

Prospects for a Gd-loaded Super-K



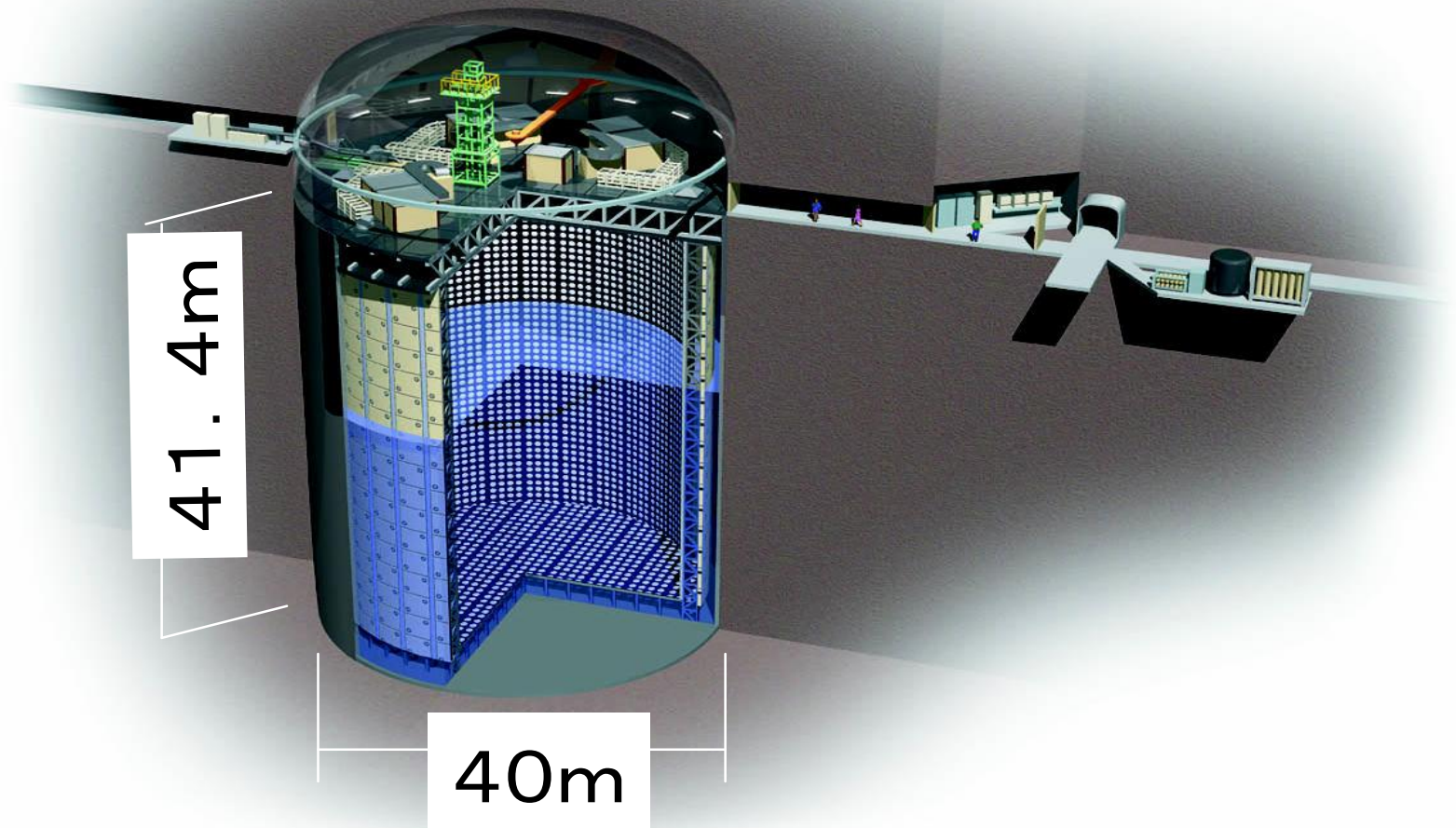
Mark Vagins

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Prospects of Neutrino Physics

Kavli IPMU, Kashiwa

April 8, 2019

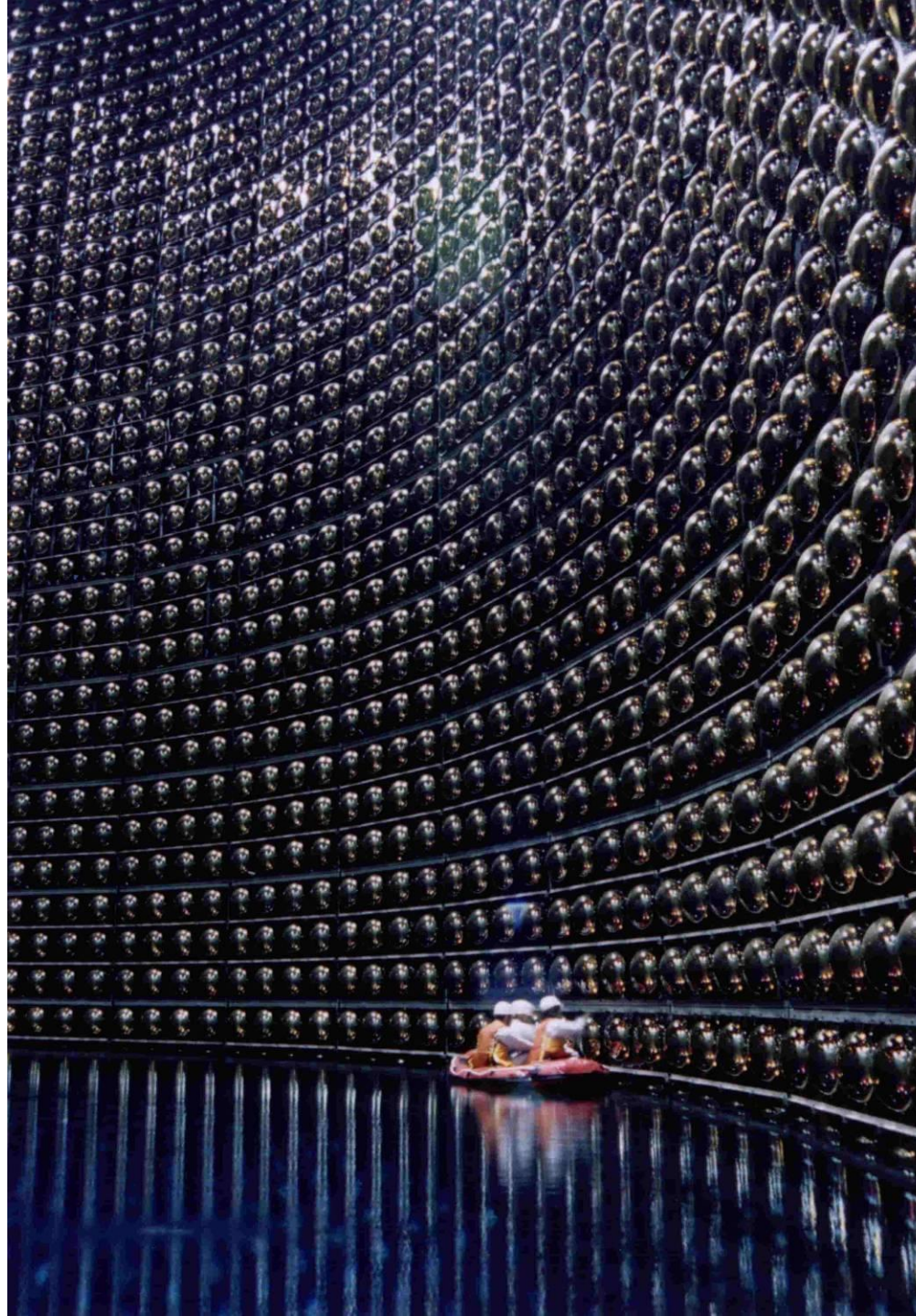


The Super-Kamiokande neutrino detector, in Mozumi, Japan.

50,000 tons
of ultra-pure
 H_2O

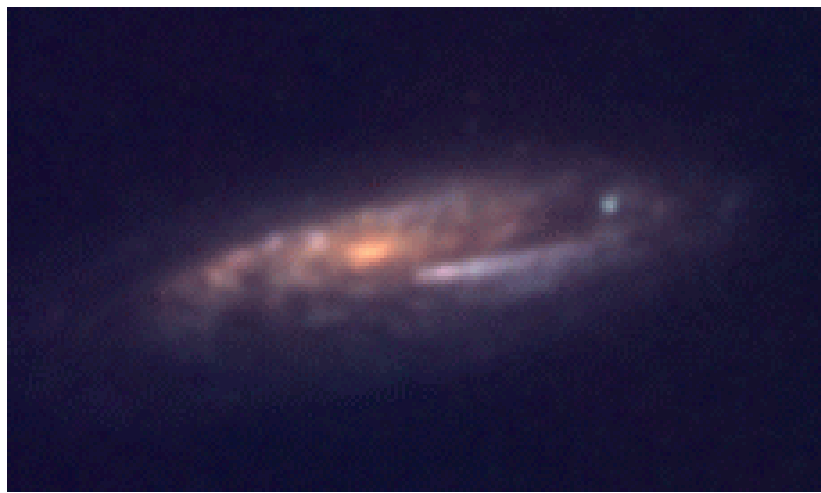
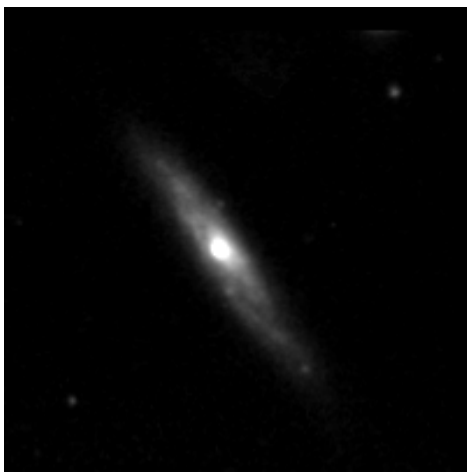
13,000
light
detectors

One kilometer
underground



Since 1996,
observes
particles
from the
sun and
cosmic rays

Also looks
for proton
decay
and
supernova
neutrinos



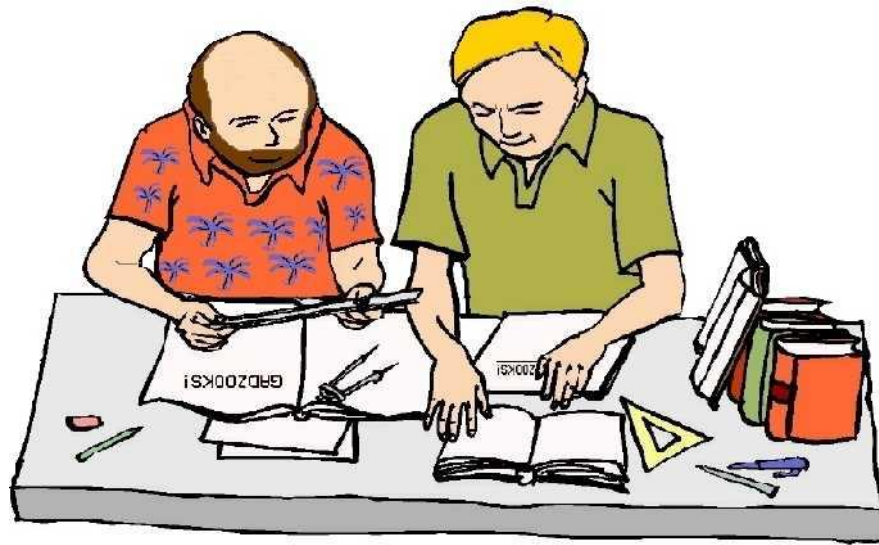
Here's how most
supernovas look
to us
(video is looped).

There is about one SN
explosion per second in
the universe as a whole!



These produce the as-
yet unobserved diffuse
supernova neutrino
background [DSNB],
also known as the
supernova relic
neutrinos [SRN].





Motivated by detecting the DSNB, theorist John Beacom and I wrote the original **GADZOOKS!**

(**G**adolinium **A**ntineutrino **D**etector **Z**ealously
Outperforming **O**ld **K**amiokande, **S**uper!) paper.

It proposed loading big WC detectors, specifically Super-K, with water soluble gadolinium, and evaluated the physics potential and backgrounds of a giant antineutrino detector.

[Beacom and Vagins, *Phys. Rev. Lett.*, **93**:171101, 2004]

(359 citations → one every 15 days for fifteen years)



Reactor antineutrino oscillations and gadolinium loaded Super-Kamiokande detector

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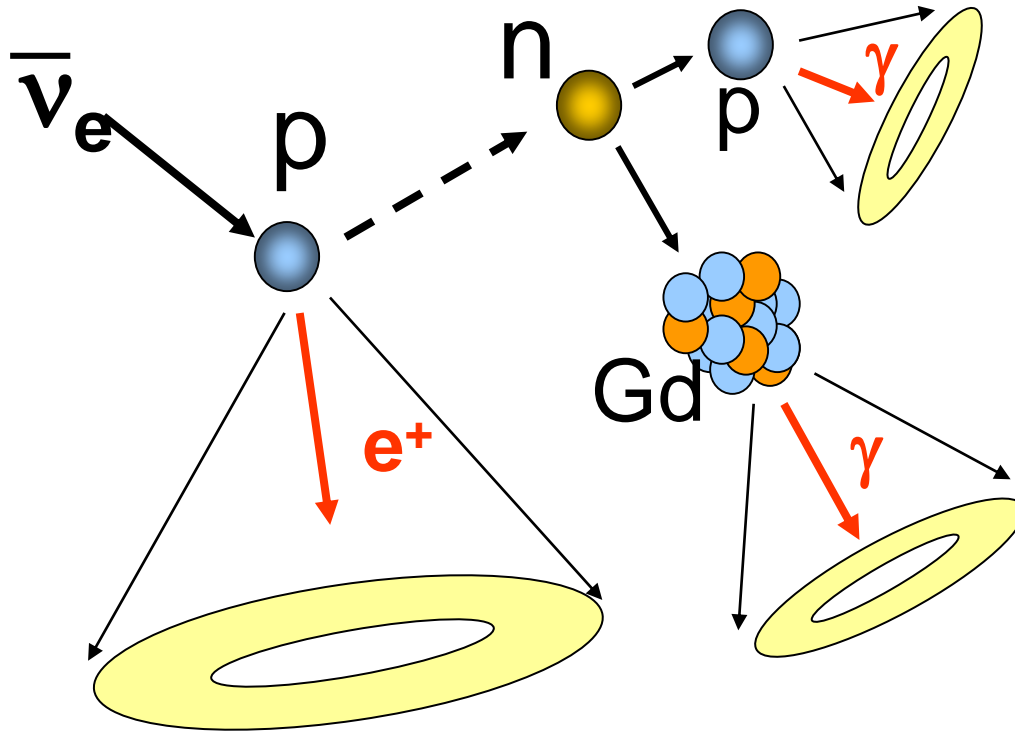
Editor: G.F. Giudice

Abstract

We explore the potential of measuring the solar neutrino oscillation parameters in the proposed gadolinium loaded Super-Kamiokande (SK-Gd) detector. Gadolinium dissolved in water can detect neutrons much more efficiently than pure water. This imparts the detector the ability to observe electron type antineutrinos, transforming Super-Kamiokande into a huge reactor antineutrino detector with an event rate approximately 43 times higher than that observed in KamLAND. We simulate the reactor antineutrino data expected in this high statistics detector. We use these prospective data to study the precision with which the solar neutrino oscillation parameters, Δm_{\odot}^2 and $\sin^2 \theta_{\odot}$, can be determined (i) with the SK-Gd detector, and (ii) by combining the SK-Gd data with the global data on solar neutrino oscillations. For comparison and completeness the allowed regions of Δm_{\odot}^2 and $\sin^2 \theta_{\odot}$ expected to be obtained from the data of the solar neutrino and KamLAND experiments, are also presented. We find that the SK-Gd experiment could provide one of the most precise (if not the most precise) determination of the solar neutrino oscillation parameters Δm_{\odot}^2 and $\sin^2 \theta_{\odot}$.

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Inverse Beta Decay with Gadolinium



Positron and gamma ray
vertices are within ~50 cm.

$\bar{\nu}_e$ can be identified by delayed coincidence.

Possibility 1: 10% or less

$n+p \rightarrow d + \gamma$
2.2 MeV γ -ray

Possibility 2: 90% or more

$n+\text{Gd} \rightarrow \sim 8\text{MeV } \gamma$
 $\Delta T = \sim 30 \text{ } \mu\text{sec}$

Neutron Captures on Gd vs. Concentration

Captures on Gd

100%
80%
60%
40%
20%
0%

0.1% Gd gives
90% efficiency
for n capture
on Gd; the
remaining 10%
capture on H

In the 50-kton
Super-Kamiokande
this means ~100
tons of water
soluble
 $\text{Gd}_2(\text{SO}_4)_3$

0.01% Gd gives
50% efficiency
for n capture
on Gd; the
remaining 50%
capture on H

Thermal
neutron
capture
cross
section
(barns)

Gd = 49700

S = 0.53

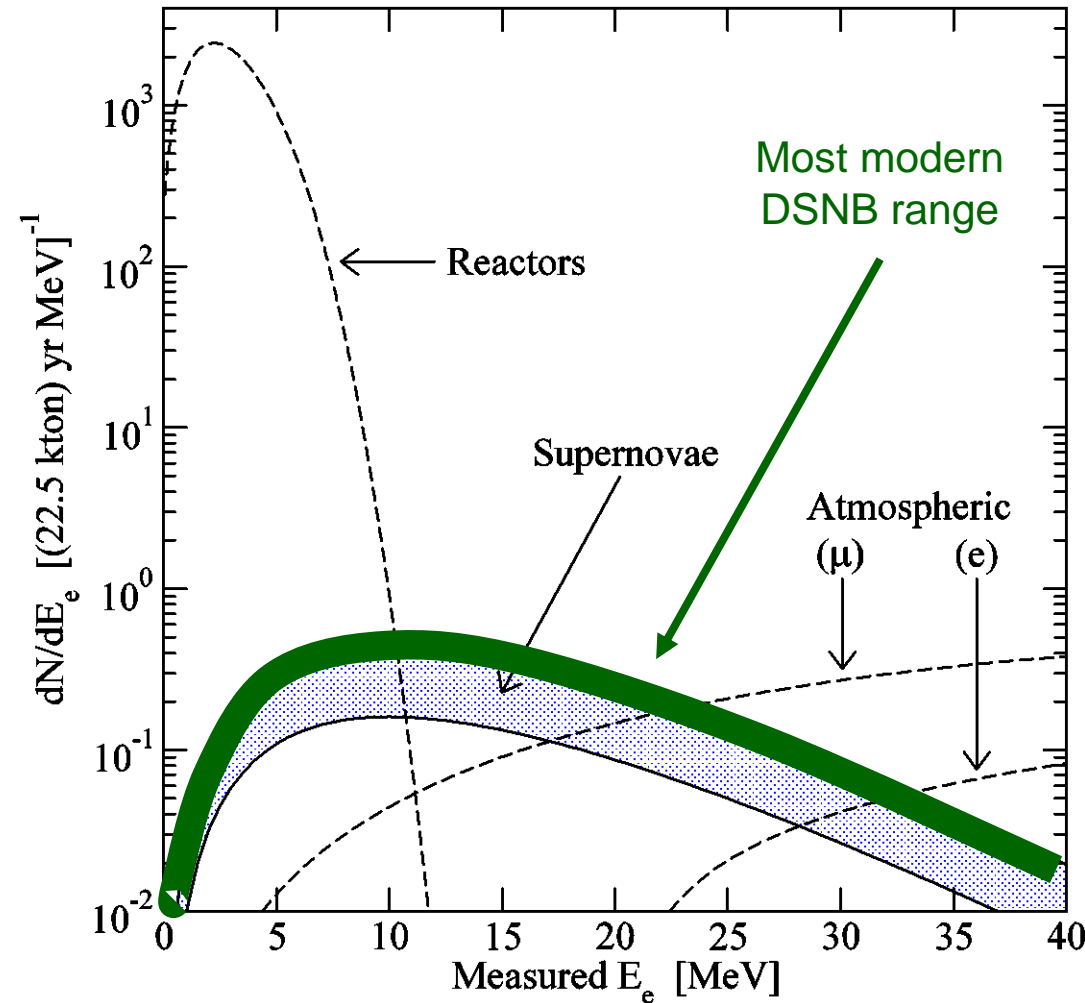
H = 0.33

O = 0.0002

0.0001% 0.001% 0.01% 0.1% 1%

Gd in
Water

Here's what the coincident signals in Super-K
with $\text{Gd}_2(\text{SO}_4)_3$ will look like
(energy resolution is applied):



spatial and
temporal separation
between prompt e^+
Cherenkov light and
delayed Gd neutron
capture gamma
cascade:

$$\lambda \sim 4 \text{ cm}, \tau \sim 30 \mu\text{s}$$

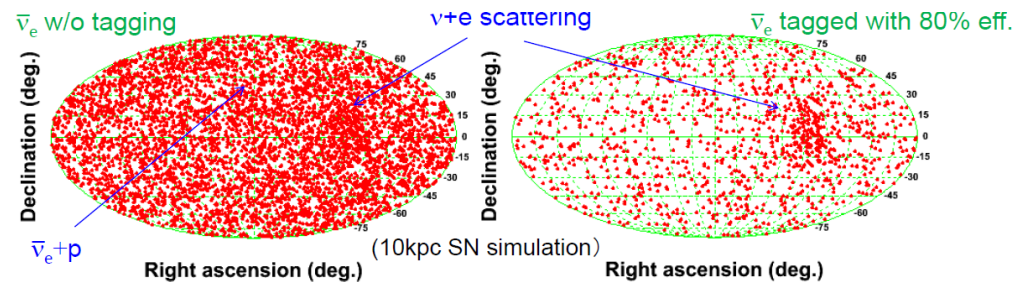
→ A few clean events/yr
in Super-K with Gd

In the case of a galactic supernova, having $\text{Gd}_2(\text{SO}_4)_3$ in Super-K will provide many important benefits:

- Allows the exact $\bar{\nu}_e$ flux, energy spectrum, and time profile to be determined via the extraction of a tagged, pure sample of inverse beta events.

- Instantly identifies a burst as genuine via “Gd heartbeat”.

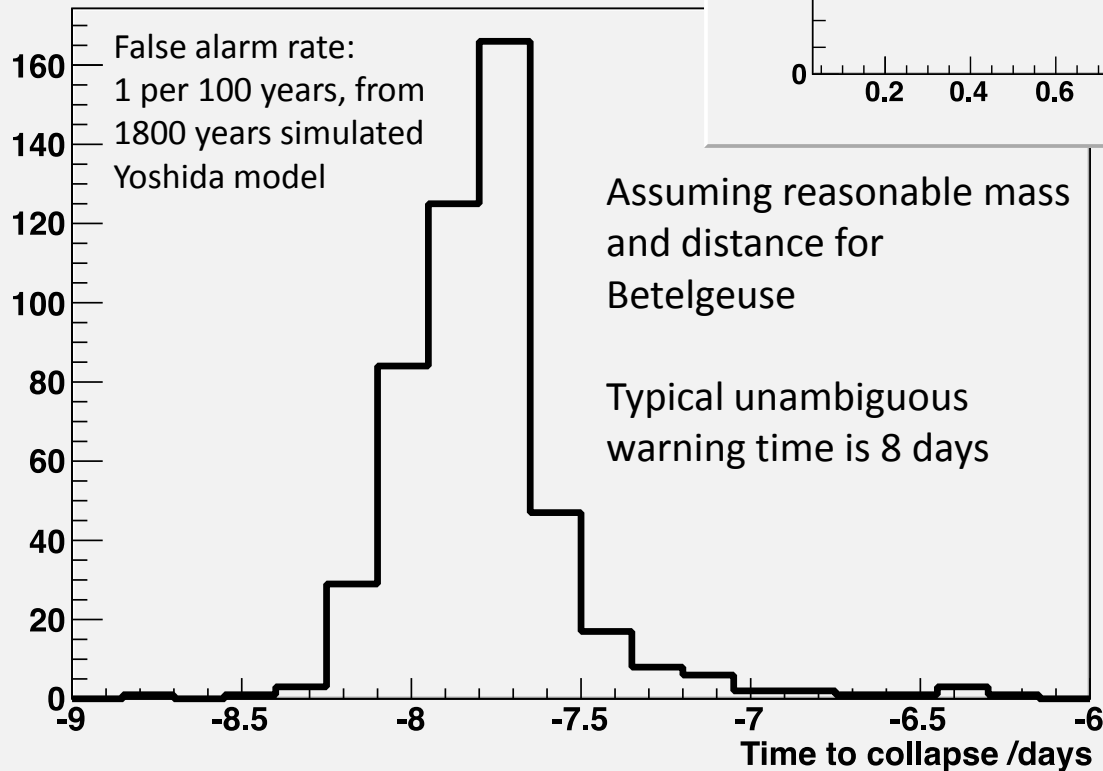
- Doubles the ES pointing accuracy. Error circle cut by 75%.



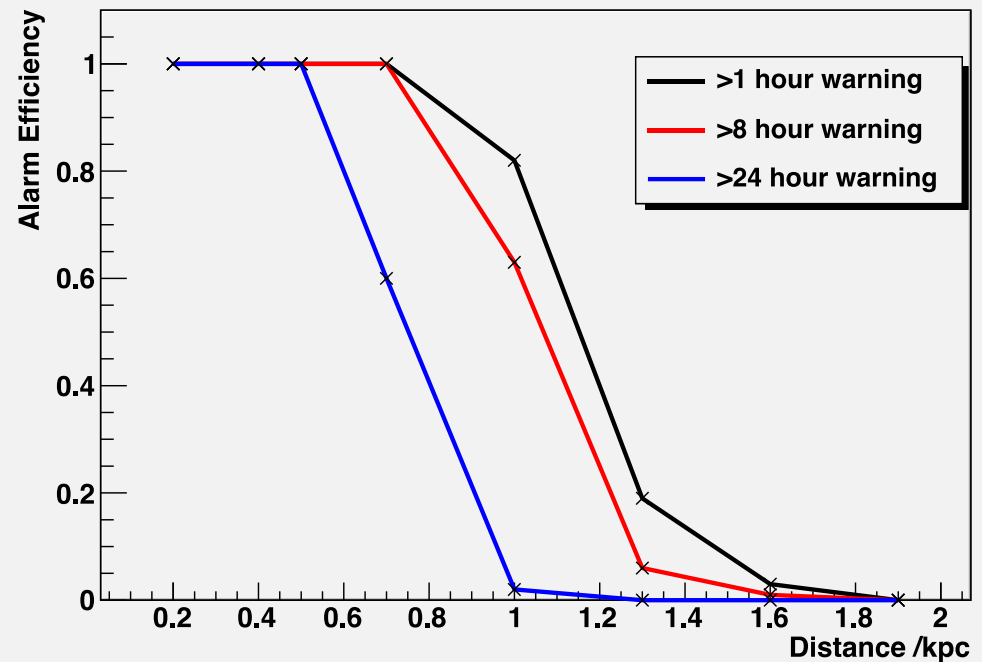
- Helps to identify the other neutrino signals, especially the weak neutronization burst of ν_e .
- Enables a search for very late time black hole formation.
- Provides for very early warning of the most spectacular, nearby explosions so we can be sure not to miss them.

Gd-loaded Super-Kamiokande's Sensitivity to pre-SN ν 's (Super-K internal study)

Warning times for $12M_{\odot}$ at 0.2kpc



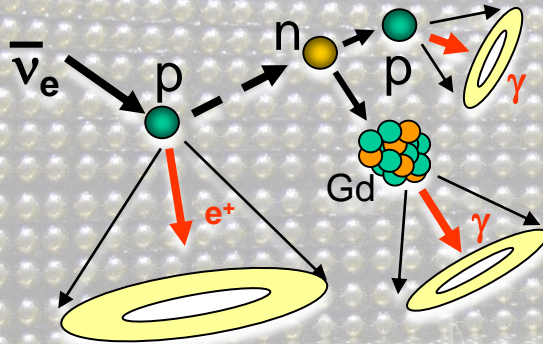
Alarm efficiencies against distance, 1 false per 100 years



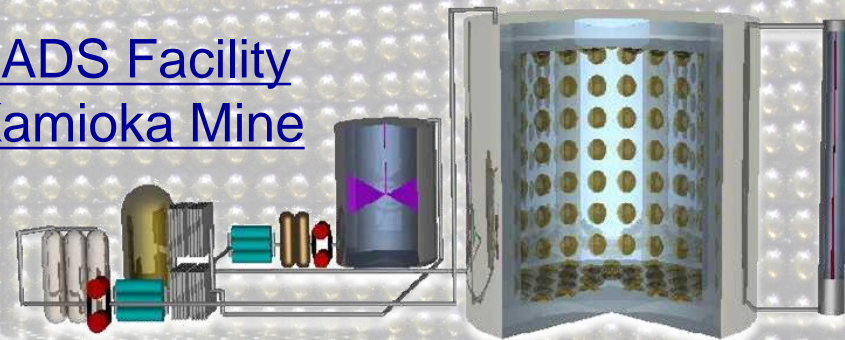
[C. Simpson]

EGADS → Gd-loaded Super-K

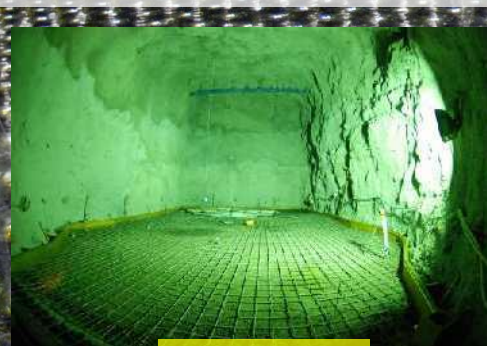
Adding **water soluble gadolinium** to **Super-K** will greatly enhance its ability to detect **supernova neutrinos** (and help with many other physics topics like **proton decay**). **EGADS** is a dedicated gadolinium demonstrator which includes a working 200 ton scale model of SK.



EGADS Facility
in Kamioka Mine



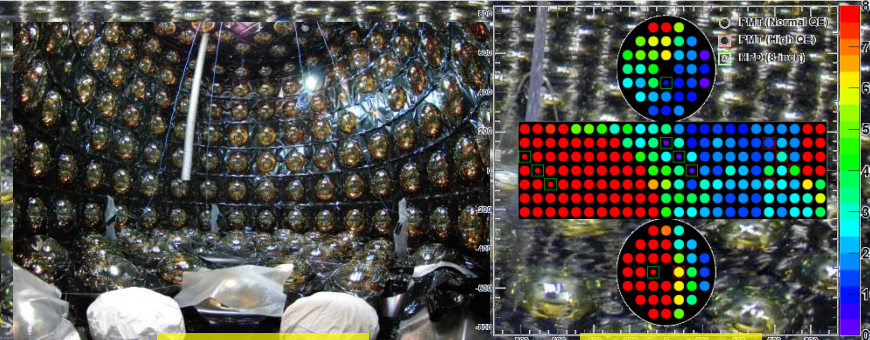
Beacom and Vagins, *Phys. Rev. Lett.*, 93:171101, 2004 [359 citations]



12/2009



11/2011



8/2013

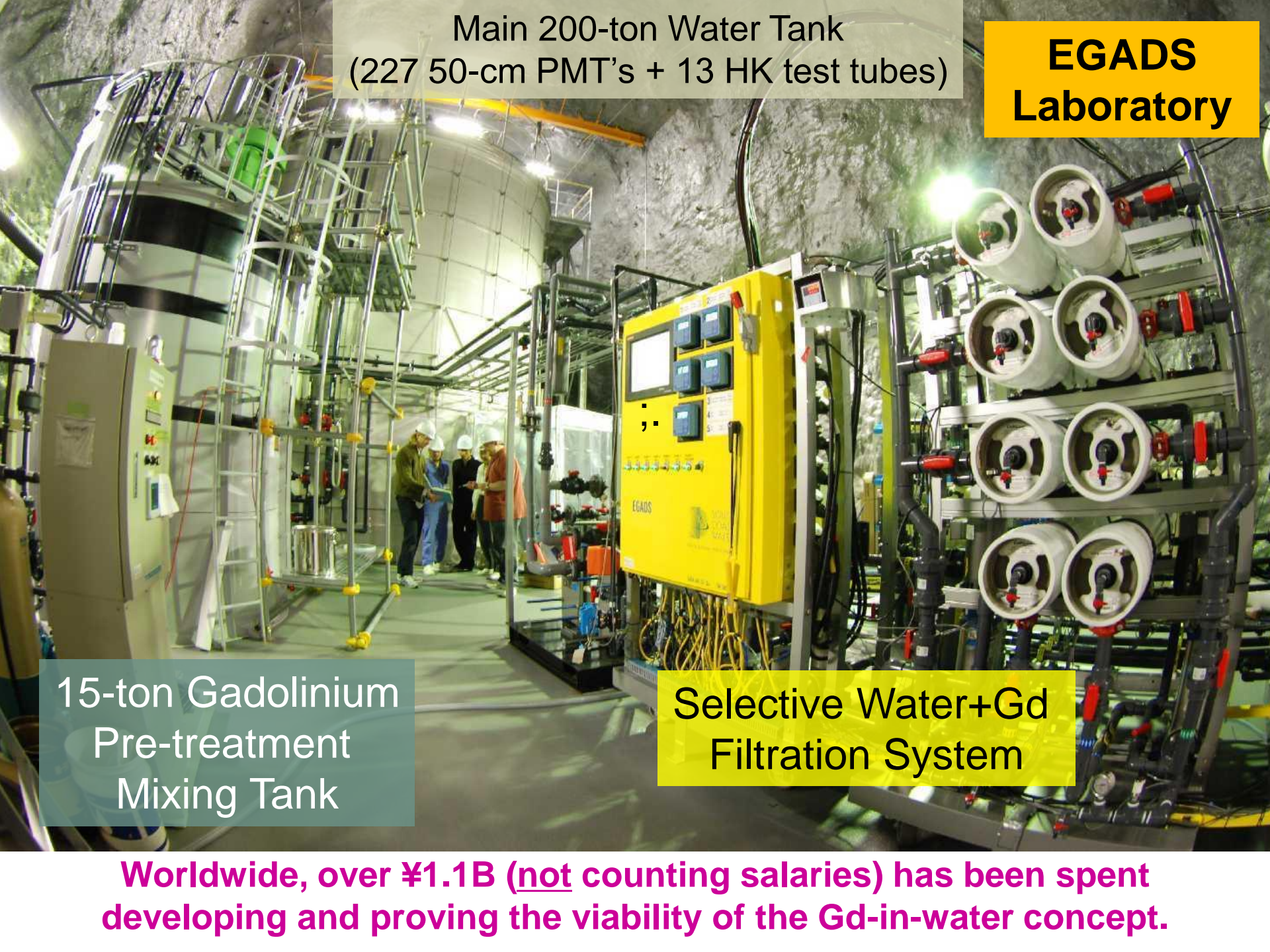
6/2015



Maintaining good water quality in the presence of dissolved gadolinium required the development of an entirely new technology: true selective filtration. I call my resulting system a “molecular band-pass” filter.

This electro-mechanical system continuously circulates the Gd-loaded water and removes every impurity *except* $\text{Gd}_2(\text{SO}_4)_3$.



The image shows a large, dimly lit underground laboratory. On the left is a massive, cylindrical white tank with a complex metal scaffolding and ladders around it. In the center, a group of four people wearing hard hats and safety gear are standing and looking at a document. To the right of the people is a large yellow control cabinet with a screen and several buttons. Further right is a complex filtration system with multiple white cylindrical filters mounted on a metal frame. The background shows the rough, rocky walls of the underground facility.

Main 200-ton Water Tank
(227 50-cm PMT's + 13 HK test tubes)

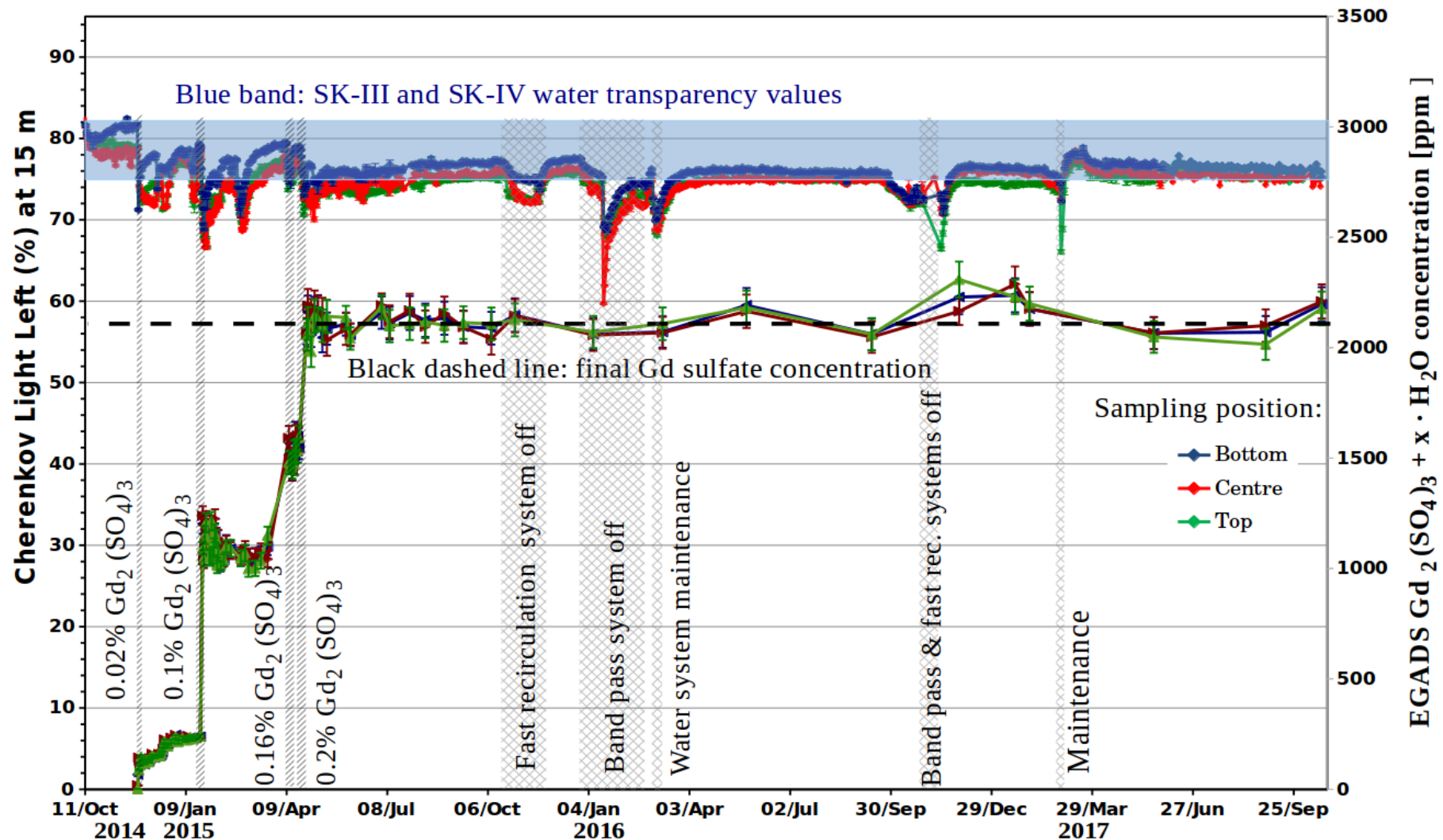
**EGADS
Laboratory**

15-ton Gadolinium
Pre-treatment
Mixing Tank

Selective Water+Gd
Filtration System

Worldwide, over ¥1.1B (not counting salaries) has been spent developing and proving the viability of the Gd-in-water concept.

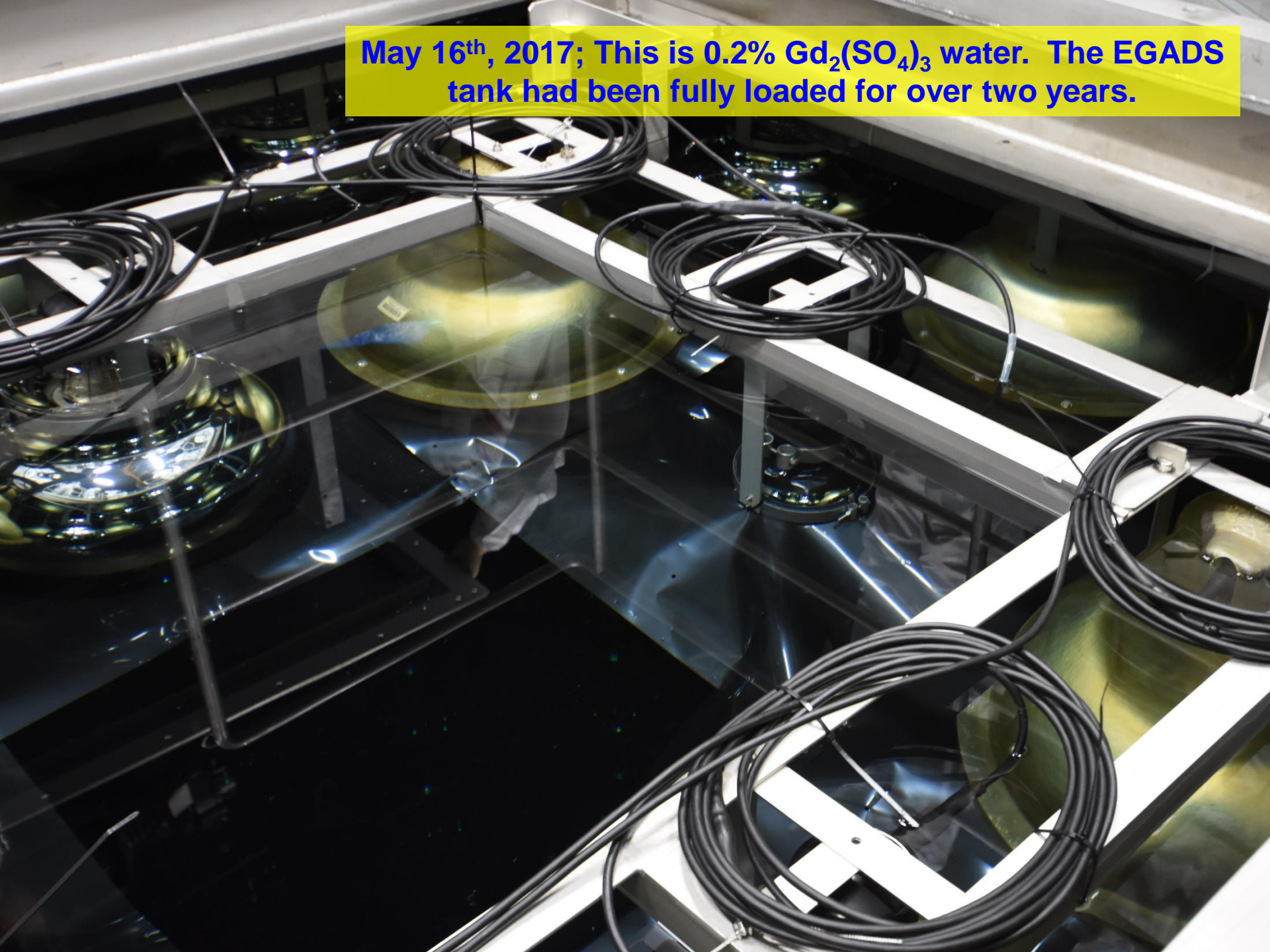
Light @ 15 meters and Gd conc. in the 200-ton EGADS tank



After two and a half years at full Gd loading, during stable operations EGADS water transparency remains within the SK ultrapure range.

→ No detectable loss of Gd after more than 650 complete turnovers. ←

May 16th, 2017; This is 0.2% $\text{Gd}_2(\text{SO}_4)_3$ water. The EGADS tank had been fully loaded for over two years.





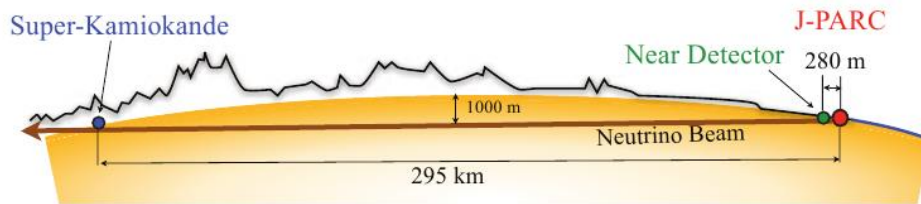
November 6th, 2017; This view is directed up the side wall from the bottom of the 200-ton tank. Looks great after 2.5 years of exposure to 0.2% $\text{Gd}_2(\text{SO}_4)_3$ water!

After years of testing and study
– culminating in these powerful EGADS results –
no technical showstoppers had been encountered. And so...

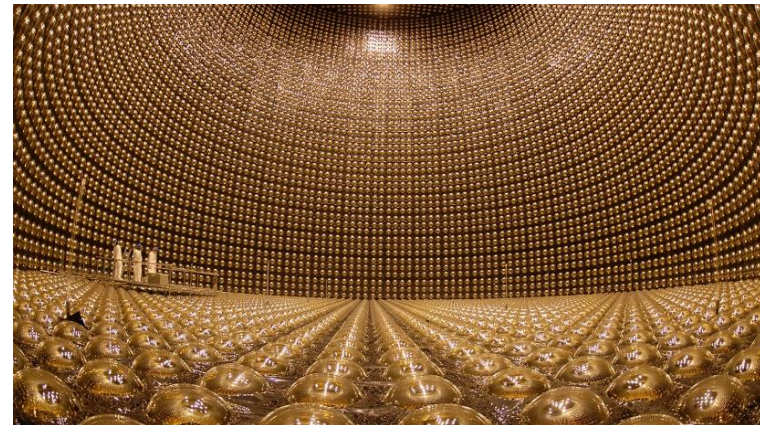
June 27, 2015: The Super-Kamiokande Collaboration approved the addition of gadolinium to the detector, pending discussions with T2K.



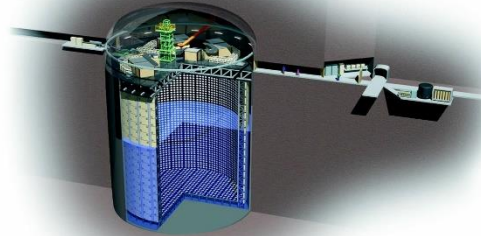
January 30, 2016: The T2K Collaboration approved addition of gadolinium to Super-Kamiokande, with the precise timing to be jointly determined based on the needs of both projects.



July 26, 2017: The official start time of draining the SK tank to prepare for Gd loading was decided → June 1, 2018.



Super-Kamiokande



Original Super-K Water System

New Gadolinium Water System "Hall G"
(4000 m³)

EGADS Hall
(2500 m³)

Gadolinium Removal
System "Hall H"
(1000 m³)

100m

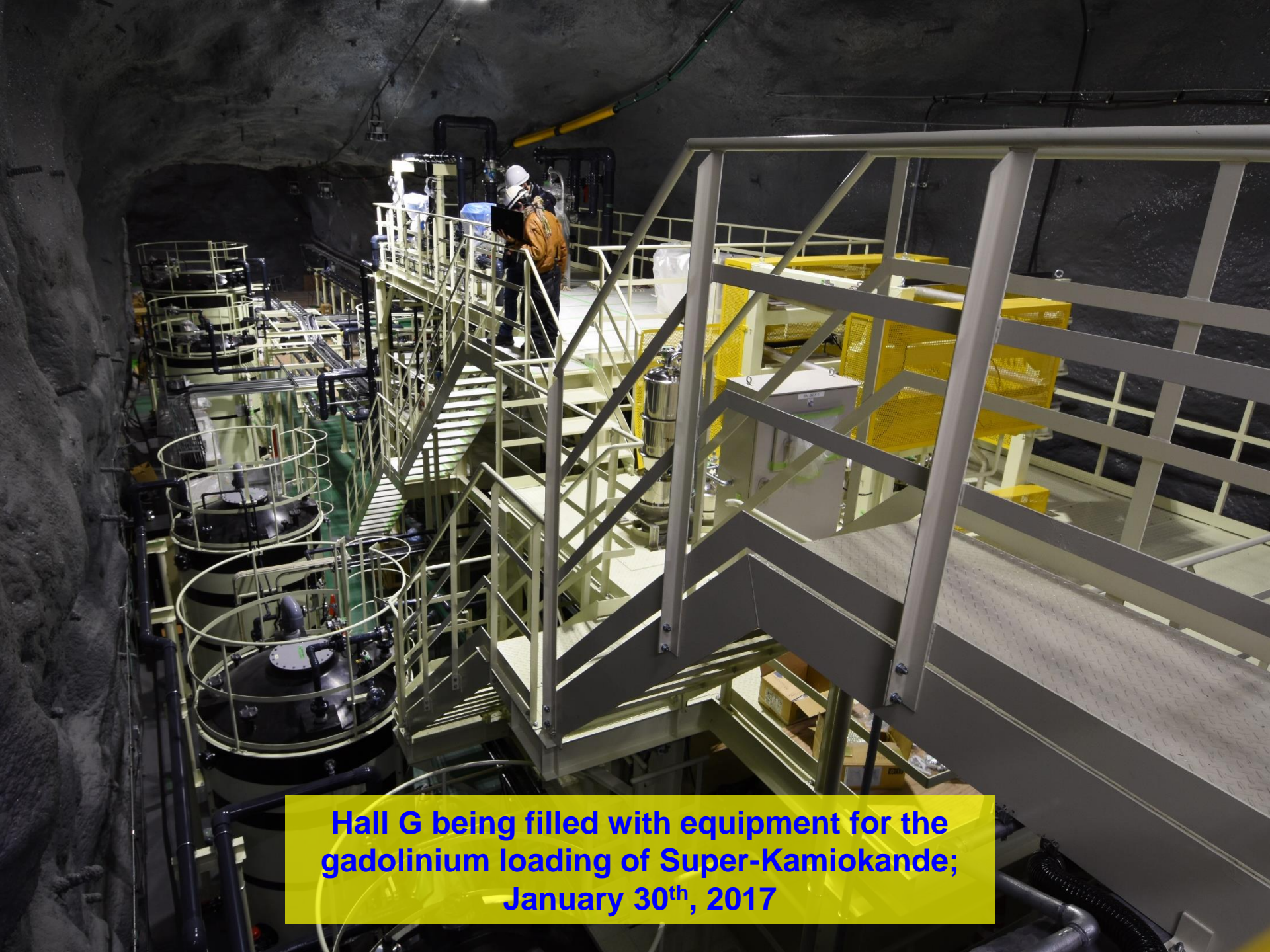
The Kamioka Observatory in the Mozumi Mine



**New gadolinium water system hall (“Hall G”);
September 10st, 2015**



**Hall G ready for occupancy;
April 22nd, 2016**



**Hall G being filled with equipment for the
gadolinium loading of Super-Kamiokande;
January 30th, 2017**

**Prior to Gd loading we must be prepared
to completely remove and capture the Gd
→ New system needed**



In Hall H; March 24th, 2018



In Hall H; March 30th, 2018



**Completed gadolinium removal system (62 tons of ion exchange resin)
in Hall H; April 1st, 2018**



Main jobs to get ready for Gd loading:

1) Fix SK leak

2) Clean up interior

3) Replace dead PMTs

4) Augment internal plumbing

Entering Super-K for the first time since 2006; June 1st, 2018

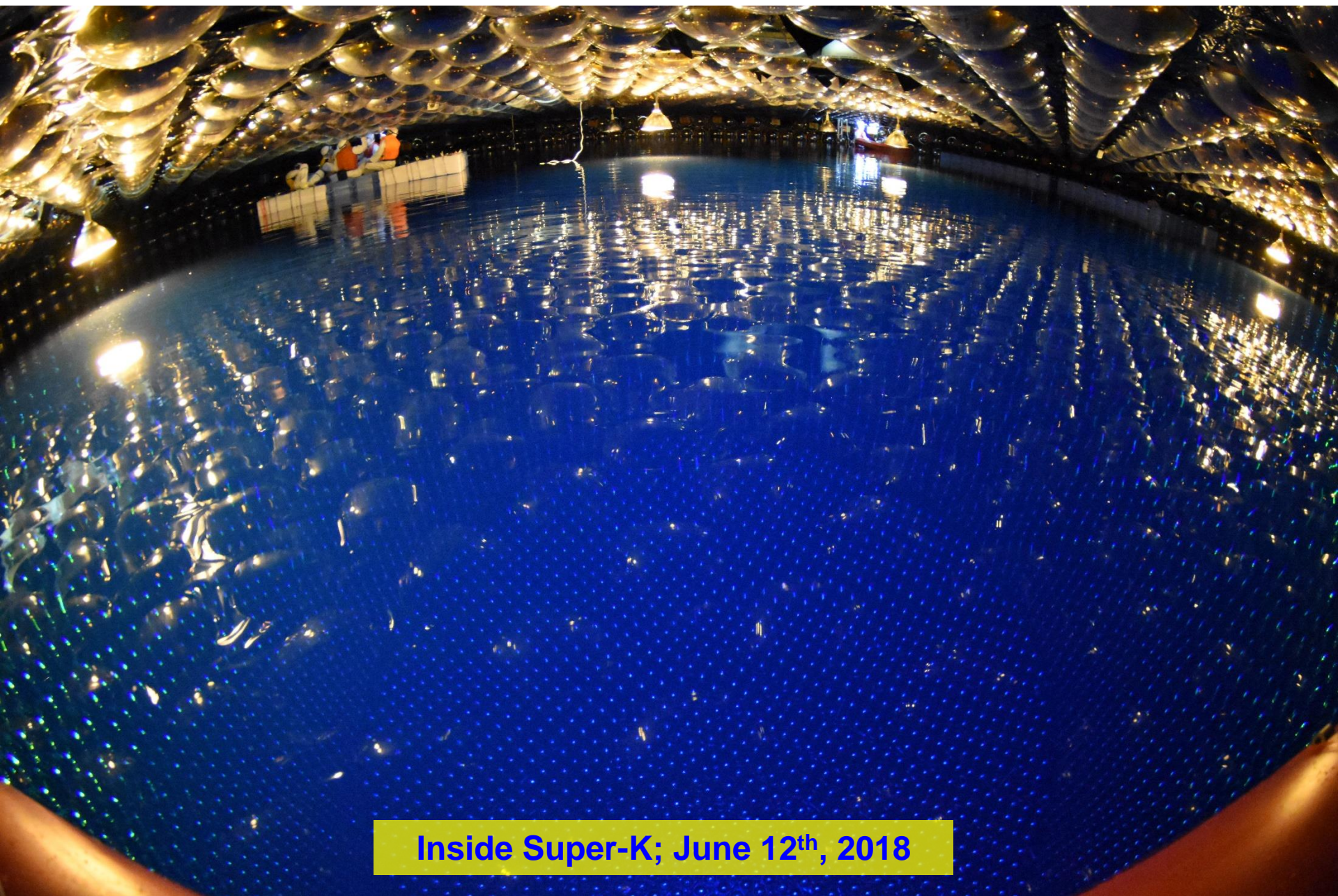


**From March 2018 → October 2018,
2683 person-days of work were required!**

Inside Super-K veto region (top); June 6th, 2018



Super-K veto region (side) with floating floor; June 6th, 2018

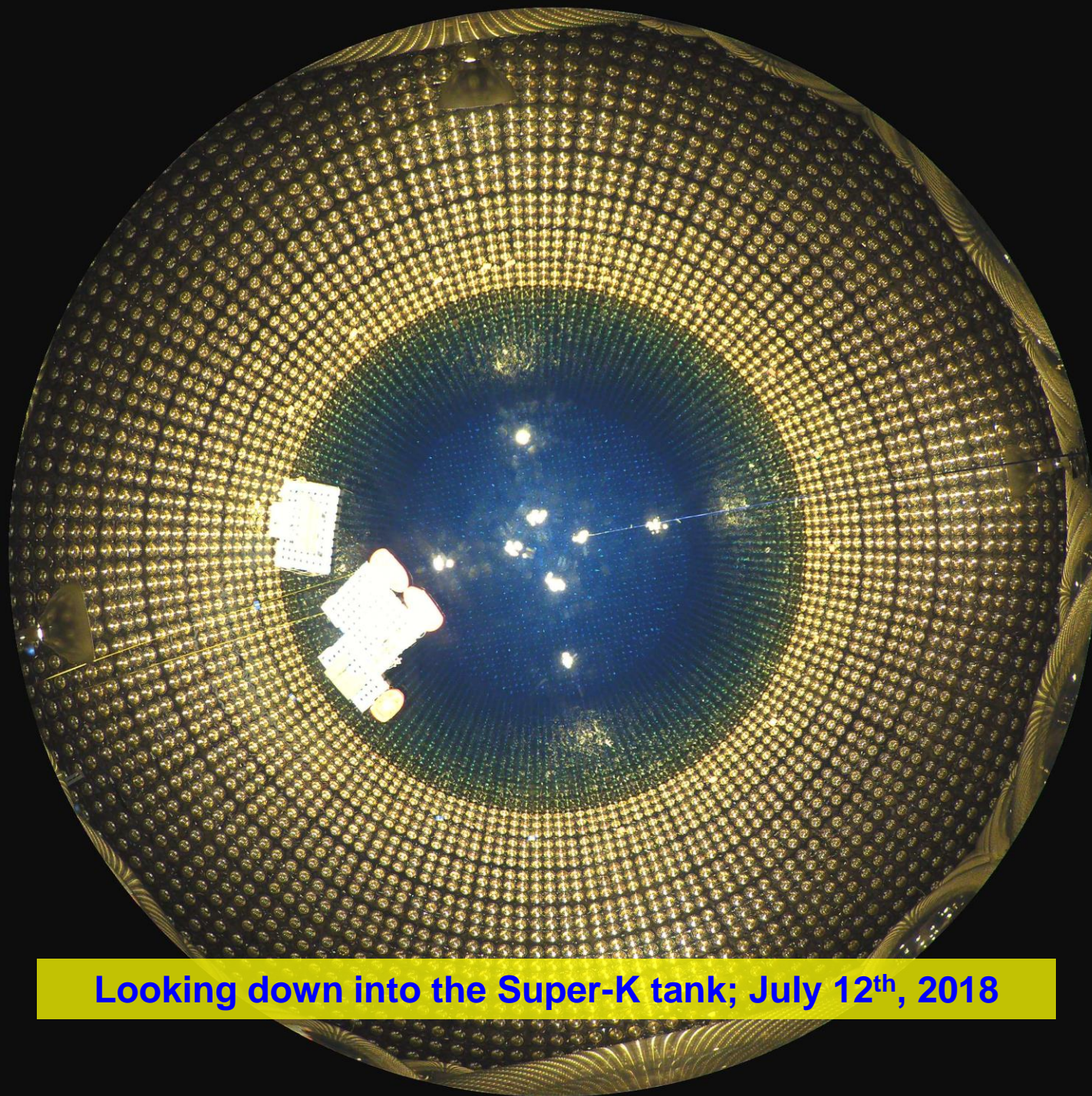


Inside Super-K; June 12th, 2018

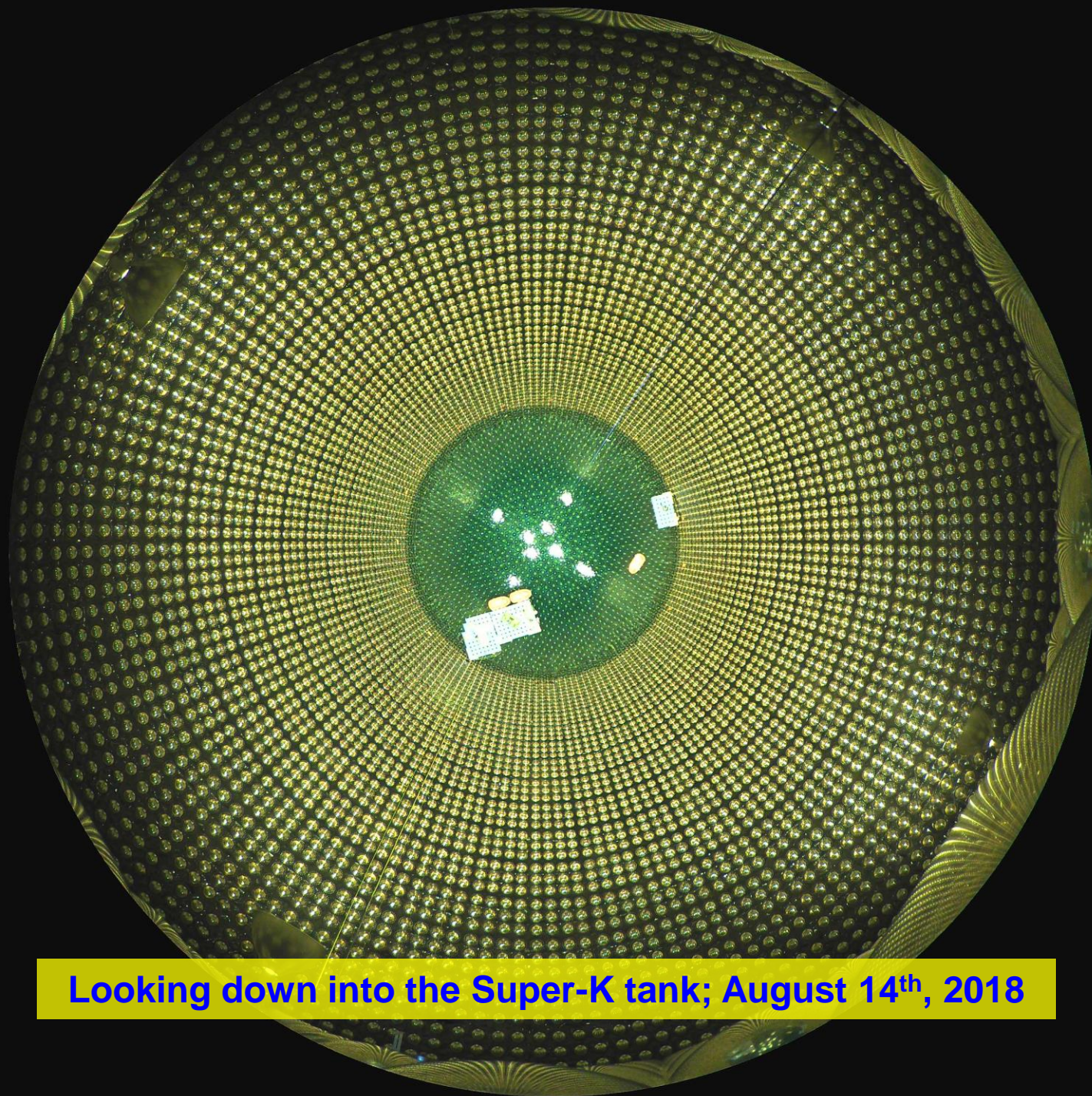
A photograph of a worker in a white protective suit and blue helmet applying sealant to a wall in a large underground cavern. The worker is holding a blue bucket and a spray gun. The wall is covered with a white plastic sheeting secured by blue tape. In the background, another worker is visible. The floor is covered with a white plastic sheeting. A yellow and black striped caution tape is on the floor. A yellow and black striped caution tape is on the floor. A yellow and black striped caution tape is on the floor.

Applying special low-background
MineGuard sealant

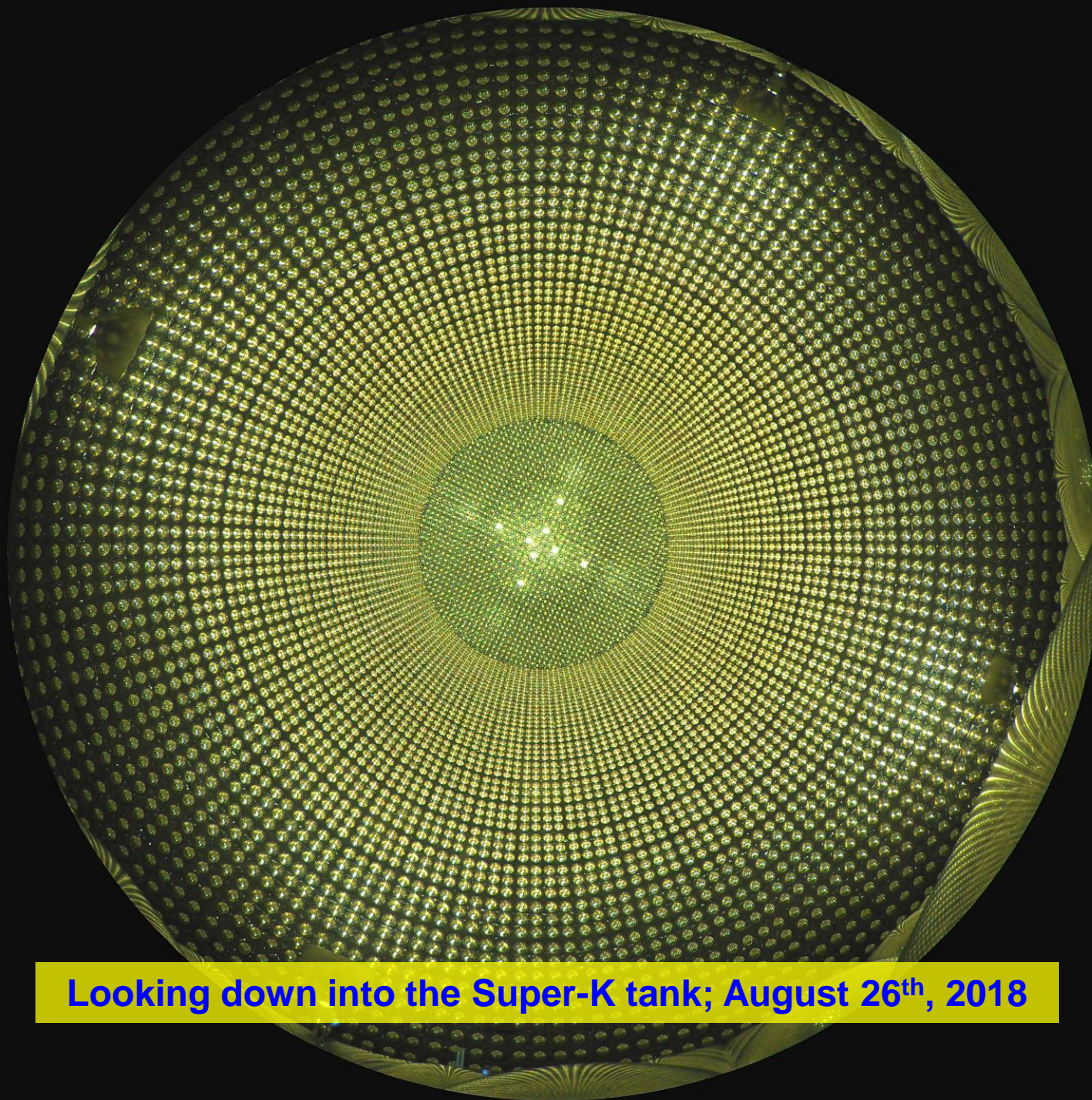
Super-K veto region (side) with floating floor; June 23rd, 2018



Looking down into the Super-K tank; July 12th, 2018



Looking down into the Super-K tank; August 14th, 2018



Following
~3000 person-
days of
refurbishment
work, as of
Feb. 2019 the
detector is now
refilled with
pure water and
taking data,
ready for the
addition of
gadolinium!

Looking down into the Super-K tank; August 26th, 2018

“Currently we do not observe any water leakage from the SK tank within the accuracy of our measurement, which is less than 0.017 tons per day. This is less than 1/200th of the leak rate observed before the 2018/2019 tank refurbishment.”



*Leak sealing work
is a success!*

Developing special low-background $\text{Gd}_2(\text{SO}_4)_3$

Radioactive chain	Part of the chain	mBq/kg
^{238}U	^{238}U	50
	^{226}Ra	5
^{232}Th	^{228}Ra	10
	^{228}Th	100
^{235}U	^{235}U	32
	$^{227}\text{Ac} / ^{227}\text{Th}$	300

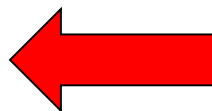
Radio isotopes in
“typical” off-the-shelf
gadolinium sulfate

We need
from 1-4
orders of
magnitude
reduction
in RI

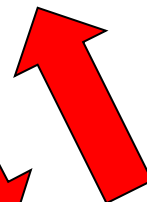
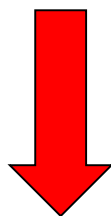


What the
physics
requires

Radioactive chain	Part of the chain	SRN (mBq/kg)	Solar ν (mBq/kg)
^{238}U	^{238}U	< 5	-
	^{226}Ra	-	< 0.5
^{232}Th	^{228}Ra	-	< 0.05
	^{228}Th	-	< 0.05
^{235}U	^{235}U	-	< 3
	$^{227}\text{Ac} / ^{227}\text{Th}$	-	< 3



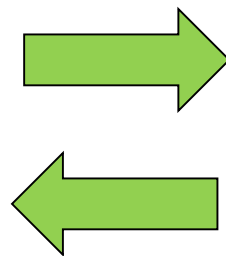
Low background
Ge counters
at Canfranc
underground
laboratory
(Spain)



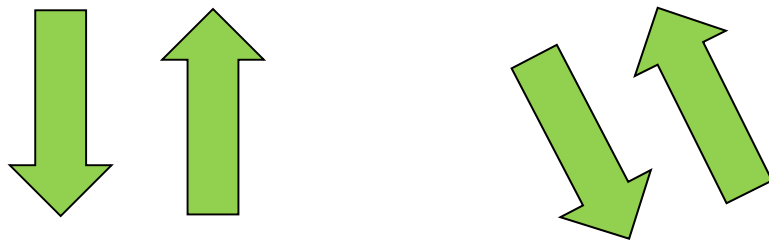
Kamioka Ge
counter (Japan)

Boulby Underground Germanium Suite (UK)





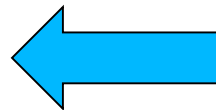
Low background
Ge counters
at Canfranc
underground
laboratory
(Spain)



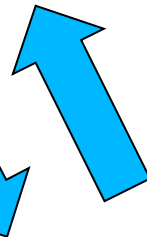
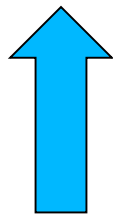
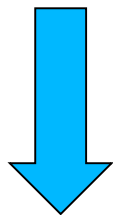
Boulby Underground Germanium Suite (UK)



Kamioka Ge
counter (Japan)



Low background
Ge counters
at Canfranc
underground
laboratory
(Spain)



Kamioka Ge
counter (Japan)

Boulby Underground Germanium Suite (UK)



After this iterative purification campaign, one company was able to produce gadolinium sulfate with the required radiopurity.

We recently ordered three 0.5 ton batches from them and tested samples from these batches at Kamioka to be sure everything was okay:

Goal (mBq/kg)		Batch 1	Batch 2	Batch 3
238U < 5	ICPMS	< 0.02	0.02	0.04
	Ge	<9.45	<9.89	<28.4
232Th < 0.05	ICPMS	0.04	0.02	0.04
	Ge	<0.20	<0.21	0.16
226Ra < 0.5	Ge	0.46 ± 0.24	<0.33	<0.20

In 2019 we will purchase ~13 tons of ultrapure gadolinium sulfate, enough for the first step in loading Super-Kamiokande.

→ *It will be the largest order of gadolinium in human history!* ←

Expected timeline for SK-Gd



Schedule
Approved



Install New SK
Water Systems, Computing, Calibration



SK In-Tank Upgrade Work



SK Pure Water Running



SK Running with 0.01% Gd (50% eff.)



Increased Loading, up to 0.1% Gd (90% eff.)



*We expect to have collected the
world's first diffuse supernova
neutrinos before 2022!*

