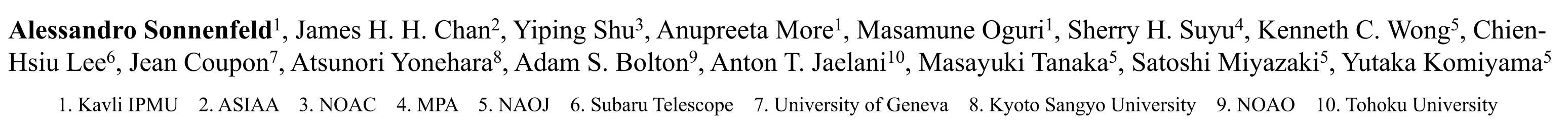
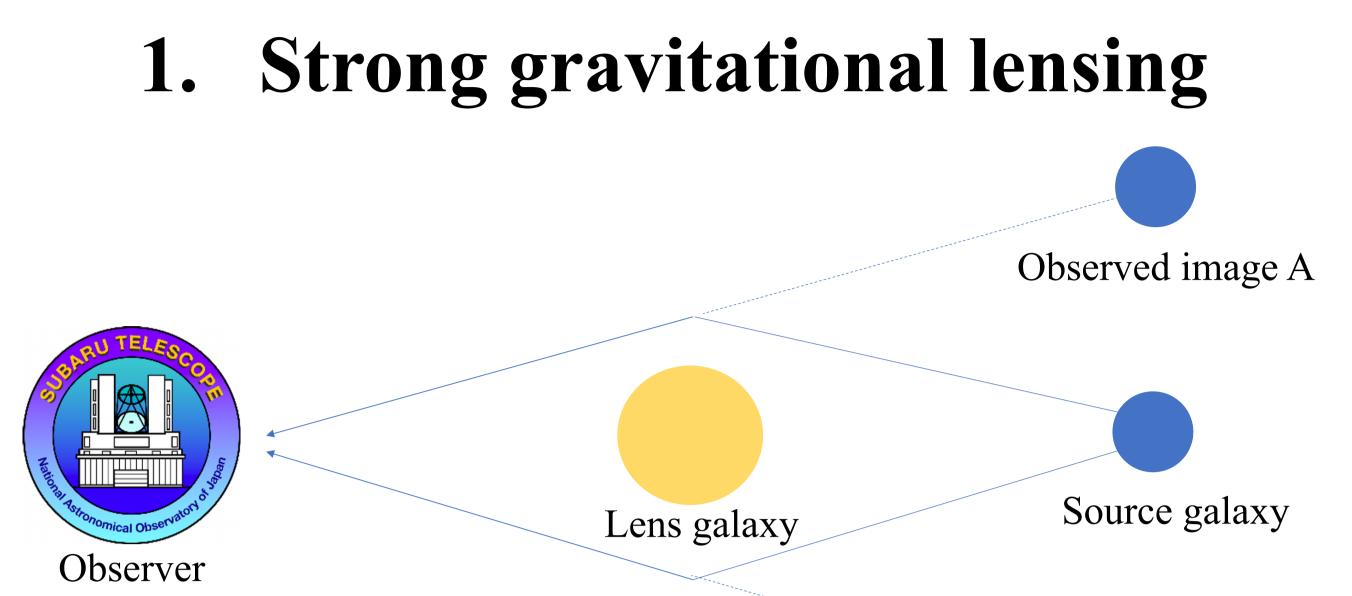


# **Survey of Gravitationally-lensed Objects in HSC Imaging:** SuGOHI





### **3. New lenses in the HSC survey**

The Hyper Suprime-Cam (HSC), recently installed on Subaru, is currently surveying an area of 1400 square degrees of sky, with excellent image quality and great depth, ideal conditions for a lens search.

We used YattaLens, among with two other lens search algorithms, to look for lenses among 37,000 massive galaxies with BOSS spectroscopy in ~400 square degrees of imaging data from the HSC survey. We found:

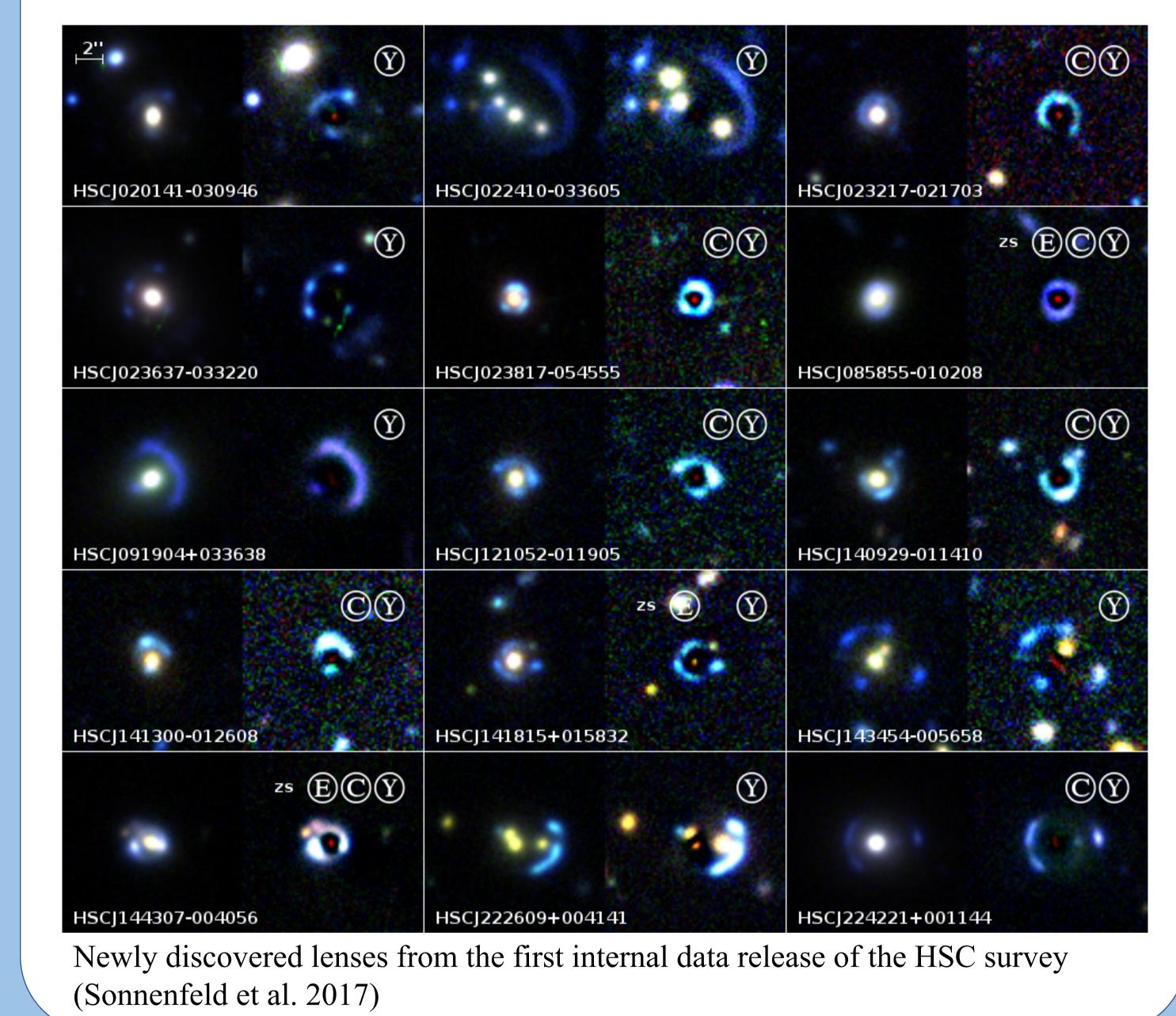
#### Observed image B

Schematics of a strong lens system: light from a background galaxy is deflected by a foreground massive object, such that multiple images of the same galaxy are formed

- Very powerful tool to study the structure of galaxies: few % precision on total projected mass within Einstein radius
- Rare phenomenon: requires almost perfect alignment between lens and source. Roughly one in a thousand galaxies is a lens.
- Only a few hundred galaxy-scale strong lens systems currently known
- Can be used to answer the following questions:
  - How do massive elliptical galaxies grow in time?
  - How does their stellar content change as a result of mergers with smaller galaxies?
  - How does the distribution of dark matter respond to the infall of gas or the presence of a central black hole?

- 15 grade A (definite) lenses
- 36 grade B (probable) lenses
- 282 grade C (possible) lenses

These lenses form the first sample of the Survey of Gravitationally-lensed Objects in HSC Imaging (SuGOHI)



• More lenses are needed, particularly at redshift z > 0.5.

## 2. Automatic detection of strong lenses

We developed a lens finding algorithm, named YattaLens, to look for lenses among massive galaxies in multi-band imaging data. YattaLens articulates itself over the following steps:

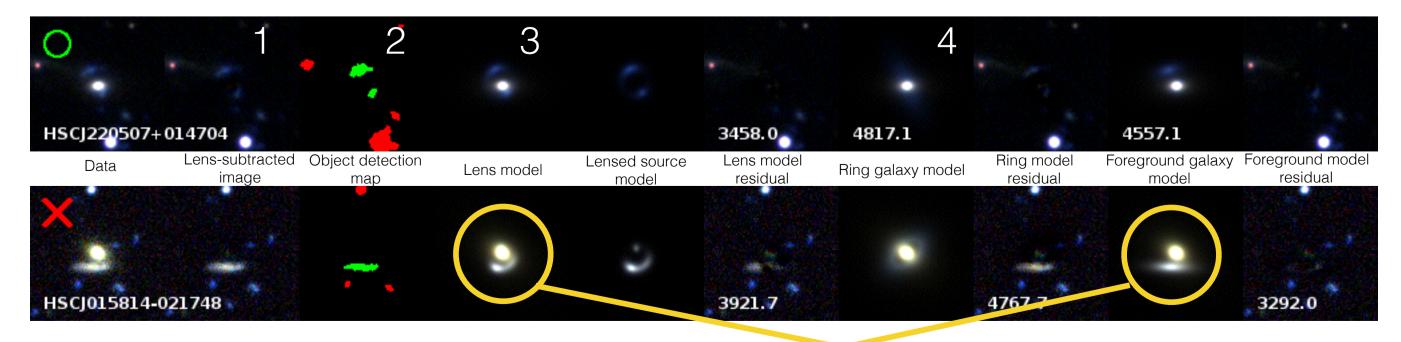
- 1. For each galaxy, fits a model surface brightness profile to remove from the image the contribution of the lens light
- 2. Looks for blue tangentially elongated objects around the galaxy center
- 3. If an arc is detected (roughly, this happens for  $\sim 10\%$  of the objects), fits a lens model to the image
- 4. Compares the best-fit lens model with alternative non-lens

4. Future prospects

- Current size of SuGOHI sample: ~100 lenses  $\bullet$
- By the end of HSC survey (2019): a few hundred. Will be **the** lacksquarelargest sample of lenses from a single survey.
- Ongoing spectroscopic follow-up with X-Shooter on VLT
- Will allow us to study the evolution in the inner structure of  $\bullet$

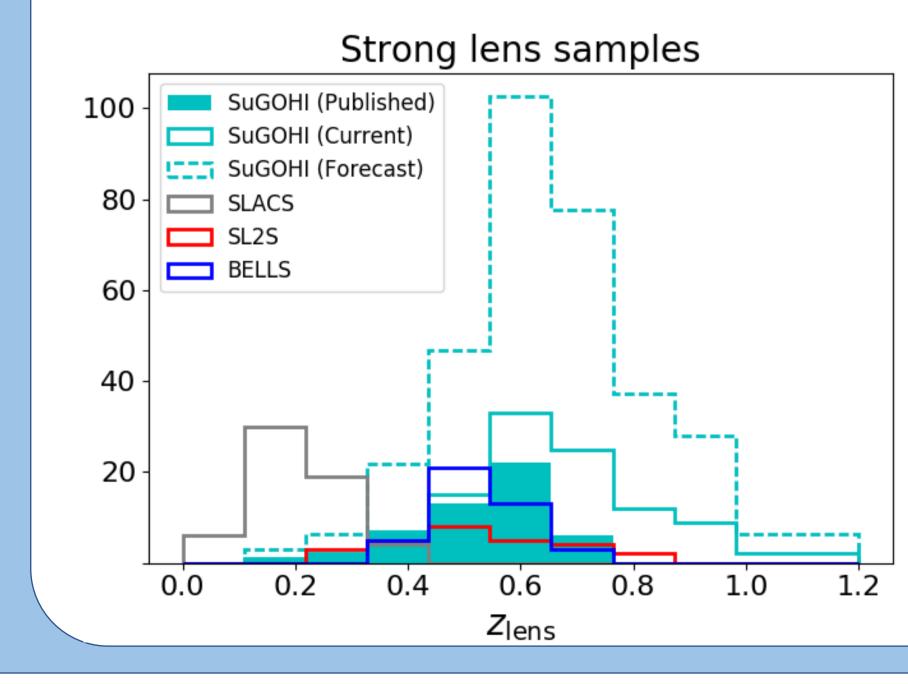
### models, to determine the likelihood of it being a lens

**References:** 

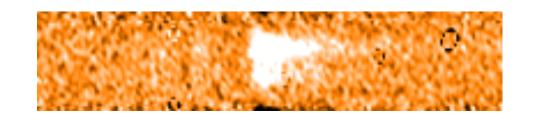


This system is better described by a model with a foreground object, compared to a model with a lensed background source, therefore the candidate is rejected

massive galaxies from z=1 to the present



*Left*: redshift distribution of lenses in the SuGOHI sample, compared to existing strong lens samples. *Bottom*: X-Shooter 2d spectrum of a doubly imaged Ly-alpha emission from a lensed source in the SuGOHI sample.



Sonnenfeld A., Chan H. H. J., Shu, Y., et al. 2017, PASJ in press, arXiv:1704.01585