

Primordial black hole tower: Dark matter, earth-mass, and LIGO black holes

Yuichiro TADA (Nagoya U.) w/S. Yokoyama PRD 100, no. 2, 023537 (2019) Inomata, Kawasaki, Kusenko, Mukaida, Yanagida 2016,2017

Primordial Black Hole : well-known to you

- PBH formation in RD $\zeta_{th} \sim 1$

PBH tower in multi-phase inflation

$\zeta_{\rm th} \sim 1$ \checkmark $\mathcal{P}_{\zeta}(k_{\rm PBH}) \sim 10^{-2}$

Yuichiro Tada





- Dynamical Friction
- Accretion Y
- Prim. PTB
- Hawking Y

Yuichiro Tada



Massive than stellar BHs found

small spin

nLQ		Event	$m_1/{ m M}_{\odot}$	$m_2/{ m M}_{\odot}$	${\cal M}/{ m M}_{\odot}$	χ eff
	-	GW150914	$35.6^{+4.8}_{-3.0}$	$30.6^{+3.0}_{-4.4}$	$28.6^{+1.6}_{-1.5}$	-0.01
		GW151012	$23.3^{+14.0}_{-5.5}$	$13.6^{+4.1}_{-4.8}$	$15.2^{+2.0}_{-1.1}$	0.04^{+0}_{-0}
		GW151226	$13.7^{+8.8}_{-3.2}$	$7.7^{+2.2}_{-2.6}$	$8.9^{+0.3}_{-0.3}$	0.18^{+0}_{-0}
	-	GW170104	$31.0^{+7.2}_{-5.6}$	$20.1^{+4.9}_{-4.5}$	$21.5^{+2.1}_{-1.7}$	-0.04
LSS		GW170608	$10.9^{+5.3}_{-1.7}$	$7.6^{+1.3}_{-2.1}$	$7.9^{+0.2}_{-0.2}$	0.03_
		GW170729	$50.6^{+16.6}_{-10.2}$	$34.3^{+9.1}_{-10.1}$	$35.7^{+6.5}_{-4.7}$	0.36_
	-	GW170809	$35.2^{+8.3}_{-6.0}$	$23.8^{+5.2}_{-5.1}$	$25.0^{+2.1}_{-1.6}$	0.07^{+0}_{-0}
		GW170814	$30.7^{+5.7}_{-3.0}$	$25.3\substack{+2.9\\-4.1}$	$24.2^{+1.4}_{-1.1}$	0.07^{+0}_{-0}
	-	GW170817	$1.46^{+0.12}_{-0.10}$	$1.27^{+0.09}_{-0.09}$	$1.186^{+0.001}_{-0.001}$	0.00^{+0}_{-0}
		GW170818	$35.5_{-4.7}^{+7.5}$	$26.8^{+4.3}_{-5.2}$	$26.7^{+2.1}_{-1.7}$	-0.09
µ-dist		GW170823	$39.6^{+10.0}_{-6.6}$	$29.4_{-7.1}^{+6.3}$	$29.3^{+4.2}_{-3.2}$	0.08^{+0}_{-0}
10 ⁴⁵	10 ⁵⁰	LIGN/\/irgo 2018				

LIGO/Virgo 2018

4

Yuichiro Tada

10¹⁵

10¹⁰

mLQ

µ-dist



+0.12-0.13 +0.28-0.19 +0.20-0.12+0.17 -0.20+0.19-0.07+0.21-0.25+0.16-0.16+0.12-0.11 +0.02-0.01 3+0.18-0.21 +0.20-0.22



Yuichiro Tada



Yuichiro Tada

Motivations of PBH



PBH tower in multi-phase inflation

arXiv: 1909.11090

What if Planet 9 is a Primordial Black Hole?

Jakub Scholtz¹ and James Unwin²

Yuichiro Tada

¹Institute for Particle Physics Phenomenology, Durham University, Durham, DH1 3LE, United Kingdom ²Department of Physics, University of Illinois at Chicago, Chicago, IL 60607, USA; & Department of Physics, University of California, Berkeley & Theoretical Physics Group, LBNL & Mathematics Sciences Research Institute, Berkeley, CA 94720, USA

1. Introduction. As of this year, two gravitational anomalies of similar mass but very different origins remain to be explained. First, there is a growing body of observational anomalies connected to the orbits of trans-Neptunian Objects (TNOs) [1–3]. These observations have been taken as evidence of a new ninth planet in our solar system, called Planet 9 (P9), with mass $M_9 \sim 5 - 15 M_{\oplus}$ and orbiting around the Sun at a distance of 300–1000 AU [4]. Second, there is set of gravitational anomalies recently observed by the Optical Gravitational Lensing Experiment (OGLE). OGLE reported an excess of six ultrashort microlensing events with crossing times of 0.1 - 0.3 days [5]. The lensing objects are located towards the galactic bulge, roughly 8kpc away. These events correspond to lensing by objects of mass $M \sim 0.5 M_{\oplus} - 20 M_{\oplus}$ [6] and could be interpreted as an unexpected population of free floating planets (FFPs) or as Primordial Black Holes (PBHs).

SUPPLEMENTARY MATERIAL

A. SIZE OF THE PBH

The Schwarzschild radius of a black hole is given by

$$r_{\rm BH} = rac{2GM_{\rm BH}}{c^2} \simeq 4.5 {\rm cm} \left(rac{M_{\rm BH}}{5M_\oplus}
ight) \;.$$

In Figure 1 we provide an exact scale image of a $5M_{\oplus}$ PBH. The associated DM halo however extends to the stripping radius $r_{t,\odot} \sim 8$ AU, this would imply a DM halo which extends roughly the distance from Earth to Saturn (both in real life and relative to the image)







How can we realize these HIERARCHICAL mass spectra?

PBH tower in multi-phase inflation

Yuichiro Tada

8 /17 ____



Yuichiro Tada



 $\mathcal{P}_{\mathcal{R}} \sim 10^{-2}$

10 /17

Yuichiro Tada

Double Inflation



PBH tower in multi-phase inflation

Kumekawa, Moroi, Yanagida 1994 Izawa, Kawasaki, Yanagida 1997 Kawasaki, Sugiyama, Yanagida 1998

not short inflation $\mathbf{1}$ large PTB at the onset

11 /17

Yuichiro Tada

 χ

Extreme Case



PBH tower in multi-phase inflation



start phase-(i+1)

12 /17 ____

Yuichiro Tada



Yuichiro Tada

Testability



PBH tower in multi-phase inflation

Yuichiro Tada



14 /17

Implication to String Theory?

dS swampland conjecture Ooguri & Vafa+ 2018

"dS vacua will be unstable in UV-complete theories"

$$\epsilon_V = \frac{M_{\rm Pl}^2}{2} \left(\frac{V'}{V}\right)^2 \gtrsim \mathcal{O}(1) \quad \text{or} \quad \eta_V = M_{\rm Pl}^2 \frac{V''}{V} \lesssim -\mathcal{O}(1)$$

each infl. phase CANNOT continue lon

- CMB scale? PTB by inflaton



PBH tower in multi-phase inflation

60 e-folds in total by multi-phase

$$r_{\rm CMB}) = n_{\rm s} - 1 \simeq -6\epsilon_V + 2\eta_V \sim \mathcal{O}(1)$$

Yuichiro Tada



Implication to String Theory?

dS swampland conjecture Ooguri & Vafa+ 2018

"dS vacua will be unstable in UV-complete theories"

$$\epsilon_V = \frac{M_{\rm Pl}^2}{2} \left(\frac{V'}{V}\right)^2 \gtrsim \mathcal{O}(1) \quad \text{or} \quad \eta_V = M_{\rm Pl}^2 \frac{V''}{V} \lesssim -\mathcal{O}(1)$$

each infl. phase CANNOT continue long

- CMB scale? PTB by curvaton Kogai, YT in prep.

$$-0.035 \simeq \frac{d \log \mathscr{P}_{\zeta}}{d \log k} (k_{\text{CMB}}) = n_{\text{s}} - 1 \simeq -2\epsilon_{V} + \frac{2}{3} \frac{m_{\sigma}^{2}}{H^{2}}$$
Planck 2018

PBH tower in multi-phase inflation

60 e-folds in total by multi-phase

Yuichiro Tada



Conclusions

interesting mass regions for PBH are hierarchical
 multi-phase inflation can realize them simultaneously
 cf. dS swampland conjecture may support multi-phase inflation
 testable by GW