Search for physics beyond Standard Model using ultra-radio-pure Nal(TI) crystals

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Dark Matter search and test of DAMA/LIBRA observations

Weakly interacting massive particles (WIMPs) are one of the most promising dark matter candidates. The direct detection of WIMPs is performed by measuring the recoil energy deposited by elastic and inelastic scattering between WIMPs and target nucleus. One of the most promising signal for WIMPs is the **annual modulation in energy spectrum due to** Earth's motion around the Sun.









2. Precise measurement of the Weinberng angle

Coherent Elastic Neutrino-Nucleus Scattering (CEvNS) is a **fundamental process** recently observed by the COHERENT collaboration. The CEvNS cross-section can be described with the formula

$$\sigma_{coh} \sim \frac{G_f^2 E^2}{4\pi} (Z(4 \sin^2 \theta_w - 1) + N)^2$$



where G_f is the Fermi constant, E is the energy of the neutrino, ϑ_w is the Weinberg angle, Z is the number of protons and N is the numbers of neutrons in the target nucleus. Since σ_{tot} depends on Weinberg angle ϑ_{W} , it opens a way to measure weak mixing angle $\sin^2 \vartheta_w$ using massive detectors.



COHERENT detectors at the SNS

The first possible annual modulation observation was reported by DAMA/LIBRA using a large volume and highly radiopure Nal(Tl) scintillators. They claim to observe an annual modulating signal in event rates between 2 keV_{ee} and 6 keV_{ee}.

We aim to test the annual modulation signal observed by DAMA/LIBRA using our recently developed highly radiopure NaI(TI) crystal scintillators. Moreover, we will try to find possible correlations between the several keV signals in NaI(TI) detectors, the radon activity in the mine air and the neutron flux. Experiment to search for dark matter is currently in development and will be performed in Kamioka underground laboratory.

We investigate possibility to use our recently developed ultra-radio-pure NaI(TI) crystals to measure Weinberg angle at the Spallation Neutron Source (SNS) in Oak Ridge, Tennessee. The SNS provides an intense flux of neutrinos in the few tens-of-MeV range, while our ultra-radio-pure Nal(Tl) crystals allow to observe events at low threshold (below 1keV) and has high light yield.

Development of ultra-radio-pure Nal(Tl) crystals 3.

The development of **highly radiopure NaI(TI) crystal scintillator** is a **crucial point** for the experiments. We have performed many tests and R&D to successfully develop highly radiopure NaI(TI) scintillators. The radioactive impurities in raw powder were removed by using **ion exchange resin**. The contamination from the crucible was prevented by selecting the **highest** purity of the crucible material. The Bridgman method was applied to grow the NaI(TI) crystals.



Comparison of radioactive contamination of **Nal(Tl) crystals** for different experiments

Impurity	DAMA/LIBRA	DM-ICE	ANAIS	KIMS	Our experiment*
natK [ppb]	< 20	558	20 ~ 46	40 ~ 50	125
Th-chain [ppt]	0.5 ~ 7.5	13	0.8 ± 0.3	0.5 ± 0.3	0.3 ± 0.5
²²⁶ Ra [µBq/kg]	21.7 ± 1.1	900	10 ± 0.2	< 1	58 ± 4
²¹⁰ Pb [µBq/kg]	24.2 ± 1.6	1500	600 ~ 800	470 ± 10	30 ± 7

* Radioactive impurity levels are presented for \emptyset 4 × 3 inch crystal

• K reduction resin

technique

Improved purification

found inside grease

Cryostat window

Al thickness 0.8 mm

Ge crystal

2016

Ø 4 × 3 inch





2014-2015 Ø 3 × 3 inch

- Ra reduction resin
- N₂ bubbling on purification
- OFHC copper housing

• Rehoused after shock absorber was the main source of radioimpurity





2017 Ø 5 × 5 inch

- High-purity graphite crucible
- TI concentration was controlled
- Improved purification technique for Pb and K • Optimized drying

Experimental facility at Kamioka mine

Experimental facility for Nal(Tl) crystals is located in KamLAND area of Kamioka underground laboratory at the depth of 2700m.w.e. It consists of the several detectors located in the Class-10 clean room.



Nal(TI) crystals were tested in the lowbackground set-up. Crystal housed in a copper holder and viewed by 3" PMT Hamamatsu R11065-20-mod2 was surrounded by a temporary shield made of **5 cm of OFHC copper** and 20 cm of high-purity old lead. The shielding was purged by the pure nitrogen gas to reduce the concentration of radon









Energy [keV

Development of

new ultra-low-

background 4"

PMT is in progress

Radiopurity of the detector materials is controlled by using Canberra GC6520 **HPGe detector**



New shielding with fresh

OFHC copper

and cleaned

Seasonal variations of thermal neutron flux are measured by using ⁶Li loaded ZnS(Ag) EJ-426 scintillator sheets with the size of 50x50 cm

ZnS(Ag)+ ⁶Li

Vindow electroo

Ge dead laver 0.48 mm

ner electrod e dead laver 0.3 µm

Outer electrode

dead laver 0.48 mm



Radon activity variations in the underground air will be measured using constructed highresolution radon gas detector







Future prospects on dark matter search

We plan to finish all of the R&D and preparations during the next year and to start the experiment with 9 Nal(TI) crystal modules in 2019. Subsequent scaling of the detector can be done in a couple of years. Large-scale experiment using KamLAND detector can be considered in a future.



Conclusions 6.

- We have successfully developed ultra-radiopure Nal(TI) crystals with radioactive contamination levels close to DAMA crystals. New Ø 5 × 5 inch Nal(Tl) crystal was produced recently and will be tested using upgraded low-background set-up at **Kamioka underground laboratory**
- The experiment to search for dark matter and to test the annual modulation signal of **DAMA/LIBRA is in preparation**. We plan to **start measurements with 54 kg of NaI(TI)** crystal modules in 2019. Subsequent annual modulation test with 250 kg of NaI(TI) crystals is expected to start in 2021
- The application of ultra-radiopure NaI(TI) detectors for the measurement of Weinberg angle at the Spallation Neutron Source is under discussion. Measurement of the CEvNS by using ~15 kg of the developed ultra-radiopure NaI(TI) detectors is under consideration

Kavli IPMU 10th anniversary symposium 16 – 18 October 2017, Kashiwanoha Conference Center, Kavli IPMU, Japan