



Results and Prospects with NOvA

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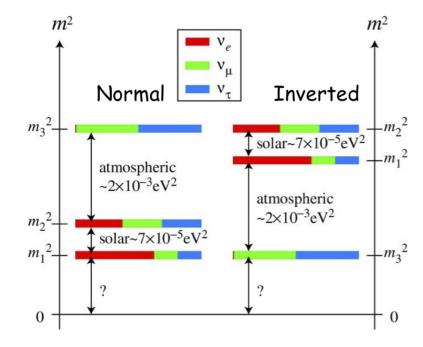
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Prospects of Neutrino Physics, Kavli IPMU, Kashiwa, Japan

NOvA Physics Goals

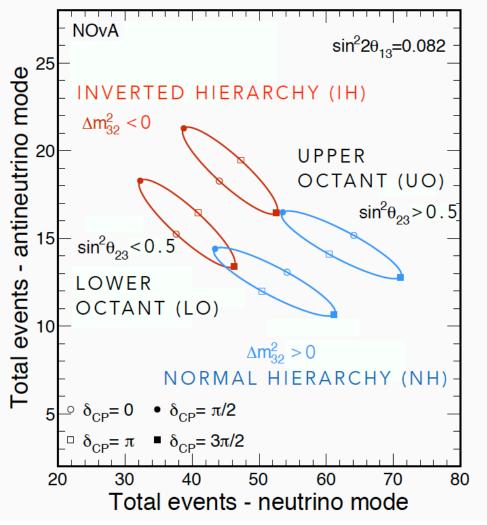
- v_e appearance + v_{μ} disappearance
 - Mass hierarchy: m₃>m_{1,2} or m_{1,2}>m₃? Implications for absolute neutrino masses, unified theories and neutrino-less double beta decay searches
 - CP phase δ_{CP} : whether neutrinos and antineutrinos behave the same way in oscillation? Implications for matter-antimatter asymmetry
 - Octant of θ_{23} : Is θ_{23} exactly 45°? Is v_3 more strongly coupled to v_{τ} or v_{μ} ?
- NC disappearance
 - Sterile neutrino search: are there other neutrinos beyond the three known active flavors?
- Also, cross sections, exotic phenomena and nonbeam physics

This talk: ν_e and ν_μ oscillation results with NOvA's first antineutrino data



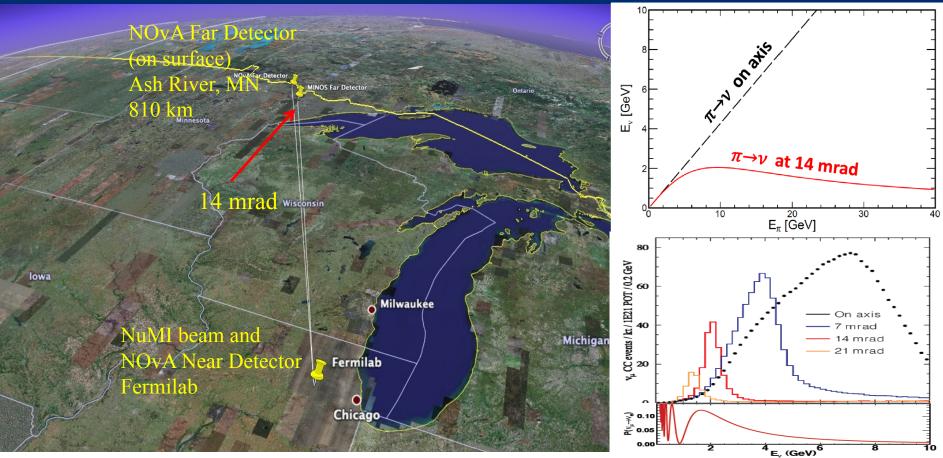
Appearance and Disappearance at NOvA

v_e/\bar{v}_e Appearance event counts



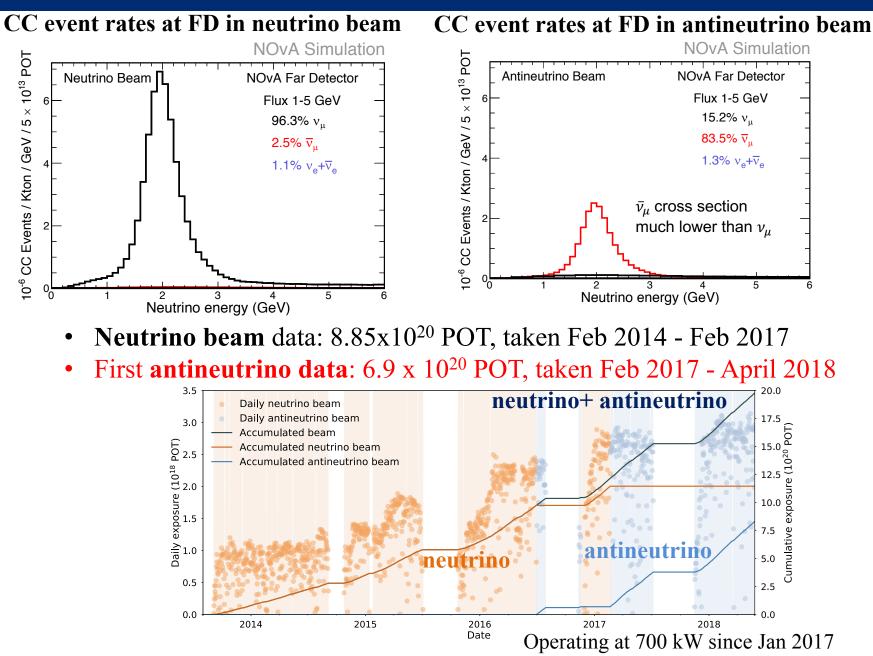
- Measuring v_e and \bar{v}_e appearance probabilities with v_{μ} and \bar{v}_{μ} beam
- When other parameters fixed, $\nu_{\mu} \rightarrow \nu_{e}$ and $\bar{\nu}_{\mu} \rightarrow \bar{\nu}_{e}$ appearance probabilities depend on
 - sign of Δm^2_{32}
 - $-\delta_{CP}$
 - octant of θ_{23}
- v_{μ} and \bar{v}_{μ} disappearance provides high precision Δm_{32} and $\sin^2 2\theta_{23}$, constrain θ_{23} octant

NuMI Off-Axis v_e Appearance Experiment



- Upgraded NuMI muon neutrino beam at Fermilab (700 kW design goal achieved)
- Longest baseline in operation (810 km), large matter effect (±30%), sensitive to mass hierarchy
- Far/Near detector sited 14 mrad off-axis, narrow-band beam around oscillation maximum, small wrong sign components ($\bar{\nu}$ in ν beam or ν in $\bar{\nu}$ beam)

Beam Performance

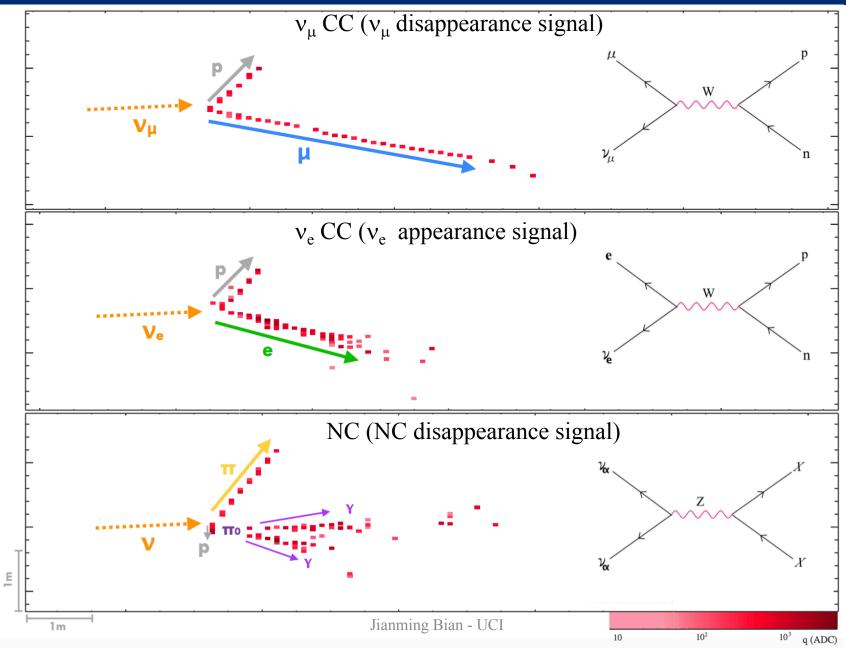


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The NOvA Detectors

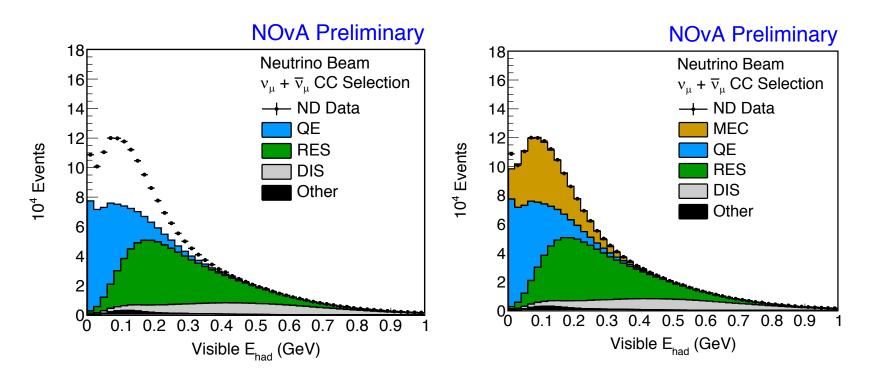
- 14-kton Far Detector 344,064 detector cells 0.3-kton functionally identical Near Detector To APD Readout 18,432 cells Far Detector, 14 kt, 60 m x 15.6 m x 15.6 m Scintillation Light Particle Trajectory 3.87 cm 14.3mx4.1mx4.1m Waveshifting Near Detector Plane of vertical cells Fiber Loop Plane of horizontal cells 6.0 cm 3.9 cm
- Composed of PVC modules extruded to form long tube-like cells
- Each cell: filled with liquid scintillator, has wavelength-shifting fiber (WLS) routed to Avalanche Photodiode (APD)
- Cells arranged in planes, assembled in alternating vertical and horizontal directions
- Low-Z and low-density, each plane just 0.15 X_0 , great for $e^- vs \pi^0$ separation

NOvA Event Topologies



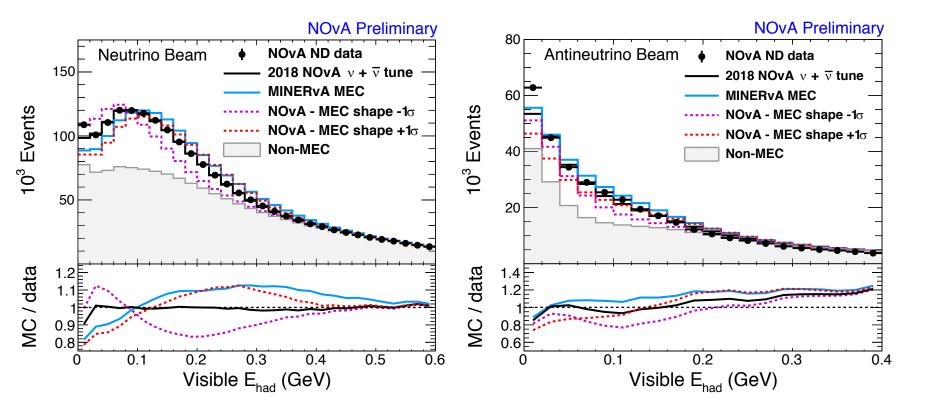
Neutrino Interaction Tuning

- QE, RES tuned to consider long-range nuclear correlations using València model via work of R. Gran (MINERvA) [https://arxiv.org/abs/1705.02932]
- DIS at high invariant mass ($W>1.7 \text{ GeV/c}^2$) weighted up 10% based on NOvA data
- Empirical MEC (Meson Exchange Current) model for Multi-nucleon ejection (2p2h) [T. Katori, AIP Conf. Proc. 1663, 030001 (2015)], amount tuned in 2D 3-momentum and energy transfers $(q_0 = E_v - E_\mu, |q| = |p_v - p_\mu|)$ space to match ND data

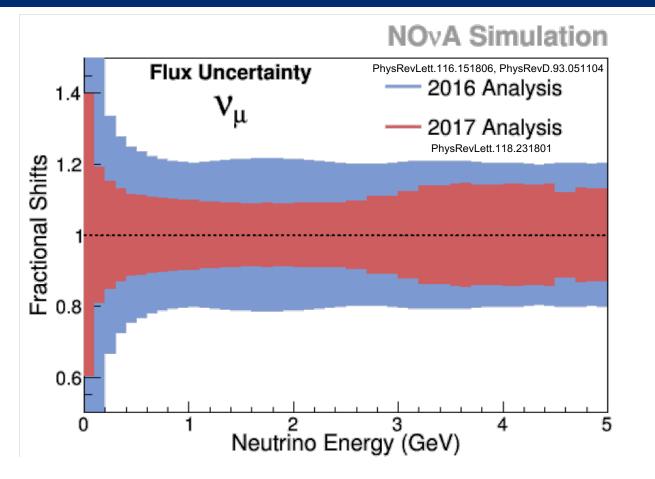


Neutrino Interaction Tuning

• MEC shape systematic estimated by re-fitting using models with QE and RES related systematic shifts



Improved Flux Model

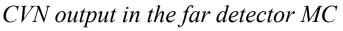


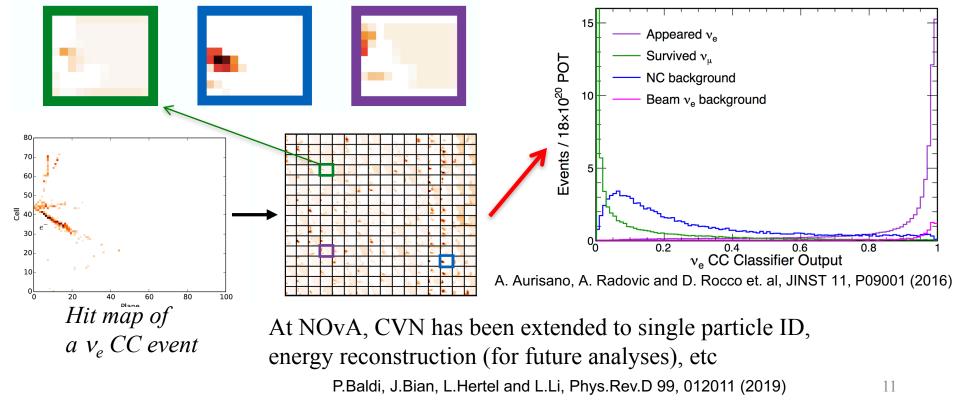
- Package to Predict the Flux (PPFX) from MINERvA (Phys. Rev. D 94, 092005. 2016).
 - Based on thin target hadron production data from NA49 and MIPP.
- Significantly reduced systematic uncertainties.
 - Central values also changed within prior systematics, but not shown here.

Deep-Learning based PID for ν_e and ν_μ Analyses

- CVN: a convolutional neural network (CNN), based on modern image recognition technology
- Introduce convolution filters to extract features from the hit map for the training of the neural net
- Statistical power equivalent to 30% more exposure than previous v_e PIDs
- v_e , v_{μ} and NC analyses all use CVN as event selector

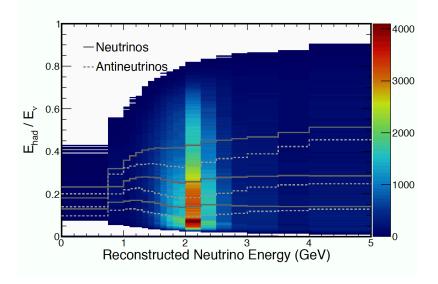
Outputs of convolutional filters (features)

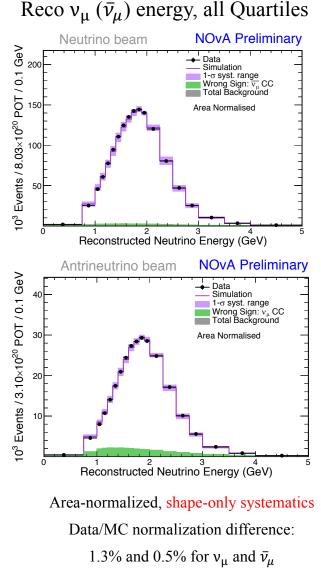




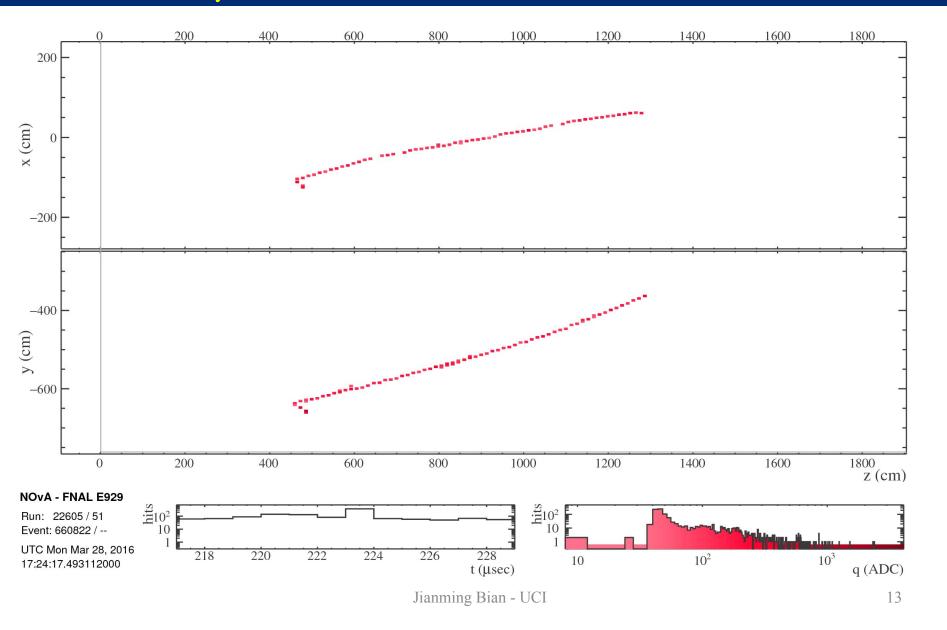
Near Detector Spectrum (v_{μ} disappearance)

- Select v_{μ} (\bar{v}_{μ}) CC in ND from neutrino (antineutrino) beam, wrong sign contamination 3% (11%)
- $E_{\nu} = E_{\mu} + E_{had}$, data split in 4 equal energy quartiles based on E_{had}/E_{ν} , resolution varies from 5.8% (5.5%) to 11.7% (10.8%) for neutrino (antineutrino) beam.
- Normalize ND MC to data in each E_v bin, then extrapolate the 4 quantiles to FD



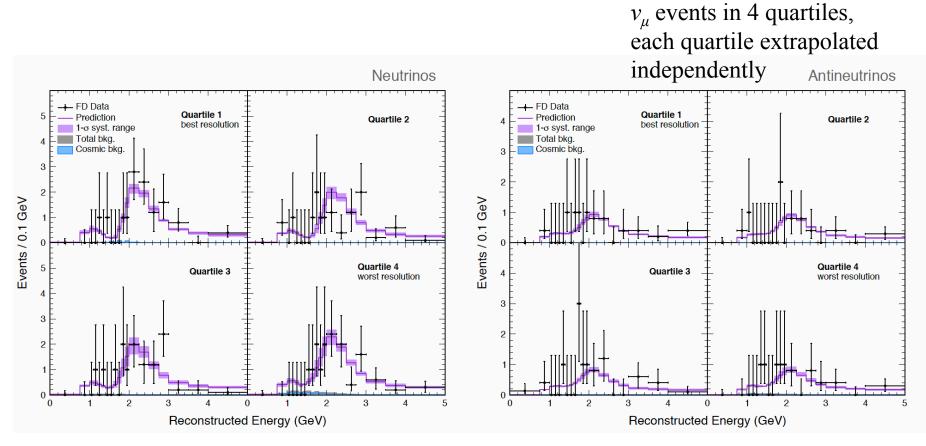


 v_{μ} Data at Far Detector



v_u Data at Far Detector

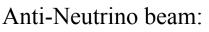
- FD selection:
 - Additional Boost Decision Tree (BDT) to reduce cosmic backgrounds
 - Estimate cosmic background rate from timing sidebands of the NuMI beam triggers and cosmic trigger data



v_{μ} Data at Far Detector

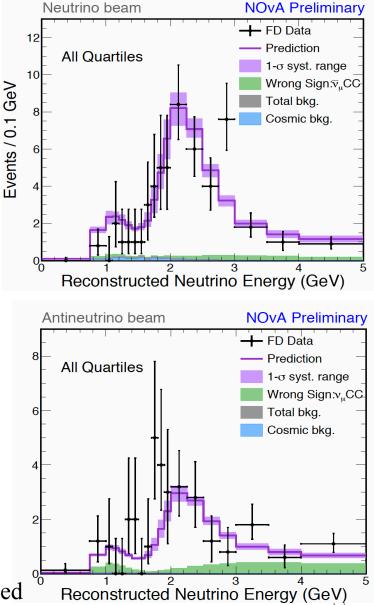
Neutrino beam:

- Observe 113 events
- Prediction at best fit: 121
- Cosmic background: 2.1
- Beam background: 1.2
- Expect 730 +38/-49(syst.) w/o oscillations



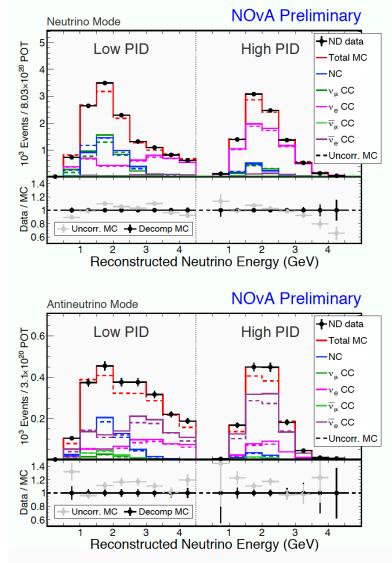
- Observe 65 events
- Prediction at best fit 50
- Cosmic background 0.5
- Beam background 0.6
- Expect 266 +12/-14(syst.) w/o oscillations

4 quartiles combined



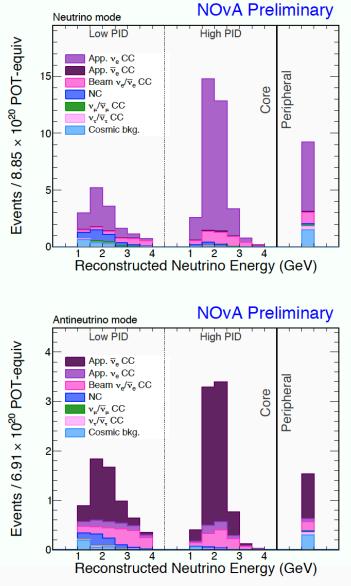
Near Detector Spectrum (v_e appearance)

- Select $v_e(\bar{v}_e)$ CC in ND from neutrino (antineutrino) beam
- $E_v = f(E_e, E_{had})$, data split into low and high particle ID (purity) range
- For neutrino beam:
 - Contained and uncontained v_{μ} events constrain the π/K contributions to the beam v_e 's.
 - Michel electrons constrain NC/ v_{μ} CC balance in each E_v bin
- For antineutrino beam, scale all components evenly to match data
- ND→FD extrapolation: Each component propagated independently in energy and PID bins



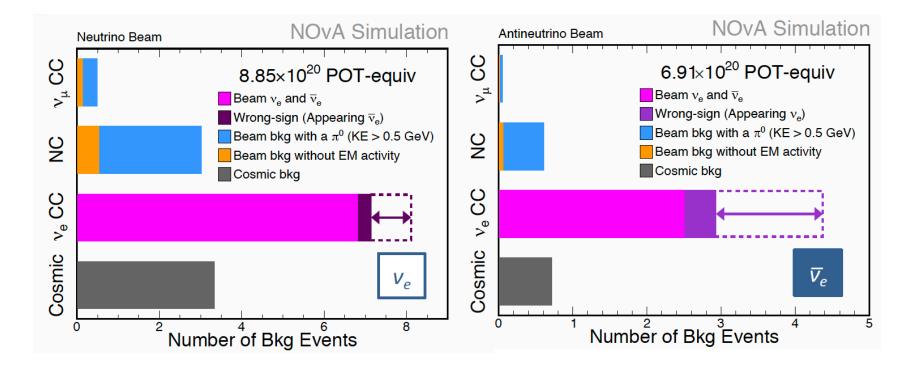
v_e Far Detector Prediction

- FD selection:
 - Add a one-bin peripheral with less stringent containment selection to include more signal
 - Use location dependent BDT and tight PID cuts to recover signal events in this peripheral bin
- ND→FD extrapolation: Each component propagated independently in energy and PID bins
 - ND v_{μ} sample to predict FD v_{μ} background and appearance v_{e} (signal+wrong-sign)
 - ND v_e -like sample to predict FD beam v_e backgrounds

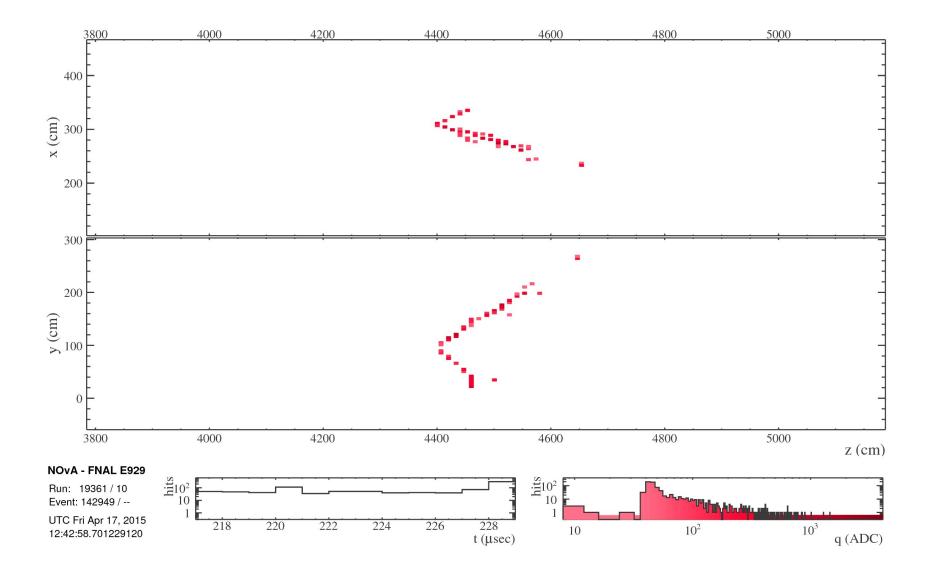


v_e Far Detector Backgrounds

- Major backgrounds from beam v_e
- Wrong sign background depends on oscillation



v_e Data at Far Detector



v_e Data at Far Detector

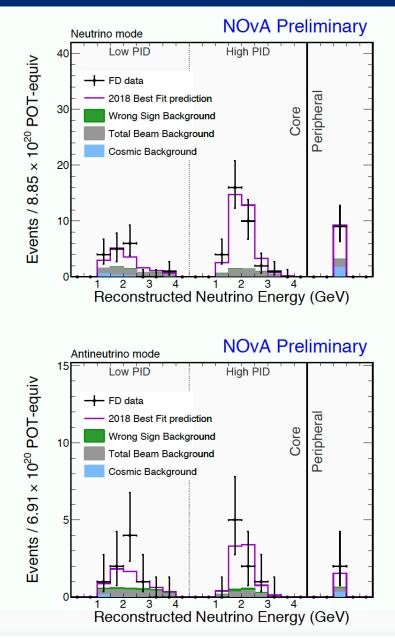
Neutrino beam:

- Observe 58 events
- Prediction at best fit 59.0
- Appearance v_e (signal) 43.9
- Total background 15.1
 - Appearance $\bar{\nu}_e$ (wrong-sign) 0.7
 - Beam $v_e + \bar{v}_e$ 6.8
 - Cosmic background 3.3

Anti-Neutrino beam:

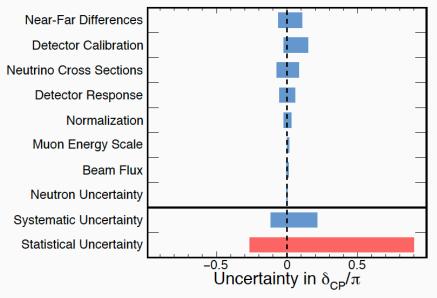
- Observe 18 events
- Prediction at best fit 16.1
- Appearance $\bar{\nu}_e$ (signal) 10.6
- Total background 5.5
 - Appearance v_e (wrong-sign) 1.1
 - Beam $v_e + \bar{v}_e$ 2.8
 - Cosmic background 0.7

> $4\sigma \bar{\nu}_e$ appearance

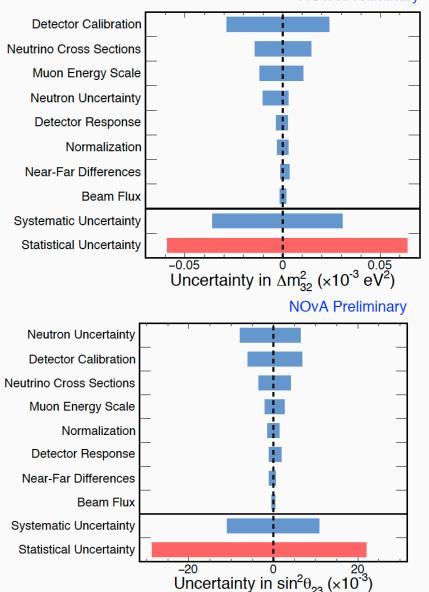


Systematic Uncertainties (Joint fit)

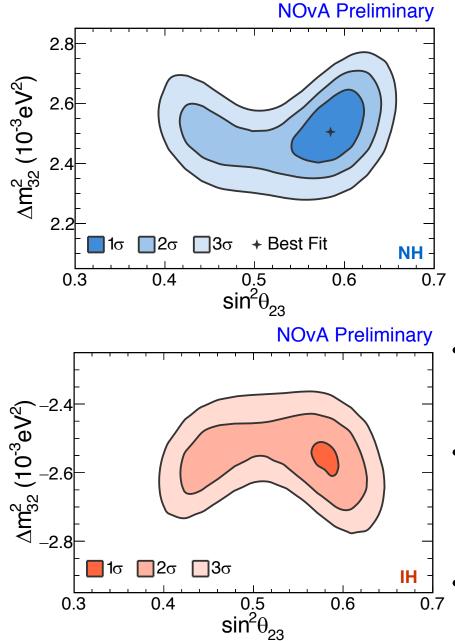
NOvA Preliminary

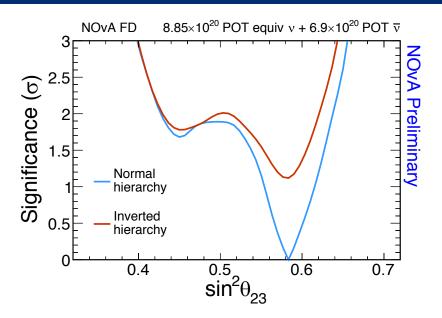


NOvA Preliminary



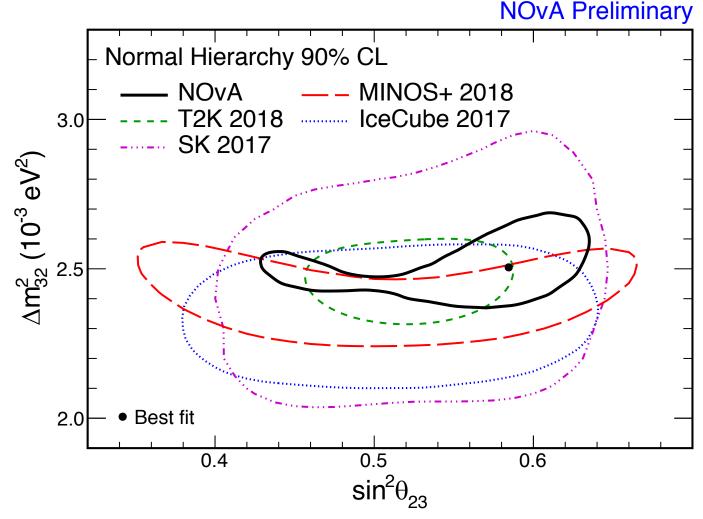
- Largest systematics for v_{μ} and v_{e} are **calibration, muon energy scale** and **cross-sections**.
- Neutron uncertainty new with v's
- Upcoming NOvA test beam program will address calibration and detector response uncertainties



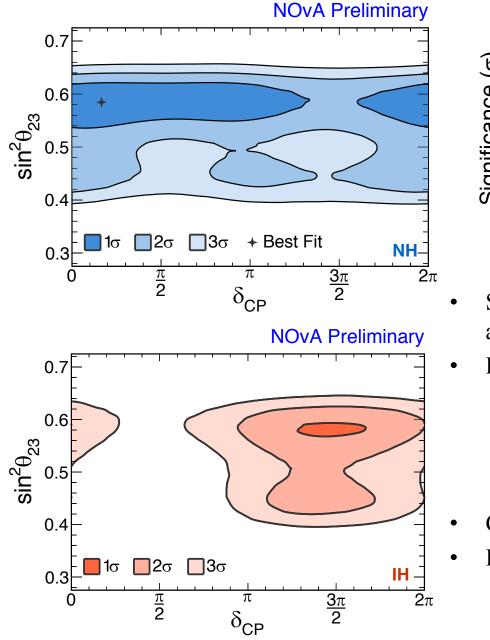


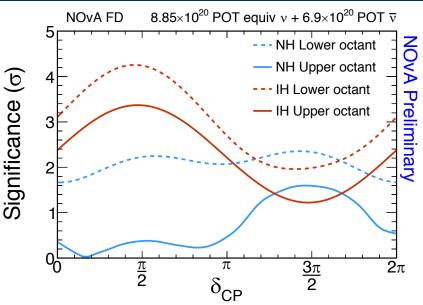
- Statistically limited, largest systematics for v_{μ} and v_{e} are calibration and crosssections.
- Best fit:
 - Normal Hierarchy
 - $\sin^2 \theta_{23} = 0.58 \pm 0.03 \text{ (UO)}$
 - $\Delta m_{32}^2 = (2.51 + 0.12 0.08) * 10^{-3} \text{ eV}^2$
 - Prefer non-maximal at 1.8σ, favor upperoctant at similar level22

NOvA's allowed 90% C.L. regions are compatible to other experiments



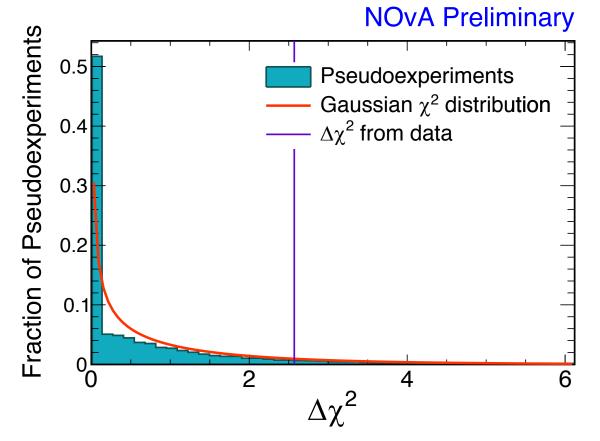
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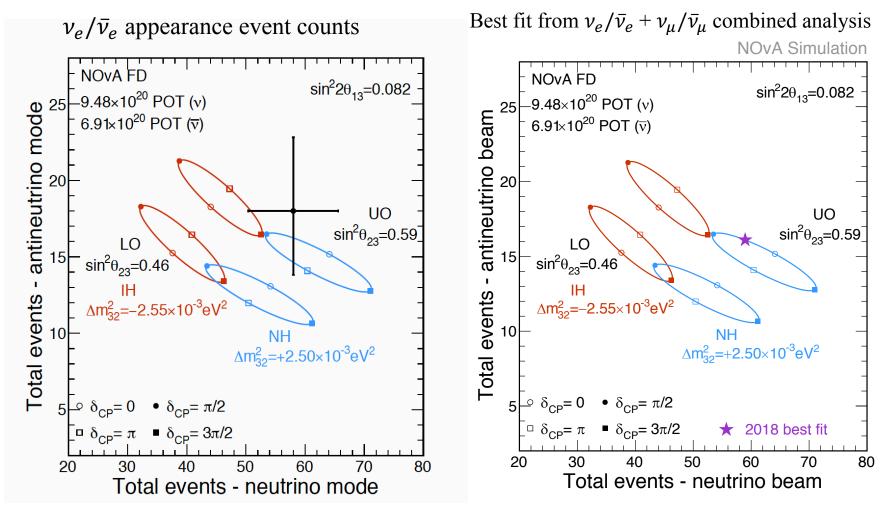
- Statistically limited, largest systematics for v_{μ} and v_e are calibration and cross-sections
- Best fit:
 - Normal Hierarchy
 - $\delta_{CP} = 0.17\pi$
 - $\sin^2 \theta_{23} = 0.58 \pm 0.03$ (UO)
 - $\Delta m_{32}^2 = (2.51 + 0.12 0.08) * 10^{-3} \text{ eV}^2$
- Consistent with all δCP values in NH at $< 1.6\sigma$
- Exclude $\delta = \pi/2$ in IH at $> 3\sigma$

Mass Hierarchy Preference



Feldman-Cousins pseudo-experiments: assume IH, set other oscillation parameters to best fit values in data under IH hypothesis **Fit:** allow CP, octant and other parameters to float

From data $\Delta \chi^2 = \chi^2 (IH) - \chi^2 (NH) = 2.47$ p-value is 0.076 \rightarrow Prefer NH at 1.8 σ



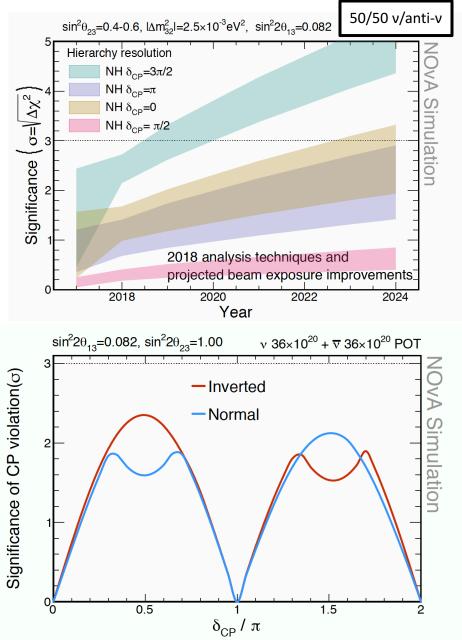
Error bars represent counting uncertainty of v_e/\bar{v}_e appearance, full power from joint fit to $v_e/\bar{v}_e + v_\mu/\bar{v}_\mu$ energy/PID spectra

- Prefer non-maximal at 1.8σ, favor upper octant Consistent with all δ_{CP} values in NH at < 1.6σ
- Exclude $\delta = \pi/2$ in IH at $> 3\sigma$
- Prefer NH at 1.8σ

Looking Forward

- Taking antineutrino data since 2017-2018, switched back to neutrinos in Feb 2019, run 50% neutrino, 50% antineutrino
- Extended running through 2025, test
 beam program and potential accelerator
 improvement to enhance ultimate reach
- If $\delta_{CP} = 3\pi/2$, 3 σ sensitivity to MH by 2020, ~5 σ by 2024
- 3 σ to MH for 30-50% (depending on octant) of δ_{CP} range by 2024
- 2+ σ to CP at $\delta_{CP}=3\pi/2$ or $\delta_{CP}=\pi/2$ by 2024

Thank you!

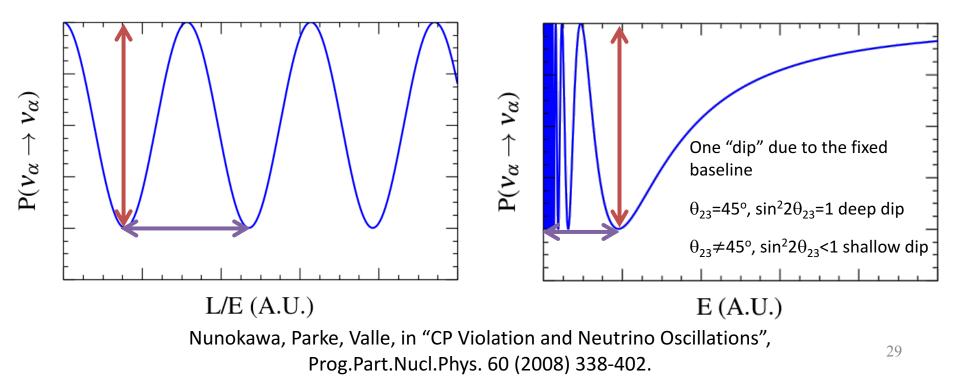




v_{μ} disappearance

$$P(\mu\mu) \approx 1 - \sin^2 2\theta_{23} \sin^2 \left(\frac{\Delta m_{32}^2 L}{4E}\right)$$

 v_{μ} disappearance: High precision Δm_{32} and $\sin^2 2\theta_{23}$, constrain octant



v_e appearance

$$P(v_{\mu} \rightarrow v_{e}) \approx \sin^{2} 2\theta_{13} \sin^{2} \theta_{23} \frac{\sin^{2} (A-1)\Delta}{(A-1)^{2}}$$

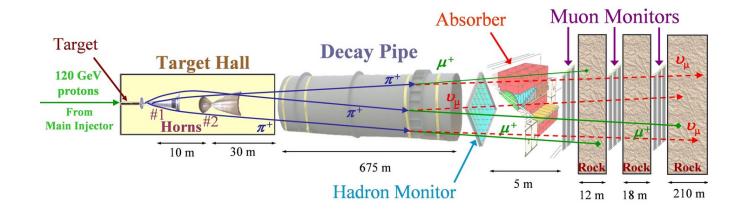
+
$$2\alpha \sin\theta_{13} \cos\delta_{CP} \sin 2\theta_{12} \sin 2\theta_{23} \frac{\sin A\Delta}{A} \frac{\sin(A-1)\Delta}{(A-1)} \cos\Delta$$

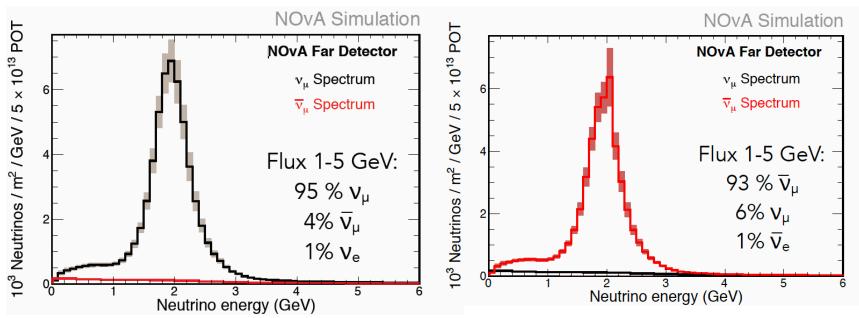
$$- 2\alpha \sin \theta_{13} \sin \delta_{CP} \sin 2\theta_{12} \sin 2\theta_{23} \frac{\sin A\Delta}{A} \frac{\sin(A-1)\Delta}{(A-1)} \sin \Delta$$

$$\alpha = \frac{\Delta m_{21}^2}{\Delta m_{31}^2} \qquad \Delta = \frac{\Delta m_{31}^2 L}{4E} \qquad A = +G_f N_e \frac{L}{\sqrt{2}\Delta}$$

- Measuring mass hierarchy (sign of Δ value), δ_{CP} and octant of θ_{23} with v_e appearance,
- $P(\nu_{\mu} \rightarrow \nu_{e})$ difference between $\Delta > 0$ and $\Delta < 0$ enlarged by matter effect A ($\propto L$ when fix L/E to oscillation maximum)

NuMI Off-Axis v_e Appearance Experiment





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Systematic Error in Calibration

• Our calibration is built on dE/dx from stopping cosmic muons.

 θ_{12}

Data

300

NOvA Preliminary

MC π^0 signal

MC bkgd

Data µ: 134.2 ± 2.9 MeV

Data σ : 50.9 ± 2.1 MeV MC μ : **136.3** ± 0.6 MeV MC σ : 47.0 ± 0.7 MeV

400

- Control samples for calibration uncertainty
 - π^0 mass peak in ND

 $m_{\pi^0}^2 = 2E_{\gamma_1}E_{\gamma_2}(1 - \cos\theta_{12})$

NC π^0

events

100

200

M_{vv} (MeV)

300

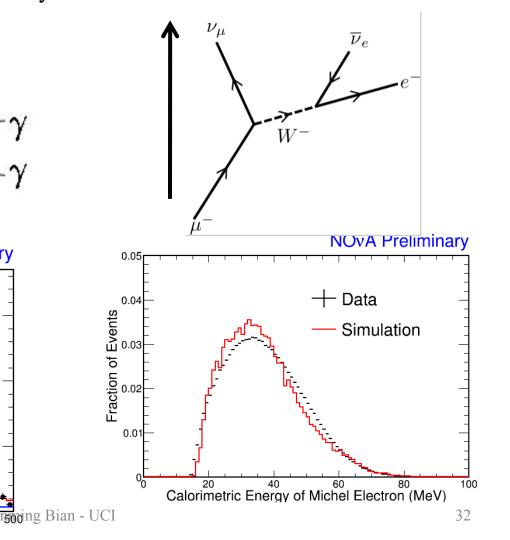
200

100

Events

- Michel electrons in ND and FD

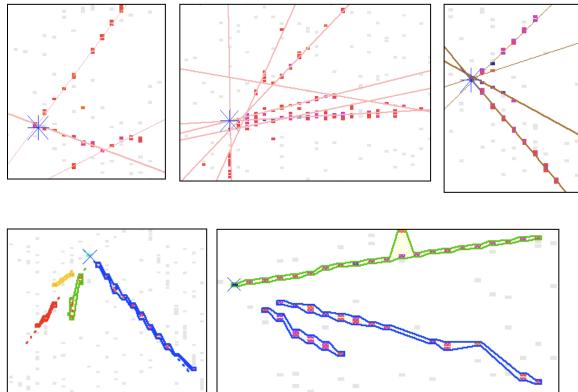
Michel electrons from muon decays



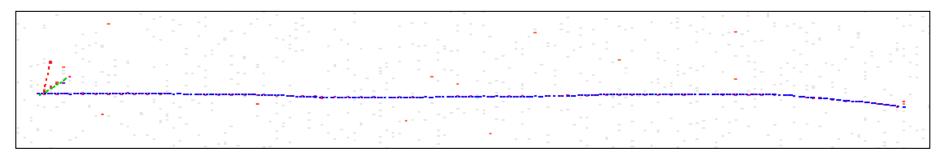
Prong/track Reconstruction

<u>Vertexing:</u> Find lines of energy depositions with Hough transform. Then determine the vertex that all lines converge to

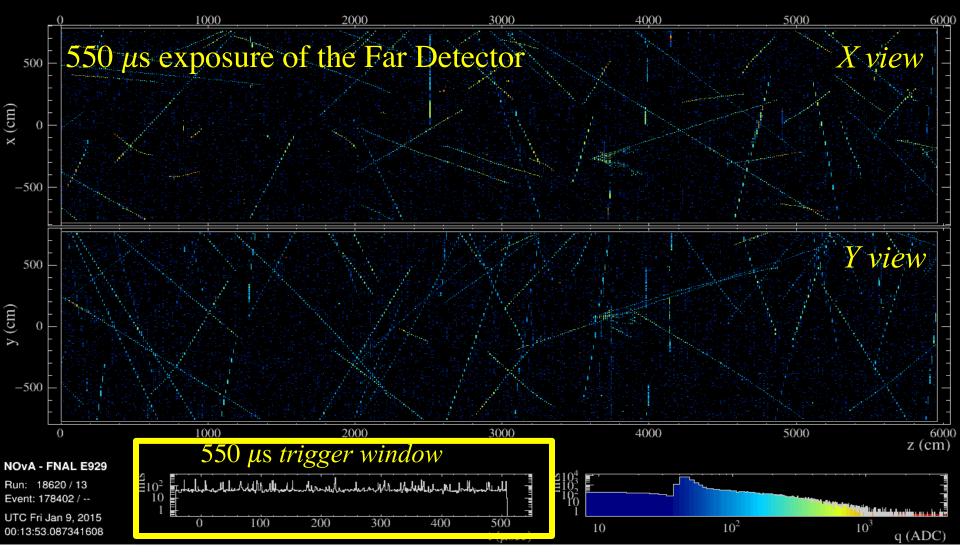
Shower Clustering: Based on the vertex and the lines, showers are reconstructed by angular clustering



<u>**Tracking:**</u> Trace particle trajectories with **Kalman filter** tracker (below). Also have a **cosmic ray tracker** that reconstructs cosmic tracks with high speed.

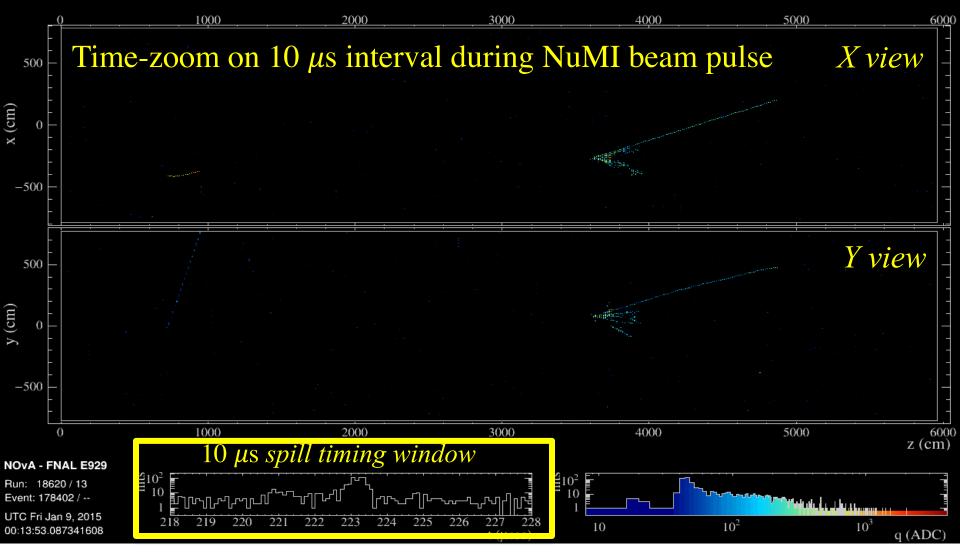


Event clustering



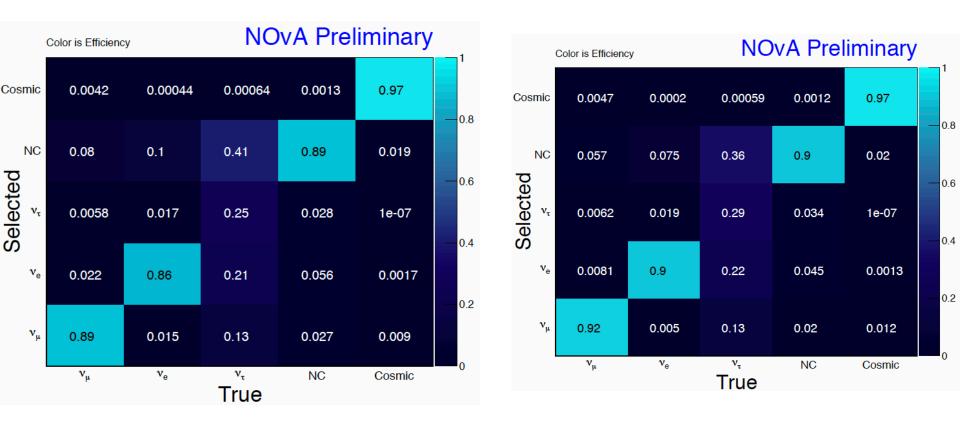
Because NOvA is on surface, hits in a trigger window are a combination of cosmic and beam events.
 First step in reconstruction is to cluster hits by space-time coincidence to separate neutrino hits and cosmic hits.

Event clustering



Event clusters that contain neutrino interactions can be correctly selected in the neutrino spill timing window
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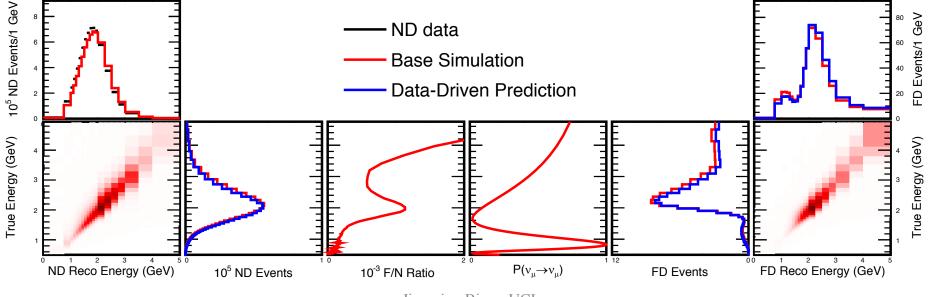
PID efficiencies



Analysis Strategy

- Separate $v_{\mu}/v_{e}/NC$ signal from beam backgrounds
- Extrapolate observed ND spectrum to FD, reject cosmic rays in FD, make FD unoscillated prediction
- Measure shapes and yields of signal events in energy/PID bins in FD to determine oscillation parameters

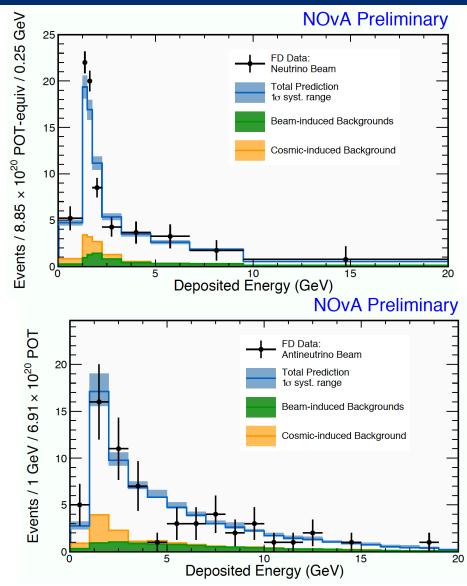




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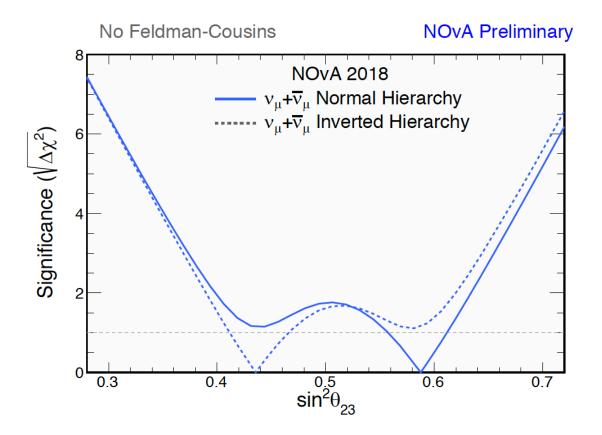
Observed NC events in Far Detector

- FD selection:
 - NC CVN selection applied
 - Additional Deep-learning based cosmic rejection
- Neutrino beam:
 - Observe 201 events, predict 188 ± 13 (syst.) events (38 bkg.)
- Antineutrino beam:
 - Observe 61 events, predict 69 ± 8 (syst.) events (16 bkg.)
- No significant suppression for NC observed, consistent with 3-flavor oscillation



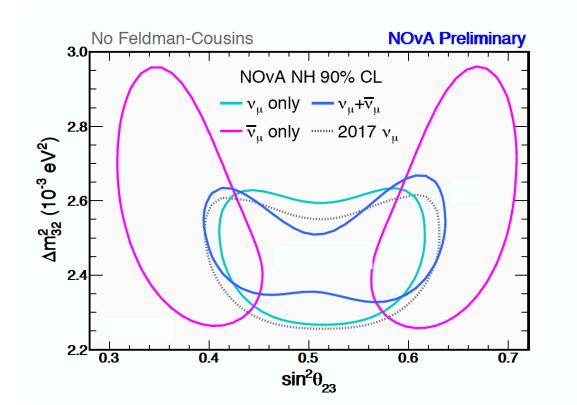
v_{μ} appearance fit

- Combined data of neutrino and antineutrino beams fitted assuming CPT invariance
- If fit separately, $\bar{\nu}_{\mu}$ data prefers non-maximal while ν_{μ} prefers maximal
- χ^2 s consistent with combined fit oscillation parameters with p > 4%
- Matter effects introduce small asymmetry in the point of maximal disappearance, ~1σ prefers Upper (Lower) Octant in NH (IH)

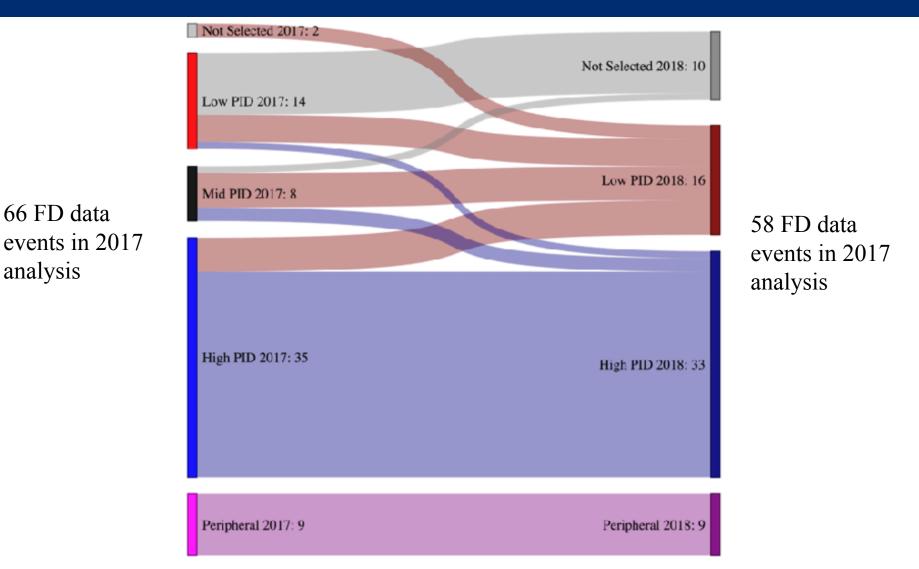


v_{μ} appearance fit

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- Matter effects introduce small asymmetry in the point of maximal disappearance



2017/2018 RHC v_e FD Data



Change in data events after retraining of PID, new training improved bkg rejection

Systematic Uncertainties (Joint Fit)

Source of Uncertainty	$\sin^2 \theta_{23} \; (\times 10^{-3})$	δ_{CP}/π	$\Delta m_{32}^2 \; (\times 10^{-3} \; {\rm eV}^2)$
Beam Flux	+0.42 / -0.48	+0.0088 / -0.0048	+0.0016 / -0.0015
Detector Calibration	+6.9 / -6.1	$+0.15 \; / \; -0.023$	+0.024 / -0.029
Detector Response	+1.9 / -0.99	$+0.055 \ / \ -0.054$	+0.0027 / -0.0034
Muon Energy Scale	+2.6 / -2.1	$+0.015 \; / \; -0.0026$	$+0.01 \ / \ -0.012$
Near-Far Differences	+0.56 / -1.1	+0.11 / -0.064	+0.0033 / -0.0013
Neutrino Cross Sections	+4.2/-3.5	+0.085 / -0.072	+0.015 / -0.014
Neutron Uncertainty	+6.4 / -7.9	+0.002 / -0.0052	+0.0028 / -0.01
Normalization	+1.4 / -1.5	+0.031 / -0.024	+0.0029 / -0.0027
Systematic Uncertainty	+9.6 / -11	+0.21 / -0.11	+0.032 / -0.035
Statistical Uncertainty	+22 / -29	$+0.9 \; / \; -0.27$	+0.064 / -0.059