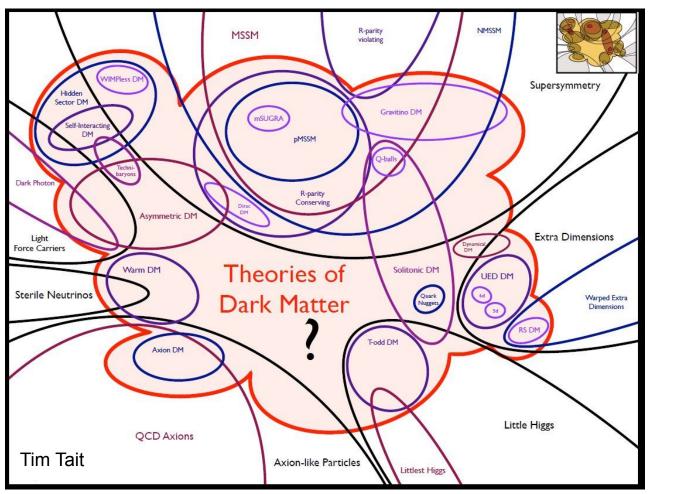
Gaining Insight into PBH Dark Matter

with Compact Stars

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+ PBH!



PBH as DM

- Black holes
 - \circ astrophysical \rightarrow old stars
 - $\circ \quad primordial \rightarrow early \ Universe \ [Zeldovich, Novikov, 1967; Hawking, 1971; Carr, Hawking, 1974]$

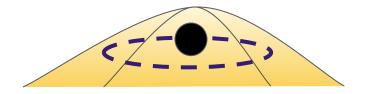
- Why get excited about PBH DM ?
 - no clear signs of particle DM
 - GW astronomy [Bird+ 2016; Sasaki, Thorne+ 1997...] [Riotto...]
 - generic in many BSM models [Fuller, Carr, Dolgov, Kawasaki, Kusenko, Gregory*..]
 - help solve astro puzzles (e.g. seed SMBH [Kusenko, Kawasaki...])
 - o already might appear in standard cosmology (but unlikely)

Many expert talks...

PBH formation

• "Standard" scenario: large perturbations ($\delta \sim 1$) enter horizon \rightarrow collapse

[Kawasaki, Sasaki ...]



- Need to fine tune inflaton potential
 - \rightarrow sensitive to restrictions on field behavior
 - Example: "string swampland conjectures" [Kawasaki, VT, PRD, 2018]

PBH formation

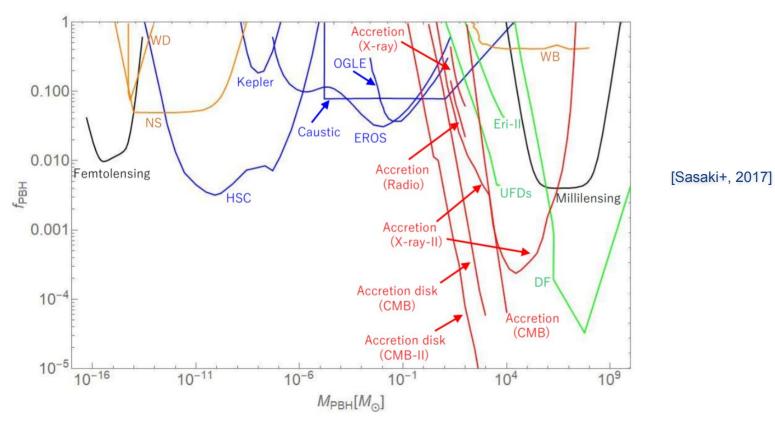
● Attractive new general scenario: scalar field fragmentation → A. Kusenko

[Cotner, Kusenko, PRL, 2016] + [Cotner, Kusenko, VT, PRD, 2018] [Cotner, Kusenko, Sasaki, VT, JCAP, 2019]

... fragmentation could be very generic, in line with lore that gravity is weakest force [Kusenko, VT, Yamada, Yamazaki, 2019]

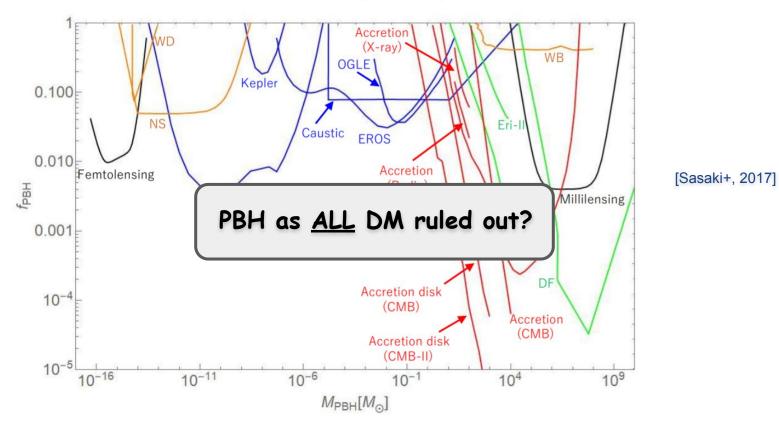
Status of Affairs

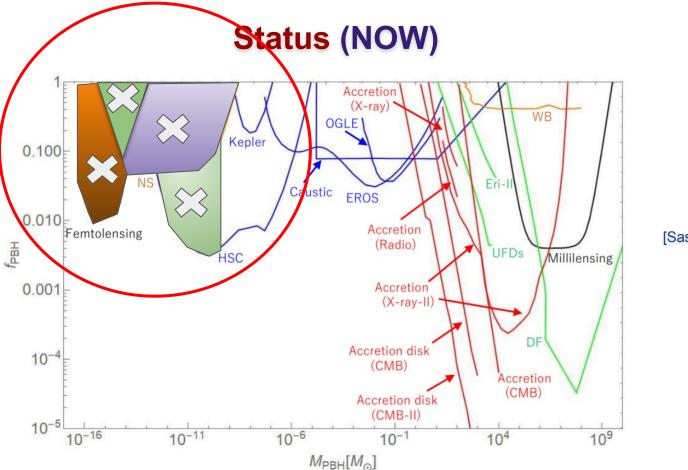
Status (2017)



→ important to consider particle DM + PBH mix [Koushiappas]

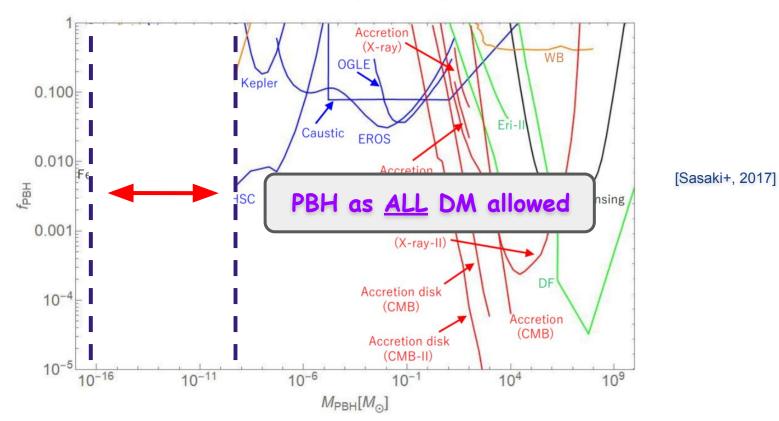
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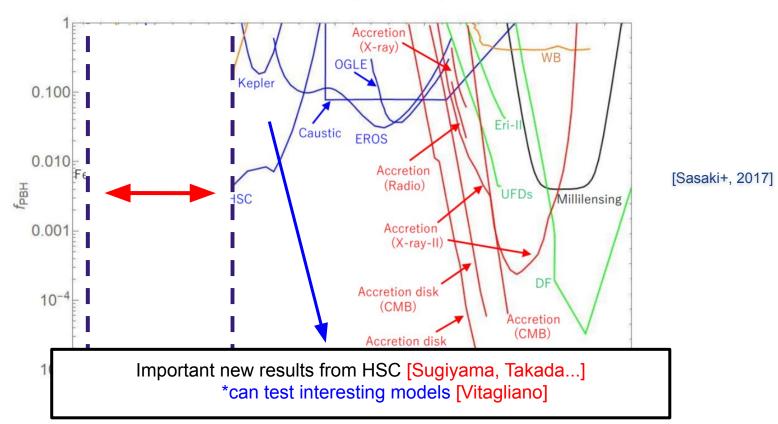


[Sasaki+, 2017]

Status (NOW)



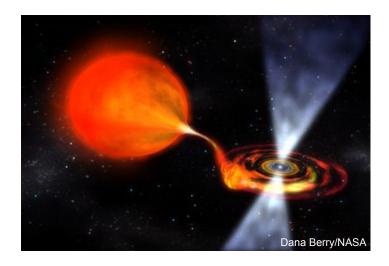
Status (NOW)



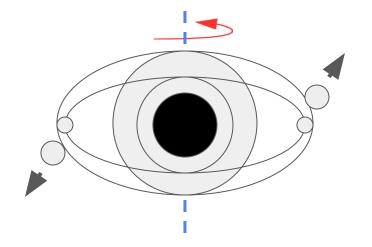
...can we gain some insight into this interesting open region?

Compact stars as PBH laboratories

- Small PBHs can be effectively captured by NS/WD in DM-rich environments (e.g. Galactic Center)
- Captured PBH settle and grow inside, destroy star
 → new signals, potential open problem solutions
 - → r-process nucleosynthesis, 511 keV, FRBs [Fuller, Kusenko, VT, PRL, 2017] + *Viewpoint Highlight* by H.-T. Janka
 - → solar-mass BHs, GRBs, microquasars [VT, PLB, 2017; VT, PLB, 2018]



PBH in millisecond pulsar

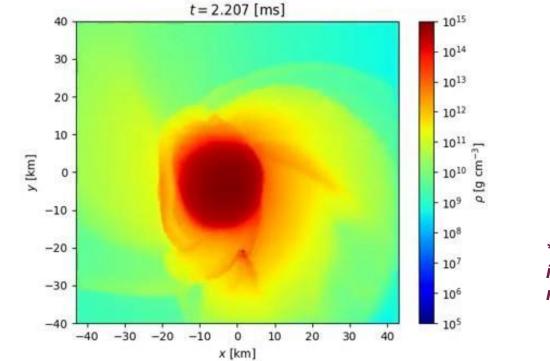


- MSP spins near mass-shedding limit
 → stretched spheroid (analytic Roche lobe model)
- Add PBH: star consumed \rightarrow contracts \rightarrow spins up
- Matter exceeds escape velocity at equator

 \rightarrow neutron-rich material ejecta

[Fuller, Kusenko, VT, PRL, 2017]

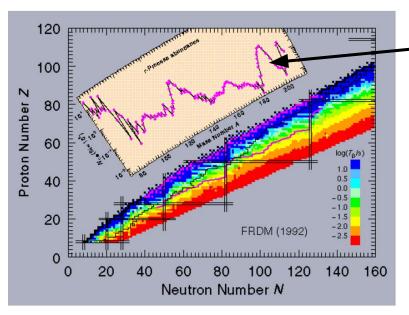
Neutron-rich material emission



preliminary simulations by David Radice (Penn State)

*** sensitive to input conditions, need more studies

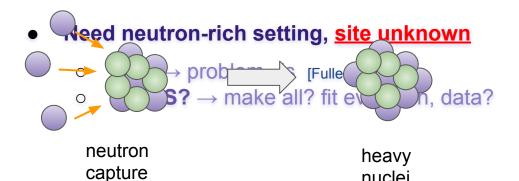
R-process nucleosynthesis: stellar gold factories



Los Alamos, Nuclear Data Group



- Neutrons *r*apidly captured before \Box -decay
- Main producer of heavy astro-elements



nuclei

Making gold with black holes

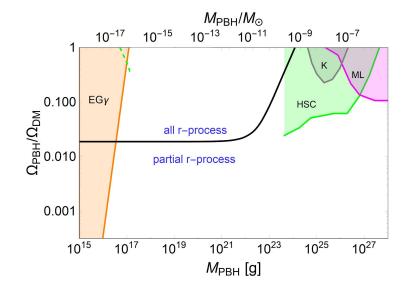
Heavy element abundance

- Milky Way contains 10⁴ M_o
- UFDs 1 in 10 (Reticulum II) shows

 \rightarrow can explain with PBH-NS !

EXCESS [Ji+, Nature 2016]





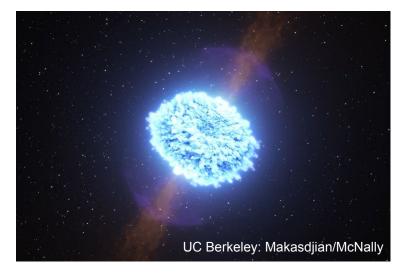
[Fuller, Kusenko, VT, PRL, 2017]

PBH-NS laboratory: orphan kilonovae

Kilonova: afterglow from ejecta

PBH-NS vs. mergers

→ "orphan kilonova" (w/o merger GWs)

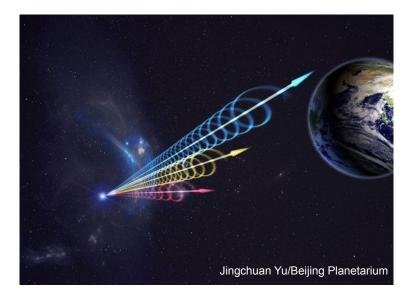


[Fuller, Kusenko, VT, PRL, 2017]

PBH-NS laboratory: fast radio bursts

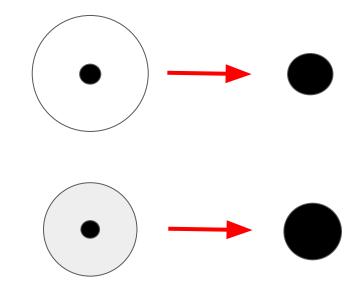
• Fast radio bursts: brief radio pulse

- origin unknown
- >100 found, few repeaters [Amiri+, Nature 2019]
- Release % of NS magnetic field energy due to PBH-NS consumption as radio
 - → non-repeating FRB



PBH-NS/WD laboratory: new solar-mass BHs

- No astro BHs ≤ 2 M₀
- PBH + NS/WD: new ~0.5-2 M
 BHs
- Novel double binary signals possible (e.g. double kilonova)
- Small population of solar-mass BHs made in late Universe, on top of asteroid-mass PBH DM



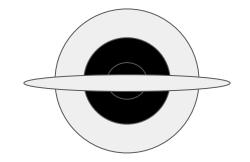
PBH-NS laboratory: orphan GRBs

• "Standard" short gamma-ray burst progenitor: BH + disk

 \rightarrow disk accreted, binding energy released

• If disk forms, could be from PBH-NS

 \rightarrow "orphan GRB" (w/o merger GWs)



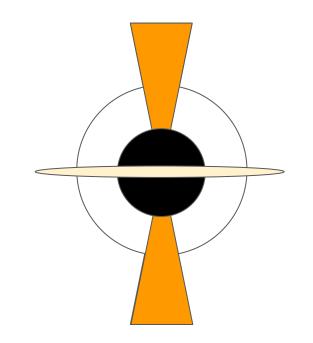
[VT, PLB, 2018]

PBH-WD laboratory: baby microquasars

• WDs have non-relativistic jets $L_{\rm jet} \sim \frac{1}{R}$

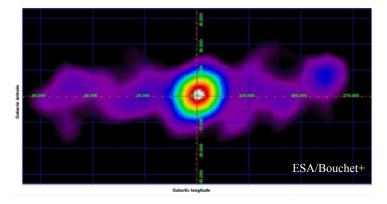
- WD + PBH → solar-mass BH accretor
 - o radius ♥, luminosity ⁰

→ continuous jet *"baby" microquasar*



511 keV Galactic Center excess

- Extensive observations (SPI/INTEGRAL) show Galactic Center shines in 511 keV γ-rays
 - consistent w/ e+ annihilation [Beacom, Yuksel, 2006]
 - source unknown



 \rightarrow can be explained by ejecta production from PBH-NS!

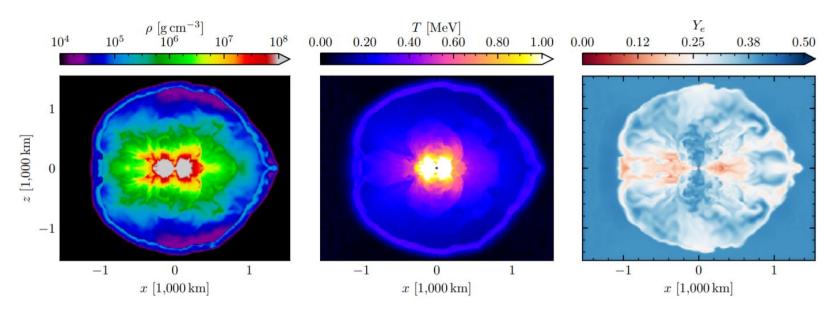
[Fuller, Kusenko, VT, PRL, 2017]

Novel generic signal of neutron star mergers

511 keV radiation

[Fuller, Kusenko, Radice, VT, PRL, 2019]

Start with NS-NS simulations

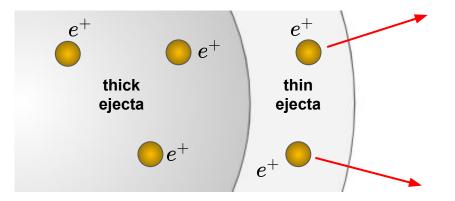


t = 10 ms

[Fuller, Kusenko, Radice, VT, PRL, 2019]

Positron production

- Expanding ejecta heated to \sim MeV \rightarrow lots of thermal positrons produced
- Magnetic confinement in ejecta not perfect
 → some positrons escape from *"optically thin"* outer layers

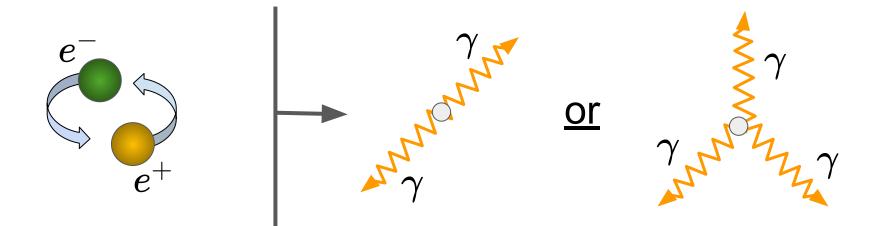


[Fuller, Kusenko, Radice, VT, PRL, 2019]

511 keV radiation

Escaping ~MeV positrons annihilate via positronium bound state formation ✓

 (as desired for GC excess) → 511 keV radiation



[Fuller, Kusenko, Radice, VT, PRL, 2019]

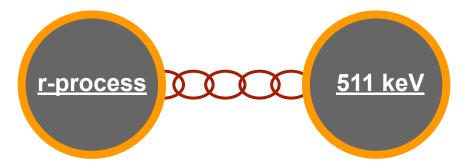
Galactic Center emission

• Take LIGO merger rates \rightarrow 511 keV emission consistent with GC excess

NS binary kicks → expect some signal in Galactic disk, not only bulge
 ○ consistent with GC excess, <u>difficulty for other proposals</u>

A smoking gun signal !

• Proposal directly links r-process and 511 keV



• Observations of Reticulum II dwarf galaxy show heavy elements AND 511 keV [Ji, Frebel+, Nature, 2016; Siegert+ 2016]

 \rightarrow new smoking gun signal of merger emission !

[Fuller, Kusenko, Radice, VT, PRL, 2019]

Summary

• Renaissance era in PBH research → synergy with multi-messenger astronomy

• Compact stars as PBH laboratories

- new observables to study DM
- new venues to pursue open problems

Astroparticle physics naturally connects early Universe and observations