

INSTITUTE OF
PARTICLE
PHYSICS

Neutrino Interactions and Future LBL experiments

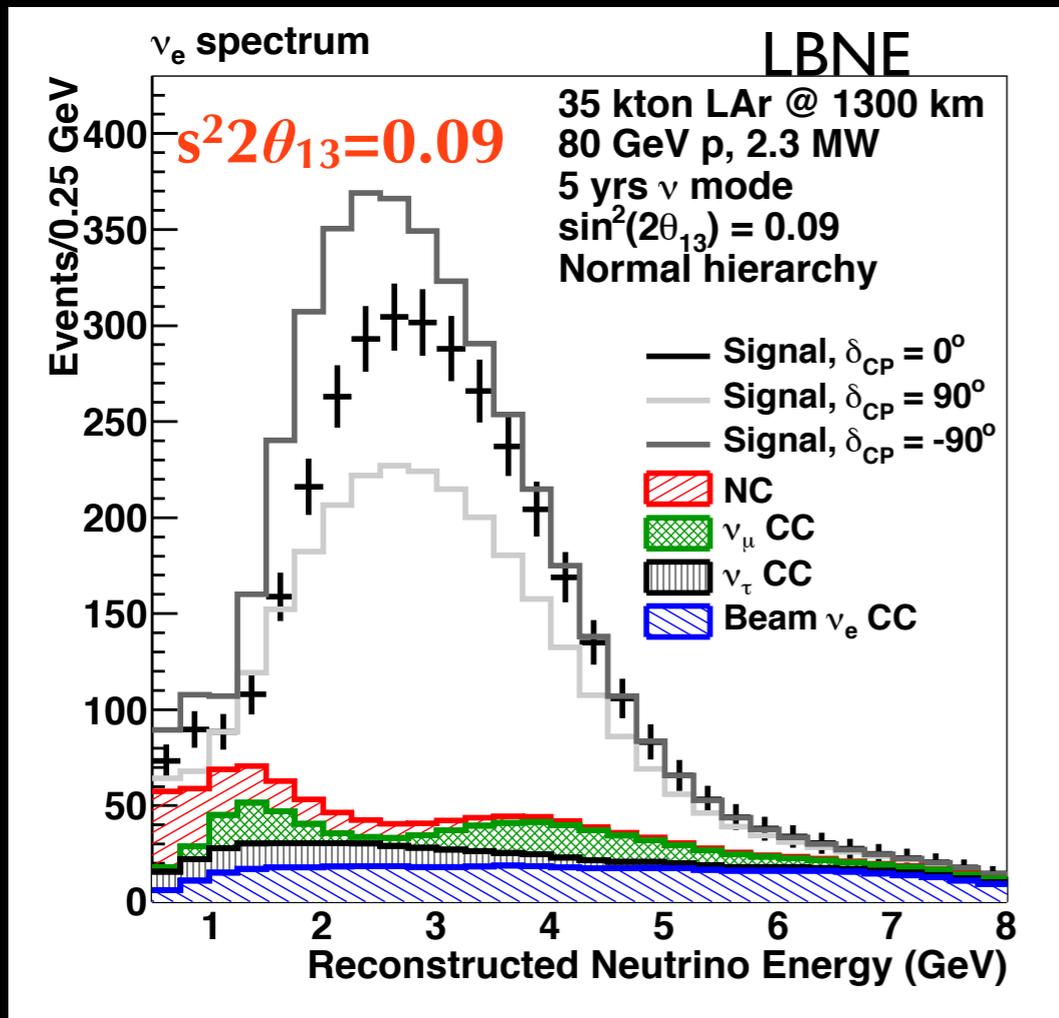
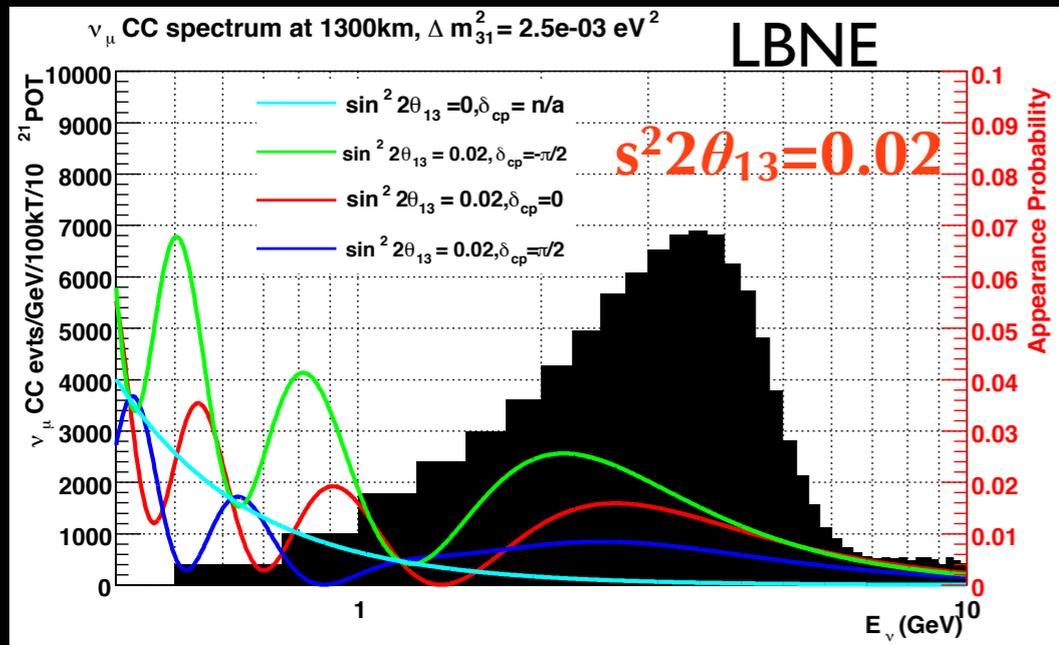
H. A. Tanaka (UBC/IPP)

International Workshop Next-Generation Nucleon Decay and Neutrino Detectors
Kavli-IPMU, Kashiwa, Japan, 11 Nov 2013

Overview

- Some generic comments on ν -int issues for LBL experiments
 - selective examples to illustrate general issues
- Some thoughts on next-generation neutrino near detectors
 - (a new workshop series?)
 - some other interesting efforts and activities
- Some sociological commentary
 - A pitch
- In the background: physics at the terascale
 - 1 MWatt x 1 MTon = 1 Tera Watt-ton
 - 1 G\$ x 1 kPerson = 1 Tera \$-Person
 - how do we ensure that we get the most out of this?

The Mixed Blessing of large θ_{13}



- CP asymmetry/variation of oscillation probability is smaller for larger θ_{13}
- small θ_{13} :
 - smaller rate/larger asymmetry
 - background systematics important
- large θ_{13} :
 - larger rate/smaller asymmetry
 - need/opportunity to exploit spectrum
 - “signal” ($\nu_{e/\mu}$ CC) systematics important

In my opinion:

Large θ_{13} is definitely a “blessing” overall

Oscillation Measurements

$$P(\nu_\mu \rightarrow \nu_e) \sim \frac{1}{2} \sin^2 2\theta_{13} \times \sin^2 \Delta$$

$$\Delta \equiv \frac{\Delta m_{31}^2 L}{4E}$$

$$P(\nu_\mu \rightarrow \nu_\mu) \sim 1 - \sin^2 2\theta_{23} \times \sin^2 \Delta$$

- With θ_{13} ~large and non-zero, we are in the era of full “3-flavor mixing”
 - all parameters (and their precision) matter
- Entering an era of precision where 3 flavor mixing effects matter
 - current $\sin^2 \theta_{23}$ uncertainty in $\nu_\mu \rightarrow \nu_e$ appearance
 - $\sin^2 \theta_{13}$ in ν_μ disappearance

N.B.
above equations are
only for illustration

Oscillation Measurements

$$P(\nu_\mu \rightarrow \nu_e) \sim \frac{1}{2} \sin^2 2\theta_{13} \times \sin^2 \Delta$$

$$P(\nu_\mu \rightarrow \nu_e) \sim \sin^2 2\theta_{13} \times \sin^2 \theta_{23} \frac{\sin^2[(1-x)\Delta]}{(1-x)^2} \quad (\equiv P_0)$$

$$-\alpha \sin 2\theta_{13} \times \sin \delta \sin 2\theta_{12} \sin 2\theta_{23} \sin \Delta \frac{\sin(x\Delta)}{x} \frac{\sin[(1-x)\Delta]}{(1-x)} \quad (\equiv P_1)$$

$$+\alpha \sin 2\theta_{13} \times \cos \delta \sin 2\theta_{12} \sin 2\theta_{23} \cos \Delta \frac{\sin(x\Delta)}{x} \frac{\sin[(1-x)\Delta]}{(1-x)} \quad (\equiv P_2)$$

$$+\alpha^2 \times \cos^2 \theta_{23} \sin^2 2\theta_{12} \frac{\sin^2(x\Delta)}{x^2} \quad (\equiv P_3)$$

M. Freund, Phys.Rev. D64 (2001) 053003

$$\alpha \equiv \frac{\Delta m_{21}^2}{\Delta m_{31}^2} \sim \frac{1}{30} \quad \Delta \equiv \frac{\Delta m_{31}^2 L}{4E} \quad x \equiv \frac{2\sqrt{2}G_F N_e E}{\Delta m_{31}^2}$$

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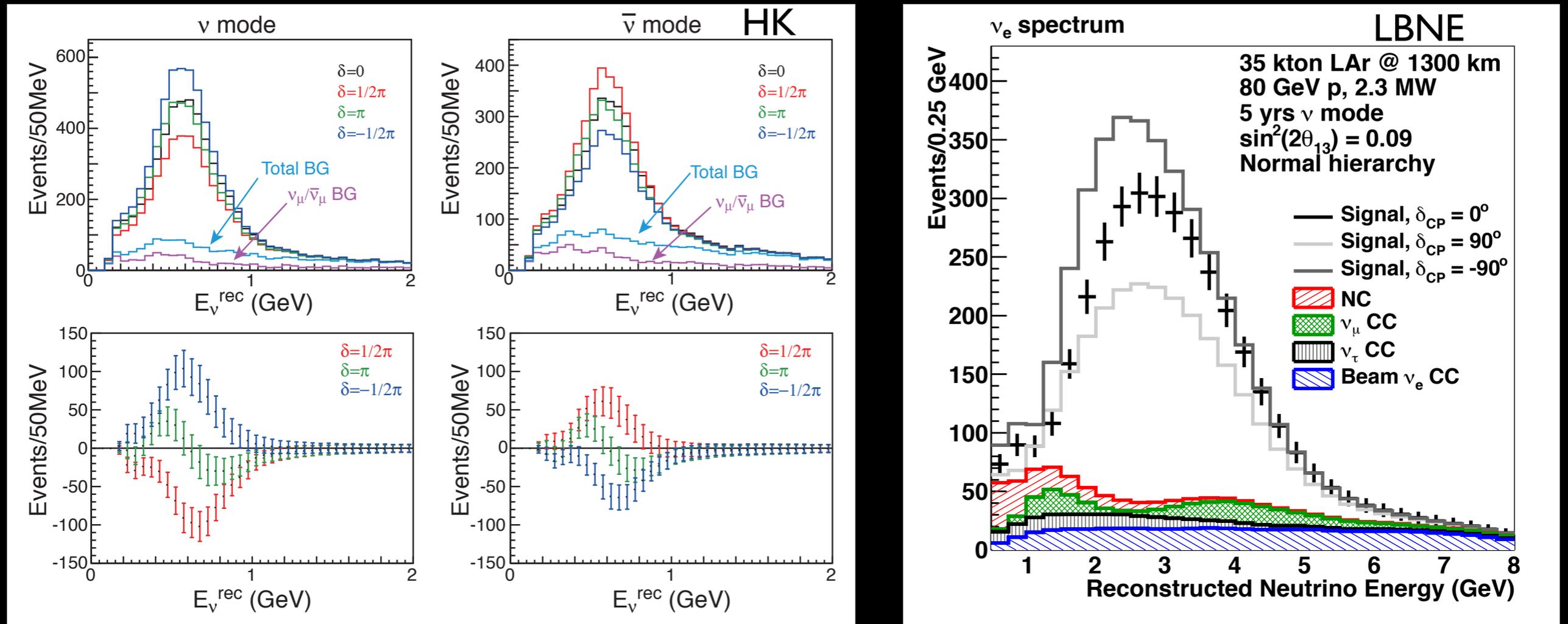
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$$P(\nu_\mu \rightarrow \nu_\mu) \sim 1 - (\cos^4 \theta_{13} \sin^2 2\theta_{23} + \sin^2 2\theta_{13} \sin^2 \theta_{23}) \sin^2 \Delta$$

- With $\theta_{13} \sim$ large and non-zero, we are in the era of full “3-flavor mixing”
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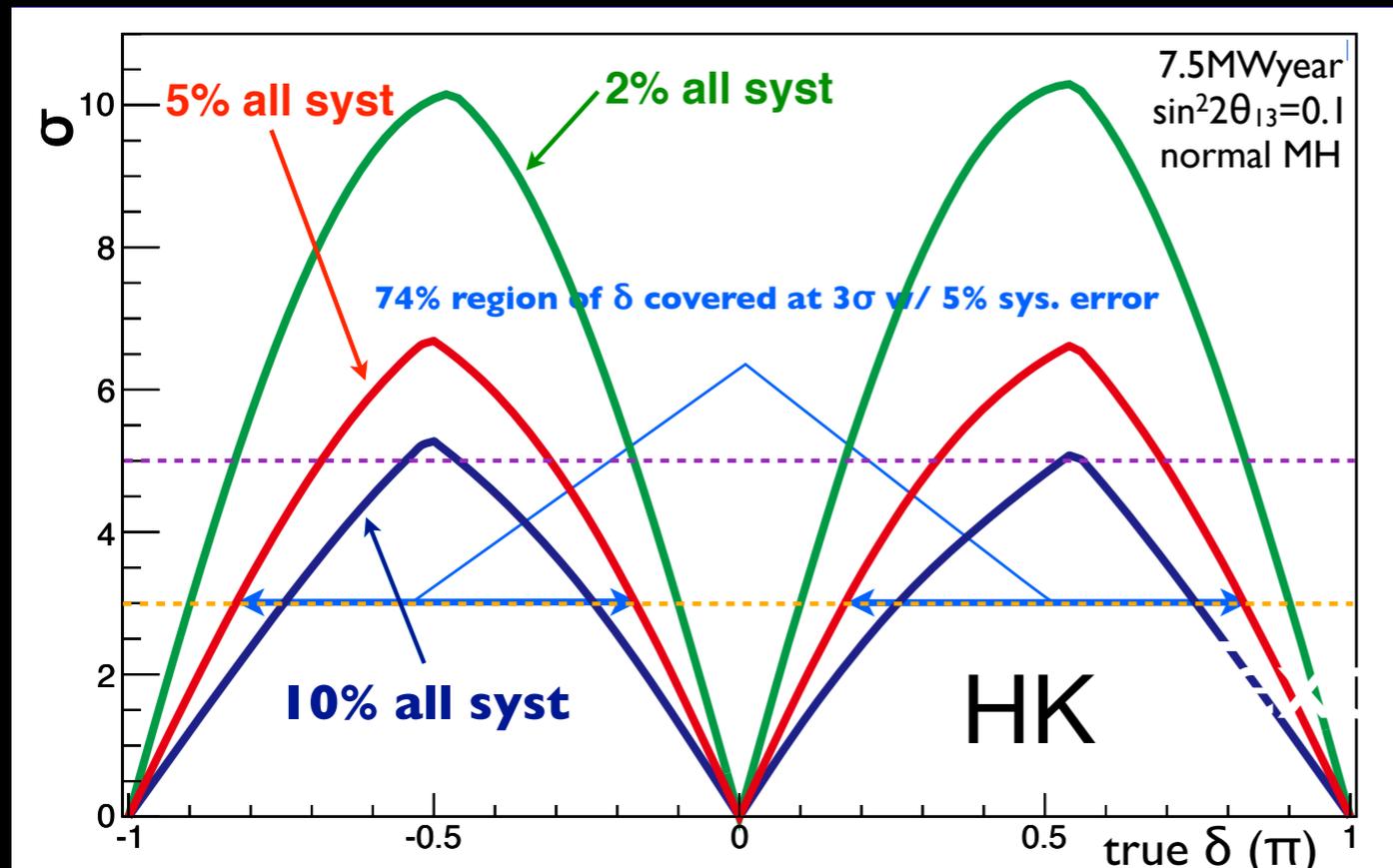
Spectrum Information



- At both HK and LBNE, spectrum information is important
- In principle, we need a thorough and precise understanding of
 - energy dependence of cross sections
 - energy reconstruction (relation between E_{ν} and outgoing particles).
 - “kinematic”: assume underlying mechanism and use $E_{\nu}(p_l, \theta_l)$
 - “calorimetric”: sum energy outgoing particles.

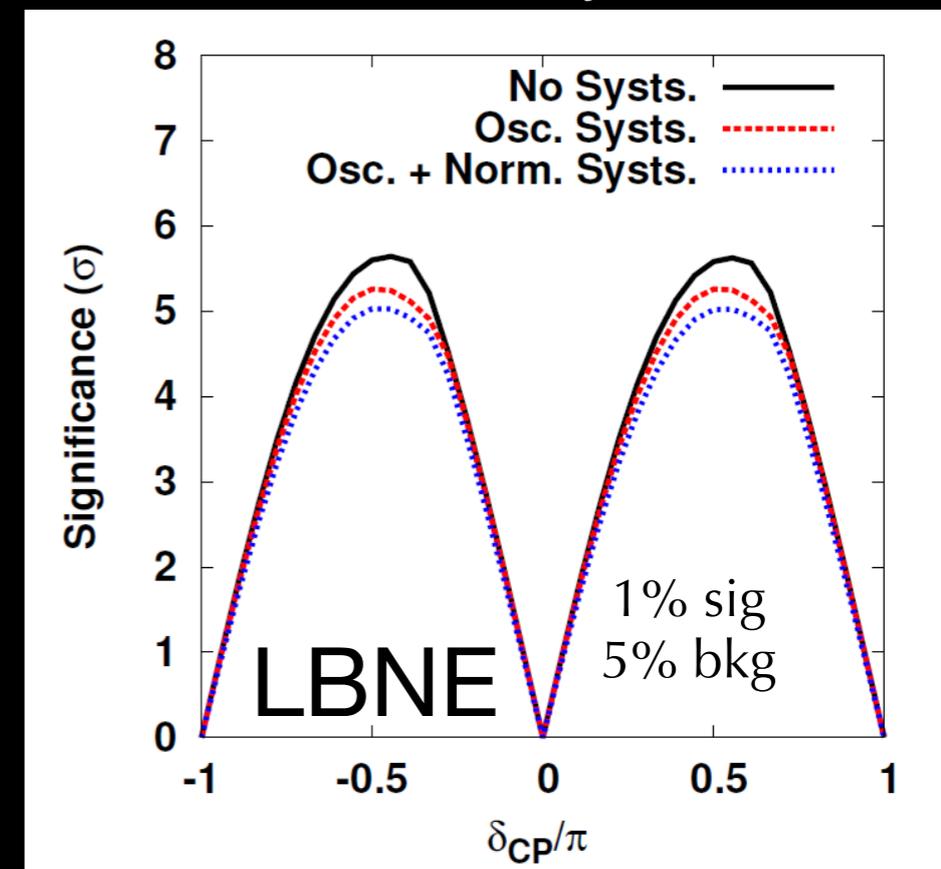
Impact of systematics

HK LOI



M. Bass et al.

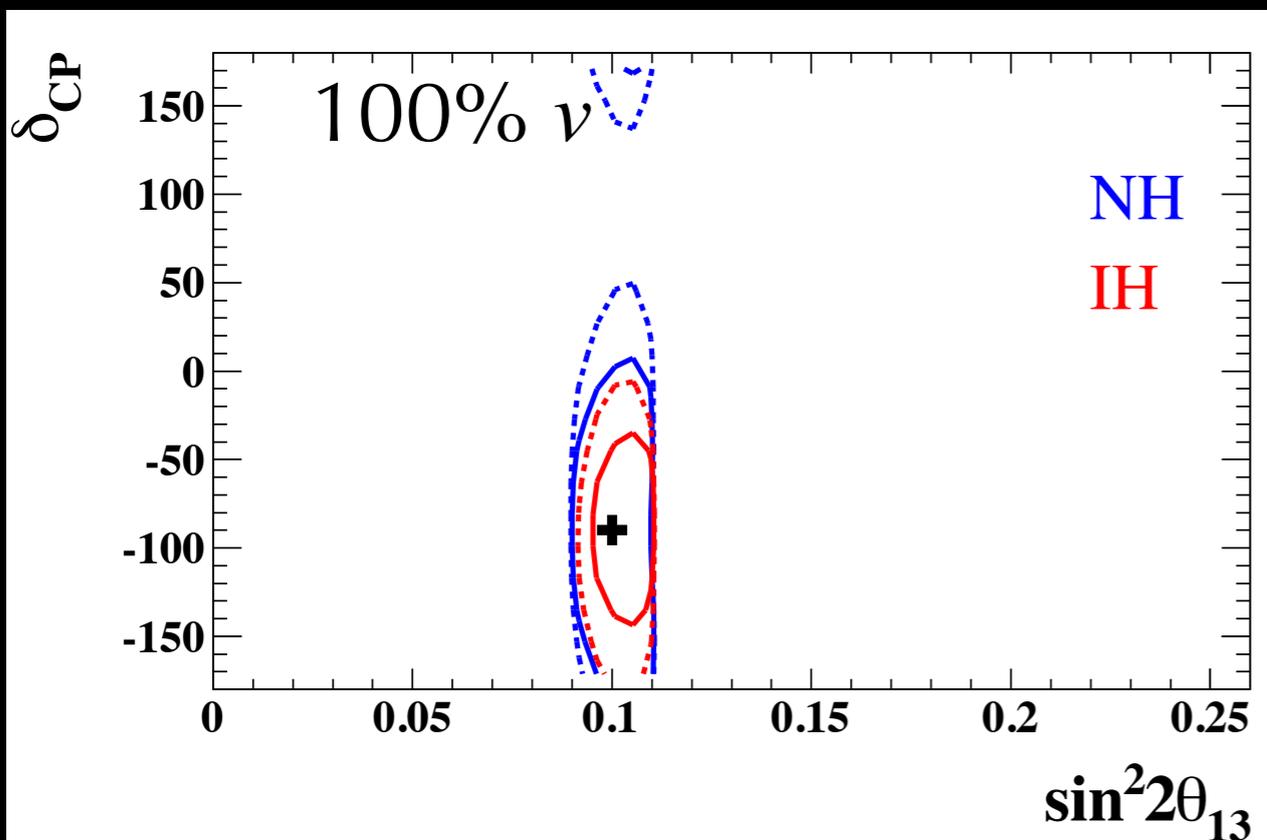
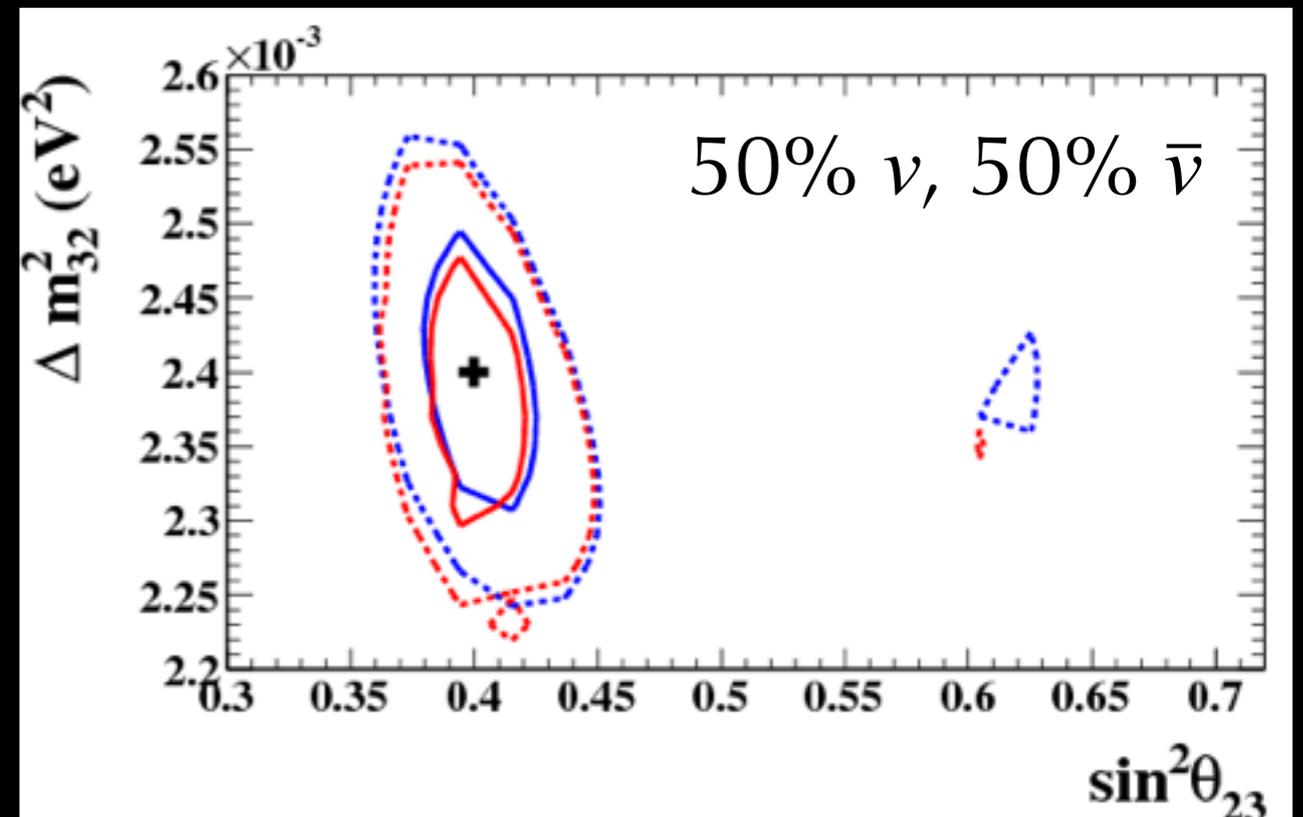
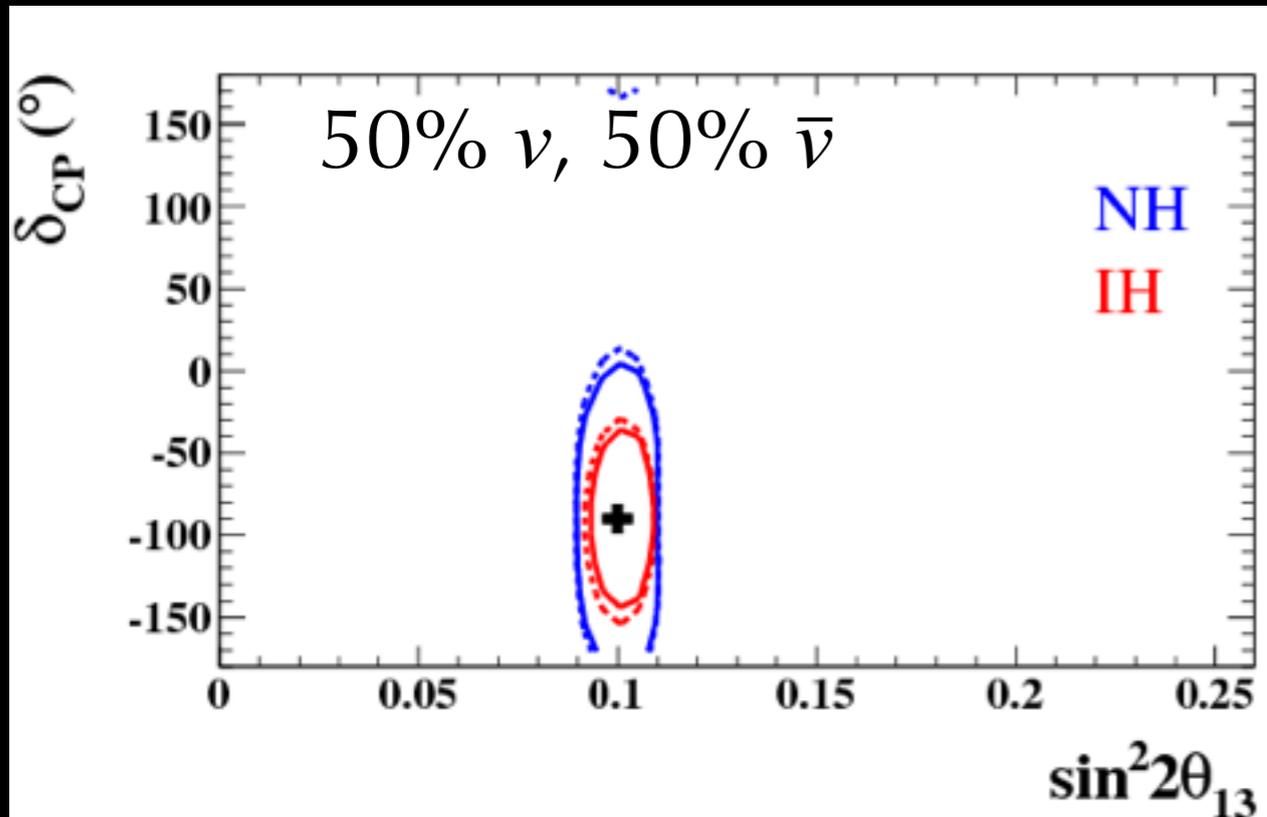
LBNE/LBNO joint studies



- Many studies have assumed “normalization” systematics that scale overall event yields but preserve spectrum shape.
 - doesn’t capture the full picture, but still very useful
 - estimates are becoming more sophisticated
 - Even few % systematic errors can have significant impact
 - how to incorporate potential ND measurements into sensitivity

Recent T2K study:

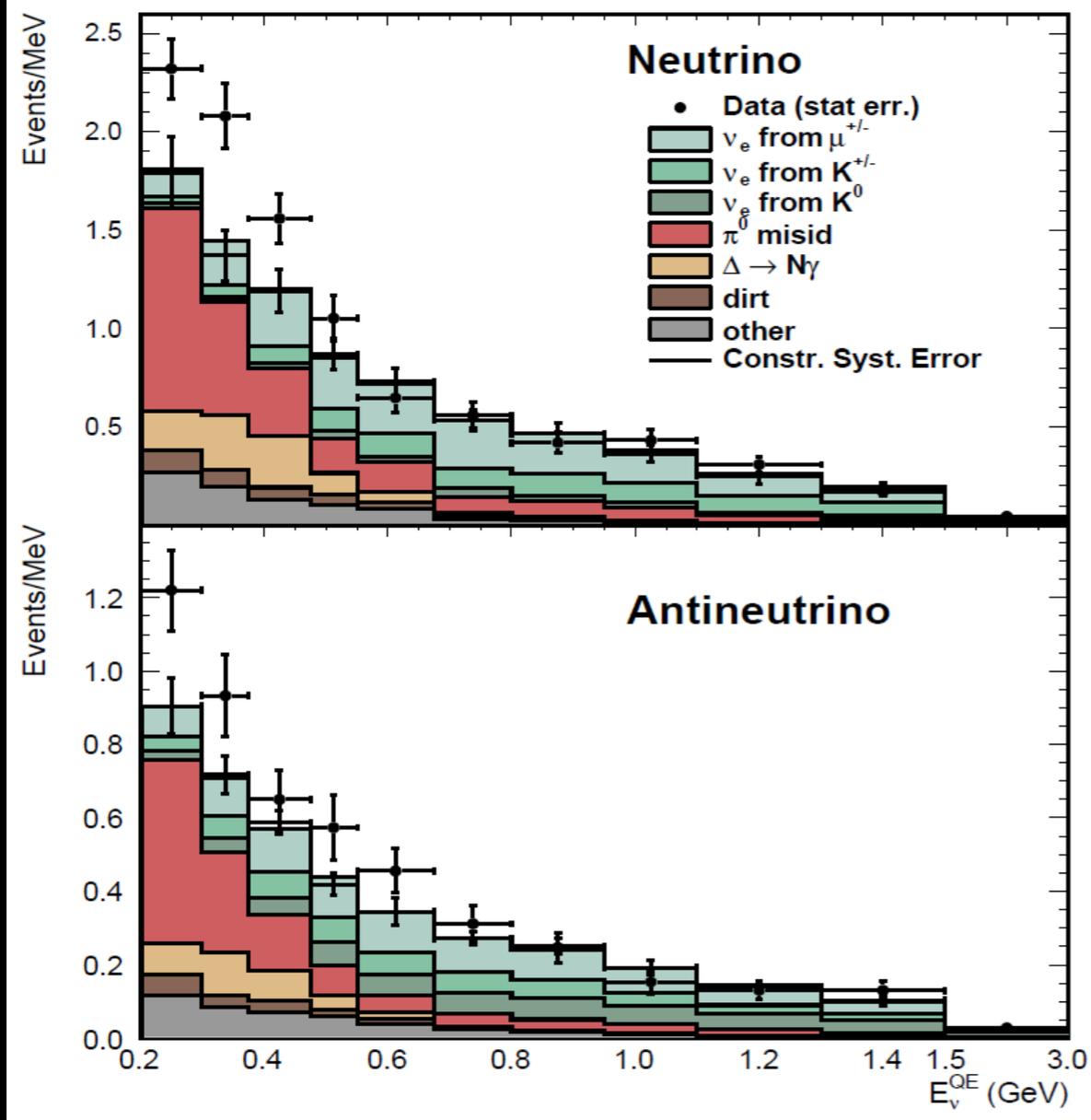
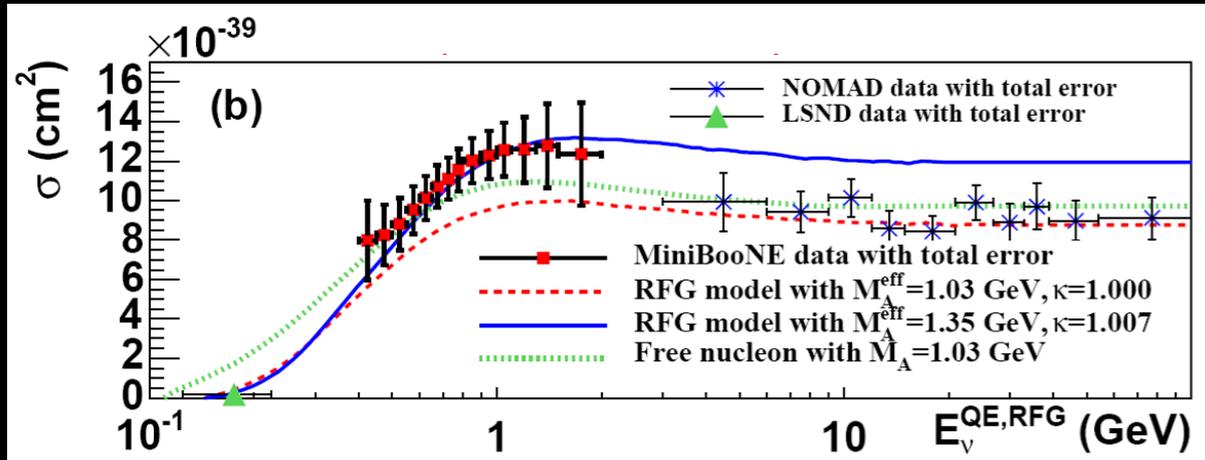
more details in D. Cherdack's talk



- Explore sensitivity to δ_{CP} , θ_{23} in most favorable cases
 - 90% CL contours for “full” T2K statistics (7.8×10^{21} POT)
 - $\delta = -\pi/2$, normal hierarchy
 - reactor constraint
 - systematics as of 2012

Issues

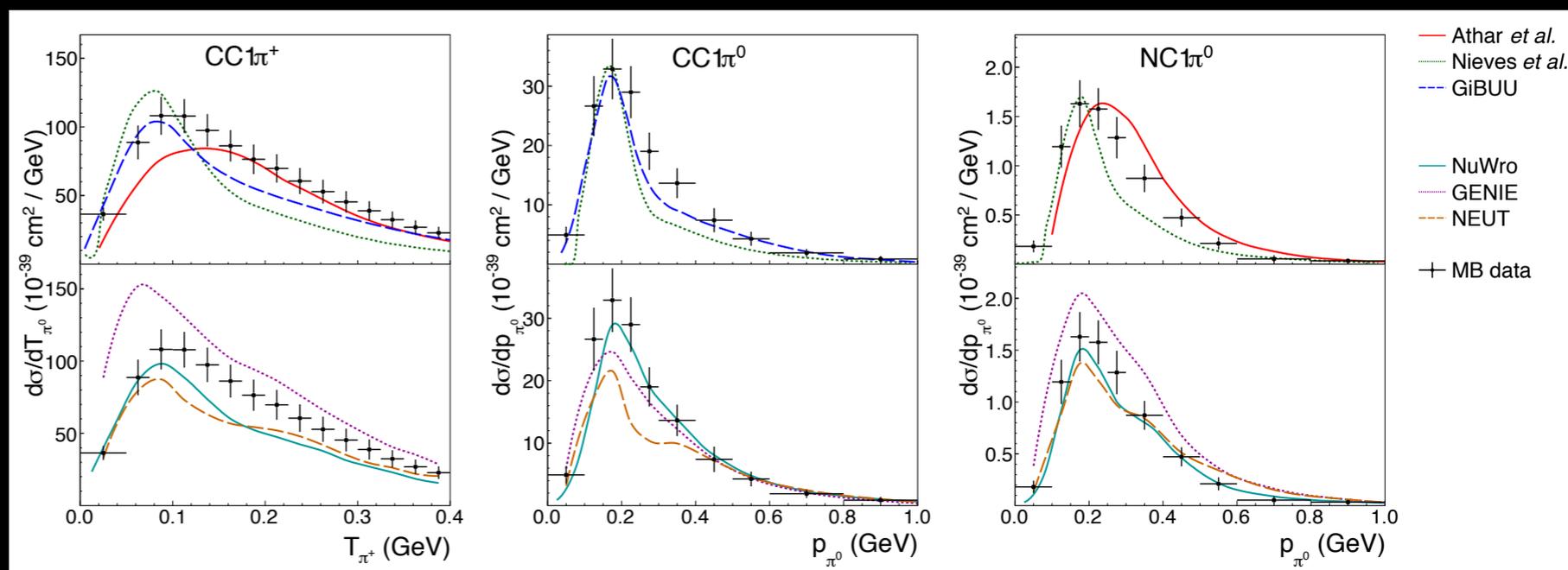
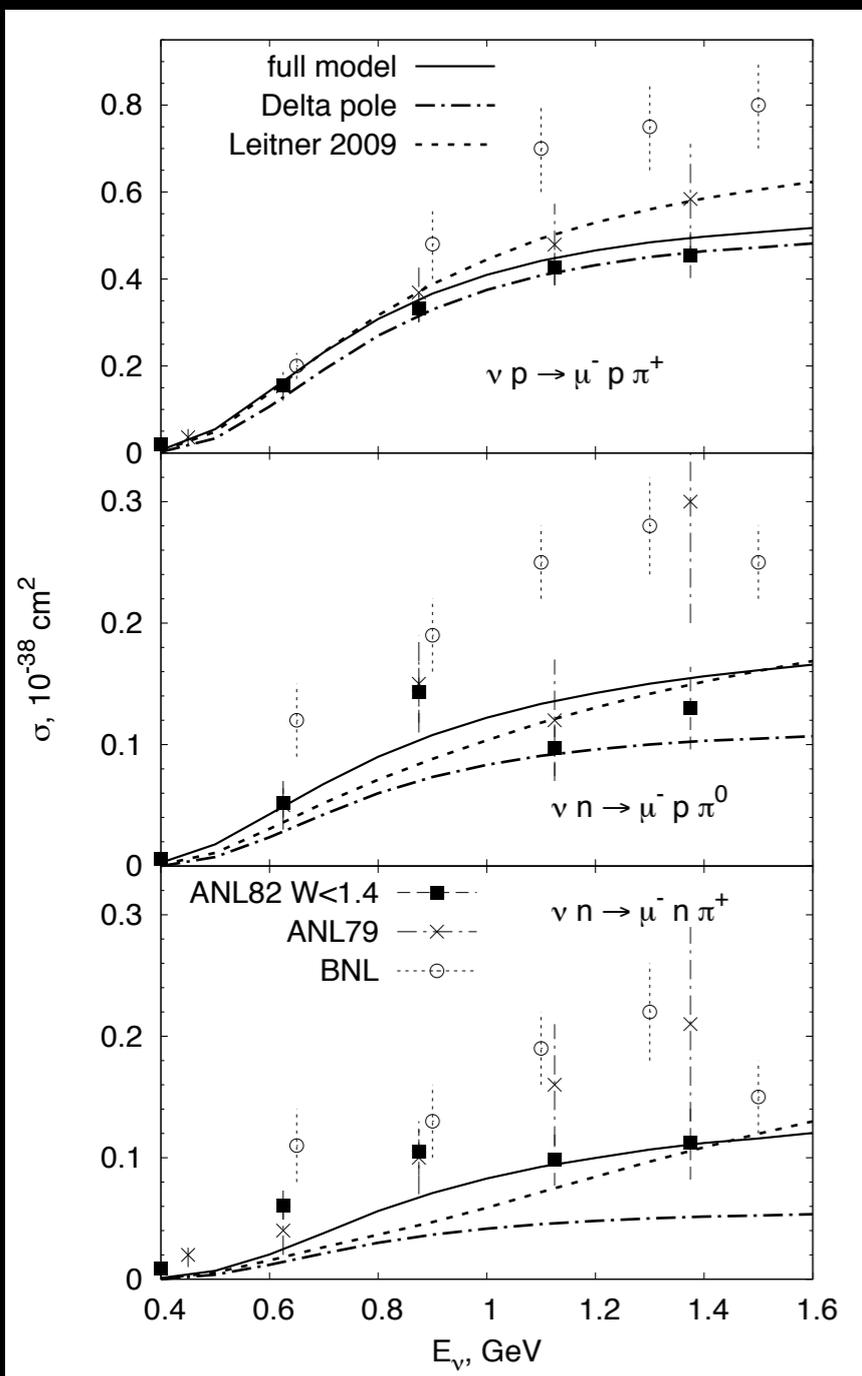
Recent issues



- *ab initio* ν -int uncertainties are driven to a large extent by:
 - data/model discrepancies
 - model inadequacies
- Examples:
 - MiniBooNE “CCQE” σ
 - low energy excess in ν_e
- Model inadequacy means:
 - “inflated” parameter errors
 - “if the data don’t fit, you must . . .”
 - explore other models to span the space of ignorance/mismodeling

Other issues

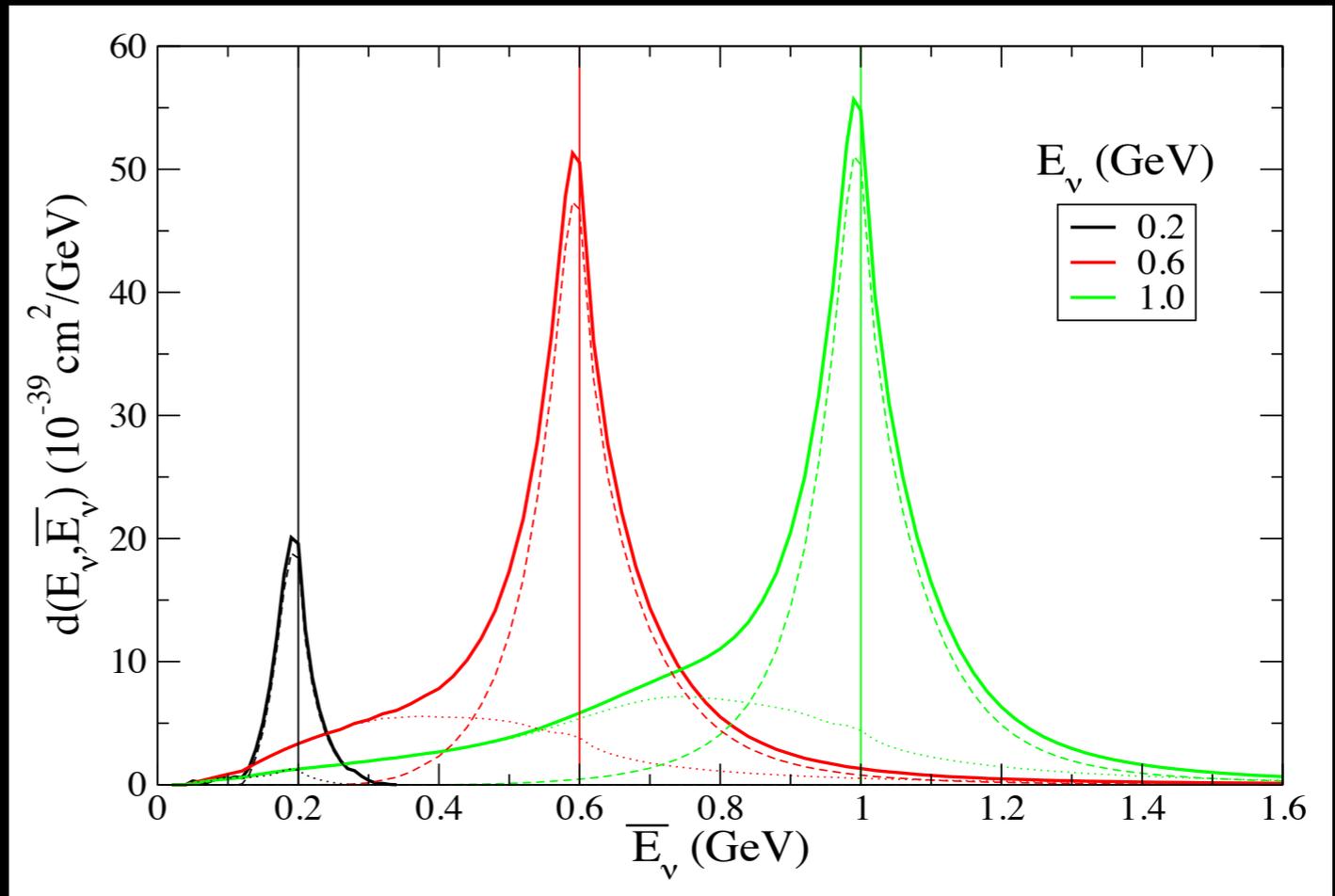
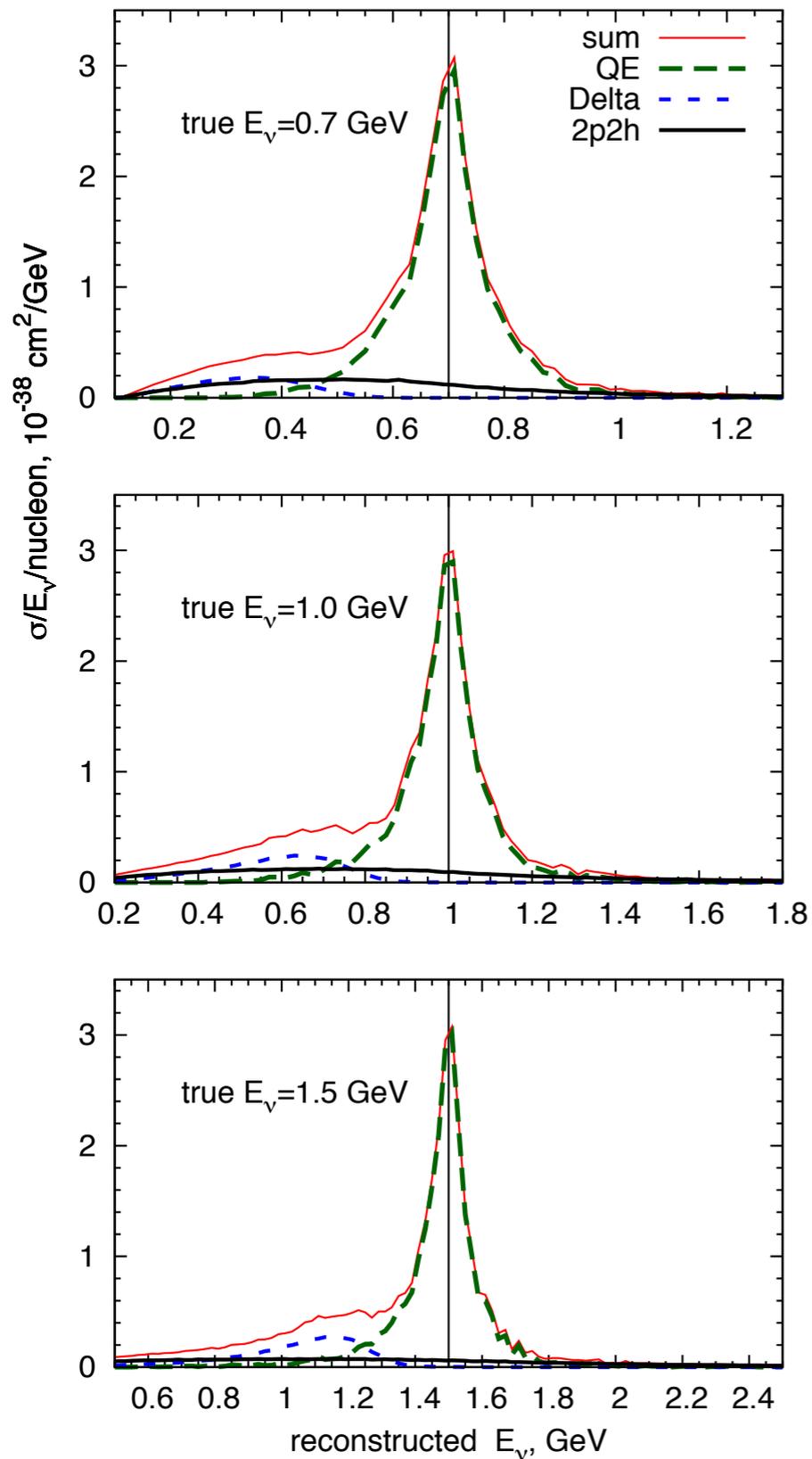
P. Rodriguez
NuInt 2012



- Most model issues are with the nuclear environment
 - but even nucleon-level scattering data has serious discrepancies.
 - Revisit them?
- Need to think “differentially”
 - kinematic distributions important in addition to overall cross section
 - sometimes n-fold-differential
 - cross channel correlations
 - same physics in different ν -int channels

O. Lalakulich et al.
hep-ex/1007.0925v2

Energy reconstruction



- To fully utilize spectral information, we have to confidence in final state kinematics
- Nuclear dynamics to a large extent determines neutrino energy reconstruction
 - not just “new” effects like MEC/multi-N but “old” stuff like π absorption, final state interactions

Near Detector

Near Detector

Far ($\theta_{ij}, \Delta m^2_{ij}$)
 $\nu_\mu \rightarrow \nu_e$ (θ_{23}, θ_{13})
 $\nu_\mu \rightarrow \nu_{\mu/\tau}$ ($\theta_{23}, \Delta m^2_{32}$)
 ν_μ, ν_e backgrounds

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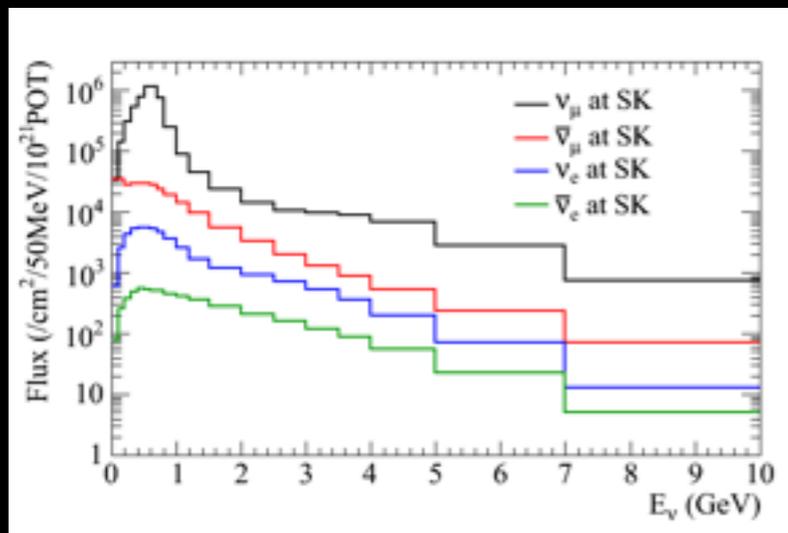
$$\Phi_{\nu} \cdot \sigma_{\nu} \cdot \epsilon_{\text{FAR}} \cdot P_{\text{osc}}$$

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$$\Phi_{\nu} \cdot \sigma_{\nu} \cdot \epsilon_{\text{FAR}} \cdot P_{\text{OSC}}$$

Φ_{ν}



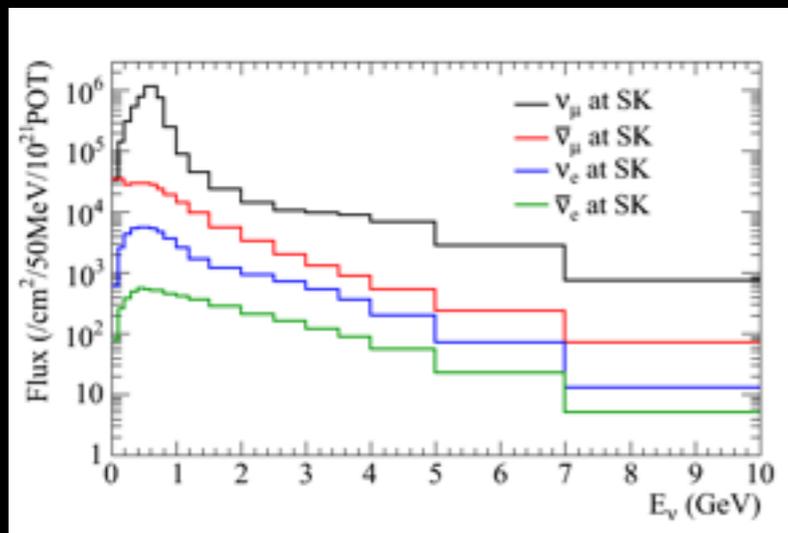
MC simulation of neutrino
beamline tuned with external
data + operational parameters

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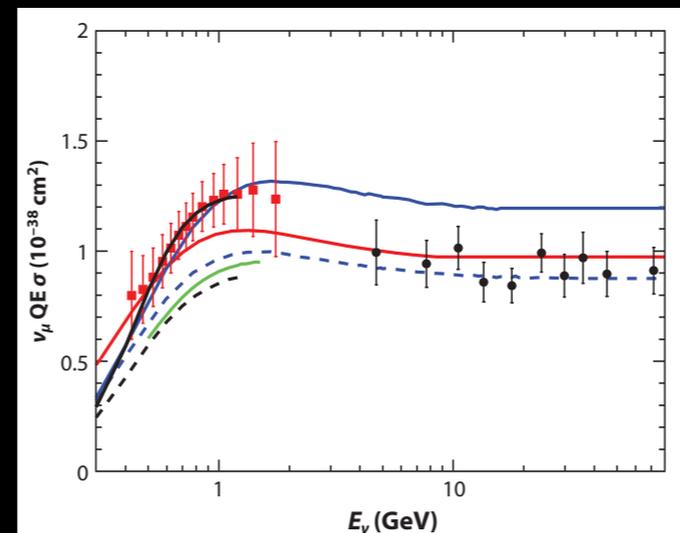
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Φ_ν



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σ_ν



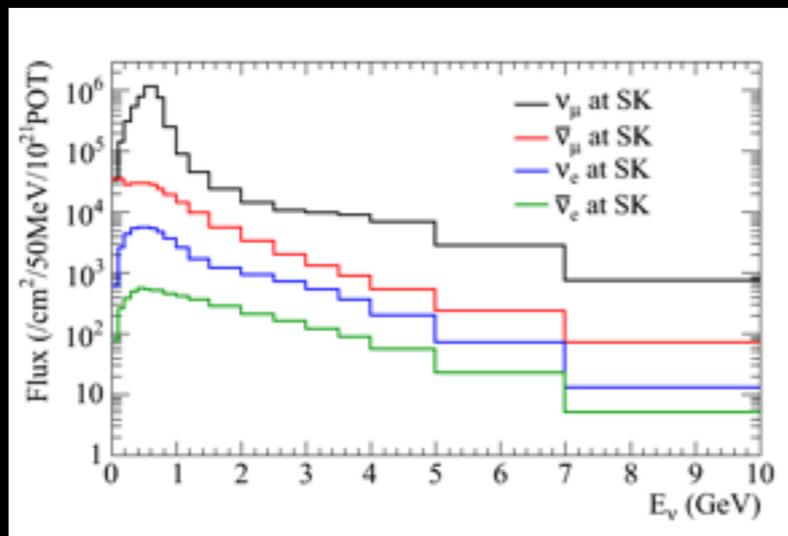
Neutrino cross section and interaction model tuned to external measurements

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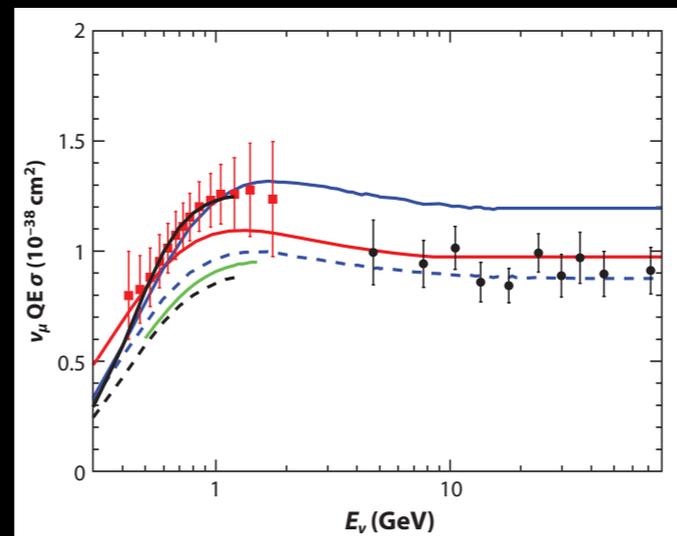
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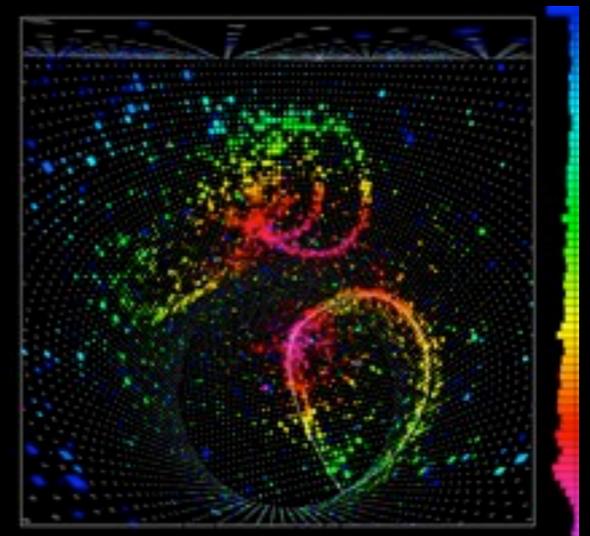
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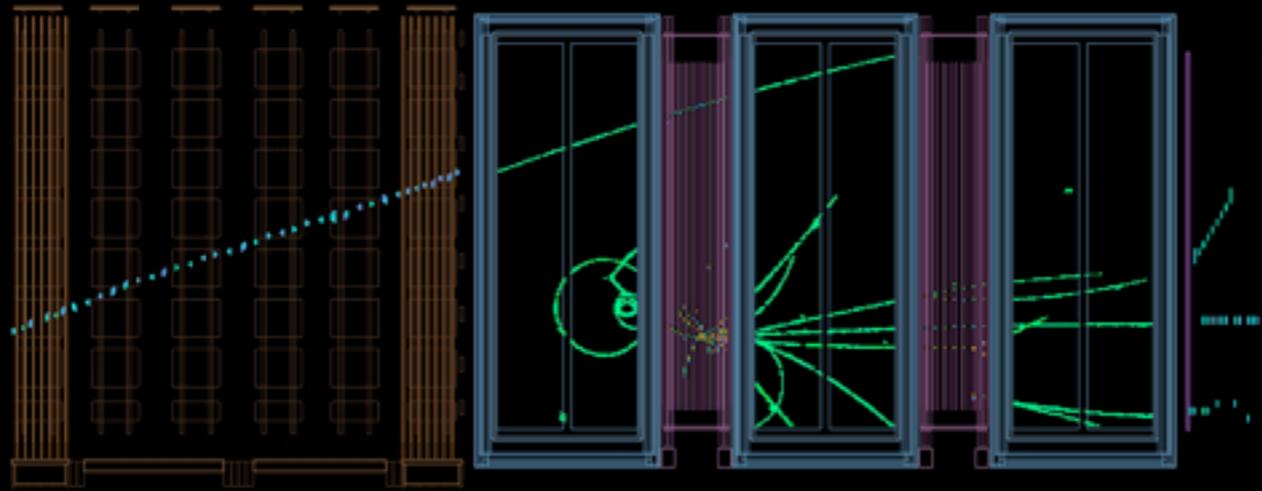
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ϵ_{FAR}



Detector simulation to determine efficiencies/ backgrounds

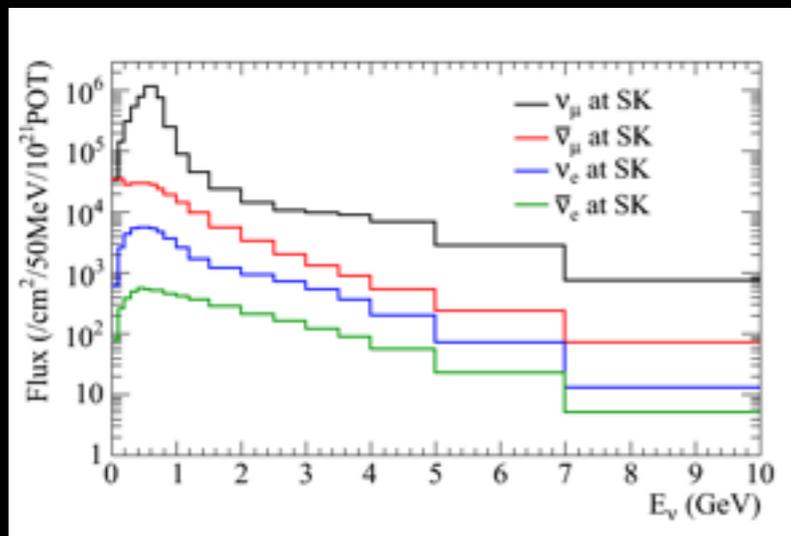
Near Detector



Near detector observes the same neutrinos prior to neutrino oscillations

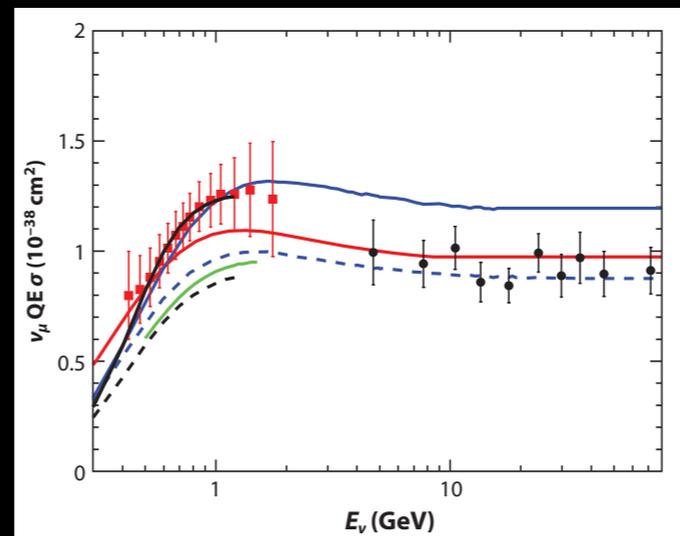
$$\Phi_\nu \cdot \sigma_\nu \cdot \epsilon_{\text{NEAR}}$$

Φ_ν



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σ_ν

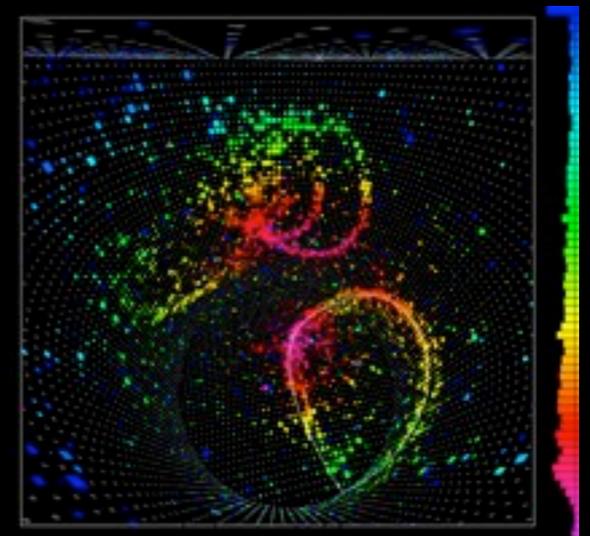


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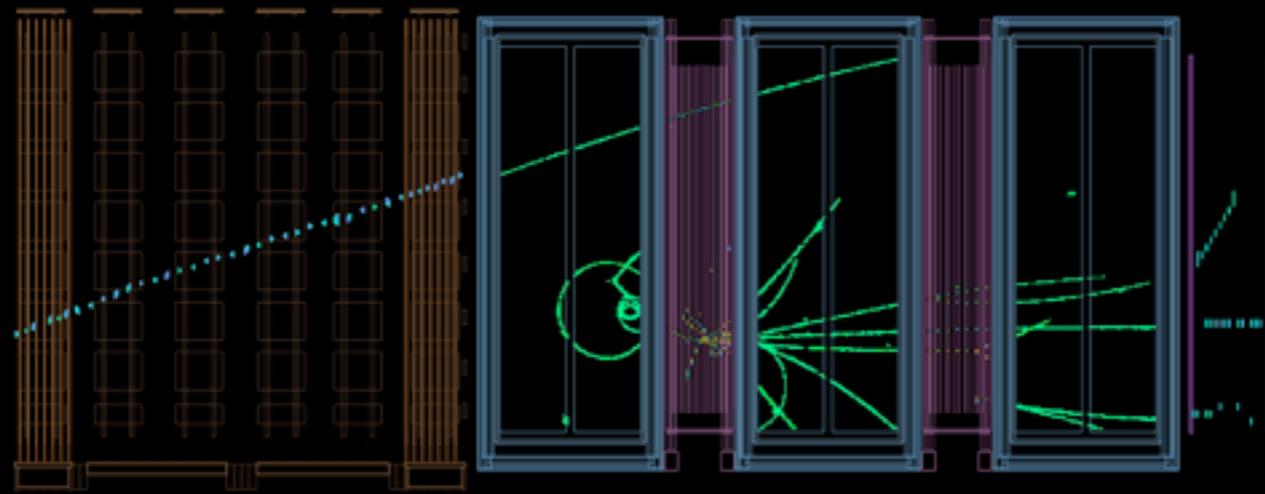
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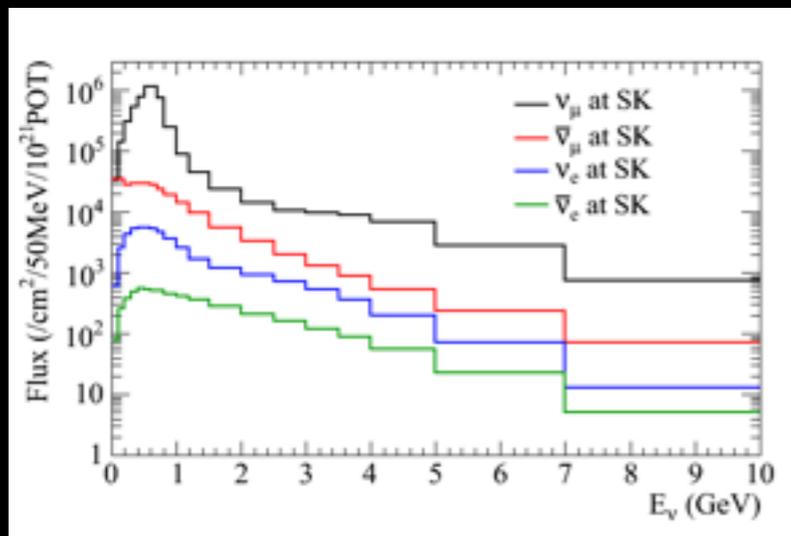
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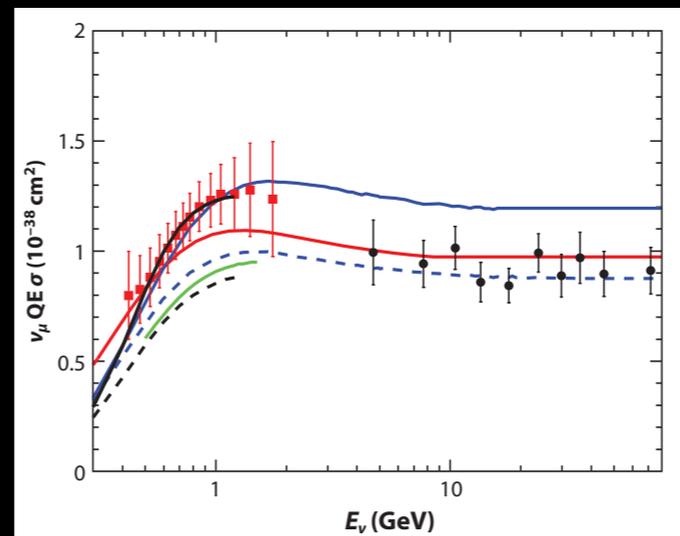
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Φ_ν



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σ_ν

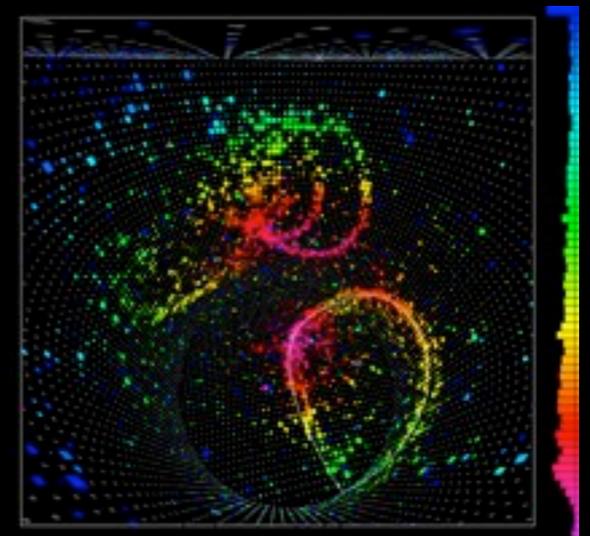


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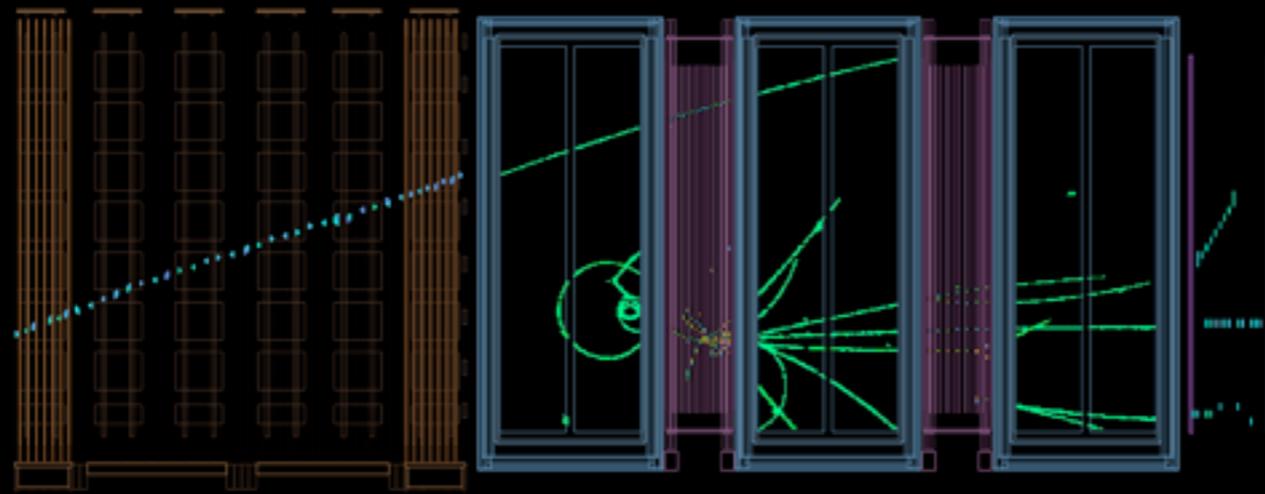
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ϵ_{FAR}



Detector simulation to determine efficiencies/ backgrounds

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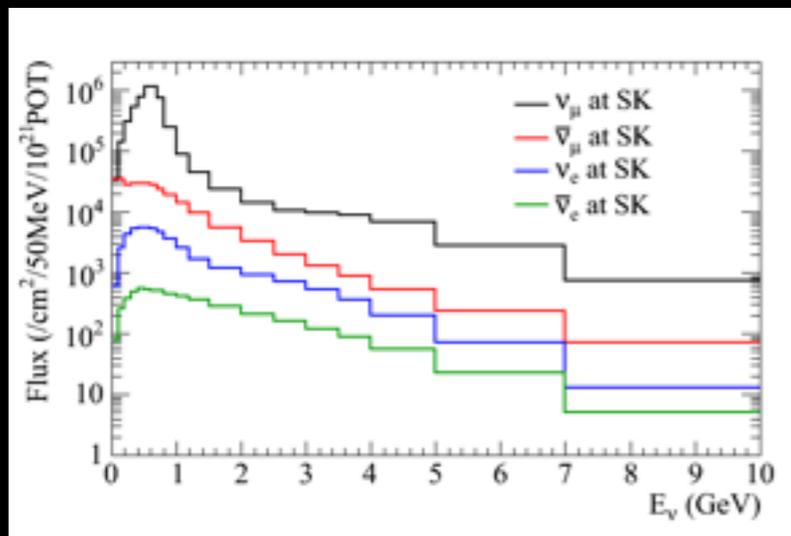


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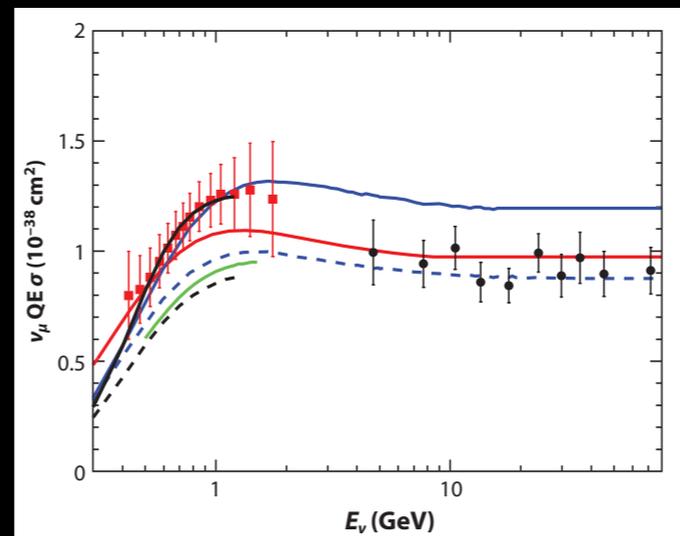
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Φ_ν

σ_ν



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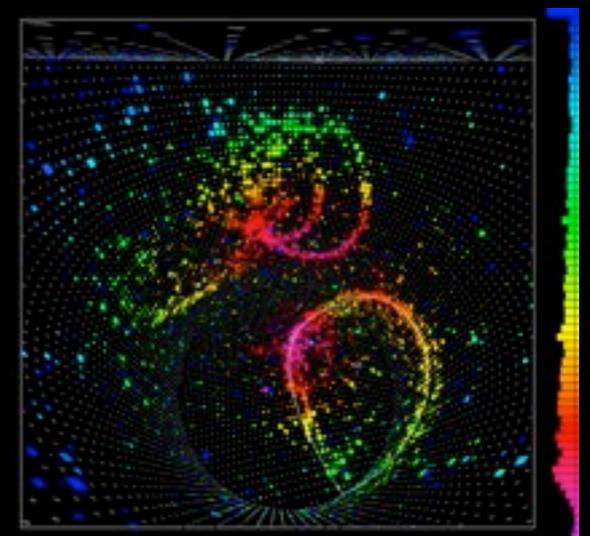


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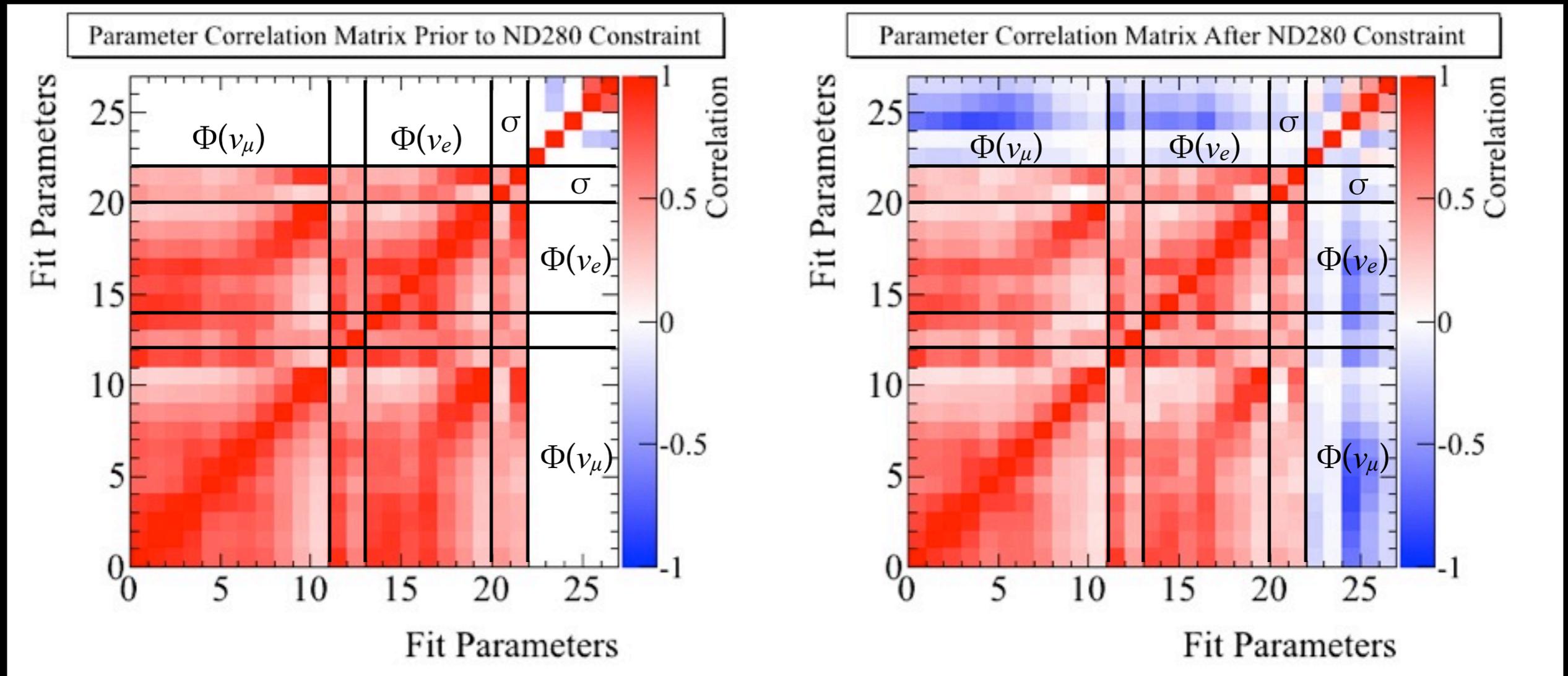
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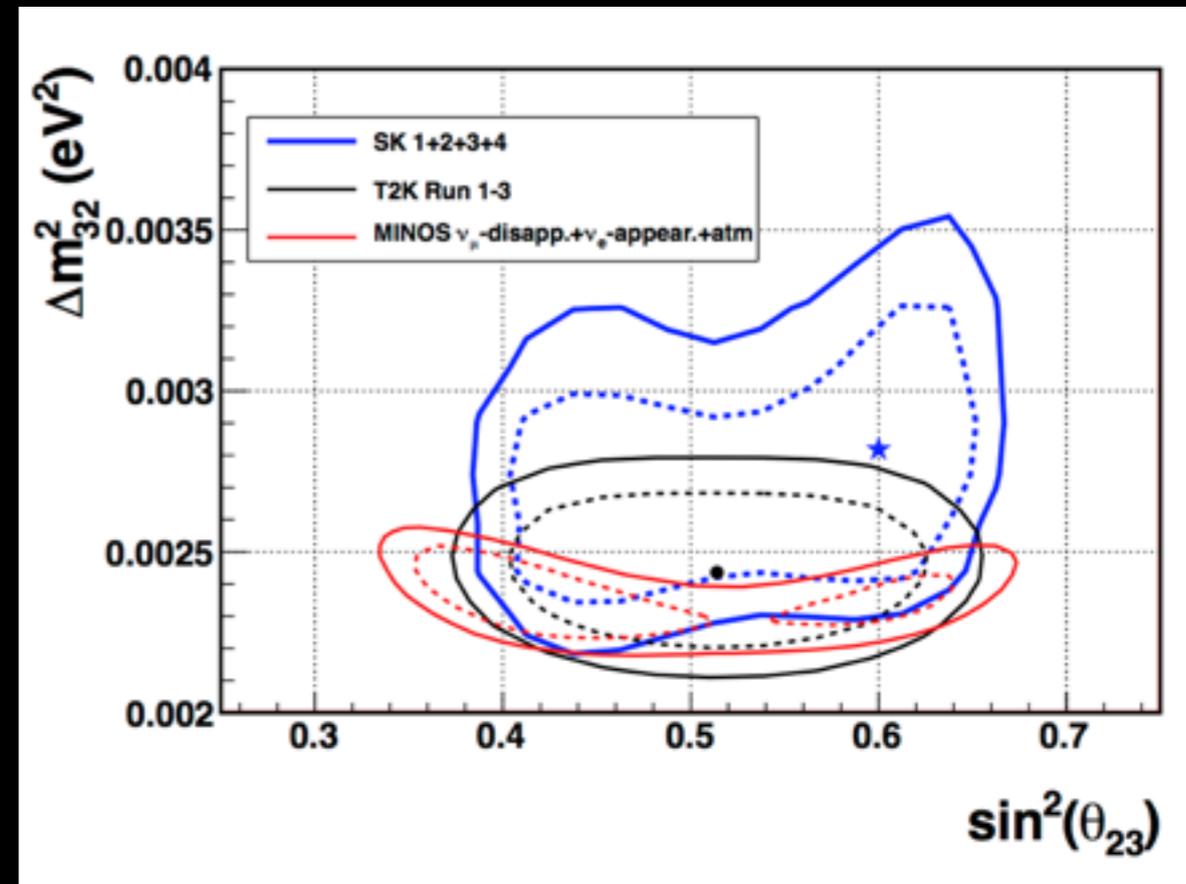
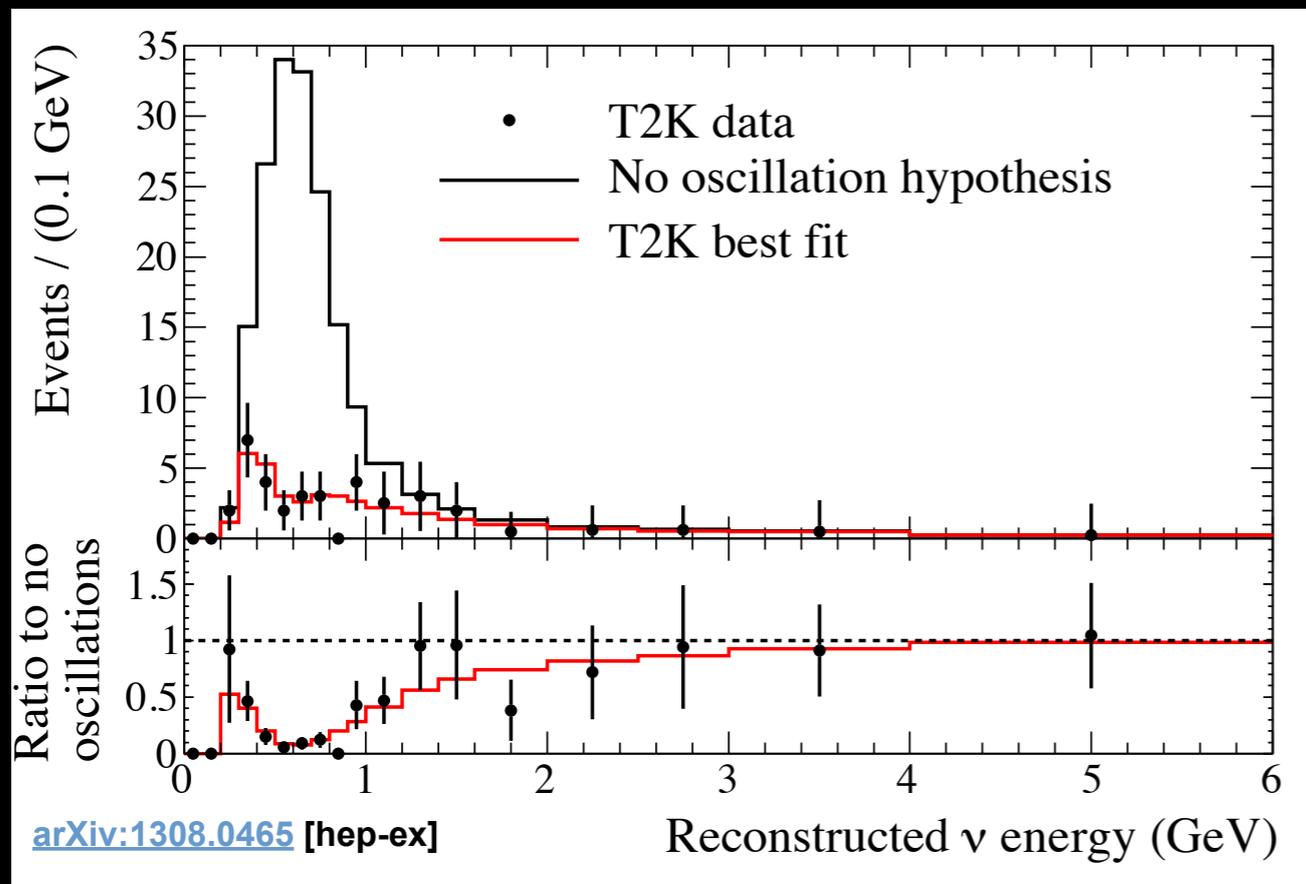
T2K ND fit

more details in D. Cherdack's talk



- Fit to ND data introduces anti-correlations between flux normalization (Φ) and ν cross section (σ) parameters
- partial cancellation in predicted rate from $\sim 28\%$ to $\sim 9\%$
 - $\sim 3\%$ due to parameters directly constrained by the ND data
 - question of “dominant” uncertainties is non-trivial

ν_μ disappearance



$$P(\nu_\mu \rightarrow \nu_\mu) = 1 - \left(\cos^4 \theta_{13} \sin^2 2\theta_{23} + \sin^2 2\theta_{13} \sin^2 \theta_{23} \right) \sin^2 \Delta m_{31}^2 \frac{L}{4E}$$

- High precision means we must consider higher order terms
 - θ_{13} : now known to be not so small $\sin^2 \theta_{13} \sim 0.03$
- Precision ν_μ disappearance will be a testing ground for latest developments in ν -int theory/modeling and associated systematics

Near Detector

Top/down

- Design near detector to
 - maximize near/far cancellation
 - minimize ν -int model dependence
- Advantages:
 - go directly to the issue without dealing with detailed ν -int physics
 - usually calls for similar/same near/far detector, simplifies analysis.
- Disadvantages:
 - there will always be some model dependence
 - data may not enlighten us on how to improve model (TLI)
 - how to estimate systematics with model you don't believe

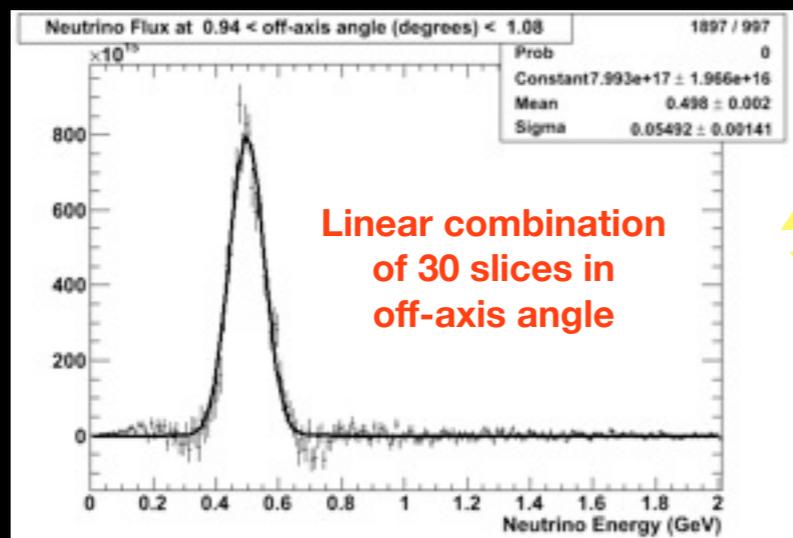
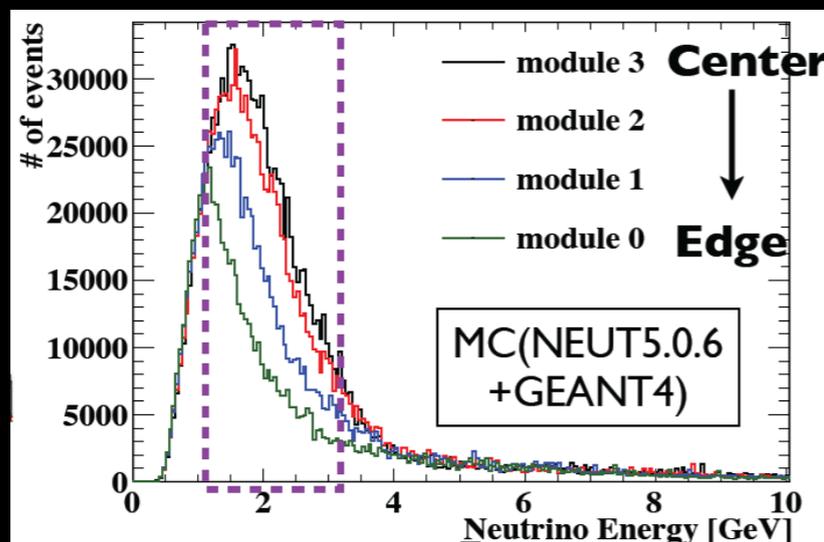
- Which is better? “both” (and not exclusive)

Bottom/up

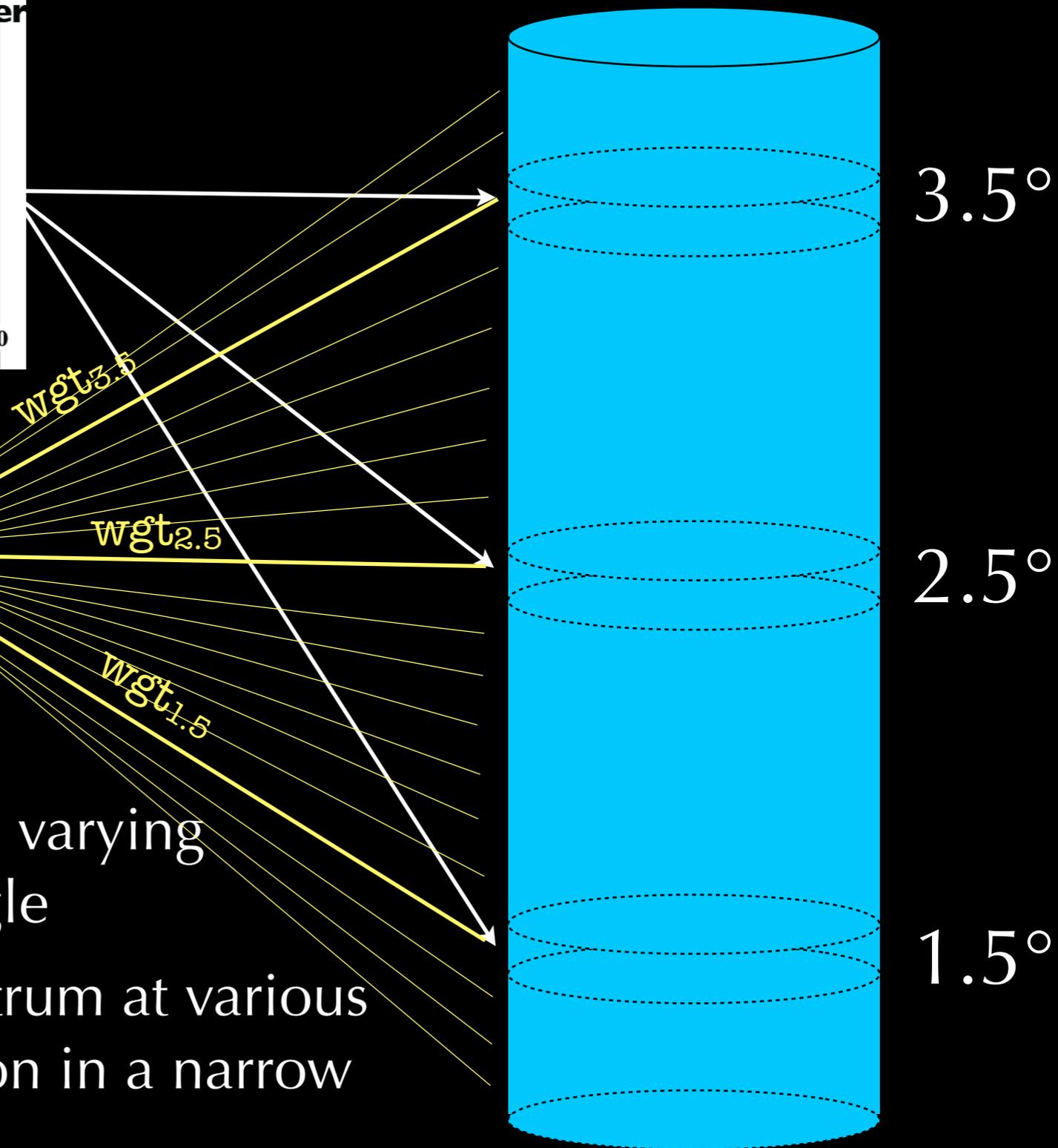
- Design near detector to
 - study details of ν -int interactions
 - advance theory and modeling
- Advantages
 - more freedom to design detector to optimize “information”
 - more information to understand and verify ν -int model
- Disadvantages:
 - may not tell you what you need to impact osc. physics immediately
 - detector/analysis may be more complicated (TMI), substantially different from far detector

Neutrino Prism:

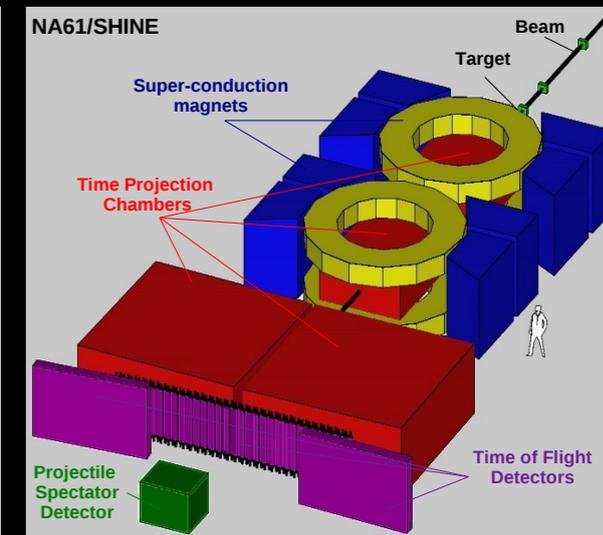
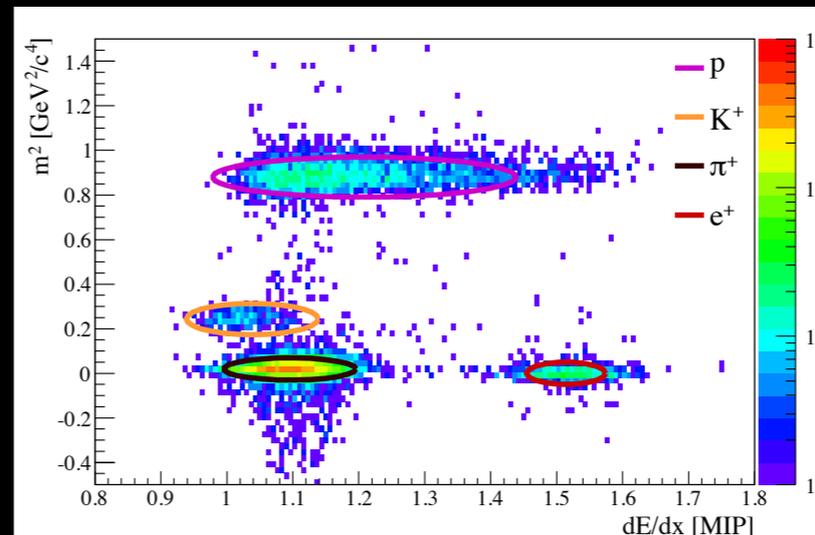
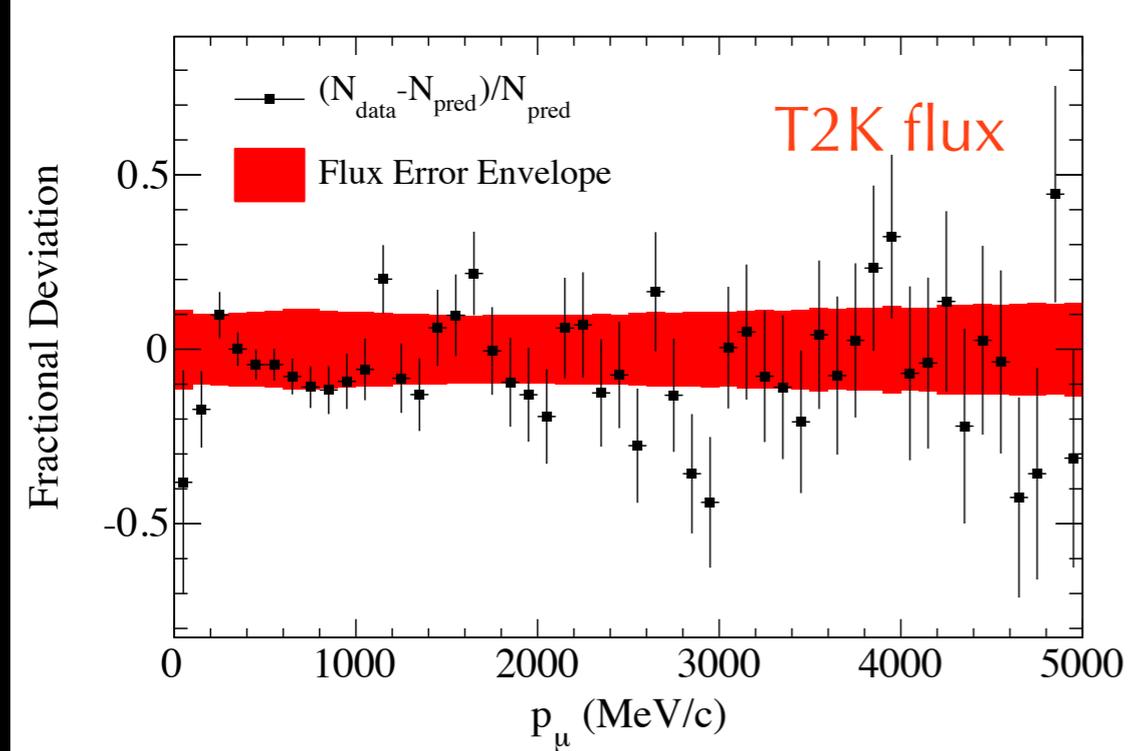
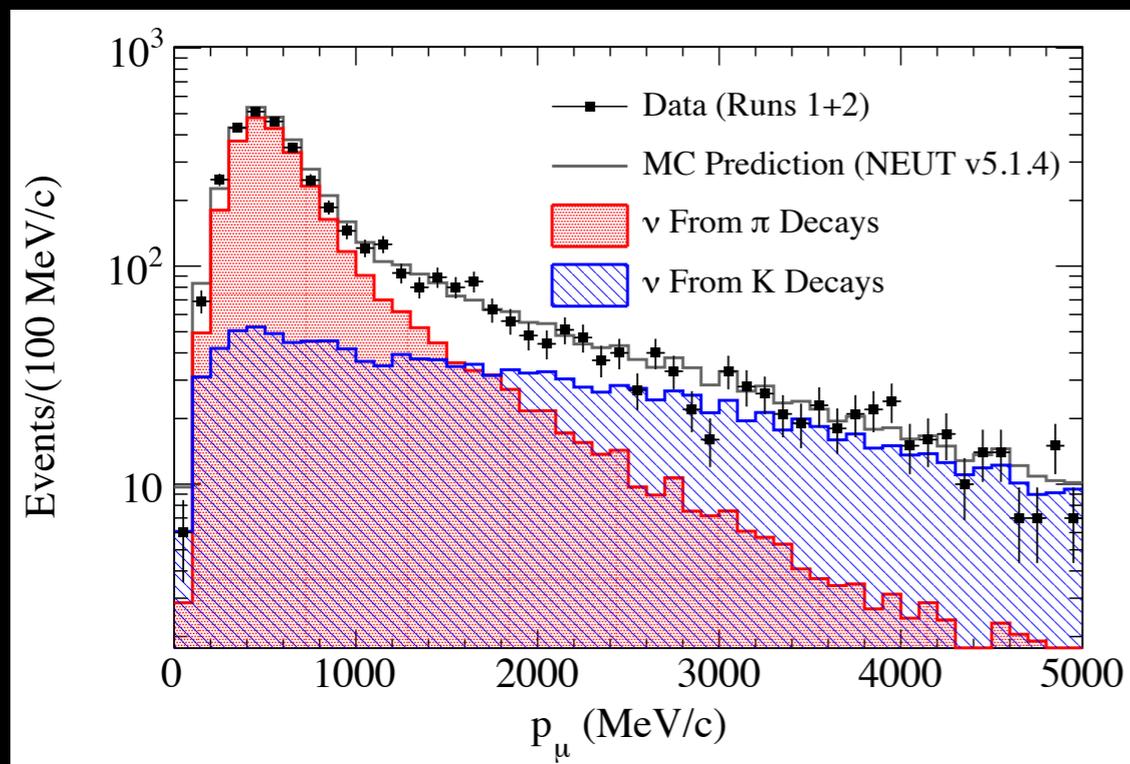
see poster by M. Hartz, M. Wilking
for more details



- Off-axis effect gives continuously varying neutrino energy spectrum vs. angle
- Linear combinations of flux spectrum at various angles maps kinematic distribution in a narrow neutrino energy band
- “Model independent” measurement

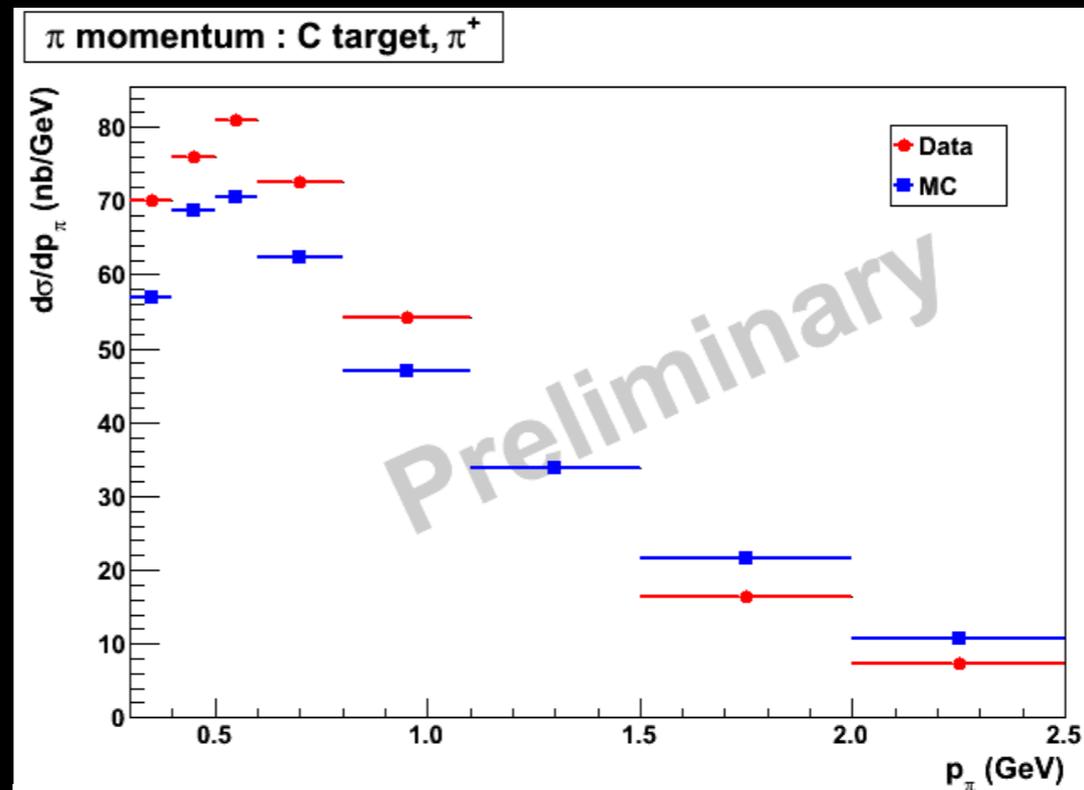
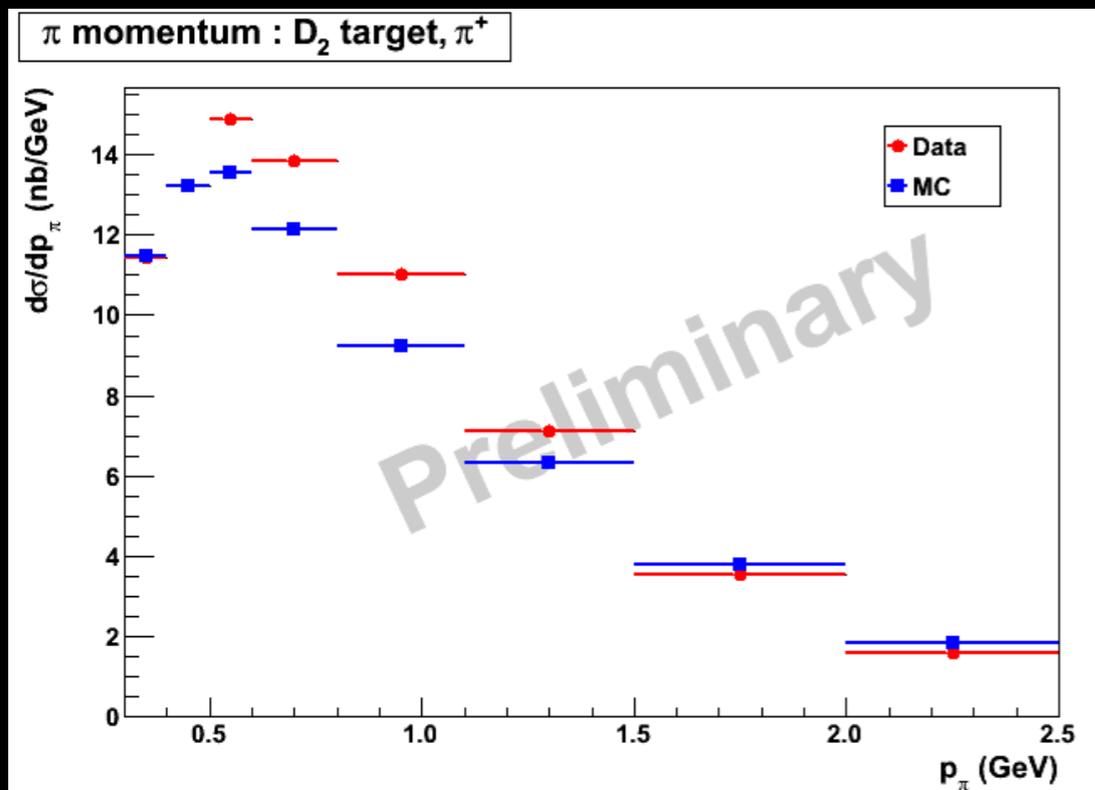


NA61

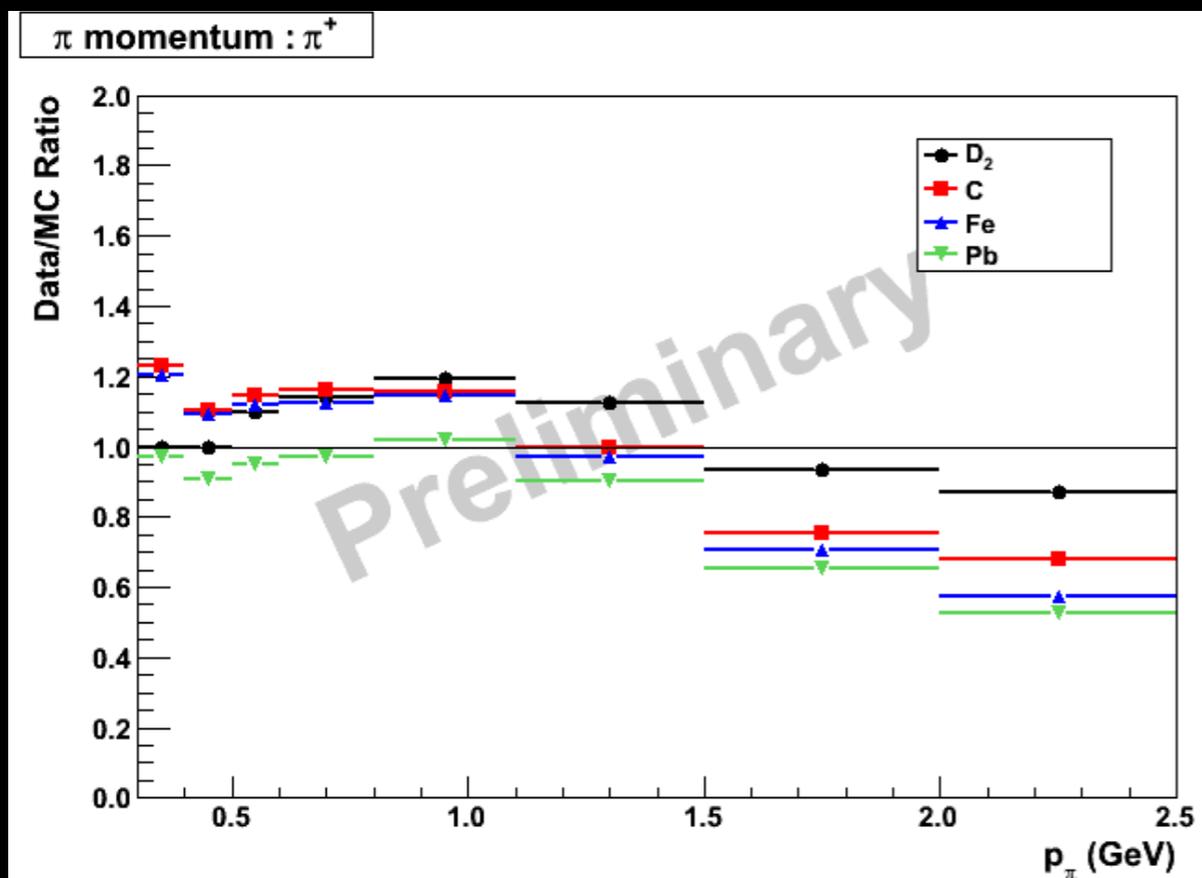


- Flux predictions will continue(?) to be a fundamental pillar for neutrino cross section/interaction physics
- Hadron production experiment(s) are essential for accurate predictions
- Can we make NA61 the ultimate hadron production experiment?
 - Are there other measurements needed/possible to reduce errors?
 - (pion scattering/absorption . . .)
 - test beams for detector performance.

e-A scattering and ν STORM



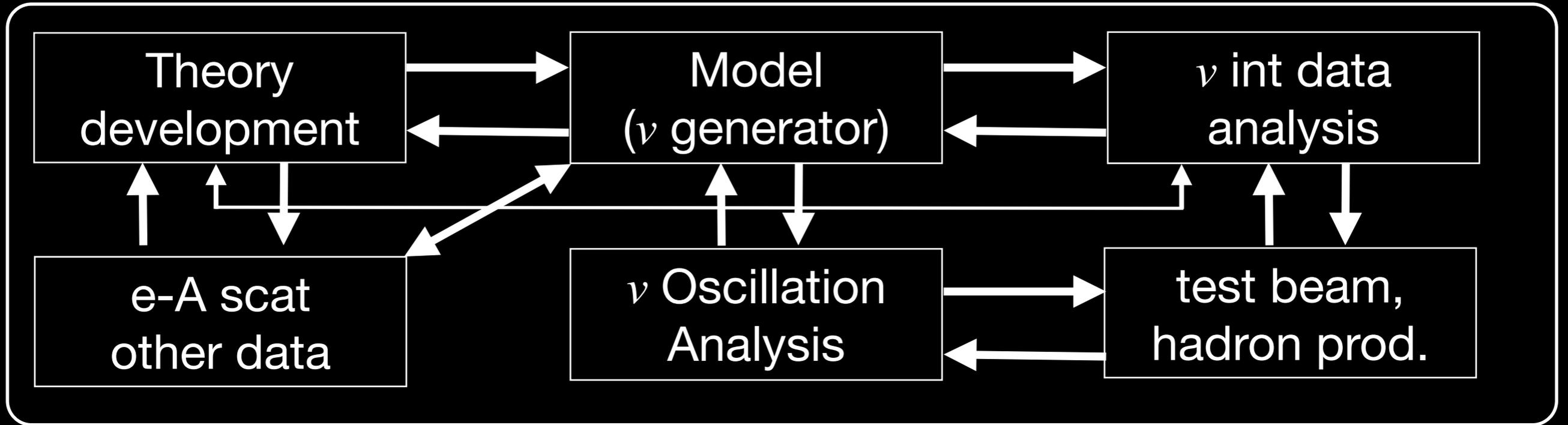
S. Manly
NuFact 2013
comparison to
Genie e-A scat.



- e-A scattering data may be a fruitful testing ground of FSI models
 - data over large range of target nuclei
 - first results from CLAS
 - an enormous effort to understand CLAS analysis tools
- Potential of μ SR-based neutrino sources for ν -int should be further explored

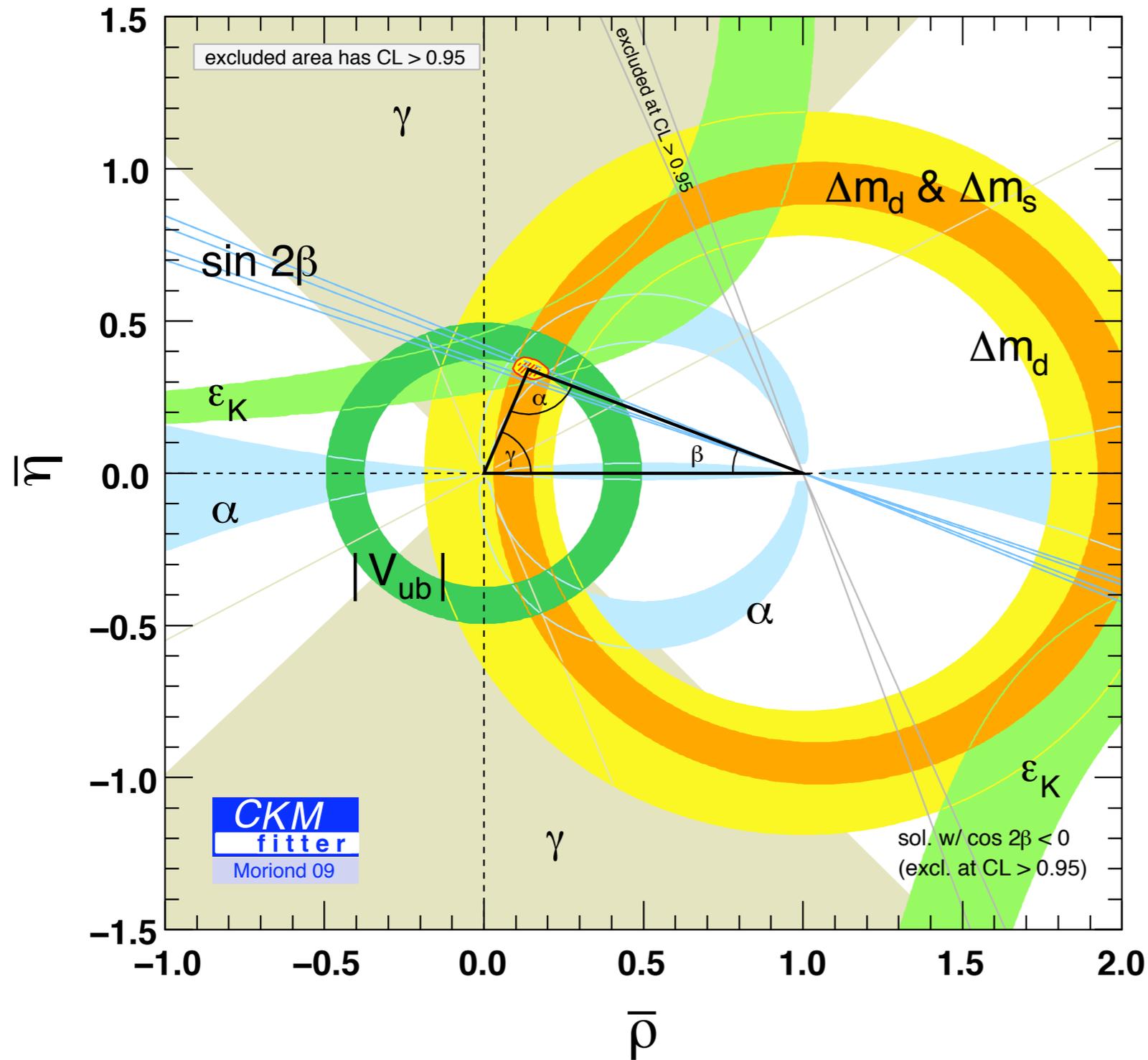
Synthesis

adapted from D. Schmitz
(FNAL W&C)



- better theory may not immediately result in “improvement”
 - worse model may “effectively” better reproduce the data
 - better theory may improve certain parts, but leave out other issues
 - example: spectral function?
- as a community we need to improve the flow. . . .
 - smoother/porous interface at theory/generator interface
 - lots of progress recently, but we need to continue to improve and support

Looking ahead:



We need to do this in the lepton sector!

A. Gouvêa
ISOUPS 2013

"this" = "overconstrain" the mixing matrix
find/rule out new physics

Theory

- To look for “new” physics, we need know the “old” physics.
 - a few “golden modes” (e.g. $B \rightarrow \psi K$) with minimal theoretical issues
 - beyond that, QCD corrections were often very complicated and difficult
 - even “silver” modes like $b \rightarrow s \gamma$ required extremely complicated calculations to predict (beyond-)SM branching fractions
- Enormous effort invested into developing the necessary theoretical tools
 - lattice QCD (masses, form factors)
 - heavy quark effective theory and other factorization schemes
 - sum rules, etc.
 - other “work arounds” like ratios, isospin decomposition, etc.

circa 2002

Theory

Experiment

$B^0 \rightarrow K^{*0} \gamma$	$B^+ \rightarrow K^{*+} \gamma$	$B^0 \rightarrow \rho^0 \gamma$	$B^+ \rightarrow \rho^+ \gamma$
7.1 ± 2.5	7.5 ± 2.5		0.16 ± 0.05
7.9 ± 3.5	7.9 ± 3.5		
7.2 ± 2.7	7.2 ± 2.7	0.049 ± 0.017	0.085 ± 0.30
$4.55^{+0.72}_{-0.68} \pm 0.34$	$3.76^{+0.89}_{-0.83} \pm 0.28$	< 1.7	< 1.3
$4.96 \pm 0.67 \pm 0.45$	$3.89 \pm 0.93 \pm 0.41$	< 1.06	< 0.99

Quarks and Leptons

- LBL experiments should make a similar investment into fundamental neutrino interaction physics to deliver on our promises
- The quark sector has a few advantages:
 - QCD is “particle physics”.
 - Much of ν -int physics is separated by particle/nuclear divide
 - cultural/social/funding/hosting issues
 - To a large extent, theory/experimental effort factorized
 - i.e. detailed theory wasn't needed to make the measurement
 - we are heavily dependent on the theory (via generators) to tell us what to measure, estimate systematic errors, etc.
- Where there is a clear opportunity, theory has delivered
 - fully embrace and support the theory we need, articulate the scientific opportunities, and push for it!

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We need to do this in
the lepton sector!

Conclusions

- Many exciting upcoming opportunities in neutrino physics
 - CP violation, mass hierarchy, θ_{23} octant . . . and beyond?
 - Our detailed understanding of O(1 GeV) neutrino interactions will play a crucial role in getting at the “particle” physics
 - can we ensure that we will not be hampered by ν -int uncertainties?
 - Good news is that T2K/NOvA give us regular “check-ups” on progress
- Exciting opportunity to make fundamental advances in ν -int
 - can we actually reach a denouement on some of the issues?
 - forge /cultivate long-term/continuous collaboration with nuclear theory
 - ensure related/necessary measurements/analyses are carried through
 - hadron production, ν STORM?
 - test beams, electron/photoproduction data?
 - explore broadly new near detector concepts
 - start now . . . ?

T2K ν_e events

	$\sin^2 2\theta_{13}=0.1$		$\sin^2 2\theta_{13}=0$	
	ν_e Prediction (Events)	Error from Constrained Parameters	ν_e Prediction (Events)	Error from Constrained Parameters
No ND280 Constraint	22.6	26.5%	5.3	22.0%
ND280 Constraint (2012, Runs 1-3)	21.6	4.7%	5.1	6.1%
ND280 Constraint (this analysis)	20.4	3.0%	4.6	4.9%
Total error (all sources)		8.8%		11.1%

- Near detector constrain can significantly reduce overall rate uncertainty
- Uncertainties now dominated by:
 - uncorrelated/unpropagated ν interaction uncertainties
 - detector uncertainties
 - final state/secondary interactions uncertainties
- “Work in progress”: will continue to improve, but where do we bottom out?

T2K error budget

Error source	$\sin^2 2\theta_{13} = 0$		$\sin^2 2\theta_{13} = 0.1$	
	w/o ND280 fit	w/ ND280 fit	w/o ND280 fit	w/ ND280 fit
Beam only	10.6	10.8	7.3	7.5
M_A^{QE}	15.6	9.5	2.4	4.0
M_A^{RES}	7.2	4.5	2.1	3.9
CCQE norm. ($E_\nu < 1.5$ GeV)	7.1	4.9	4.8	3.8
CC1 π norm. ($E_\nu < 2.5$ GeV)	4.9	5.1	2.4	3.5
NC1 π^0 norm.	2.7	7.9	1.9	7.3
CC other shape	0.3	0.2	0.3	0.2
Spectral Function	4.7	3.3	4.8	3.3
p_F	0.1	0.3	0.1	0.3
CC coh. norm.	0.3	0.2	0.3	0.2
NC coh. norm.	1.1	2.1	1.1	2.0
NC other norm.	2.3	2.6	2.2	2.6
$\sigma_{\nu_e}/\sigma_{\nu_\mu}$	2.4	1.8	2.4	1.8
W shape	1.0	1.9	1.0	1.9
pion-less Δ decay	3.3	0.5	3.1	0.5
SK detector eff.	5.7		5.6	
FSI	3.0		3.0	
PN	3.6		3.5	
SK momentum scale	1.5		1.5	
Total	24.5		11.1	

Black: 2013

Blue: 2012