Secondary CMB Anisotropies from Large Scale Structure: Science & Foregrounds

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observable	science	foreground	concern
$\langle BB \rangle$	r	$\langle BB \rangle$ (κ, g, τ)	$\langle BB_{g,\tau} \rangle$ $\langle EBEB_{g,\tau} \rangle$ (delensing)
$I_\nu(n)$, $\delta_g(n,z)$ (LSS delensing)	r, N_{eff}	galactic dust & interlopers	ρ(LSS, κ) < 1
$\langle \kappa\kappa \rangle$	σ_8, ΣM_ν	$\langle I_\nu I_\nu \kappa \rangle$, $\langle I_\nu I_\nu I_\nu I_\nu \rangle$	CIB, tSZ, kSZ non-Gaussianity & SED
$\langle \delta_g \kappa \rangle$, $\langle \gamma\kappa \rangle$	w, f_{nl}, ΣM_ν, b_g	$\langle g I_\nu I_\nu \rangle$	CIB, tSZ, kSZ non-Gaussianity & SED
$\langle \delta_g \Delta T_{kSZ} \rangle$	f(R), ΣM_ν galaxy formation	$\langle \Delta T_p \Delta T_p \rangle$	alignment effects & feedback uncertainties
$\langle \Delta T_{kSZ} \Delta T_{kSZ} \rangle$	eor, galaxy formation	$\langle I_\nu y \rangle$	CIB, tSZ, kSZ non-Gaussianity & SED
$\langle yy \rangle$	σ_8, galaxy formation	$\langle I_\nu y \rangle$	alignment effects & feedback uncertainties
$n_c(z,M)$	σ_8, ΣM_ν, galaxy formation	$\langle I_\nu y \rangle$	alignment effects & feedback uncertainties