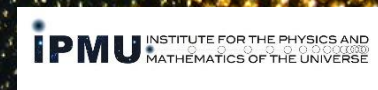


# SK RESULTS ON SOLAR NEUTRINO AND PROSPECTS WITH HK



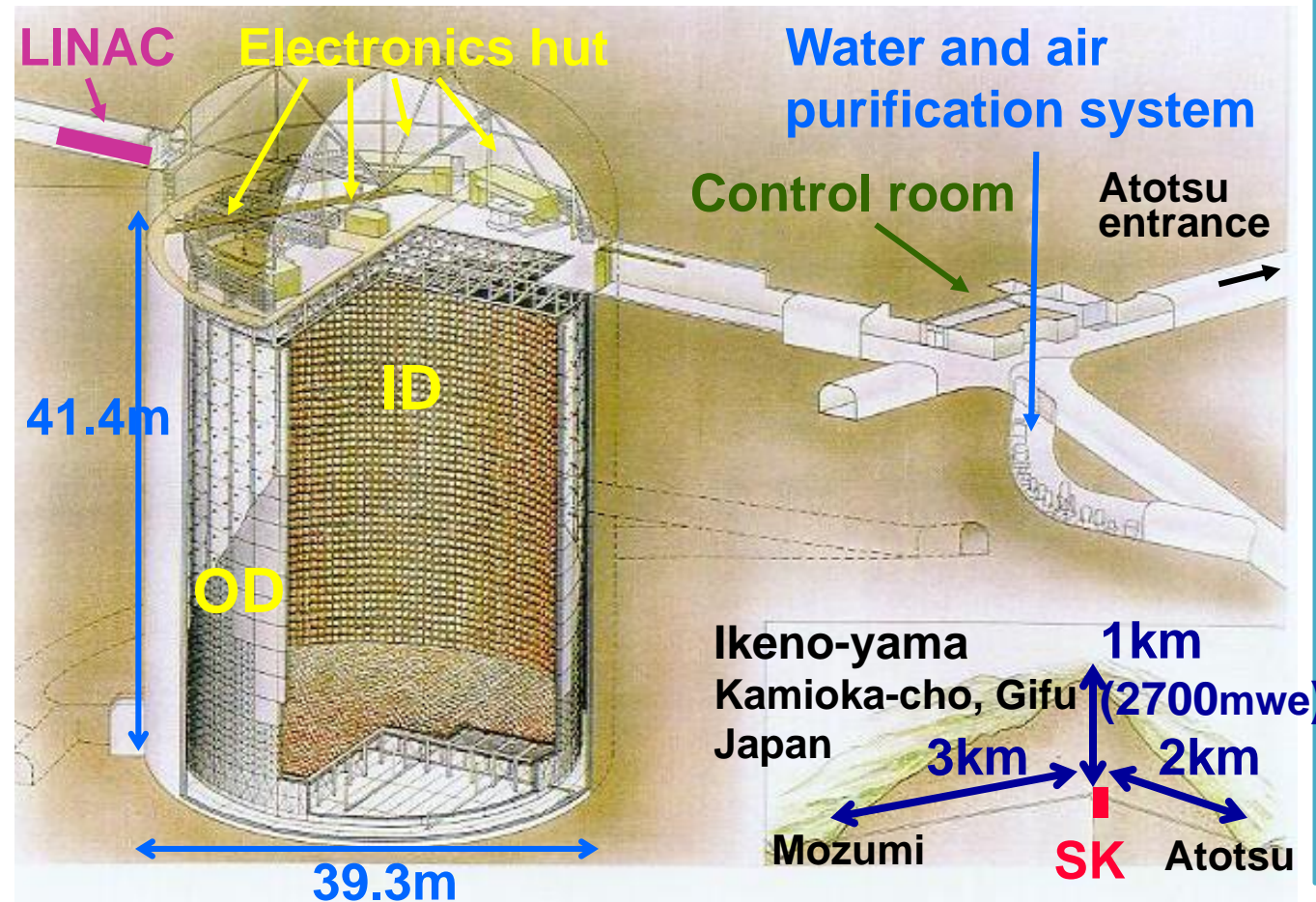
Yasuo Takeuchi  
(Kobe University / Kavli IPMU)

Inside of SK detector during  
refurbishment work (July 15, 2018)

# Outline

- **SK detector**
- **Solar neutrinos**
- **Recent results from SK**
- **Prospects with HK**
- **Summary**

# Super-Kamiokande detector



- 50 kton water
- ~2m OD viewed by 8-inch PMTs
- 32kt ID viewed by 20-inch PMTs
- 22.5kt fid. vol. (2m from wall)
- SK-I: April 1996~
- SK-V is running
- For Solar  $\nu$ :  

$$\nu_x + e^- \rightarrow \nu_x + e^-$$

Inner Detector (ID) PMT: ~11100 (SK-I,III,IV,V), ~5200 (SK-II)  
 Outer Detector (OD) PMT: 1885

# History & Plan of Super-Kamiokande



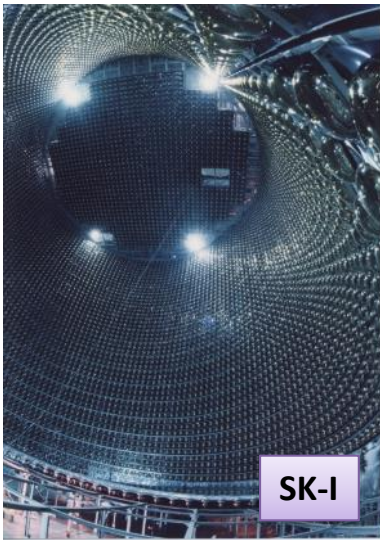
SK-I

SK-II

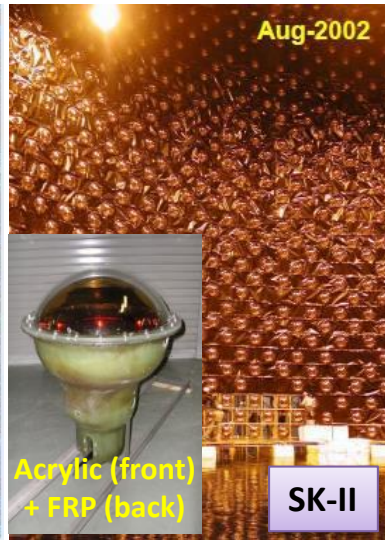
SK-III

SK-IV

SK-V →  
SK-Gd

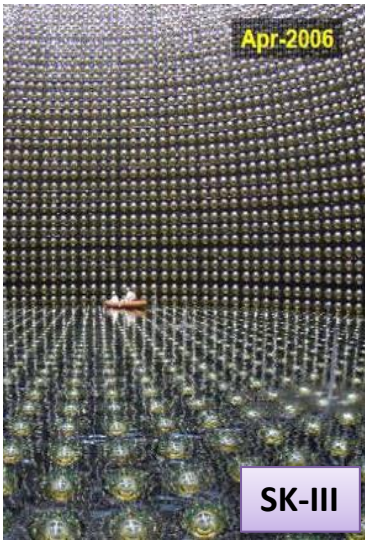


SK-I



Acrylic (front) + FRP (back)

SK-II



SK-III



SK-IV



Water system For SK-Gd

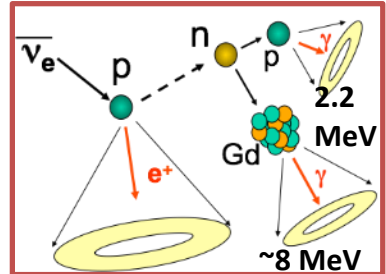
11146 ID PMTs  
(40% coverage)  
4.5 MeV  
1496 days

5182 ID PMTs  
(19% coverage)  
6.5 MeV  
791 days

11129 ID PMTs  
(40% coverage)  
4.5 MeV  
548 days

Electronics Upgrade  
3.5 MeV  
2860 days (~Dec. 2017)

Neutron tagging with Gd



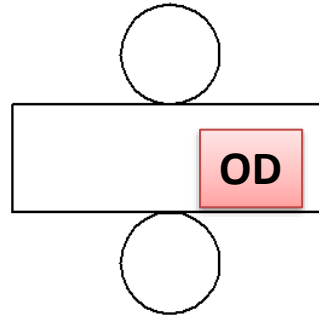
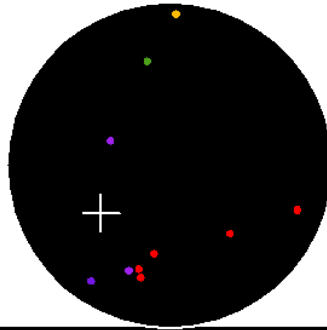
Current total: 5695 days

- Analysis energy threshold (recoil electron kinetic energy)
- Live time for solar neutrino analysis

# Typical low-energy event

## Super-Kamiokande

Run 1742 Event 102496  
 96-05-31:07:13:23  
 Inner: 103 hits, 123 pE  
 Outer: -1 hits, 0 pE (in-time)  
 Trigger ID: 0x03  
 E = 9.086 GDN=0.77 COSSUN= 0.949  
 Solar Neutrino

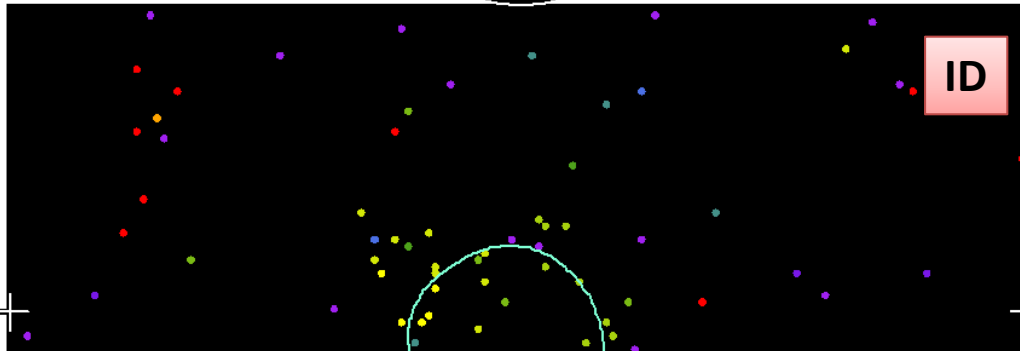


(for solar neutrinos)

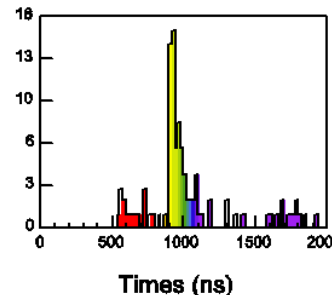
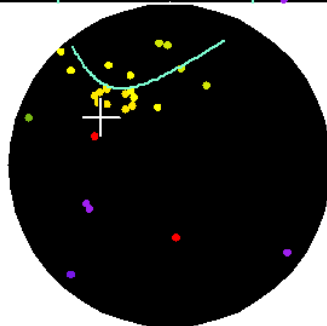
- Timing information
  - ➔ vertex position
- Ring pattern
  - ➔ direction
- Number of hit PMTs
  - ➔ energy

Time (ns)

- < 815
- 815- 835
- 835- 855
- 855- 875
- 875- 895
- 895- 915
- 915- 935
- 935- 955
- 955- 975
- 975- 995
- 995-1015
- 1015-1035
- 1035-1055
- 1055-1075
- 1075-1095
- >1095



(color: time)



$E_{e,\text{total}} = 9.1 \text{ MeV}$   
 $\cos\theta_{\text{sun}} = 0.95$

**~6 hit / MeV**  
 (SK-I, III, IV, V)

<b>Resolutions (for 10 MeV electrons)</b>			(software improvement)
<b>Energy: 14%</b>	<b>Vertex: 87cm</b>	<b>Direction: 26° SK-I</b>	
<b>Energy: 14%</b>	<b>Vertex: 55cm</b>	<b>Direction: 23° SK-III, IV, V</b>	

# Low-energy backgrounds in SK-I

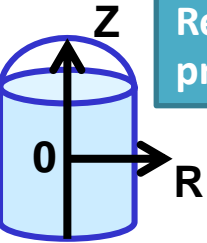


Misfit

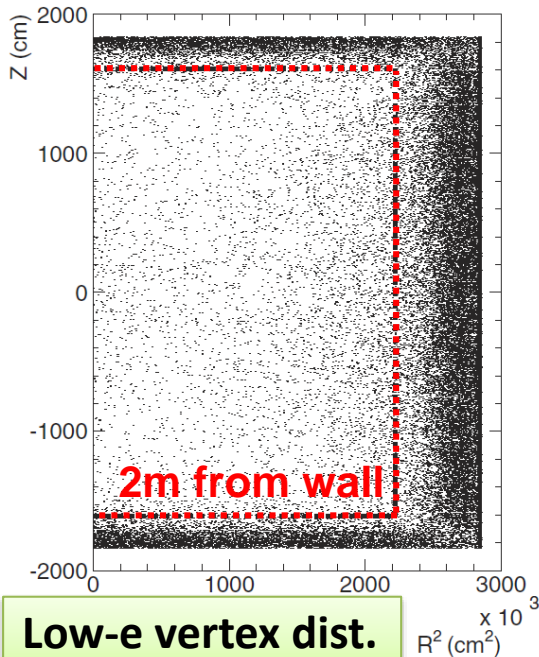
External gamma

Radon, Misfit,  
External gamma, etc.

Remaining spallation  
products (probably)

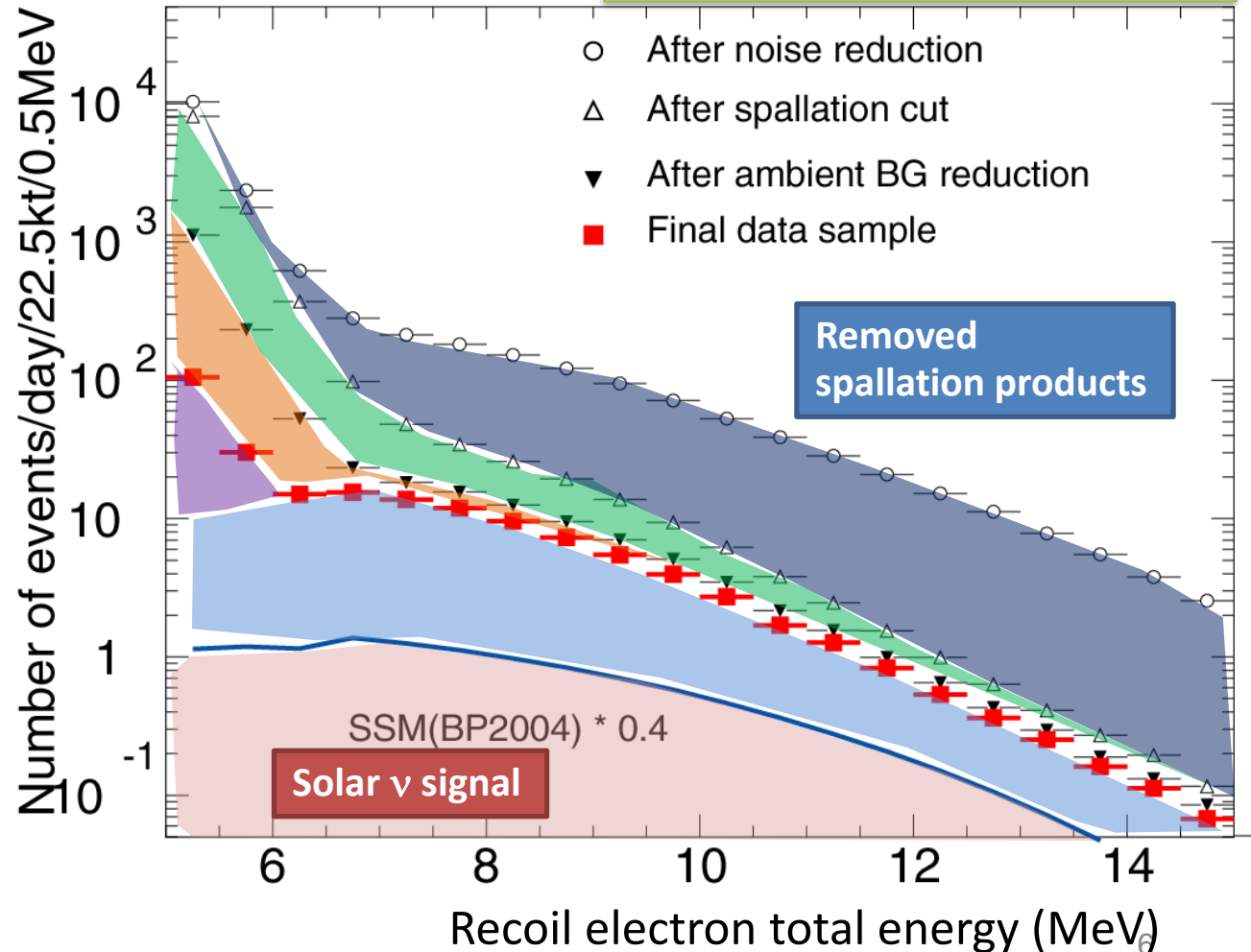


SK-I: PRD73, 112001



Spallation is dominant BG source in  $\sim 6.5\text{-}20\text{MeV}$

Solar  $\nu$  data reduction in SK-I  
(in 2m fiducial volume, 22.5kt)



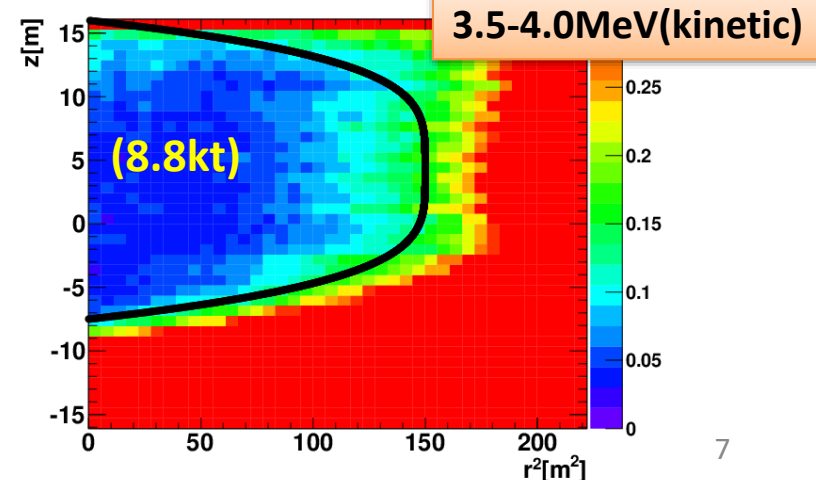
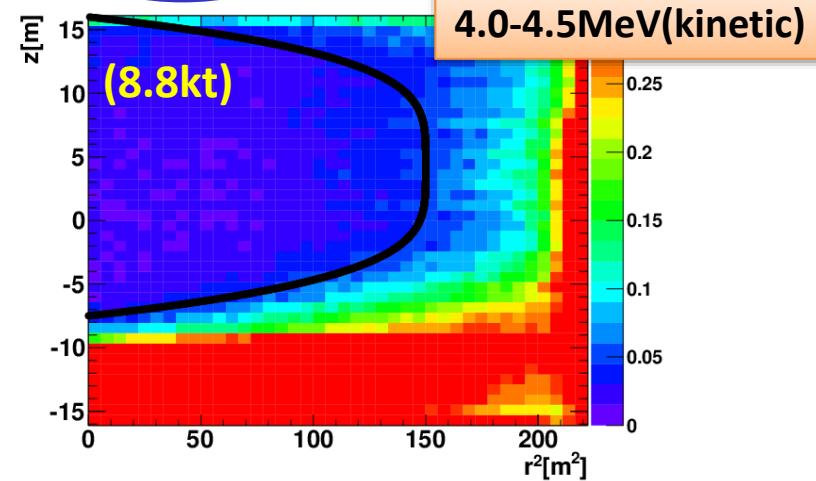
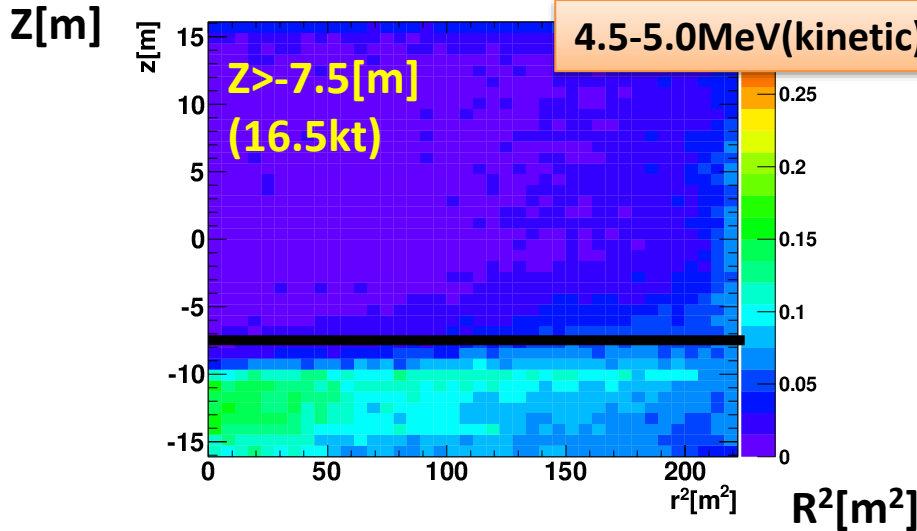
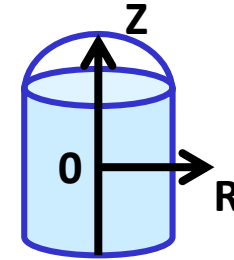
# Vertex distribution in SK-IV

June 2017

Preliminary

SK-IV 2645 days

Color : Events/day/bin low  $\rightarrow$  high



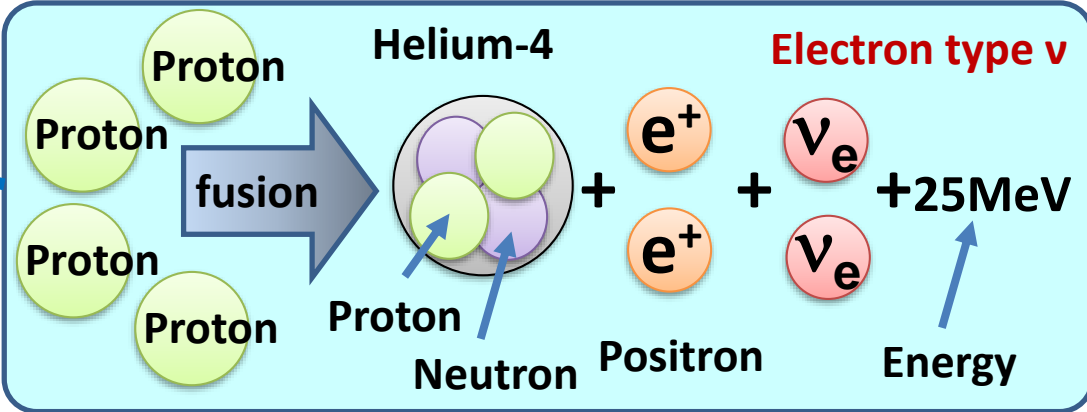
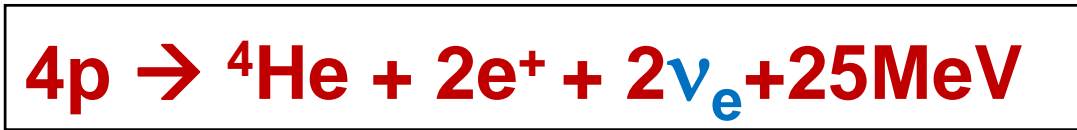
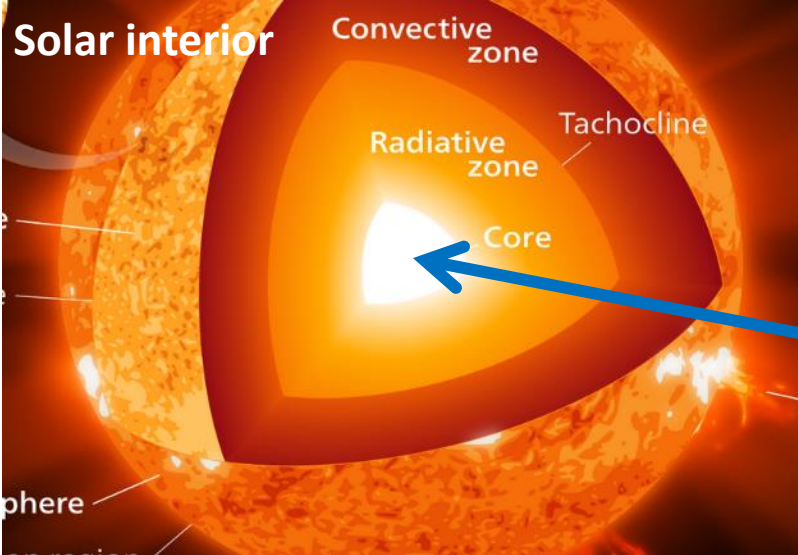
- Whole area in these plots corresponds to 22.5kton
- Above 5.0MeV(kinetic energy), fiducial volume is 22.5kton
- Below 5.0 MeV(kinetic), tight fid. vol. cut is applied.

# Outline

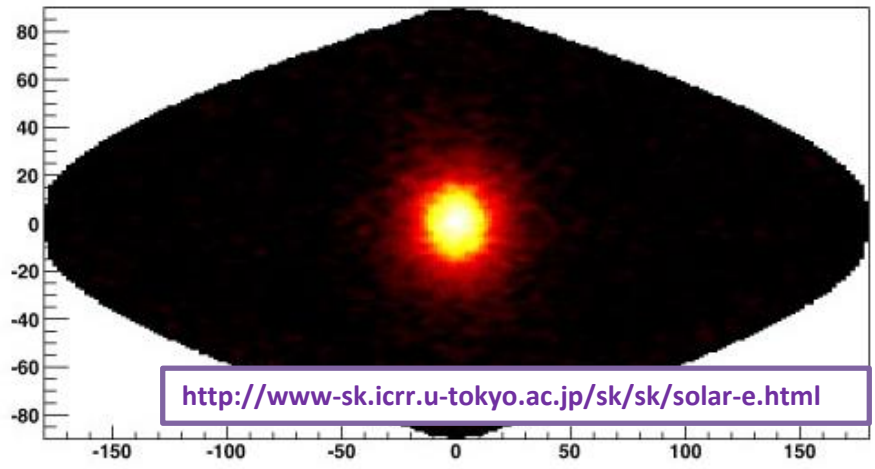
- SK detector
- **Solar neutrinos**
- Recent results from SK
- Prospects with HK
- Summary



# Solar neutrino



[http://en.wikipedia.org/wiki/Sun#/media/File:Sun\\_poster.svg](http://en.wikipedia.org/wiki/Sun#/media/File:Sun_poster.svg)



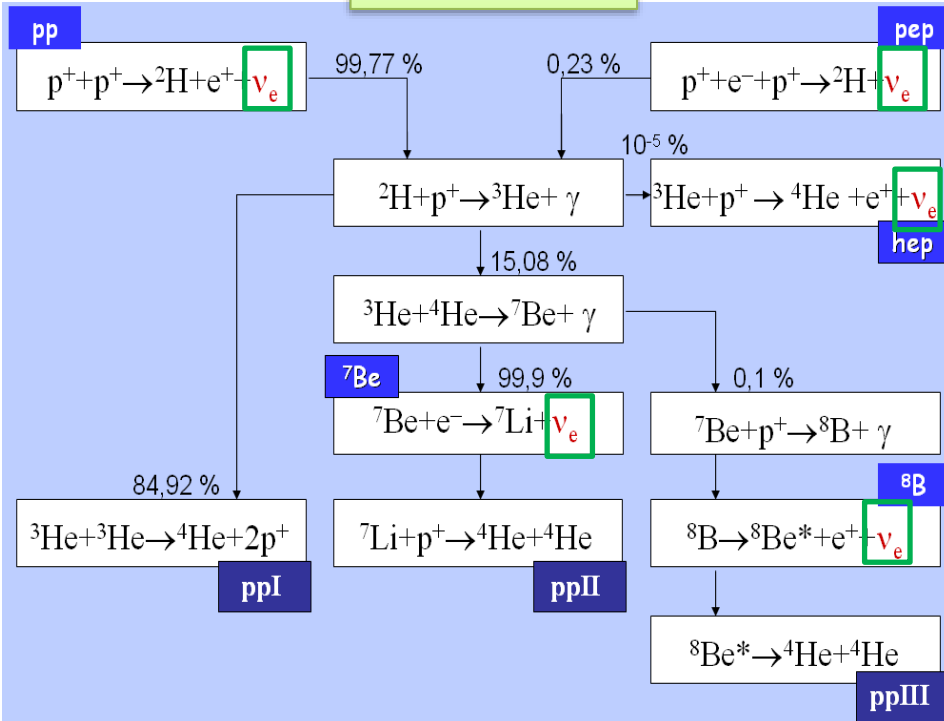
The Sun seen with neutrinos in SK. The coordinate system in which the Sun is placed at the center is used. Color means event rates.

- **Standard Solar Model (SSM)** predicts neutrino fluxes
- Most strong ν source on Earth  
at Earth: ~66 billion ν/sec/cm<sup>2</sup>
- Photon: only surface
- ν: direct observation of interior of the *present* Sun

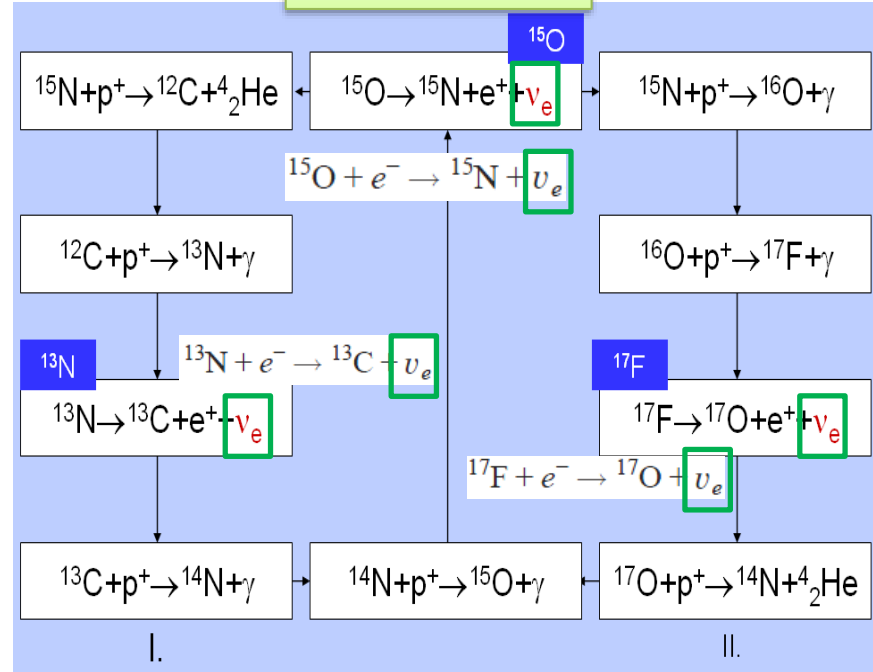
# Solar neutrinos in SSM

Overall reaction:  $4p \rightarrow {}^4\text{He} + 2e^+ + 2\nu_e + 25\text{MeV}$

## p-p chain



## CNO cycle



- Solar fusion cross sections: **SF-II** (Rev. of Mod. Phys. 83 (2011) 195)
- **B16 SSM**: A New Generation of Standard Solar Models (ApJ 835 (2017) 202)
- **Problem**: Solar abundances: **GS98** (High metallicity), **AGSS09** (Low metallicity)

# Expected solar neutrino spectrum

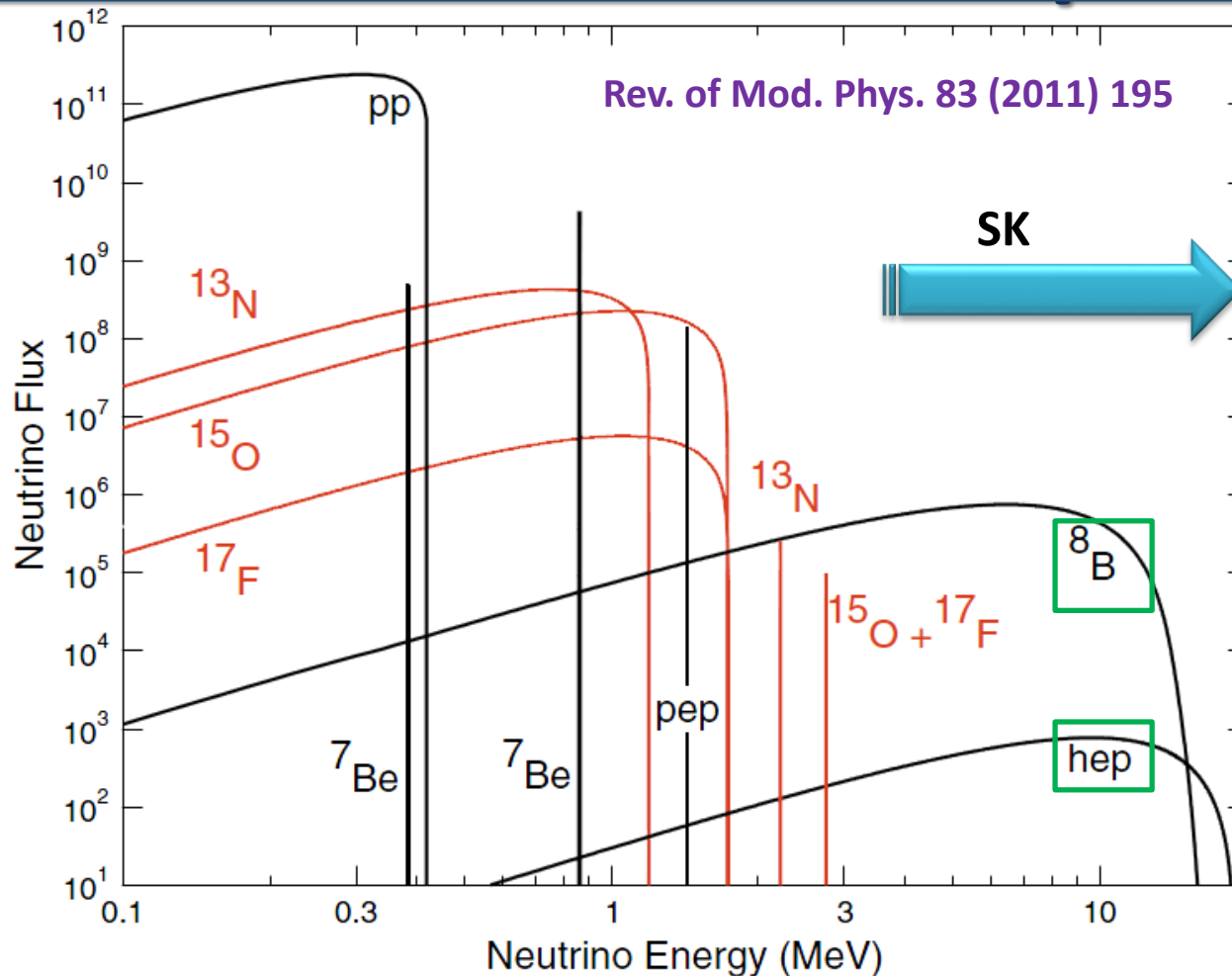


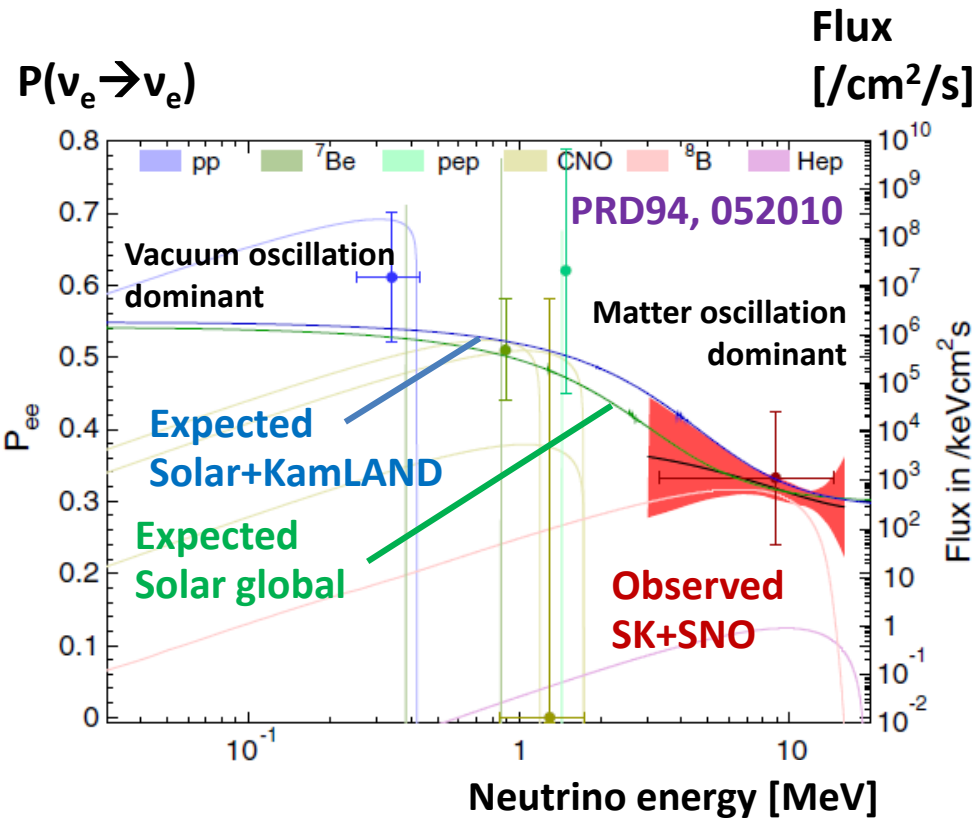
FIG. 7 (color online). Solar neutrino fluxes based on the “OP” calculations of Bahcall *et al.* (2005), with the addition of the new line features from CNO reactions. Line fluxes are in  $\text{cm}^{-2} \text{s}^{-1}$  and spectral fluxes are in  $\text{cm}^{-2} \text{s}^{-1} \text{MeV}^{-1}$ . From Stonehill *et al.*, 2004.

# Solar neutrinos at SK/HK

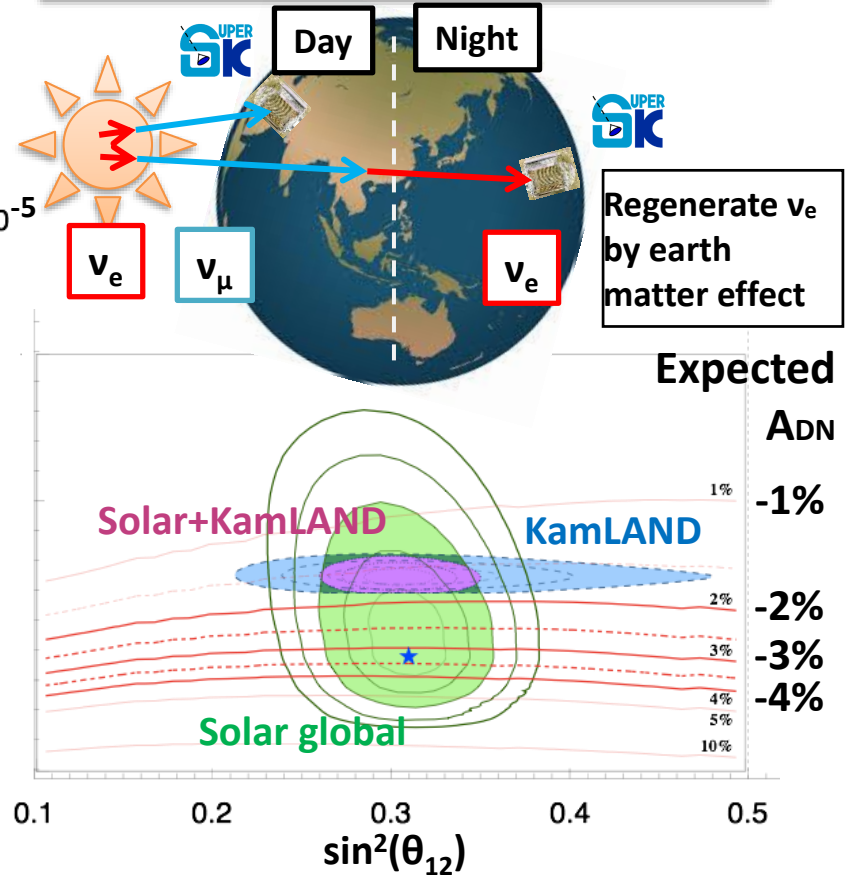
- High statistics measurement of  $^8\text{B}$  solar neutrinos
  - Possible time variation of the flux
  - **Energy spectrum distortion** due to solar matter effect
  - **Day-night flux asymmetry** due to earth matter effect

$$A_{DN} = \frac{(\text{Day} - \text{Night})}{(\text{Day} + \text{Night})/2}$$

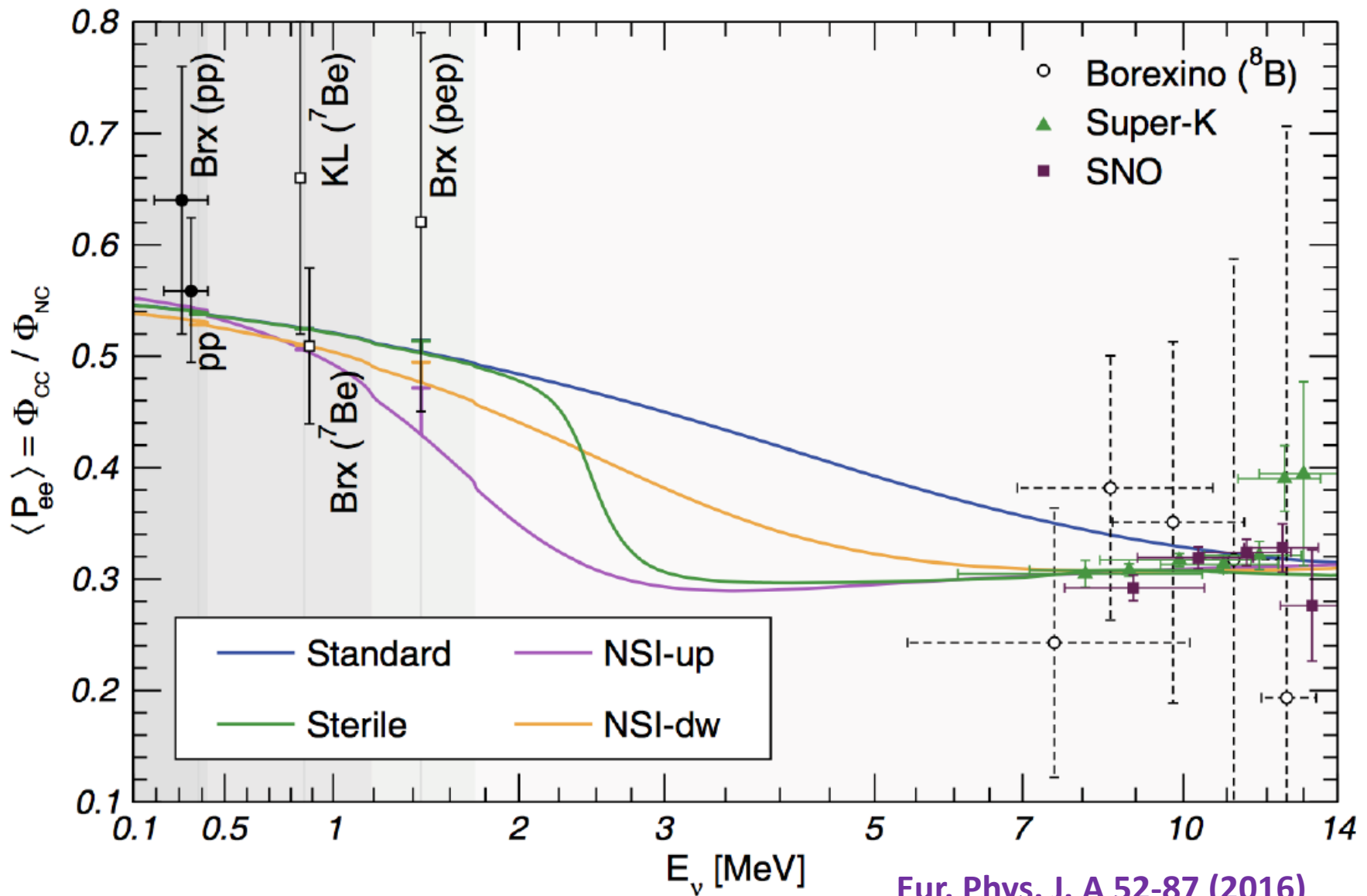
## Spectrum distortion



## Day-Night flux asymmetry



# A theoretical expectation



# Outline

- SK detector
- Solar neutrinos
- **Recent results from SK**
- Prospects with HK
- Summary

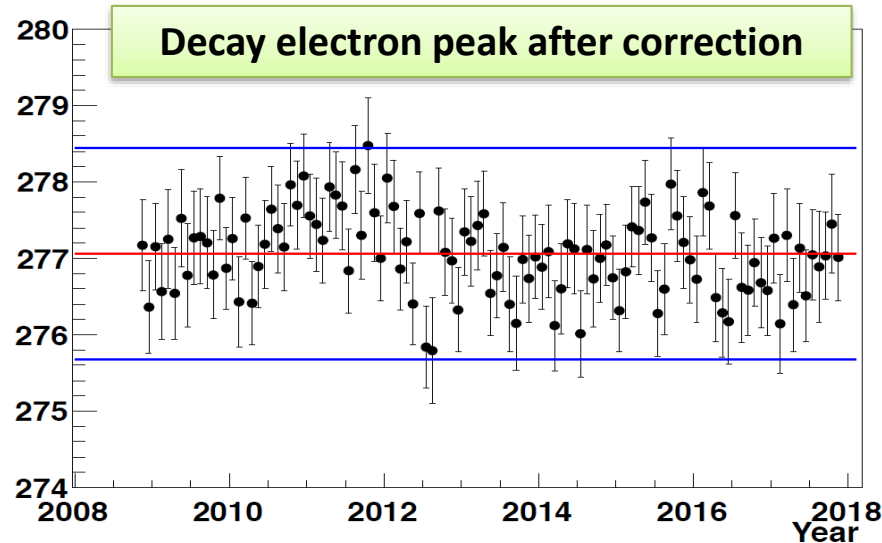
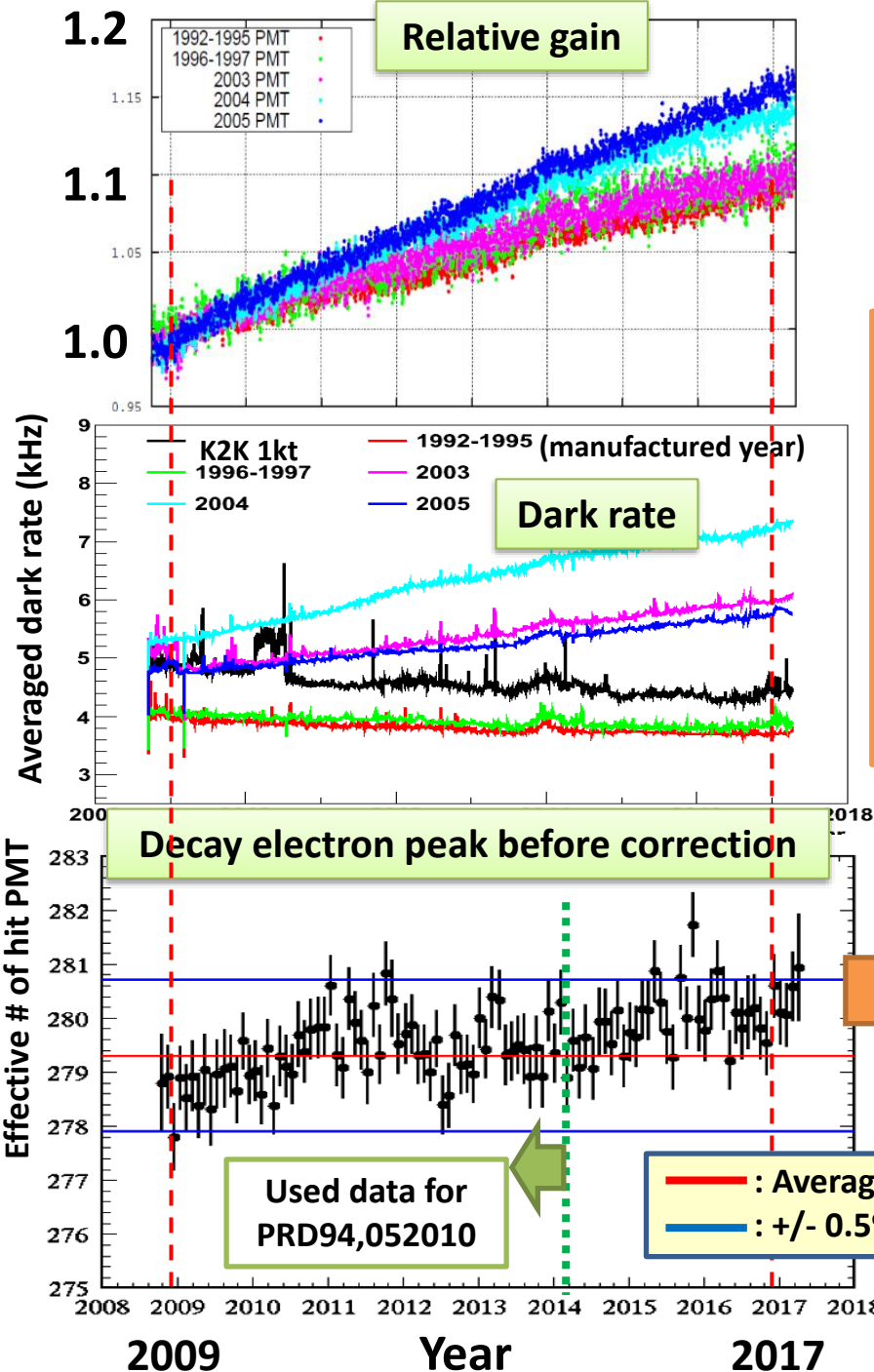
# Recent activities on solar $\nu$ analysis



- ➔ ■ **Energy scale improvements** Apr 2018
  - Taking into account PMT gain & dark rate effects
- ➔ ■ **Preliminary update of spectrum analysis** Apr 2018
  - Total live time 5695 days (May 1996 - Dec. 2017)
  - SK-I (1496 days), SK-II (791 days), SK-III (548 days), SK-IV (**1664 days** → **2860 days**)
- **Update of day/night analysis: On going (1664 days → ?)**
- **Preliminary periodic analysis in SK-IV**
  - Using same data set as PRD94, 052010
- **NSI analysis: On going**
- **Solar  $\nu$ -e-bar: On going**
- **Study of spallation BG**
  - Using neutron events (2.2 MeV  $\gamma$  from n+p) in SK-IV
- **Study of radon BG**
  - “Measurement of Radon Concentration in Super-Kamiokande's Buffer Gas”, NIM A (DOI: 10.1016/j.nima.2017.04.037 )

# Energy scale improvement

- PMT gain and PMT dark rate are changing, and it caused a slight drift of energy scale as a function of time.
- A new energy scale with gain and dark rate correction is implemented.
- It enables low-energy analysis of **whole SK-IV period** with a same systematic uncertainty.





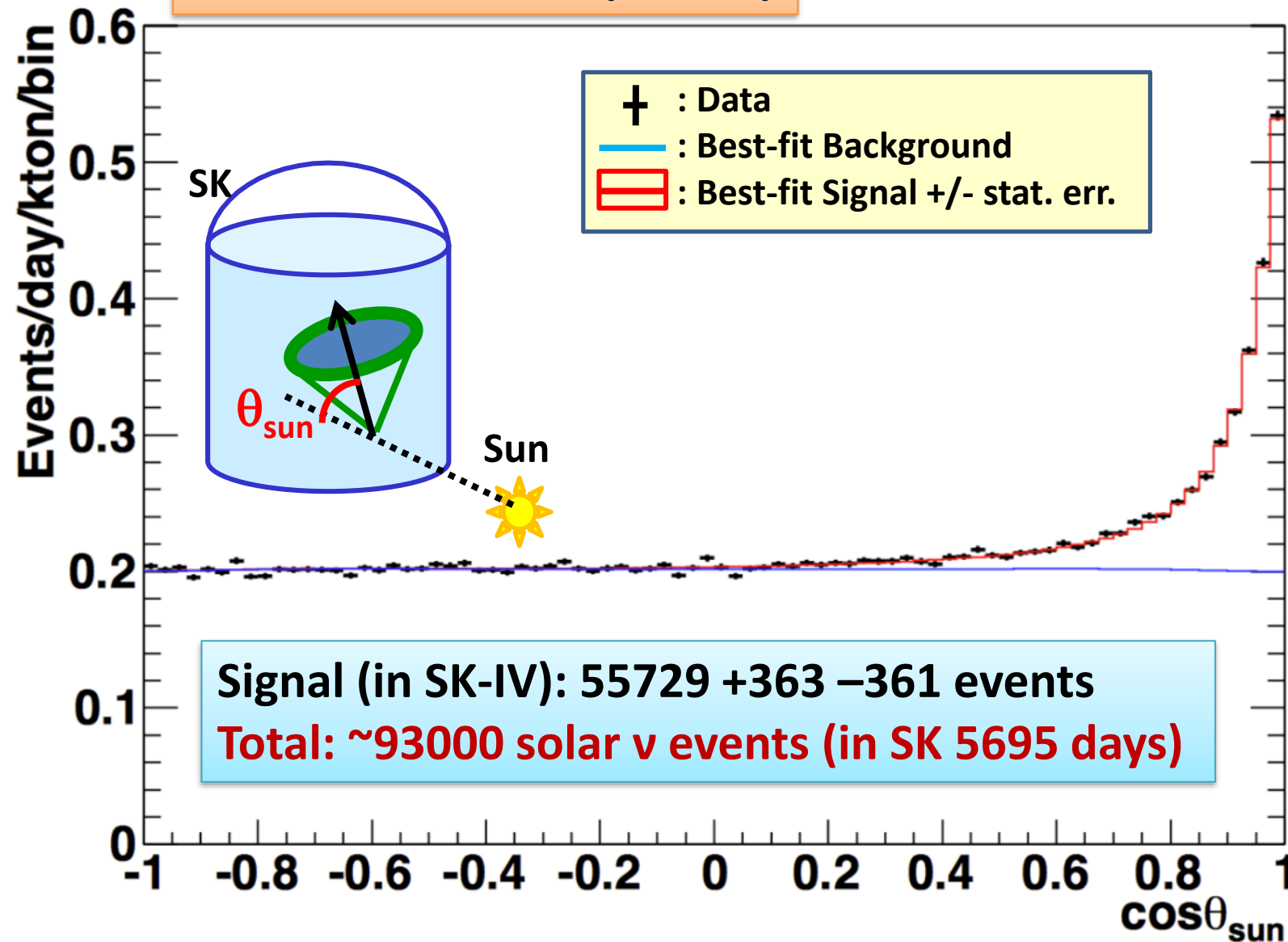
# SK-IV solar neutrino signal

Apr 2018

Preliminary

SK-IV 2860days

SK-IV 3.5-19.5 MeV(kinetic)

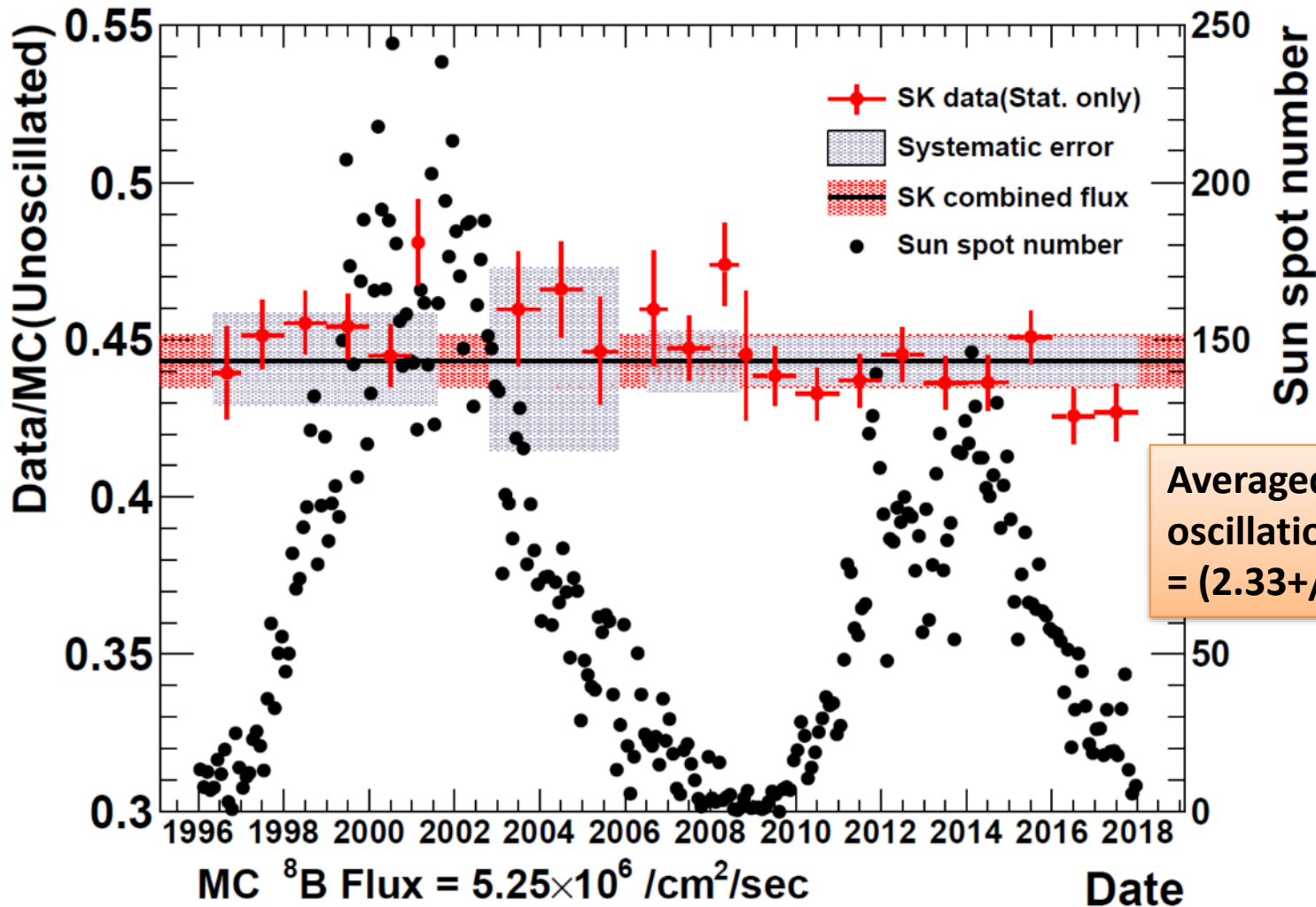


# $^8\text{B}$ solar neutrino flux: Yearly plot

Apr 2018

Preliminary

SK 5695 days



Averaged  $^8\text{B}$  flux with no oscillation  
=  $(2.33 \pm 0.04) \times 10^6 / \text{cm}^2 / \text{s}$

MC  $^8\text{B}$  Flux =  $5.25 \times 10^6 / \text{cm}^2 / \text{sec}$

$\chi^2 = 21.57 / 21 \text{ d.o.f.} \rightarrow$  Confidence level = 41.4 %

Super-K solar rate measurements are fully consistent with a constant solar neutrino flux emitted by the Sun.

Sun spot number:  
WDC-SILSO, Royal  
Observatory of  
Belgium, Brussels

# Solar $\nu$ oscillation results

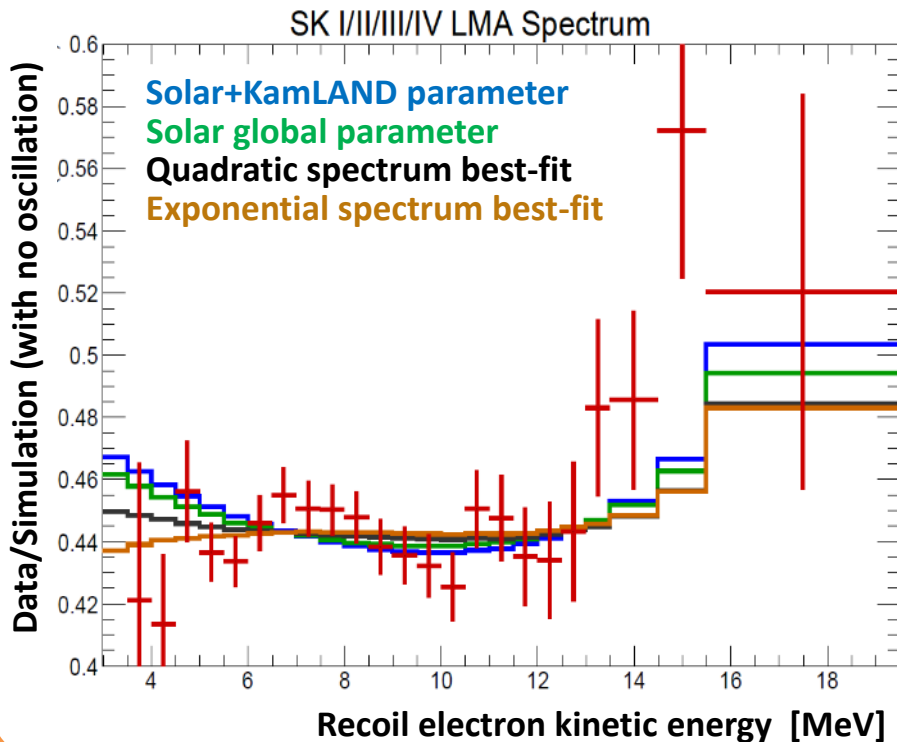
Apr 2018

Preliminary

SK 5695 days

- Quadratic fit of SK spectrum is consistent with solar  $\Delta m_{21}^2$  within  $\sim 1.2 \sigma$  and disfavors KamLAND  $\Delta m_{21}^2$  by  $\sim 2.0 \sigma$ .
- $\sim 2.0 \sigma$  level tension in  $\Delta m_{21}^2$  between solar global analysis and KamLAND is still remaining.

## Solar $\nu$ energy spectrum

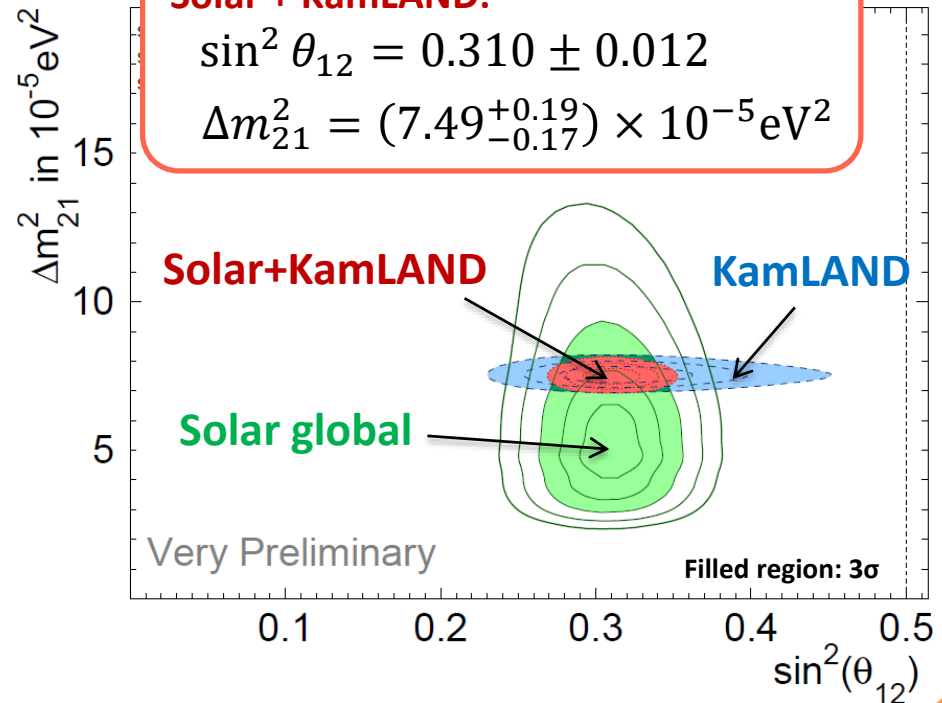


## Solar $\nu$ oscillation parameters

**Solar + KamLAND:**

$$\sin^2 \theta_{12} = 0.310 \pm 0.012$$

$$\Delta m_{21}^2 = (7.49^{+0.19}_{-0.17}) \times 10^{-5} \text{eV}^2$$



# Super-K Spectral Data

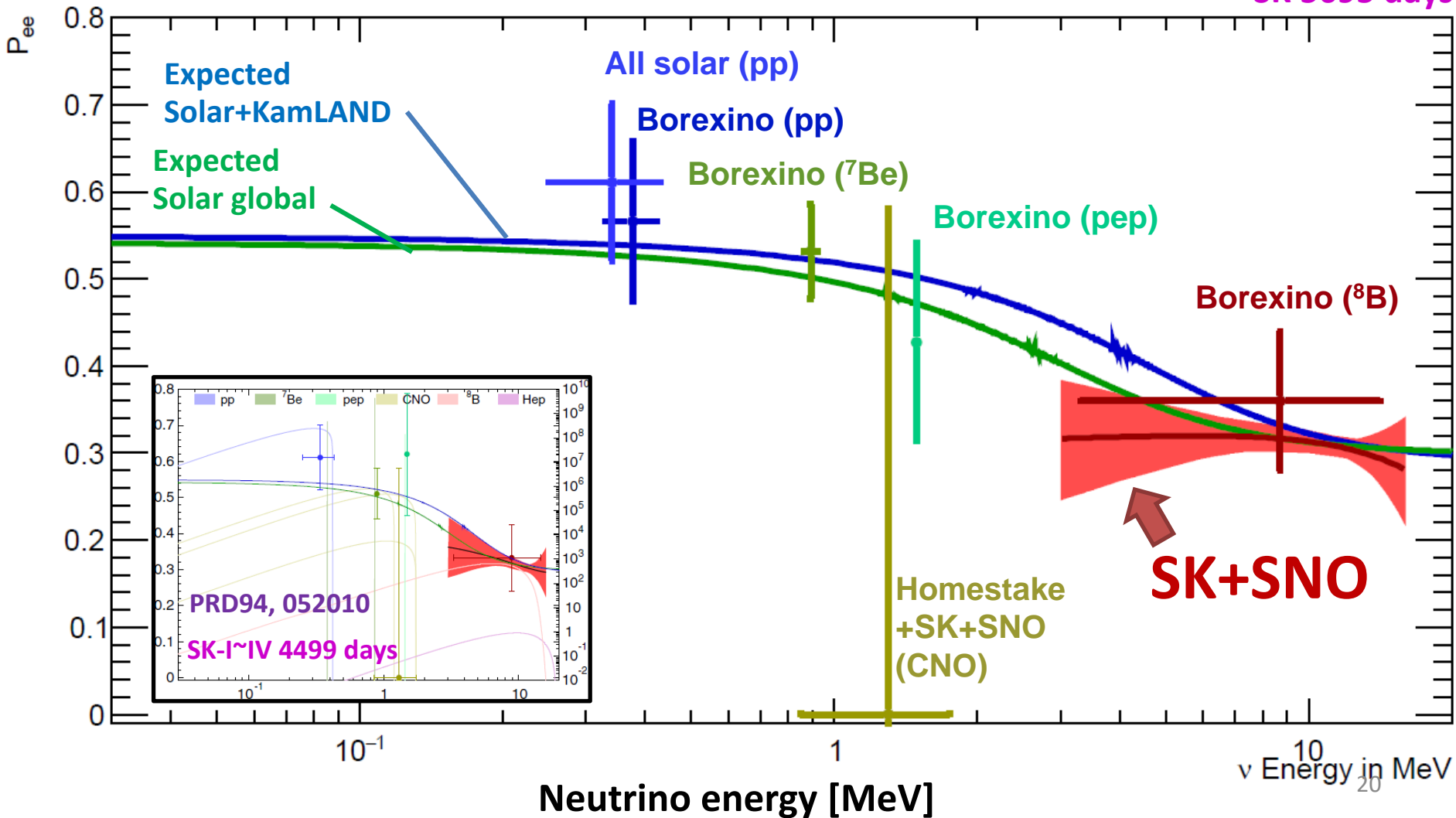
Apr 2018

Preliminary

SK 5695 days

$\nu_e$  survival probability

$$P(\nu_e \rightarrow \nu_e)$$



# Outline

- SK detector
- Solar neutrinos
- Recent results from SK
- **Prospects with HK**
- Summary

# Hyper-Kamiokande (HK)



Design Report 2018: arXiv:1805.04163

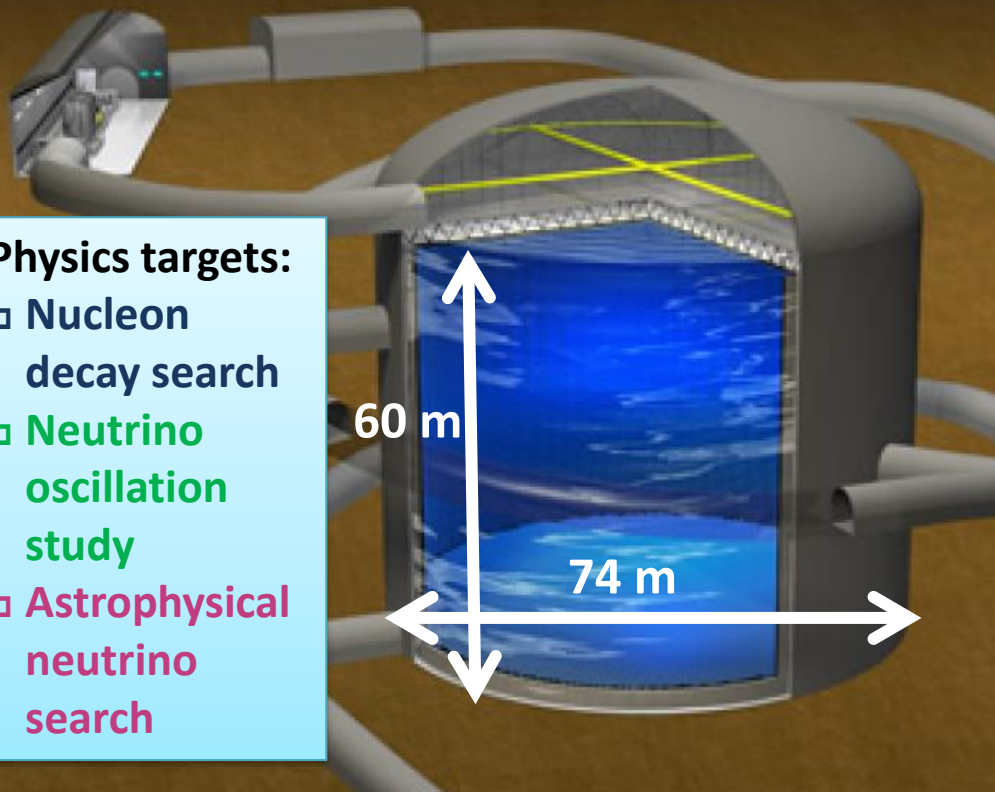
- Gigantic neutrino and nucleon decay detector in Kamioka, Japan
- **187 kton fiducial mass**:  $\sim 8 \times$  SK
- **x 2 higher photon sensitivity** than SK
- MW-class world-leading  $\nu$ -beam by upgraded J-PARC



Hyper-K



J-PARC  
Accelerator Complex



## Physics targets:

- Nucleon decay search
- **Neutrino oscillation study**
- **Astrophysical neutrino search**

## Current status:

- HK is a priority project by MEXT's Roadmap 2017
- **MEXT allocated a seed budget of HK in FY2019.** (In SK, after the one-year seed budget, full budget was allocated)
- **U. Tokyo decided to start HK construction in April 2020.**
- To enhance neutrino oscillation physics, a 2<sup>nd</sup> detector in Korea is under study

# Solar neutrinos in HK



Design Report 2018: arXiv:1805.04163

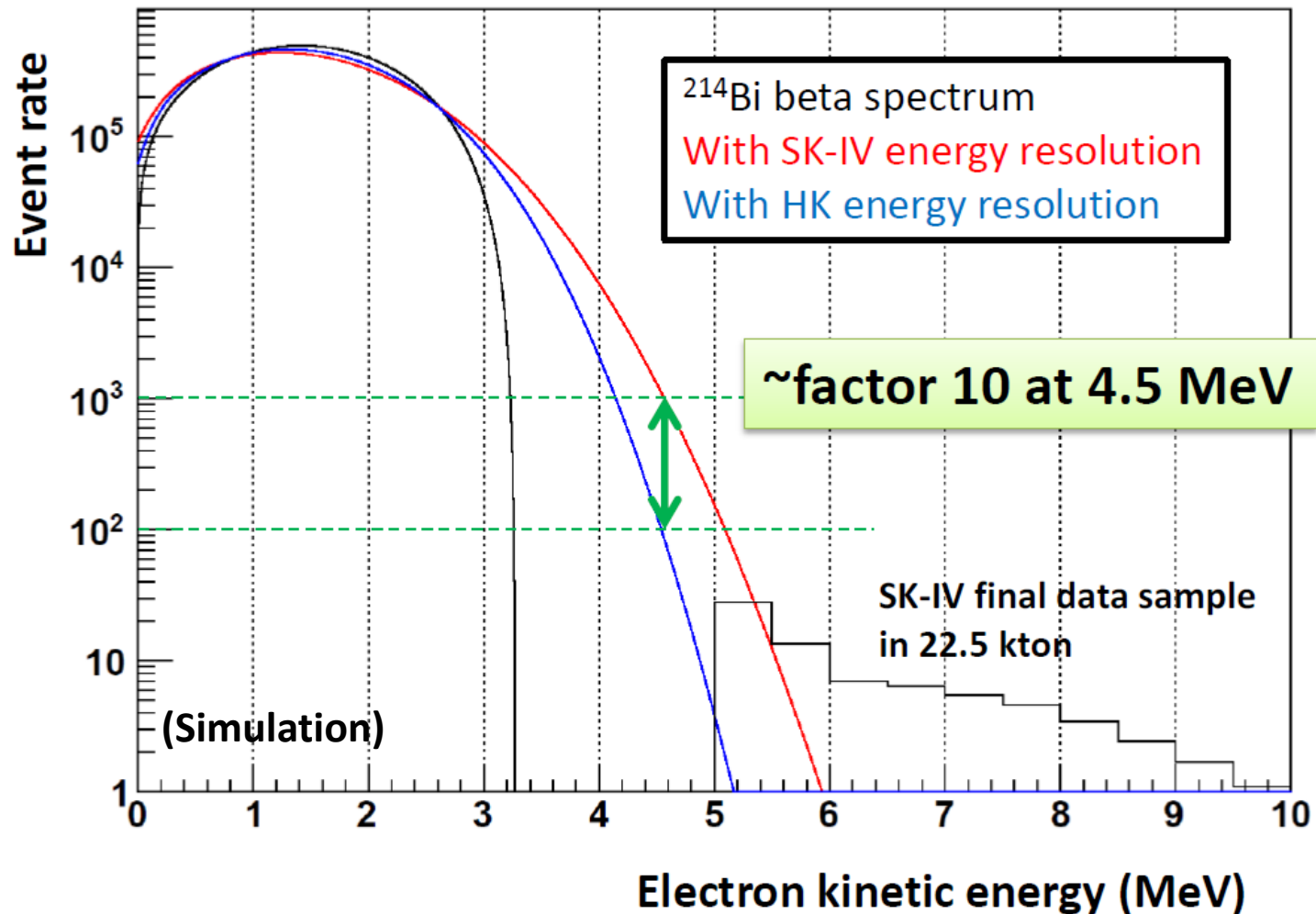
- **Expected event rate**
  - $^8\text{B}$  solar neutrino: **~130 events/day**
  - 4.5 MeV threshold (visible energy)
  - with oscillation, scaled from SK rate
- **Neutrino physics**
  - Oscillation parameters
  - NSI analysis
- **Solar physics**
  - *Hep* solar neutrino
  - Core temperature monitor
  - Solar g-mode oscillation
  - Neutrino from solar flare

# An estimation of Rn background



Design Report 2018: arXiv:1805.04163

- Radon events could be reduced in HK thanks to **better energy resolution**.



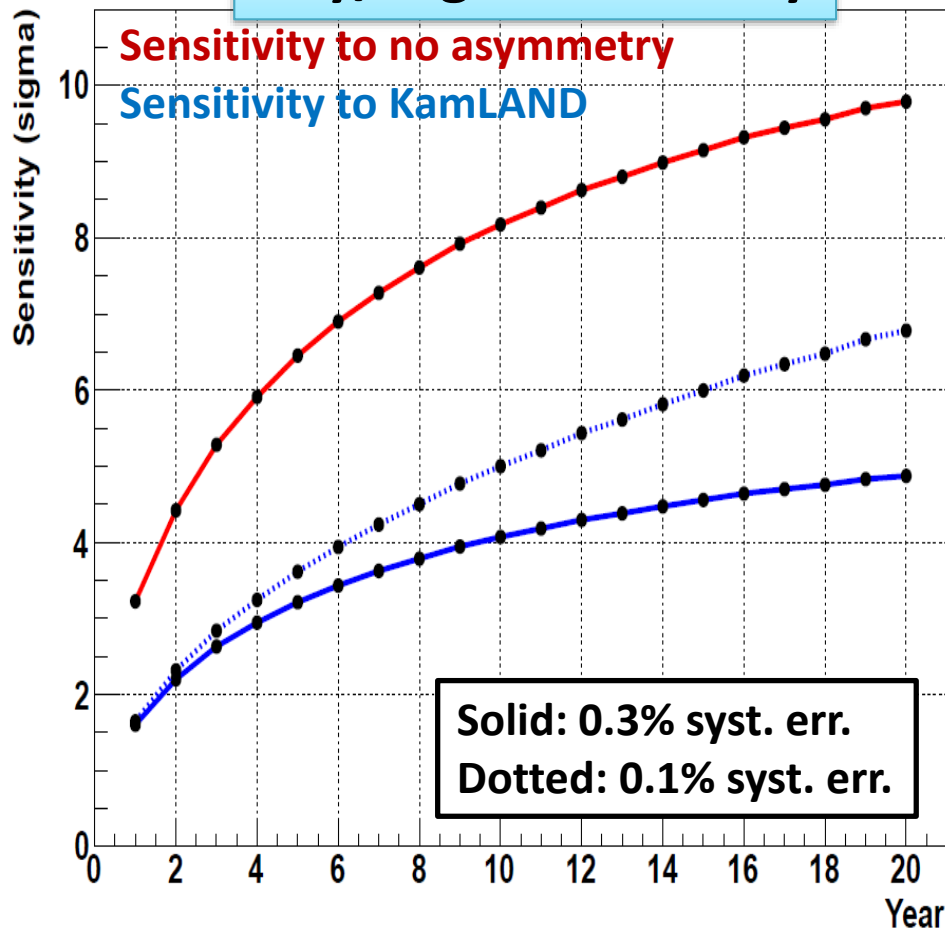


# $^8\text{B}$ Solar neutrino measurements

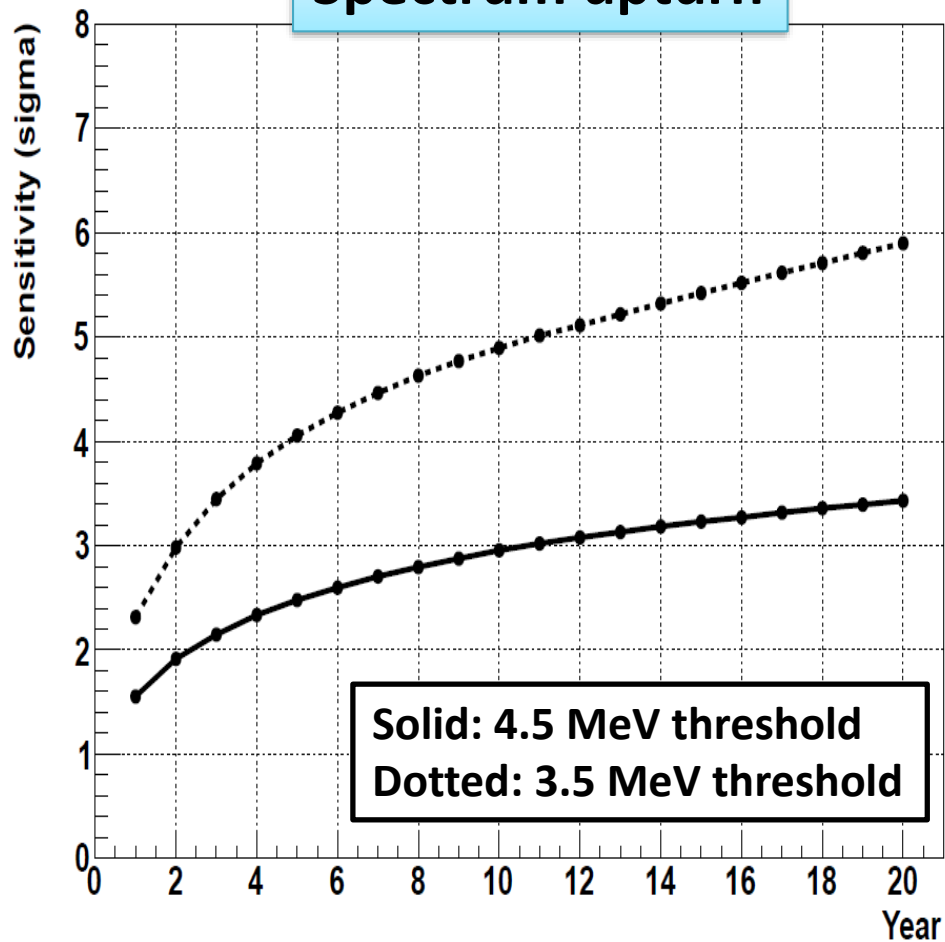


Design Report 2018: arXiv:1805.04163

## Day/Night sensitivity



## Spectrum upturn



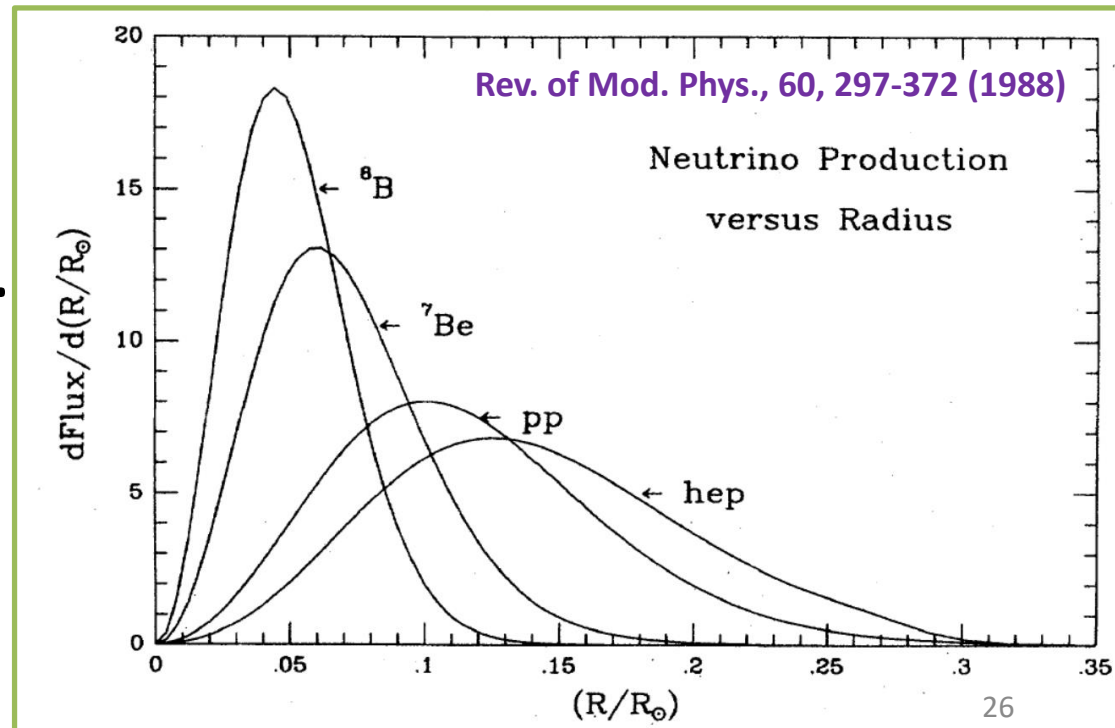
- Day/Night (solar vs reactor): 4~5 sigma in 10 years
- Spectrum upturn: ~3 sigma in 10 years

# Hep solar neutrino



Design Report 2018: arXiv:1805.04163

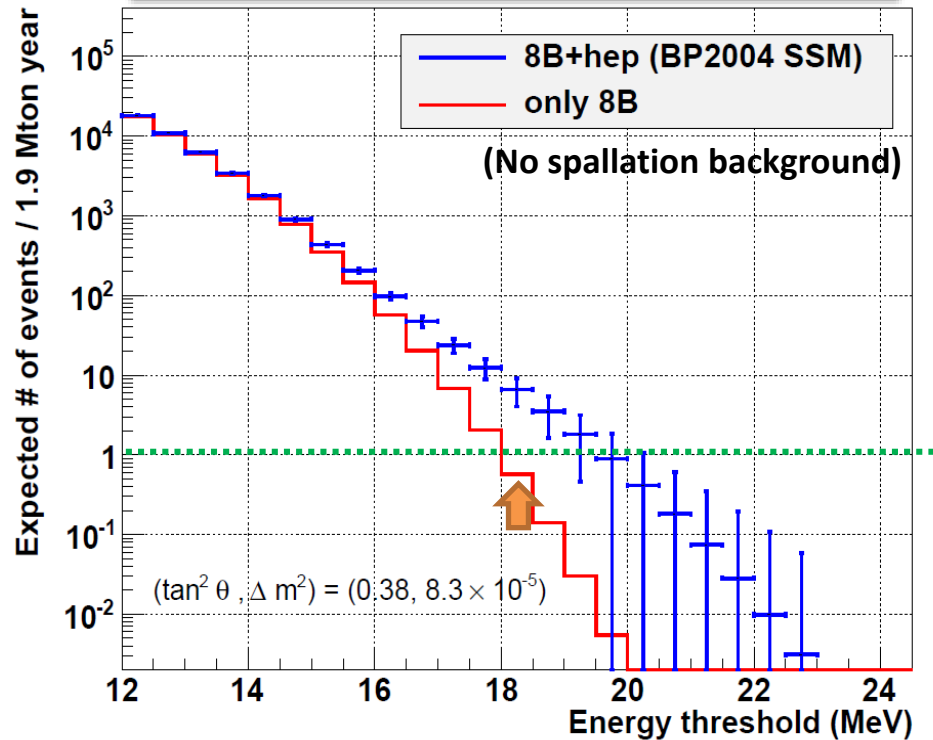
- The last piece of the solar neutrinos in pp-chain.
- Theoretical calculation is difficult.
  - → Better understanding of solar physics
- Production area is different.
  - → A new probe of the solar interior around core region
- Non-standard solar models predict the potential enhancement of the hep neutrino flux.
- Could be additional input to the solar chemical composition problem.



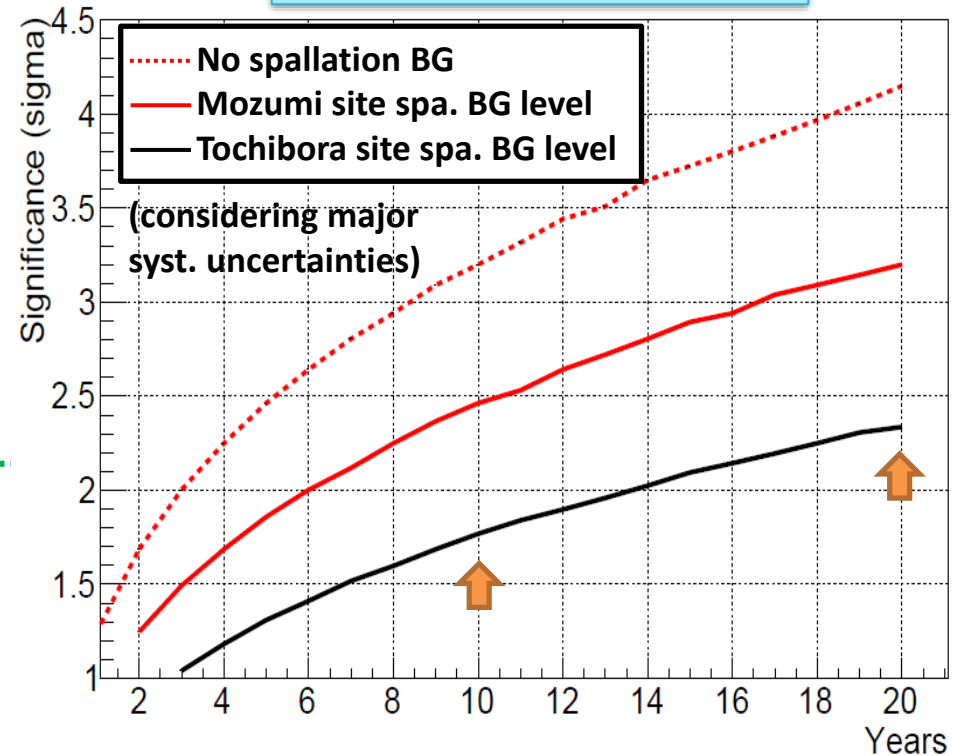
# Hep solar neutrino at HK



Integrated # of expected solar neutrino events



Expected non-zero significance

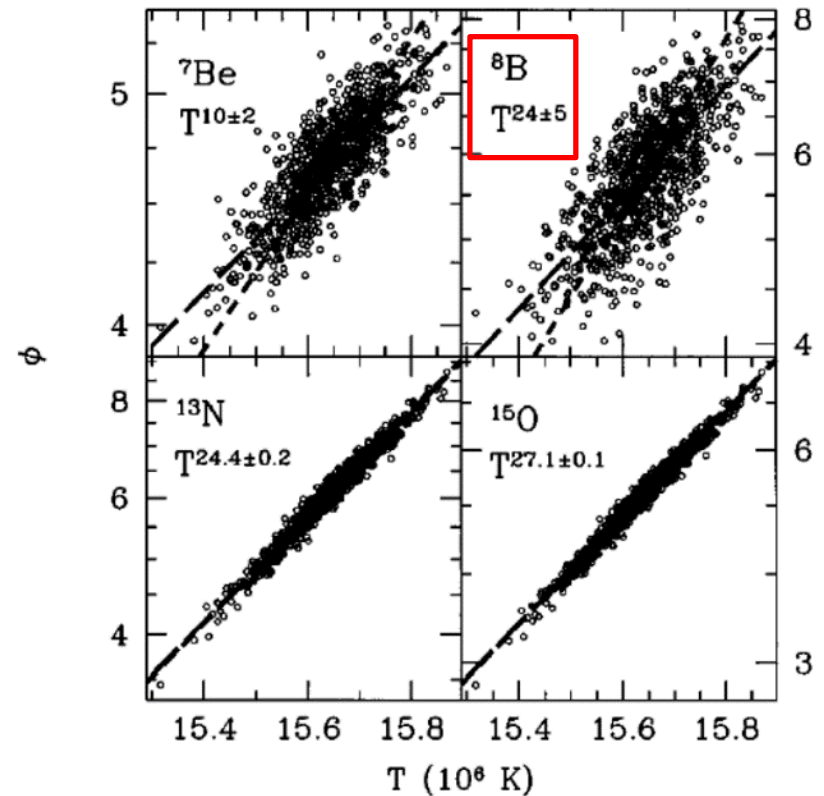
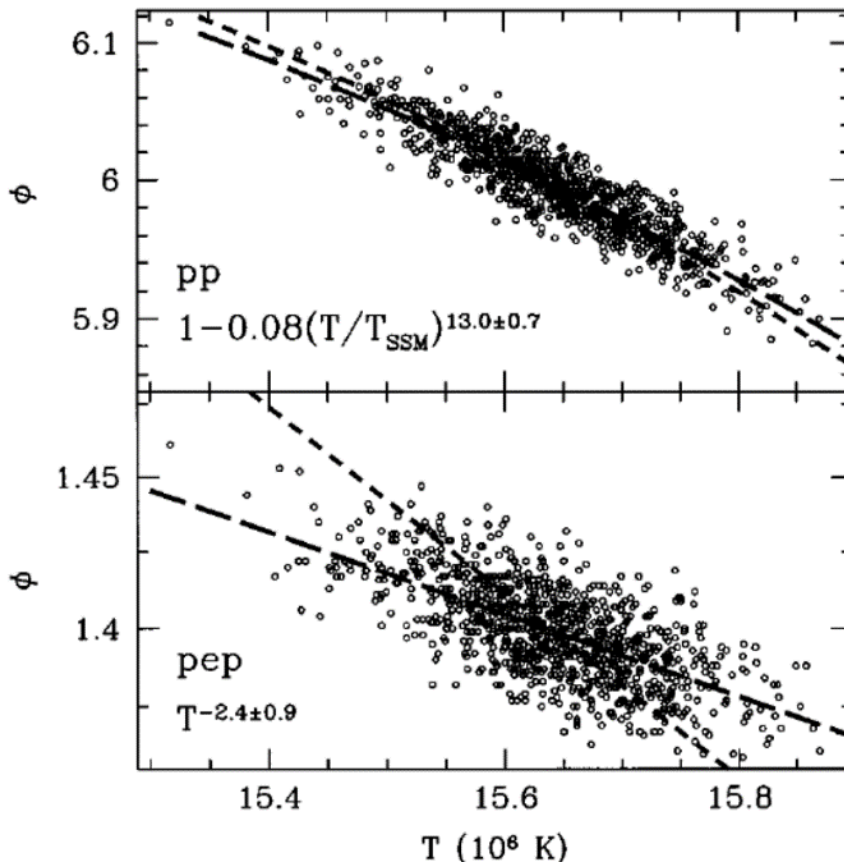


	$E_{\text{total}}$ [MeV]	$^8\text{B}$ [/1.9Mt yr]	Hep [/1.9Mt yr]	Hep / $^8\text{B}$
SK	19.5-25.0	0.77	3.03	3.9
HK	18.0-25.0	0.56	6.04	10.6

- First direct observation.
- Energy resolution & spallation BG reduction are essential.

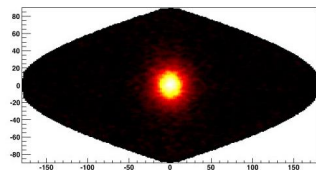
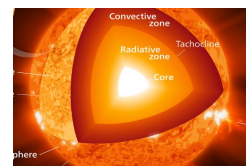
# Temperature dependences of $\nu$ flux from SSM

J. N. Bahcall and A. Ulmer, Phys. Rev. D 53 (1996) 4202.



${}^8\text{B}$  flux is very sensitive to solar core temperature (flux  $\propto T^{24}$ )  
 $\rightarrow$  study of possible time variation of  ${}^8\text{B}$  flux (Solar physics)

(Accuracy of core temperature measurement:  $\sim 0.8\%$  day by day at HK)



# Summary



- Solar neutrino has been observed since 1970's
  - It could be used as a probe of solar physics
  - Agreement between SSM +  $\nu$  oscillation and solar neutrino experiments looks good, but there is **some room for new physics**
    - Tension between solar and reactor experiments
    - Energy spectrum in vacuum-matter transition region
  - A very precise solar neutrino measurements could be done **with HK for both  $\nu$  properties and solar physics**
    - Oscillation parameters, *hep* neutrinos, short time variations, solar flares, ...