

The Unique Role of Future High-Energy pp Colliders

Andrea Wulzer



UNIVERSITÀ
DEGLI STUDI
DI PADOVA



Ideology

HEP before the LHC



HEP before the F.C.

Ideology

HEP before the LHC



HEP before the F.C.



Ideology

HEP before the LHC



HEP before the F.C.



Particle physics is not **validation** anymore, rather it is **exploration of unknown territories** *

* Not necessarily a bad thing. Columbus left for his trip just because he had no idea of where he was going !!

Ideology

No single experiment can explore all directions at once ...



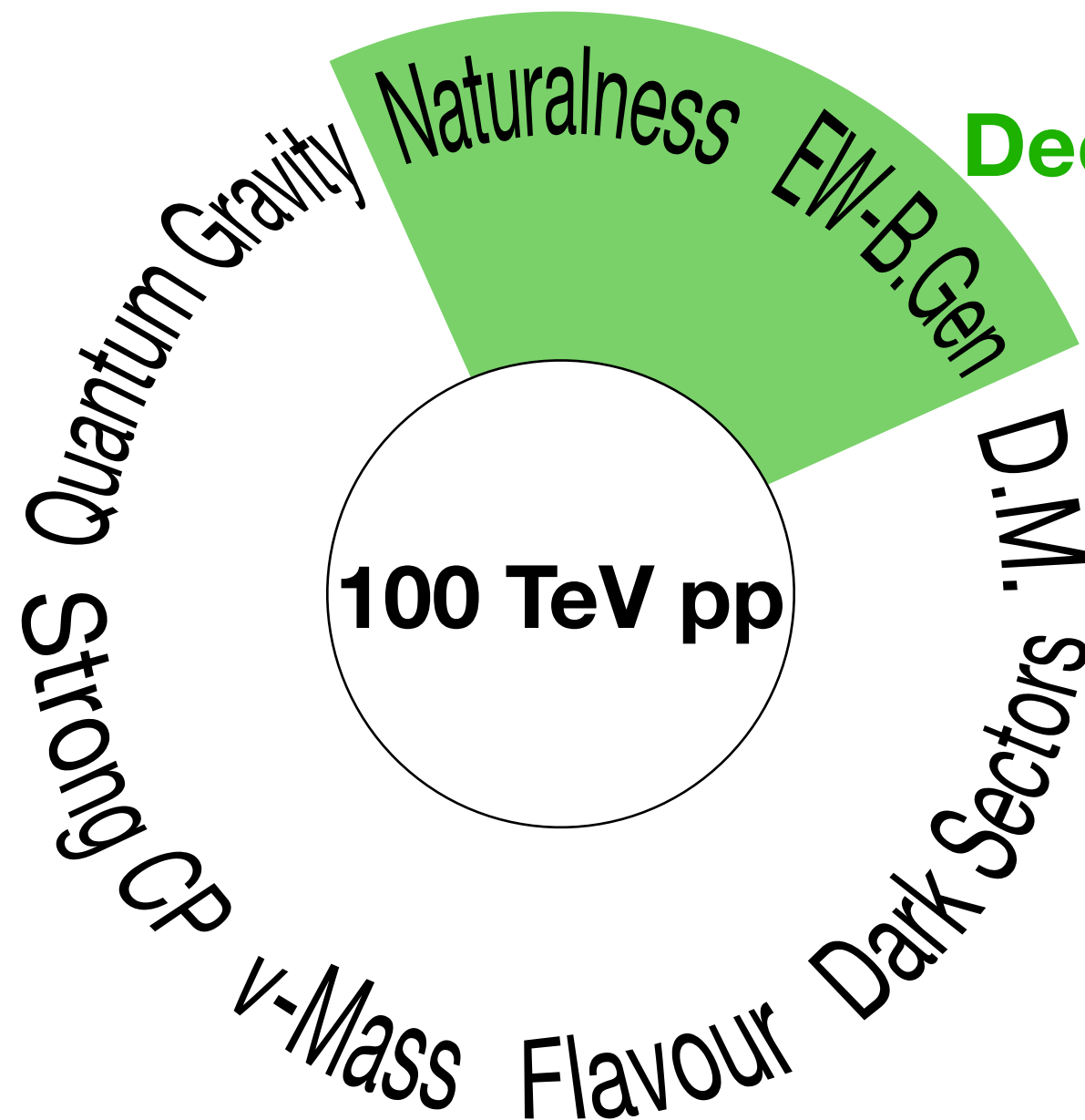
Ideology

No single experiment can explore all directions at once ...
but **many directions** explored by a **100 TeV pp collider**



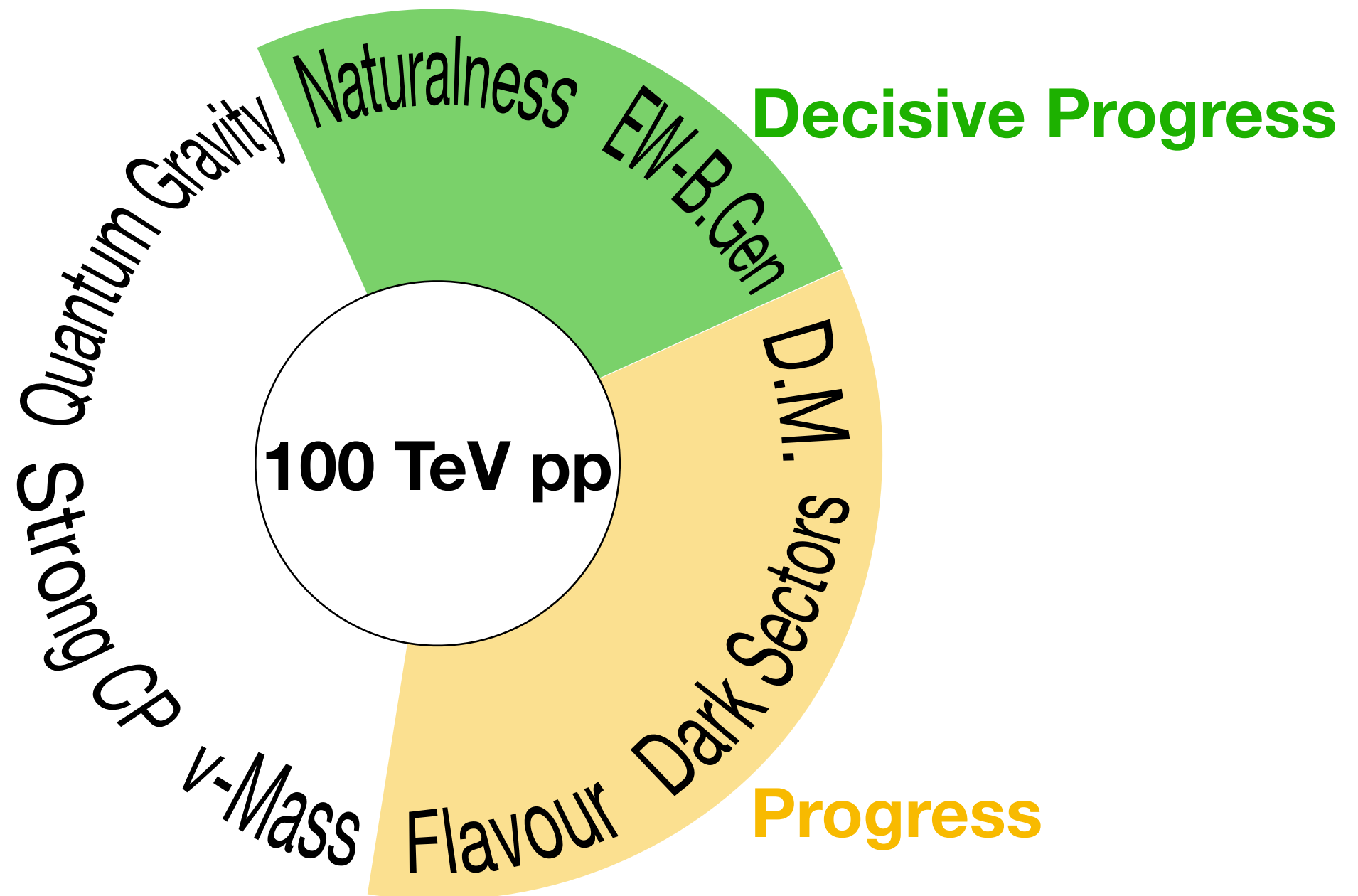
Ideology

No single experiment can explore all directions at once ...
but **many directions** explored by a **100 TeV pp collider**



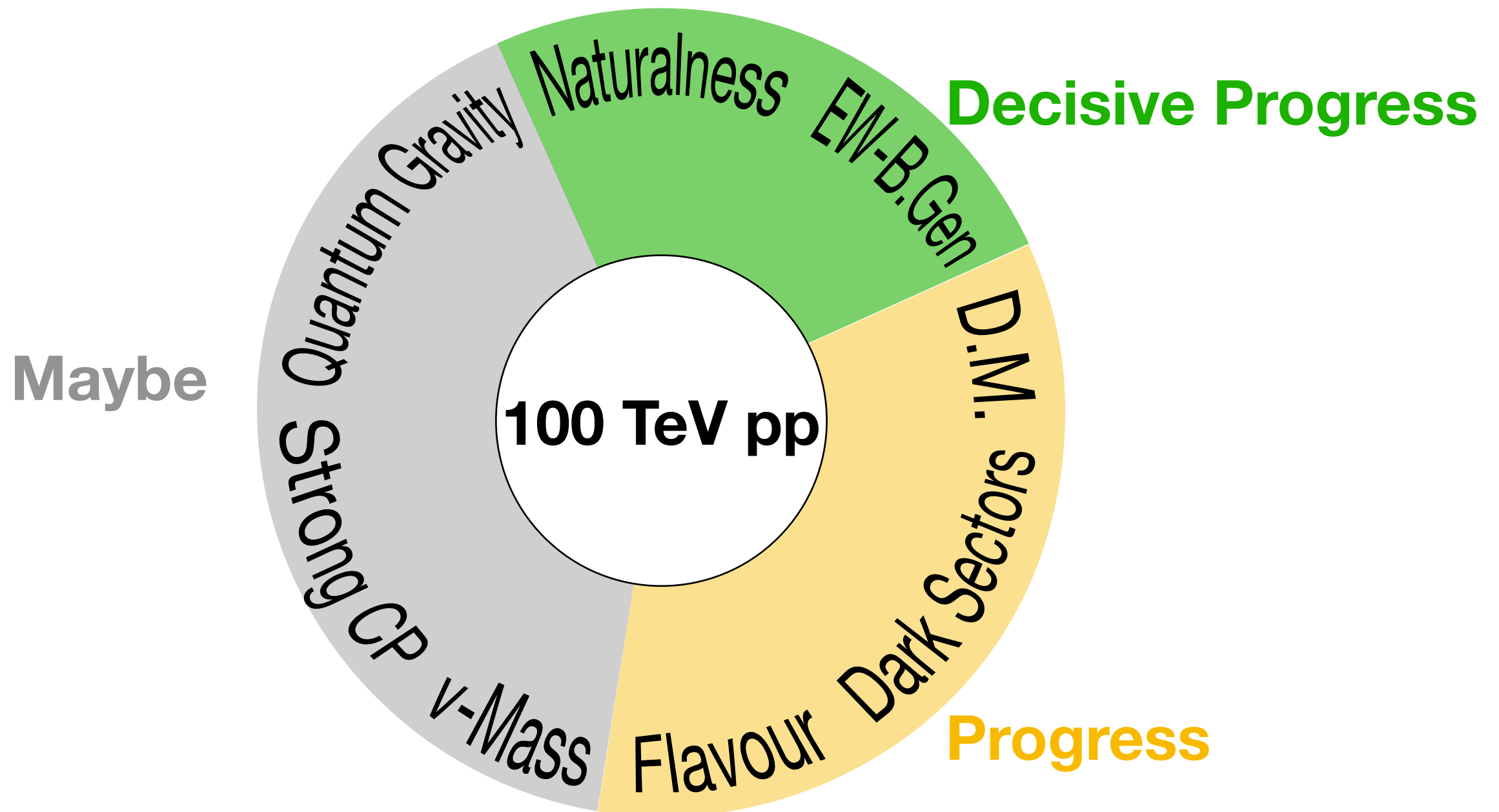
Ideology

No single experiment can explore all directions at once ...
but **many directions** explored by a **100 TeV pp collider**



Ideology

No single experiment can explore all directions at once ...
but **many directions** explored by a **100 TeV pp collider**



Naturalness

“Is m_H Unnatural?” = “Is m_H Unpredictable?”

$$(m_H^2)_{Phys.} = \int_0^\infty F_{true}(E; g_{true})$$

SM Contribution

$$\delta m_H^2 = \frac{3y_t^2}{8\pi^2} \Lambda_{SM}^2$$

$$= \int_0^{\Lambda_{SM}} (\dots) + \int_{\Lambda_{SM}}^\infty (\dots)$$

UV Contribution

$$c \Lambda_{SM}^2$$

Fine Tuning: $\Delta \geq \frac{\delta m_H^2}{m_H^2} \simeq \left(\frac{126 \text{ GeV}}{m_H} \right)^2 \left(\frac{\Lambda_{SM}}{500 \text{ GeV}} \right)^2$

Measures how much Unpredictable m_H is.

Naturalness

“Is m_H Unnatural?” = “Is m_H Unpredictable?”

$$(m_H^2)_{Phys.} = \int_0^\infty F_{true}(E; g_{true})$$

SM Contribution
 $\delta m_H^2 = \frac{3y_t^2}{8\pi^2} \Lambda_{SM}^2$

UV Contribution
 $c \Lambda_{SM}^2$

$$= \int_0^{\Lambda_{SM}} (\dots) + \int_{\Lambda_{SM}}^\infty (\dots)$$

Fine Tuning: $\Delta \geq \frac{\delta m_H^2}{m_H^2} \simeq \left(\frac{126 \text{ GeV}}{m_H} \right)^2 \left(\frac{\Lambda_{SM}}{500 \text{ GeV}} \right)^2$

Measures how much Unpredictable m_H is.

Unnaturalness is a challenge to **Reductionism**

Naturalness

“Is m_H Unnatural?” = “Is m_H Unpredictable?”

$$(m_H^2)_{Phys.} = \int_0^\infty F_{true}(E; g_{true})$$

SM Contribution
 $\delta m_H^2 = \frac{3y_t^2}{8\pi^2} \Lambda_{SM}^2$

UV Contribution
 $c \Lambda_{SM}^2$

$$= \int_0^{\Lambda_{SM}} (\dots) + \int_{\Lambda_{SM}}^\infty (\dots)$$

Fine Tuning: $\Delta \geq \frac{\delta m_H^2}{m_H^2} \simeq \left(\frac{126 \text{ GeV}}{m_H} \right)^2 \left(\frac{\Lambda_{SM}}{500 \text{ GeV}} \right)^2$

Measures how much Unpredictable m_H is.

Unnaturalness is a challenge to **Reductionism**

Dramatic paradigm shift. E.g. Anthropic or Dynamical

Naturalness

$$\Delta \geq \frac{\delta m_H^2}{m_H^2} \simeq \left(\frac{126 \text{ GeV}}{m_H} \right)^2 \left(\frac{\Lambda_{\text{SM}}}{500 \text{ GeV}} \right)^2$$

LHC may push conventional Natural models to

$$\Lambda_{\text{SM}} \gtrsim 2 \text{ TeV} \longrightarrow \Delta \gtrsim 10$$

Naturalness

$$\Delta \geq \frac{\delta m_H^2}{m_H^2} \simeq \left(\frac{126 \text{ GeV}}{m_H} \right)^2 \left(\frac{\Lambda_{\text{SM}}}{500 \text{ GeV}} \right)^2$$

LHC may push conventional Natural models to

$$\Lambda_{\text{SM}} \gtrsim 2 \text{ TeV} \longrightarrow \Delta \gtrsim 10$$

Still Naturalness might be there in the form of:

Partial Unnaturalness

$$\Delta \sim 100$$



$$\Lambda_{\text{SM}} \sim 5 \text{ TeV}$$

Neutral Naturalness

Naturalness

$$\Delta \geq \frac{\delta m_H^2}{m_H^2} \simeq \left(\frac{126 \text{ GeV}}{m_H} \right)^2 \left(\frac{\Lambda_{\text{SM}}}{500 \text{ GeV}} \right)^2$$

LHC may push conventional Natural models to

$$\Lambda_{\text{SM}} \gtrsim 2 \text{ TeV} \longrightarrow \Delta \gtrsim 10$$

Still Naturalness might be there in the form of:

Partial Unnaturalness

$$\Delta \sim 100$$



$$\Lambda_{\text{SM}} \sim 5 \text{ TeV}$$

Neutral Naturalness

$$\Delta \sim \text{few}$$



$$\Lambda_{\text{SM}}^{\text{neut.}} \lesssim 1 \text{ TeV}$$

Naturalness

$$\Delta \geq \frac{\delta m_H^2}{m_H^2} \simeq \left(\frac{126 \text{ GeV}}{m_H} \right)^2 \left(\frac{\Lambda_{\text{SM}}}{500 \text{ GeV}} \right)^2$$

LHC may push conventional Natural models to

$$\Lambda_{\text{SM}} \gtrsim 2 \text{ TeV} \longrightarrow \Delta \gtrsim 10$$

Still Naturalness might be there in the form of:

Partial Unnaturalness

$$\Delta \sim 100$$



$$\Lambda_{\text{SM}} \sim 5 \text{ TeV}$$

Neutral Naturalness

$$\Delta \sim \text{few} \longrightarrow \Lambda_{\text{SM}}^{\text{col.}} \sim 5 \text{ TeV}$$



$$\Lambda_{\text{SM}}^{\text{neut.}} \lesssim 1 \text{ TeV}$$

Naturalness

$$\Delta \geq \frac{\delta m_H^2}{m_H^2} \simeq \left(\frac{126 \text{ GeV}}{m_H} \right)^2 \left(\frac{\Lambda_{\text{SM}}}{500 \text{ GeV}} \right)^2$$

LHC may push conventional Natural models to

$$\Lambda_{\text{SM}} \gtrsim 2 \text{ TeV} \longrightarrow \Delta \gtrsim 10$$

Still Naturalness might be there in the form of:

Partial Unnaturalness

$$\Delta \sim 100$$



$$\Lambda_{\text{SM}} \sim 5 \text{ TeV}$$

Neutral Naturalness

$$\Delta \sim \text{few} \longrightarrow \Lambda_{\text{SM}}^{\text{col.}} \sim 5 \text{ TeV}$$



$$\Lambda_{\text{SM}}^{\text{neut.}} \lesssim 1 \text{ TeV}$$

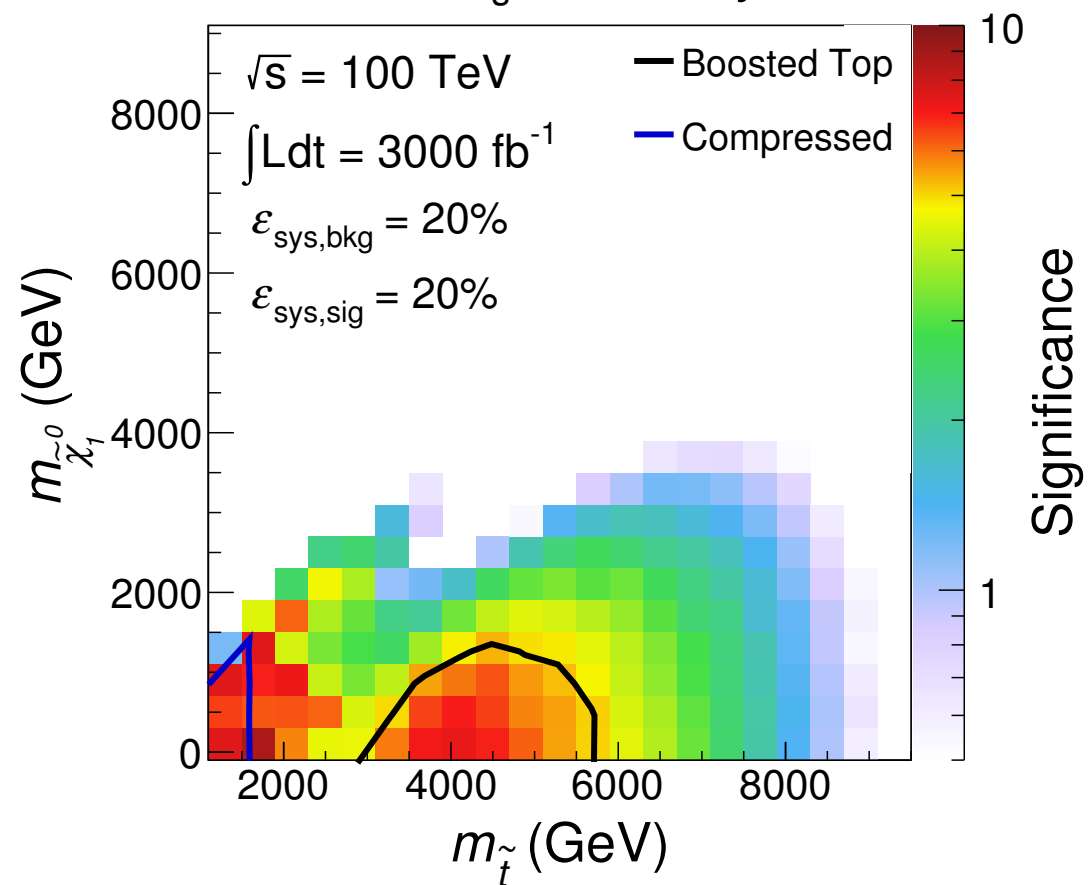
Need 5 TeV reach on ordinary Top Partners

Naturalness

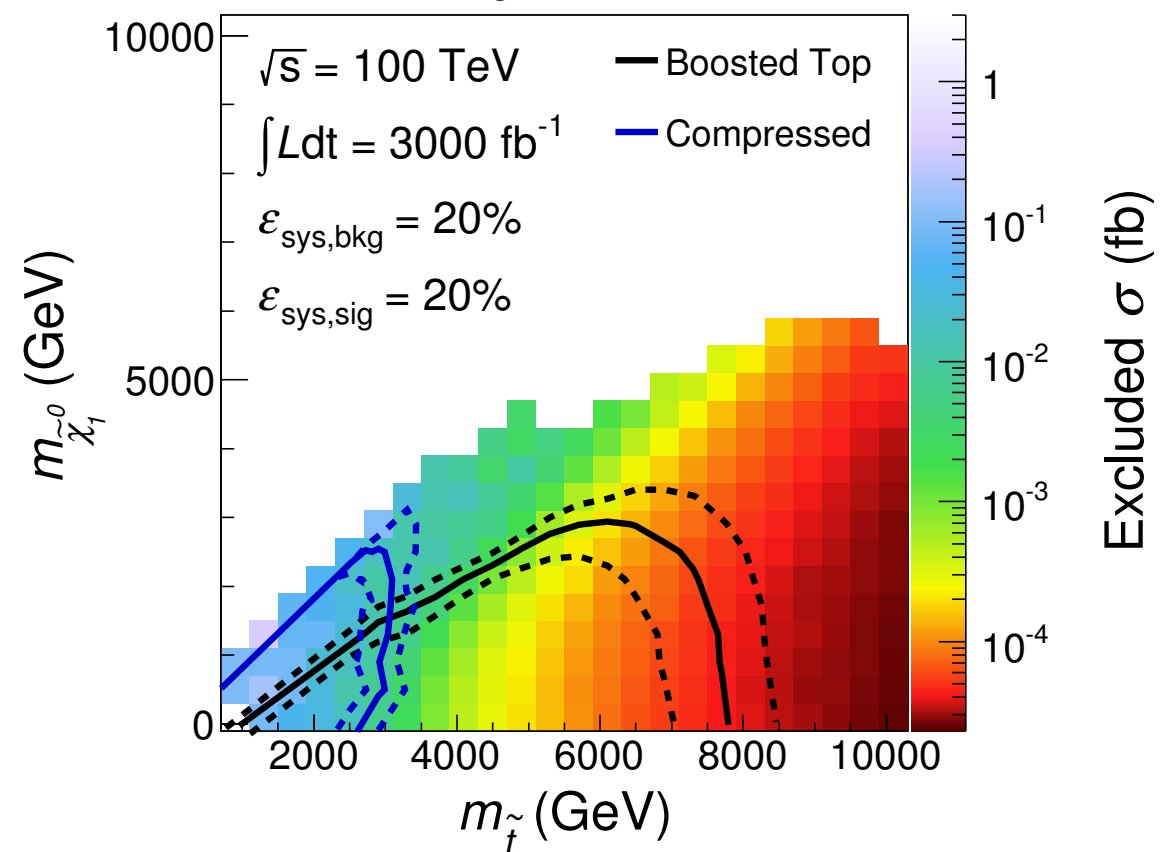
Stop to top + Neutralino:

[arXiv:1406.4512]

CL_s Discovery



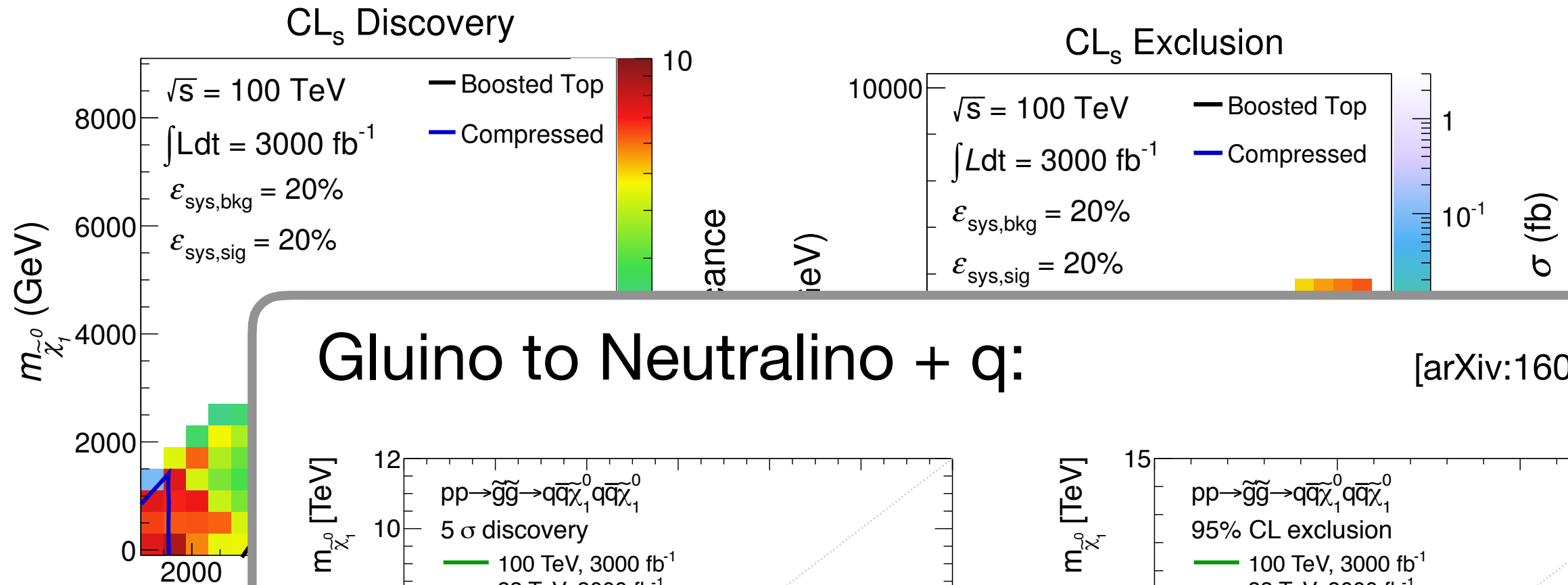
CL_s Exclusion



Naturalness

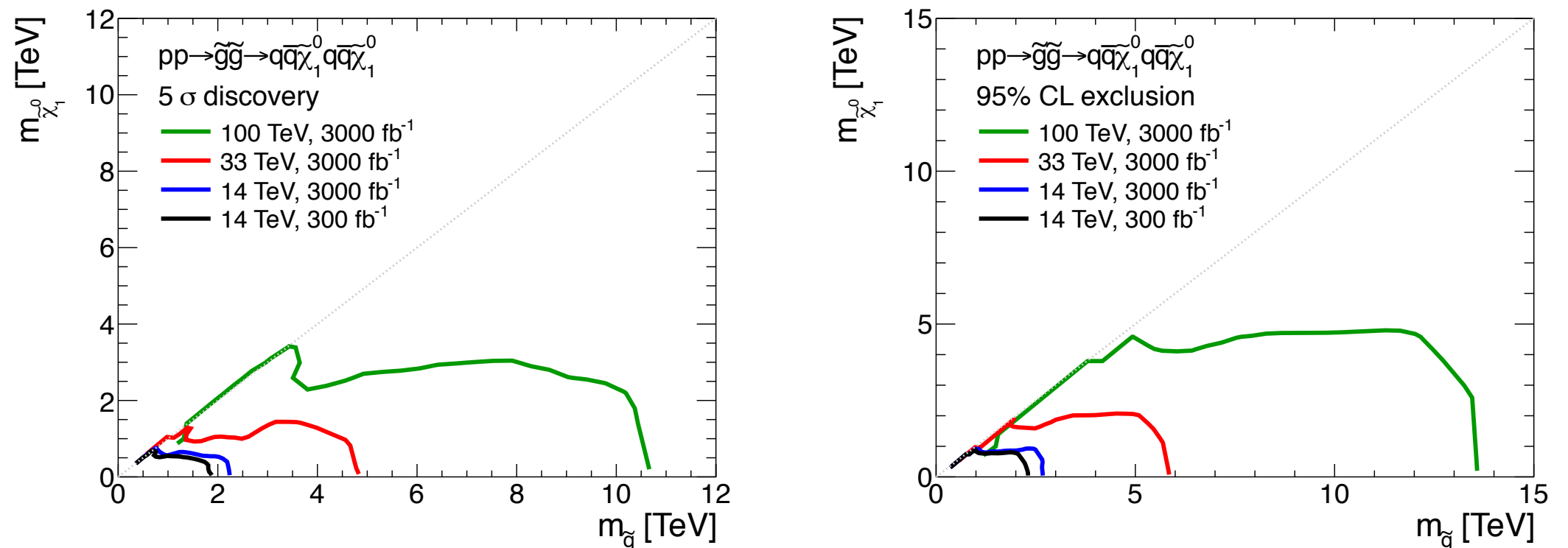
Stop to top + Neutralino:

[arXiv:1406.4512]



Gluino to Neutralino + q:

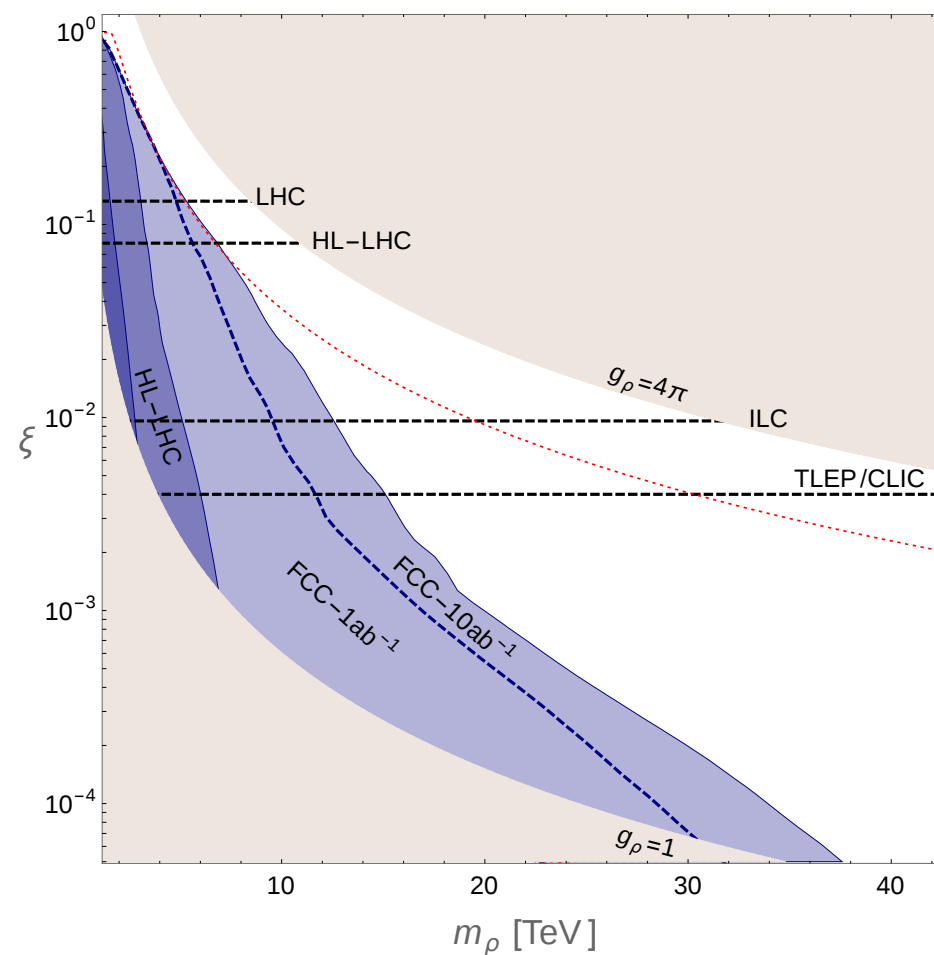
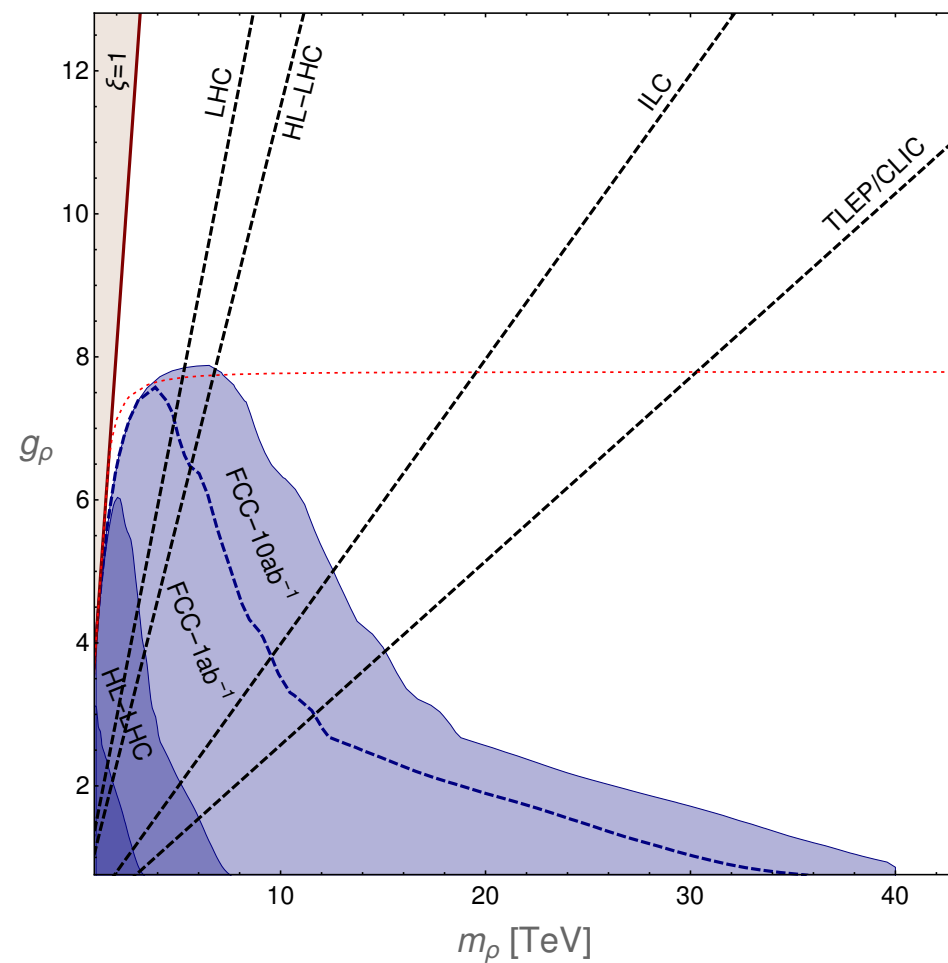
[arXiv:1606.00947]



Naturalness

Spin-1 CH resonances (or V'):

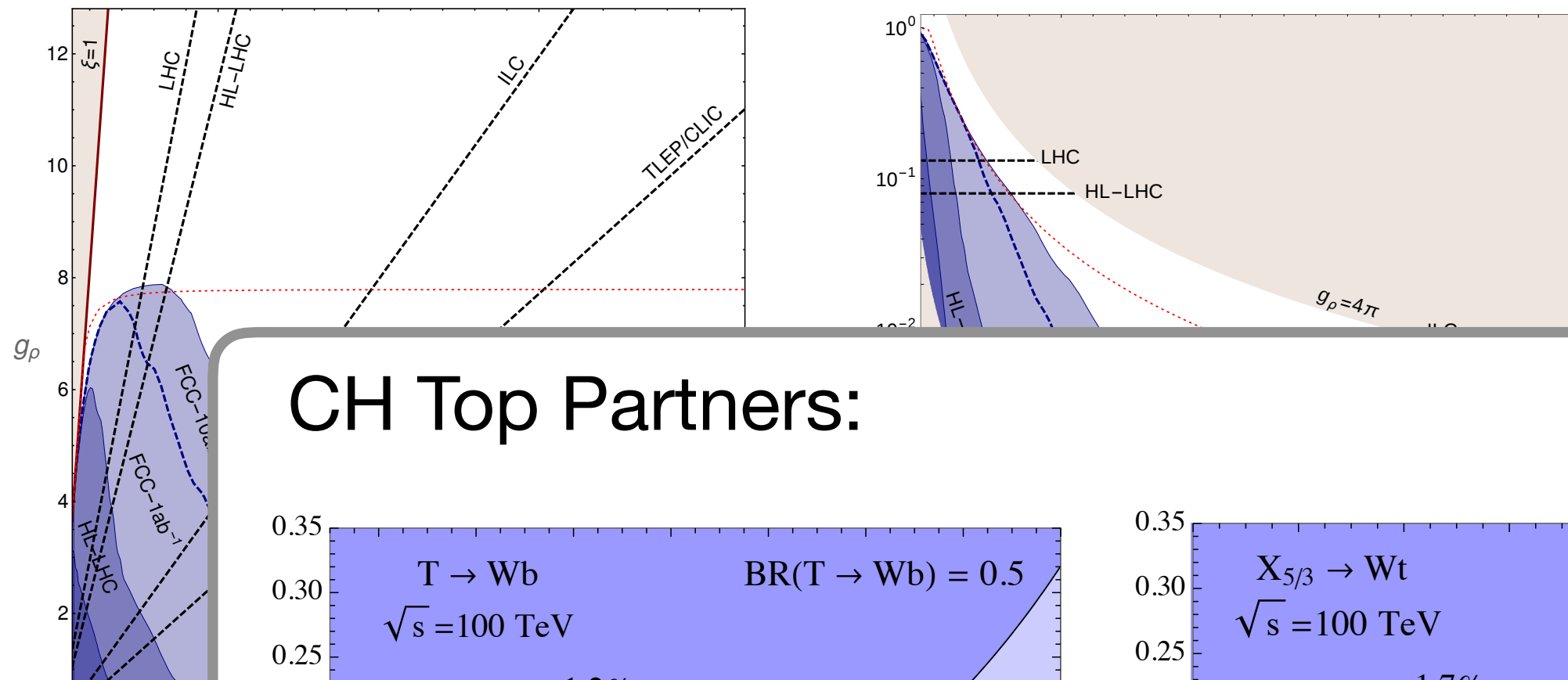
[arXiv:1502.01701]



Naturalness

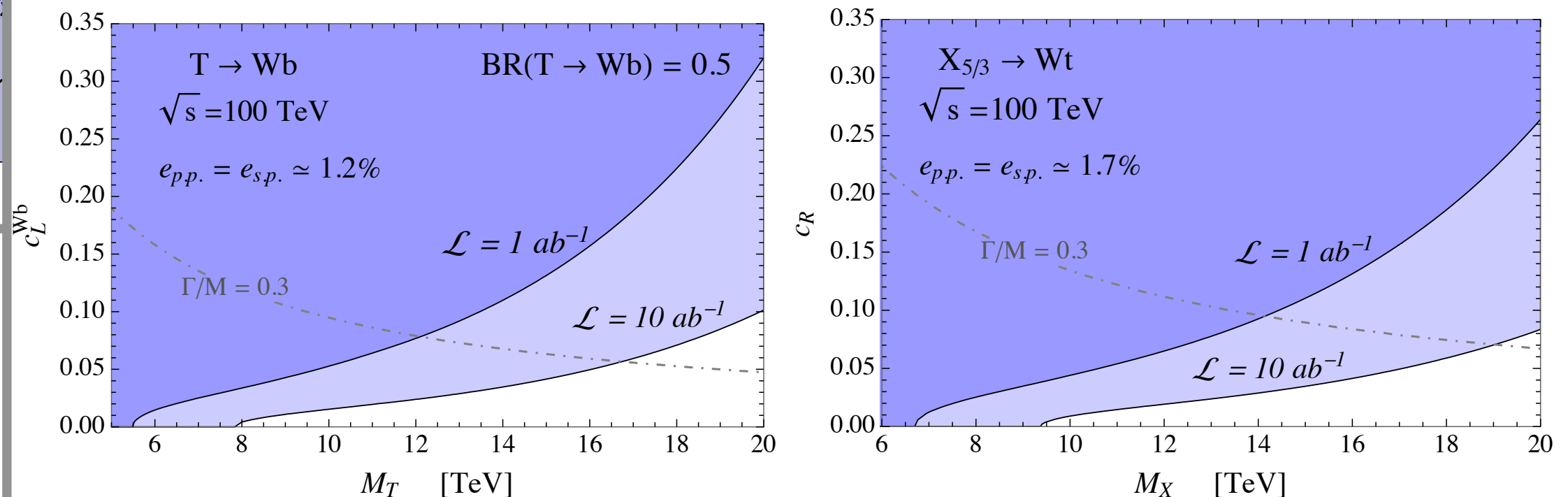
Spin-1 CH resonances (or V'):

[arXiv:1502.01701]



CH Top Partners:

[arXiv:1409.0100]

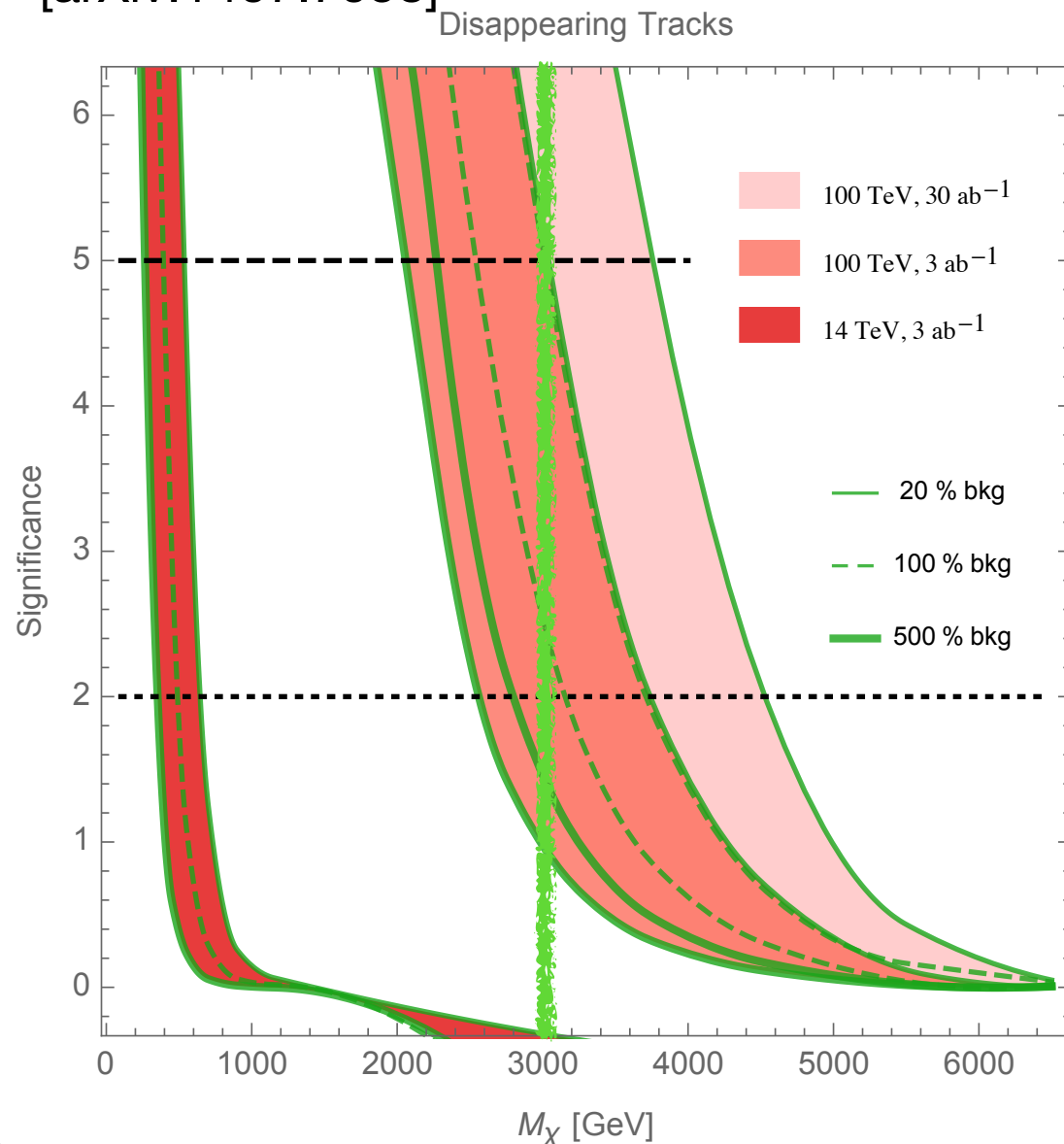


Dark Matter

100 TeV can probe WIMP DM inaccessible to current and future Direct Detection

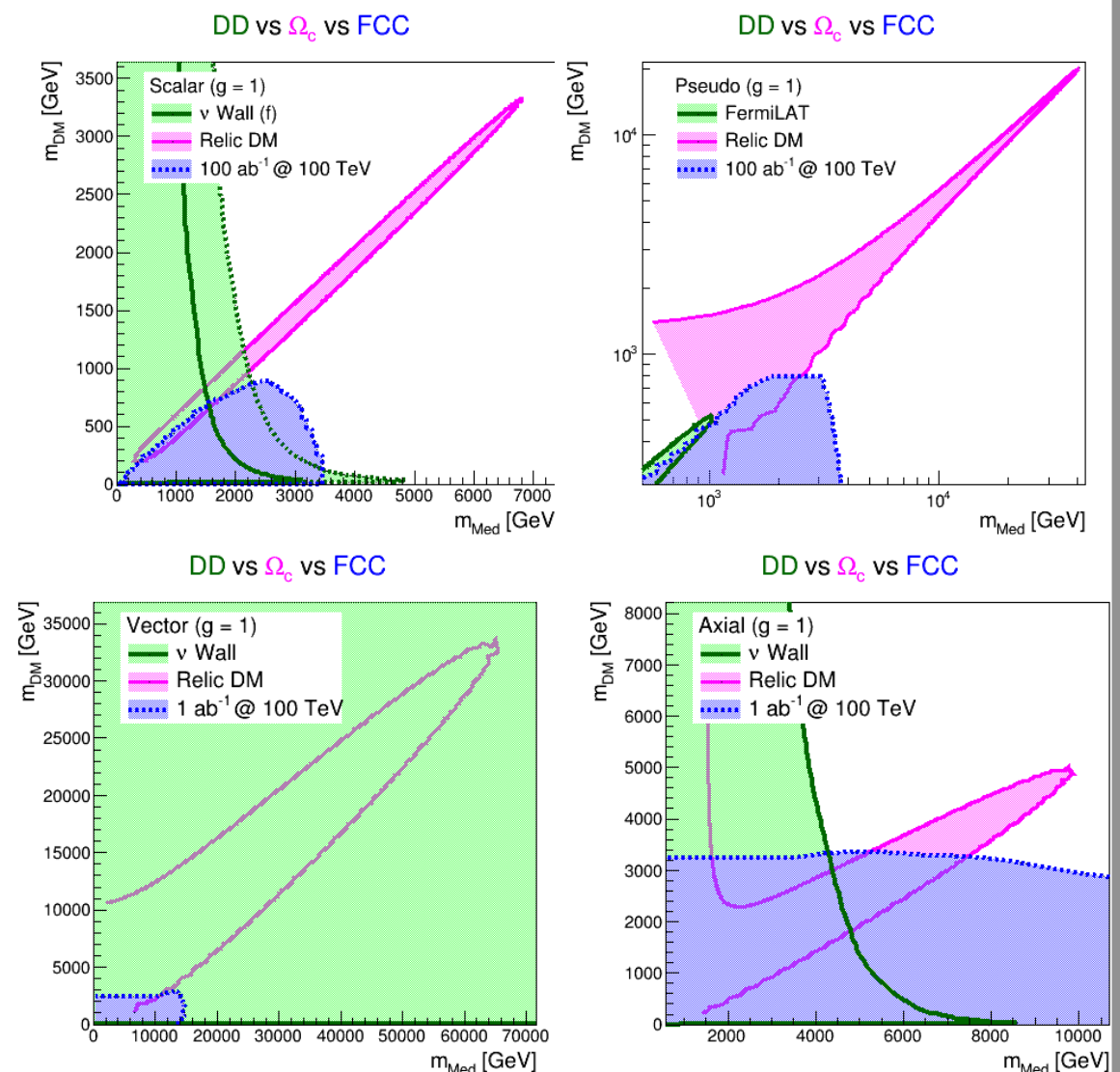
Wino DM. (disappearing tracks:)

[arXiv:1407.7058]



Crossing neutrino wall:

[arXiv:1606.00947]



Dark Matter

100 TeV can probe WIMP DM inaccessible to current and future Direct Detection

Wino DM. (disappearing tracks:)

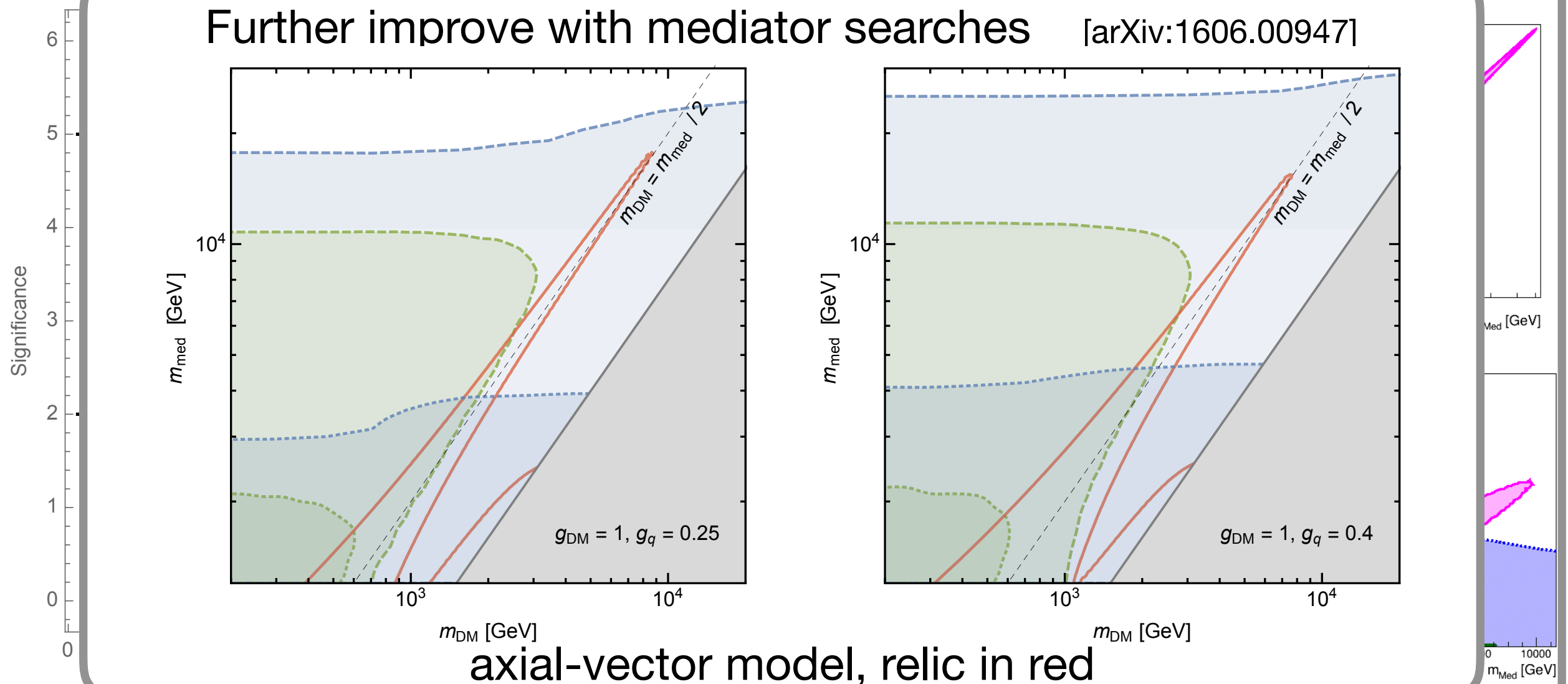
[arXiv:1407.7058]

Crossing neutrino wall:

[arXiv:1606.00947]

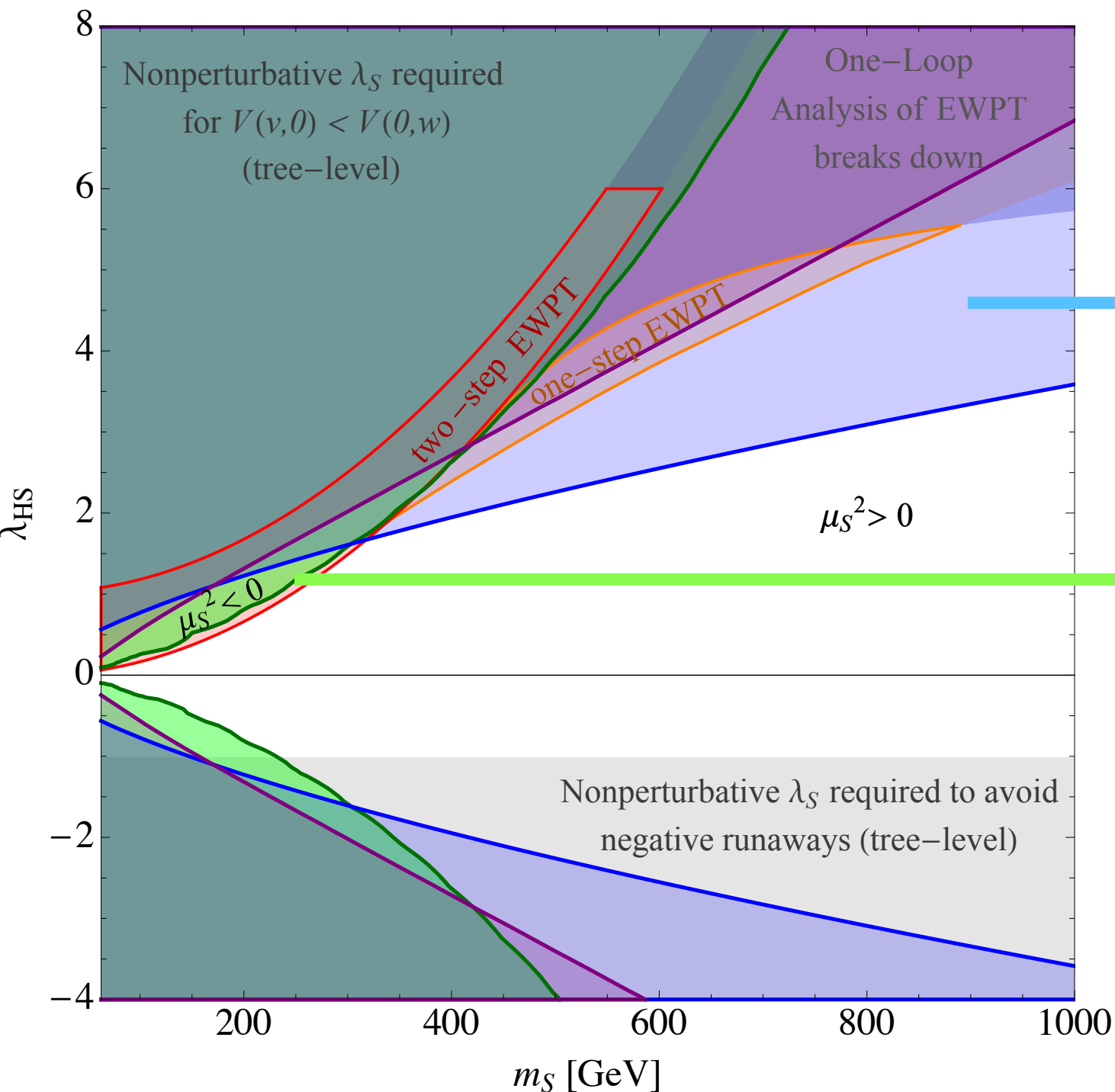
Further improve with mediator searches

[arXiv:1606.00947]



EW Baryogenesis

Direct reach on new states and few-percent 3H coup.
can probe **strong first order EW phase transition**



Z2-odd singlet (nightmare scenario?)
[arXiv:1409.0005]

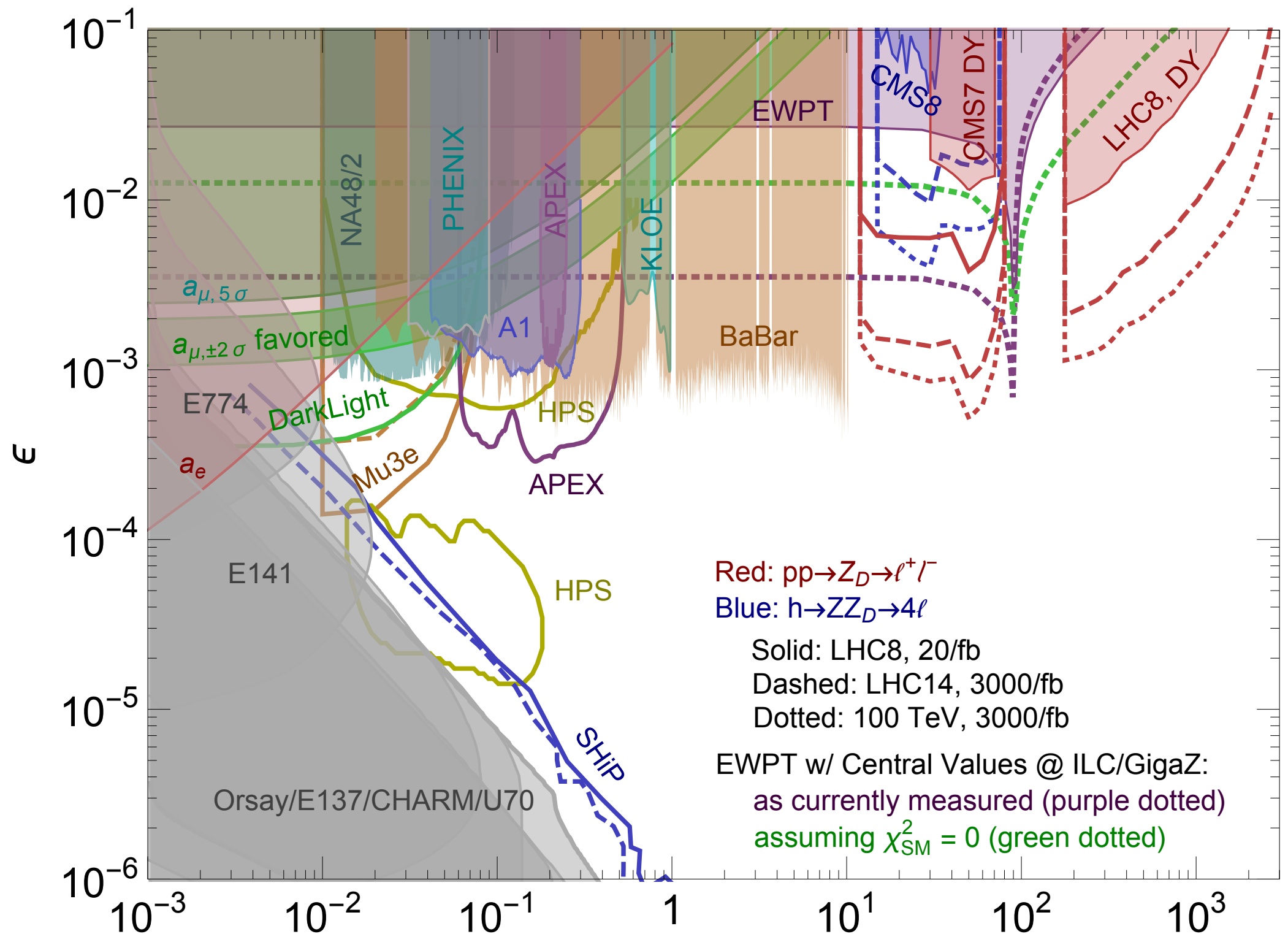
$$\delta g_{3H} / g_{3H} \sim 3, 4\%$$

Direct singlet search
(VBF $h^* \rightarrow SS$)

Intensity/Accuracy Frontier @ 100 TeV

Huge “low-mass” rate allows to produce Dark Sec. part. ...

Dark photon:
[arXiv:1412.0018]

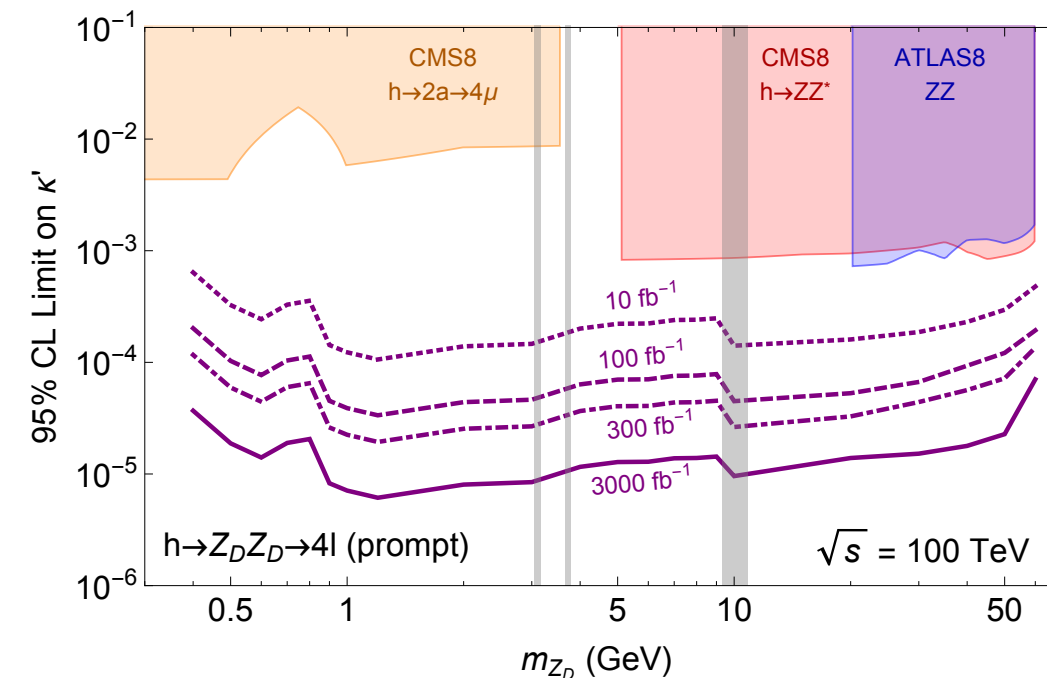
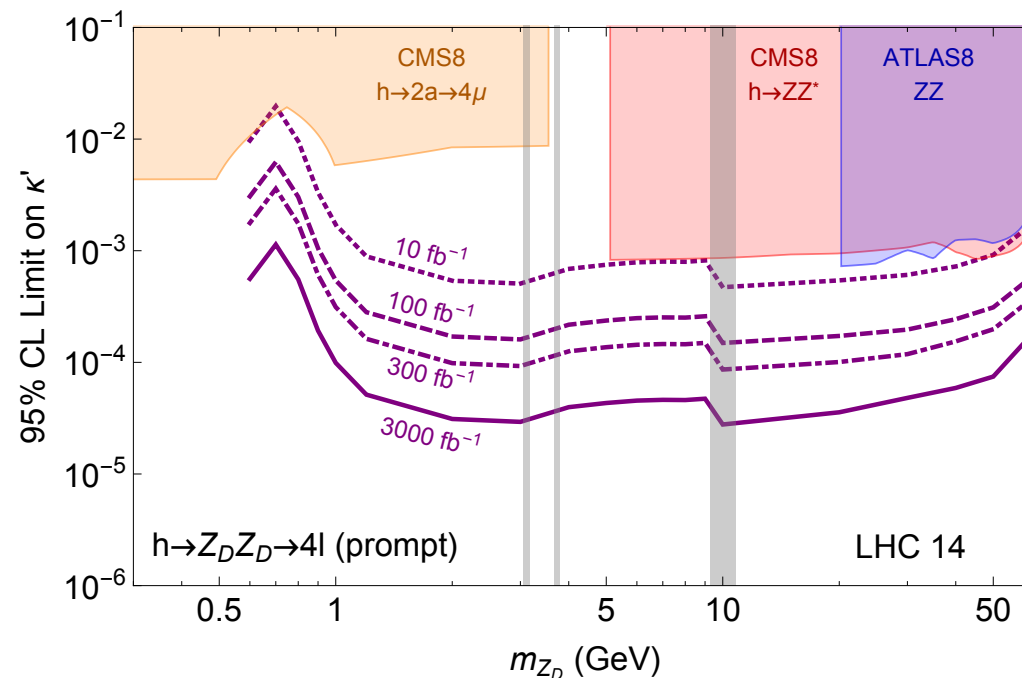
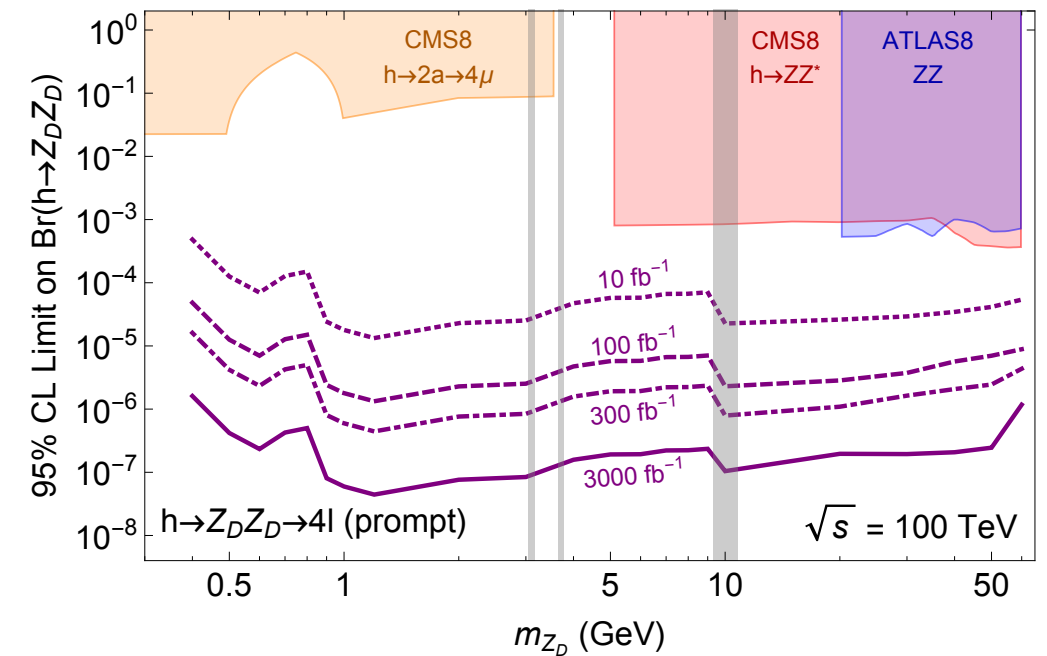
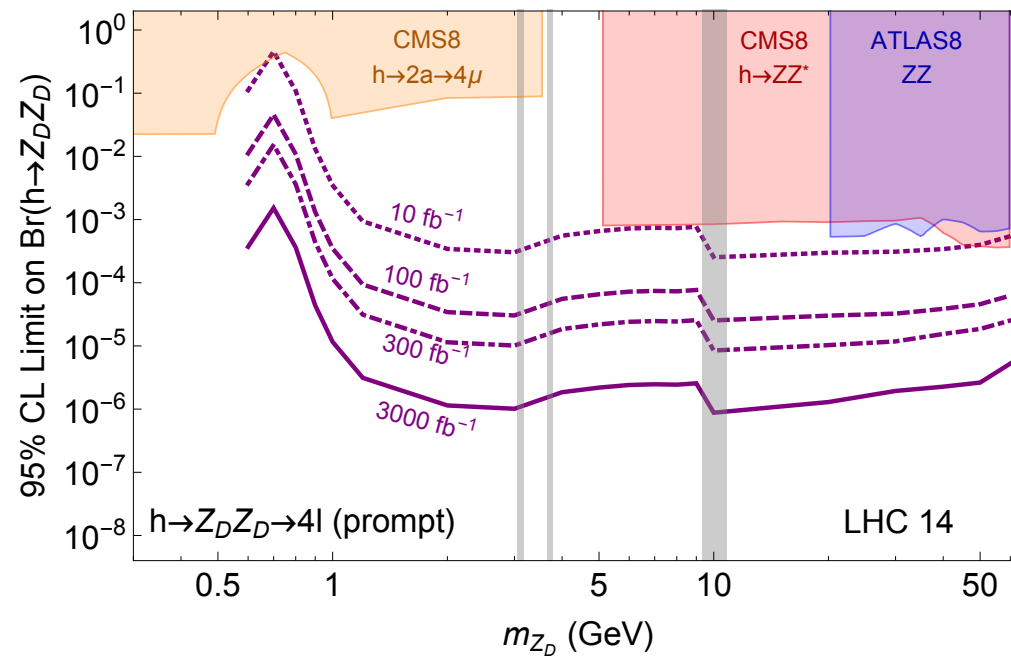


Intensity/Accuracy Frontier @ 100 TeV

Huge “low-mass” rate allows to produce Dark Sec. part. ...

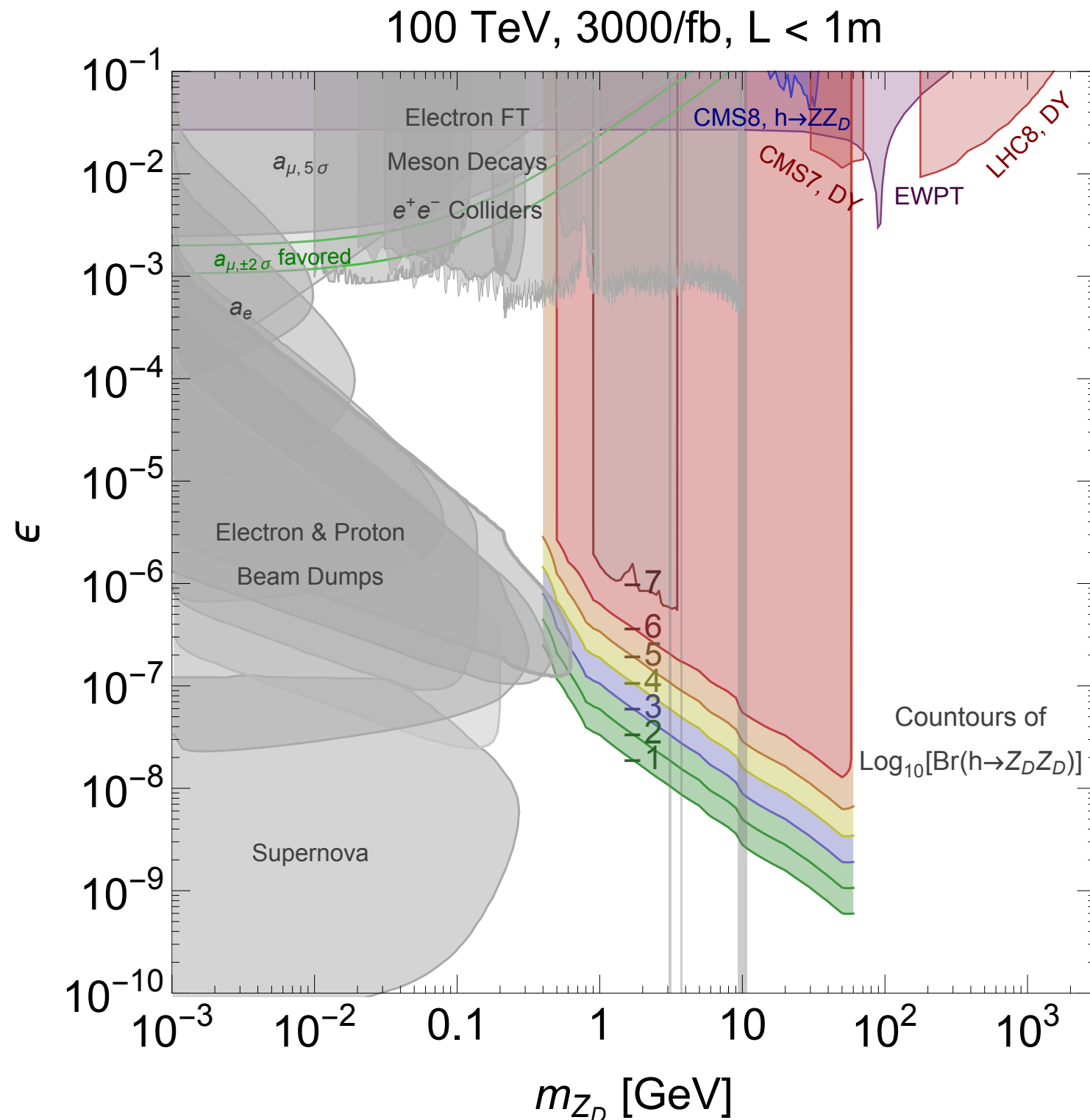
Dark photon:
[arXiv:1412.0018]

Dark photon
from H-portal:
[arXiv:1412.0018]



Intensity/Accuracy Frontier @ 100 TeV

Huge “low-mass” rate allows to produce Dark Sec. part. ...



Intensity/Accuracy Frontier @ 100 TeV

Huge “low-mass” rate allows to produce Dark Sec. part. ...
... and many SM particles as well (10^{10} Higgs, 10^{12} tops, ...)

Accuracy performances still to be fully assessed.
Known examples include:

- H-trilinear @ 3-4%, thanks to high rate (**Unique**)
- Top Yukawa @ 1% from ttH/ttZ [arXiv:1507.08169]

Intensity/Accuracy Frontier @ 100 TeV

Huge “low-mass” rate allows to produce Dark Sec. part. ...
... and many SM particles as well (10^{10} Higgs, 10^{12} tops, ...)

Accuracy performances still to be fully assessed.
Known examples include:

- H-trilinear @ 3-4%, thanks to high rate (**Unique**)
- Top Yukawa @ 1% from ttH/ttZ [arXiv:1507.08169]

Under investigation:

- Flavour Physics?
- Top FCNC? (e.g. $BR(Zq) \sim 10^{-7}$)
- Light Yukawa from Exotic Exclusive H decays?

Energy and Accuracy Frontier @ 100 TeV

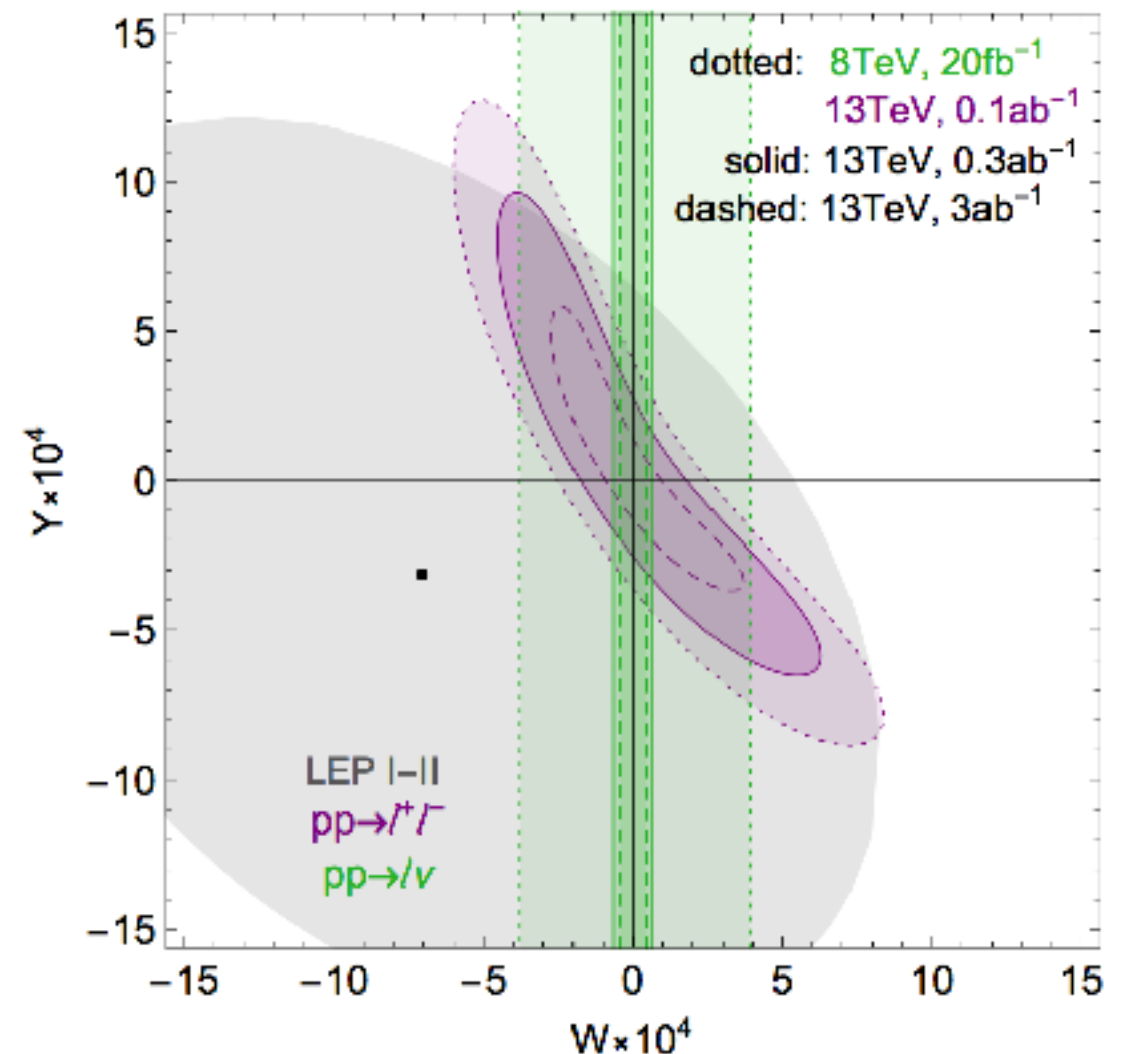
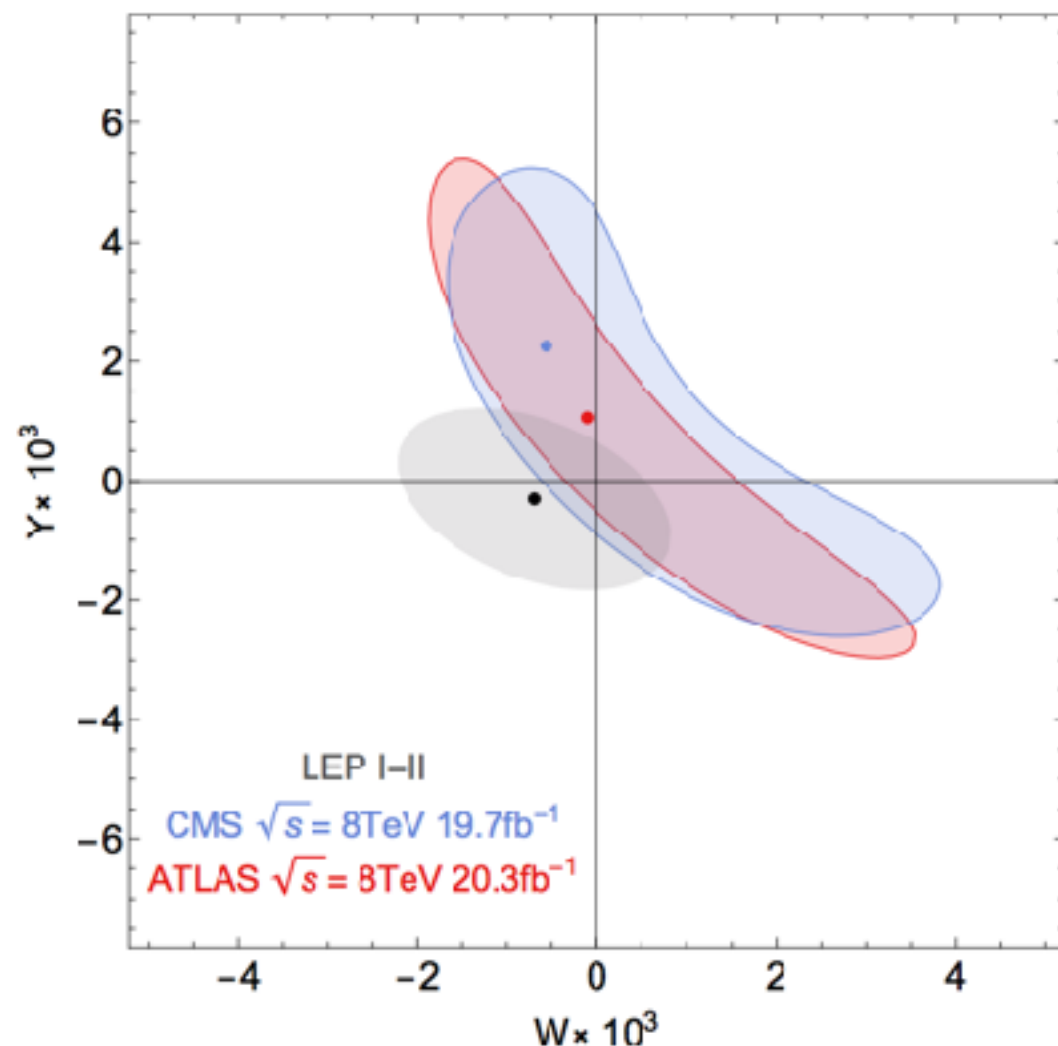
Enhanced indirect NP effects in high mass tails



No need of extreme accuracy for indirect NP probe

EWPT @ hadron colliders: (W and Y oblique par.s)

[arXiv:1609.08157]



Energy and Accuracy Frontier @ 100 TeV

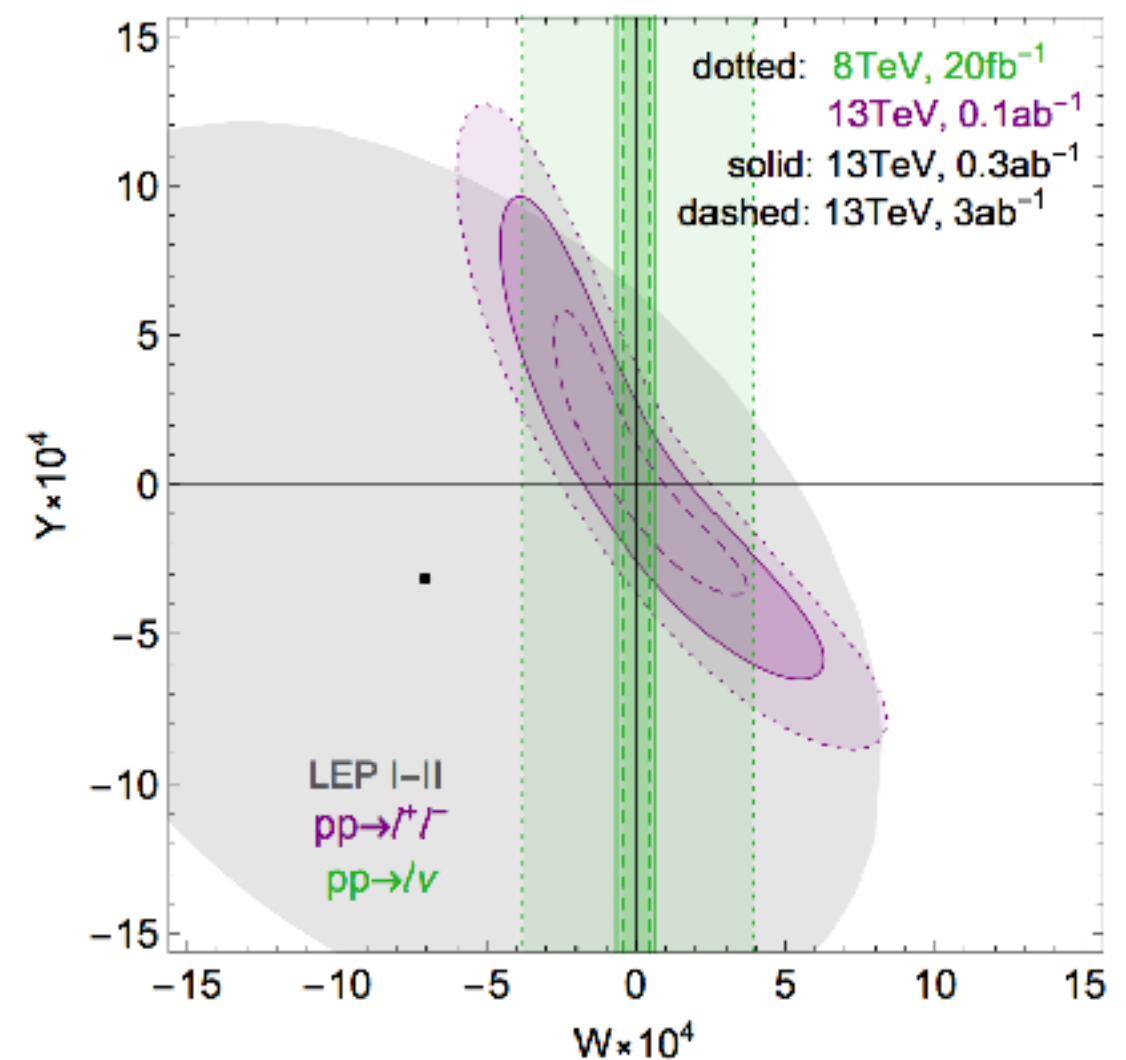
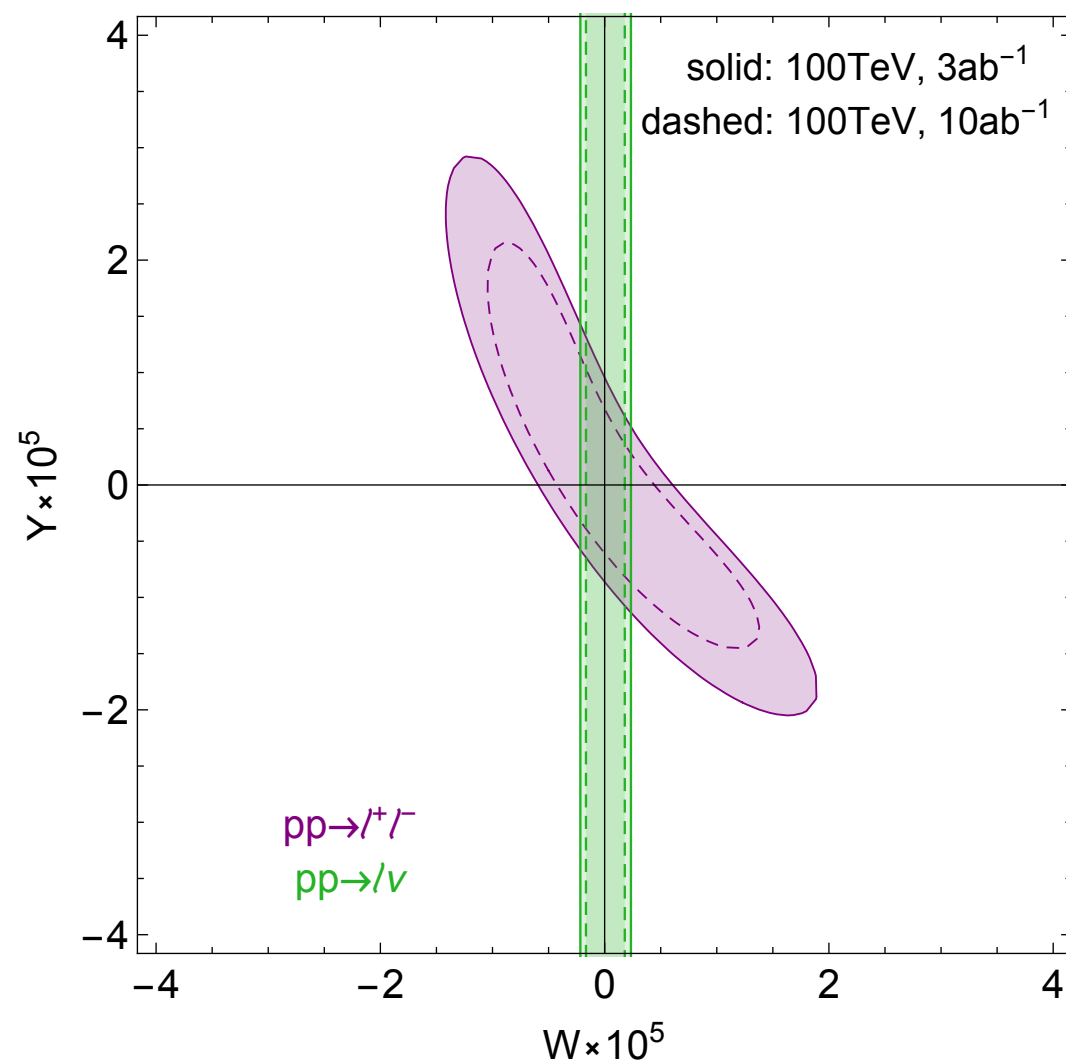
Enhanced indirect NP effects in high mass tails



No need of extreme accuracy for indirect NP probe

EWPT @ hadron colliders: (W and Y oblique par.s)

[arXiv:1609.08157]



Energy and Accuracy Frontier @ 100 TeV

Enhanced indirect NP effects in high mass tails



No need of extreme accuracy for indirect NP probe

EWPT @ hadron colliders: (W and Y oblique par.s) [arXiv:1609.08157]

Only CLIC@3TeV can be competitive:

		LEP	ATLAS 8	CMS 8	LHC 13		100 TeV	ILC	TLEP	ILC 500 GeV
luminosity		$2 \times 10^7 Z$	19.7 fb^{-1}	20.3 fb^{-1}	0.3 ab^{-1}	3 ab^{-1}	10 ab^{-1}	$10^9 Z$	$10^{12} Z$	3 ab^{-1}
NC	$W \times 10^4$	$[-19, 3]$	$[-3, 15]$	$[-5, 22]$	± 1.5	± 0.8	± 0.04	± 3	± 0.7	± 0.3
	$Y \times 10^4$	$[-17, 4]$	$[-4, 24]$	$[-7, 41]$	± 2.3	± 1.2	± 0.06	± 4	± 1	± 0.2
CC	$W \times 10^4$	—	± 3.9		± 0.7	± 0.45	± 0.02	—	—	—

Energy and Accuracy Frontier @ 100 TeV

Enhanced indirect NP effects in high mass tails



No need of extreme accuracy for indirect NP probe

EWPT @ hadron colliders: (W and Y oblique par.s) [arXiv:1609.08157]

Other example is top couplings from HE ttW/ttZ [arXiv:1511.03674]

Diboson final states or boosted Higgs under exploration

Summary

Unique Features of Future pp colliders:

1. **High Mass Reach** in many channels offering a broad coverage of multi-TeV NP. **NP characterisation!**

Summary

Unique Features of Future pp colliders:

1. **High Mass Reach** in many channels offering a broad coverage of multi-TeV NP. **NP characterisation!**
2. **Huge Variety** of NP searches, covering Energy, Intensity, Accuracy (and Energy) frontiers. **NP exploration**

Summary

Unique Features of Future pp colliders:

1. **High Mass Reach** in many channels offering a broad coverage of multi-TeV NP. **NP characterisation!**
2. **Huge Variety** of NP searches, covering Energy, Intensity, Accuracy (and Energy) frontiers. **NP exploration**
3. **Cost and technological challenges**

Summary

Unique Features of Future pp colliders:

1. **High Mass Reach** in many channels offering a broad coverage of multi-TeV NP. **NP characterisation!**
2. **Huge Variety** of NP searches, covering Energy, Intensity, Accuracy (and Energy) frontiers. **NP exploration**
3. **Cost and technological challenges**

But, **we must dream !**

