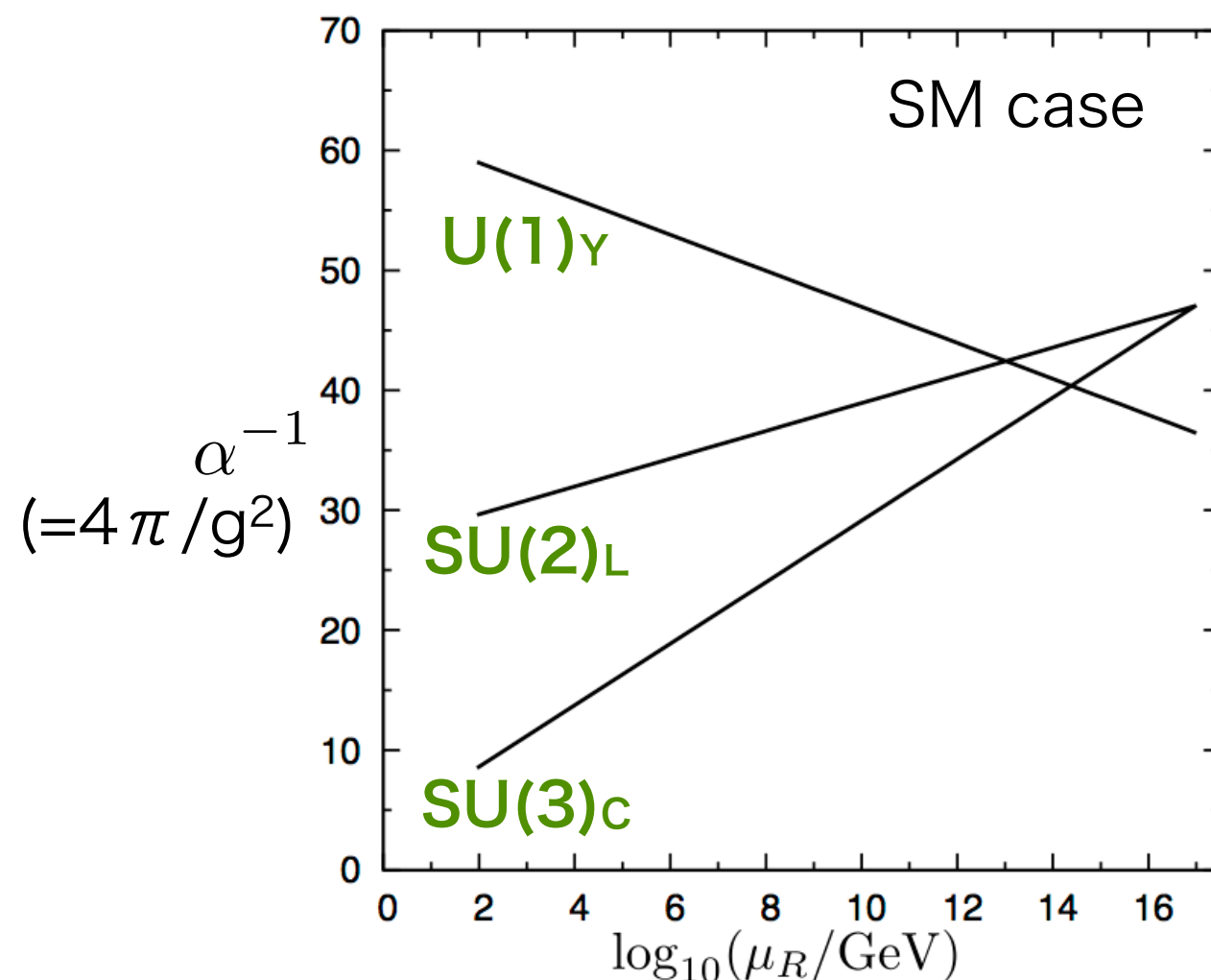
Fuminobu Takahashi, Masaki Yamada, **N.Y.** arXiv:1604.07145, PLB

Ryuji Daido, Fuminobu Takahashi, **N.Y.** arXiv:1610.00631, PLB

Motivations to go beyond the SM

- Dark matter

Unification of SM forces and matters,
and charge quantization

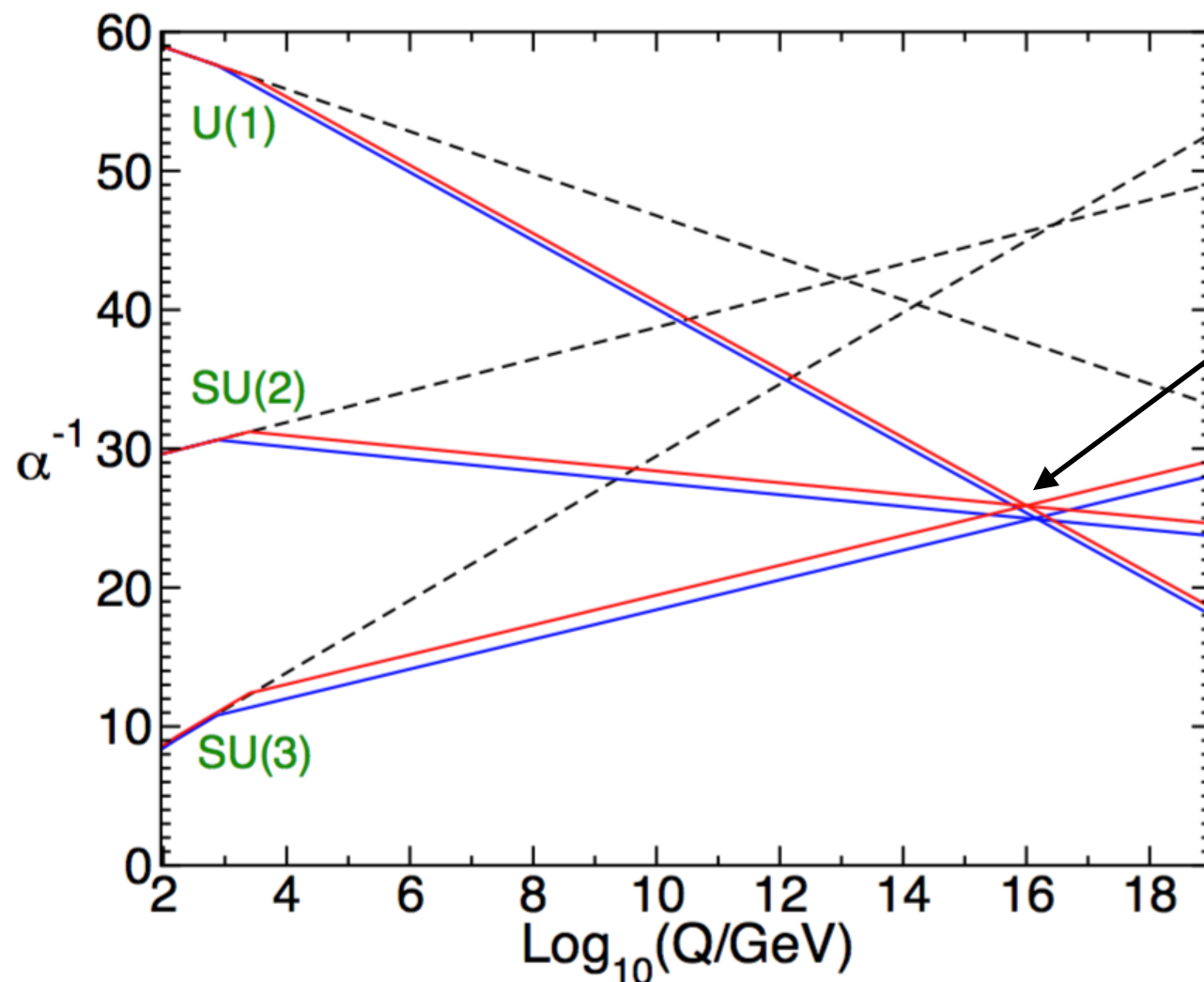


In SM, the unification of the gauge couplings fails apparently.

Motivations to go beyond the SM

- Dark matter

Unification of SM forces and matters,
and charge quantization

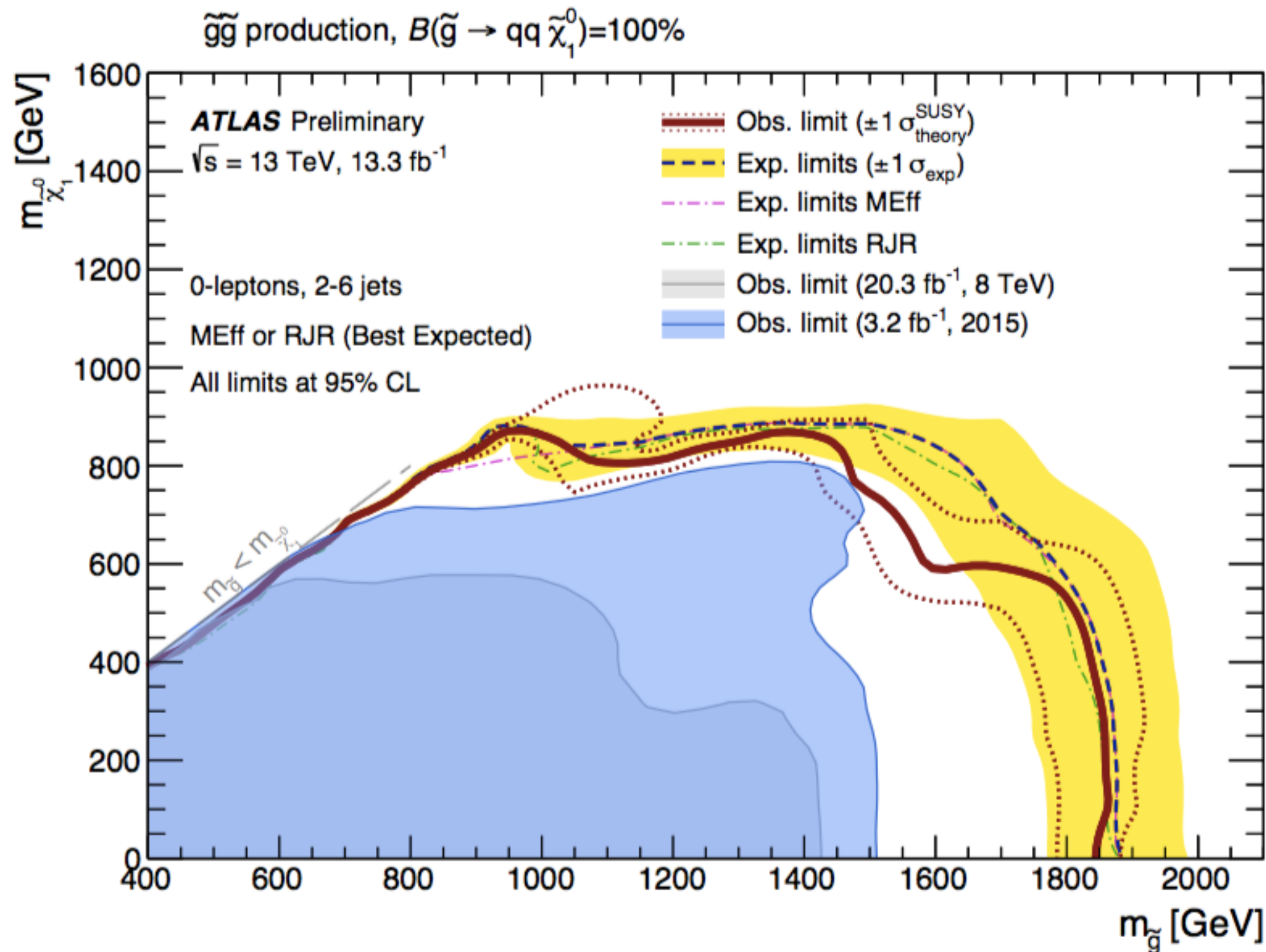


MSSM case

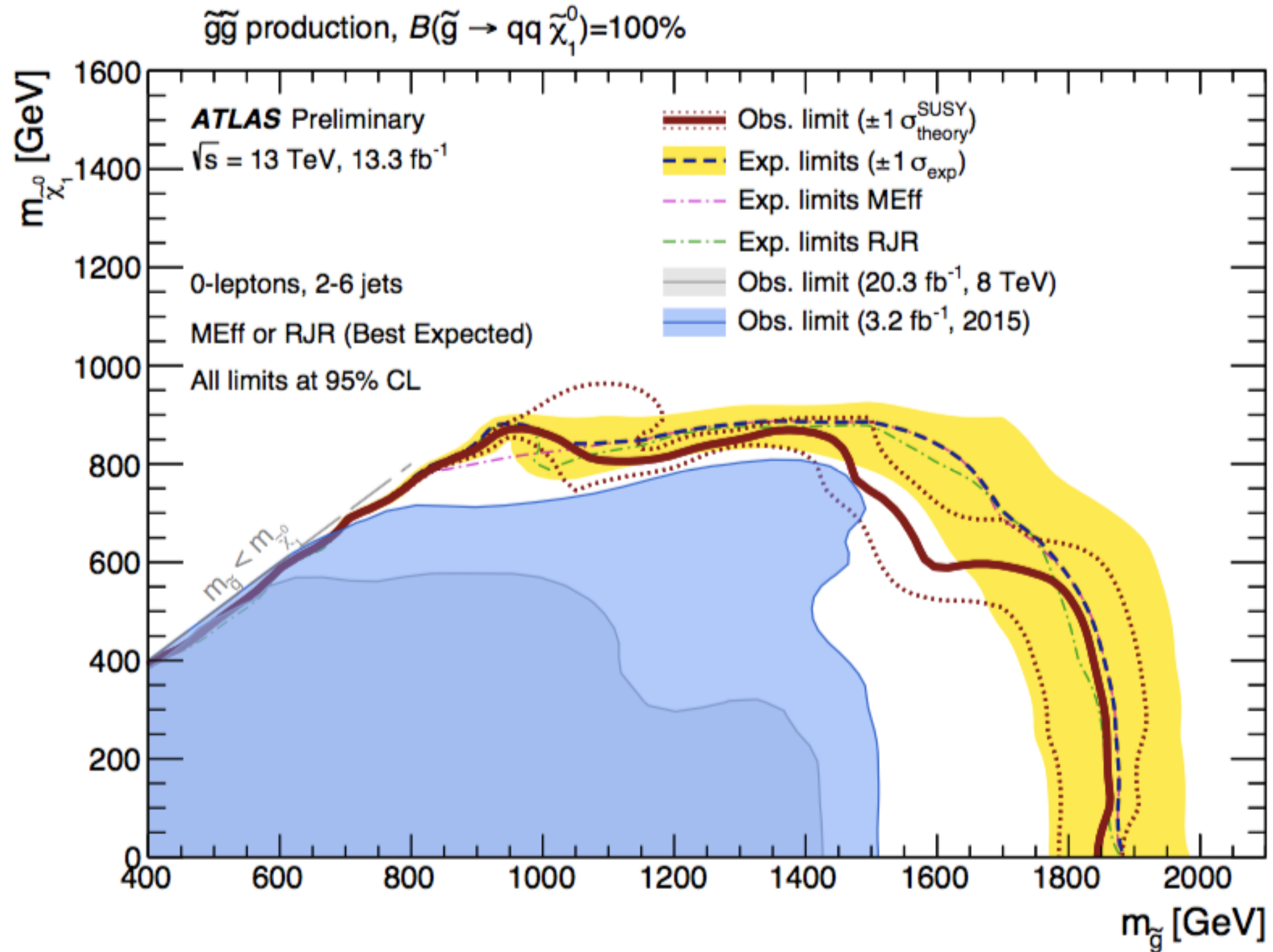
Supersymmetry might
be a solution

[Fig. from SUSY primer, S. Martin]

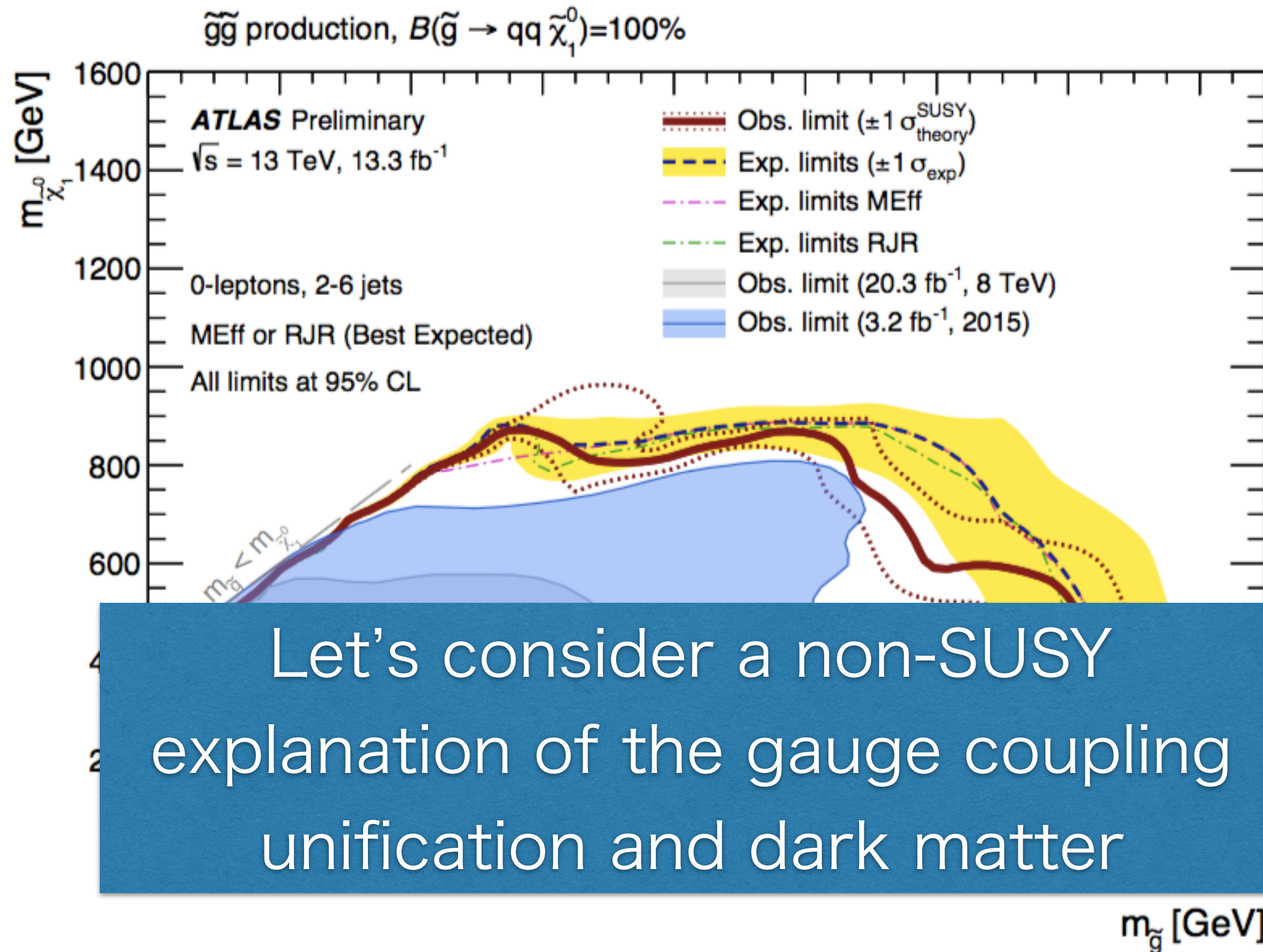
However, SUSY particles have not been found



Colored SUSY particles should be heavier than 1.4-1.8 TeV.

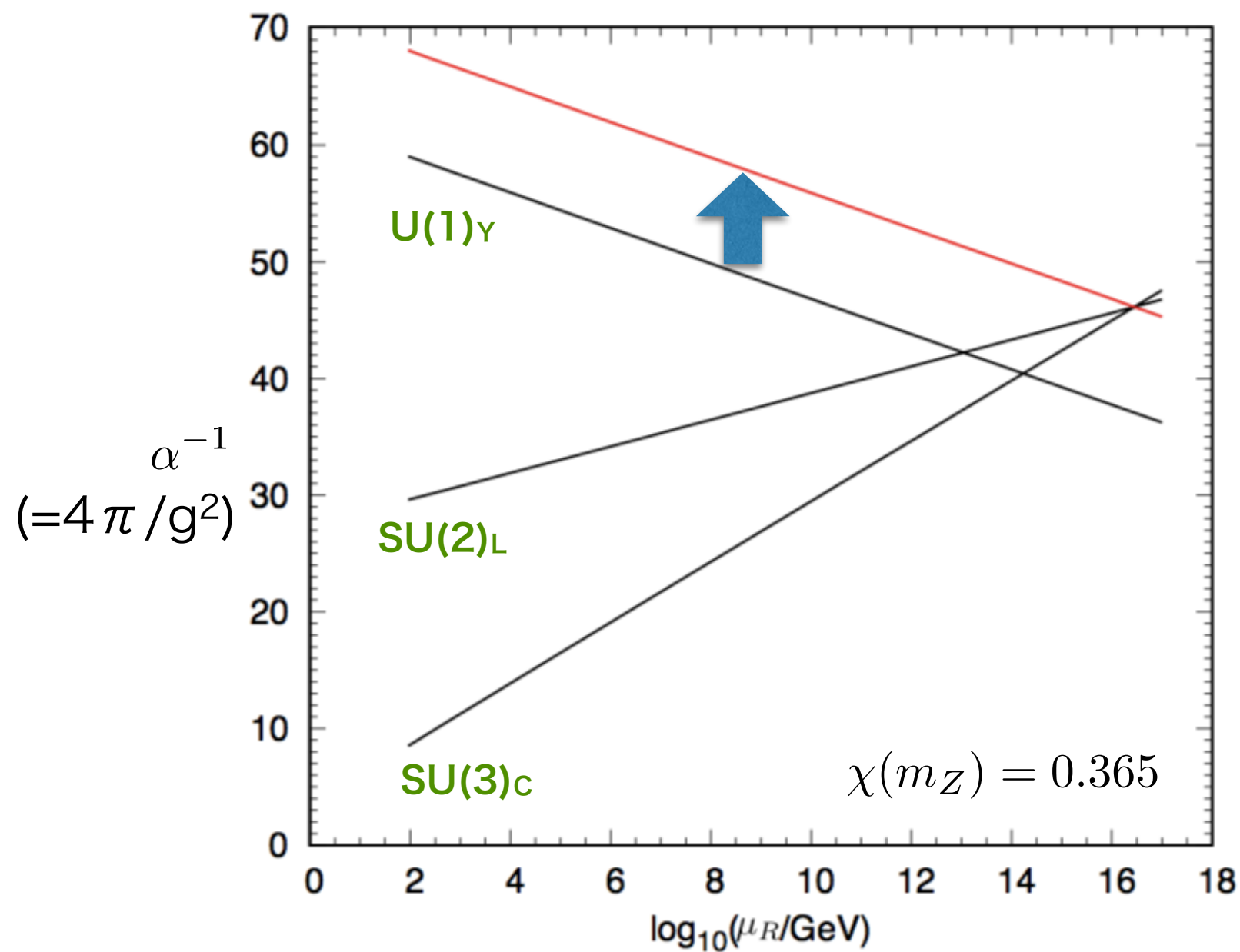


Colored SUSY particles should be heavier than 1.4-1.8 TeV.



We introduce a **massless hidden photon** and **hidden matter** to SM

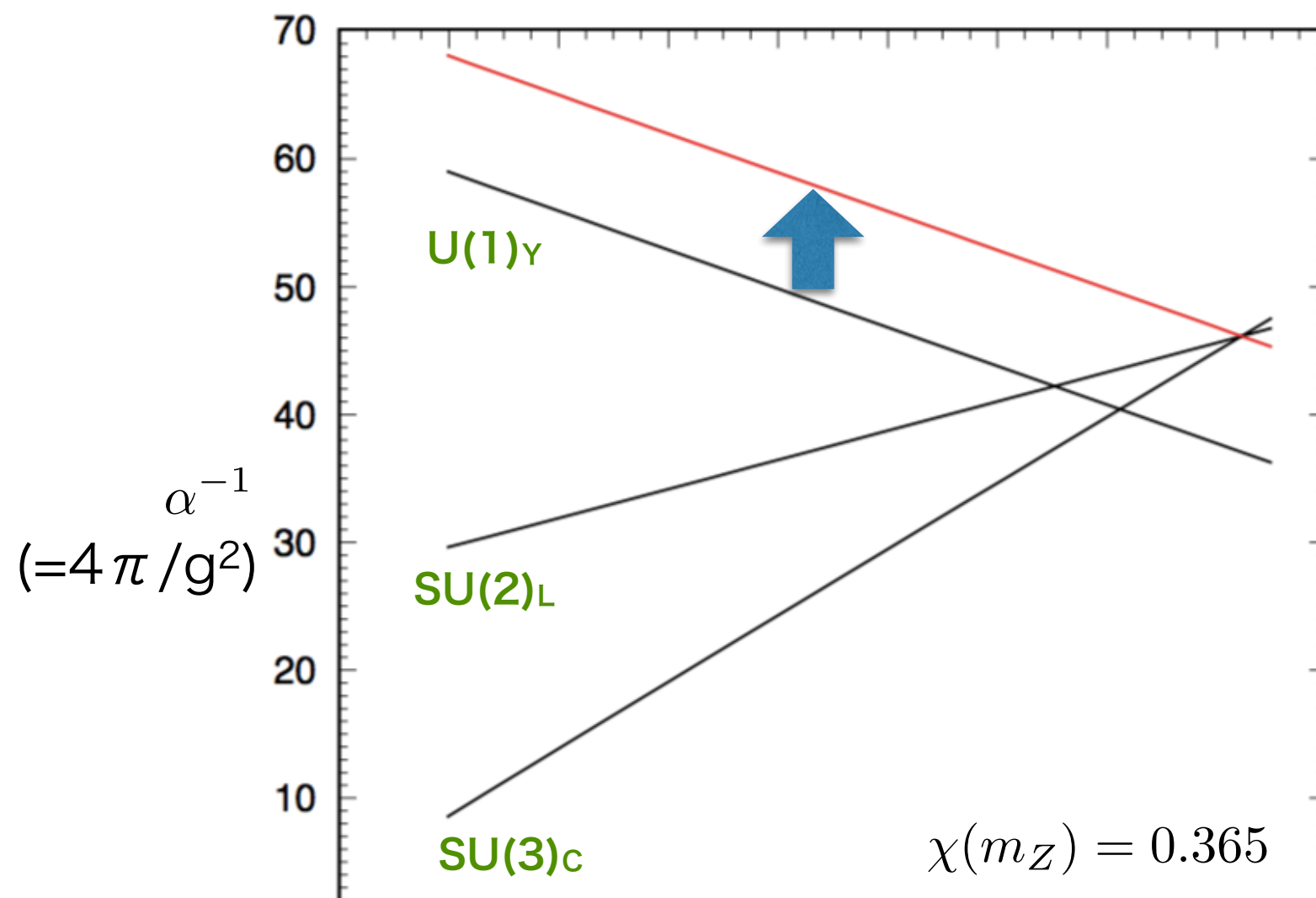
The normalization of the $U(1)_Y$ gauge coupling is modified by the kinetic mixing with a massless hidden photon.



[Redondo, 2008; Takahashi, Yamada, N.Y., 2016; Daido, Takahashi, N.Y., 2016]

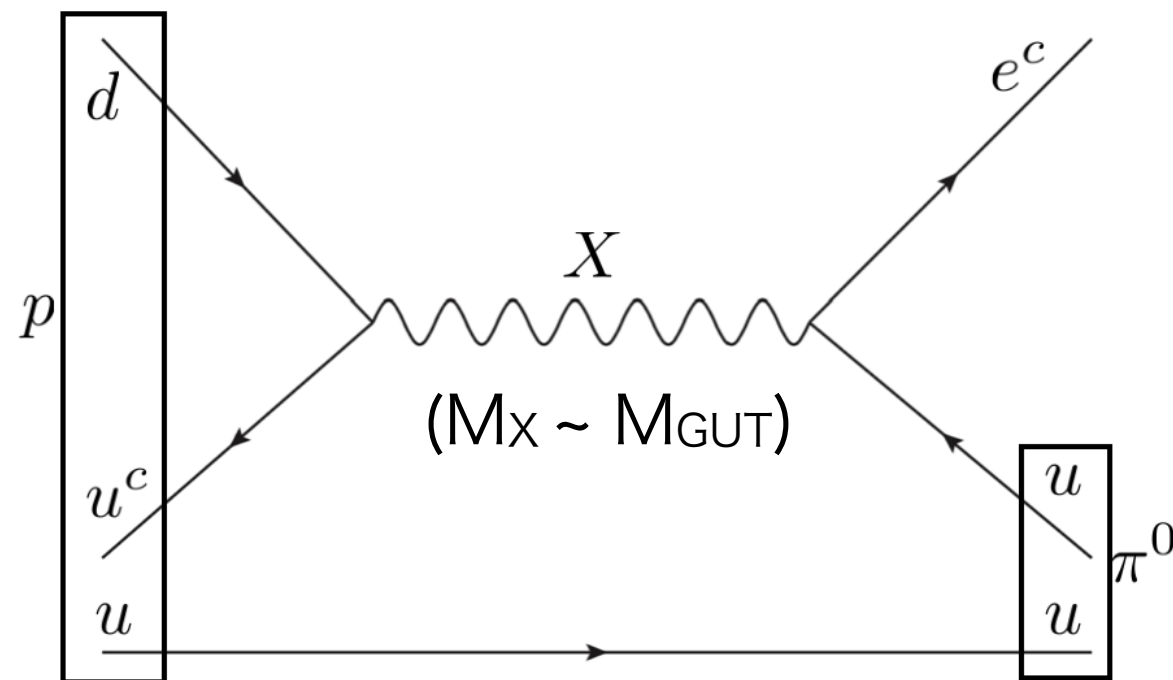
We introduce a **massless hidden photon** and **hidden matter** to SM

The normalization of the $U(1)_Y$ gauge coupling is modified by the kinetic mixing with a massless hidden photon.



The gauge coupling unification is achieved
without Supersymmetry

- The unification scale is $\sim 10^{16.5}$ GeV, avoiding the rapid proton decay



$\sim 10^{37}$ years ($p \rightarrow \pi^0 e^+$)

exp: $> 1.7 \times 10^{34}$ years

[Takhistov, 2016]

The unbroken $U(1)_H$ explains the stability of the DM (hidden matter)

A model with a hidden
photon ($U(1)_H$ gauge boson)

unbroken

Consider $U(1)_Y \times U(1)_H$ model with a kinetic mixing

$$\mathcal{L} = -\frac{1}{4}F_Y'^{\mu\nu}F_{Y\mu\nu}' - \frac{1}{4}F_H'^{\mu\nu}F_{H\mu\nu}' - \frac{\chi}{2}F_Y'^{\mu\nu}F_{H\mu\nu}'$$

$$F_i'^{\mu\nu} \equiv \partial^\mu A_i'^\nu - \partial^\nu A_i'^\mu \quad (i = Y, H)$$

[Holdom, 1986]

Here, we consider unbroken $U(1)_H$

Consider $U(1)_Y \times U(1)_H$ model with a kinetic mixing

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$$F_i'^{\mu\nu} \equiv \partial^\mu A_i'^\nu - \partial^\nu A_i'^\mu \quad (i = Y, H)$$

By the field redefinitions, we can go to the canonical basis

$$A_Y^{\mu'} = \frac{A_Y^\mu}{\sqrt{1-\chi^2}}$$

$$A_H^{\mu'} = A_H^\mu - \frac{\chi}{\sqrt{1-\chi^2}}A_Y^\mu$$



$$\mathcal{L} = -\frac{1}{4}F_Y^{\mu\nu}F_{Y\mu\nu} - \frac{1}{4}F_H^{\mu\nu}F_{H\mu\nu}$$

Consider $U(1)_Y \times U(1)_H$ model with a kinetic mixing

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$$F_i'^{\mu\nu} \equiv \partial^\mu A_i'^\nu - \partial^\nu A_i'^\mu \quad (i = Y, H)$$

Let's consider a matter field charged under $U(1)_H$

$$\begin{aligned} & \bar{\Psi}_i \gamma_\mu (g_H' q_{Hi} A_H'^\mu) \Psi_i \\ = & \bar{\Psi}_i \gamma_\mu \left(\boxed{-\frac{q_{Hi} g_H \chi}{\sqrt{1 - \chi^2}} A_Y^\mu} + g_H q_{Hi} A_H^\mu \right) \Psi_i \end{aligned}$$

The hidden matter obtains fractional $U(1)_Y$ charge
proportional to g_H and χ

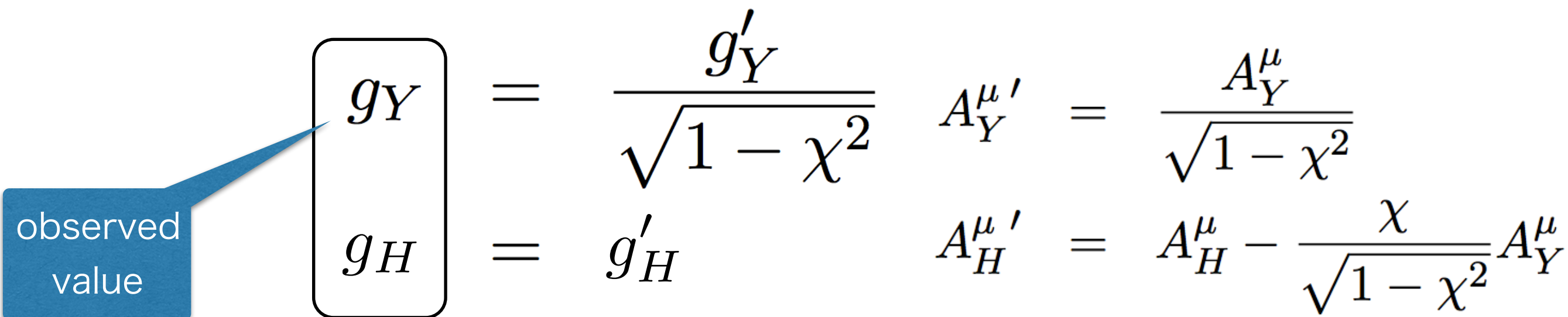
Consider $U(1)_Y \times U(1)_H$ model with a kinetic mixing

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By the field redefinitions, we can go to the canonical basis

Couplings are related to



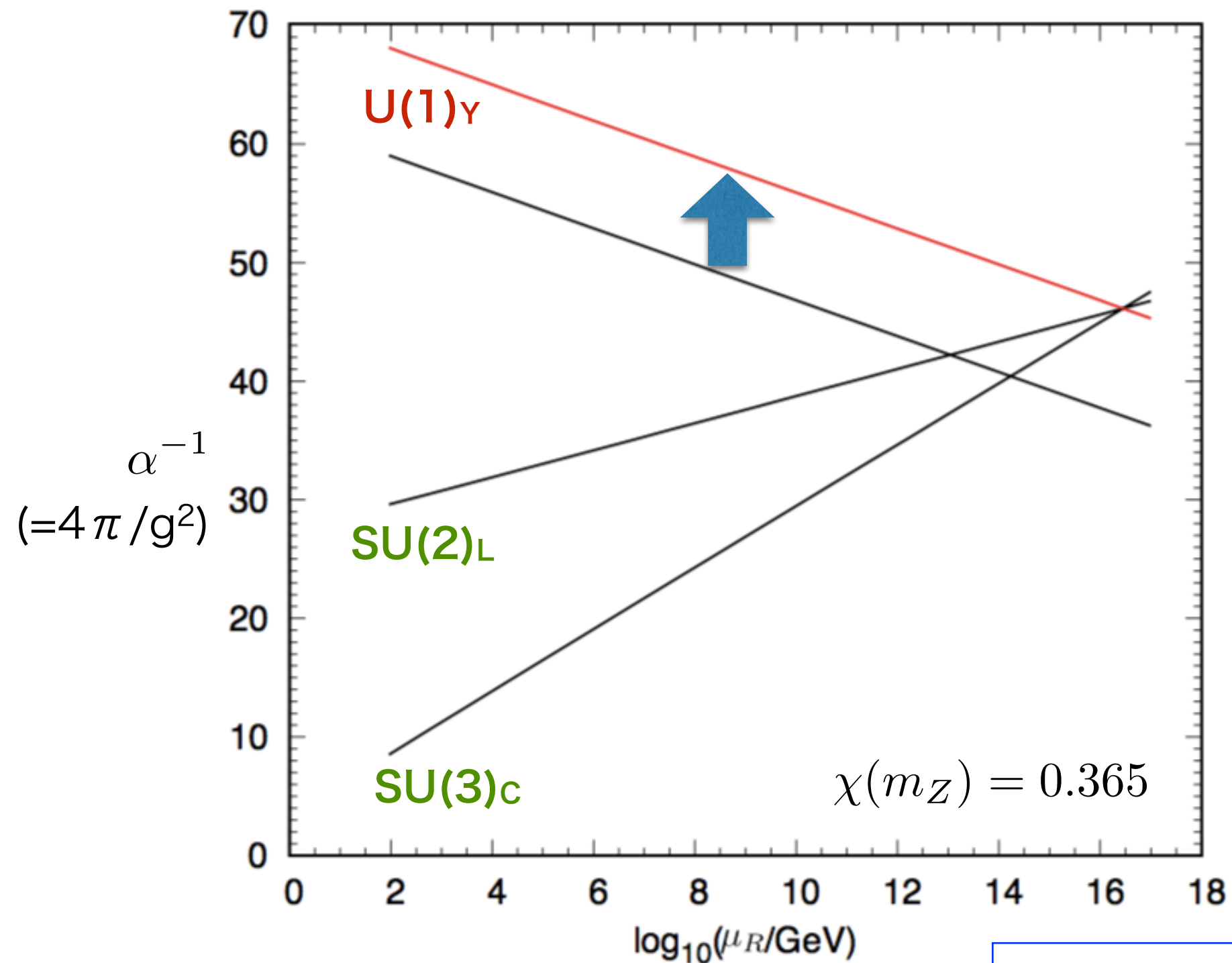
$$\begin{array}{l} \boxed{\begin{array}{l} g_Y \\ g_H \end{array}} = \begin{array}{l} \frac{g_Y'}{\sqrt{1-\chi^2}} \\ g_H' \end{array} \quad \begin{array}{l} A_Y^{\mu'} = \frac{A_Y^\mu}{\sqrt{1-\chi^2}} \\ A_H^{\mu'} = A_H^\mu - \frac{\chi}{\sqrt{1-\chi^2}} A_Y^\mu \end{array} \end{array}$$

couplings in the
canonical basis

Normalization of $U(1)_Y$ changes

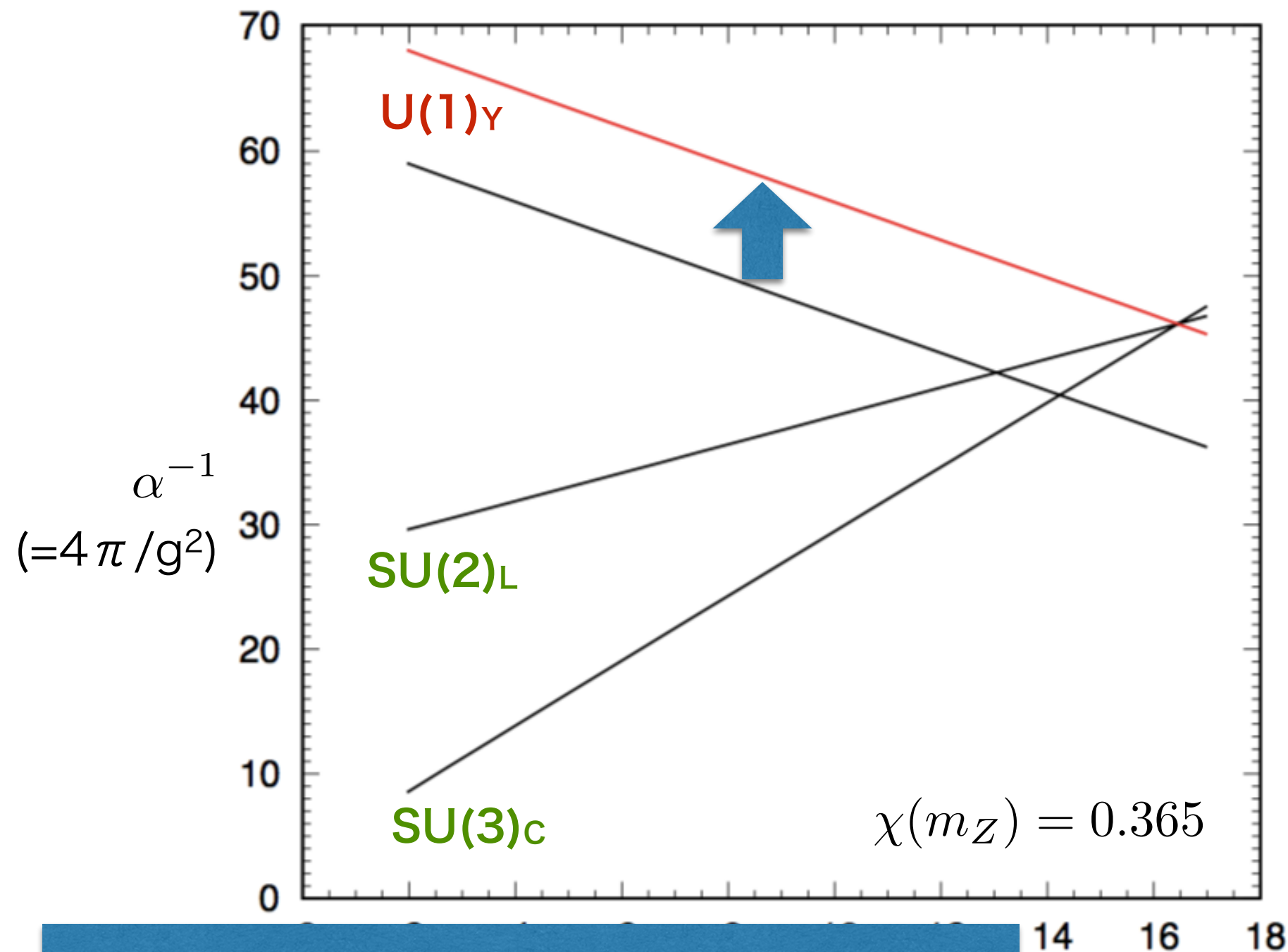
Gauge coupling unification

Without hidden matters



$$\alpha_1^{-1} = \frac{(\alpha_1)_{\text{measured}}^{-1}}{(1 - \chi^2)}$$

Without hidden matters



This is just the change of the normalization of $U(1)_Y$

$$\alpha_1^{-1} = \frac{(\alpha_1)_{\text{measured}}^{-1}}{(1 - \chi^2)}$$

In other words, there are no phenomenological implications
at the low-energy

$$\bar{\Psi}_i \gamma_\mu (g'_Y Q_i A'^\mu_Y) \Psi_i = \bar{\Psi}_i \gamma_\mu (g_Y Q_i A^\mu_Y) \Psi_i$$

(original basis) (canonical basis)

$$g_Y = \frac{g'_Y}{\sqrt{1 - \chi^2}}, \quad \text{with} \quad \begin{aligned} A^{\mu'}_Y &= \frac{A^\mu_Y}{\sqrt{1 - \chi^2}} \\ A^{\mu'}_H &= A^\mu_H - \frac{\chi}{\sqrt{1 - \chi^2}} A^\mu_Y \end{aligned}$$

The hypercharges are identical: the basis of GUT is not fixed.

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Once we have a hidden matter ...

The unification basis is fixed.

We have a dark matter candidate!

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Does it spoil the unification?

In other words, there are no phenomenological implications
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$$\bar{\Psi}_i \gamma_\mu (g'_Y Q_i A'^\mu_Y) \Psi_i = \bar{\Psi}_i \gamma_\mu (g_Y Q_i A^\mu_Y) \Psi_i$$

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The hypercharges are identical: the basis of GUT is not fixed.

Once we have a hidden matter ...

The unification basis is fixed.

We have a dark matter candidate!

Does it spoil the unification?

No. We have checked it using 2-loop RGEs.

Results with 2-loop RGEs

Case with a hidden matter which is a singlet of SU(5)

$$\mathcal{L} = -\frac{1}{4}F'_{\mu\nu}F'^{\mu\nu} - \frac{1}{4}F'_{H\mu\nu}F'^{\mu\nu}_H - \frac{\chi}{2}F'_{H\mu\nu}F'^{\mu\nu}$$

$$-M_0\bar{\Psi}_0\Psi_0 + (\text{other terms in SM Lagrangian})$$

↖
No SM charges

With a hidden charge of unity

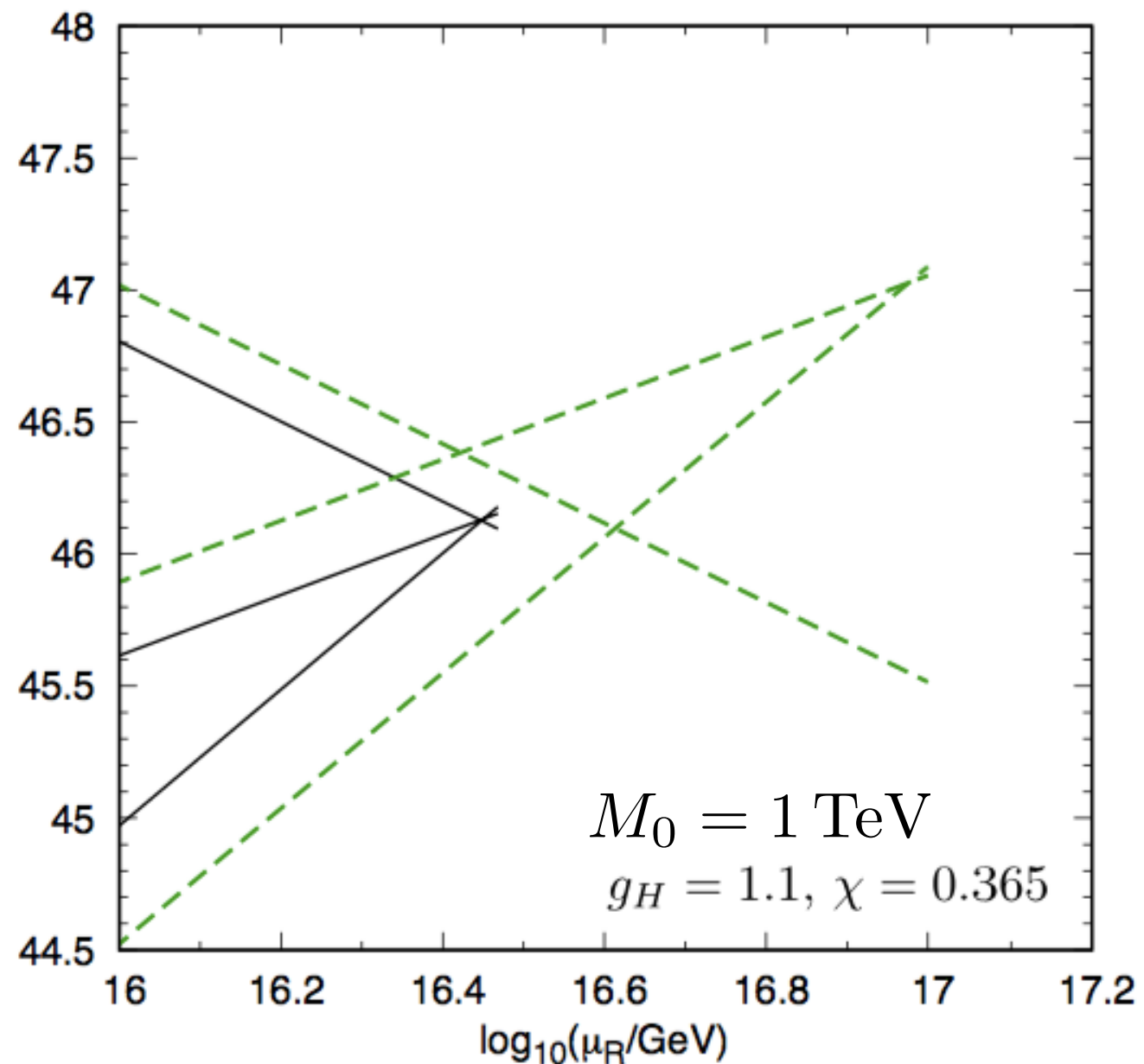
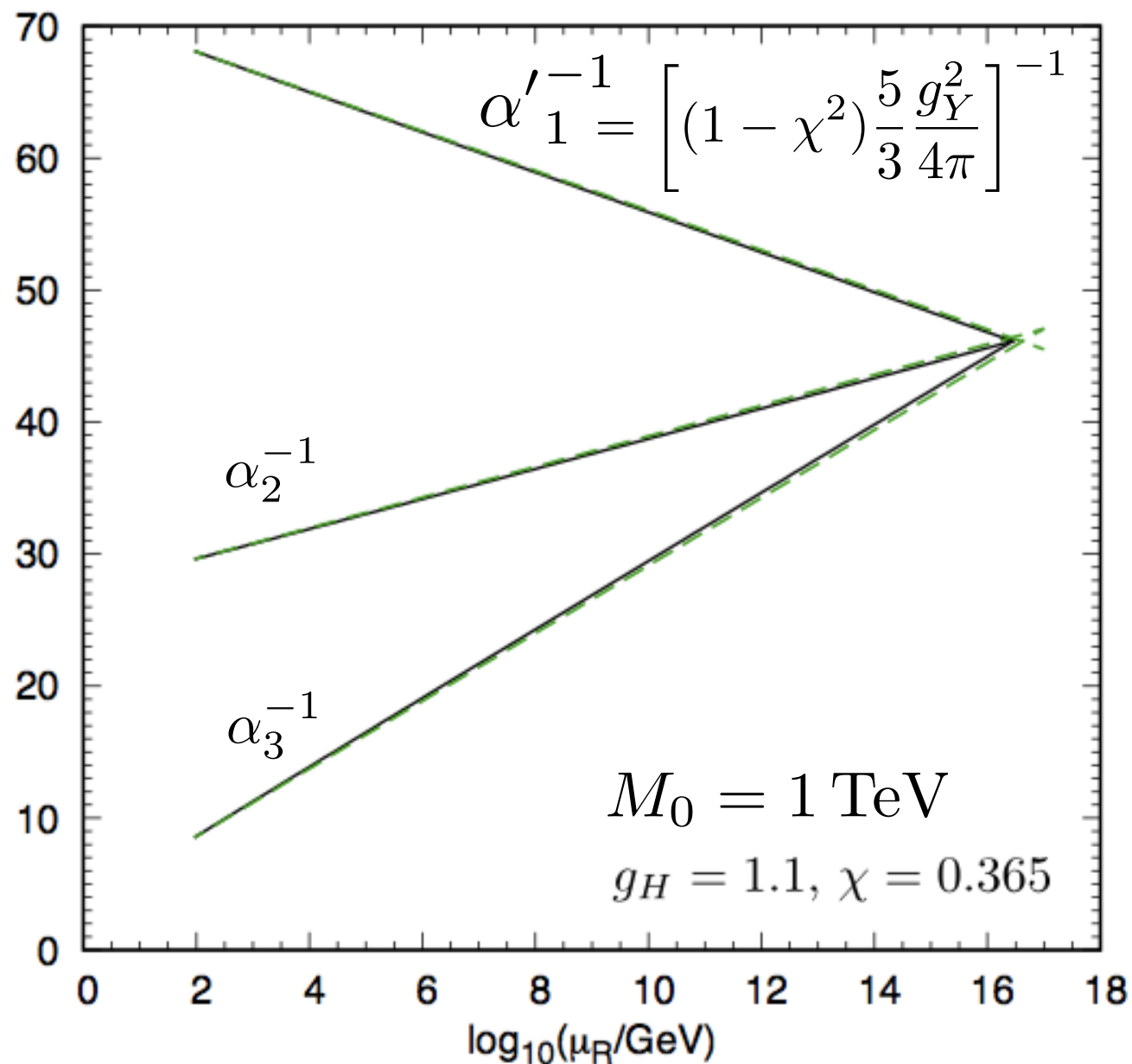
Case with a hidden matter which is a singlet of SU(5)

$$\mathcal{L} = -\frac{1}{4}F'_{\mu\nu}F'^{\mu\nu} - \frac{1}{4}F'_{H\mu\nu}F'^{\mu\nu}_H - \frac{\chi}{2}F'_{H\mu\nu}F'^{\mu\nu} - M_0\bar{\Psi}_0\Psi_0 + (\text{other terms in SM Lagrangian})$$

$$\frac{dg_Y}{d\ln\mu_R} = \frac{1}{16\pi^2} \left(\frac{41}{6}g_Y^3 + \frac{4}{3}g_Y g_H^2 \frac{\chi^2}{1-\chi^2} \right) \text{ canonical basis}$$

$$\frac{dg'_Y}{d\ln\mu_R} = \frac{1}{16\pi^2} \left(\frac{41}{6}g_Y'^3 \right) \text{ original basis}$$

$$\alpha^{-1} (=4\pi/g^2)$$

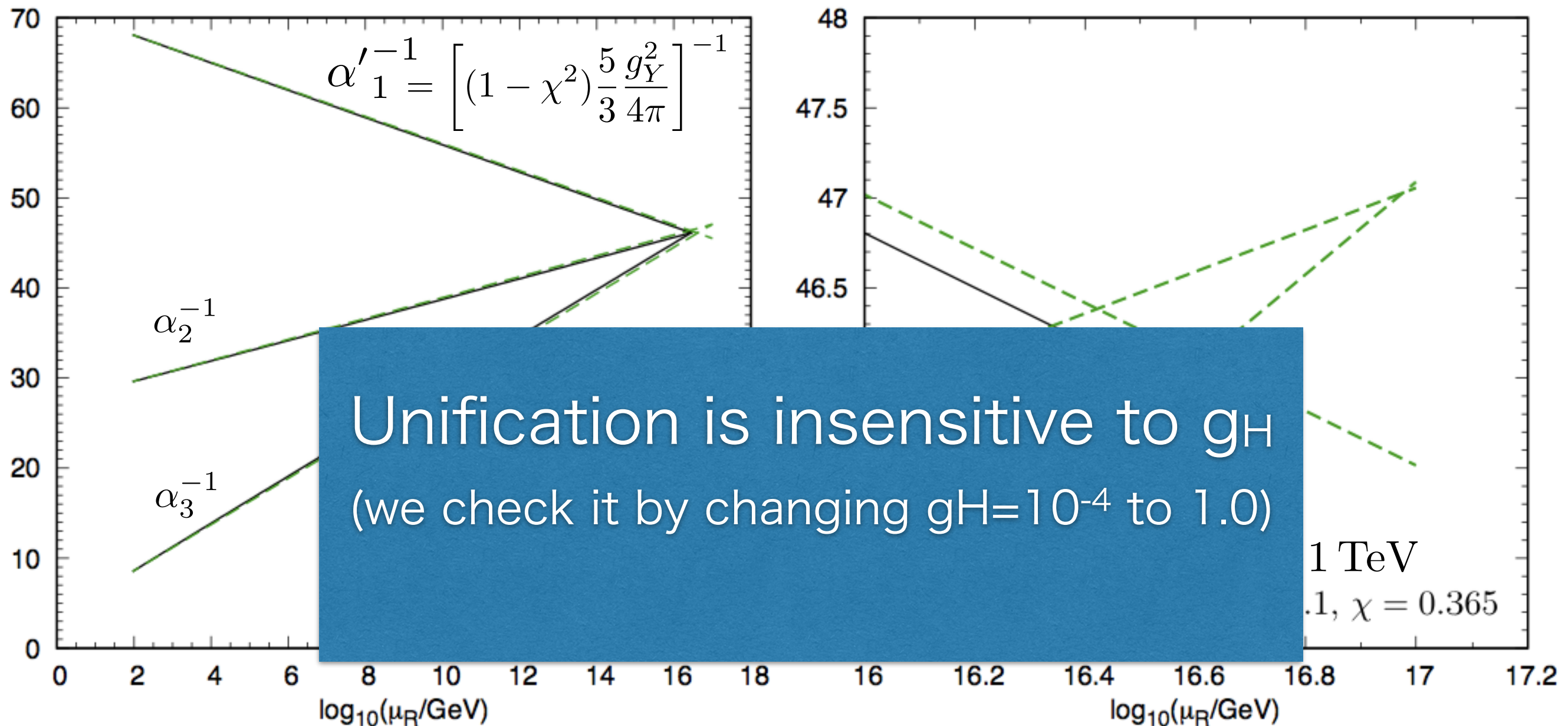


green dashed: one-loop

black solid: two-loop

The unification is fixed by χ (mz)

$$\alpha^{-1} (=4\pi/g^2)$$



green dashed: one-loop

black solid: two-loop

The unification is fixed by χ (mz)

With N_{bi} pairs of $SU(5)$ multiplets having $U(1)_H$ charges

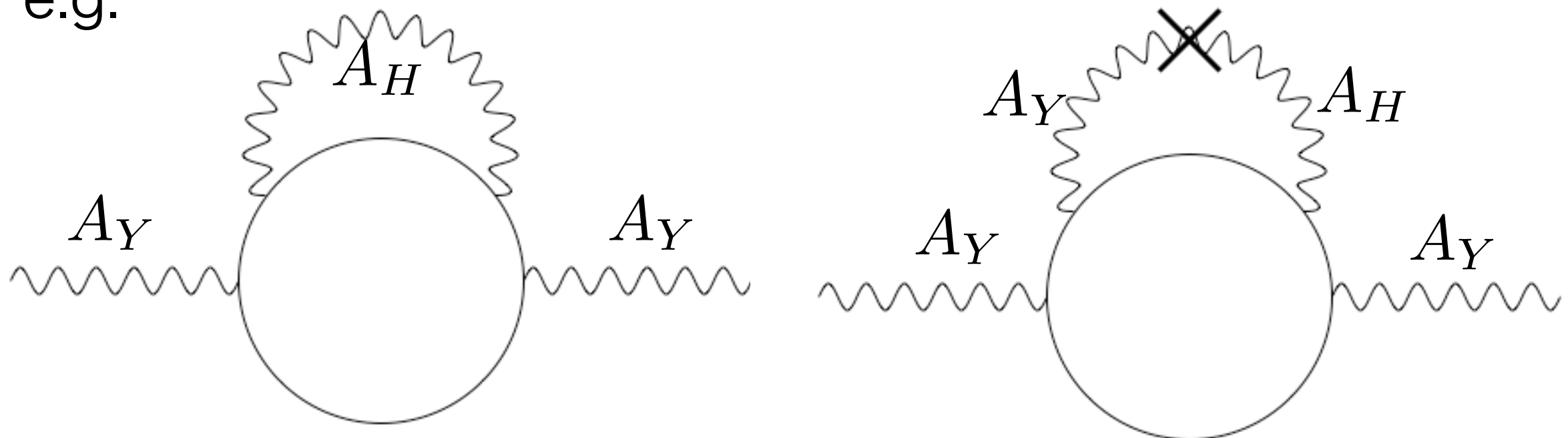
$$\mathcal{L} = -\frac{1}{4}F'_{\mu\nu}F'^{\mu\nu} - \frac{1}{4}F'_{H\mu\nu}F'^{\mu\nu}_H - \frac{\chi}{2}F'_{H\mu\nu}F'^{\mu\nu}$$

$$-M_V \sum_{i=1}^{N_{\text{bi}}} (\bar{\Psi}_{L,i} \Psi_{L,i} + \bar{\Psi}_{\bar{D},i} \Psi_{\bar{D},i}),$$

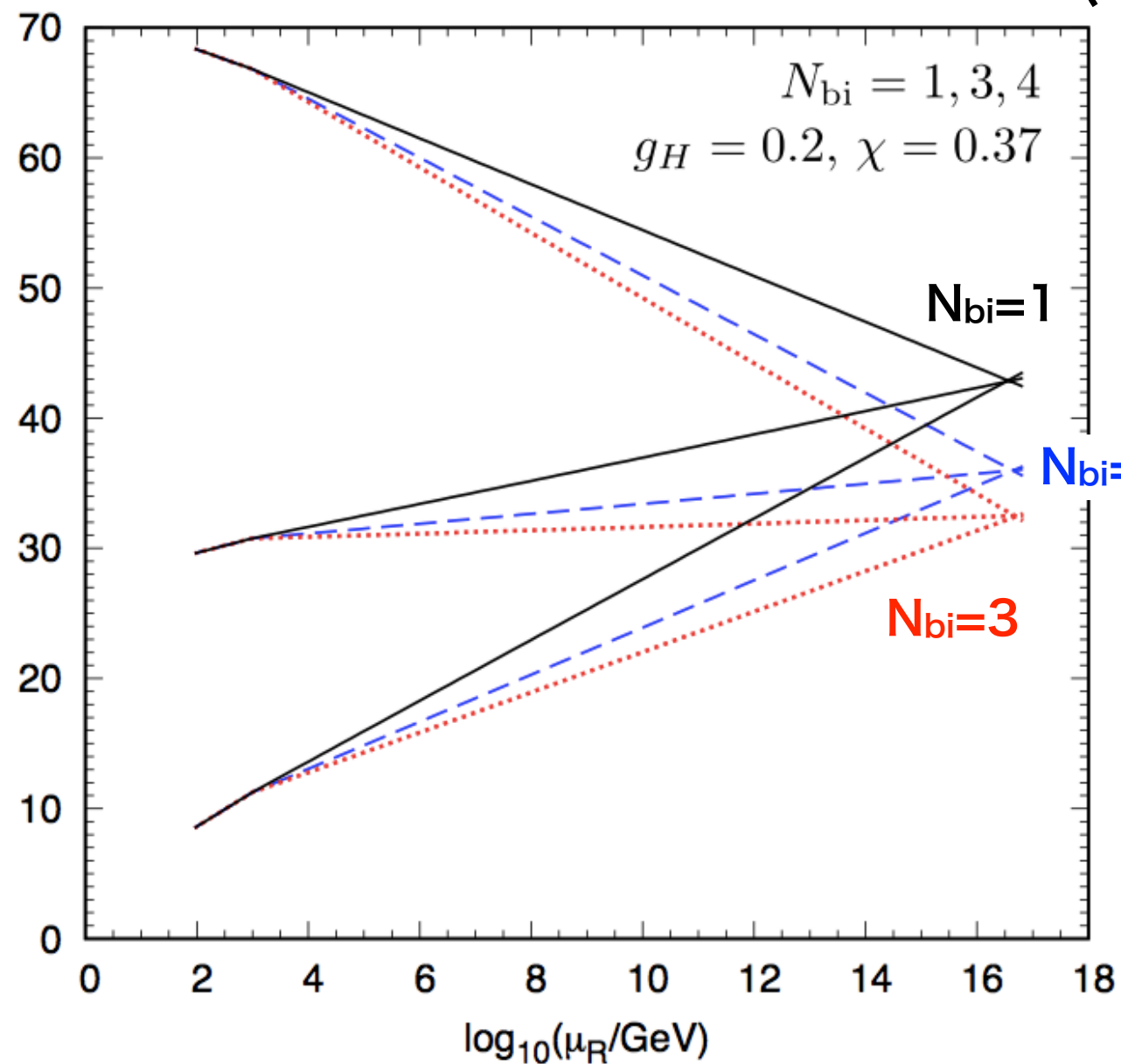
bi-charged fields

$$\text{with } (Q_Y, q_H) = \begin{cases} (-\frac{1}{2}, 1) & \text{for } \Psi_{L,i} \\ (\frac{1}{3}, 1) & \text{for } \Psi_{\bar{D},i} \end{cases}$$

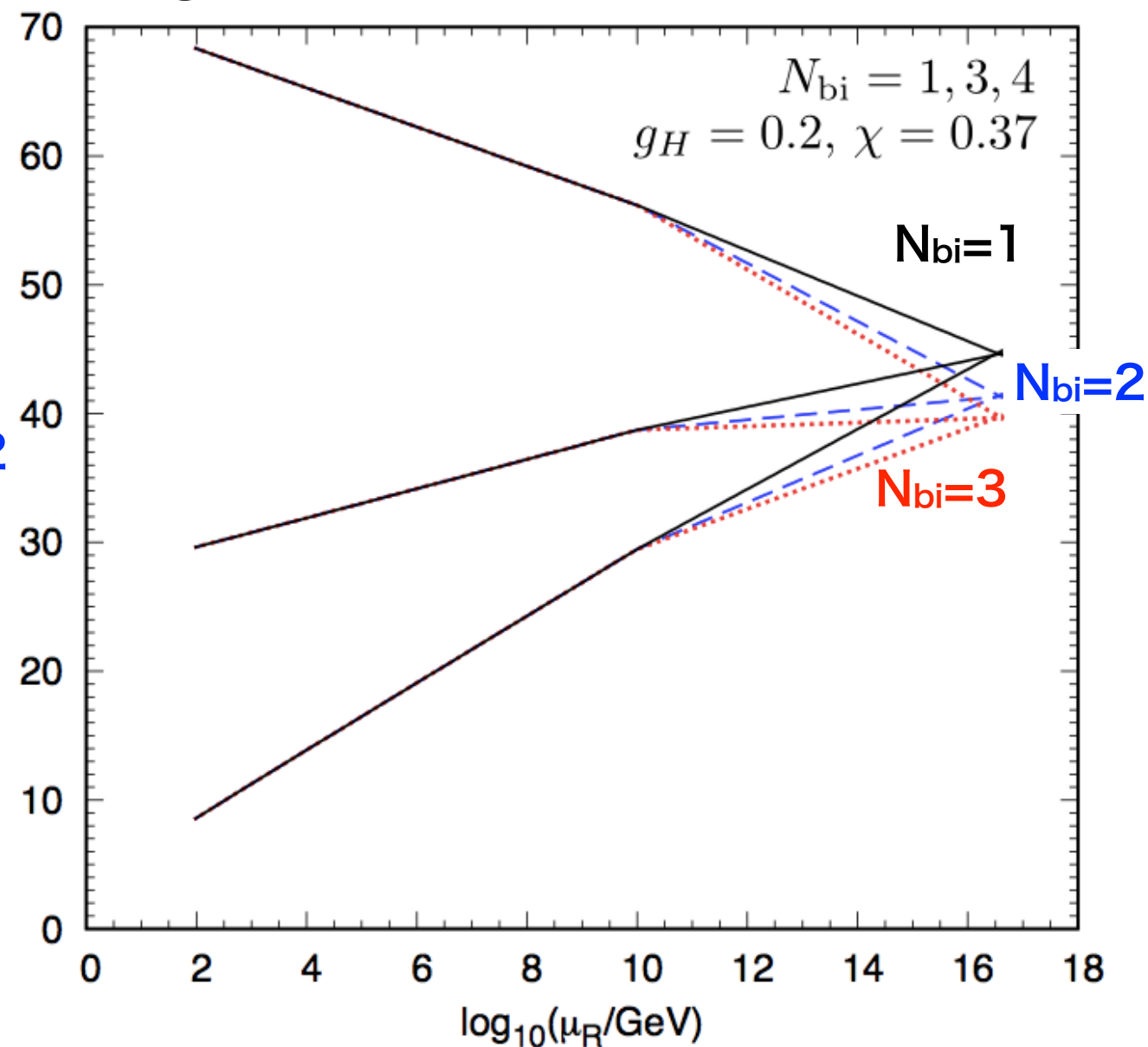
e.g.



$$\alpha^{-1} (=4\pi/g^2)$$



$$M_V = 1 \text{ TeV}$$



$$M_V = 10^{10} \text{ GeV}$$

Almost insensitive to N_{bi} , g_H and M_V

Outcomes

- The gauge coupling unification is achieved with the kinetic mixing between $U(1)_Y$ and $U(1)_H$ of ~ 0.37 .
- The unification behavior does not depend on the size of hidden gauge coupling
- Also, the existence of matter fields charged under $SU(5)$ and/or $U(1)_H$ does not affect the unification.
- We can simply includes the $SU(5)$ multiplets at an arbitrary scale.
- For instance, we can accommodate the PQ symmetric solution to the strong CP problem.

Outcomes

- Due to the kinetic mixing of ~ 0.37 , a hidden matter has an electric charge.
- For small hidden gauge coupling
→ **Minicharged dark matter (having tiny electric charge)**
- **The stability of the DM** is ensured by the unbroken $U(1)_H$.

Minicharged Dark Matter

- For tiny g_H , the hidden matter with $Q_Y=0$ has a non-zero but tiny electric charge in the canonical basis

$$\bar{\Psi} \gamma_\mu (g'_H A'^\mu_H) \Psi = \bar{\Psi} \gamma_\mu \left(-\frac{g_H \chi}{\sqrt{1 - \chi^2}} A_Y + g_H A^\mu_H \right) \Psi$$

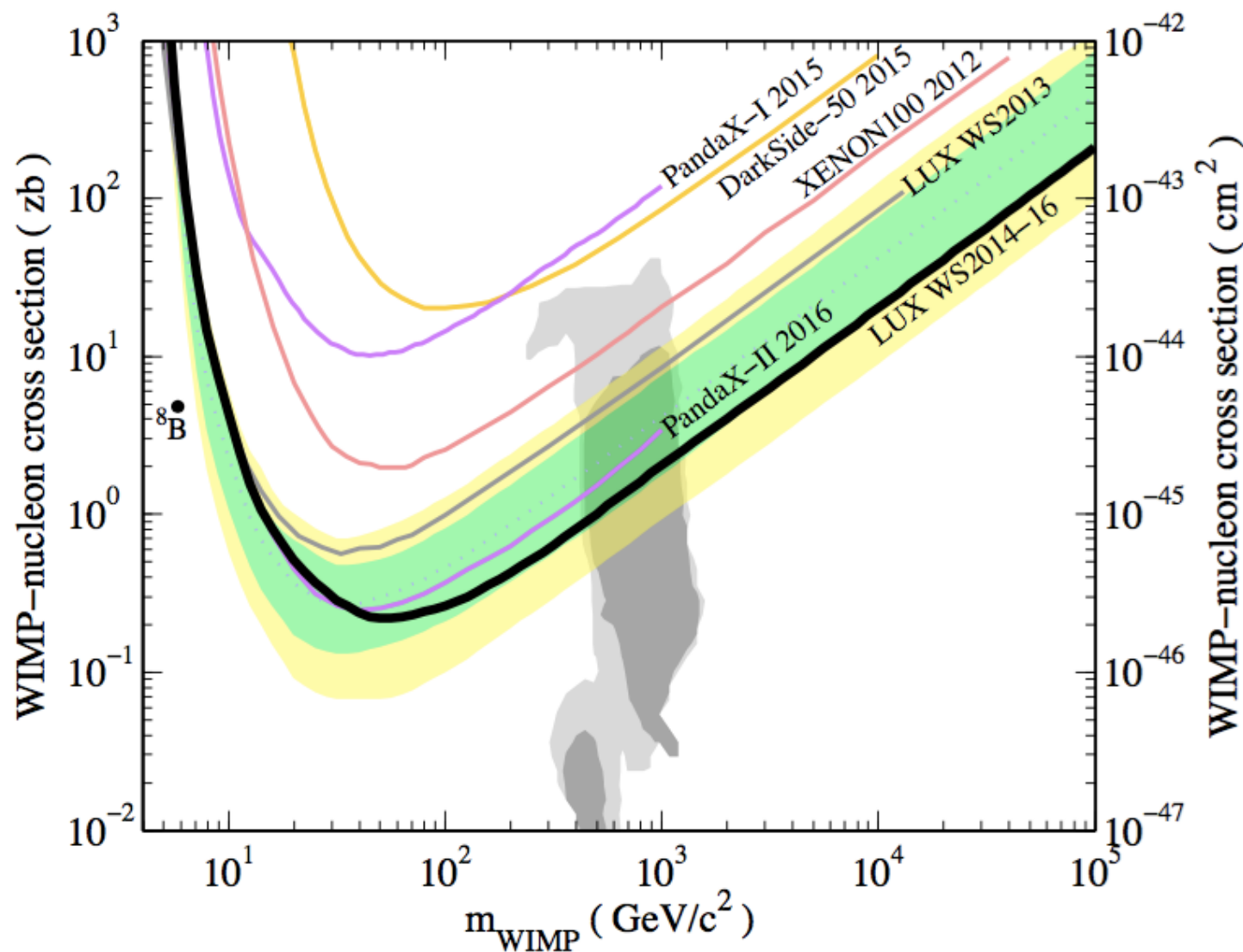
$$q_e = -\frac{g_H}{g_Y} \frac{\chi}{\sqrt{1 - \chi^2}}$$

→ Stable; a candidate for dark matter
(minicharged dark matter)

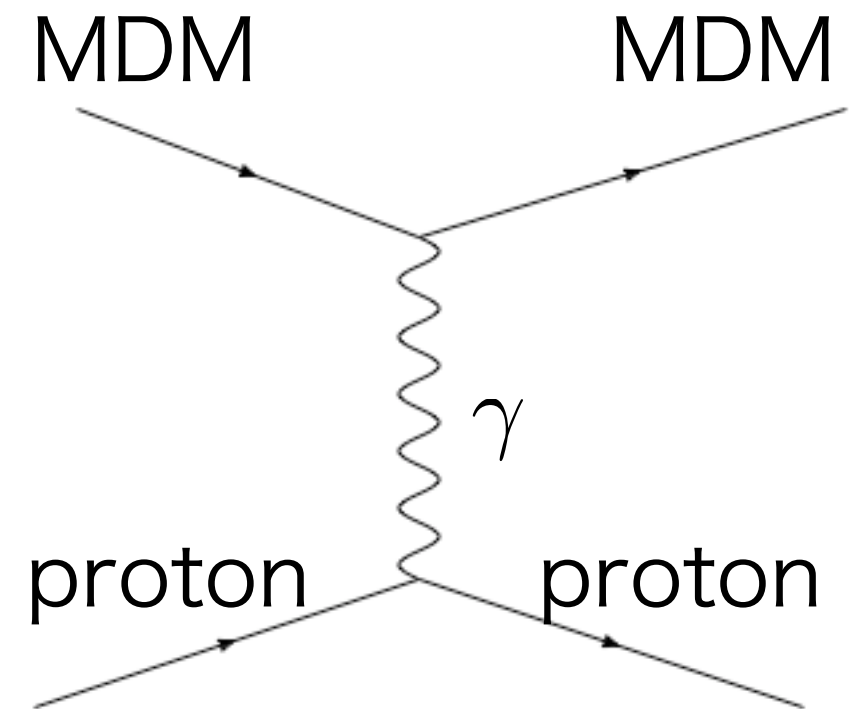
This minicharged DM is constrained by cosmological and astrophysical observations.

Constraints on minicharged DM

Stringent constraint comes from the direct detection experiment.



[Lux, 2016]

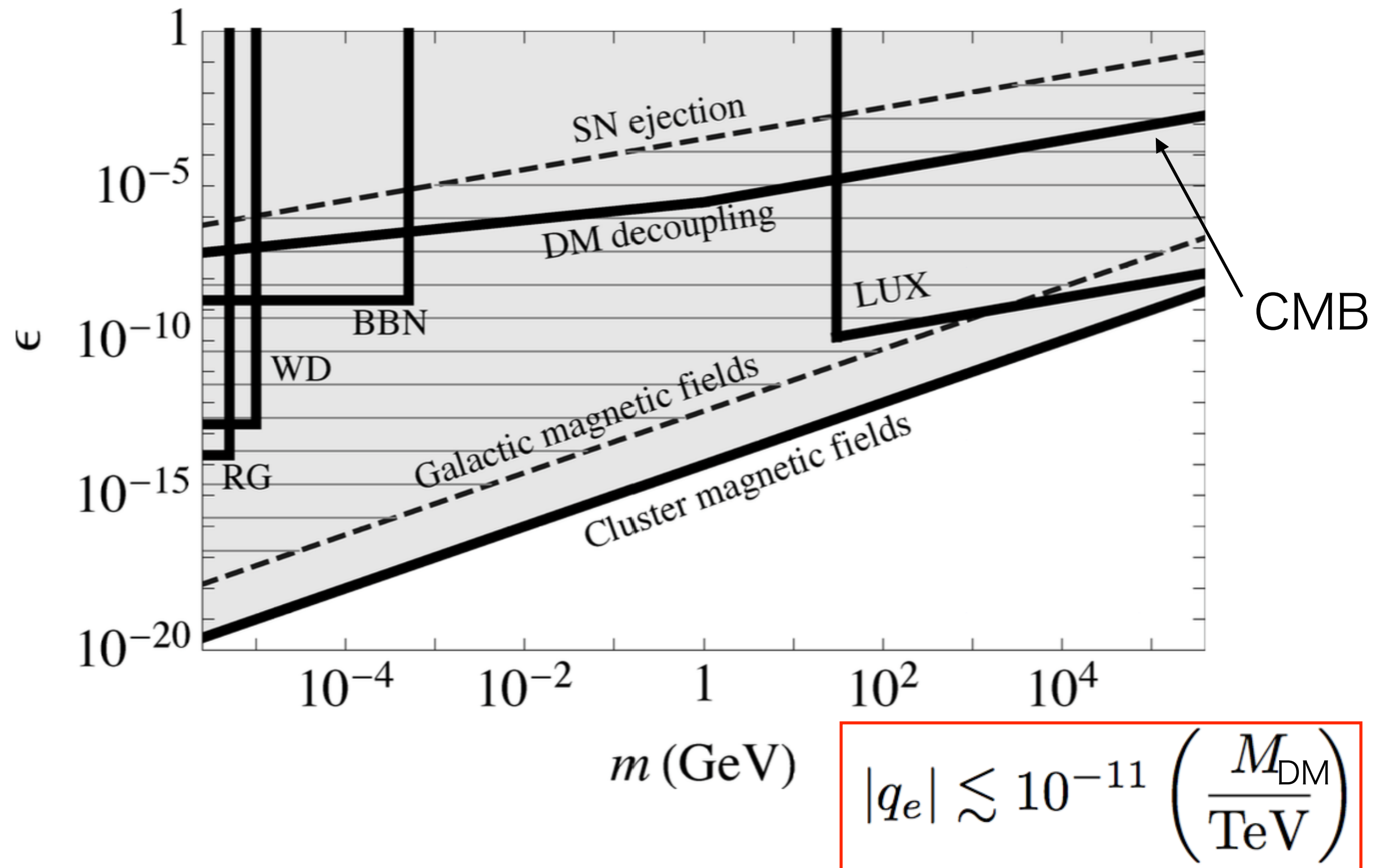


$$|q_e| \lesssim 3.6 \times 10^{-10} \left(\frac{M_{\text{DM}}}{1 \text{ TeV}} \right)^{1/2}$$

g_H should be tiny as 10^{-10}

[Nobile, Nardecchia, Panci, 2015]

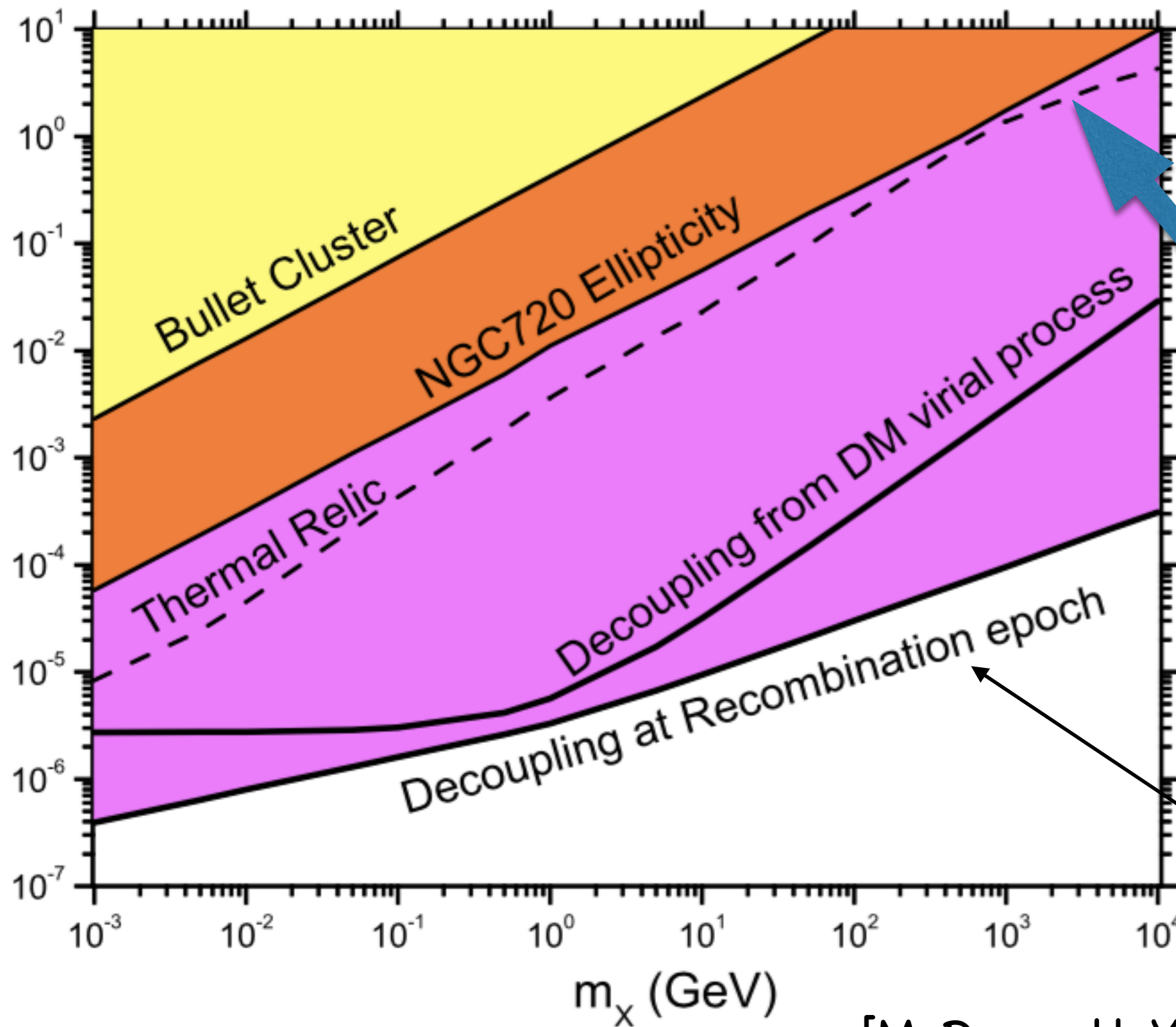
Magnetic fields in galaxy clusters provide more stringent constraint for the small mass region. (changing DM profile)



[Kadota, Sekiguchi, Tashiro, 1602.04009]

electric charge

ω



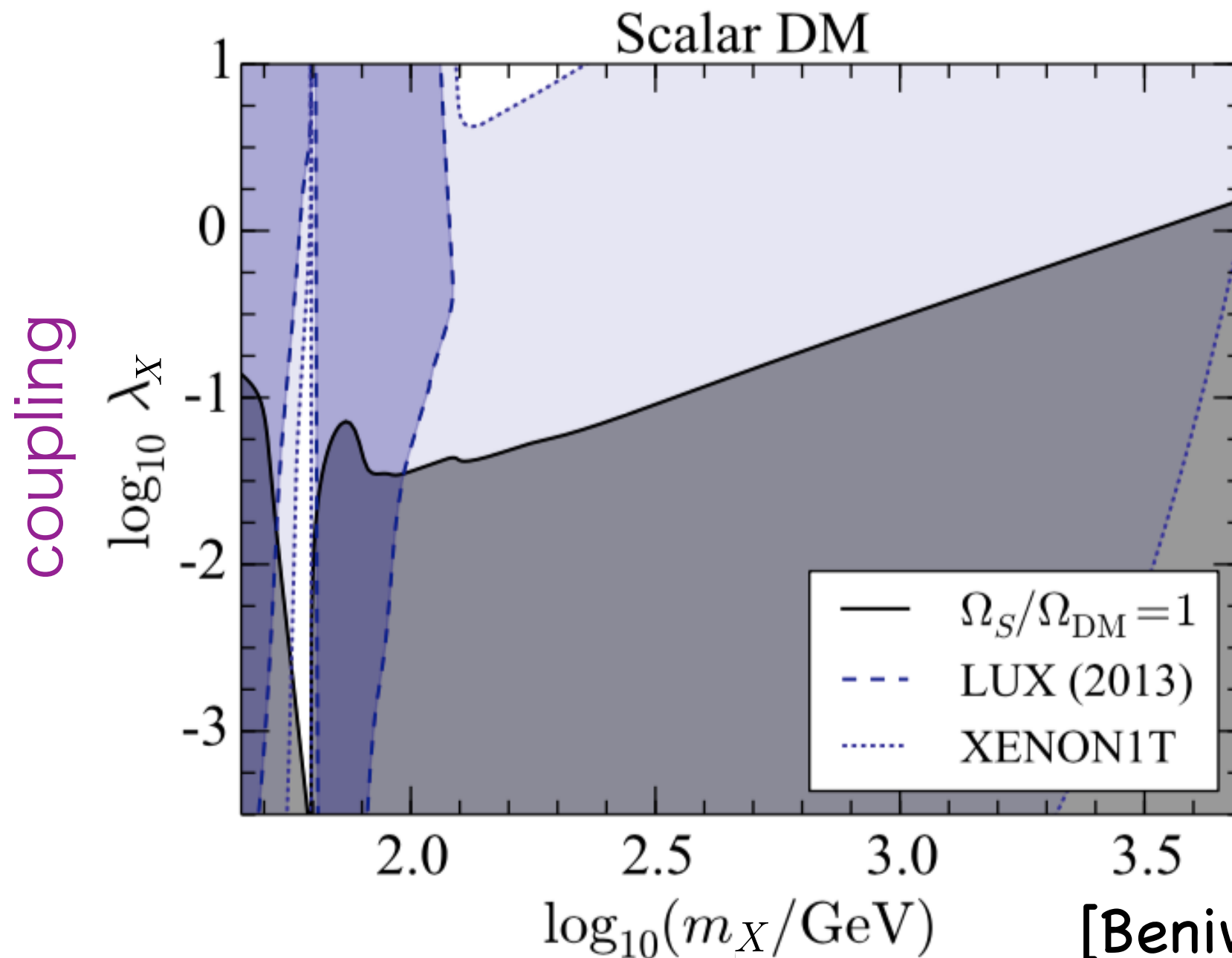
DM mass

[McDermott, Yu, Zurek, 2010]

Required size of the electric charge is much larger than the constraints from LUX, CMB and cluster magnetic fields.

Correct thermal relic abundance can be obtained if there is a Higgs portal interaction (or non-thermal production)

$$\mathcal{L} \ni -m_X^2 |X|^2 + \lambda_X |H|^2 |X|^2 \quad \text{or} \quad -m_Y \bar{\Psi}_Y \Psi_Y + \frac{\lambda_Y}{\Lambda_Y} |H|^2 \bar{\Psi}_Y \left(\frac{1 - \gamma_5}{2} \right) \Psi_Y + h.c.,$$



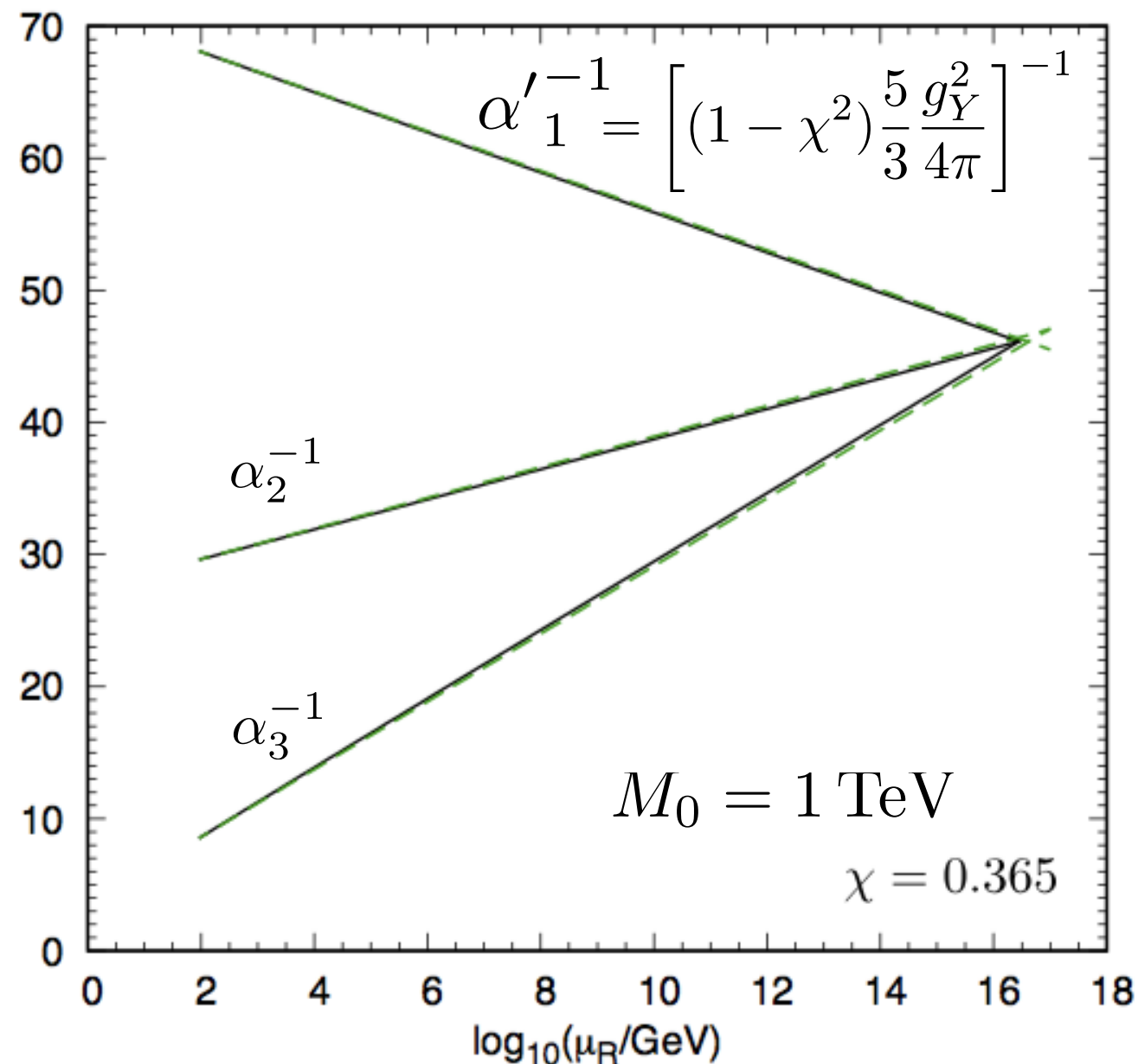
$m_X > \sim 0.6 \text{ TeV}$
with Lux2016

[Beniwal et.al. 2015]

Summary

- Massless hidden photon with the kinetic mixing term can achieve the gauge coupling unification **without Supersymmetry**
- The unification scale is rather high as $10^{16.5}$ GeV (no rapid proton decay)
- For tiny g_H , we have the interesting dark matter candidate: **minicharged dark matter**
- Stability is ensured by the hidden gauge symmetry
- The correct relic abundance is obtained via the Higgs portal.

SM + hidden photon + kinetic mixing → GUT + minicharged dark matter!



green dashed: one-loop

black solid: two-loop

**Thank you for your
attention!**

The RG equations

RGEs in the canonical basis

$$\begin{aligned}\frac{dg_Y}{dt} &= \frac{1}{16\pi^2}(b_Y g_Y^3 + b_H g_Y g_{\text{mix}}^2 + 2b_{\text{mix}} g_Y^2 g_{\text{mix}}), \\ \frac{dg_H}{dt} &= \frac{1}{16\pi^2} b_H g_H^3, \\ \frac{dg_{\text{mix}}}{dt} &= \frac{1}{16\pi^2}(b_Y g_{\text{mix}} g_Y^2 + 2b_H g_{\text{mix}} g_H^2 + b_H g_{\text{mix}}^3 + 2b_{\text{mix}} g_Y g_H^2 + 2b_{\text{mix}} g_Y g_{\text{mix}}^2),\end{aligned}$$

$$g_{\text{mix}} = -\frac{g_H \chi}{\sqrt{1 - \chi^2}}$$

[Babu, Kolda, March-Russell, 1996]

$t = \ln \mu_R$ (μ_R is a renormalization scale)

$$b_Y = \frac{41}{6} + \frac{4}{3} \sum_i Q_i^2, \quad b_H = \frac{4}{3} \sum_i q_{H_i}^2, \quad b_{\text{mix}} = \frac{4}{3} \sum_i Q_i q_{H_i}$$

RGE of $U(1)_Y$ is trivial only if $b_H=0$ and $b_{\text{mix}}=0$

i.e. if there are no hidden matters

The RG equations

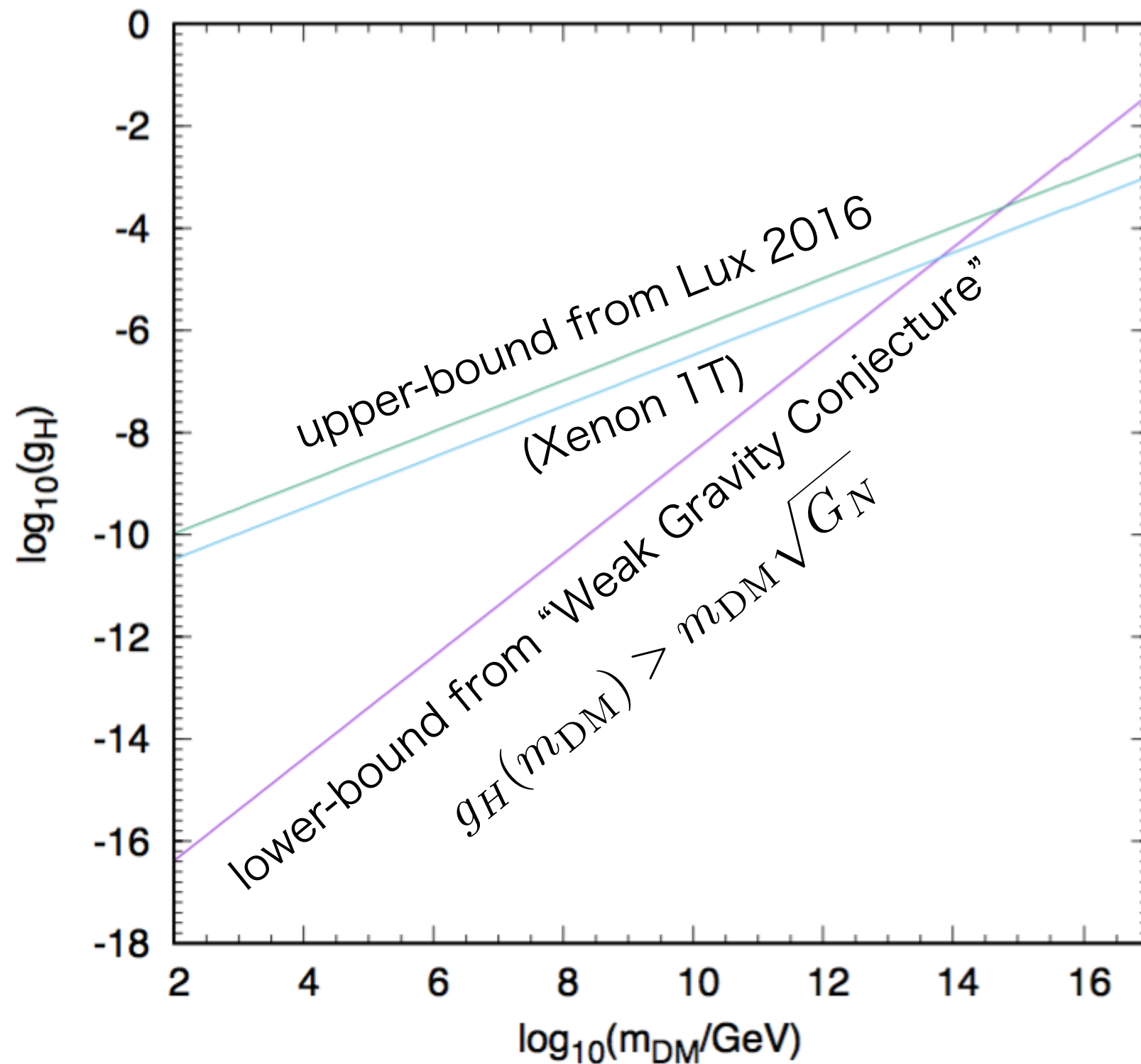
RGEs in the original basis are very simple:

$$\begin{aligned}\frac{dg'_Y}{dt} &= \frac{1}{16\pi^2} b_Y g'^3_Y, \\ \frac{dg_H}{dt} &= \frac{1}{16\pi^2} b_H g^3_H, \\ \frac{d\chi}{dt} &= \frac{1}{16\pi^2} [\chi(b_Y g'^2_Y + b_H g^2_H) - 2b_{\text{mix}} g'_Y g_H]\end{aligned}$$

$$g_Y = \frac{g'_Y}{\sqrt{1 - \chi^2}}$$

RGE running is trivial at the one-loop level in this basis.

Bound from WGC



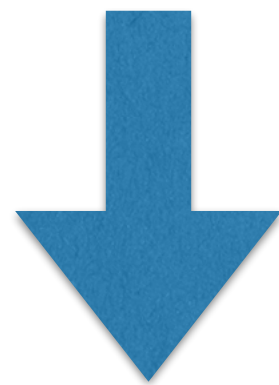
The small coupling is close to the lower-bound from the WGC.

[WGC: Arkani-Hamed, Motl, Nicolis, Vafa, 2006]

Origin of the Kinetic mixing

The kinetic mixing arises via a higher dimensional operator

$$\mathcal{L} \ni \frac{k_0}{M_*} \text{Tr}(\langle \Sigma_{24} \rangle F_{5\mu\nu}) F_H^{\mu\nu},$$

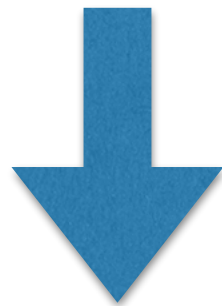


$$\langle \Sigma_{24} \rangle = \frac{v_{24}}{2\sqrt{15}} \begin{pmatrix} 2 & & & & \\ & 2 & & & \\ & & 2 & & \\ & & & -3 & \\ & & & & -3 \end{pmatrix}$$

$$\chi = -k_0 \frac{v_{24}}{M_*},$$

There may exists an operator

$$\mathcal{L} \ni \frac{k_1}{2M_*} \text{Tr}(\langle \Sigma_{24} \rangle F_{5\mu\nu} F_5^{\mu\nu}),$$



$$\frac{\Delta\alpha_1^{-1}}{\alpha_5^{-1}} = \sqrt{\frac{1}{60}} \frac{k_1}{k_0} \chi, \quad \frac{\Delta\alpha_2^{-1}}{\alpha_5^{-1}} = \sqrt{\frac{3}{20}} \frac{k_1}{k_0} \chi, \quad \frac{\Delta\alpha_3^{-1}}{\alpha_5^{-1}} = -\sqrt{\frac{1}{15}} \frac{k_1}{k_0} \chi,$$

We can not realize the consistent GUT only with the above threshold correction, due to the proton decay constraint

$$k_1/k_0 \lesssim 6 \rightarrow (\Delta\alpha_2^{-1} - \Delta\alpha_3^{-1})/\alpha_5^{-1}, \text{ is within 5\% level}$$