


Beyond 2-point Statistics

Daan Meerburg



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- If we want to learn anything new **probably have to seek beyond 2-point statistics**

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Provide **powerful tests of the inflationary paradigm** (Maldacena 2002, Creminelli and Zaldarriaga 2004 ++)

———— Beyond 2-point statistics: including tensors ————

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Put in (optimistic!) another way: Smoking gun for new
physics (Lee, Baumann & Pimentel 2016, Baumann et al in prep)

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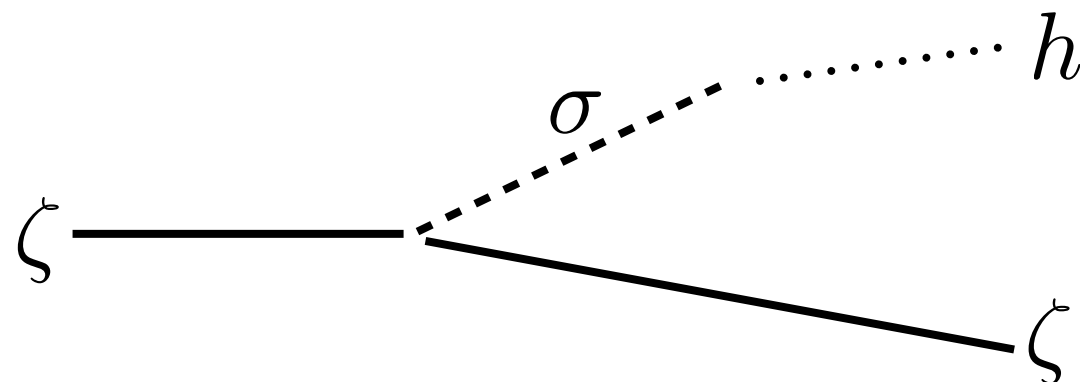
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Massive higher order spin fields (with 'free' coupling, Lee, Baumann and Pimentel 2016, Baumann et al in prep) e.g.



SEE ALSO ARKANI-HAMED & MALDACENA (2015)

Beyond 2-point statistics: simple ansatz

Example, correlating two scalar fluctuations with one tensor (model independent):

$$\langle \zeta \zeta h \rangle \propto \sqrt{r} f_{\text{NL}}^{\zeta \zeta h} \delta \left(\sum \vec{k}_i \right) \mathcal{I}(k_1, k_2, k_3) \epsilon_{ij}(k_3) \hat{k}_1^i \hat{k}_2^j$$

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Note that we assume $\sqrt{\langle hh \rangle} \propto \mathcal{O}(\sqrt{r})$

$$f_{\text{NL}}^{\zeta \zeta h} \equiv \langle \zeta \zeta h \rangle / (P_{\zeta}^{3/2} P_h^{1/2})$$

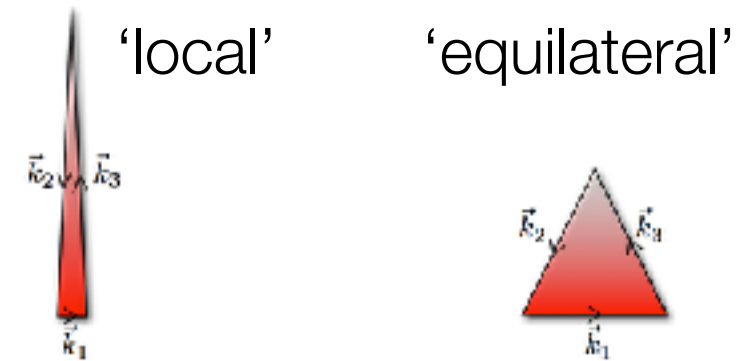
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Compute **observable bispectrum** $B_{\ell_1 \ell_2 \ell_3}$ for different shapes $\mathcal{I}(k_1, k_2, k_3)$

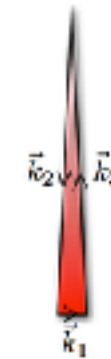


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Forecast detectability $F_{ii} = \sum_{\text{all modes}} \frac{(B^{BTT})^2}{\text{variance}}$



'local'



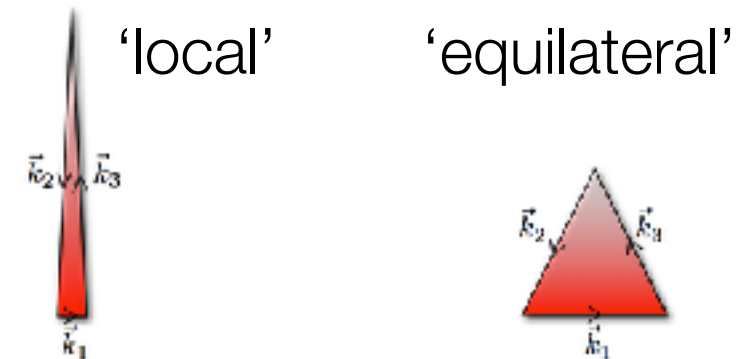
'equilateral'

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Future constraints on scalar NGs (TTT, TTE, TEE, EEE)

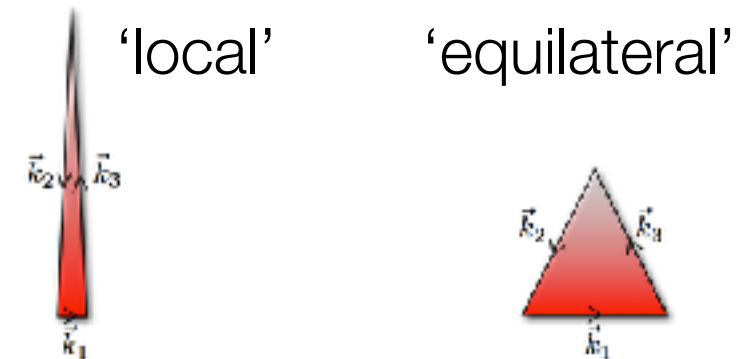
Type	<i>Planck</i> actual (forecast)	CMB-S4	CMB-S4 + low- ℓ <i>Planck</i>	Rel. improvement
Local	$\sigma(f_{\text{NL}}) = 5$ (4.5)	$\sigma(f_{\text{NL}}) = 2.6$	$\sigma(f_{\text{NL}}) = 1.8$	2.5
Equilateral	$\sigma(f_{\text{NL}}) = 43$ (45.2)	$\sigma(f_{\text{NL}}) = 21.2$	$\sigma(f_{\text{NL}}) = 21.2$	2.1
Orthogonal	$\sigma(f_{\text{NL}}) = 21$ (21.9)	$\sigma(f_{\text{NL}}) = 9.2$	$\sigma(f_{\text{NL}}) = 9.1$	2.4

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Future constraints on tensor NGs (BTT only)

Type	Planck	CMB-S4	rel. improvement
local	$\sigma(\sqrt{r} f_{\text{NL}}) = 15.2$	$\sigma(\sqrt{r} f_{\text{NL}}) = 0.3$	50.7
equilateral	$\sigma(\sqrt{r} f_{\text{NL}}) = 200.5$	$\sigma(\sqrt{r} f_{\text{NL}}) = 7.4$	27.1
local ($r = 0.01$)	$\sigma(\sqrt{r} f_{\text{NL}}) = 15.2$	$\sigma(\sqrt{r} f_{\text{NL}}) = 0.7$	25.3
equilateral ($r = 0.01$)	$\sigma(\sqrt{r} f_{\text{NL}}) = 200.8$	$\sigma(\sqrt{r} f_{\text{NL}}) = 14.7$	13.7

Beyond 2-point statistics: from space

Full-Sky	$\sum_n \ell_n = \text{even}$	$\sum_n \ell_n = \text{odd}$
Flat-Sky	Left-Handed = Right-Handed	Left-Handed = (-) Right-Handed
Non-Vanishing In Parity-Conserving Universe	$\langle TTT \rangle, \langle TEE \rangle, \langle TTE \rangle,$ $\langle EEE \rangle, \langle BBE \rangle, \langle BBT \rangle$	$\langle BTT \rangle, \langle BEE \rangle,$ $\langle BET \rangle, \langle BBB \rangle$

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Delta's, cosines and sines become **Wigner j's**

Include all possible measures and **all possible sources**:

$$\langle \zeta \zeta \zeta \rangle \rightarrow \langle TTT \rangle, \langle TTE \rangle, \langle TEE \rangle, \langle EEE \rangle$$

$$\langle h \zeta \zeta \rangle \rightarrow \langle TTT \rangle, \langle TTE \rangle, \langle TEE \rangle, \langle EEE \rangle$$

$$\langle BTT \rangle, \langle BTE \rangle, \langle BEE \rangle$$

$$\langle hh \zeta \rangle \rightarrow \langle TTT \rangle, \langle TTE \rangle, \langle TEE \rangle, \langle EEE \rangle, \langle BBT \rangle$$

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Beyond 2-point statistics: challenges

Fast and optimal estimator: Komatsu, Spergel and Wandelt (KSW) estimator (Assuming isotropic noise and diagonal covariance)

$$\hat{f}_{\text{NL}}^{X_1 X_2 X_3} = \frac{1}{N} \sum_{pqrijk} \sum_{lm} \frac{B_{m_1 m_2 m_3}^{\ell_1 \ell_2 \ell_3, pqr} X_1 X_2 X_3}{C_{\ell_1}^{pi} C_{\ell_2}^{qj} C_{\ell_3}^{rk}} a_{\ell_1 m_1}^i a_{\ell_2 m_2}^j a_{\ell_3 m_3}^k$$

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Speed depends on loops; naively ℓ^5

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$$\hat{f}_{\text{NL}}^{\zeta\zeta\zeta} \propto \int y^2 dy \int d\Omega A^\zeta(\hat{n}, y) B^\zeta(\hat{n}, y) C^\zeta(\hat{n}, y)$$

With A , B and C filtered maps. Computations per map $\mathcal{O}(\ell^3)$

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With A , B and C filtered maps. Computations per map $\mathcal{O}(\ell^3)$

Determines overall scaling algorithm, i.e. $\ell^5 \rightarrow \ell^3$

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Maps exist. Are complex maps, and estimator becomes

$$\hat{f}_{\text{NL}}^{\zeta \zeta h} \propto \sum_{M m_a m_b} \begin{pmatrix} 1 & 1 & 2 \\ m_a & m_b & M \end{pmatrix} \int_{\mathbb{R}^+} r^2 dr \int_{\mathbb{S}^2} d\Omega A_{m_a}^{\zeta}(\hat{n}, r) B_{m_b}^{\zeta}(\hat{n}, r) C_M^h(\hat{n}, r)$$

Duivenvoorden et al in prep

— Beyond 2-point statistics: outlook & conclusions —

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- Upcoming polarisation experiments hope to detect B-modes
- Similar to scalars, make sense to extend the search beyond 2-point statistics
- Contrary to scalars, not as easy to produce large violations, but at same time huge improvements on constraints are expected in the near future
- Could provide smoking gun for new physics
- Some very interesting models, such as heavy particles as relics of string theory, naturally produce effects at 3-level
- Actively building tools to apply to explore data (SO, CMBS4, litebird)
- Promising results, resolved the scaling problem of the KSW estimator