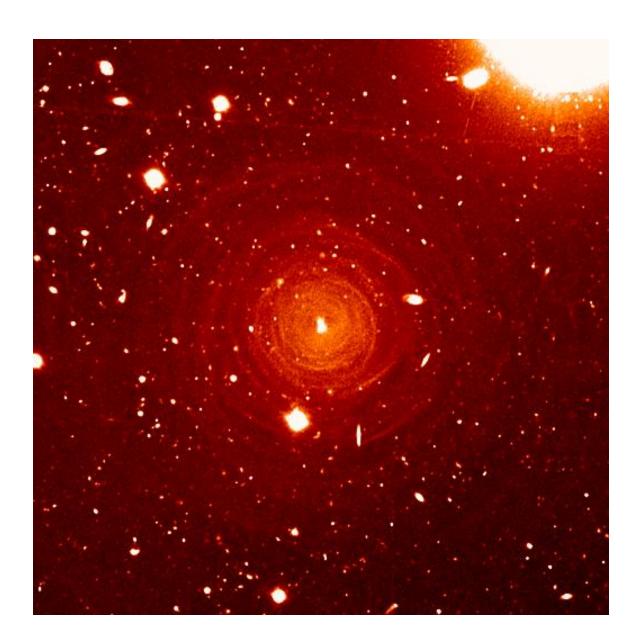
# An Abundance of Very Metal-Poor Stars in the Solar Neighborhood

**Collaborators:** 

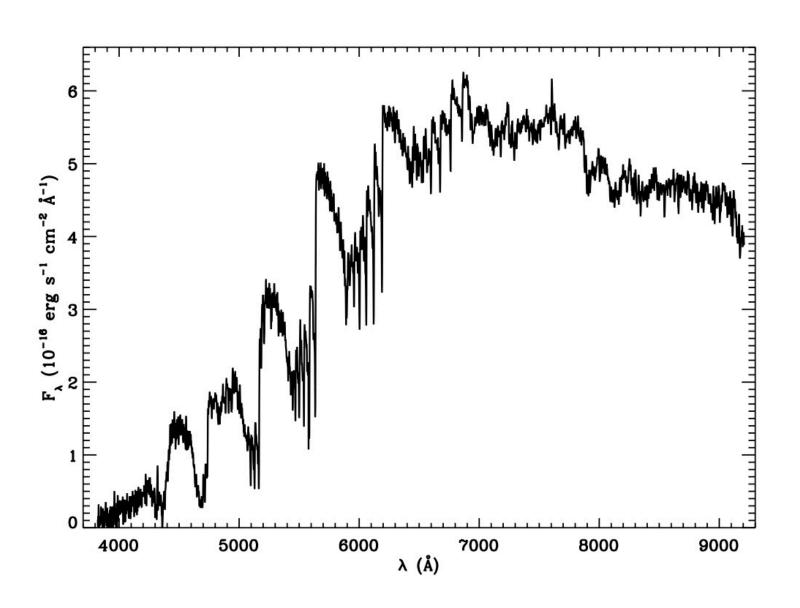
L. Whitehouse, P. Green, A. Arendt, H. Machado, J. Subasavage

J. Farihi
University College London

#### This Is A Carbon Star



### This is a Carbon Star Spectrum



#### G77-61: The First Dwarf Carbon (dC) Star

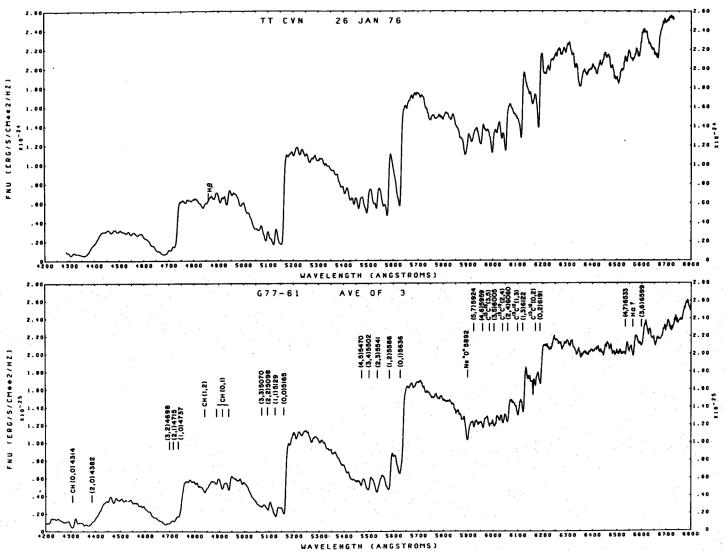


Fig. 1.—Lick 3 m IDS scans of a giant carbon star TT CVn (C3,5 CH—Victoria classification) and G77-61. The <sup>12</sup>C<sup>13</sup>C band heads of the (2,0) sequence are evident in G77-61, but appear only weakly in TT CVn.

#### dC Prototype Fact Sheet

• G77-61 discovered 1977 at Lowell (Dahn et al. 1977)

•  $M \approx 0.3 M_{\odot}$  halo star currently at 78.7pc

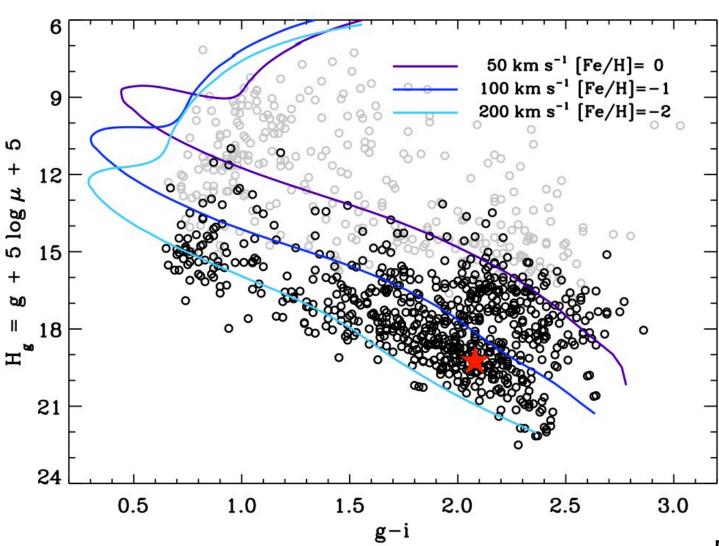
• SB1 spectroscopic binary in 245 d orbit (Dearborn et al. 1986)



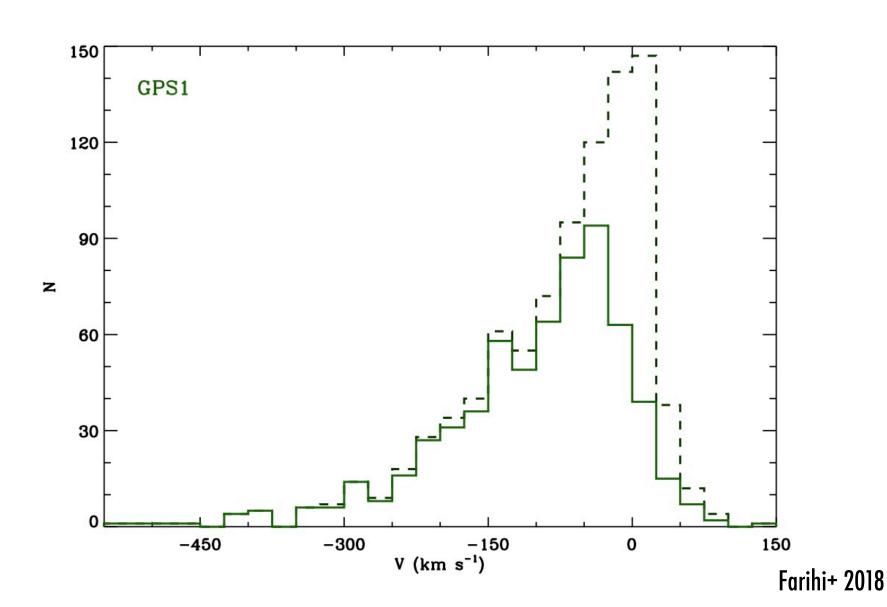
L. Whitehouse talk

[Fe/H] = -4.0 (Plez & Cohen 2005)

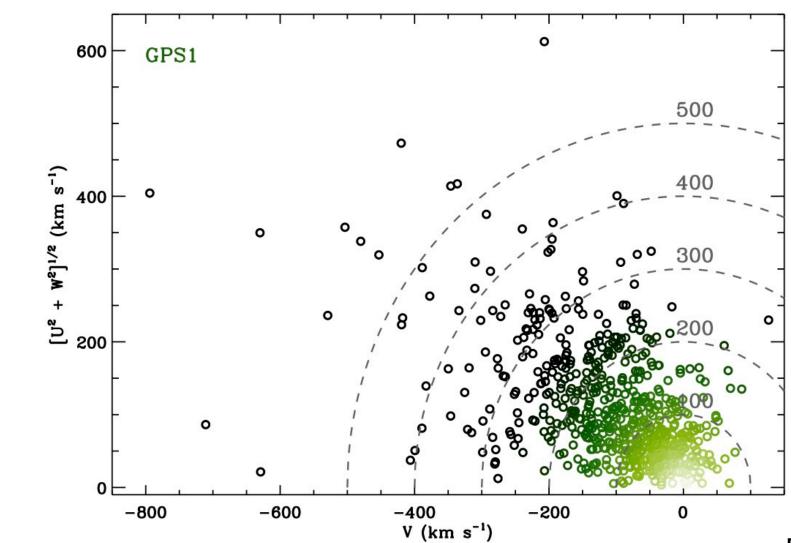
#### Gaia DR1 Reduced Proper Motions



#### dC Stars Lag Behind Disk Rotation

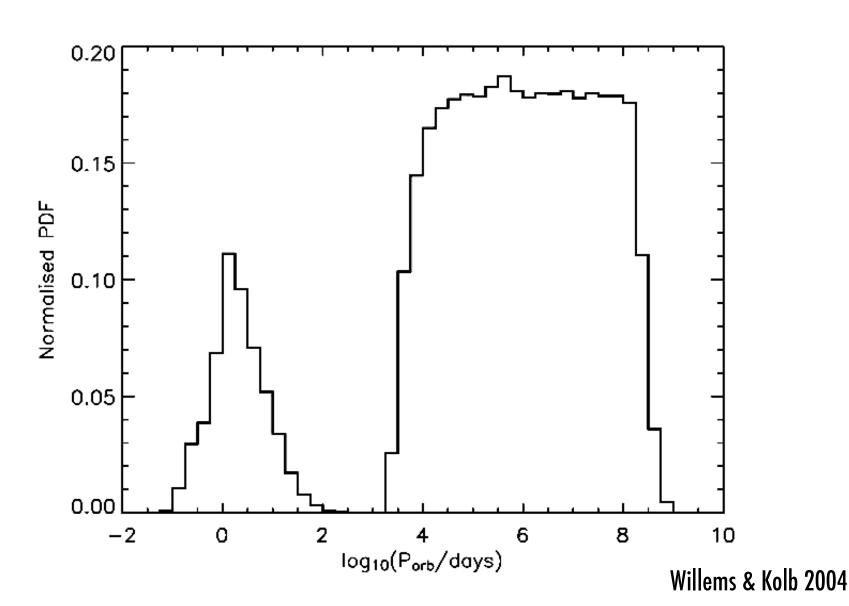


#### dC Stars Have a Significant Halo Component

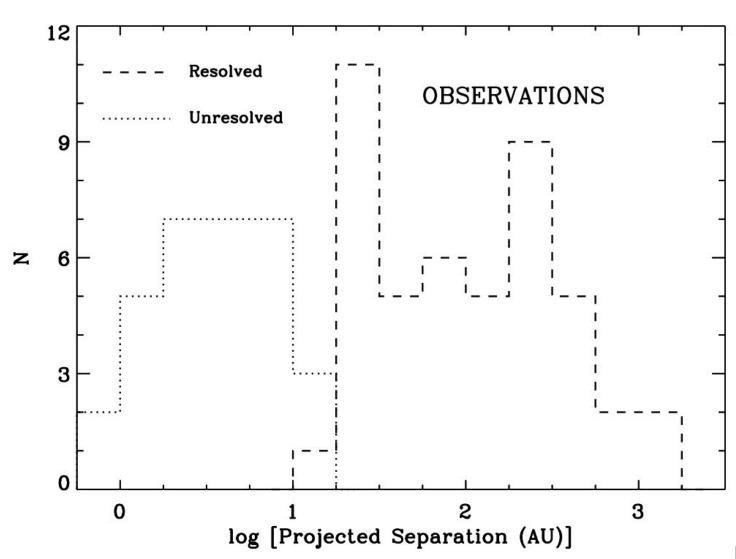


Farihi+ 2018

#### **Theoretical Models for Binary Orbits**



#### **Empirical Data for Binary Orbits**



### **Looking Ahead**

Gaia and better kinematical constraints

Fundamental properties, identification, occurrence

Atmospheric models and abundances!

Steller and binary evolution modeling

Spectroscopy at higher resolution

## Thank you very much ありがとうございます