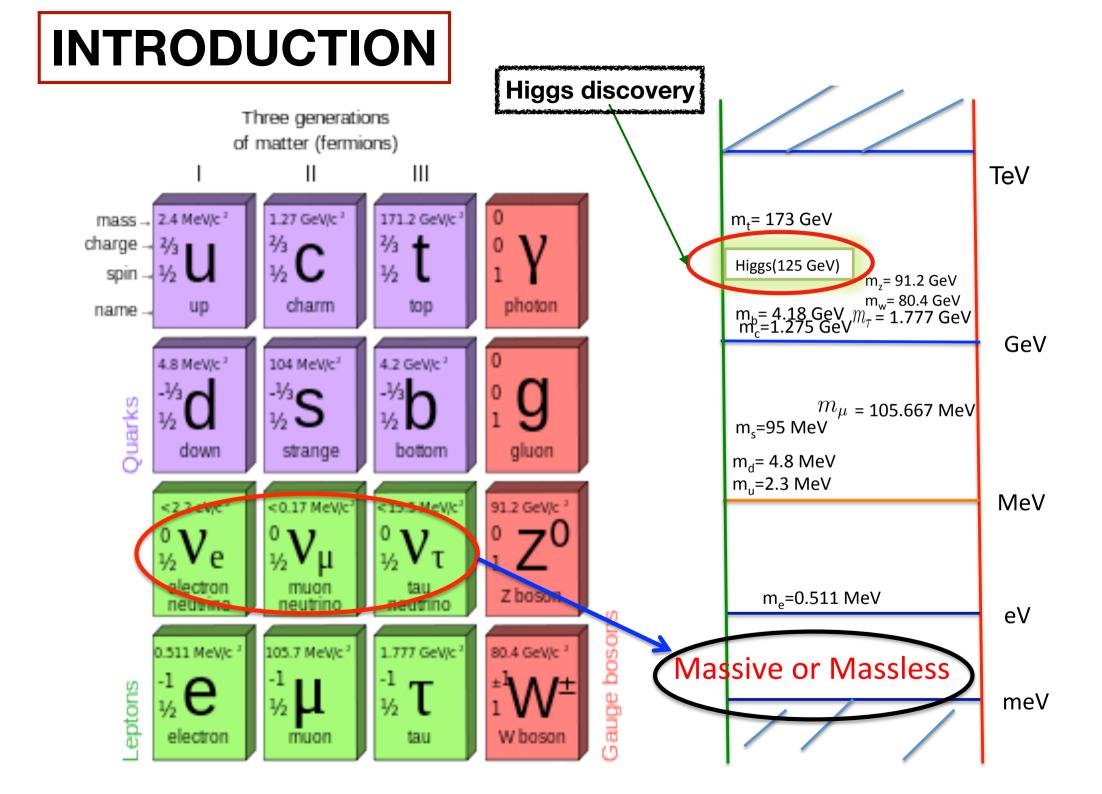
# Testing right handed neutrinos at the mear collider

# Arindam Das Osaka University

Based on: 1207.3734,1811.04291 (also see 1812.11931)



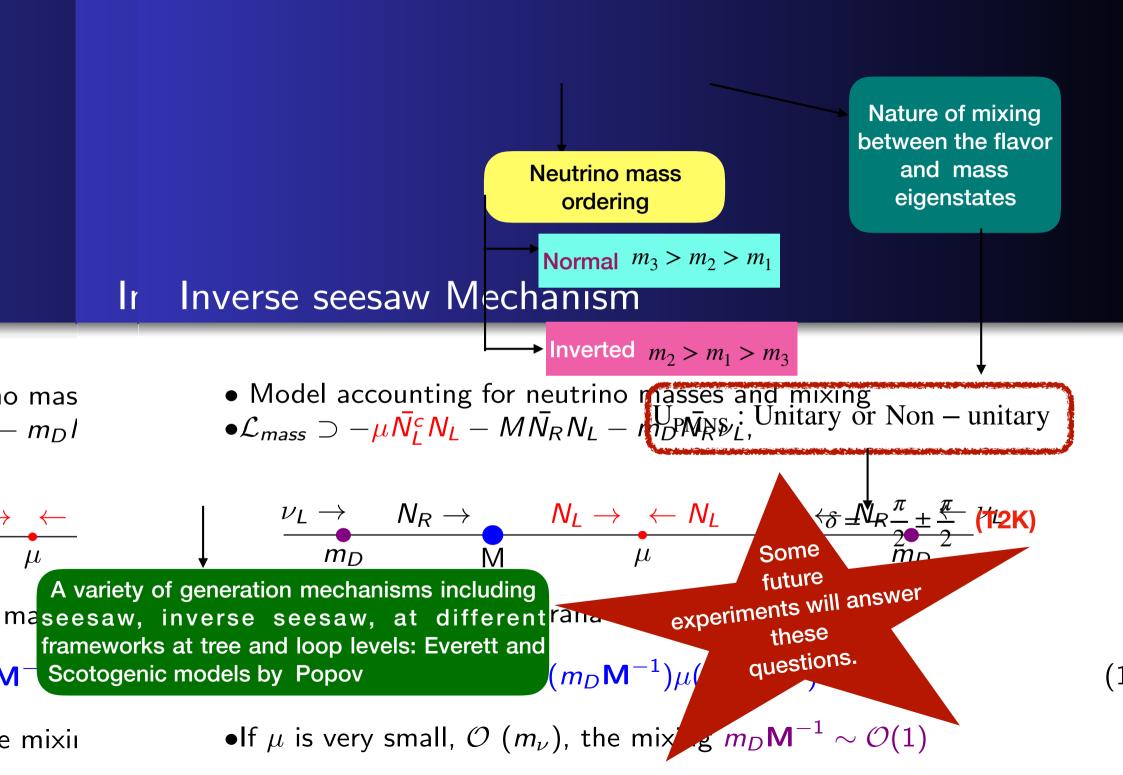
9th April, 2019 Prospect of Neutrinos, Kavli- IMPU (8th April- 12th April)



### Some interesting results in the neutrino sector

Super- Kamiokande, Sudbury Neutrino Observatory 1999, Neutrino oscillation between mass and flavor eigenstates

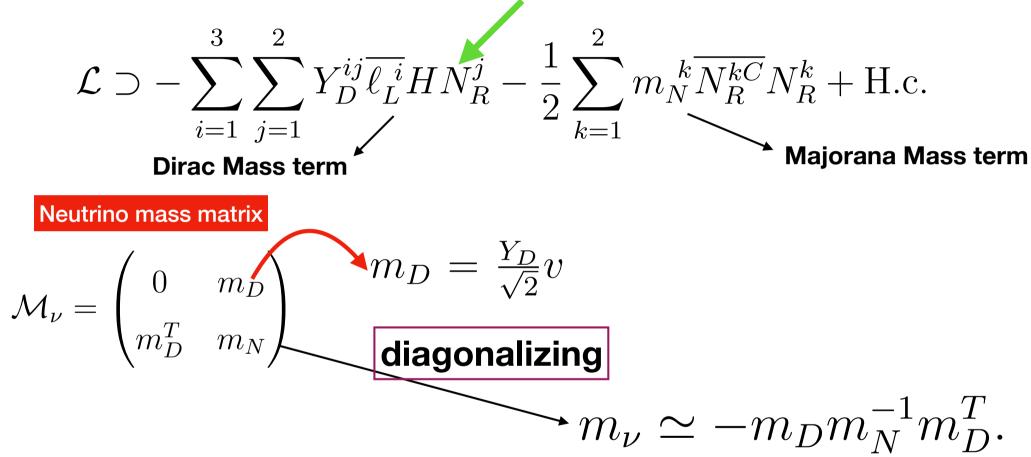
a different kind of neutrino has emerged		Physics Nobel Pr	ize 2015
TRINGS HAD MASS, BUT.		Neutrino oscillatio $7.6  imes 10^{-5} \mathrm{eV}^2$	n data SNO
Mass Found in Elusive Particle; Universe May Never Be the Same	$ \Delta m_{31} ^2$	$2.4 \times 10^{-3} \mathrm{eV}^2$	Super - K
Discovery on Neutrino Detecting Neutrinos pass trough The Earth's Survey of Control of C	$\sin^2 2\theta_{12}$	0.87	KamLAND, SNO
About All Matter By MALCOLM W. BROWNE TAKKYAMA, Japan, Jame 5 – In whether the physicists from 33 re- response to the post of the statement of the statement of the statement in a notorioally shakes is subativity and collisions and the existence of mass particle called the neutrino. These no electric charge, is so light to have no mass at all. After today's many conference of mass years to have no mass at all. After today's many conference of mass of the investment of the neutrino. The state is an electric charge, is so light to have no mass at all. After today's many conference of mass of the investment of the the statement of the statement of the statement to have no mass at all. After today's many conference of the statement of the statement of the statement to have no mass at all. After today's many conference of the statement of the statement of the statement to have no mass at all. After today's many conference of the statement of th	$\sin^2 2\theta_{23}$	0.999	T2K
		0.90	MINOS
	$\sin^2 2\theta_{13}$	0.084	DayaBay2015
	-	0.1	RENO
ecorded by successful theory of the compessition of matter known as the Standard Model. Word of the discovery had drawn some 200 physicitis here to discuss		0.09	DoubleChooz
some 340 physicids here to discuss neutrino research. Among other things, the finding of asutrino mass might affect there is about the for- region and evolution of galaxies and By analyzing the corres of light.		1	



**Seesaw Mechanism** 

Gell-Mann, Glashow, Minkowski, Mohapatra, Ramond, Senjanovic, Slansky, Yanagida

Extending the SM with SM-singlet right handed neutrino



Flavor eigenstate can be expressed in terms of the mass eigenstate

$$\nu_{\ell} \simeq U_{\ell m} \nu_m + V_{\ell n} N_n$$
PMNS matrix
$$M_D M_N^{-1}$$

$$\mathcal{L}_{mass} = \begin{pmatrix} \overline{\nu_L^c} & \overline{N_R} & \overline{N_L^c} \end{pmatrix} \begin{pmatrix} 0 & m_D & 0 \\ m_D^T & 0 & M \\ 0 & M^T & \mu \end{pmatrix} \begin{pmatrix} \nu_L \\ N_R^c \\ N_L \end{pmatrix}$$
$$m_{\nu} \simeq \mu \frac{m_D^2}{M^2} \quad \textbf{diagonalizing}$$
$$\nu_L \rightarrow \qquad N_R \rightarrow \qquad N_L \rightarrow \leftarrow N_L \quad \leftarrow N_R \leftarrow \nu_L$$
$$m_D \qquad M \qquad \textbf{d} \qquad \textbf{d}$$

$$\nu \simeq \mathcal{N} \nu_{m} + \mathcal{R} N_{m}$$

$$\mathcal{R} = m_{D} m^{-1}$$
Alternative choices
$$\overbrace{\boldsymbol{\epsilon} = \mathcal{R}^{*} \mathcal{R}^{T}}$$

$$m_{\nu} = \mathcal{R} \mu \mathcal{R}^{T} = \frac{1}{M^{2}} m_{D} \mu m_{D}^{T} = U_{\text{MNS}}^{*} D_{\text{NH/IH}} U_{\text{MNS}}^{\dagger}.$$

$$\varepsilon = \frac{1}{2} \mathcal{R}^{*} \mathcal{R}^{T}$$

$$\varepsilon = 2 \mathcal{R}^{$$

### General Parametrization of the neutrino Dirac mass matrix

From the inverse seesaw formula  

$$m_{\nu} = \mu \mathcal{R} \mathcal{R}^{T} = \frac{\mu}{M^{2}} m_{D} m_{D}^{T} = U_{\text{MNS}}^{*} D_{\text{NH/IH}} U_{\text{MNS}}^{\dagger}$$

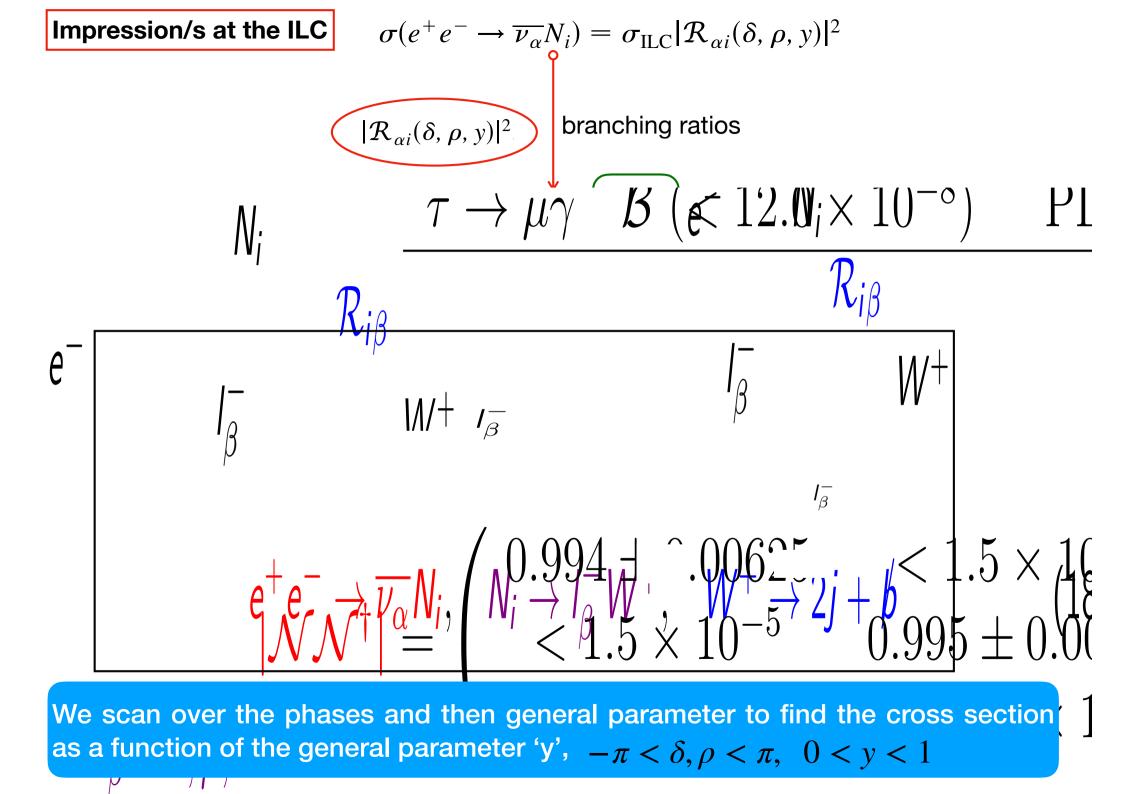
$$m_{D} \text{ carries the off-diagonal entries}$$

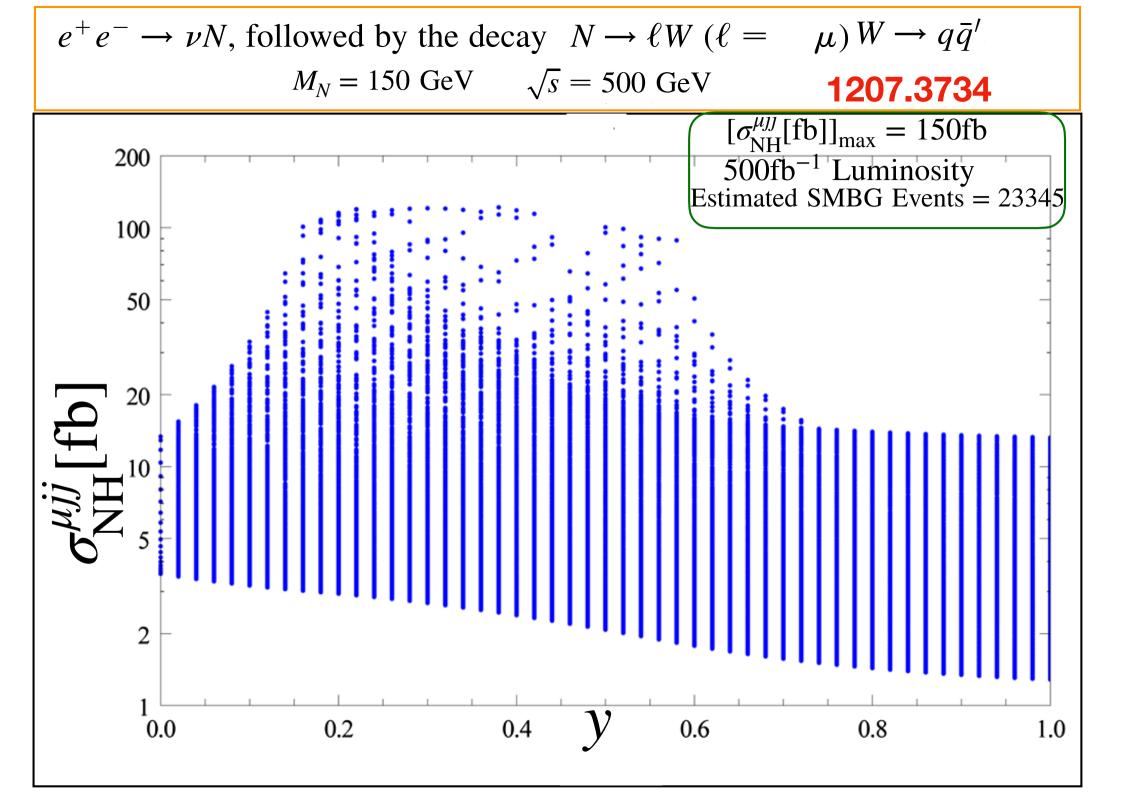
$$\mathcal{R}(\delta, \rho, x, y) = \frac{1}{\sqrt{\mu}} U_{\text{MNS}}^{*} \sqrt{D_{\text{NH/IH}}} O$$

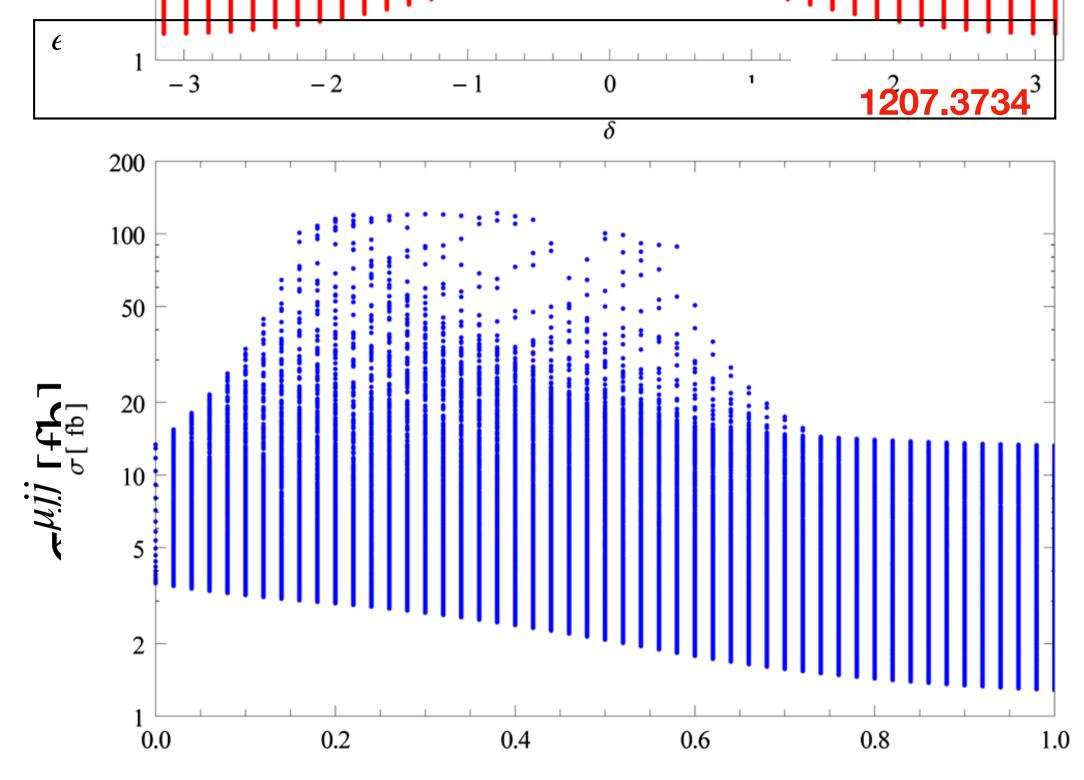
$$\mu, M \text{ diagonal entries}$$

$$\rho = \begin{pmatrix} \cos \alpha & \sin \alpha \\ -\sin \alpha & \cos \alpha \end{pmatrix} = \begin{pmatrix} \cosh y & i \sinh y \\ -i \sinh y & \cosh y \end{pmatrix} \begin{pmatrix} \cos x & \sin x \\ -\sin x & \cosh x \end{pmatrix}$$

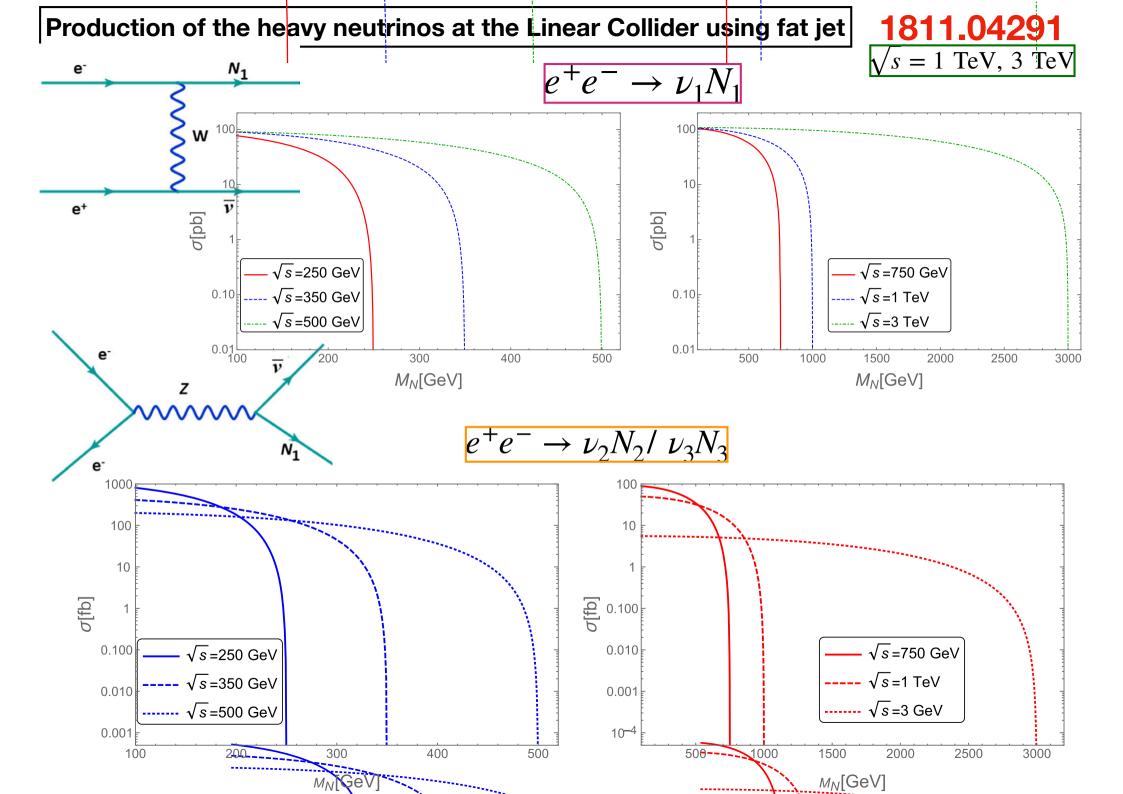
$$\epsilon(\delta, \rho, y) = \mathcal{R}^{*} \mathcal{R}^{T} = \frac{1}{\mu} U_{\text{MNS}} \sqrt{D_{\text{NH/IH}}} O^{*} O^{T} \sqrt{D_{\text{NH/IH}}}^{T} U_{\text{MNS}}^{\dagger}$$
We can parametrize the mixing in terms of the phases and the general parameter considering bounds from non-unitarity

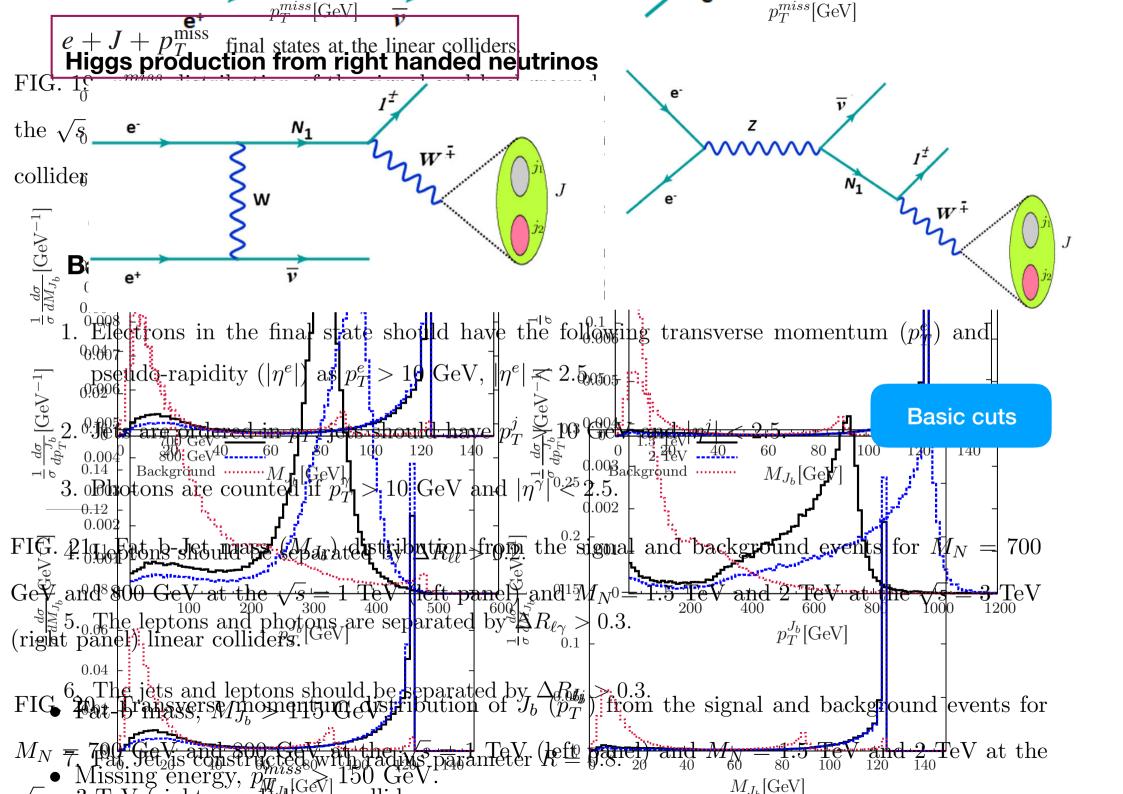


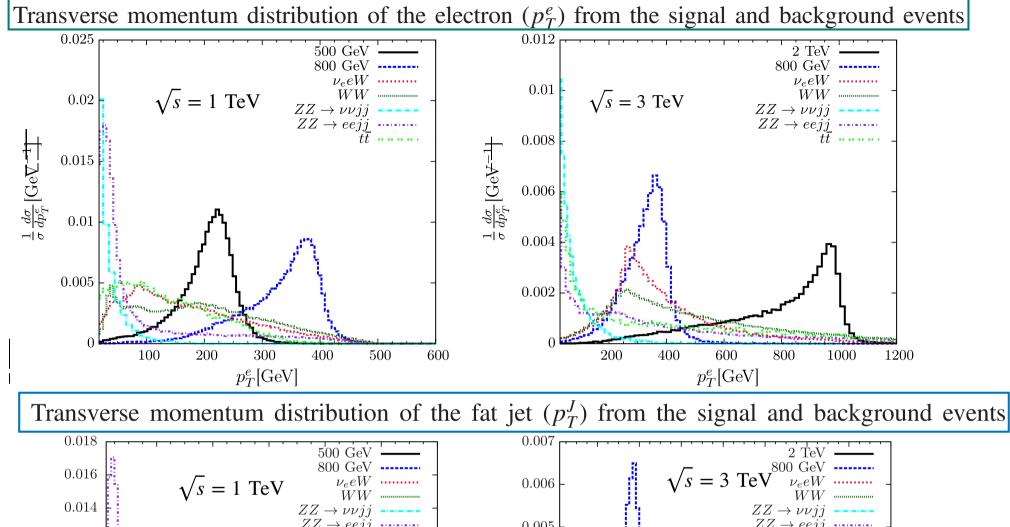


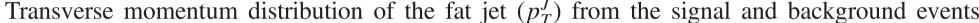


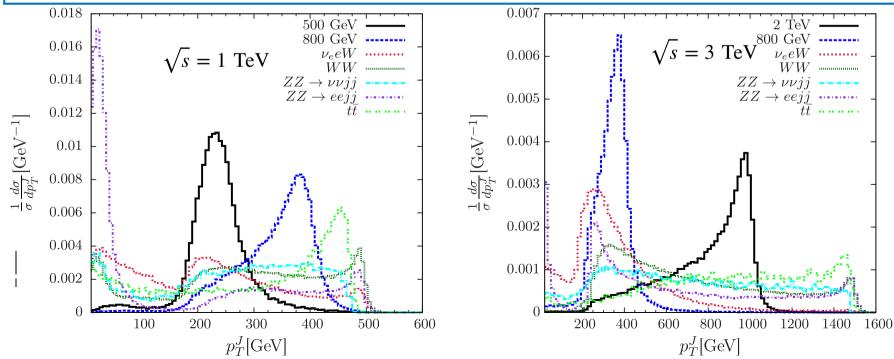
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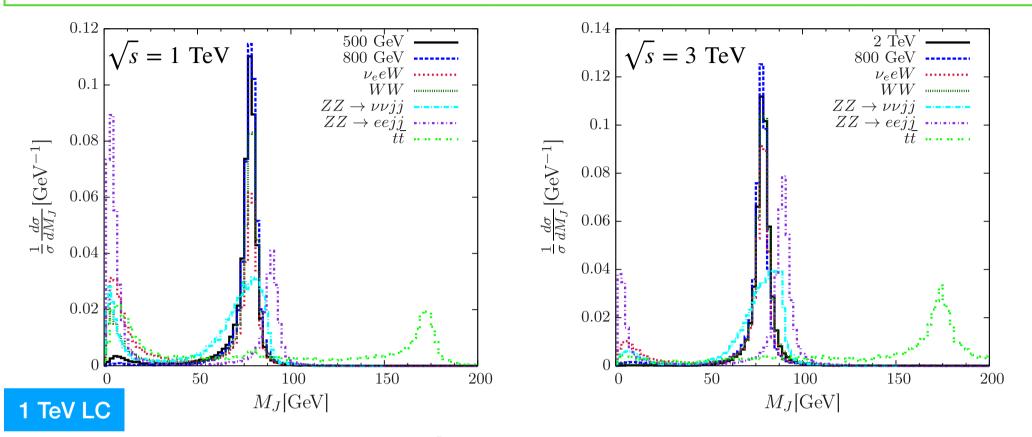






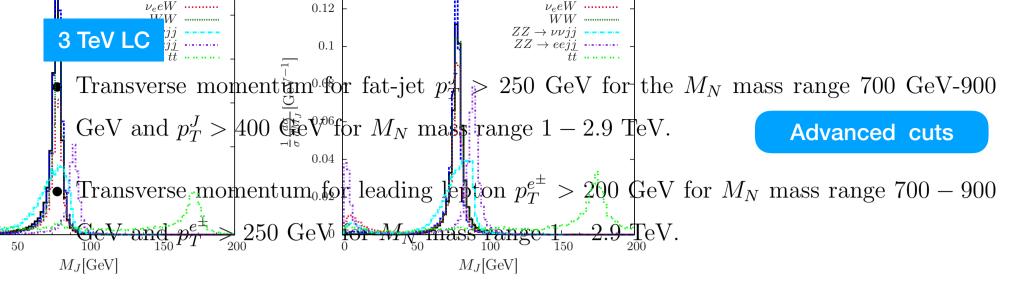


Jet mass  $(M_J)$  distribution of the fat jet from the signal and background events



• Transverse momentum for fat-jet  $p_T^J > 150$  GeV for  $M_N$  mass range 400 GeV-600 GeV and  $p_T^J > 250$  GeV for  $M_N$  mass range 700 GeV-900 GeV. Advanced cuts

- Transverse momentum for leading lepton  $p_T^{e^{\pm}} > 100 \text{ GeV}$  for  $M_N$  mass range 400 GeV-600 GeV and  $p_T^{e^{\pm}} > 200 \text{ GeV}$  for  $M_N$  mass range 700 GeV-900 GeV.
- Polar angle of lepton and fat-jet  $|\cos \theta_e| < 0.85$ ,  $|\cos \theta_J| < 0.85$ .
- Fat-jet mass  $M_J > 70$  GeV.



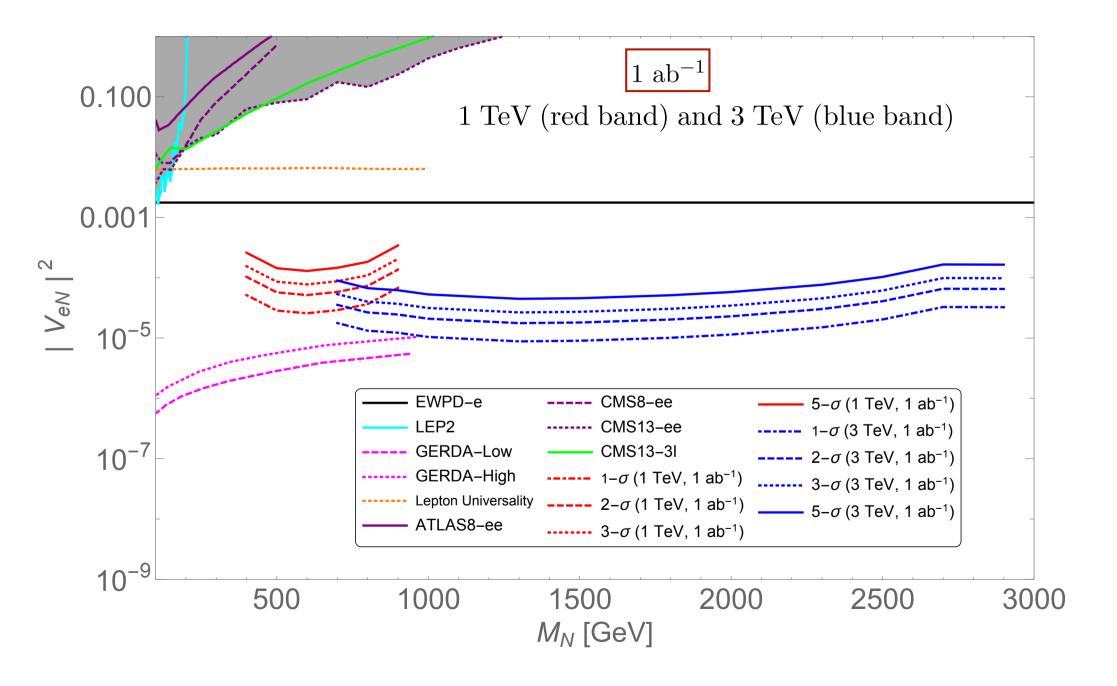
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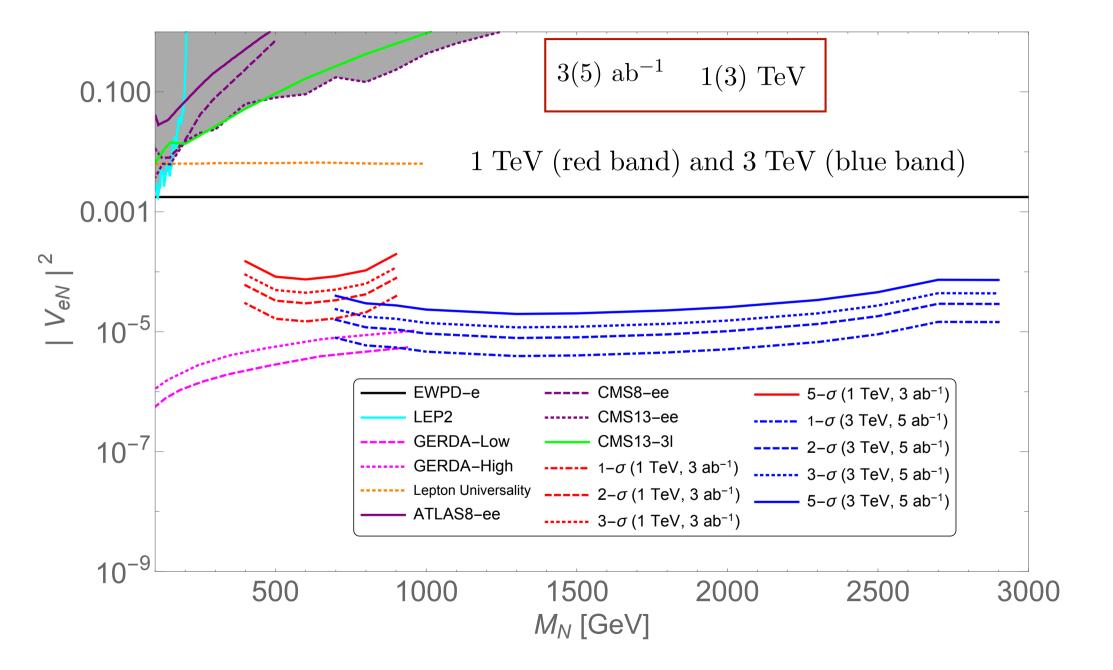
#### 500 GeV @ 1 TeV LC

#### 800 GeV @ 3 TeV LC

Cuts	Signal	]	Total			
		$\nu_e eW$	WW	ZZ	$t\bar{t}$	
Basic Cuts	12,996,200	201,586	72,244	7,200	4,300	285,330
$ \cos \theta_J  \le 0.85$	12,789,800	148,802	44,910	3,800	4,100	201,600
$ \cos \theta_e  \le 0.85$	12,671,800	79,008	40,574	2,800	3,900	126,280
$p_T^J > 150 \text{ GeV}$	12,308,300	70,669	40,490	2,300	3,200	116,660
$M_J > 70 \text{ GeV}$	10,923,100	62,303	37,043	2,100	2,300	103,700
$p_T^\ell > 100{\rm GeV}$	10,714,500	57,076	33,488	1,400	1,530	93,400

	Cuts	Signal	В	Total			
			$\nu_e eW$	WW	ZZ	$t\bar{t}$	
$\left( \right)$	Basic Cuts	21,789,900	$193,\!533$	$12,\!135$	1,361	271	207,301
	$ \cos \theta_J  \le 0.85$	13,599,300	126,980	4,766	406	215	132,367
	$ \cos \theta_e  \le 0.85$	12,163,300	21,110	4,609	390	195	26,304
	$p_T^J > 250~{\rm GeV}$	12,083,500	18,619	4,607	390	189	23,807
	$M_J > 70 { m ~GeV}$	11,287,000	17,442	4,411	385	176	22,416
$\left( \right)$	$p_T^\ell > 200{\rm GeV}$	11,094,300	$16,\!915$	4,108	343	104	21,470



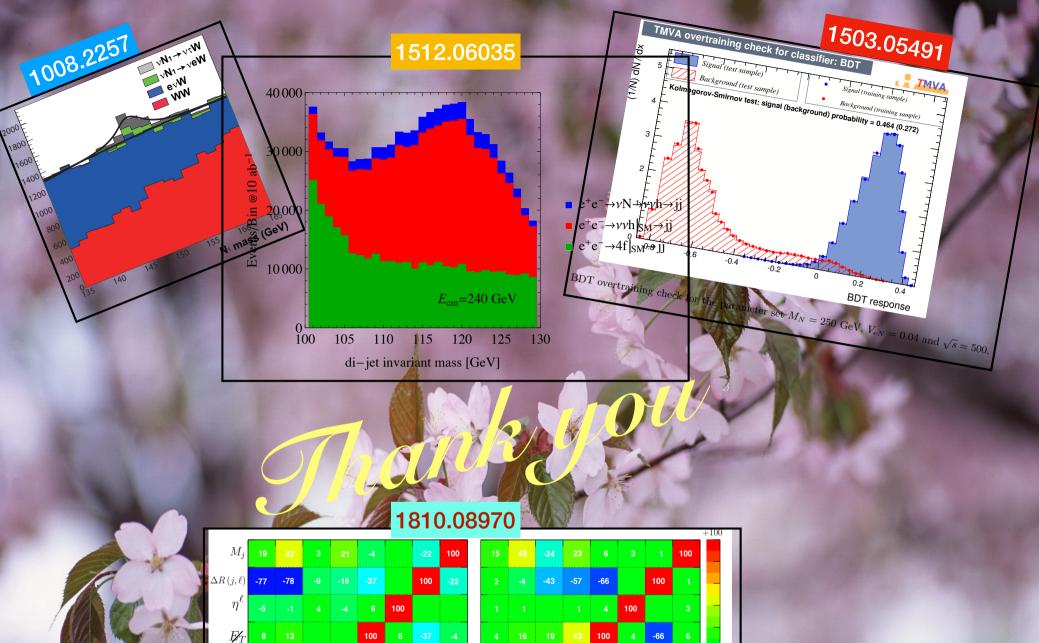


# Conclusions

The neutrino oscillation experiments confirm the existence of the tiny neutrino mass which lead to extend the Standard Model by right handed neutrino/s which mix with the SM neutrinos. There are several interesting simple (and next-to-simple) scenarios with different nice properties which efficiently explain the existence of the neutrino mass some of which which can be tested at the colliders such as at the proposed linear collider (also at the LHC) if the masses of the heavy neutrinos belong to the TeV scale or so.

In this presentation we have concentrated only on the signatures which can motivate the search of such heavy neutrinos at the linear collider. We have used the general parametrization based on the Casas-Ibarra conjecture. Using non-unitarity, neutrino oscillation data and indirect constraints from flavor violation and LEP experiments we have put bounds on the mixing angles to probe the signature of the heavy neutrinos. We have found that using the general parameters one can produce the heavy neutrinos followed by the leading decay modes up to a cross section of 150 fb including the mixing which can reach at a high discovery limit in future.

We have also studied the scenarios where the heavier neutrinos can be discovered at the linear collider from the boosted object search. In that case the heavy neutrinos will dominantly produce the boosted W boson which will further decay into a fat jet. Using jet substructure method we distinguish the signal and background. Finally we show a limit plot in the mass-mixing plane up to 5-sigma which is very strong compared to the EWPD.



$M_{j}$	19	42	3	21	-4		-22	100		15	48	-24	23	6	3	1	100	
$\Delta R\left(j,\ell\right)$	-77	-78			-37		100	-22		2		-43	-57	-66		100	1	
$\eta^{\ell}$	-5	-1		-4		100				1			1	4	100		3	
$\not\!$	8	13			100	6	-37	-4		4			43	100	4	-66	6	
$M_{\rm inv}^{\not \!\!\! Z_T,j}$	8	36	-86	100		-4		21		-33		-12	100	43	1	-57	23	0
$M_{\rm inv}^{{\bf Z}_T,\ell}$	22	-1	100	-86				3		-14	-57	100	-12	19		-43	-24	
$p_T^j$	60	100	-1	36	13		-78	42		43	100	-57	28	16		-4	48	
$p_T^\ell$	100	60	22	8	8	-5	-77	19		100	43	-14	-33	4	1	2	15	
$p_T^{\ell}  p_T^j  M_{\text{inv}}^{\boldsymbol{\mathbb{Z}}_T,\ell}  M_{\text{inv}}^{\boldsymbol{\mathbb{Z}}_T,j}  \boldsymbol{\mathbb{Z}}_T  \eta^{\ell}  \Delta R\left(j,\ell\right)  M_j \qquad p_T^{\ell}  p_T^j  M_{\text{inv}}^{\boldsymbol{\mathbb{Z}}_T,\ell}  M_{\text{inv}}^{\boldsymbol{\mathbb{Z}}_T,j}  \boldsymbol{\mathbb{Z}}_T  \eta^{\ell}  \Delta R\left(j,\ell\right)  M_j  -100$																		