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## Other New Physics and the MiniBooNE Excess? Gordan Krnjaic

+Johnathon Jordan, Yonatan Kahn, Matthew Moschella, Joshua Spitz Phys.Rev.Lett. 122 (2019) no.8, 081801 arXiv:1810.07185

Prospects of Neutrino Physics, Kavli IPMU, April 12, 2019

Overview

1) Review of the MiniBooNE Excess

2) Excluding Simple Models w/ Kinematic Distributions

3) Excluding (Nearly) All Other Models w/ Beam Dump Data

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Comparable oscillation probabilities

MiniBooNE Collaboration 1805.12028

MiniBooNE Analysis Details

Luminosity	neutrino mode $12.84 \times 10^{20} \text{ POT}$	antineutrino mode $11.27 \times 10^{20} \text{ POT}$
Reconstructed Neutrino Energy	$200 < E_{\nu}^{QE} < 1$	$250 \mathrm{MeV}$
Excess events BG subtracted	$381.2 \pm 85.2$	$79.3 \pm 28.6$

#### **Possibly Important Caveat**

Events/MeV Mild tension ~ 2+ sigma between neutrino and antineutrino mode

Updated Neutrino Mode Analysis MiniBooNE Collaboration1805.12028 Complements earlier antineutrino results collected 2002-2010

Excess

#### MiniBooNE Anomaly

Neutrino mode only Both excesses, BG subtracted Events/MeV Excess Events/MeV Data (stat err.) 1.8 → v<sub>e</sub>: 12.84×10<sup>20</sup> POT  $v_{a}$  from  $\mu^{+}$  $v_{o}$  from  $K^{+}$ 1.6 from K<sup>u</sup> misid 1.4  $\Delta \rightarrow N\gamma$ dirt 1.2 other Constr. Syst. Error 3 Best Fit 0.8 0.6 2 0.4 0.2 8.2 -0.2 0.4 1.2 0.6 0.8 1.4 1 3.0 0.4 0.2 0.6 0.8 1.2 3.0 1.4 E<sub>v</sub><sup>QE</sup> (GeV)  $E_v^{QE}$  (GeV)

$$E_{\nu}^{(\text{reconst.})} = \frac{2m_n E_e + m_p^2 - m_n^2 - m_e^2}{2(m_n - E_e + \cos\theta_e \sqrt{E_e^2 - m_e^2})}$$

Measure charged lepton energy/angle Observed ~ 400 events, PMNS predicts 0 Combined  $\nu/\bar{\nu}$  modes<sup>12.84</sup>18°  $\bar{\sigma}$  excess  $\bar{\nu}_{e}$ : 11.27×10<sup>20</sup> POT

s Events/MeV

MiniBooNE Collaboration 1805.12028

**Important Caveats** 

**1) Could be an experimental artifact** Unknown systematic or mismeasured BG

2) LSND/MiniBooNE connection is *assumes* steriles Disfavored by disappearance & cosmolgy See Marztinez-Soler and Marfatia's talks

**3) I will ignore both LSND and sterile neutrinos** Can MiniBooNE anomaly be any *other* new physics?

## Overview

1) Review of the MiniBooNE Excess

## 2) Excluding Simple Models w/ Kinematic Distributions

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#### What's a "Simple Model"?

New particle **unrelated** to neutrino oscillation or production

**Scenario A: Unstable** particle produced in target Decays **visibly** inside the detector

Scenario B:Stable particle produced in targetScatters elastically inside the detector

# Scenario A:Unstable particle produced in targetDecays visibly inside the detector



Detector can't distinguish electrons/photons Collimated particles reconstruct as one "CCQE" track



#### **Scenario A: Unstable** particle produced in target Decays visibly inside the detector

3.0



## Scenario A:Unstable particle produced in targetDecays visibly inside the detector



*E* > 200 MeV required by cuts, easy to fake a single track  $\cos \theta_{12} \approx \cos \theta_X > 0.999$ 

Model independent exclusion of visible decays

# Scenario B:Stable particle produced in targetScatters elastically inside the detector



Elastic scatter must use detector electrons as targets to fake CCQE

$$\cos \theta_e = \frac{E_X E_e - m_e (E_X + m_e - E_e)}{\sqrt{(E_X^2 - m_X^2)(E_e^2 - m_e^2)}}$$



e.g. dark matter induced

# Scenario B:Stable particle produced in targetScatters elastically inside the detector

If X is relativistically produced  $E_X \gg m_X$ 

$$\cos \theta_e = 1 - m_e \left(\frac{E_X - E_e}{E_X E_e}\right) + \mathcal{O}\left(\frac{m_e^2}{E_e^2}\right) > 0.99$$

Same problem: all events in last bin after *Ee* > 200 MeV cut

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If X is quasi-relativistic  $E_X \gtrsim m_X$ 

$$\cos \theta_e = \frac{E_X E_e - m_e (E_X + m_e - E_e)}{\sqrt{(E_X^2 - m_X^2)(E_e^2 - m_e^2)}} \implies E_e \sim m_e \quad \text{for } \cos \theta_e \sim 0$$
  
Fails selection cuts

#### Elastic scatter ruled out model independently

**Generalizing our assumptions** 

**Model Independent Arguments (Kinematic Features Only)** 

Scenario A: Unstable particle produced in target Decays visibly inside the detector

Allow decays with both visible and invisible final states

 Scenario B:
 Stable\* particle produced in target

 Scatters elastically inside the detector

Allow scattering with

1) Nuclear targets for angular distribution

2) Inelastic coupling to produce visible EM energy

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#### Neutrino Mode vs. Beam Dump Mode



Neutrino mode uses Be target and magnet focusing Charged particles collimated Neutral particles diffuse

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Beam dump avoids target and magnet Optimized for dark matter production All particles diffuse

#### Neutrino Mode vs. Beam Dump Mode



MiniBooNE Collaboration arXiv1807.06137

Continuum production Similar in both modes



Uses full beam energy Important for heavy X

Thickness irrelevant if greater than rad. length

## Null Beam Dump Mode Search: Strong BSM Bounds

#### No signal events observed above BG



 $\sim 10^{20} \text{POT} \quad \cos \theta_e > 0.9 \quad 75 \le E_{\text{vis}}^e \text{ (MeV)} \le 850$ 45 0.5 of luminosity in neutrino mode which saw ~460 events (GeV) MiniBooNE Collaboration arXiv1807.06137 What have we learned?

Scenario A: Unstable particle produced in target Decays visibly inside the detector

Scenario A'Unstable particle produced in targetEach decay has visible & invisible daughters

 Scenario B:
 Stable\* particle produced in target

 Scatters elastically inside the detector

Scenario B'Stable particle produced in targetInelastically scatters of nucleons



### Scenario A' Unstable particle produced in target Each decay has visible & invisible daughters



Scenario **Disfavored** by angular distribution, unless  $m_X > \text{GeV}$ 

### Scenario A' Unstable particle produced in target Each decay has visible & invisible daughters

#### **Bigger Problem**: can't make ~ GeV X in charged meson decays

 $m_{\pi^+} = 139.54 \text{ MeV}$  $m_{K^+} = 493.67 \text{ MeV}$ 

Need neutral mesons or continuum production

Similar production rate in beam dump mode



**Scenario A' ruled out** ~ 60 events in beam dump mode (obs ~2)

#### **Step 1)** produce *X* from charged meson decays



100x reduction in beam dump mode for all boosts  $\implies \mathcal{O}(\text{few})$  events

**Step 2)** scatter X inelastically off detector nuclei for wide angle recoils



Can't scatter off detector electrons (always recoils w/  $\cos > 0.99$ )

**Step 3)** X' injects **EM Energy** inside detector to mimic CCQE signal



EM energy must be collimated to fake single track for CCQE signal

# Scenario B'Stable particle produced in target<br/>Inelastically scatters of nucleon/nucleusStep 1Step 2Step 3 $\pi^+$ or $K^+ \to X \cdots$ $XN \to X'N \cdots$ $X' \to EM$

#### Model building challenges:

Couple to nucleons for detector upscatter satisfy angular distribution

Avoid neutral meson+continuum production ruled out by beam dump

Avoid electron scattering otherwise always forward electrons

**Example:** pseudo-dirac X & X' coupled to new force  $U(1)_{L_e-L_{\mu}}$ 



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This model is UV consistent and anomaly free



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**Inelastic X X' coupling easy** 

2 Weyl fermions w / large Dirac mass & small Majorana mass



Conclusions

4.8 sigma MB excess, simple sterile interpretations disfavored Disappearance + Cosmology

Simple other BSM (non-neutrino) models ruled out

New beam dump DM search is powerful constraint

Requires1) production from charged mesons2) inelastic scatter off nuclei3) produce new, visibly decaying particle

All will be tested with ~ 10x existing beam dump data

## Thanks!

#### Null Result Imposes Nontrivial BSM Bounds

No signal even with neutrino mode cuts in beam dump mode



 $E_p \sim 9 \text{ GeV} \sim 10^{20} \text{POT} \quad 200 < E_{\nu}^{QE} < 1250 \text{ MeV}$ 10% of luminosity in neutrino mode which saw ~460 events MiniBooNE Collaboration arXiv1807.06137

## Dark Neutrino Portal

$$\mathcal{L}_{\mathcal{D}} \supset \frac{m_{Z_{\mathcal{D}}}^{2}}{2} Z_{\mathcal{D}\mu} Z_{\mathcal{D}}^{\mu} + g_{\mathcal{D}} Z_{\mathcal{D}}^{\mu} \overline{\nu}_{\mathcal{D}} \gamma_{\mu} \nu_{\mathcal{D}} + e\epsilon Z_{\mathcal{D}}^{\mu} J_{\mu}^{\text{em}} + \frac{g}{c_{W}} \epsilon' Z_{\mathcal{D}}^{\mu} J_{\mu}^{Z}$$
broken U(1)
dark heavy
neutrino
kinetic mixing

Also add mixing between active and (unstable) dark neutrinos

$$\nu_{\alpha} = \sum_{i=1}^{3} U_{\alpha i} \,\nu_i + U_{\alpha 4} \,N_{\mathcal{D}} \,, \quad \alpha = e, \mu, \tau, \mathcal{D},$$

Bertuzzo, Jana, Machado, Funchal 1807.09877

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Beam neutrinos mix (not oscillate) to "dark" Scatter nuclei through kinetic mixing Make dark neutrino

Decays emitting Z'



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