

RECENT ADVANCES
in the
BLACK HOLE INFORMATION
PARADOX

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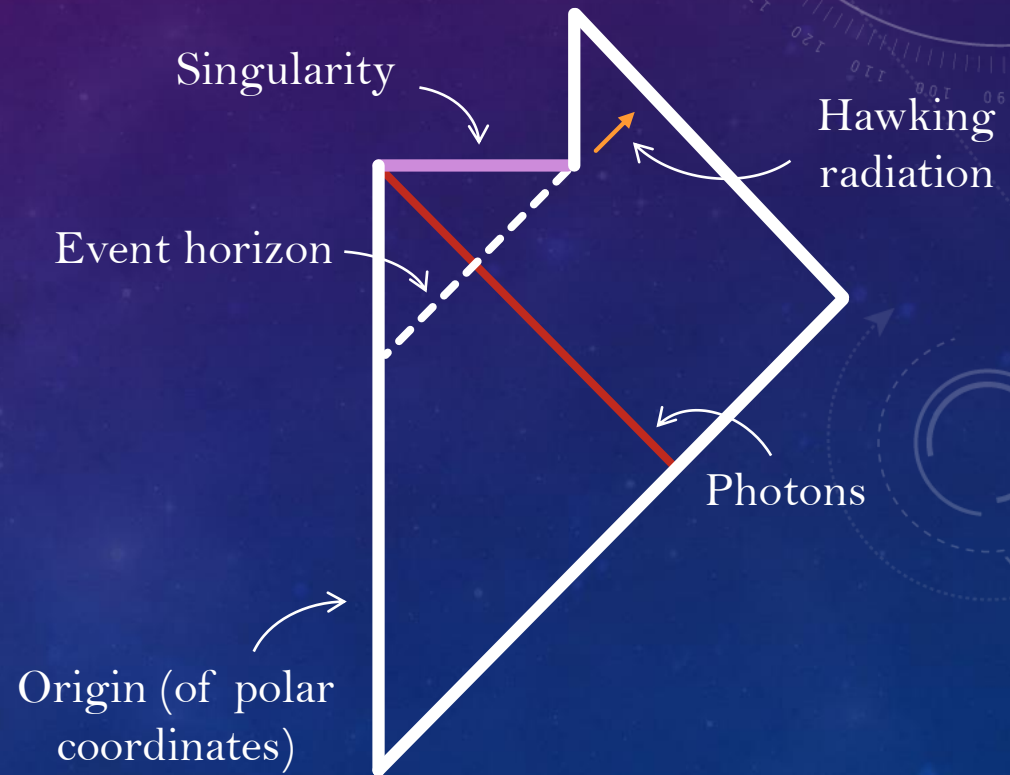
Berkeley Week, IPMU

January 17, 2020



THE INFORMATION PARADOX: BLACK HOLE EVAPORATION

Hawking (1975) showed that, semiclassically, black holes should emit radiation and eventually evaporate. This radiation comes out in a *mixed state*.



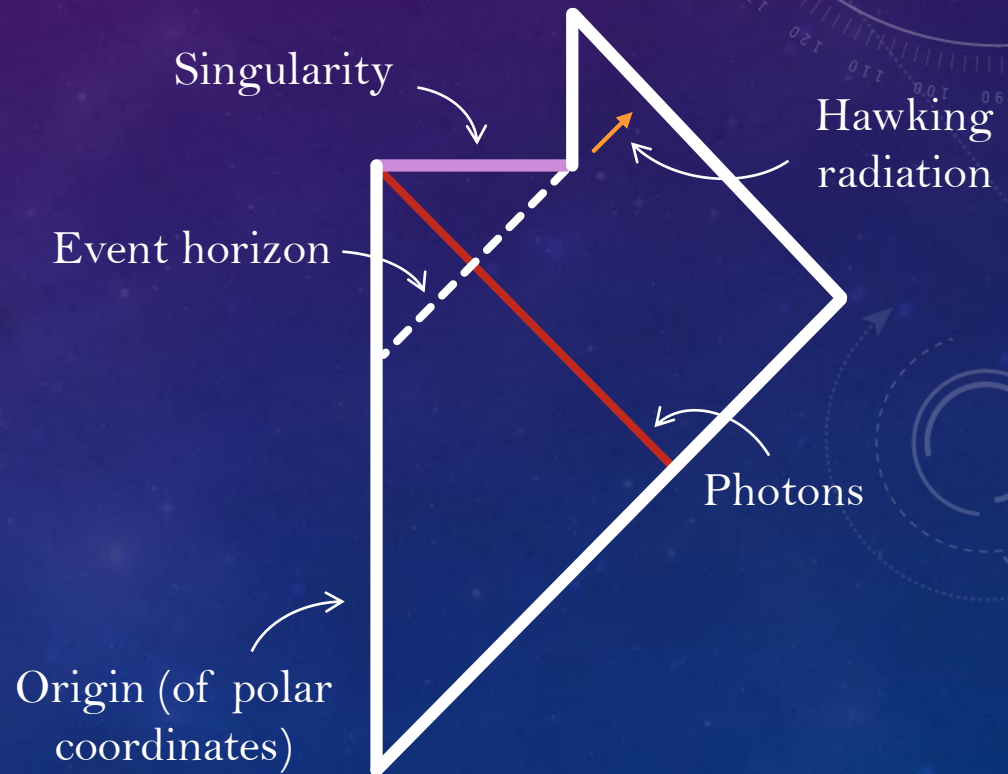
THE INFORMATION PARADOX: BLACK HOLE EVAPORATION

The Minkowski vacuum can be decomposed along any plane into an entangled pure state:

$$|\Omega\rangle = \frac{1}{\sqrt{Z}} \sum_i e^{-\pi\omega_i} |i^*\rangle_A |i\rangle_B$$

Tracing out region A \rightarrow thermal state in region B:

$$\rho_B = \frac{1}{Z} \sum_i e^{-2\pi\omega_i} |i\rangle_B \langle i|$$



THE INFORMATION PARADOX: THE S-MATRIX



$$P(\text{out}|\text{in}) = |\langle \text{out} | S | \text{in} \rangle|^2$$

If the **S-Matrix**
is unitary...



THE INFORMATION PARADOX: A STATEMENT OF THE PARADOX

1. The (unitary) S-matrix maps pure in-states to pure out-states. In other words, no information is lost.
2. We expect a process in asymptotically-flat quantum gravity to be described by an S-matrix.
3. According to Hawking's calculation, a black hole formed from matter in a pure state will evaporate into radiation in a mixed state.
4. If a black hole is formed from matter in a pure state and evaporates into a mixed state, the process cannot be described by a unitary S-matrix. The whole process involves information loss.

WHY FAVOR UNITARITY NONETHELESS?

1. We think QFT works in weakly-gravitating regions.

2. The AdS/CFT Correspondence.

THE ADS/CFT CORRESPONDENCE

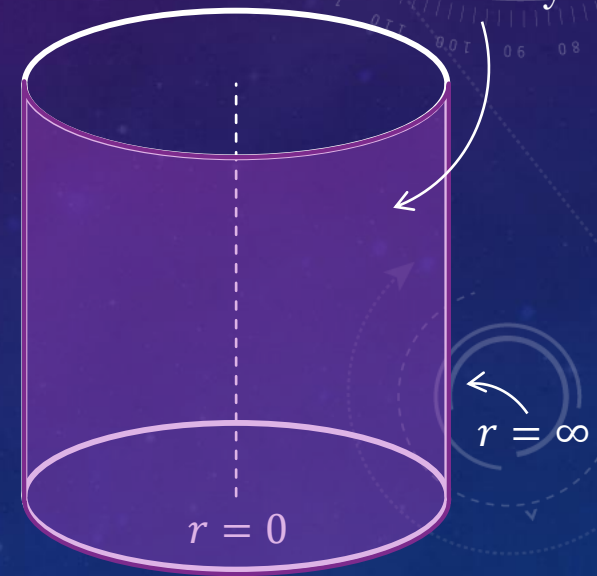
Quantum Gravity
in AdS_{d+1}



Non-Gravitational
 CFT_d

Anti de Sitter space:

Maximally symmetric
solution to Einstein's
equations with negative
cosmological constant

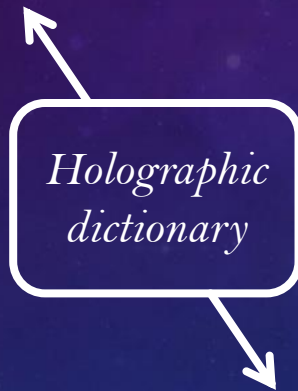


AdS Space

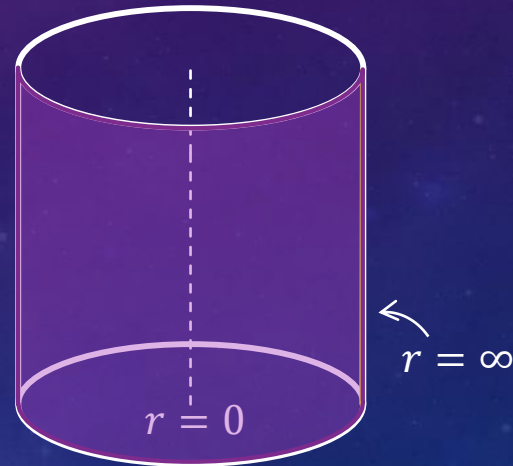
*CFT is unitary, so black hole evaporation
should be unitary*

THE AdS/CFT CORRESPONDENCE

Quantities in
the bulk



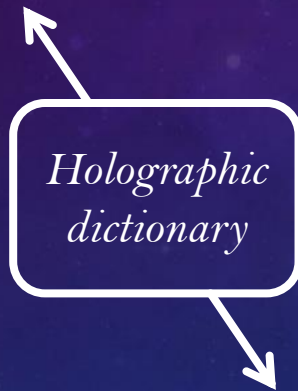
Quantities on
the boundary



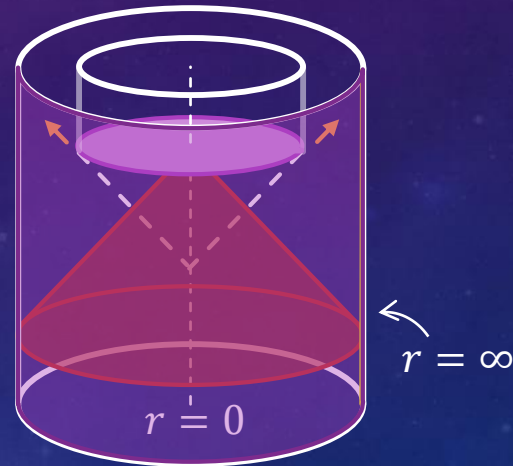
AdS Space

THE ADS/CFT CORRESPONDENCE

Quantities in
the bulk



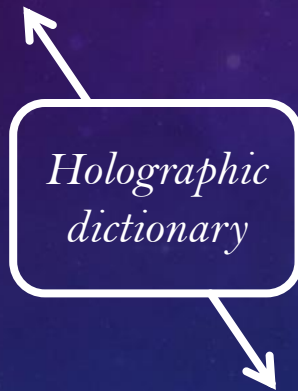
Quantities on
the boundary



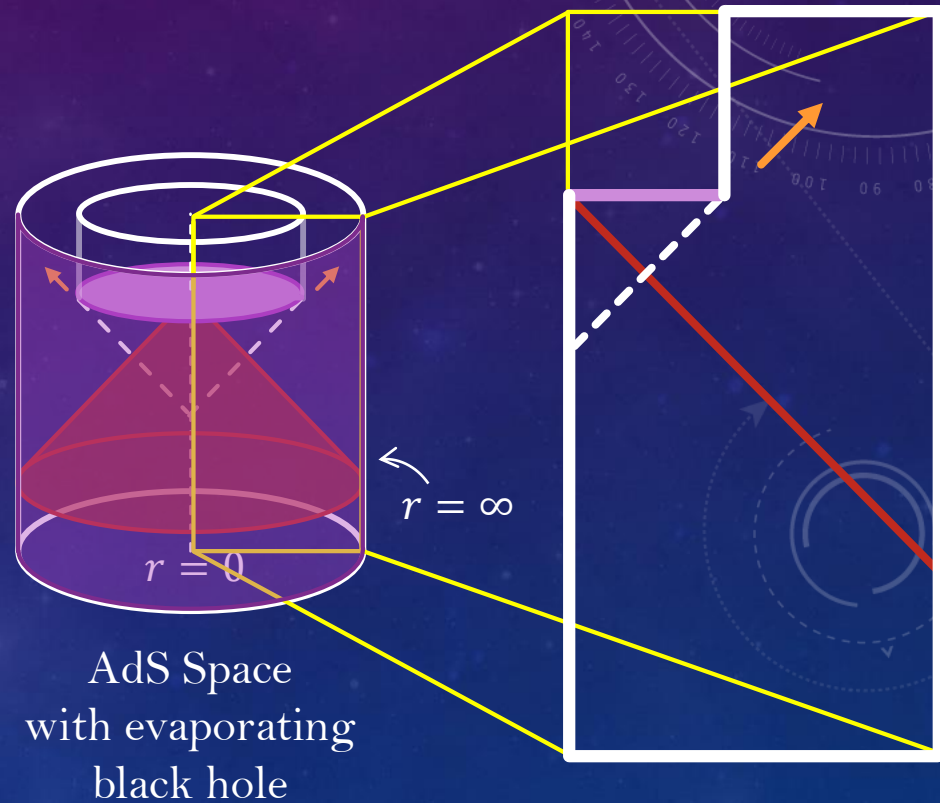
AdS Space
with evaporating
black hole

THE ADS/CFT CORRESPONDENCE

Quantities in
the bulk

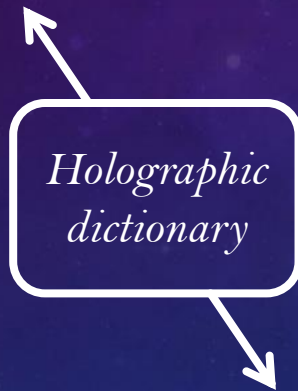


Quantities on
the boundary



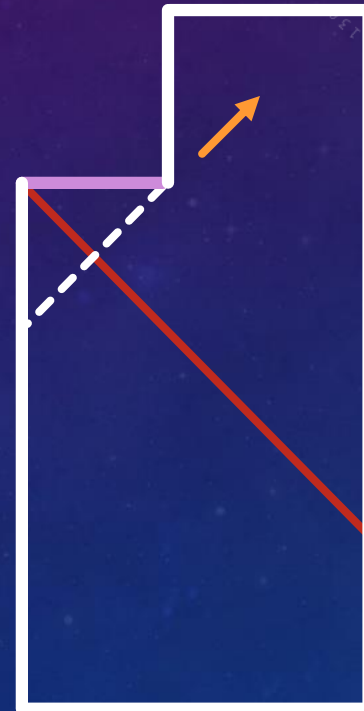
THE AdS/CFT CORRESPONDENCE

Quantities in
the bulk



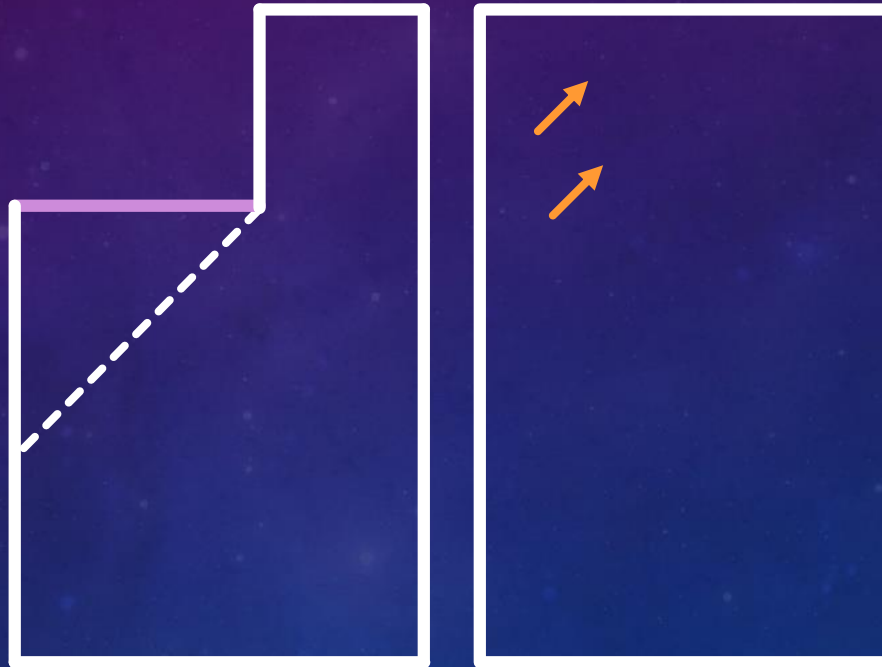
Quantities on
the boundary

Evaporating
black hole in
AdS space



A RECENT LOOK AT EVAPORATING BH'S: THE SETUP

JT (1+1D)
gravity with
Hawking State



Auxiliary
System

arXiv:1905.08255

arXiv:1905.08762

TRACKING INFORMATION IN BLACK HOLE EVAPORATION

**Von Neumann
Entropy:**

$$S(\rho) \equiv -\text{tr}(\rho \log \rho)$$



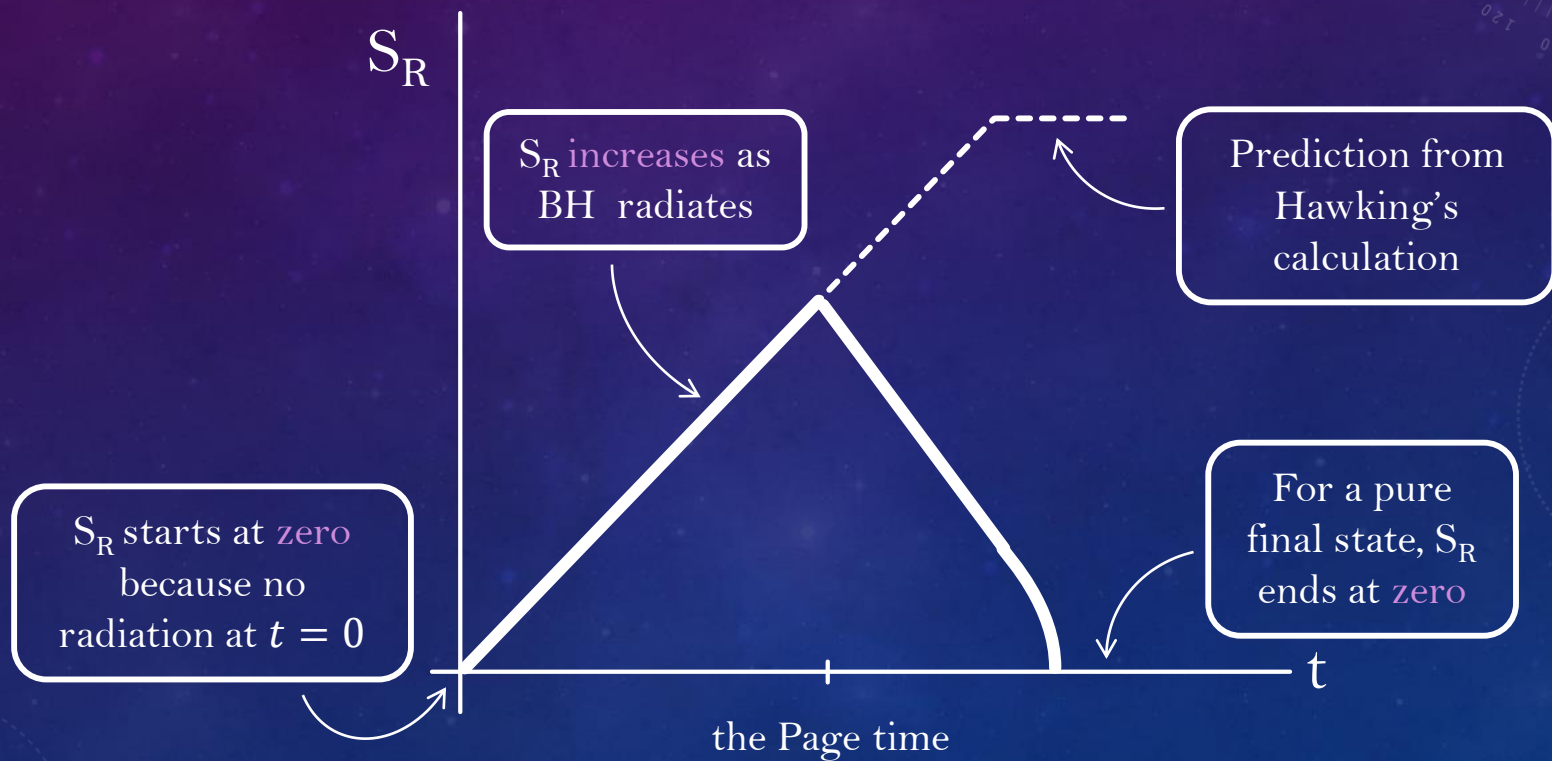
Entanglement Entropy:

$$S_A \equiv -\text{tr}(\rho_A \log \rho_A)$$

$$\langle a' | \rho_A | a \rangle \equiv \sum_b \langle a' b | \rho_{AB} | ab \rangle$$

**Pure total state
implies $S_A = S_B$**

THE PREDICTION OF UNITARY EVAPORATION: THE PAGE CURVE



A USEFUL HOLOGRAPHIC RELATIONSHIP: THE (H)RT PRESCRIPTION AND QUANTUM EXTREMAL SURFACES

Generalized Entropy:

$$S_{\text{gen}}[\sigma] = \frac{\mathcal{A}(\sigma)}{4G\hbar} + S[\text{Ext}(\sigma)]$$

**Quantum Extremal
Surface (QES):**

A surface whose generalized entropy is stationary with respect to all deformations.

A USEFUL HOLOGRAPHIC RELATIONSHIP: THE (H)RT PRESCRIPTION AND QUANTUM EXTREMAL SURFACES

Generalized Entropy:

$$S_{\text{gen}}[\sigma] = \frac{\mathcal{A}(\sigma)}{4G\hbar} + S[\text{Ext}(\sigma)]$$



Entanglement
Wedge of A

Quantum (H)RT
Prescription:

$$S_{\text{CFT}}[A] = S_{\text{gen}}[\gamma_{\text{RT}}[A]],$$

where γ_{RT} is the QES
homologous to A with the
smallest S_{gen} .

Quantum Extremal
Surface (QES):

A surface whose
generalized entropy is
stationary with respect to
all deformations.



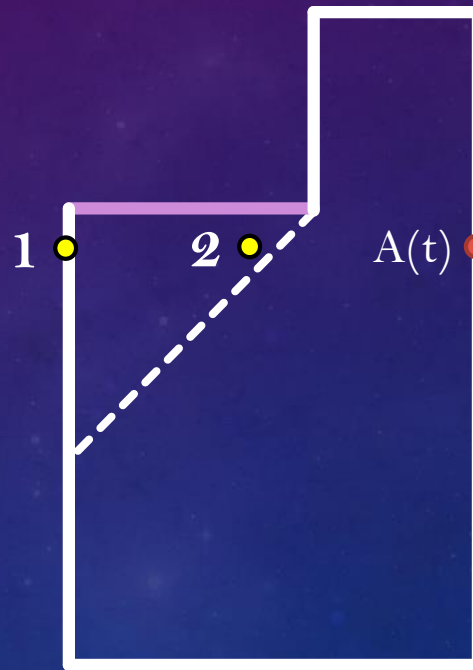
A RECENT LOOK AT EVAPORATING BH'S: THE CALCULATION

General Approach:

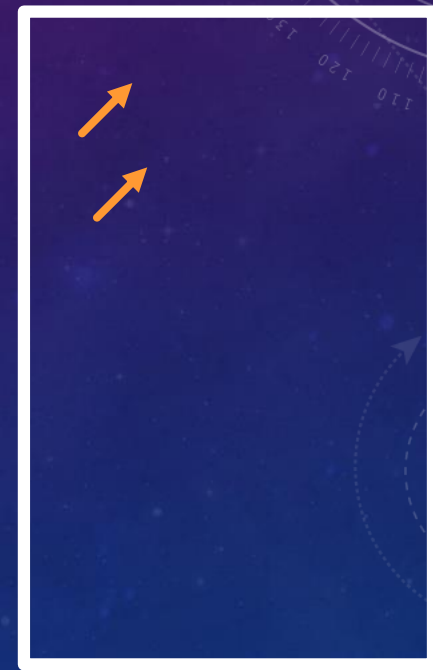
Taking the boundary region (A) to be the whole boundary at a particular time (t), find the QES's.

The QES's:

1. The empty surface
2. $\gamma(t)$



JT gravity with
Hawking State



Auxiliary System

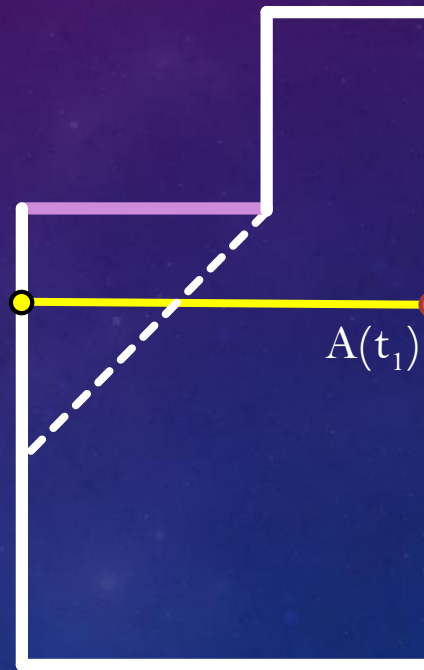
A RECENT LOOK AT EVAPORATING BH'S: THE CALCULATION

Before the Page Time:

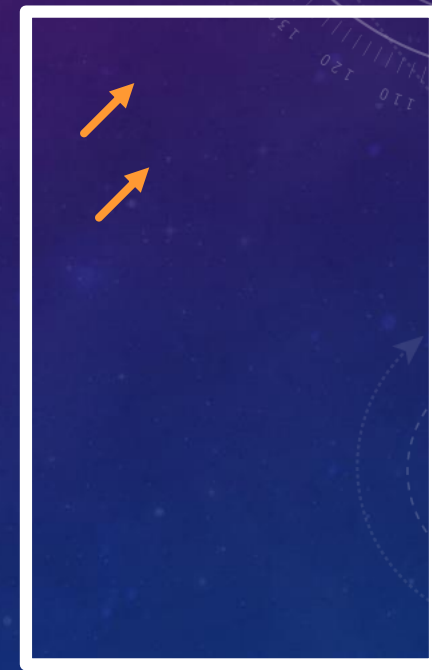
Minimal QES = empty surface

$$S_{\text{CFT}} = 0 + S_{\text{bulk}} = S_{\text{interior}}$$
$$S_{\text{interior}} = S_{\text{R}}$$

S_{CFT} increases
monotonically
before the Page
time



JT gravity with
Hawking State



Auxiliary System

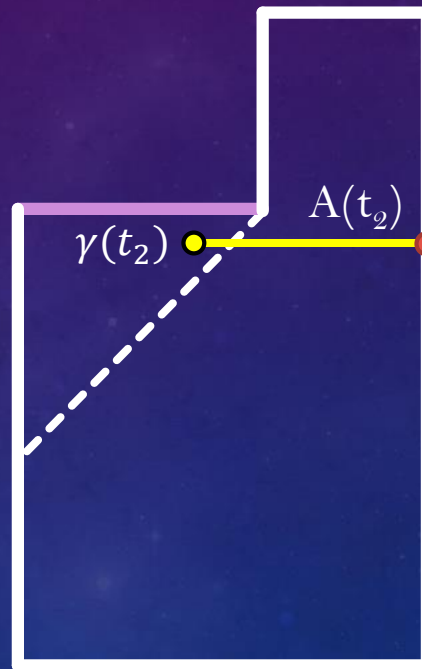
A RECENT LOOK AT EVAPORATING BH'S: THE CALCULATION

After the Page Time:

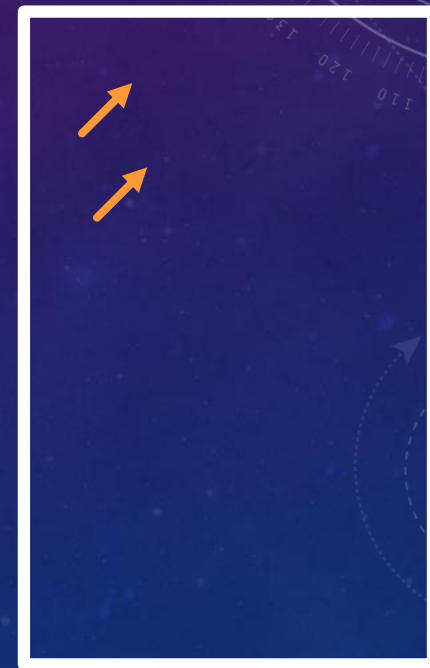
Minimal QES = $\gamma(t)$

$$S_{\text{CFT}} = \frac{\mathcal{A}(t)}{4G\hbar} + 0$$

S_{CFT} decreases
monotonically
after the Page
time



JT gravity with
Hawking State



Auxiliary System

A RECENT LOOK AT EVAPORATING BH'S: THE RESULTS

S_{CFT} starts at zero and increases monotonically until the Page time.

After the Page time, S_{CFT} decreases monotonically until it reaches zero.

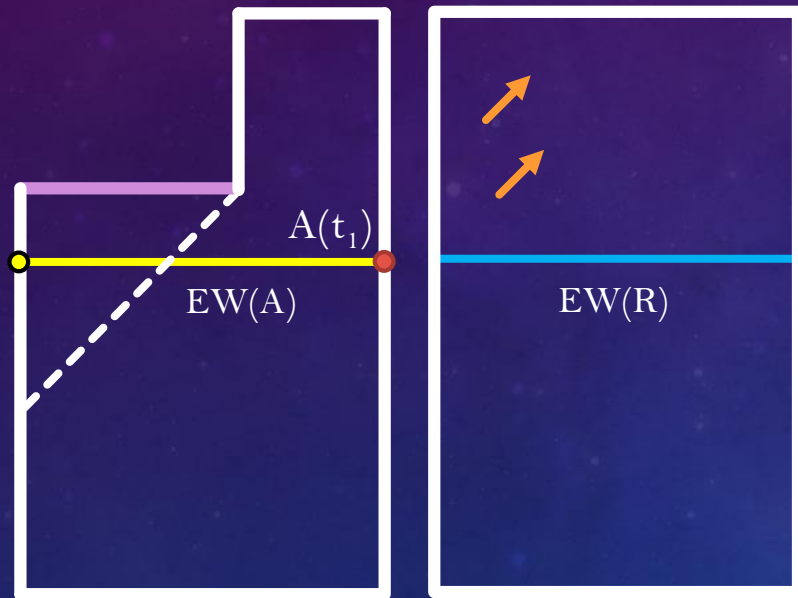
Therefore, S_{CFT} follows the Page Curve.

This is surprising because the calculation was semiclassical (Hawking's calculation) in the bulk.

WHAT ABOUT THE RADIATION?

Entanglement Wedge Complementarity

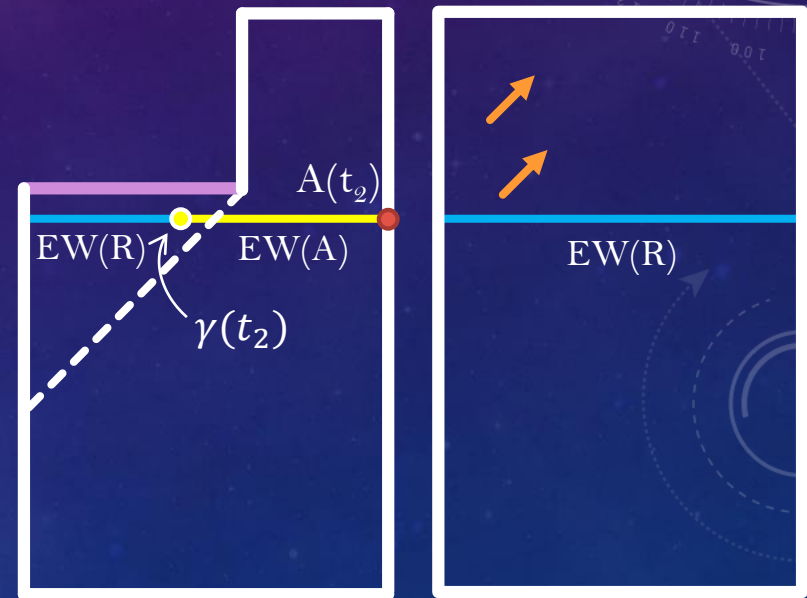
Before the Page Time



JT gravity with
Hawking State

Auxiliary System

After the Page Time



JT gravity with
Hawking State

Auxiliary System

An “island” appears in the entanglement wedge

SOME PUZZLES AND FURTHER RELEVANT WORK

What does the observer detect? Information loss or not?

Contradiction between mixed out-state and purity of the CFT?

Significance of ensemble of CFTs on the boundary?

arXiv:1911.06305.

Computing the Page curve for the radiation with double holography:

arXiv:1908.10996.

Applying the gravitational path integral (via replica wormholes) to
this setup: *arXiv:1911.12333, arXiv:1911.11977.*

Double-holography in higher dimensions: *arXiv:1911.09666.*

REFERENCES

- A. Almheiri et al. “The entropy of bulk quantum fields and the entanglement wedge of an evaporating black hole,” (2019). arXiv:1905.08762.
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- R. Bousso and M. Tomasevic. “Unitarity From a Smooth Horizon?” (2019). arXiv:1911.06305.
- D. Harlow. “Jerusalem Lectures on Black Holes and Quantum Information,” *Rev. Mod. Phys.* **88** (2016) 015002. arXiv:1405.1995.
- G. Penington. “Entanglement Wedge Reconstruction and the Information Paradox,” (2019). arXiv:1905.08255.
- G. Penington et al. “Replica wormholes and the black hole interior,” (2019). arXiv:1911.11977.
- T. Takayanagi. “Holographic dual of BCFT,” *Phys. Rev. Lett.* **107** (2011) 101602. arXiv:1105.5165.



THANK YOU FOR YOUR
ATTENTION!

EXTRA SLIDES

ADS GEOMETRY

Metric of AdS_{d+1}

$$ds^2 = \ell^2(-\cosh^2 \rho dt^2 + d\rho^2 + \sinh^2 \rho d\Omega_{d-1}^2)$$

Construction of Penrose diagram

$$d\sigma = \frac{d\rho}{\cosh \rho} \rightarrow \sigma = 2 \tan^{-1} \tanh\left(\frac{\rho}{2}\right)$$

ρ runs from 0 to ∞ , σ runs from 0 to $\frac{\pi}{2}$.

JT GRAVITY AND THE SYK MODEL

JT Gravity

Action of JT gravity: $I = I_0 + I_G + I_M$, where

$$I_0 = \frac{\phi_0}{16\pi G_N} \left[\int_{\mathcal{M}} d^2x \sqrt{-g} R + 2 \int_{\partial\mathcal{M}} K \right]$$

$$I_G = \frac{1}{16\pi G_N} \left[\int_{\mathcal{M}} d^2x \sqrt{-g} \phi (R + 2) + 2 \int_{\partial\mathcal{M}} \phi_b K \right]$$

$$I_M = I_{CFT}[g]$$

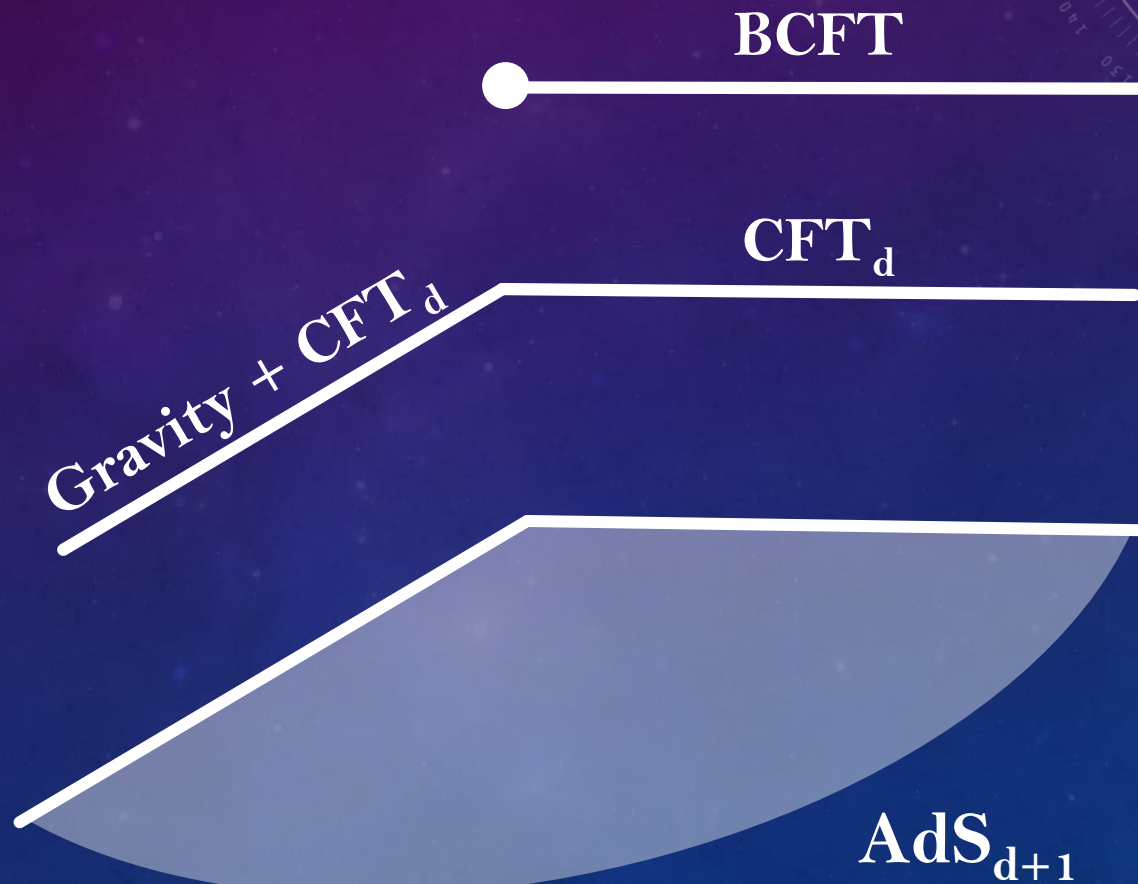
Solving the action yields a 1+1D metric that is locally AdS_2 .

SYK Model

SKY model: 0+1D QM model of N Majorana fermions (ψ_i).

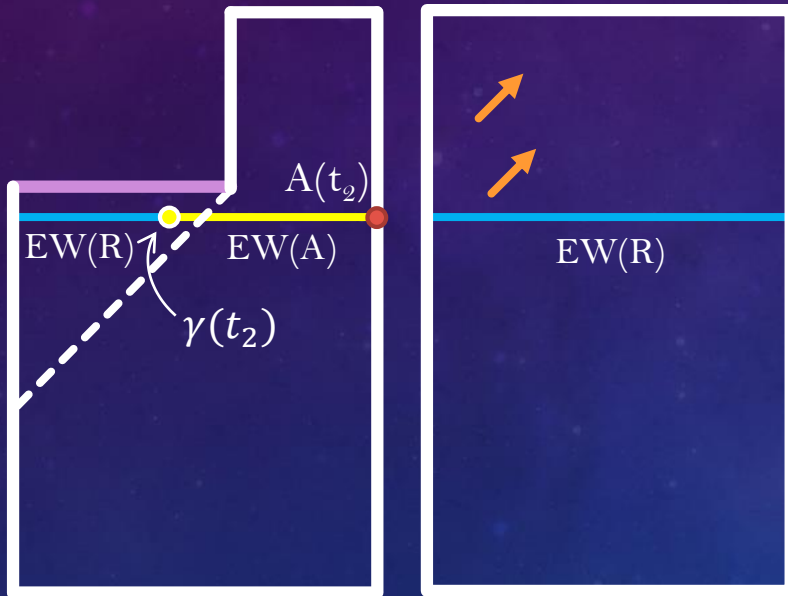
Hamiltonian: $H = \sum_{iklm} j_{iklm} \psi_i \psi_k \psi_l \psi_m$, where j_{iklm} are independent Gaussian random couplings.

DOUBLE-LAYER HOLOGRAPHY: THE BATH AS A CFT



DOUBLE-LAYER HOLOGRAPHY: THE BATH AS A CFT

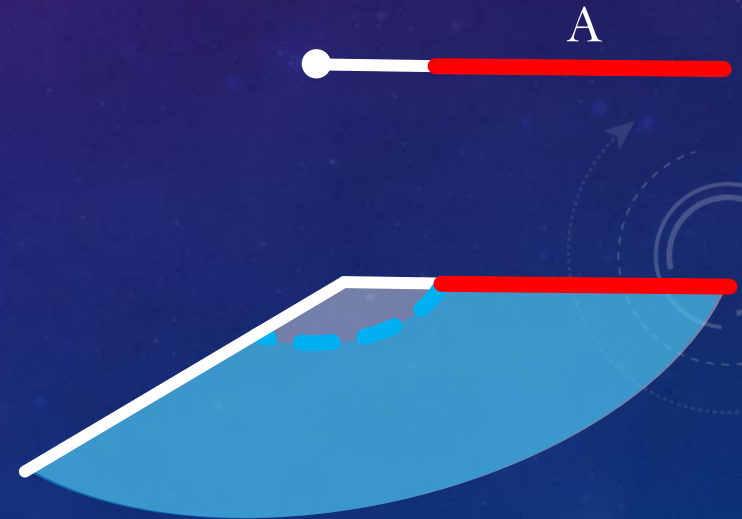
Entanglement Wedge Complementarity



JT gravity with
Hawking State

Auxiliary System

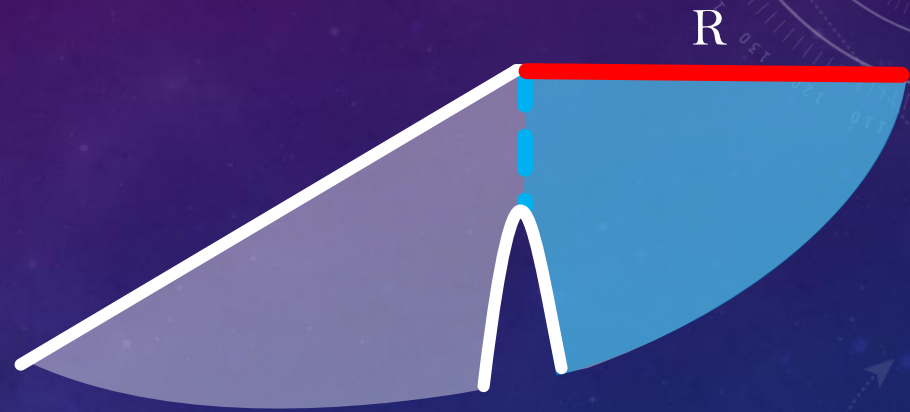
RT for BCFTs



A

DOUBLE-LAYER HOLOGRAPHY: THE PAGE CURVE OF THE RADIATION

Before
the Page Time



After
the Page Time

DOUBLE-LAYER HOLOGRAPHY: THE PAGE CURVE OF THE RADIATION

The entropy of the radiation follows the Page curve.

A New Prescription:

$$S(A) = \text{Min}_{\mathcal{J}_g} \text{Ext}_{\mathcal{J}_g} \left[S_{\text{eff}}(A \cup \mathcal{J}_g) + \frac{\text{Area}[\partial \mathcal{J}_g]}{4G_N} \right]$$

RINDLER COORDINATES

Rindler Space metric:

$$ds^2 = dR^2 - R^2 d\eta^2, R > 0, \eta = \text{anything}$$

Coordinate transformation makes the Rindler patch

more evident:

$$x = R \cosh \eta$$

$$t = R \sinh \eta$$

$$ds^2 = -dt^2 + dx^2, x > 0, |t| < x$$