

A high-angle, wide shot of the MINOS experiment's detector hall. The central feature is a long, narrow, rectangular detector structure with a metallic, reflective surface. It is flanked by two long, parallel rows of blue structural beams and yellow safety railings. The floor is a mix of grey concrete and yellow safety markings. The lighting is bright and even, highlighting the industrial scale of the facility.

# The MINOS Experiment and MINOS+

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**UCL**

**11<sup>th</sup> November 2013**

**NNN13**



# The MINOS Experiment

- ✦ Long-baseline neutrino oscillation experiment

- ✦ L/E ~500 km / GeV

- ✦ Exposed by the NuMI beam

- ✦ Runs off the Main Injector at Fermilab, USA

- ✦  $\nu_\mu$  or  $\bar{\nu}_\mu$  beam mode

- ✦ Two detector system used to minimise systematics:

- ✦ Beam flux

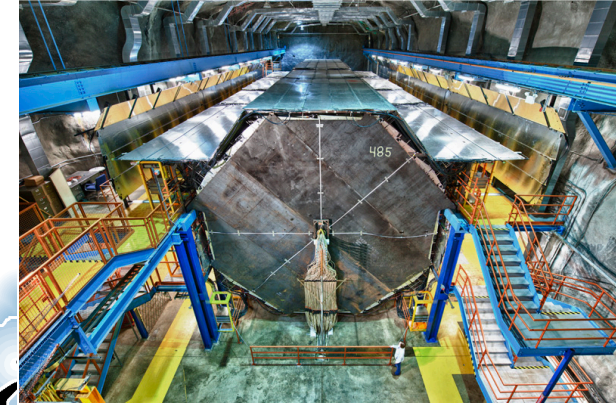
- ✦ Neutrino interaction cross-sections

- ✦ Look for:

- ✦  $\nu_\mu$  disappearance

- ✦  $\nu_e$  appearance

- ✦ Neutral current events



Far Detector (FD)



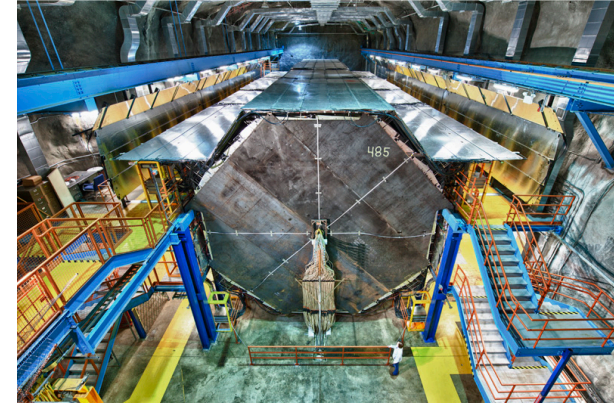
Near Detector (ND)





# The MINOS Detectors

- ✦ Magnetised steel/scintillator sampling tracking calorimeters
  - ✦ Consecutive steel / scintillator planes
  - ✦ Average 1.3T field for charge selection
  - ✦ Functionally equivalent between ND and FD

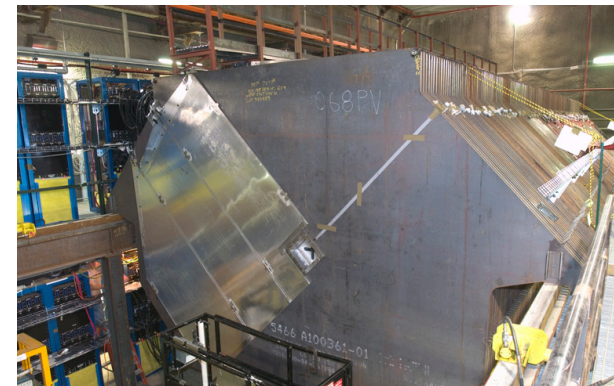


Far Detector (FD)

Detector	Baseline / km	Mass / kton	Start up
Far	735	5.4	2003
Near	1	1.0	2005

Near Detector (ND)

- ✦ ND measures beam before oscillations
- ✦ FD looks for changes in the beam relative to the ND

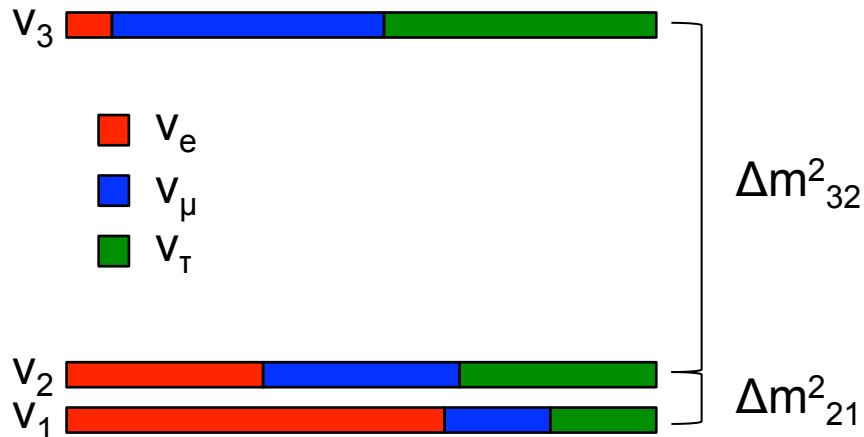


# What do we want to measure?

$$U_{PMNS} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta_{CP}} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta_{CP}} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

## ✦ MINOS can measure:

- ✦  $\theta_{23}$  from  $\nu_\mu$  disappearance
- ✦  $|\Delta m^2_{32}|$  from  $\nu_\mu$  disappearance
- ✦  $\theta_{13}$  from  $\nu_e$  appearance



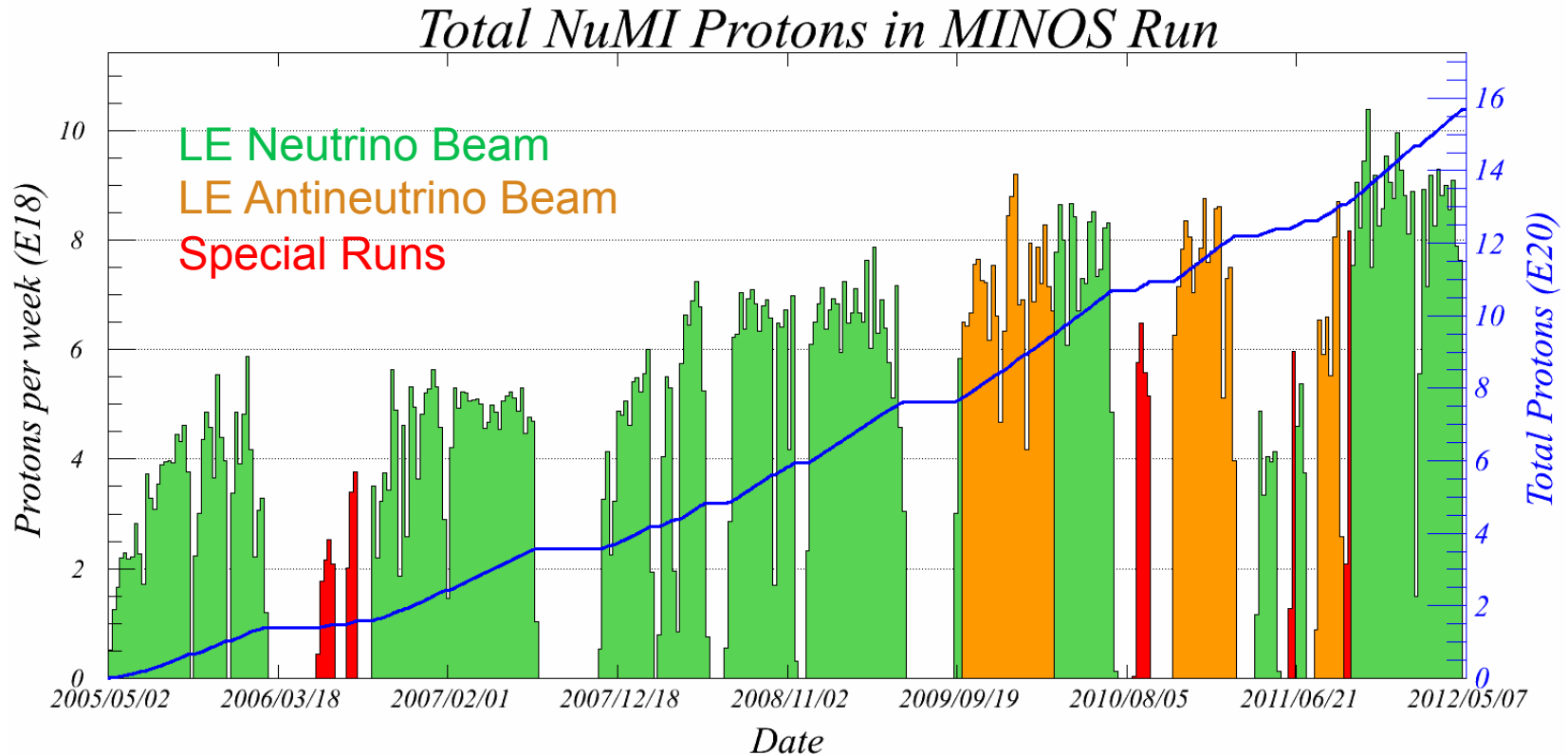
## ✦ Current unknowns:

- ✦ Octant of  $\theta_{23}$ 
  - ✦ Is it maximal?
- ✦ Sign of  $\Delta m^2_{32}$ 
  - ✦ Normal or inverted hierarchy?
- ✦ Value of  $\delta_{CP}$
- ✦ Studying three flavour oscillations will provide the answers!



# Data Samples

- ✦ Use the full MINOS data set.

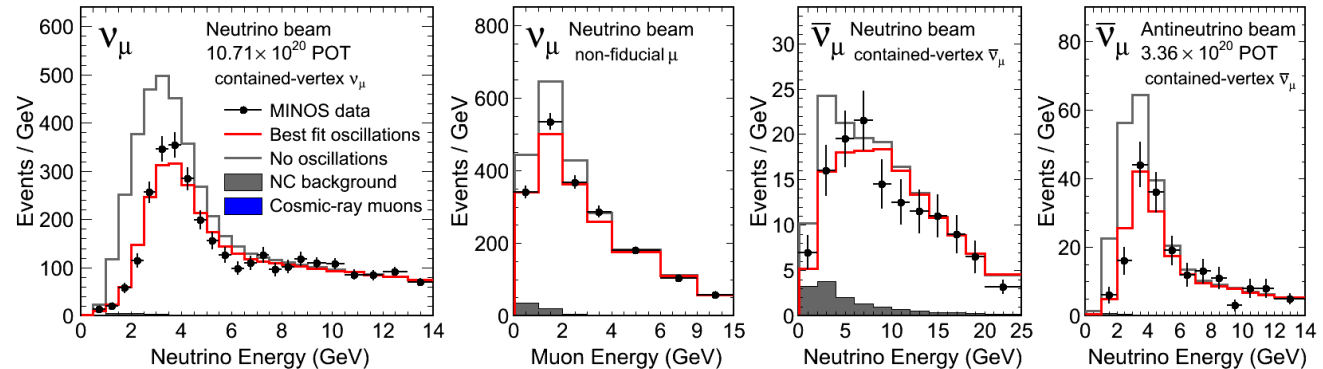


- ✦  $10.71 \times 10^{20}$  POT neutrino mode
- ✦  $3.36 \times 10^{20}$  POT antineutrino mode
- ✦ 37.88 kton years FD atmospheric neutrinos

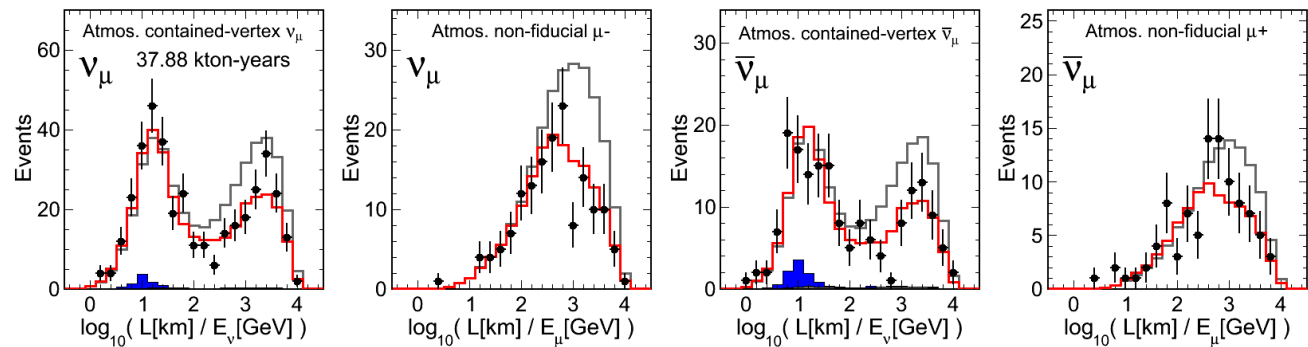
# Disappearance Measurements

- ✦ Predict the FD un-oscillated spectrum
  - ✦ Compare FD data to the prediction.

Beam  
Neutrinos



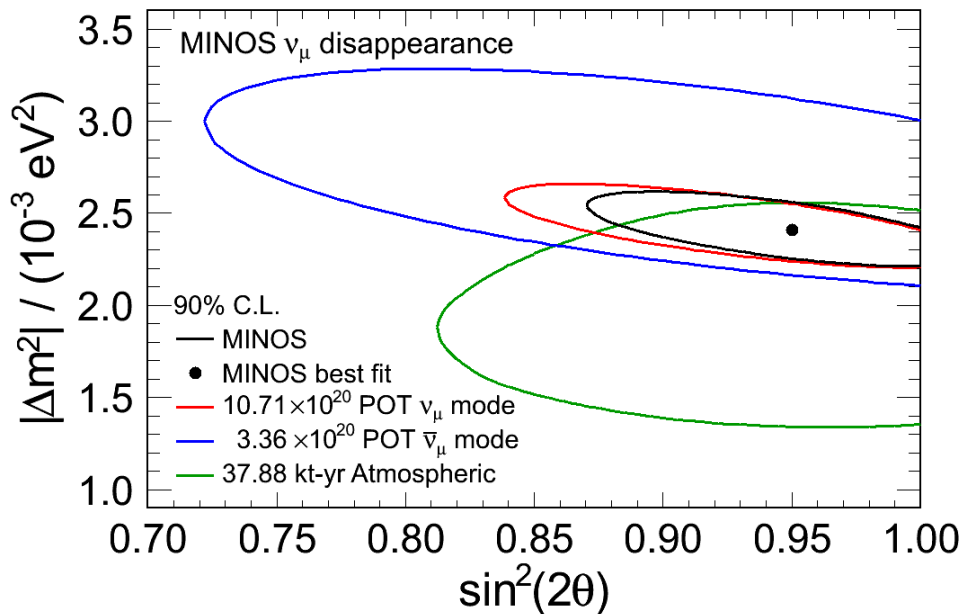
Atmospheric  
Neutrinos



- ✦ Fit the FD under the hypothesis of neutrino oscillations.
  - ✦ Two flavour approximation, or full three flavour
- ✦ Include 15 systematic parameters as nuisance terms

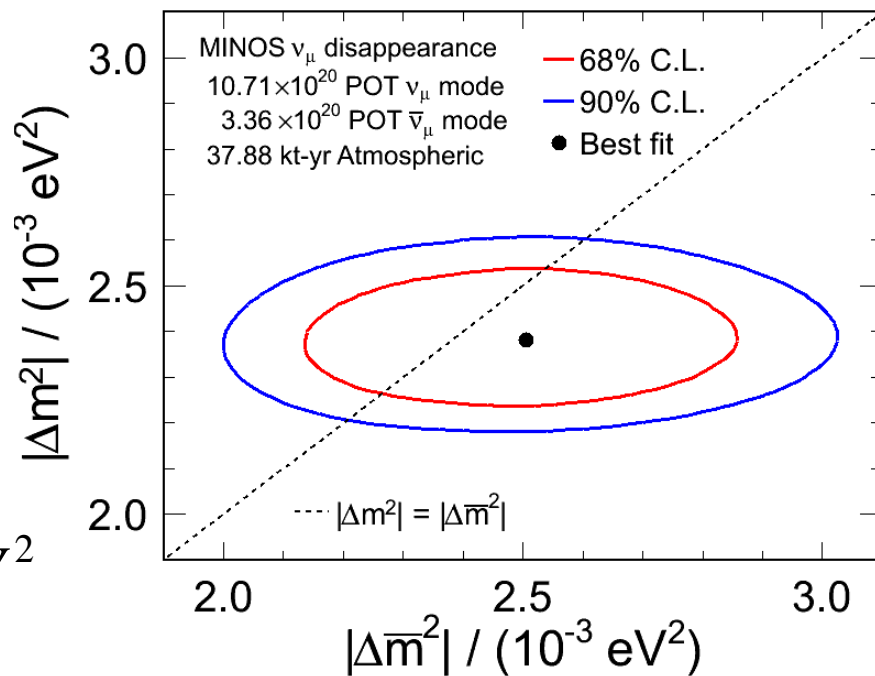
# Two Flavour Muon Neutrino Disappearance

★ Two flavour approximation:  $P(\nu_\mu \rightarrow \nu_\mu) = 1 - \sin^2(2\theta) \sin^2\left(\frac{1.27\Delta m^2 L}{E}\right)$



★ Two parameter fit:

$$\Delta m^2 = 2.41_{-0.10}^{+0.09} \times 10^{-3} \text{ eV}^2$$

$$\sin^2 2\theta = 0.950_{-0.036}^{+0.035}$$


★ Four parameter fit:

$$\left| \Delta \bar{m}^2 \right| - \left| \Delta m^2 \right| = \left( 0.12_{-0.26}^{+0.24} \right) \times 10^{-3} \text{ eV}^2$$

Phys. Rev. Lett. 110, 251801 (2013)



# Three Flavour Muon Neutrino Disappearance

$$P(\nu_\mu \rightarrow \nu_\mu) = 1 - \sin^2(2\theta_{\mu\mu}) \sin^2\left(\frac{1.27 \Delta m_{\mu\mu}^2 L}{E}\right) + \mathcal{O}\left(\Delta m_\odot^2 \frac{L}{E}\right)^2$$

$$\sin^2 \theta_{\mu\mu} = \sin^2 \theta_{23} \cos^2 \theta_{13}$$

$$\Delta m_{\mu\mu}^2 = \Delta m_{32}^2 + \frac{1 - |U_{\mu 1}|^2}{|U_{\mu 3}|^2} \Delta m_{21}^2$$

✦ Disappearance depends on:

- ✦ Solar oscillation parameters
- ✦ Mass hierarchy
- ✦  $\theta_{13}$

✦ Matter effects important for multi-GeV up-going atmospheric

✦ Make a 4D fit to the FD data in  $(\Delta m_{32}^2, \sin^2 \theta_{23}, \sin^2 \theta_{13}, \delta_{CP})$

- ✦ Constrain  $\theta_{13}$  from the reactor experiment average:

$$\sin^2 \theta_{13} = 0.0242 \pm 0.0025$$

✦ Solar parameters fixed at the global average†

$$\Delta m_{21}^2 = 7.54 \times 10^{-5} eV^2$$

$$\sin^2 \theta_{12} = 0.307$$

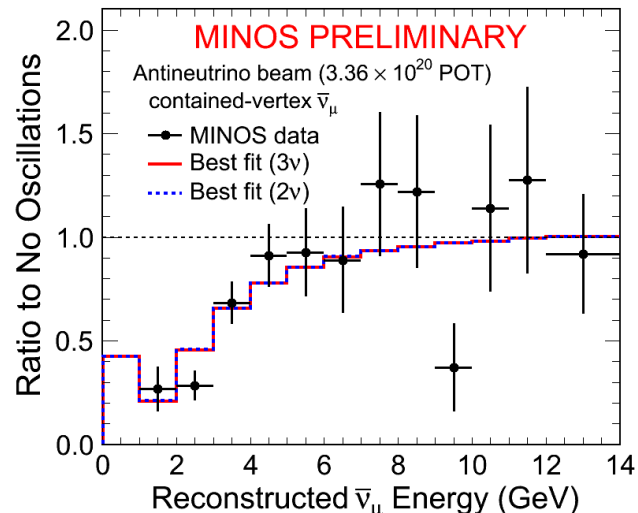
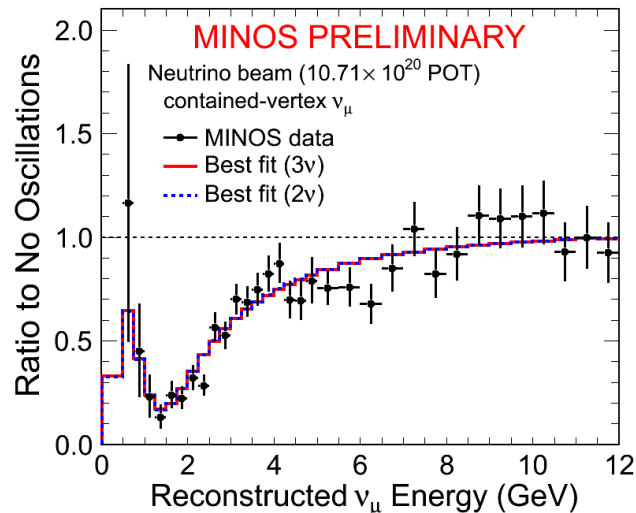
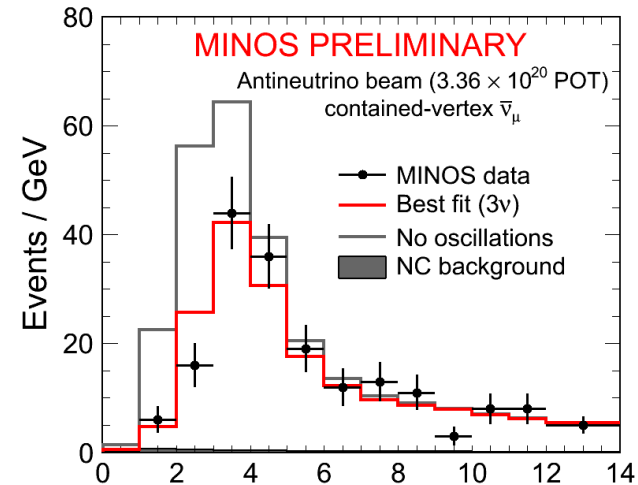
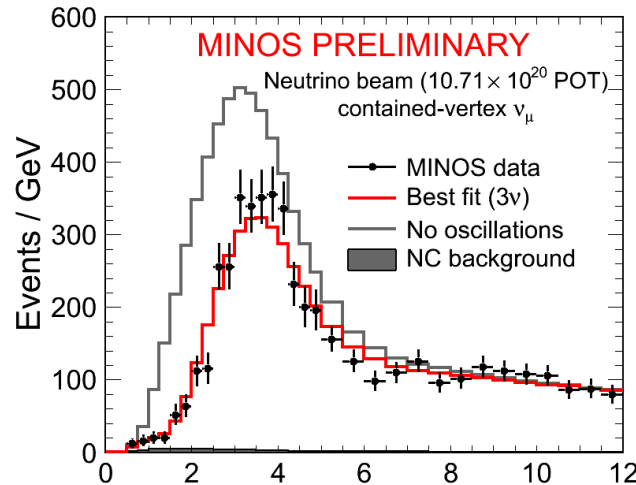
† Fogli et al. Phys. Rev. D 86, 013012 (2012)

# Fitted Far Detector Beam Samples

✦ Best fit shown for the  $\nu_\mu$  and  $\bar{\nu}_\mu$  samples.

✦ Two and three flavour fits almost indistinguishable.

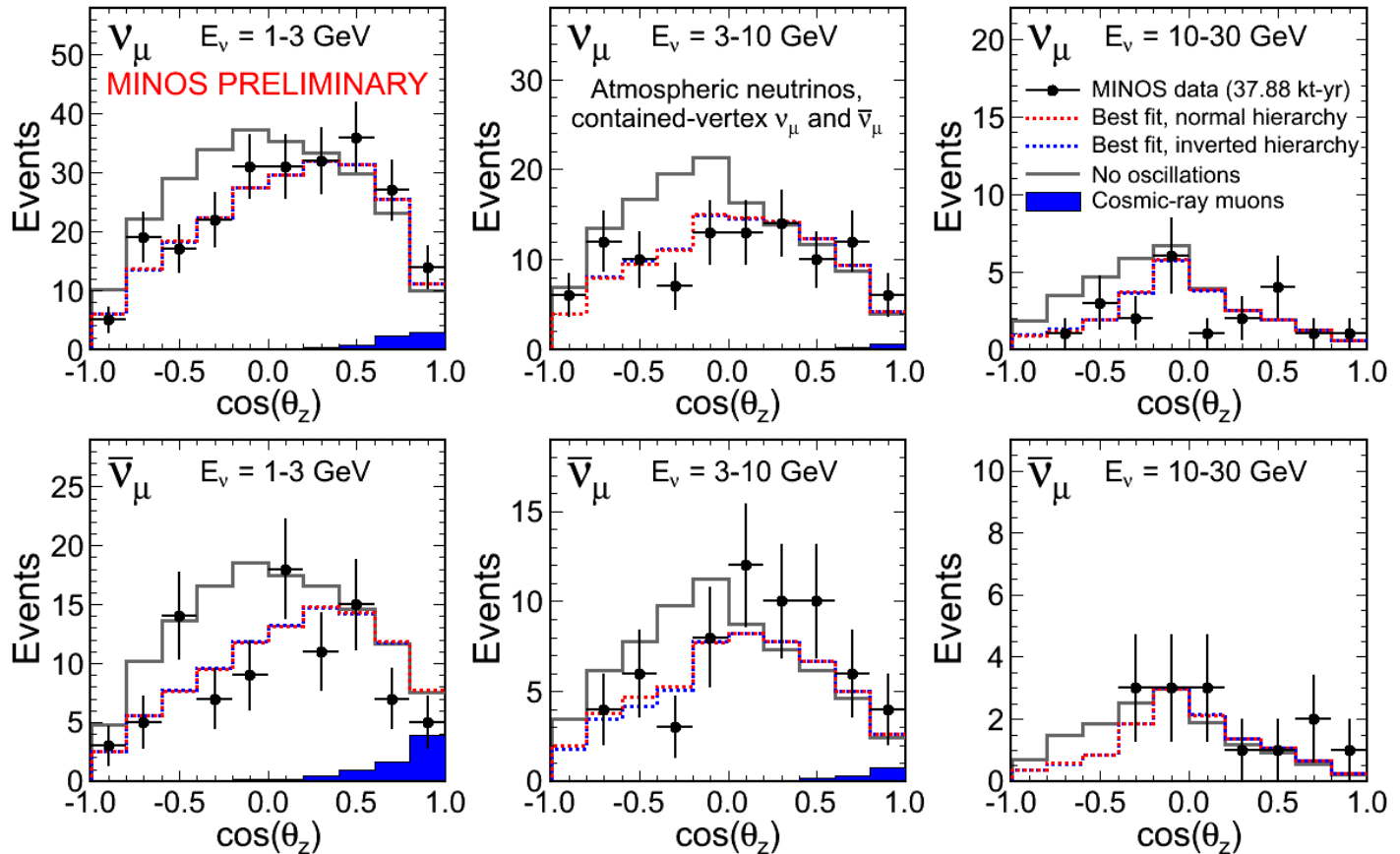
✦ Fit also includes  $\bar{\nu}_\mu$  in the neutrino beam and the non-fiducial muons.



# Fitted Far Detector Atmospheric Samples

✦ Best fit shown contained  $\nu_\mu$  and  $\bar{\nu}_\mu$  samples.

Neutrinos



Antineutrinos

✦ Fit also includes non-contained  $\mu^- / \mu^+$  and shower-like events.

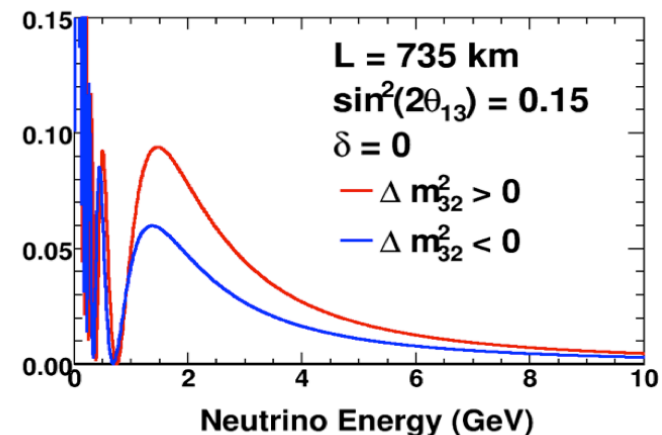
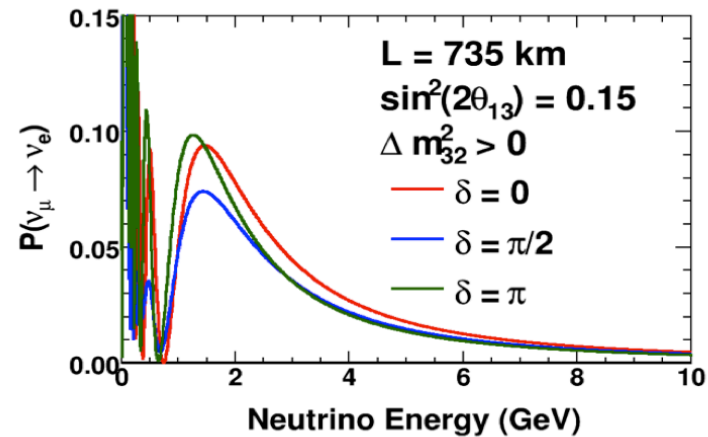


# Three Flavour Electron Neutrino Appearance

- ✦ In the three flavour fit, we can also include the  $\nu_e$  appearance
  - ✦ Helps to probe the three flavour oscillation effects
  - ✦ To first order, neglecting matter effects:

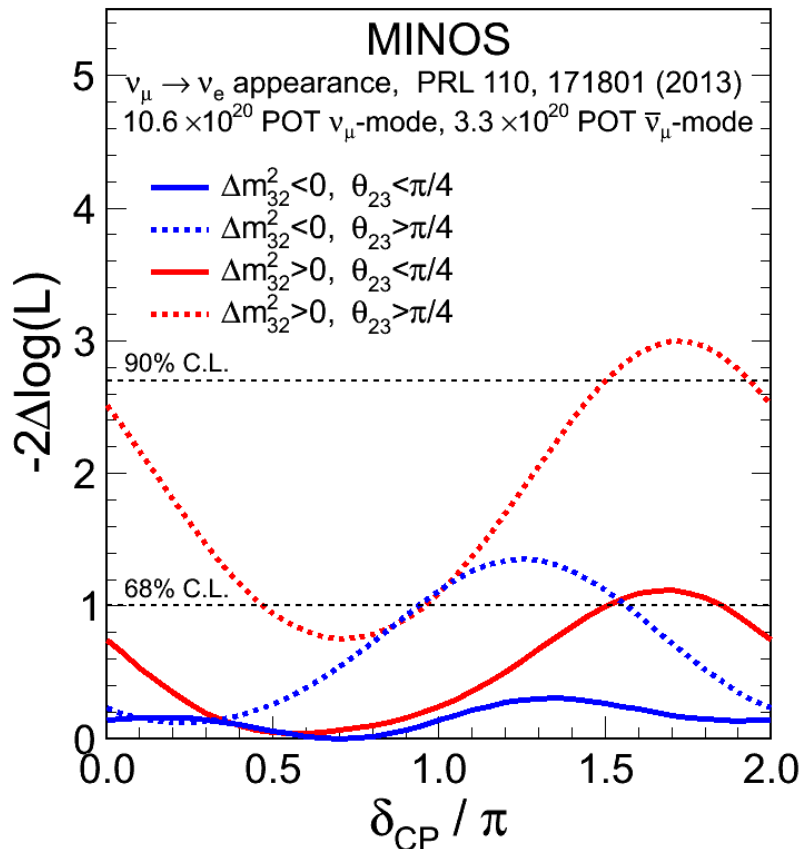
$$P(\nu_\mu \rightarrow \nu_e) = \sin^2 \theta_{23} \sin^2 2\theta_{13} \sin^2 \left( \frac{1.27 \Delta m_{31}^2 L}{E} \right)$$

- ✦ Sensitive to the value of  $\delta_{cp}$
- ✦ Matter effects play a large role
  - ✦ Modifies oscillation probability by  $\sim 30\%$
  - ✦ Term is dependent on the sign of the mass splitting
    - ✦ Provides potential to probe the mass hierarchy
- ✦ Sensitive to  $\theta_{23}$  octant

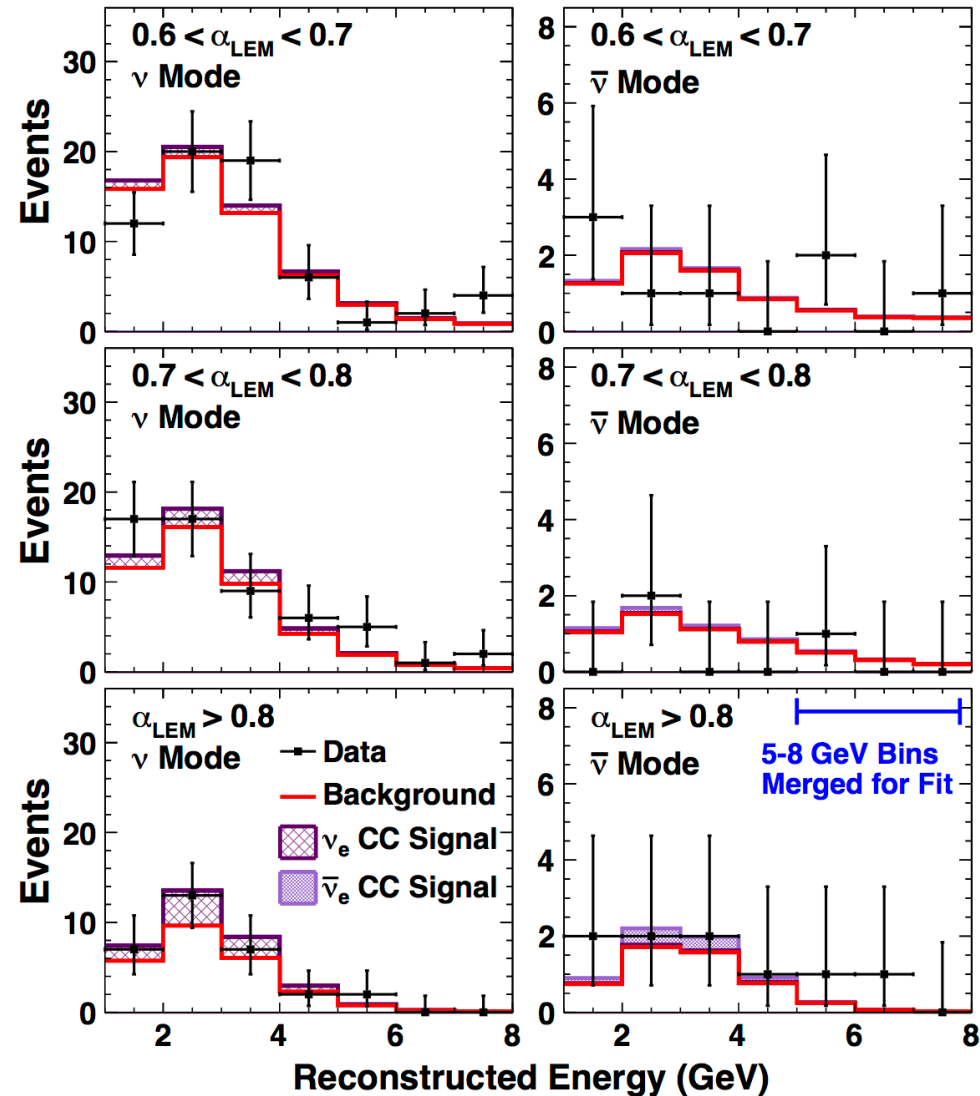


# Three Flavour Electron Neutrino Appearance

- Fit both  $\nu_e$  and  $\bar{\nu}_e$
- $\sin^2(2\theta_{13}) = 0.053 (0.094)$
- $\sin^2(2\theta_{13}) > 0 (96\%)$



## MINOS Far Detector Data



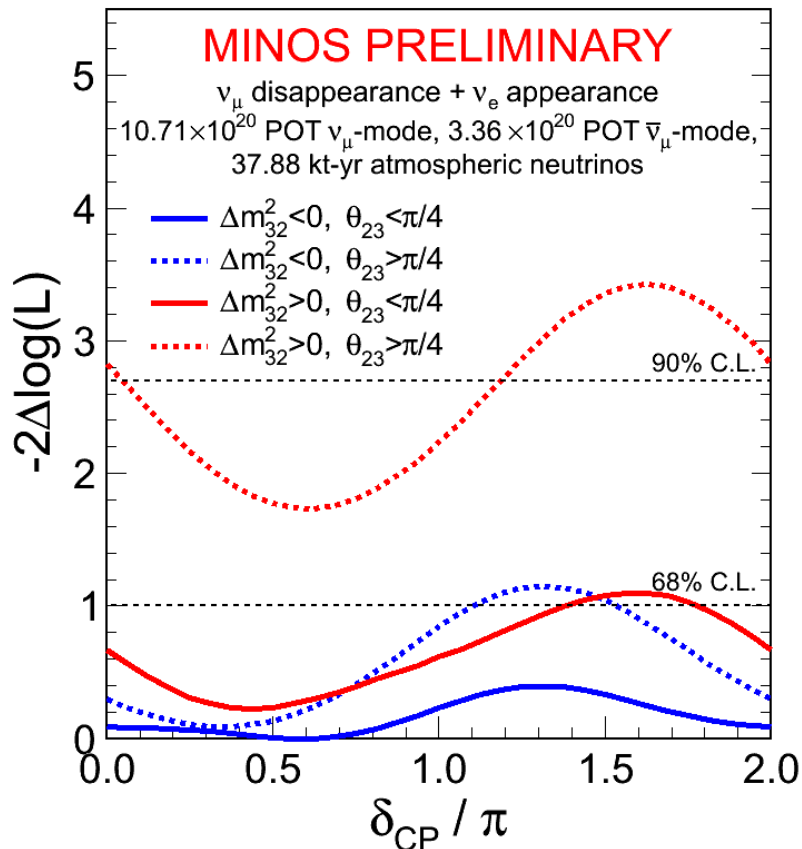
Phys. Rev. Lett. 110, 171801 (2013)

# Three Flavour Disappearance and Appearance

✦ Combine the information from both fits

✦ Each provides a 4D likelihood surface in  $\Delta m_{32}^2$ ,  $\sin^2\theta_{23}$ ,  $\sin^2\theta_{13}$  and  $\delta_{cp}$

✦ Systematics assumed to be uncorrelated

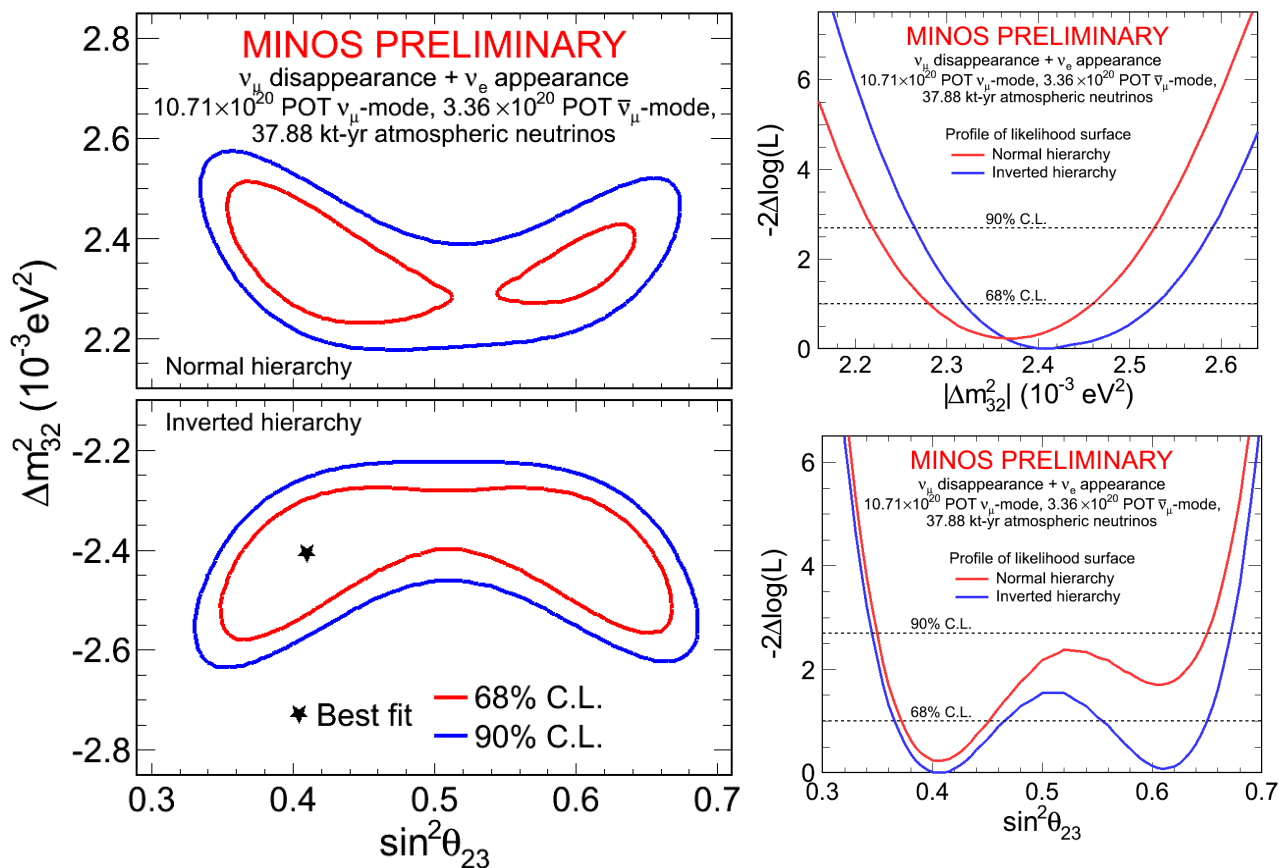


✦ Normal hierarchy, upper octant case is now further disfavoured

✦ At least 90% C.L. for half  $\delta_{cp}$  range



# Three Flavour Disappearance and Appearance



- ✦ Slight preference for the inverted hierarchy
- ✦ Normal hierarchy, higher octant disfavoured

# Three Flavour Disappearance and Appearance

✦ The four local best fit points:

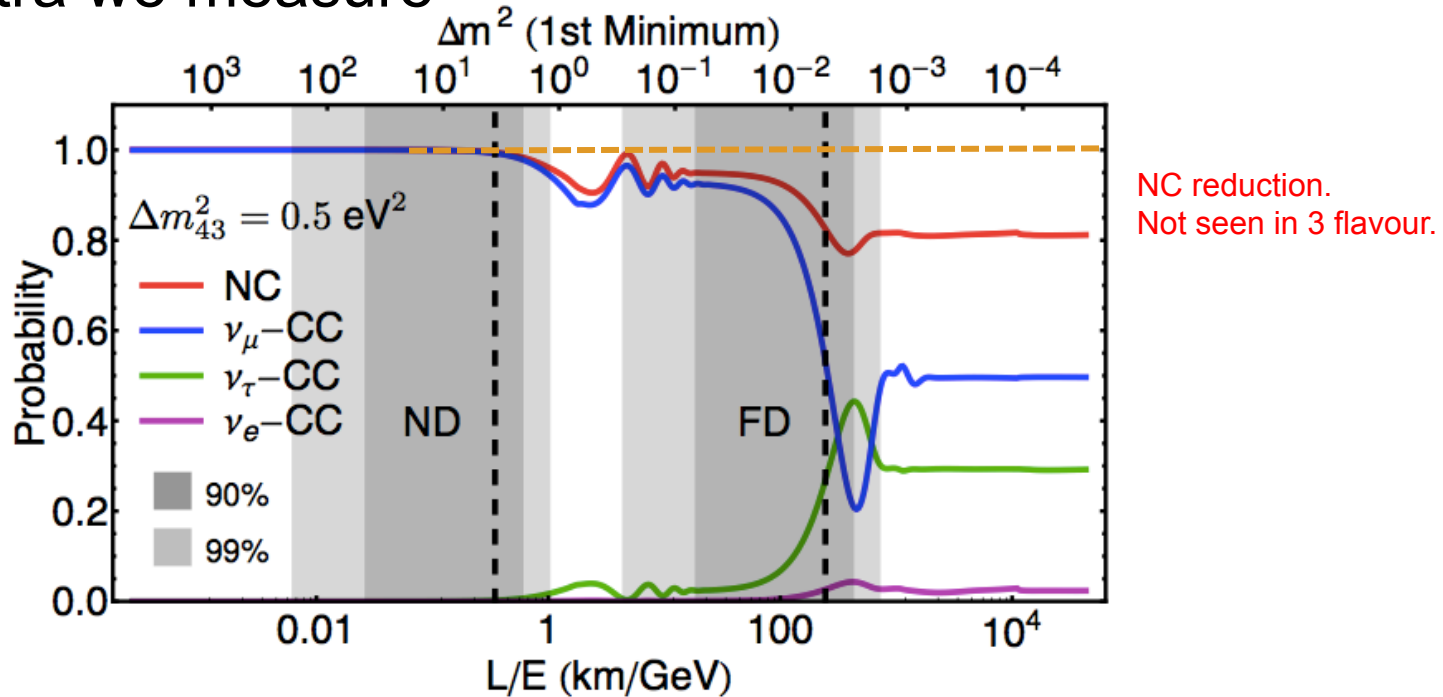
Hierarchy	Octant	Best fit parameters				$-2\Delta\log(L)$
		$\Delta m_{32}^2 / 10^{-3} \text{eV}^2$	$\sin^2\theta_{23}$	$\sin^2\theta_{13}$	$\delta_{\text{CP}} / \pi$	
Normal	Lower	2.37	0.41	0.0242	0.44	0.23
Normal	Higher	2.35	0.61	0.0238	0.62	1.74
Inverted	Lower	-2.41	0.41	0.0243	0.62	-
Inverted	Higher	-2.41	0.61	0.0241	0.37	0.09

Hierarchy	$\Delta m_{32}^2 / 10^{-3} \text{eV}^2$	$\sin^2\theta_{23}$ (90% C.L.)
Normal	$2.37^{+0.09}_{-0.09}$	$0.35 < \sin^2\theta_{23} < 0.65$
Inverted	$-2.41^{+0.12}_{-0.09}$	$0.34 < \sin^2\theta_{23} < 0.67$

✦ Prefer non-maximal mixing at 79% C.L.

# Four Flavour Oscillations

- ✦ Oscillations to a fourth sterile neutrino will change the energy spectra we measure



## ✦ 3 Regimes:

- ✦ Small  $\Delta m^2_{43}$ : Wide oscillations at the FD. Coming soon!
- ✦ Medium  $\Delta m^2_{43}$ : Rapid oscillations that average out at the FD. Done!
- ✦ Large  $\Delta m^2_{43}$ : Oscillations in the ND. Coming soon!

# Neutral Current

- ✦ All neutrino flavours undergo neutral current (NC) interactions
  - ✦ Any deficit in the number of NC events can probe sterile neutrinos

- ✦ Combined fit of the CC and NC spectra

- ✦ No evidence found for sterile neutrinos with  $\Delta m_{43}^2 = 0.5 eV^2$

Predicted background

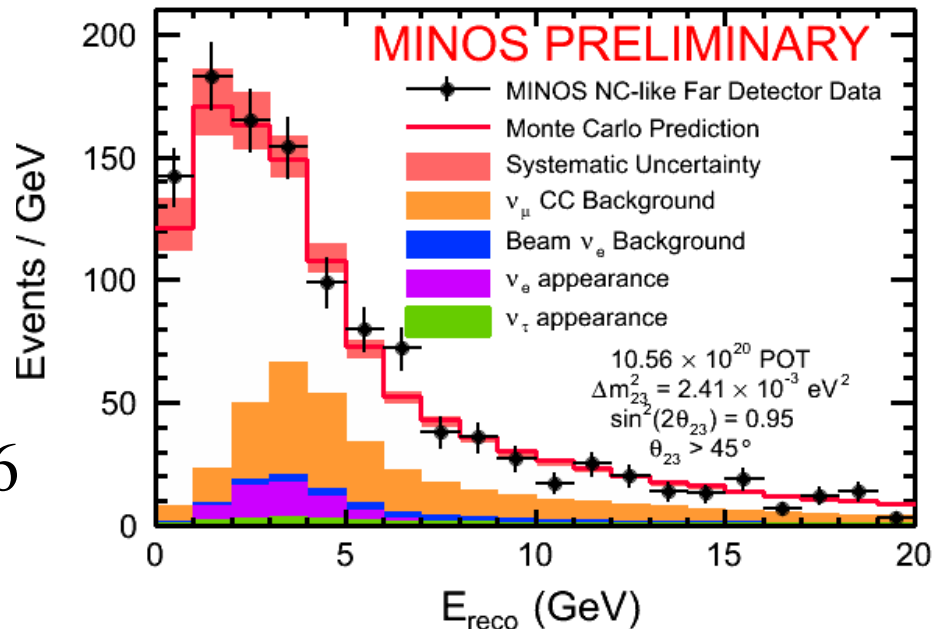
✦ Define  $R = \frac{N_{data} - N_{bkg}}{S_{NC}}$

Predicted signal  $S_{NC}$

✦ 0 – 200 GeV:  $R = 1.049 \pm 0.076$

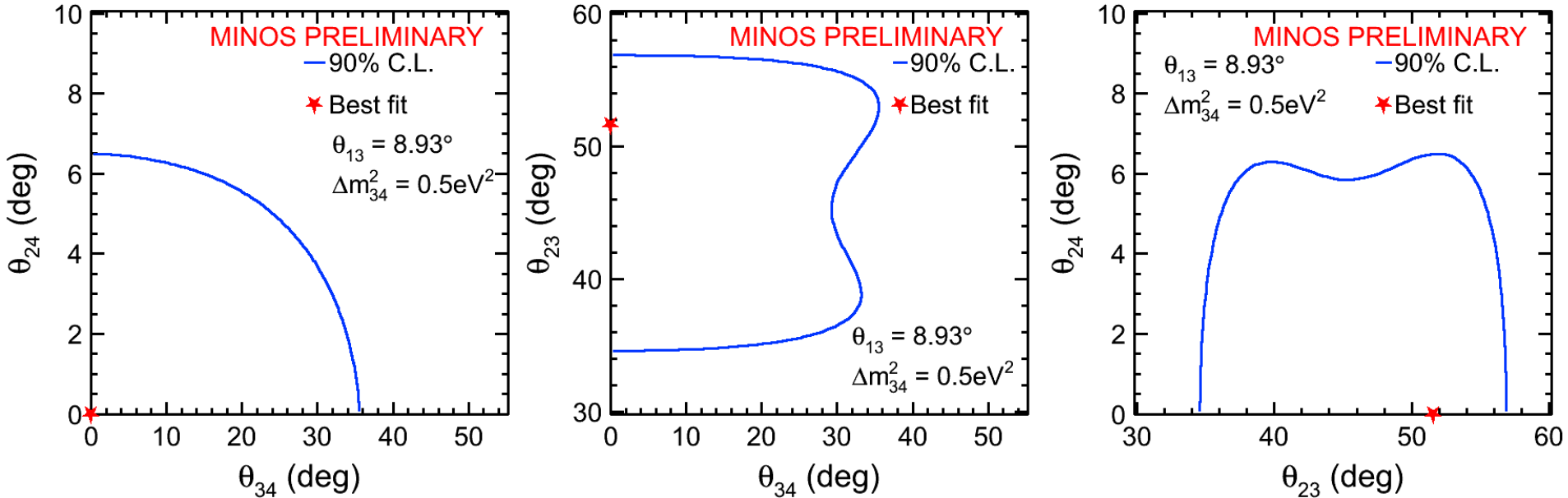
✦ 0 – 3 GeV:  $R = 1.093 \pm 0.097$

✦ 3 – 200 GeV:  $R = 1.009 \pm 0.095$



# Neutral Current – Mixing Angle Limits

✦ We can now place limits on the values of the mixing angles



✦ 90% C.L:  $\theta_{24} < 5^\circ$

✦ 90% C.L:  $\theta_{34} < 24^\circ$

✦ Very slight preference to upper octant of  $\theta_{23}$

# Neutral Current – Combined with BUGEY

✦ To compare to short-baseline appearance experiments, combine with BUGEY:

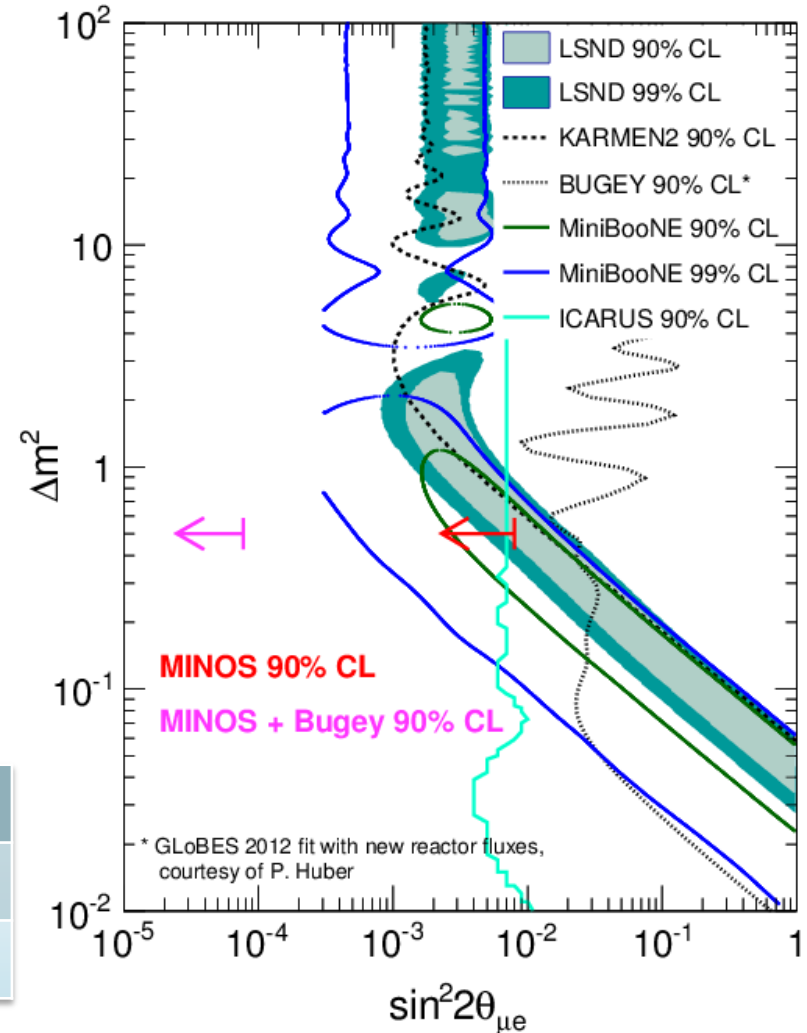
- ✦  $\sin^2 2\theta_{\mu e} \sim \sin^2 2\theta_{14} \sin^2 \theta_{24}$
- ✦ MINOS sensitive to  $\theta_{24}$  and  $\theta_{34}$
- ✦ BUGEY sensitive to  $\theta_{14}$

✦ Coming soon:

- ✦ Full contour over range of  $\Delta m^2$

✦ At  $\Delta m_{43}^2 = 0.5 eV^2$

Sample	$\sin^2 2\theta_{\mu e}$ (90% C.L.)
MINOS	$< 7.1 \times 10^{-3}$
MINOS + BUGEY	$< 7.7 \times 10^{-5}$

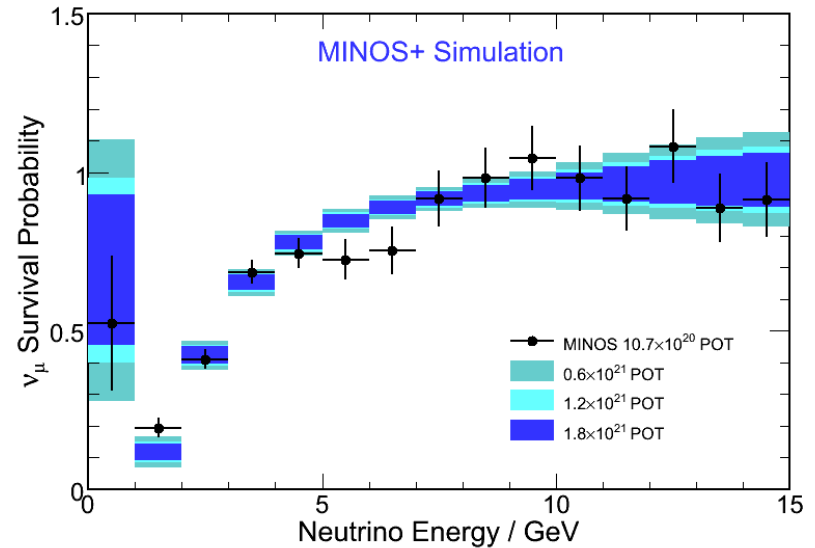
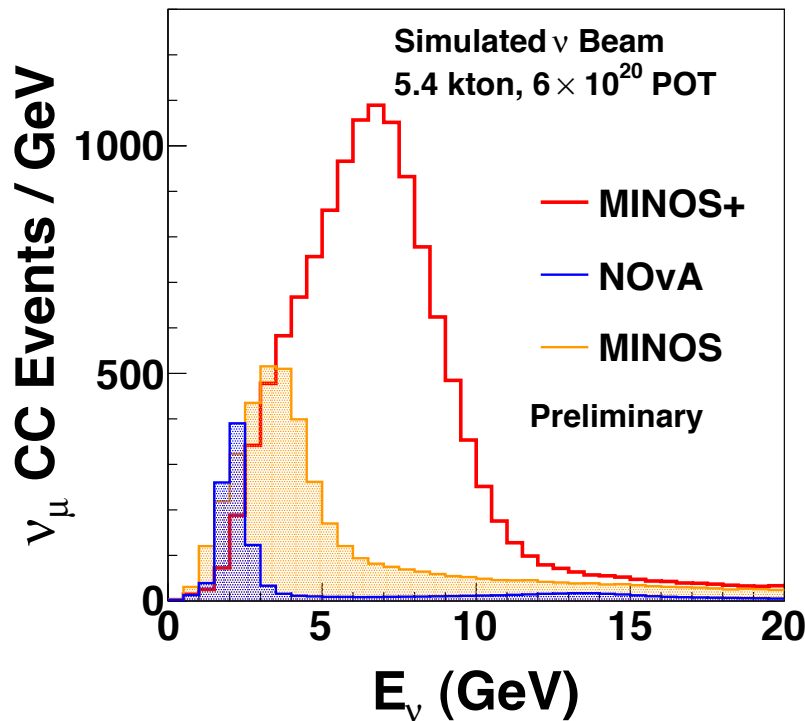




# MINOS+

✦ MINOS+ is a continuation of the MINOS experiment into the NOvA era.

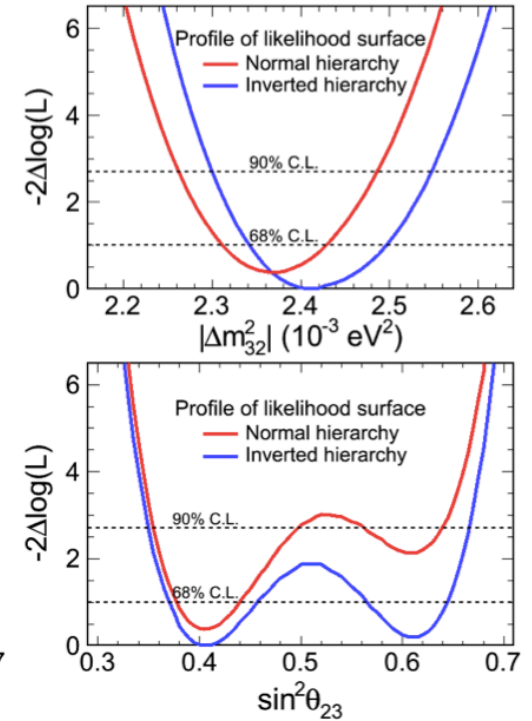
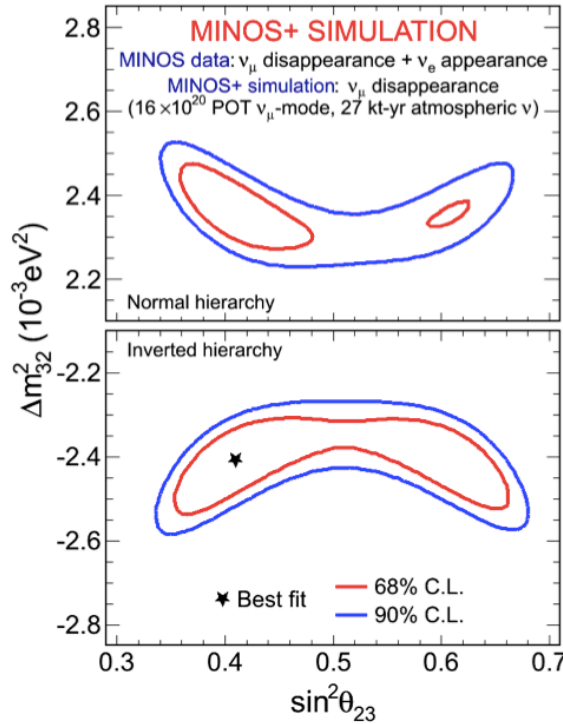
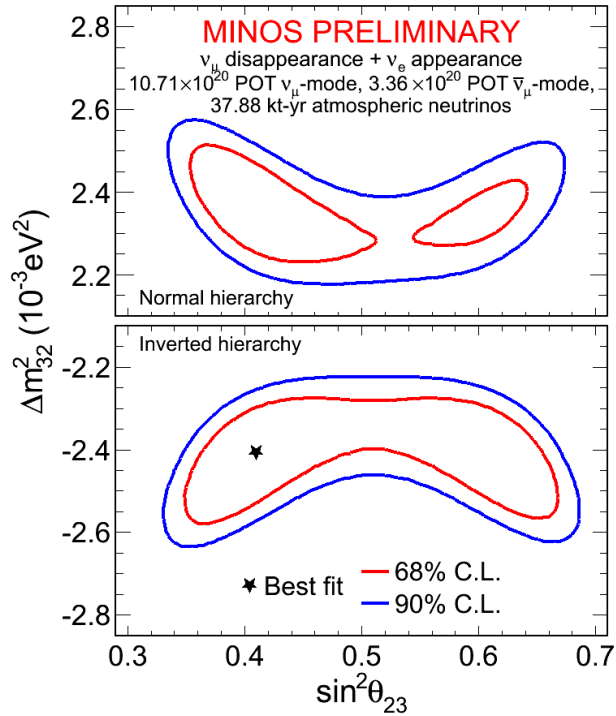
- ✦ Beam now runs in ME mode.
- ✦ ME spectrum peaks above oscillation dip.



- ✦ Expect  $\sim 4000$  events per year.
- ✦ Precision test of the oscillation hypothesis in the tail.
- ✦ Search for sterile neutrinos (using both CC and NC), extra dimensions etc.

# MINOS+ Three Flavour Sensitivity

## ✦ Sensitivity of MINOS & MINOS+



MINOS combined disappearance and appearance.

MINOS combined disappearance and appearance plus MINOS+ disappearance

# MINOS+: Up and running!

✦ As of the start of September

✦ NuMI running well at ~350kW

✦ We have FD events!

✦ NuMI expected to run at ~700kW once all upgrades are completed

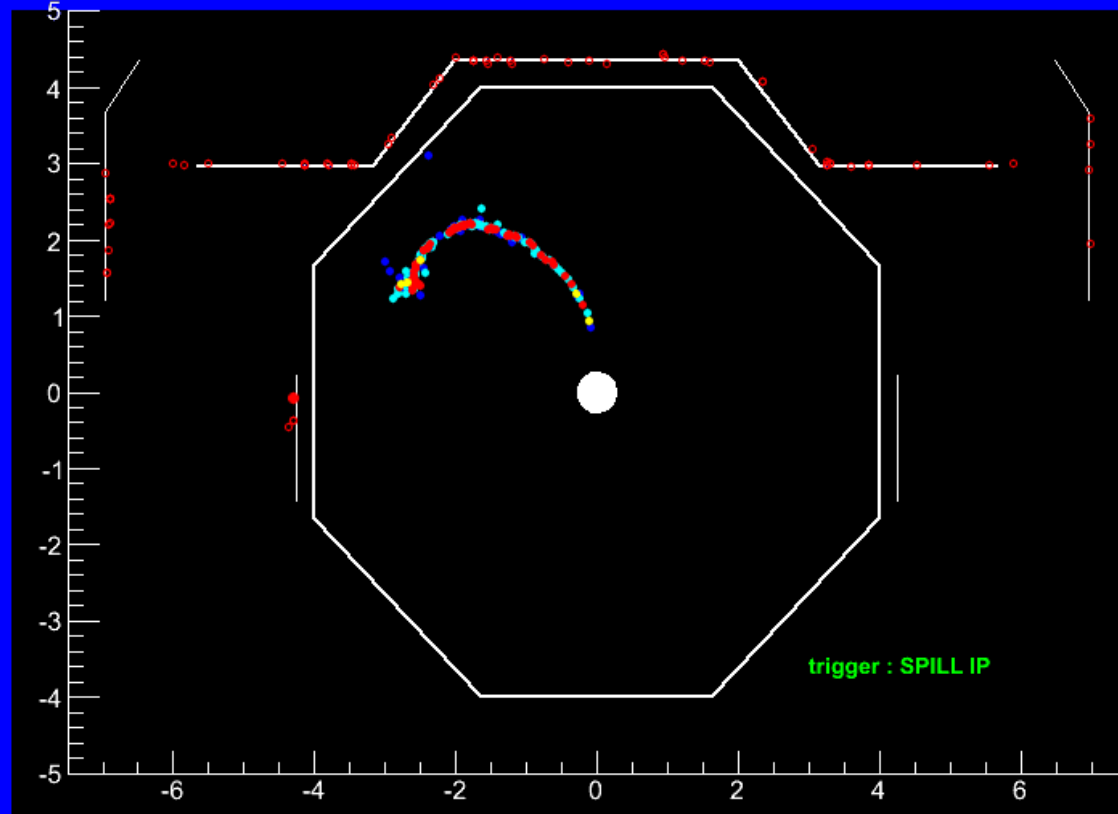
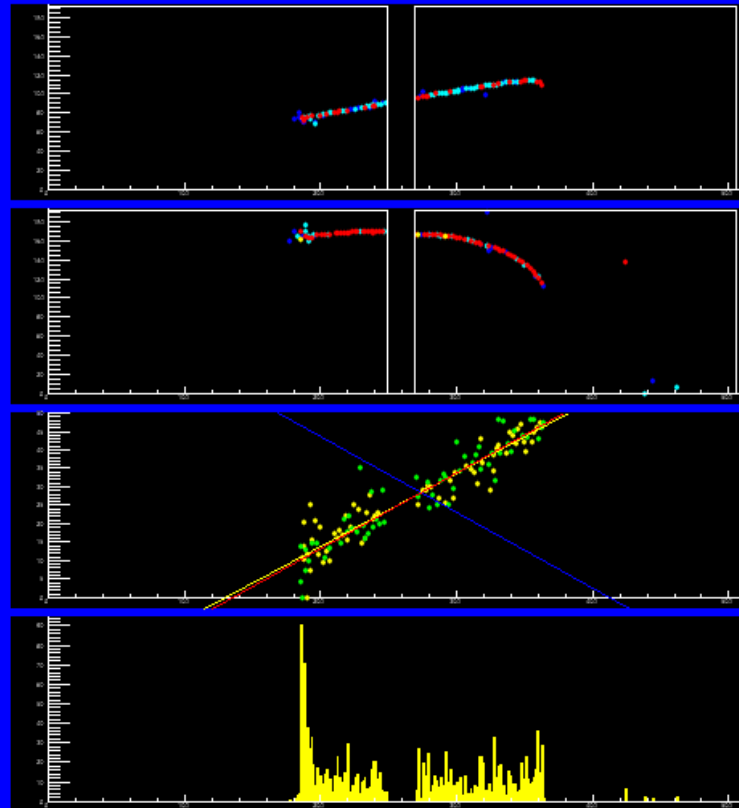
Date : 14 Sep 2013

Time : 07:04:33

Run : 52780\_15

Snarl : 151711

EventType : Golden Beam Neutrino



# Conclusions

## ✦ The final word from MINOS:

- ✦ First combined disappearance and appearance result from a long-baseline experiment

Hierarchy	$\Delta m_{32}^2 / 10^{-3} \text{ eV}^2$	$\sin^2\theta_{23}$ (90% C.L.)
Normal	$2.37^{+0.09}_{-0.09}$	$0.35 < \sin^2\theta_{23} < 0.65$
Inverted	$-2.41^{+0.12}_{-0.09}$	$0.34 < \sin^2\theta_{23} < 0.67$

- ✦ See no tension between neutrino and antineutrino oscillation parameters
- ✦ No evidence of sterile neutrinos at  $\Delta m_{43}^2 = 0.5 \text{ eV}^2$
- ✦ Look out for updates to the sterile neutrino analysis

## ✦ MINOS+ will continue where MINOS left off, providing exciting physics results in the coming years



# Thank You





A high-angle photograph of a large industrial facility, likely a particle accelerator. The central focus is a long, narrow, metallic structure with a dark, textured interior. The number "485" is visible on the right side of this structure. The facility is surrounded by complex machinery, including blue and yellow overhead beams, orange safety railings, and a network of pipes and cables. A person in a white shirt is visible on a lower level, providing a sense of scale. The overall scene is brightly lit, with a mix of industrial colors and metallic surfaces.

# Backup Slides





# Neutrino Oscillations

- ✦ The neutrino **weak flavour states** are not the same as the neutrino **mass states**.
  - ✦ Creation and detection governed by flavour states
  - ✦ Propagation governed by mass states

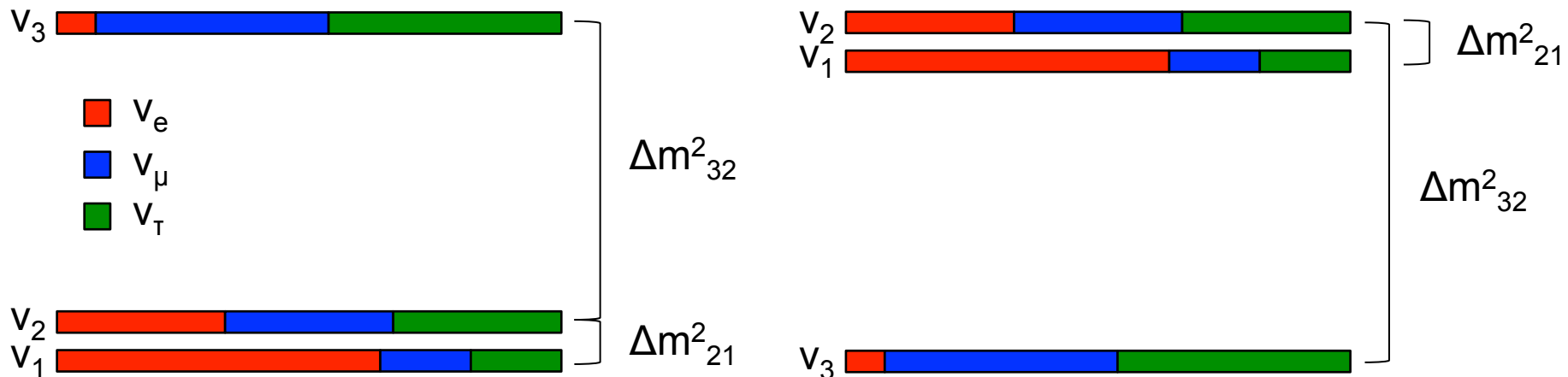
$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = U_{PMNS}^\dagger \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix} \quad P(\nu_\alpha \rightarrow \nu_\beta) = \left| \sum_j U_{\beta j}^* e^{-i\frac{m_j^2 L}{2E}} U_{\alpha j} \right|^2$$

- ✦ Matrix U governs the oscillation of a neutrino of one flavour into a different flavour.

# The PMNS Matrix and Mass Hierarchy

✦ Can be written as a product of three matrices

$$U_{PMNS} = \underbrace{\begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix}}_{\text{Atmospheric and LBL beam experiments}} \underbrace{\begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta_{CP}} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta_{CP}} & 0 & c_{13} \end{pmatrix}}_{\text{SBL reactor and LBL beam experiments}} \underbrace{\begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}}_{\text{LBL reactor and solar experiments}}$$

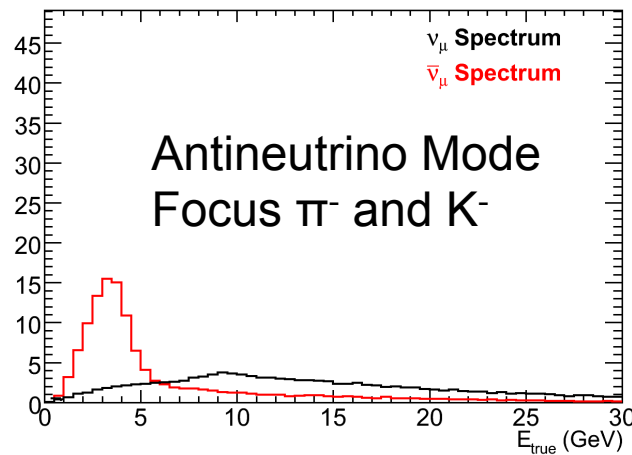
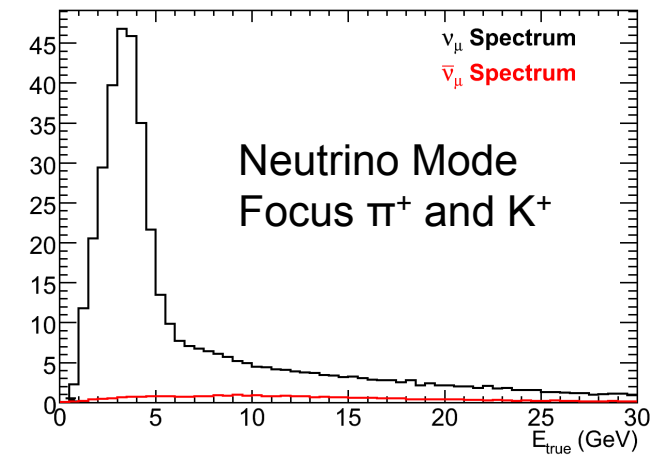
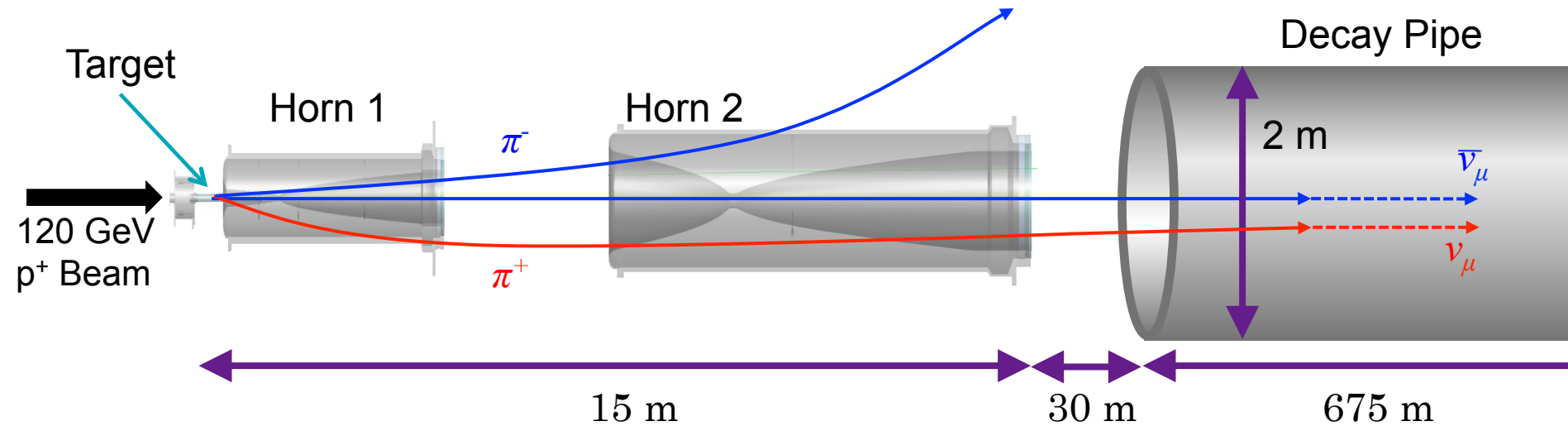


# Open Questions

- ✦ There are still lots of unknowns in neutrino physics!
  - ✦ Is  $\theta_{23} = 45^\circ$ ?
    - ✦ If not, is it higher or lower octant?
  - ✦ Is the mass hierarchy normal or inverted?
  - ✦ What is the absolute scale of neutrino mass?
  - ✦ Are neutrinos Dirac or Majorana?
    - ✦ Majorana neutrinos would open the door to the seesaw mechanism
  - ✦ Do neutrinos violate CP?
    - ✦ Is it enough to explain the matter – antimatter asymmetry of the universe?
  - ✦ Why does the PMNS matrix have the form it does?
    - ✦ Is there an underlying symmetry?
  - ✦ Why is the PMNS matrix different to the CKM matrix?
  - ✦ Are there any sterile neutrinos?

# The NuMI Beam

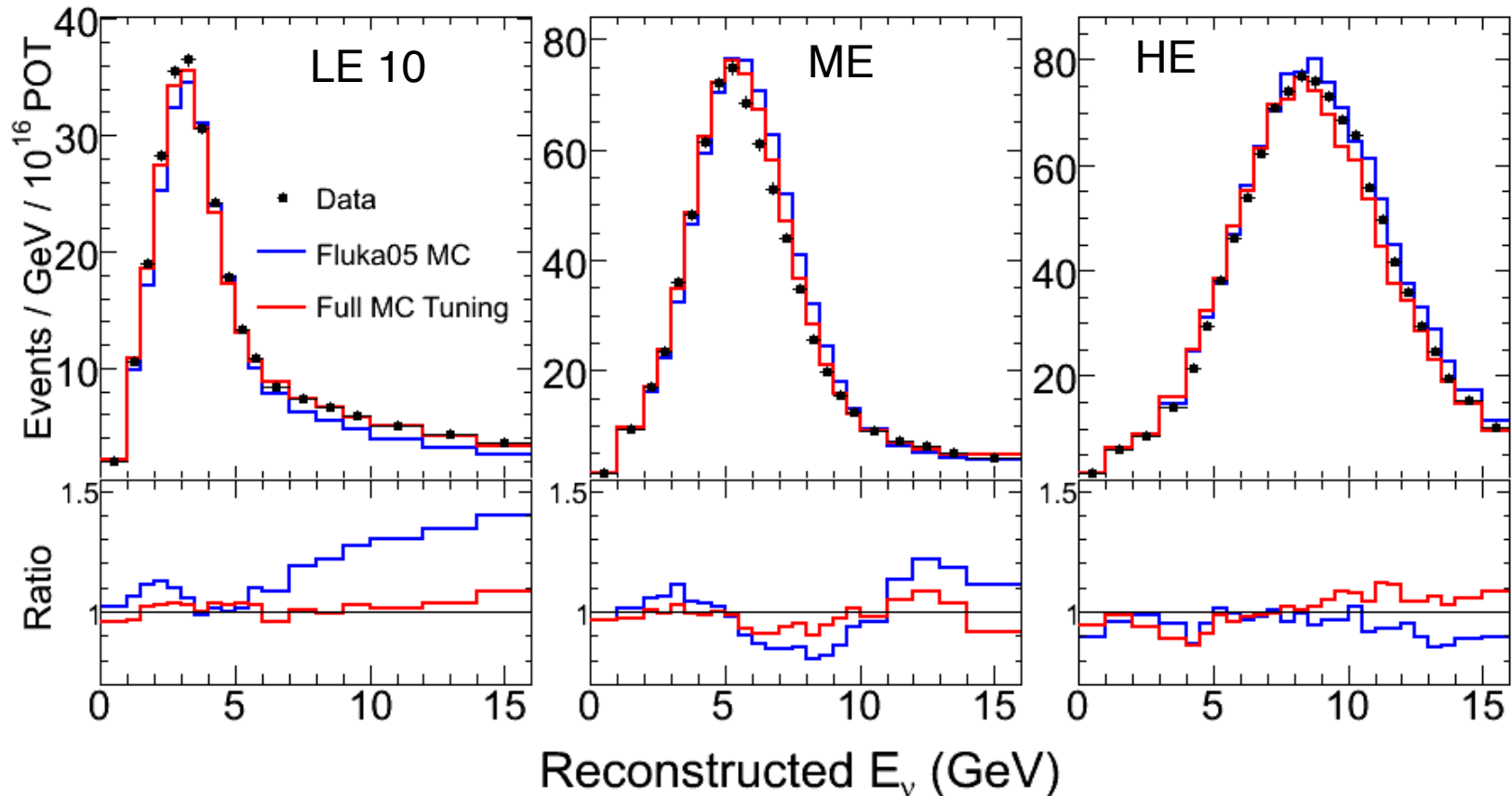
- Beam starts with 120 GeV protons from the Main Injector.



- Wrong sign high energy mesons give “wrong sign” neutrinos

# Beam Simulation

- ✦ The beam simulation is tuned using data from the ND
  - ✦ Take advantage of the different run types to constrain the simulation



# Detector Technology

- ✦ Detectors built from alternating planes:

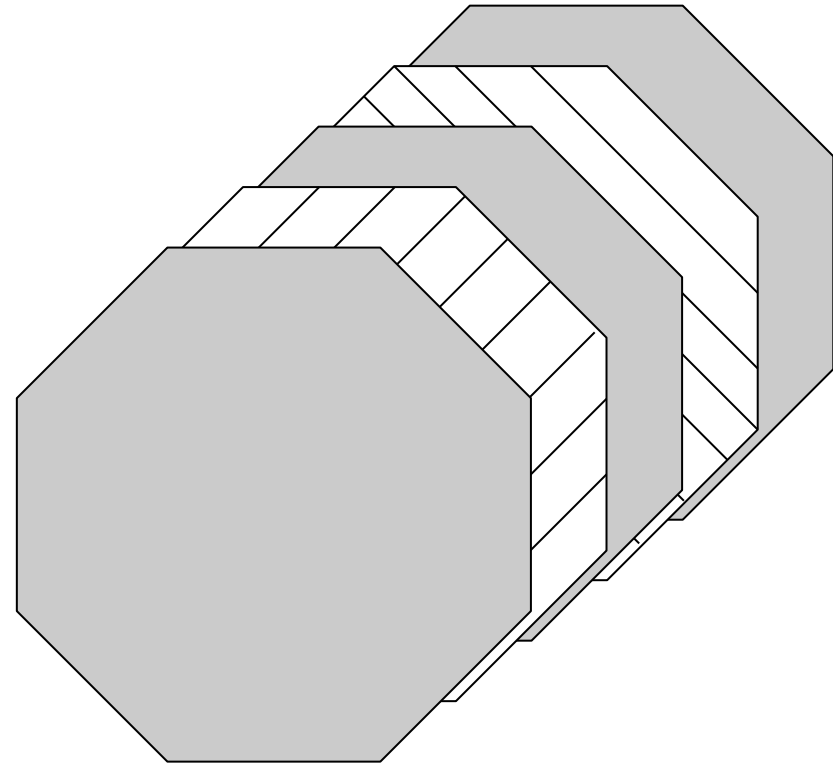
- ✦ 2.54cm steel absorber  $\sim 1.4 X_0$
- ✦ 1cm thick scintillator.

- ✦ Scintillator planes:

- ✦ Made from plastic scintillator bars, each 4.1cm wide.
- ✦ Read out by multi-anode PMTs via WLS fibres.
- ✦ Alternating layers have bars in orthogonal directions views, U and V

- ✦ Magnetic field allows for charge separation.

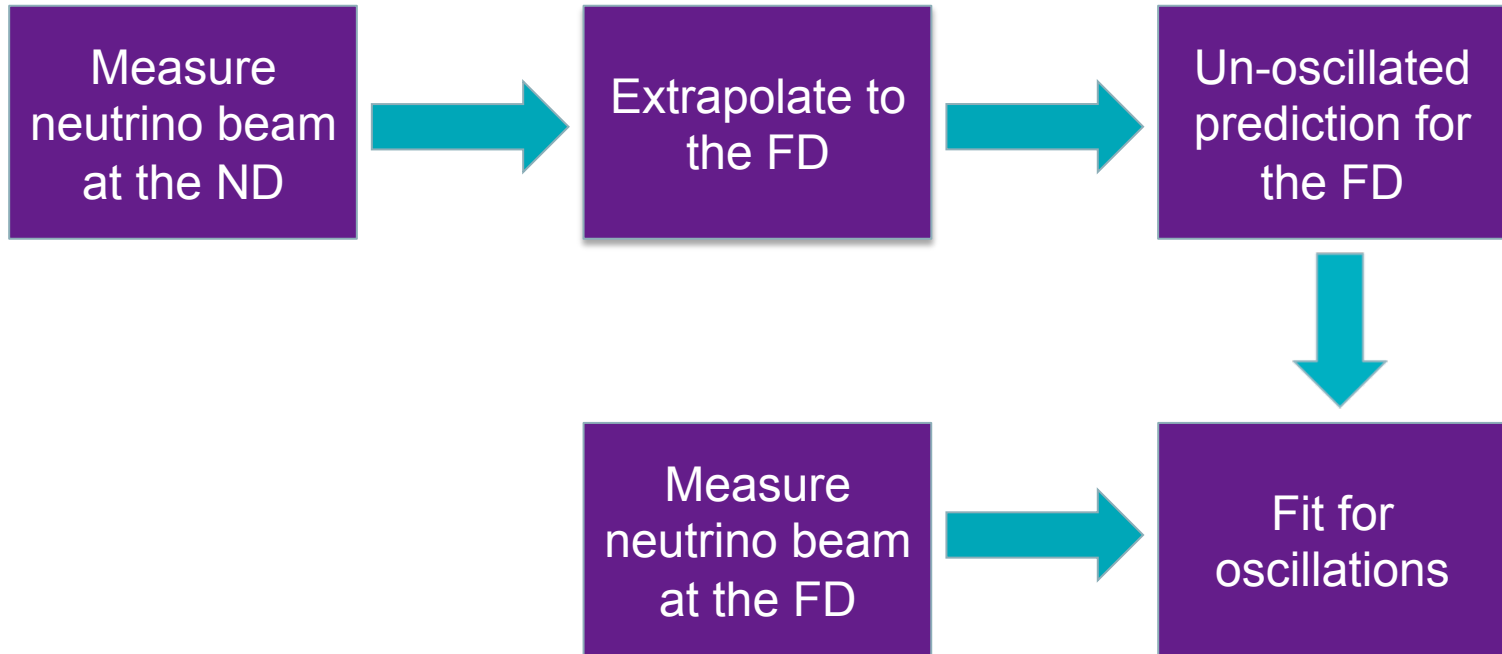
- ✦ Both detectors have average field of 1.3T





# Measurement Strategy: Overview

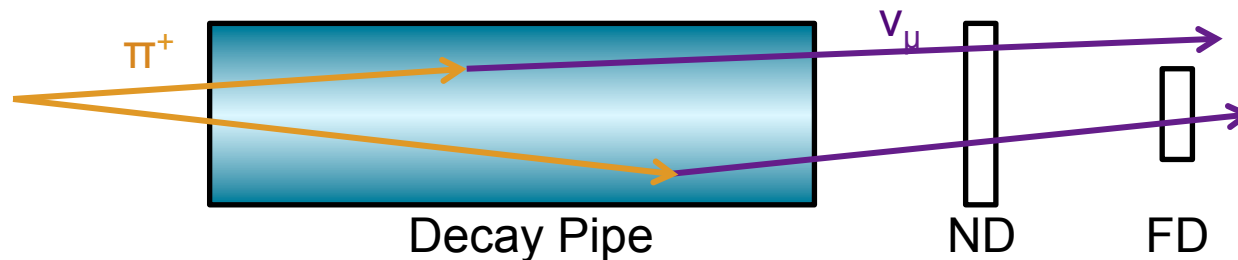
- ✦ Use the ND to predict the FD un-oscillated spectrum



- ✦ The extrapolation to the FD requires a few steps...

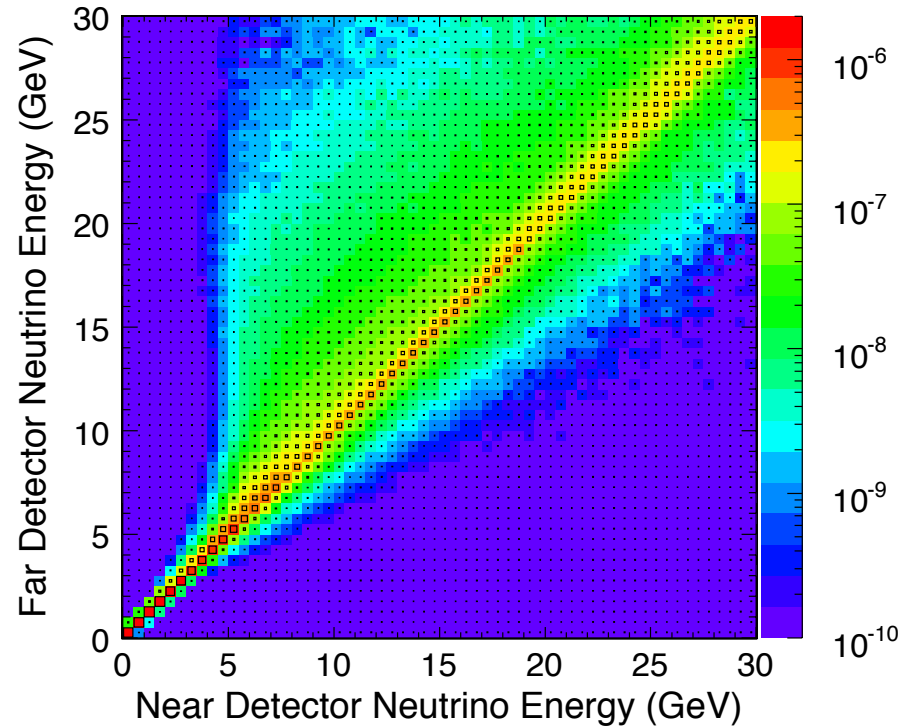
# Measurement Strategy: Extrapolation I

- ✦ Starting with the ND data:
  - ✦ Correct for ND purity and efficiency and apply reco-true matrix
  - ✦ Account for cross-sections, PoT and ND mass
- ✦ Need to account for beam differences now
  - ✦ The energy spectrum differs between the two detectors
    - ✦ Different angular acceptances
    - ✦ Low energy pions decay upstream in the decay pipe.
  - ✦ FD sees a point source
  - ✦ ND sees an extended source



# Measurement Strategy: Extrapolation II

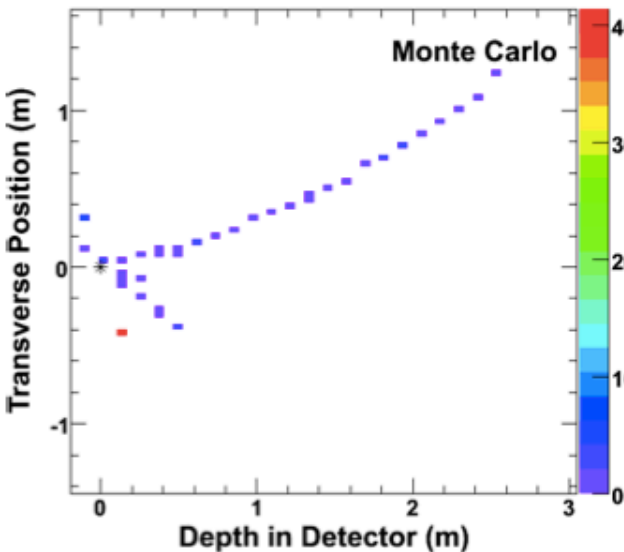
- ✦ Apply the beam matrix to extrapolate to the FD.



- ✦ Then apply the FD specific corrections.
  - ✦ These are the analogues to those shown previously for the ND
- ✦ Provides the un-oscillated prediction at the FD

# Event Topologies in MINOS

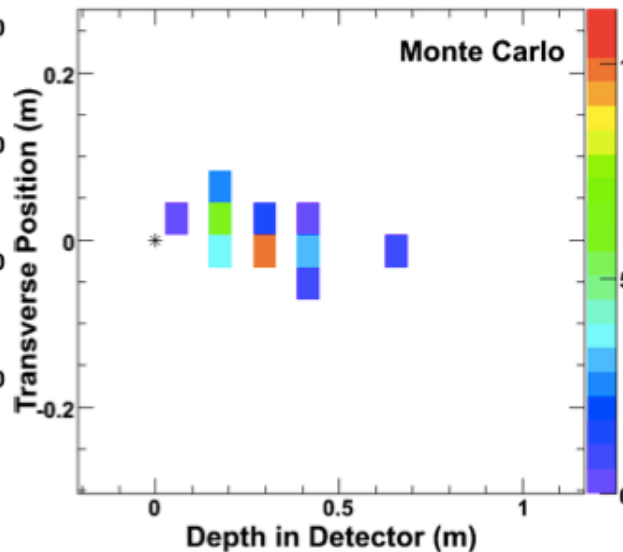
✦ Expect to see three classes of event.



CC  $\nu_\mu$

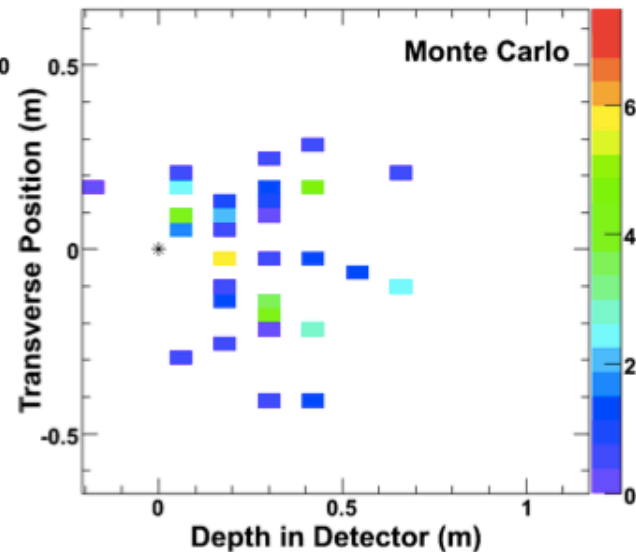
Identify muon track and use curvature to measure the sign.

Momentum comes from range or curvature (if not contained).



CC  $\nu_e$

Compact electromagnetic shower.



NC  $\nu_x$

Disperse hadronic shower energy deposits.

# Three Flavour Muon Neutrino Disappearance

$$P(\nu_\mu \rightarrow \nu_\mu) = 1 - \sin^2 2\theta_{23} \cos^2 \theta_{13} \sin^2 \left( \frac{1.27 \Delta m_{\mu\mu}^2 L}{E} \right) + \mathcal{O} \left( \Delta m_\odot^2 \frac{L}{E} \right)^2$$

✦ Disappearance depends on:

$$\Delta m_{\mu\mu}^2 = \Delta m_{32}^2 + \frac{1 - |U_{\mu 1}|^2}{|U_{\mu 3}|^2} \Delta m_{21}^2$$

- ✦ Solar oscillation parameters
- ✦  $\theta_{13}$
- ✦ Mass hierarchy
- ✦  $\theta_{23}$  octant (very weakly)

✦ Matter effects important for multi-GeV up-going atmospheric

✦ Make a 4D fit to the FD data in  $(\Delta m_{32}^2, \sin^2 \theta_{23}, \sin^2 \theta_{13}, \delta_{CP})$

✦ Constrain  $\theta_{13}$  from the reactor experiment average:

$$\sin^2 \theta_{13} = 0.0242 \pm 0.0025$$

✦ Solar parameters fixed at the global average<sup>†</sup>

$$\Delta m_{21}^2 = 7.54 \times 10^{-5} eV^2$$

$$\sin^2 \theta_{12} = 0.307$$

<sup>†</sup> Fogli et al. Phys. Rev. D 86, 013012 (2012)

# Disappearance Analysis Event Counts

✦ Predicted numbers with oscillations made assuming:

$$\Delta m^2 = 2.41 \times 10^{-3} eV^2$$

$$\sin^2 2\theta = 0.95$$

Sample	No Osc	Osc	Measured
$\nu_\mu$ from $\nu_\mu$ beam	3201	2543	2579
$\bar{\nu}_\mu$ from $\nu_\mu$ beam	363	324	312
Non-fiducial $\mu^-$ from $\nu_\mu$ beam	3197	2862	2911
$\bar{\nu}_\mu$ from $\bar{\nu}_\mu$ beam	313	227	226
Atm. contained $\nu_\mu + \bar{\nu}_\mu$	1100	881	905
Atm. Non-fiducial $\mu^- + \mu^+$	570	467	466
Atm. showers	727	724	701



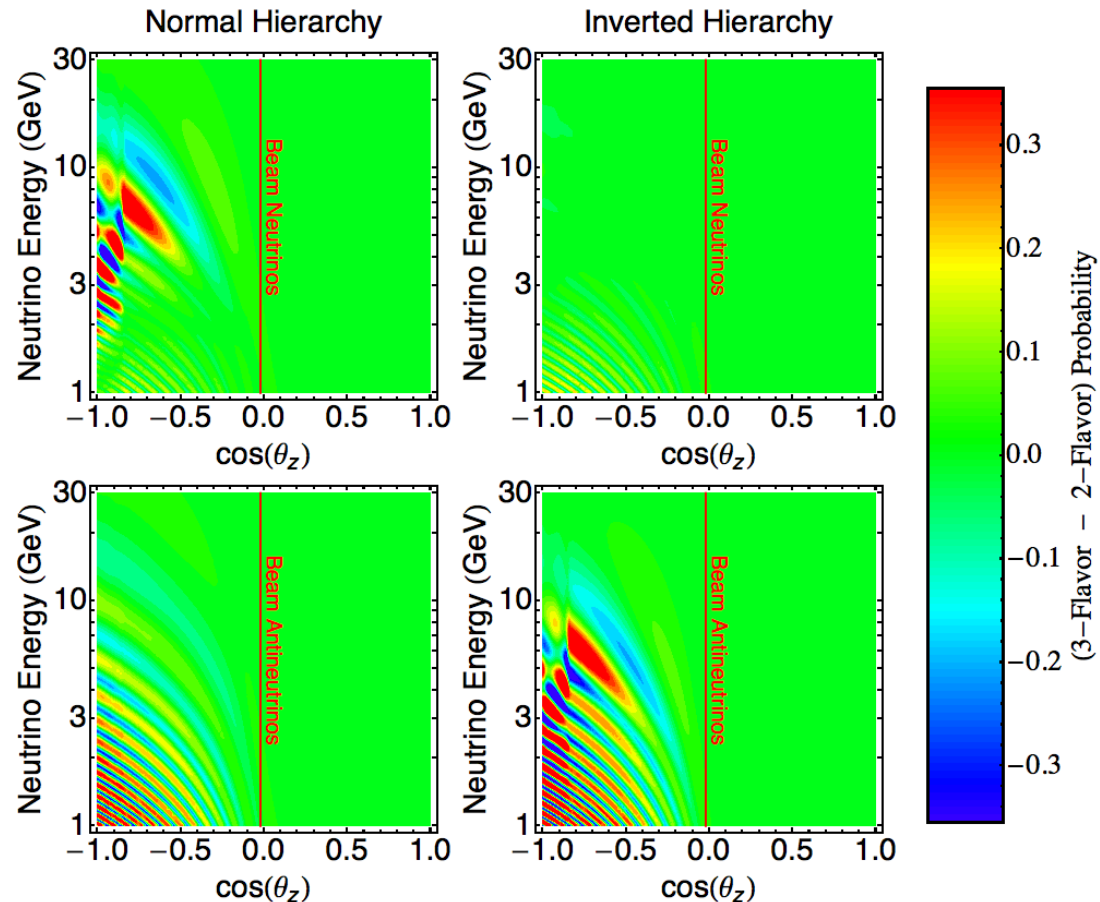
# Matter Effects (MSW)

- ✦ Interactions with matter modify the standard oscillations
  - ✦ Comparing between 2 and 3 flavour oscillations, see fairly large variation for  $\sim$ few GeV atmospheric neutrinos

- ✦ Changes the probability by up to 30%

- ✦ Gives sensitivity to the mass hierarchy

- ✦ Very small effect for the beam neutrinos



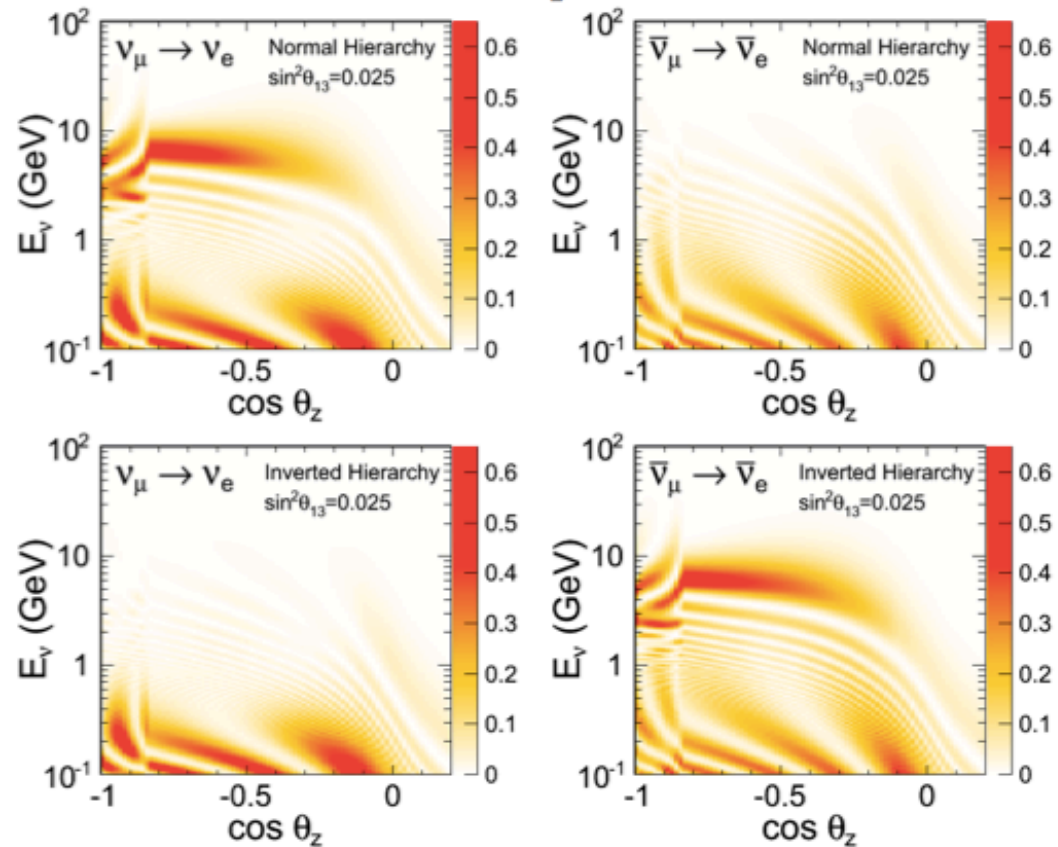
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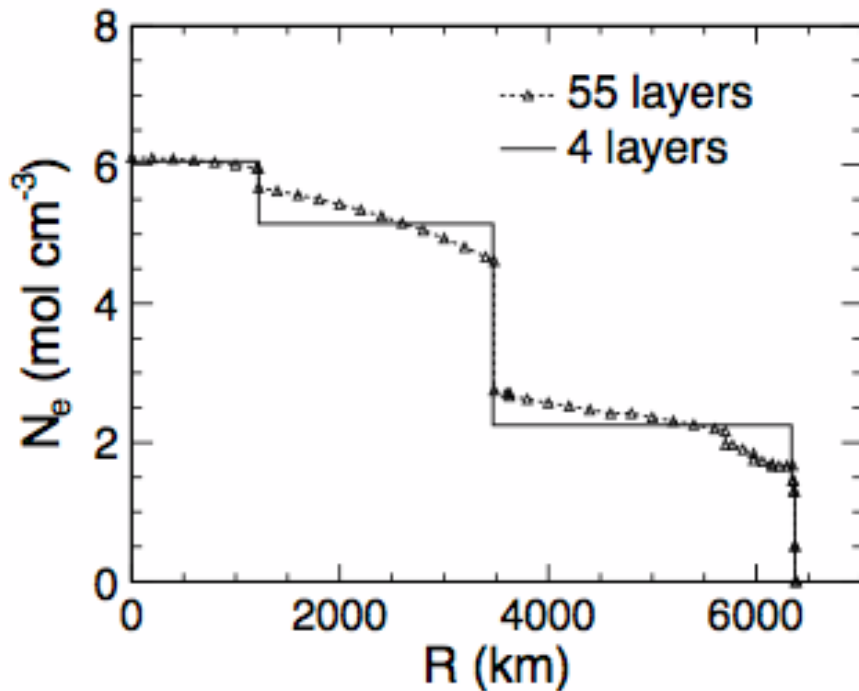
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# Earth Matter Model

- ✦ Beam: Use fixed electron density of  $1.36 \text{ mol cm}^{-3}$
- ✦ Atmospherics: Use a 4 layer model

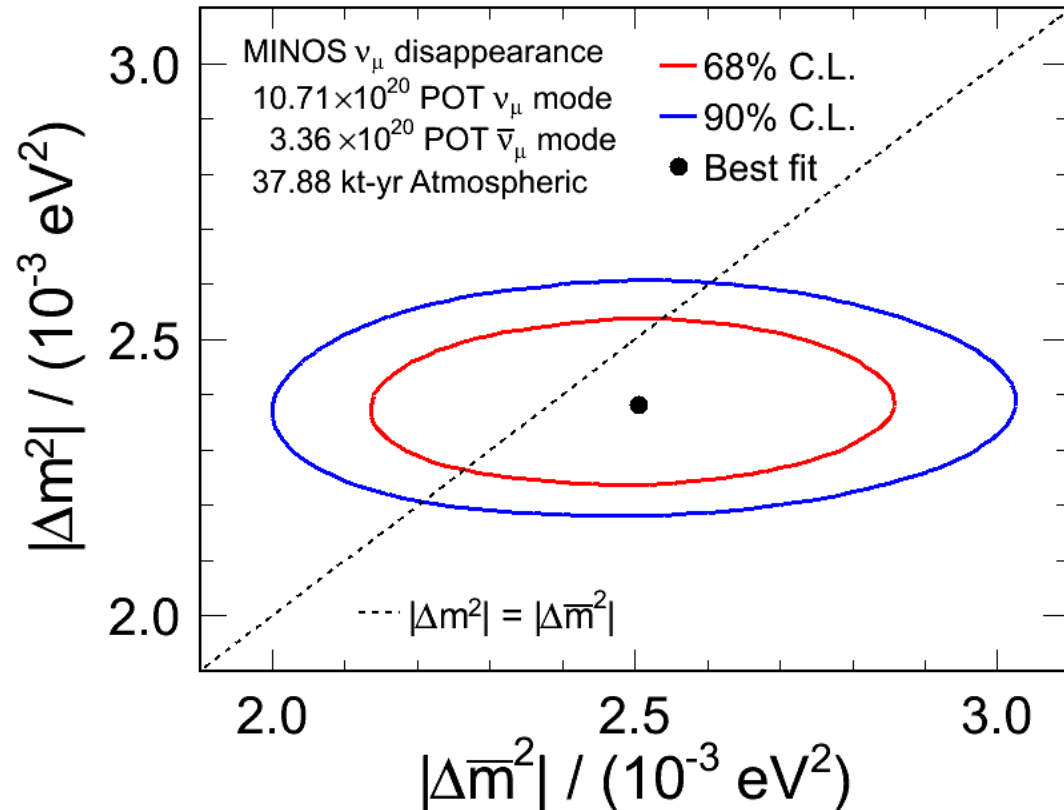


Region	Radius (km)	Earth density ( $\text{mol cm}^{-3}$ )
Crust	> 6336	1.45
Mantle	3470 – 6336	2.25
Outer Core	1220 – 3470	5.15
Inner Core	< 1220	6.05

- ✦ Negligible effect on the result between using 4 and 55 layers

# Two Flavour Disappearance: CPT

- ✦ Fitted with neutrino and antineutrino parameters separate and free to float
- ✦ The old tension between the results has vanished
  - ✦ CPT conserved



$$|\Delta \bar{m}^2| - |\Delta m^2| = \left(0.12^{+0.24}_{-0.26}\right) \times 10^{-3} eV^2$$

# Two Flavour Disappearance: Comparisons

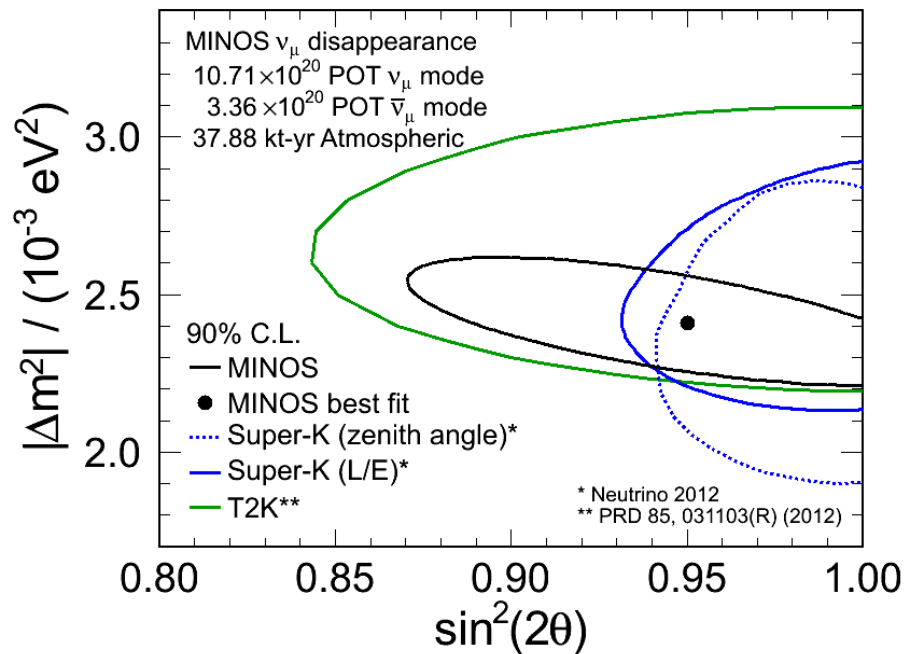
✦ Combined MINOS two-flavour contour with same oscillation parameters for neutrinos and antineutrinos.

✦ Best fit at:

$$\Delta m^2 = 2.41^{+0.09}_{-0.10} \times 10^{-3} eV^2$$

$$\sin^2 2\theta = 0.950^{+0.035}_{-0.036}$$

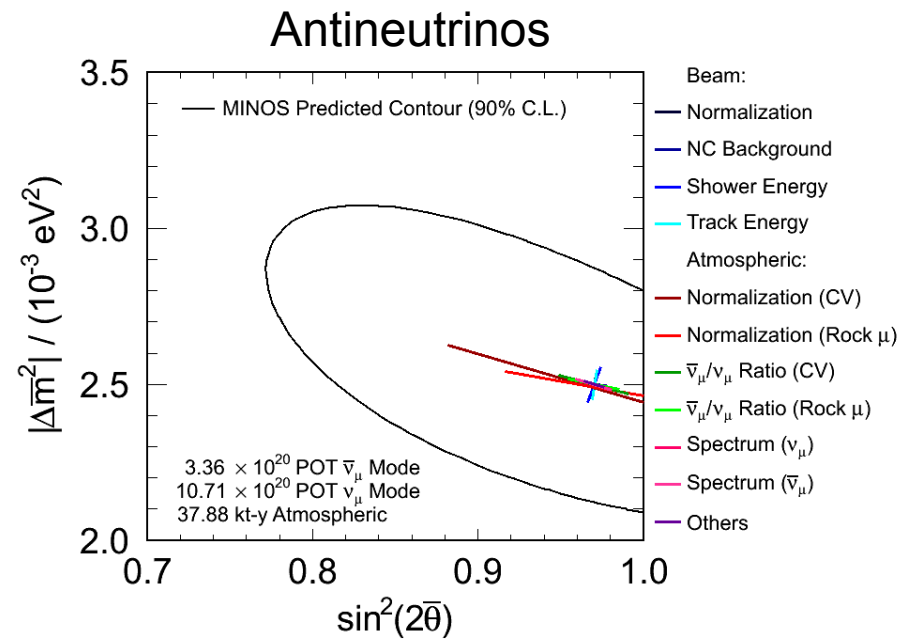
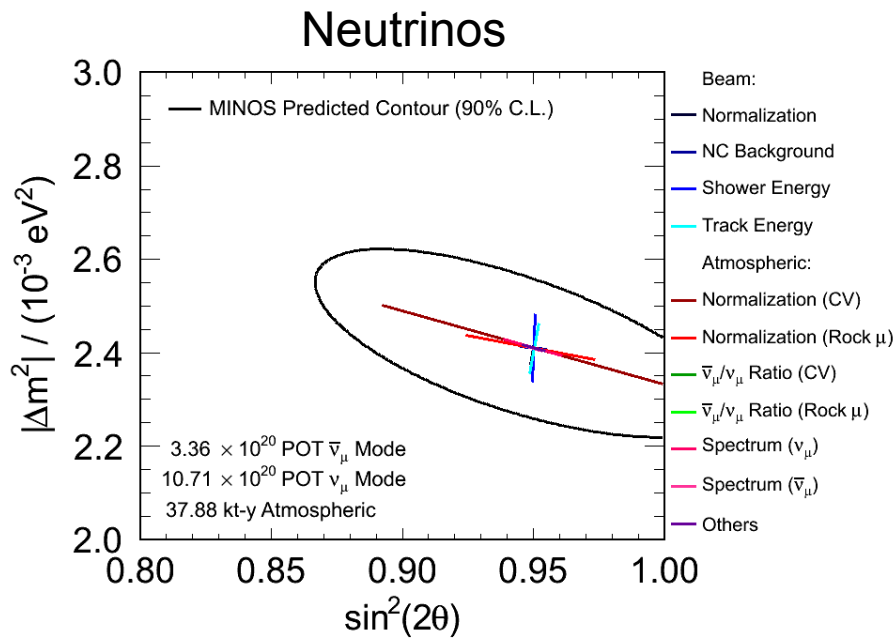
✦ Hints that mixing may not be maximal



Phys. Rev. Lett. 110, 251801 (2013)

# Disappearance: Systematic Uncertainties

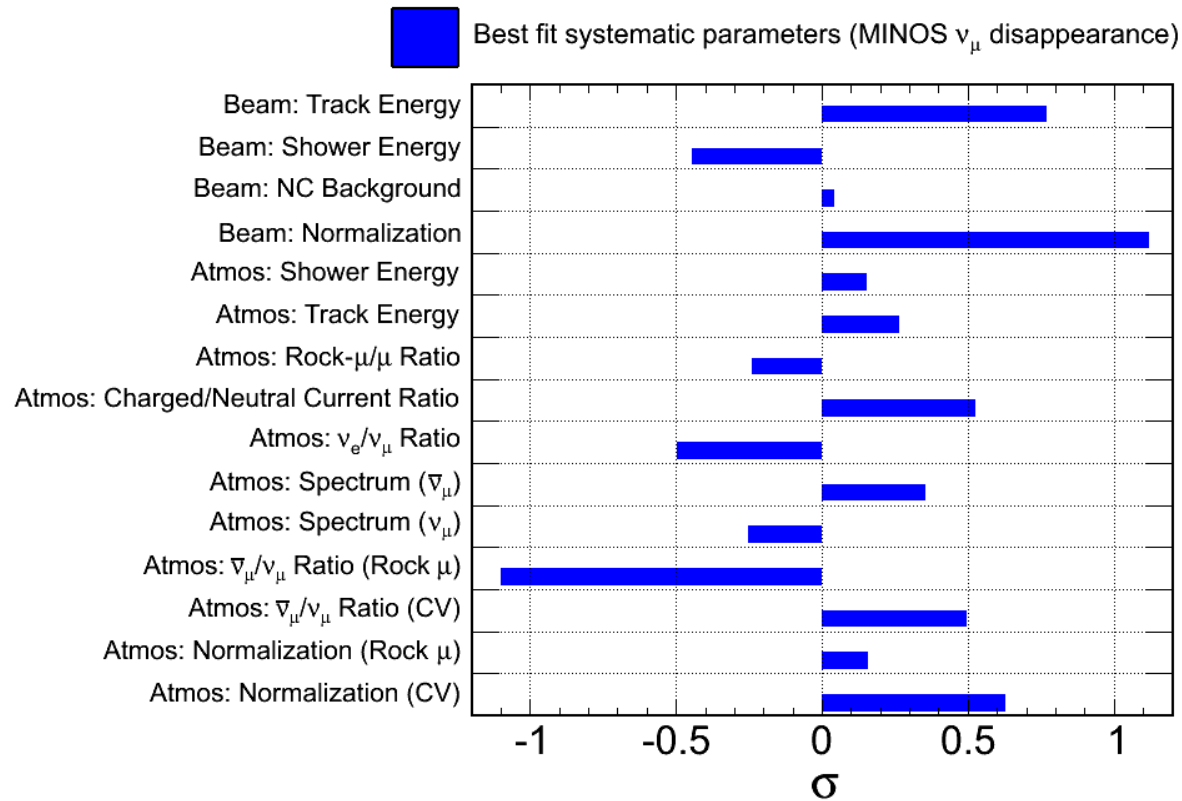
- ★ Star plots show the relative size of statistical and systematic uncertainties
  - ★ They are fitted as nuisance parameters in the fit



Phys. Rev. Lett. 110, 251801 (2013)

# Disappearance: Systematic Uncertainties

✦ Plot shows how the systematics are pulled in the fit

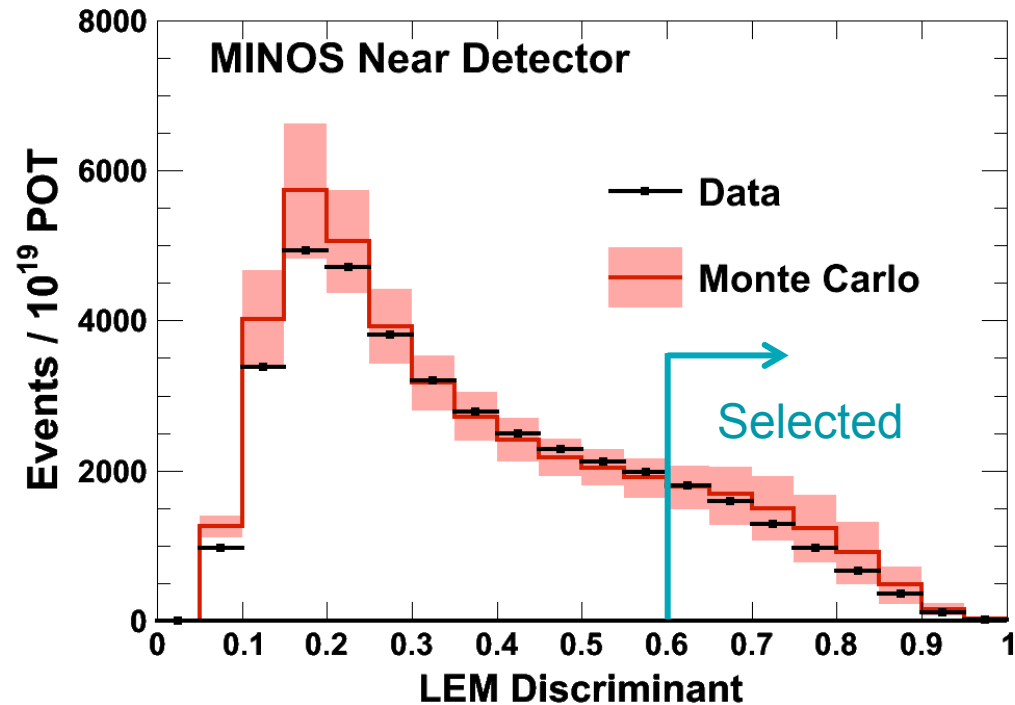


✦ The majority of these pulled by less than one sigma.



# Electron Neutrino Selection

- ✦ Very challenging given the coarse nature of the calorimeters!
- ✦ Electron neutrinos are selected using the Library Event Matching algorithm
- ✦ Candidates compared to a library of simulated signal and background events.
- ✦ The 50 best matches are used to form variables to discriminate signal and background.



Phys. Rev. Lett. 110, 171801 (2013)

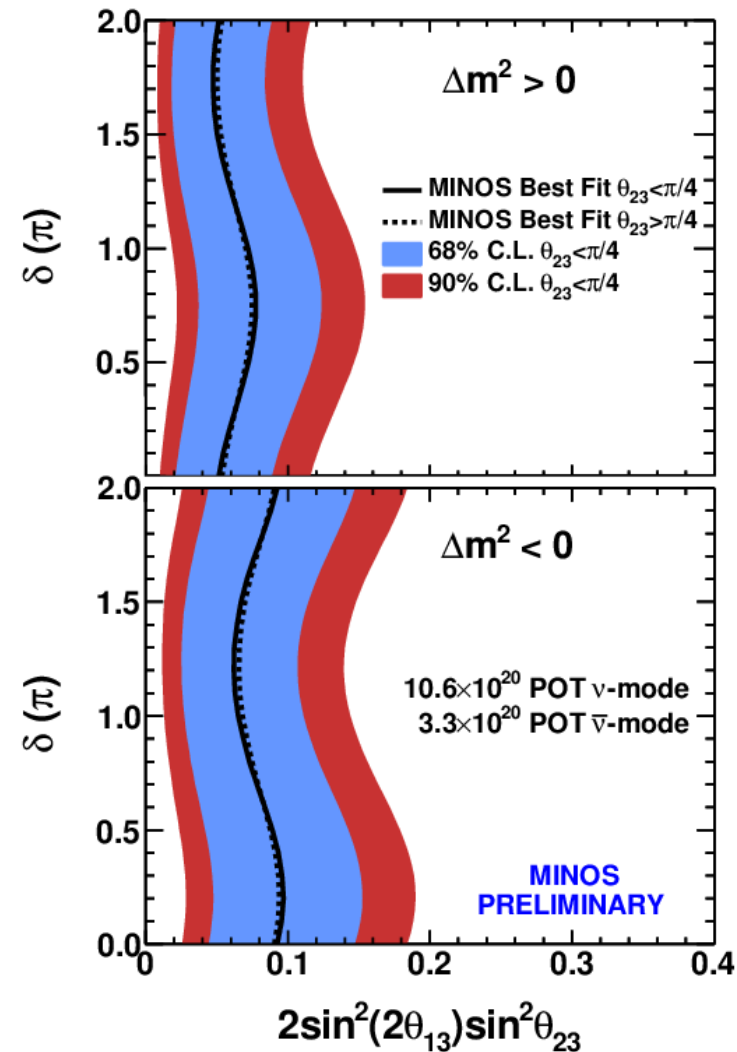
# Electron Neutrino Results

✦ Combined neutrino and antineutrino fit.

✦ Best fit, assuming  $\sin^2(2\theta_{23}) = 1, \delta = 0$  and normal (inverted) hierarchy:

$$\sin^2(2\theta_{13}) = 0.053 \text{ (0.094)}$$

$$\sin^2(2\theta_{13}) = 0 \text{ excluded at 96\%}$$



Phys. Rev. Lett. 110, 171801 (2013)

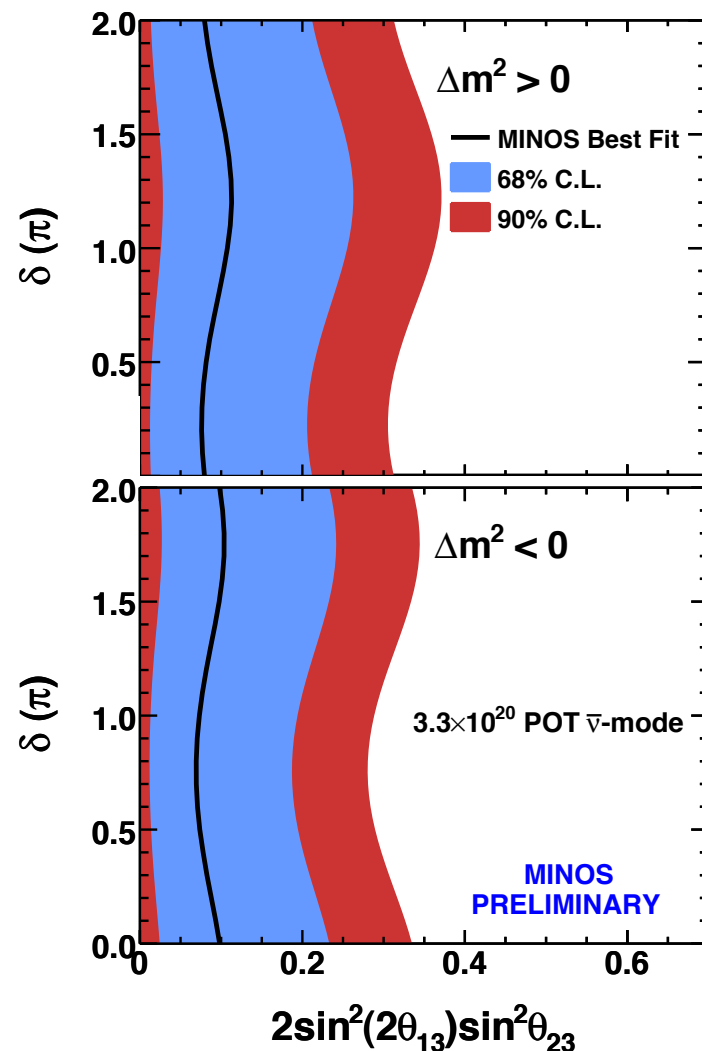
# Electron Antineutrino Results

✦ Fit to just the antineutrino events

✦ Best fit, assuming  $\sin^2(2\theta_{23}) = 1, \delta = 0$   
and normal (inverted) hierarchy:

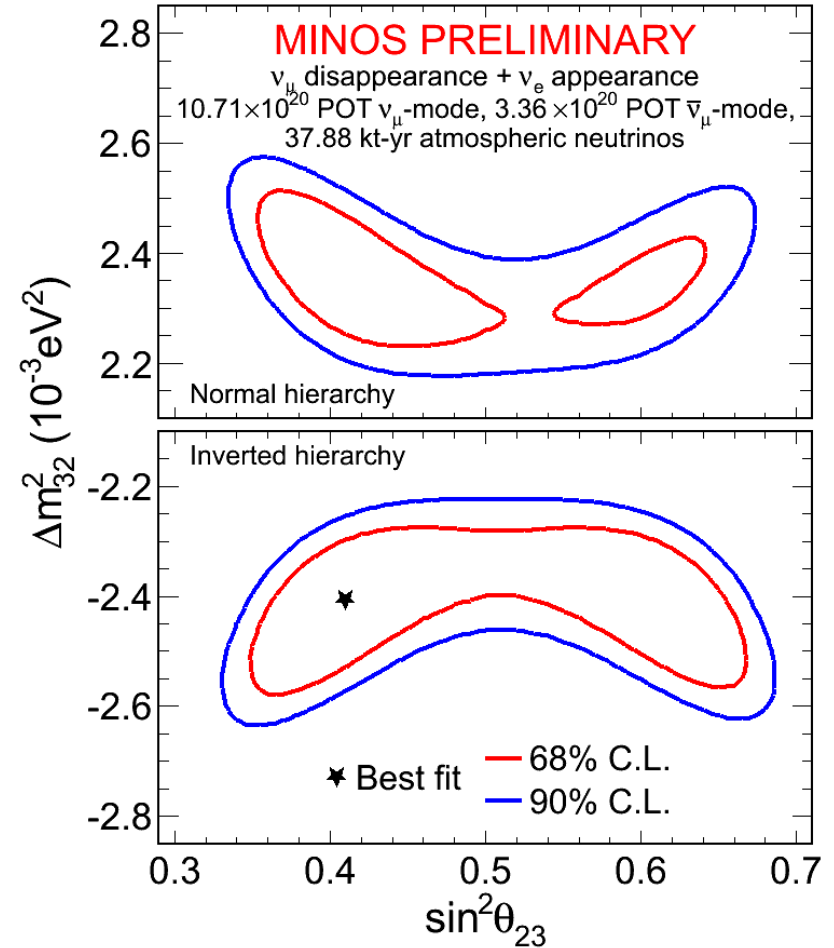
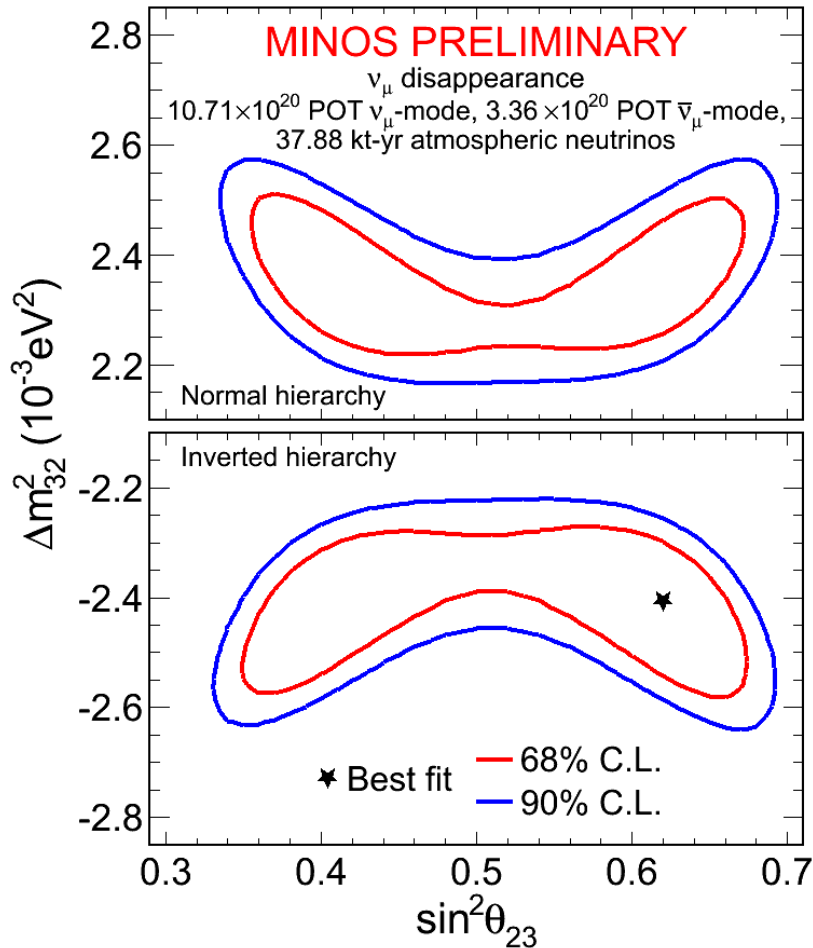
$$\sin^2(2\theta_{13}) = 0.079 (0.098)$$

$$\sin^2(2\theta_{13}) = 0 \text{ excluded at 80\%}$$



Phys. Rev. Lett. 110, 171801 (2013)

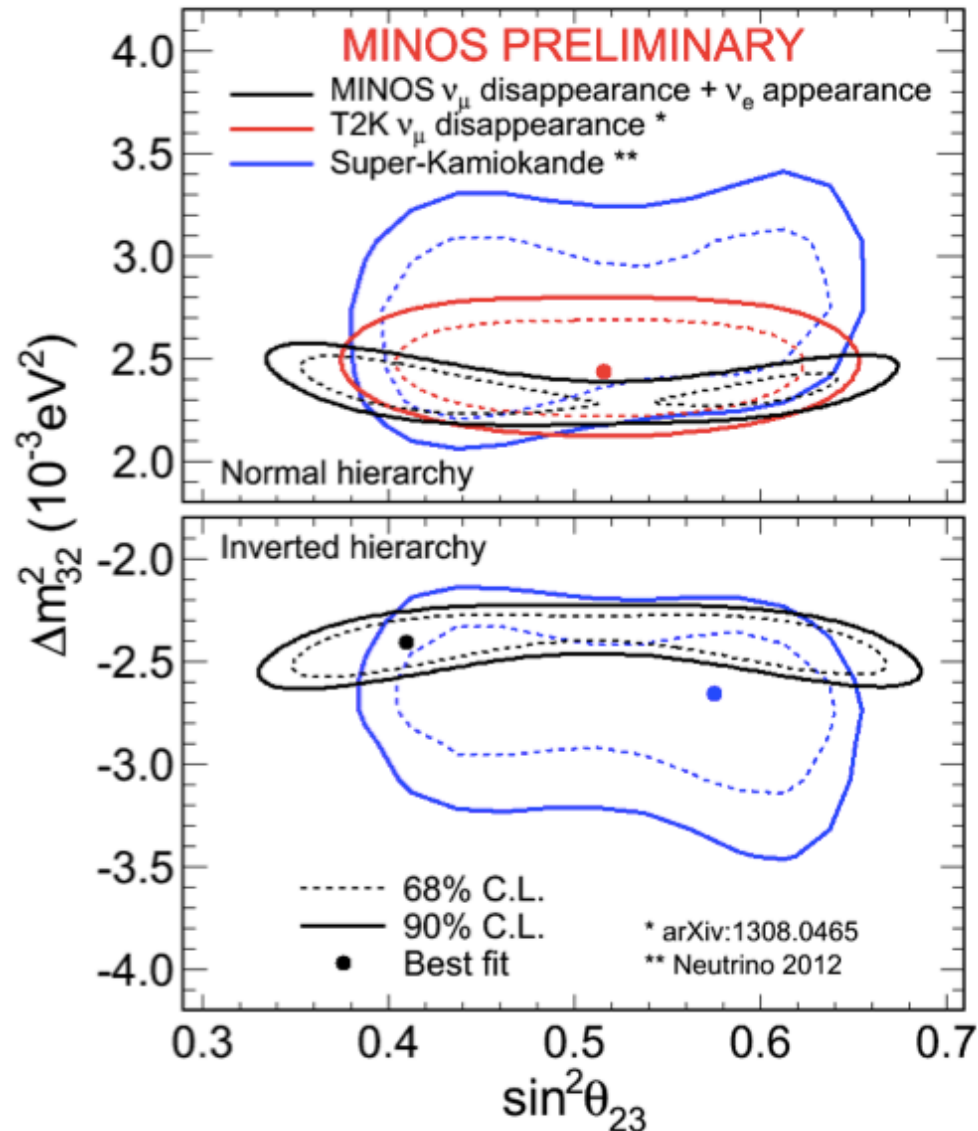
# Disappearance + Appearance



# MINOS, T2K and SK

✦ Comparison of the MINOS combined result with SK and T2K

✦ Note: T2K result converted from  $\sin^2 2\theta$  to  $\sin^2 \theta_{23}$



# Four Flavour Oscillations

✦ A fourth neutrino state adds:

✦ 3 mixing angles:  $\theta_{14}, \theta_{24}, \theta_{34}$

✦ 3 mass splittings:  $\Delta m_{43}^2, \Delta m_{42}^2, \Delta m_{41}^2$

✦ 2 additional complex phases

$$U_4 = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} & U_{e4} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} & U_{\mu4} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} & U_{\tau4} \\ U_{s1} & U_{s2} & U_{s3} & U_{s4} \end{pmatrix}$$

✦ LSND/MiniBooNE measured  $\sin^2 2\theta_{\mu e} \sim \sin^2 2\theta_{14} \sin^2 \theta_{24}$

✦ MINOS NC measures  $\sin^2 2\theta_{\mu s} \sim \sin^2 2\theta_{24} \cos^2 \theta_{34}$

✦ Reactor experiments measure  $\sin^2 2\theta_{ee} \sim \sin^2 2\theta_{14}$

✦ Therefore it makes sense to combine MINOS with a reactor experiment such as BUGEY to set a limit on  $\sin^2 2\theta_{\mu e}$



A large, industrial-scale particle detector, likely a Time-of-Flight (TOF) detector, is shown in a laboratory setting. The detector is a large, dark, rectangular structure with a complex internal structure. It is surrounded by blue and yellow metal frames and scaffolding. The number "485" is visible on the top surface of the detector. The text "TOF Analysis" is overlaid in the center of the image.



# TOF Overview

- ✦ Measure the time of flight between the ND and FD.
  - ✦ Distance of 734,286.8m
- ✦ To perform the time of flight measurement, need to know very precisely:
  - ✦ The distance between the ND and the FD.
  - ✦ All time delays:
    - ✦ Cables, GPS offsets, etc etc
- ✦ Worked with:
  - ✦ NIST Time and Frequency Division
  - ✦ USNO Time Service Department
- ✦ Perform a likelihood fit on the FD data to the time of flight

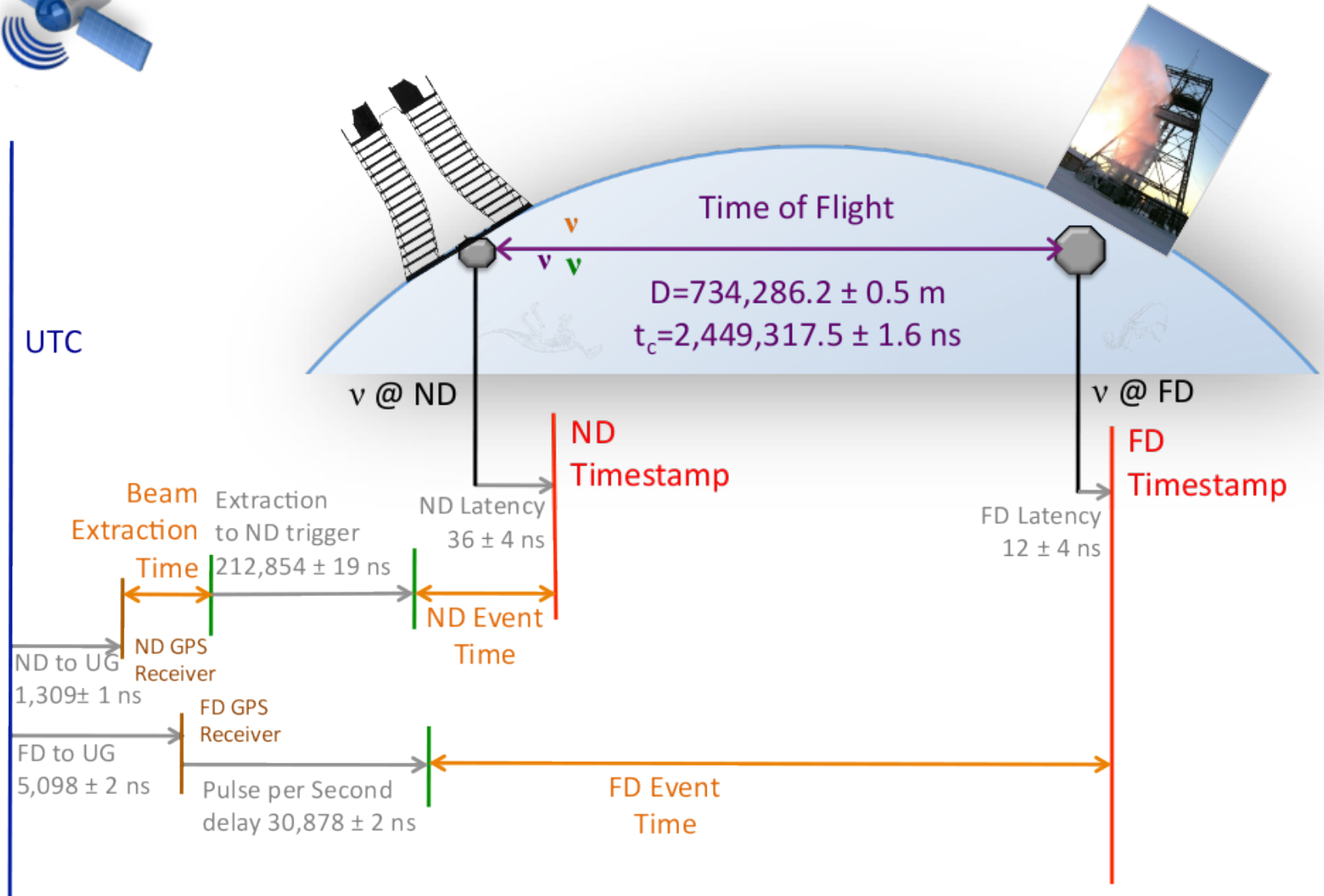
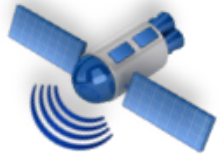


# Measuring Distances

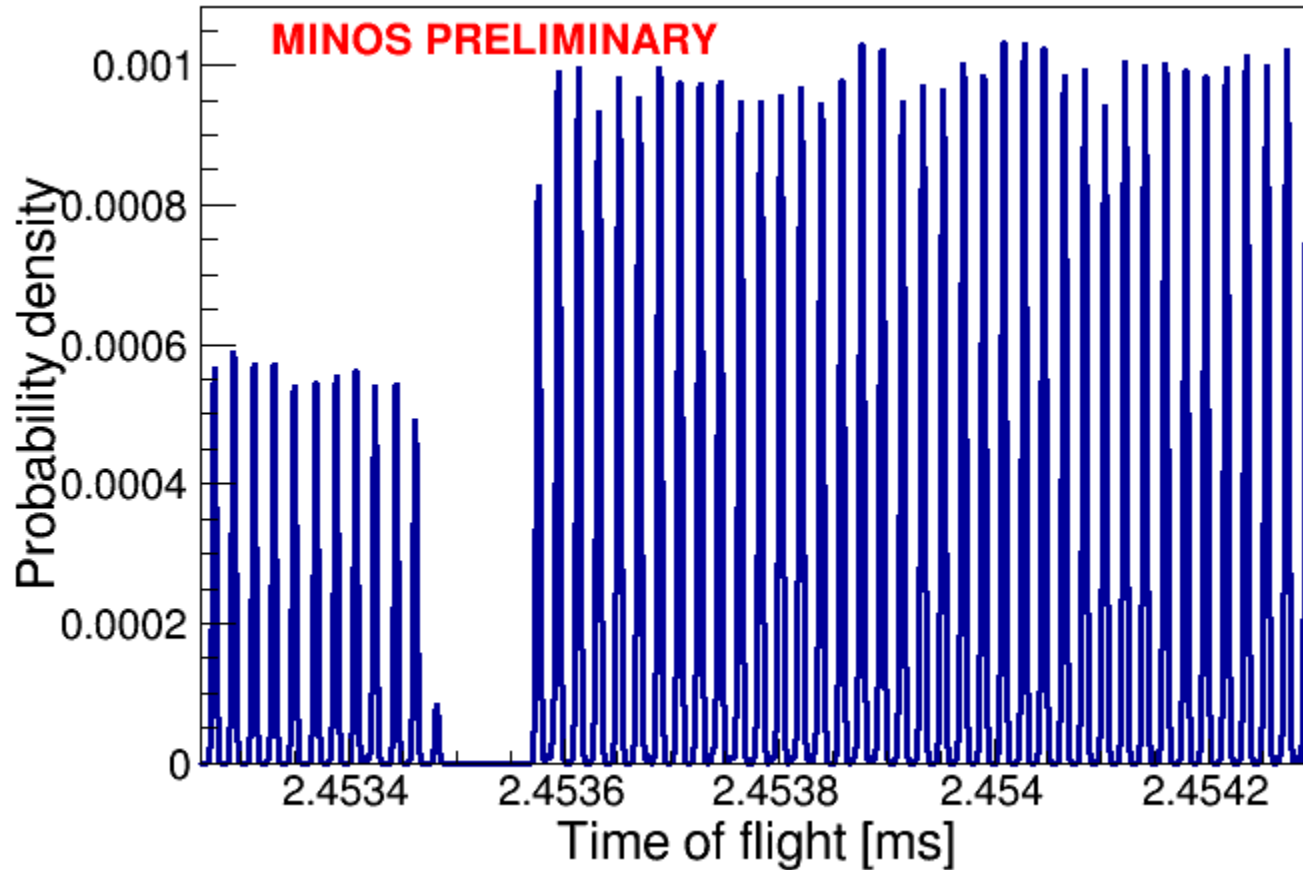
- ✦ Three components to measure:
  - ✦ Surface distance between ND and FD
  - ✦ Distance from surface to underground at ND
  - ✦ Distance from surface to underground at FD
- ✦ The last point is the most challenging
  - ✦ FD is 2341 feet down a single mineshaft
  - ✦ No line of sight from the surface to the FD
- ✦ Don't forget that the Earth rotates!



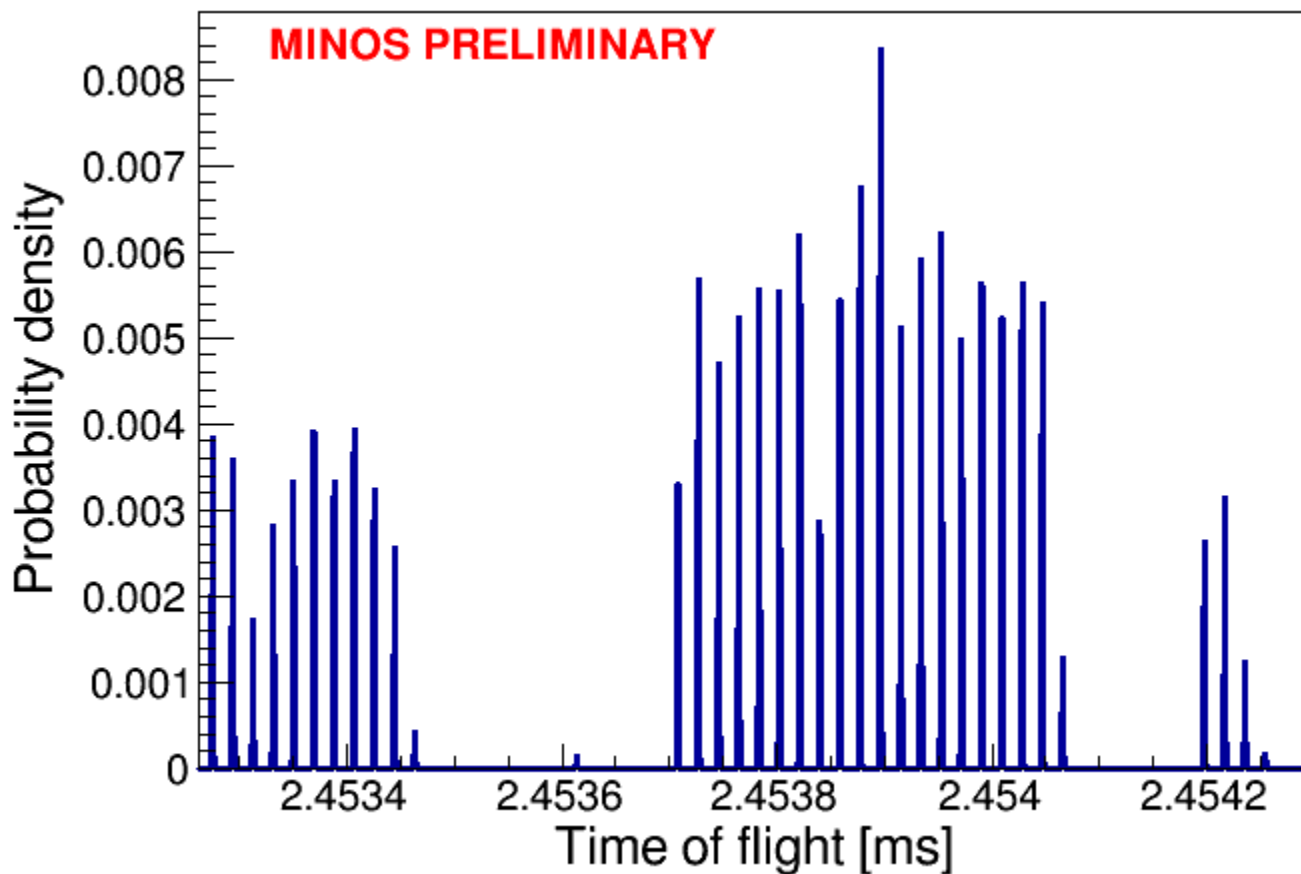
# Timing Diagram



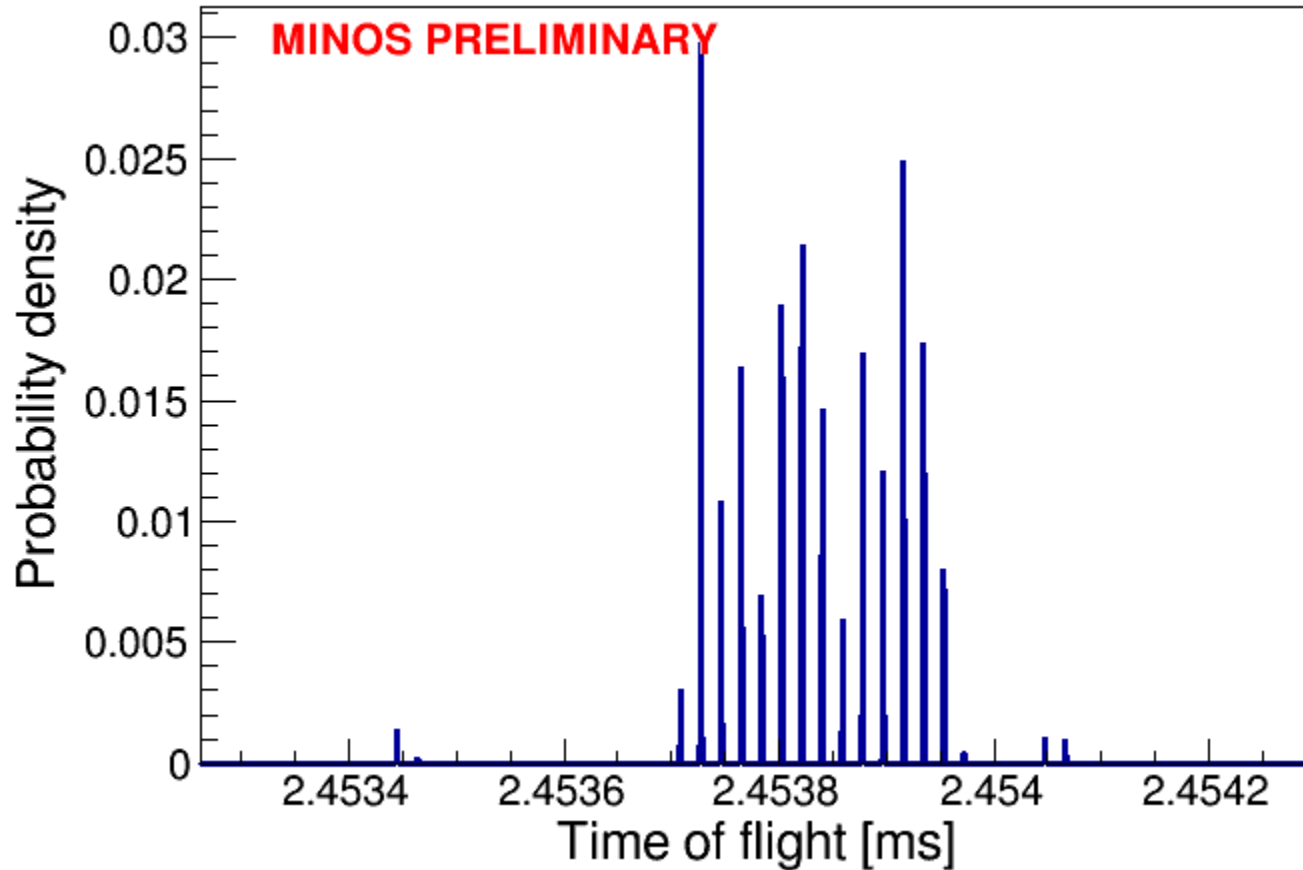
# TOF Likelihood: 1 event



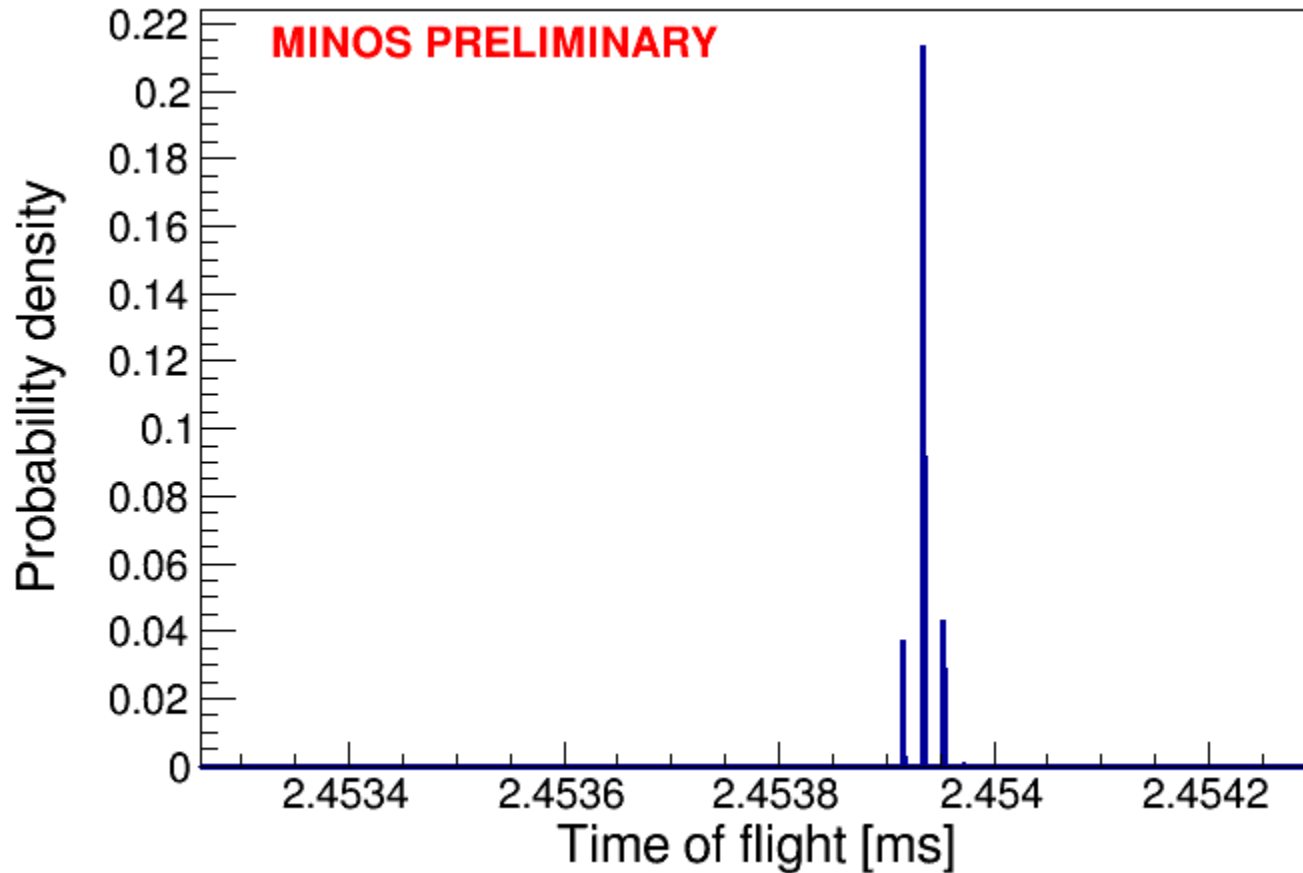
# TOF Likelihood: 10 events



# TOF Likelihood: 30 events

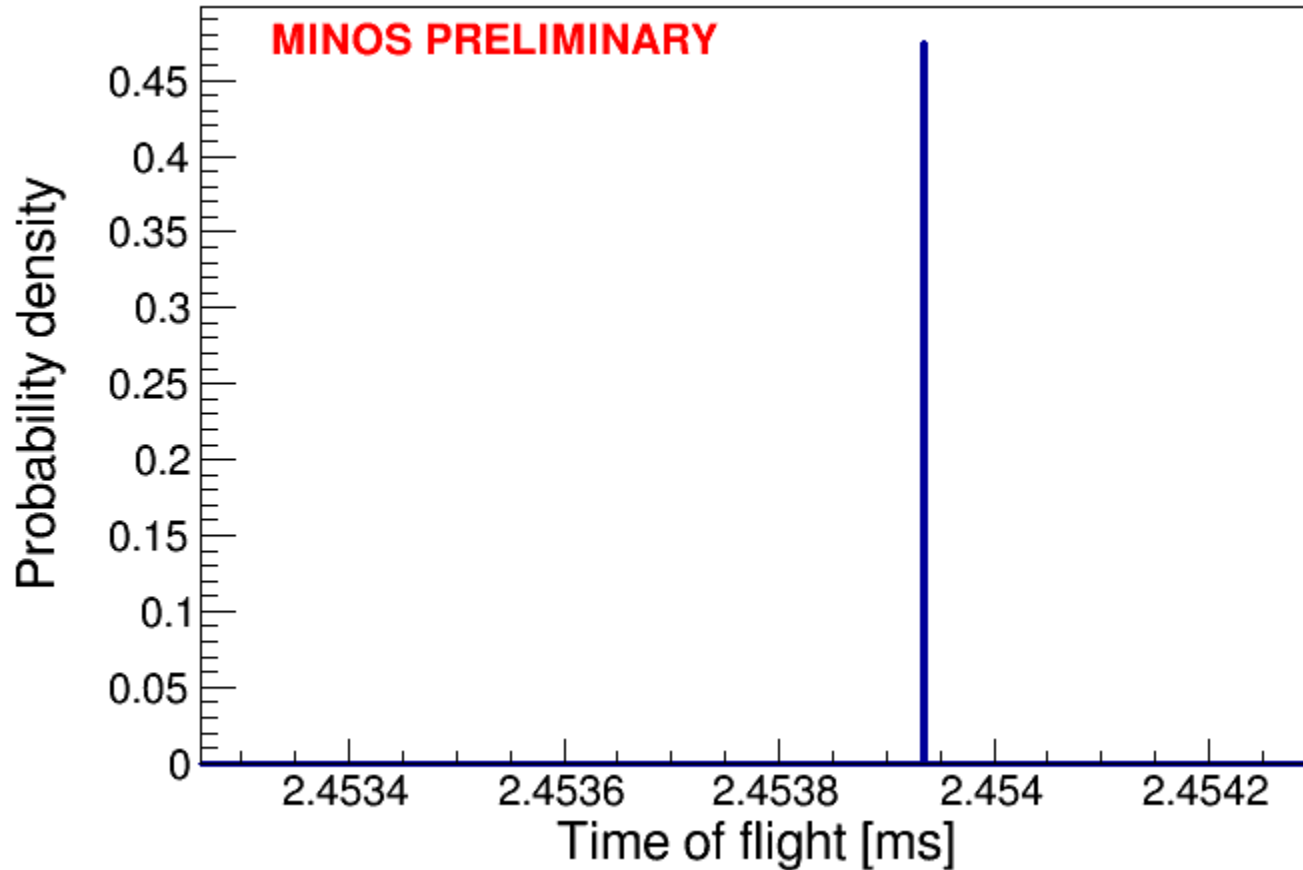


# TOF Likelihood: 60 events



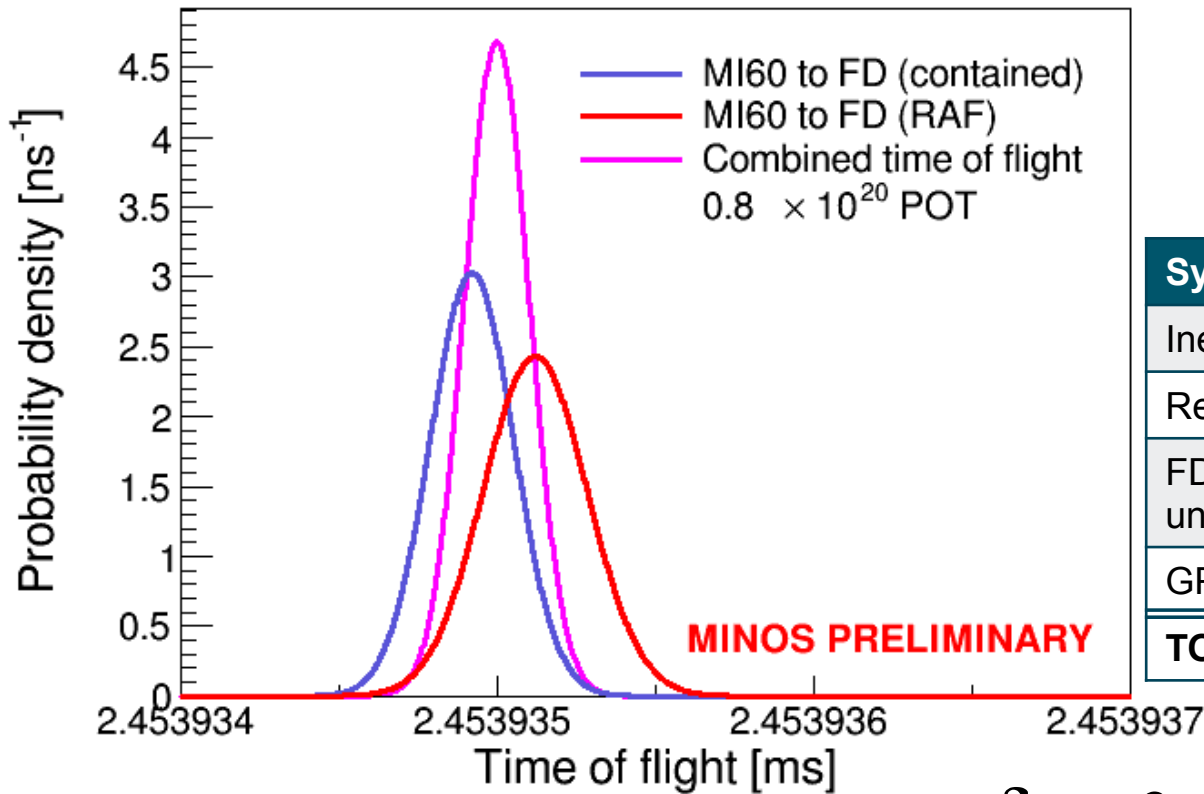


# TOF Likelihood: 195 events



# TOF Result

- ✦ Separate fits to the contained and RAF events.



Systematic uncertainty	Value
Inertial survey at FD	2.3 ns
Relative ND-FD latency	1.0 ns
FD TWTT between surface and underground	0.6 ns
GPS time transfer accuracy	0.5 ns
<b>TOTAL (quadrature sum)</b>	<b>2.6 ns</b>

- ✦ Difference from speed of light:  $\delta = -2.4 \pm 0.1(stat) \pm 2.6(syst)ns$
- $$\left(\frac{v}{c} - 1\right) = (1.0 \pm 1.1) \times 10^{-6}$$