

Morphological Classification of Veiled CEMP Stars

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We present a novel approach for categorical analysis of strongly carbon-enhanced metal-poor (CEMP) stars using medium-resolution ($R \sim 1,800$) spectra. Analysis of cool ($T_{\text{eff}} \sim 4000$ K) CEMP stars is largely inhibited by a strong depression of the underlying continuum (veiling) by extreme molecular bands, making normalization and fitting of metallic features such as Ca II H&K lines difficult.

Consequently, few metal-poor dwarf carbon stars are known, with which we can constrain the low-mass tail of the Pop II IMF. We present a new technique for spectral normalization of these stars, and explore a technique for assigning CEMP morphological types based on characteristic abundances seen in Group I, II, and III CEMP stars, using the Yoon-Beers diagram introduced in Yoon et al. (2016).

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Talk/Poster

Poster

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