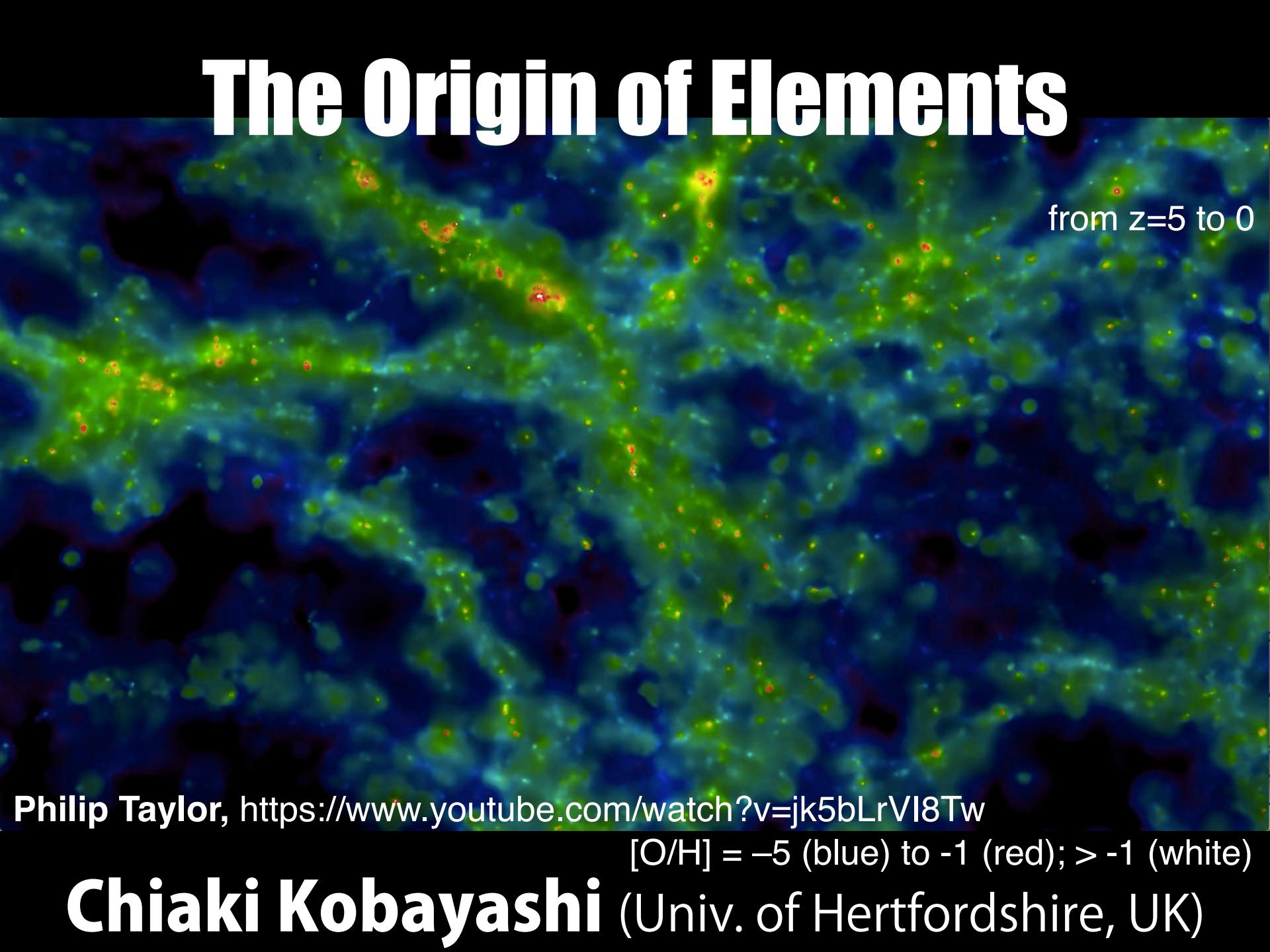


# The Origin of Elements



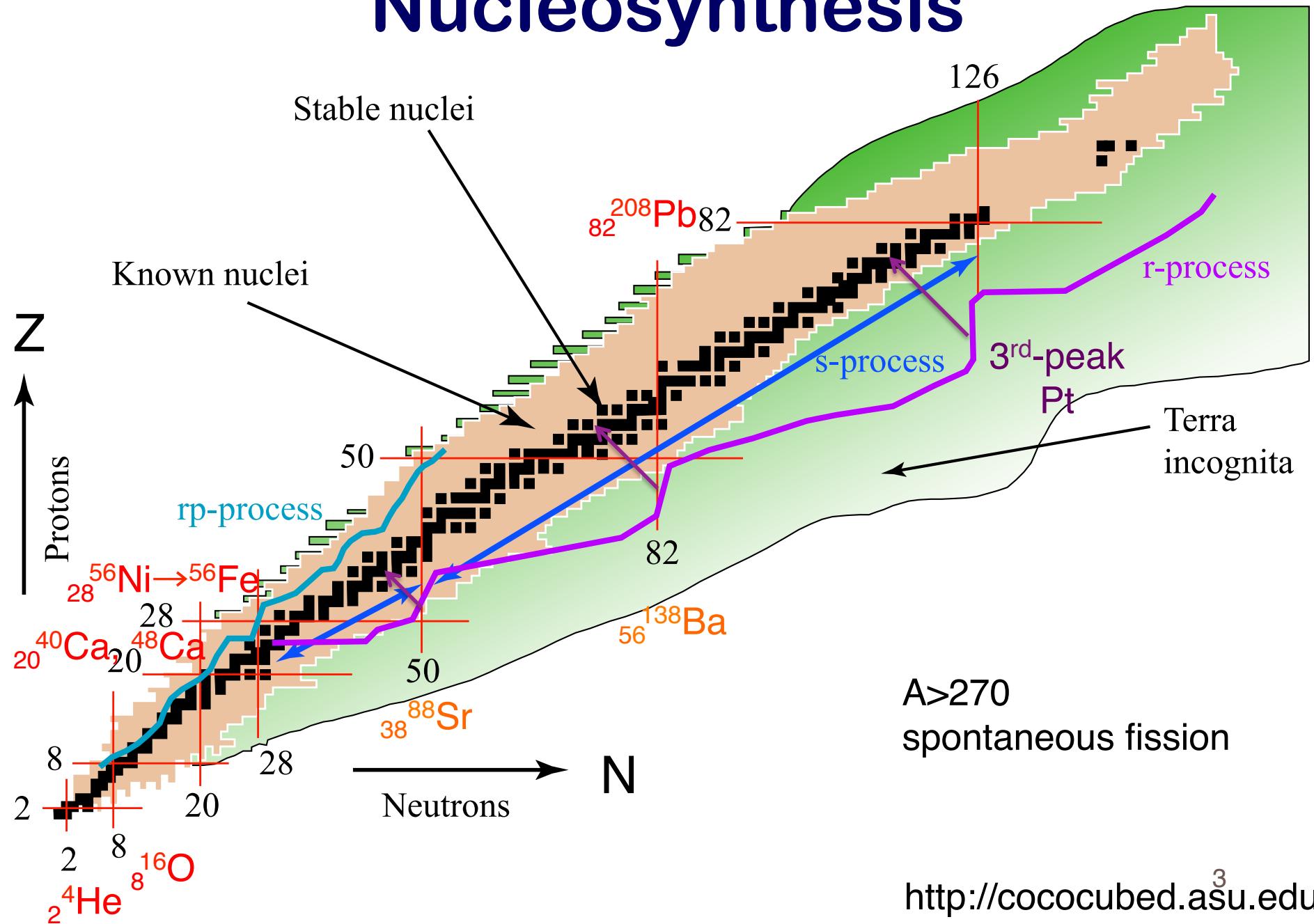
from z=5 to 0

Philip Taylor, <https://www.youtube.com/watch?v=jk5bLrVI8Tw>

[O/H] = -5 (blue) to -1 (red); > -1 (white)

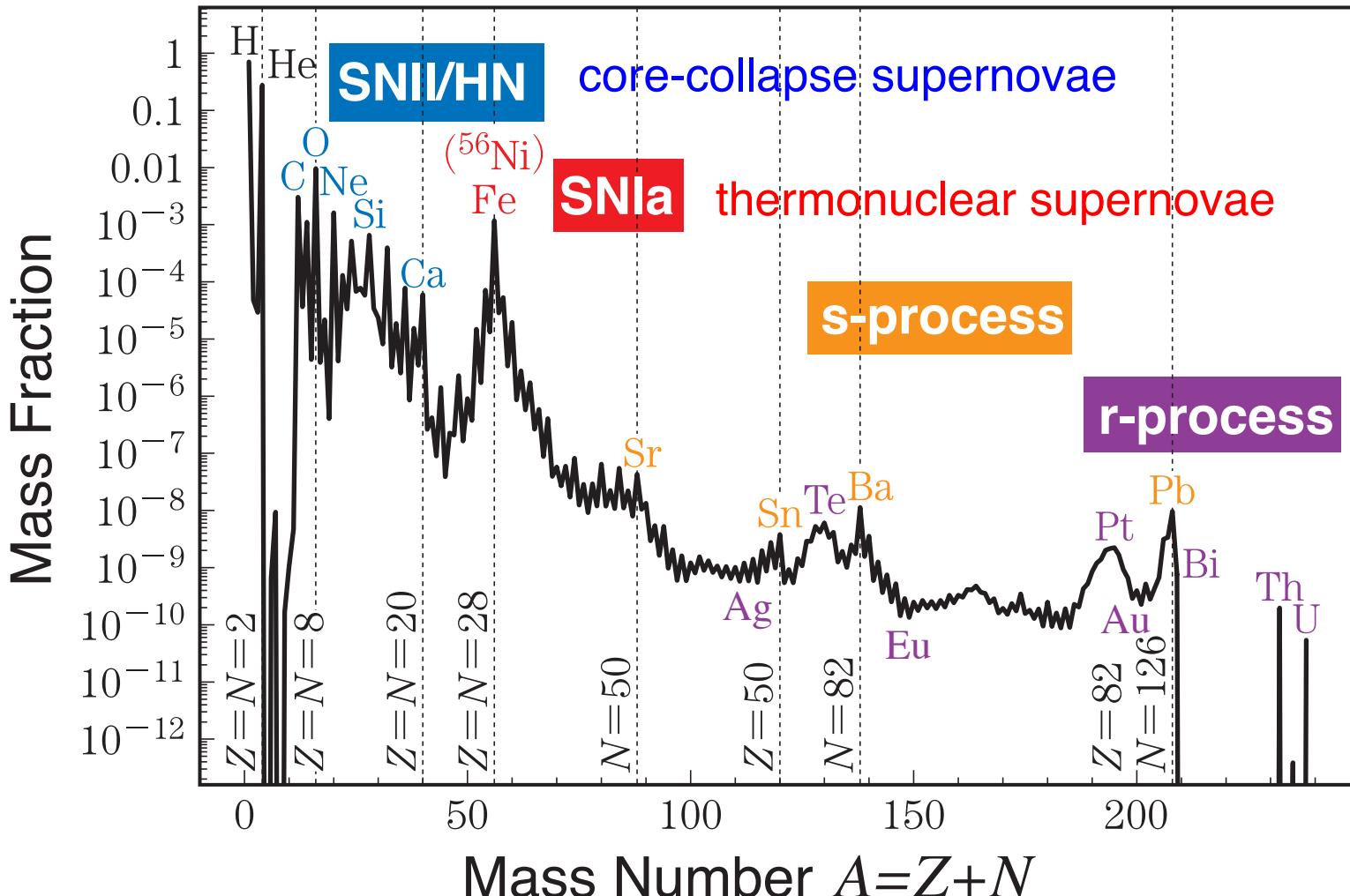
**Chiaki Kobayashi** (Univ. of Hertfordshire, UK)

# Nucleosynthesis



# Origin of Elements in the Sun

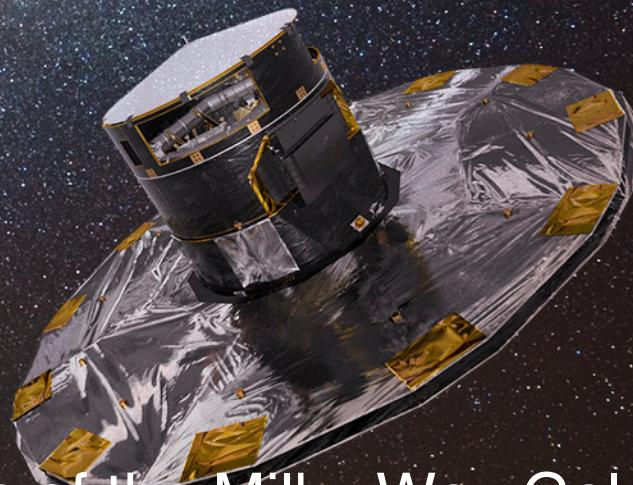
Big Bang Nucleosynthesis:  $^1\text{H}$ ,  $^2\text{H}$ ,  $^4\text{He}$ ,  $^7\text{Li}$



# Galactic Archaeology

## for Milky Way and local dwarf galaxies

- ★ Motions of one billion stars are measured with GAIA.
- ★ Ages from asteroseismology COROT, Kepler, K2, TESS...
- ★ Elemental Abundances (from Li to Eu) of one million stars will be measured with multi-object spectrographs:
  - ★ **SEGUE** (Resolution~1800) on SDSS
  - ★ **RAVE** ( $R\sim 7500$ ) on 1.2m UKST
  - ★ **HERMES** on AAT ( $R\sim 28000/50000$ )
  - ★ **APOGEE** ( $R\sim 20000$ , IR) on SDSS
  - ★ **GAIA-ESO with VLT** ( $R\sim 20000/40000$ )
  - ★ ~~WFMOS on Subaru~~
  - ★ **WEAVE** on WHT ( $R\sim 5000/20000$ )
  - ★ **4MOST** on VISTA ( $R\sim 5000/18000$ )
  - ★ **PFS** on Subaru ( $R\sim 2300-5000$ )
  - ★ **MSE** ( $R\sim 2000/6500/20000$ )
- ★ Chemical and dynamical evolution of the Milky Way Galaxy are being revealed!



# Core-collapse Supernovae

SN1987A in LMC on 2/23/1987; Betelgeuse xx/xx/2020-3020



BSG

© Anglo-Australian Observatory

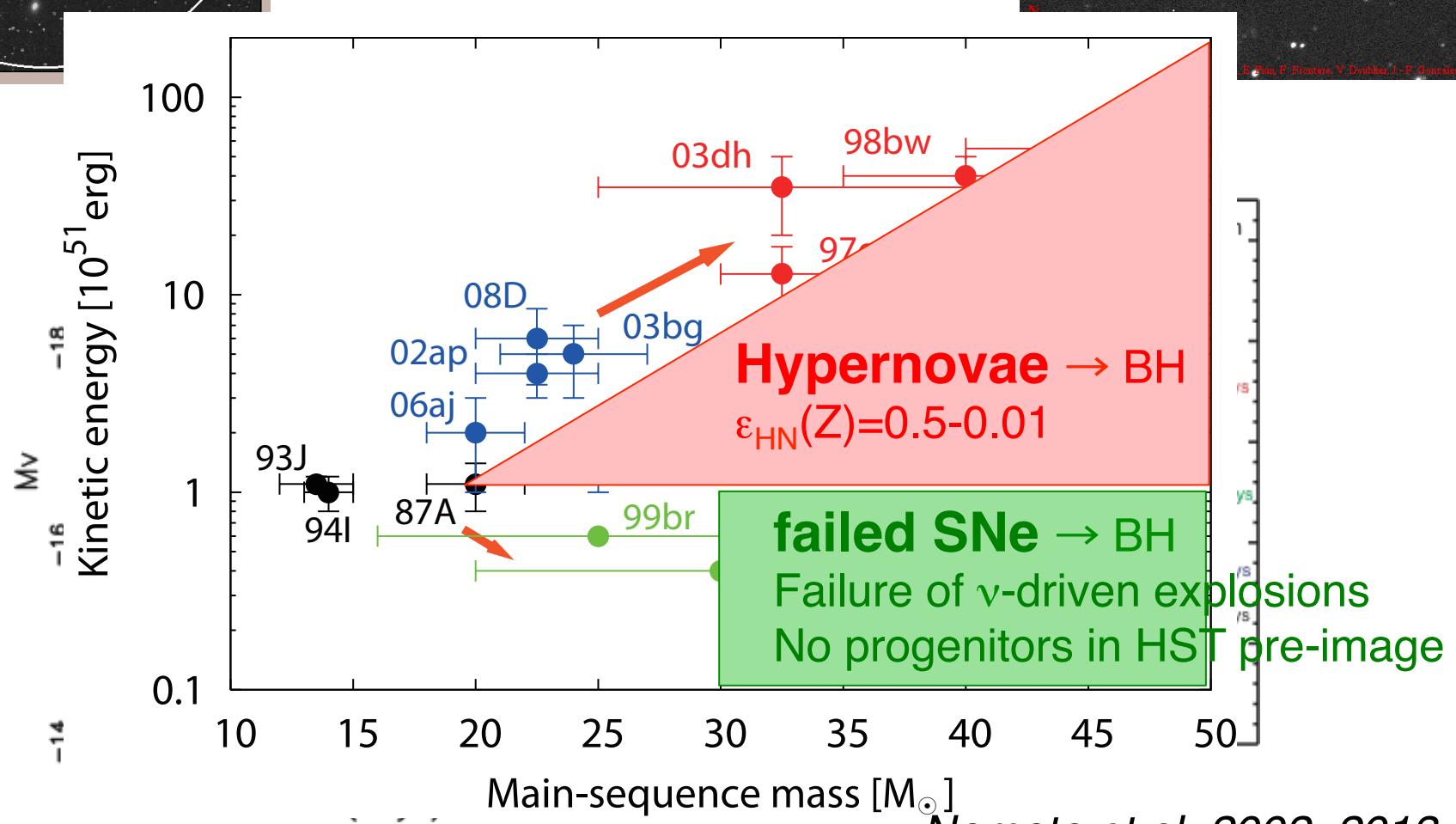
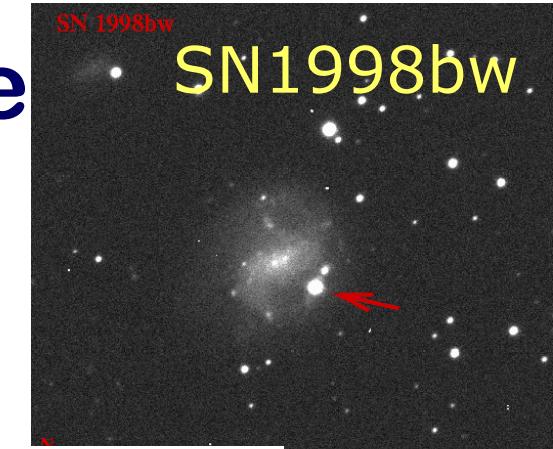
GRB980425

# Core-collapse SNe

SN 1998bw

SN1998bw

- \* SN light curves & spectra fitting →  $M$ ,  $E_{\text{kin}}$ ,  $M(\text{Fe})$
- \* Mixing-fallback model



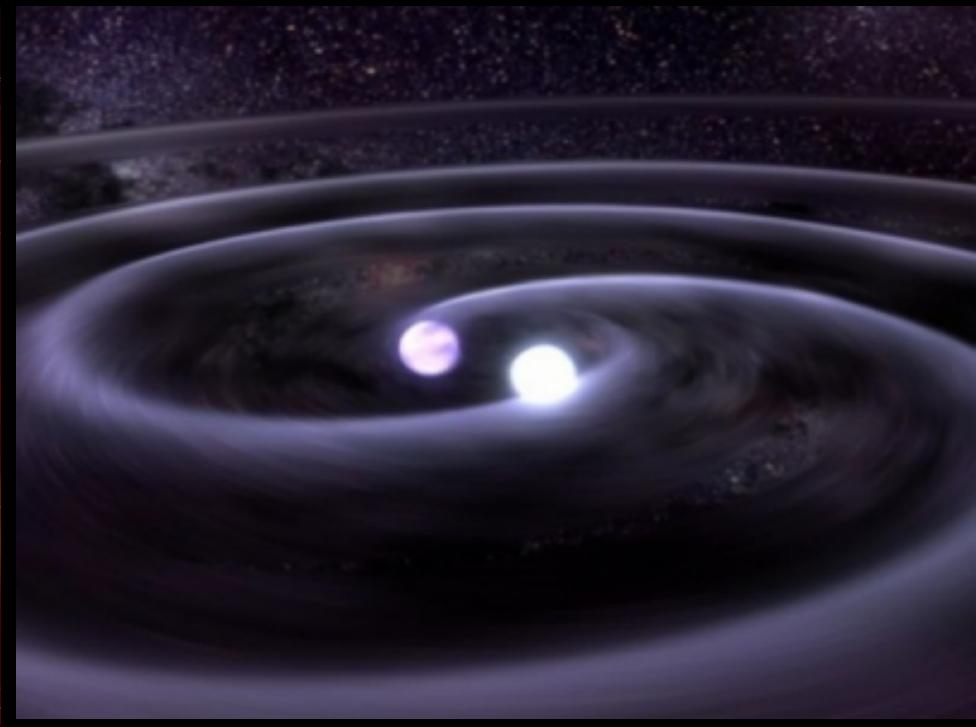
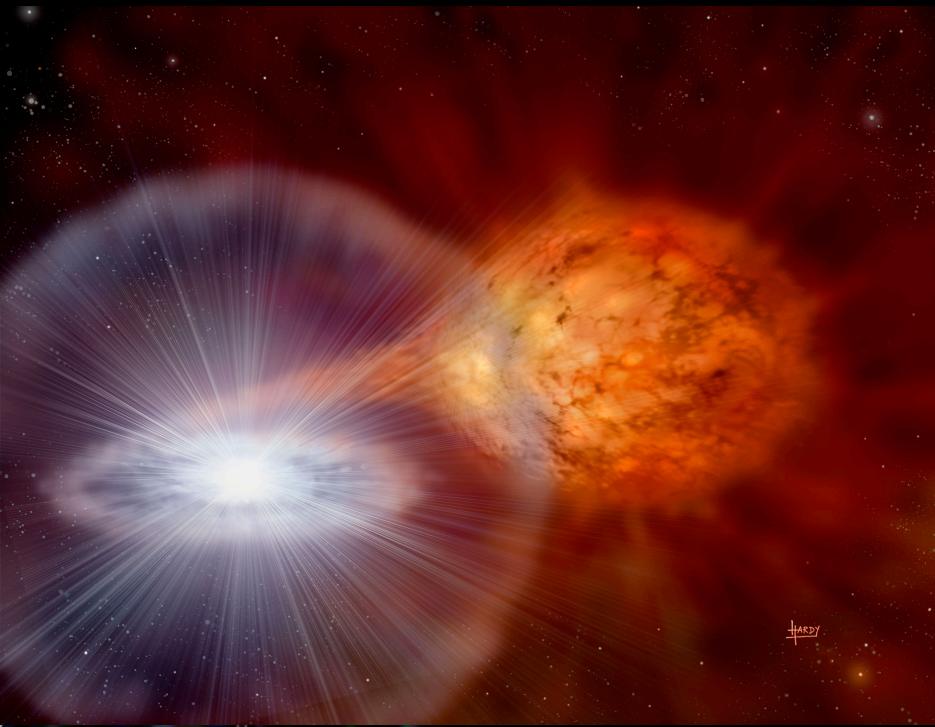
# Thermonuclear (Type Ia) Supernovae

Thermonuclear explosion in a binary with C+O white dwarf

Ch-mass explosion

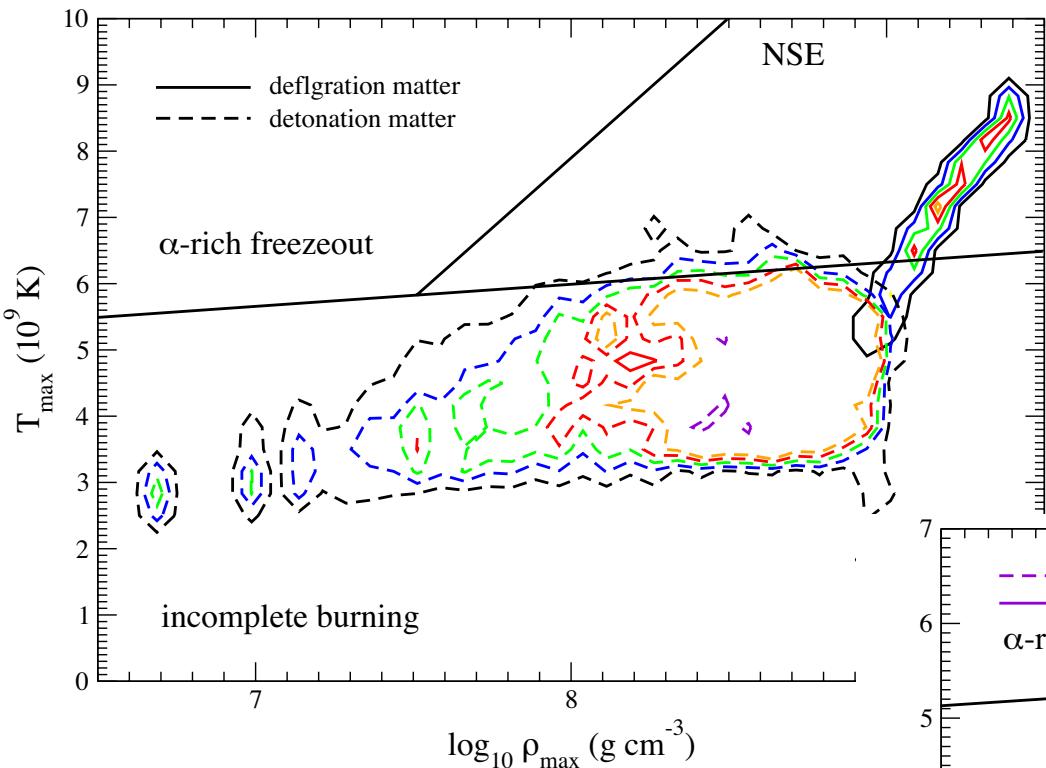
(deflagration or delayed  
detonation) possibly in SD

vs      sub-Ch mass explosion  
(double detonation) in DD and SD



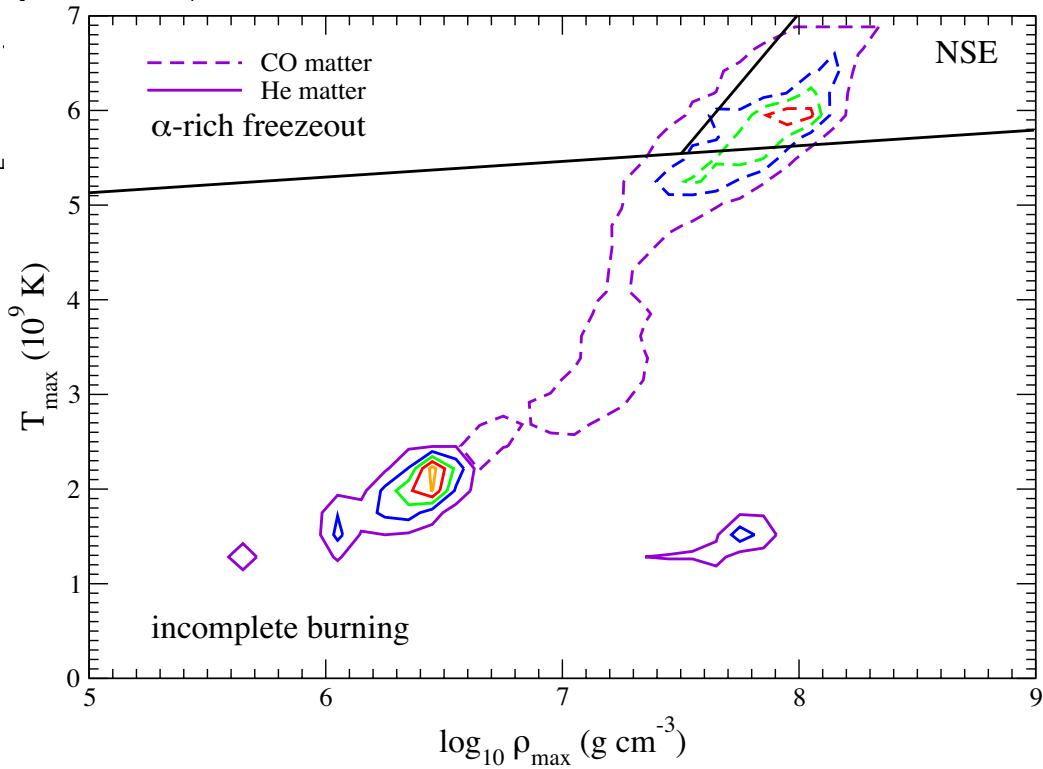
McCully+14

# 2D Nucleosynthesis: Ch vs sub-Ch



CK, Leung, Nomoto 19

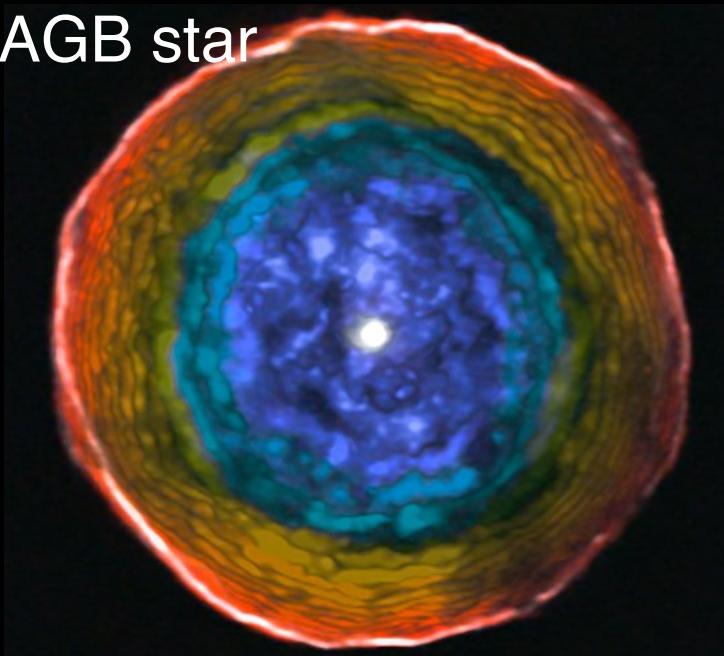
- \* **Double detonation of sub-Ch WD (below)**
- \*  $Y_e > 0.499$
- \*  $[(\text{Si}, \text{S}, \text{Ar}, \text{Ca})/\text{Fe}]$  close to the solar ratios



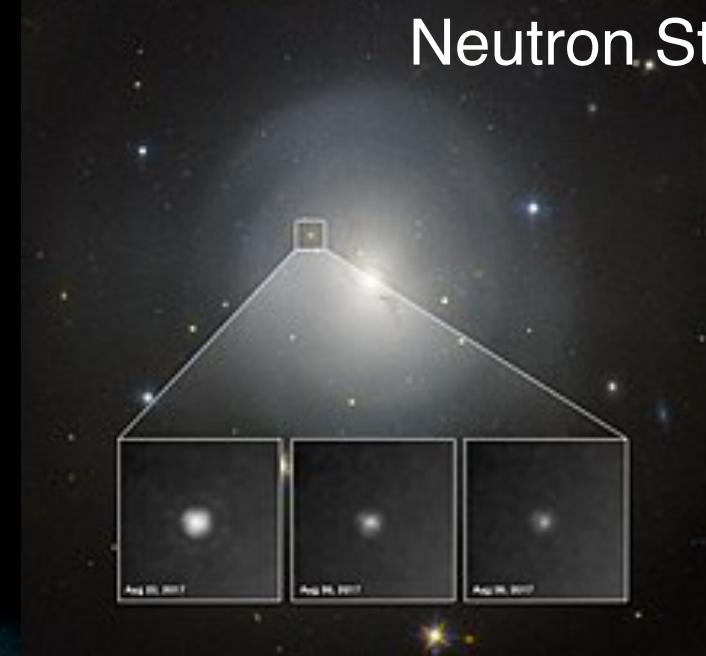
- \* **Delayed detonation of Ch WD (above)**
- \*  $Y_e$  can be as low as 0.46 for deflagration
- \*  $^{54}\text{Fe}$ ,  $\text{Mn}$ ,  $^{58}\text{Ni}$  enhanced

# Neutron-capture processes

AGB star



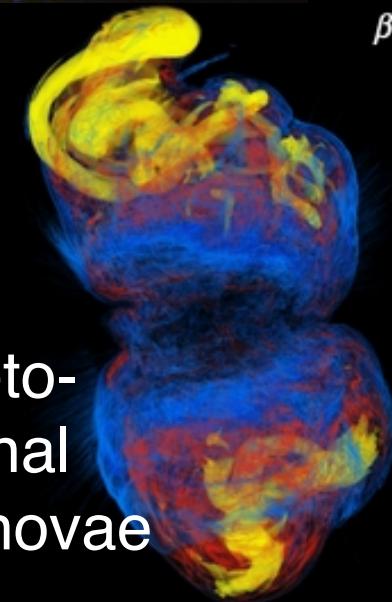
Neutron Star Merger



Electron Capture  
Supernovae



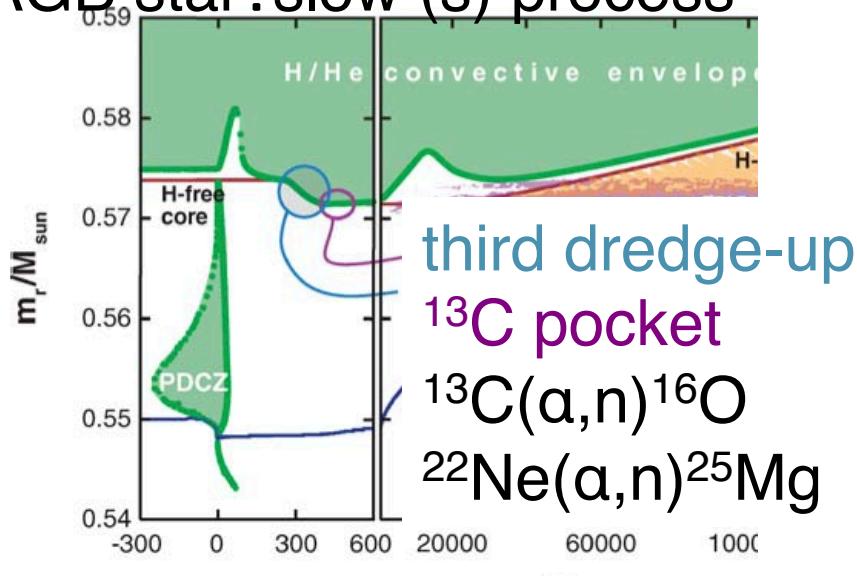
Magneto-  
rotational  
Supernovae



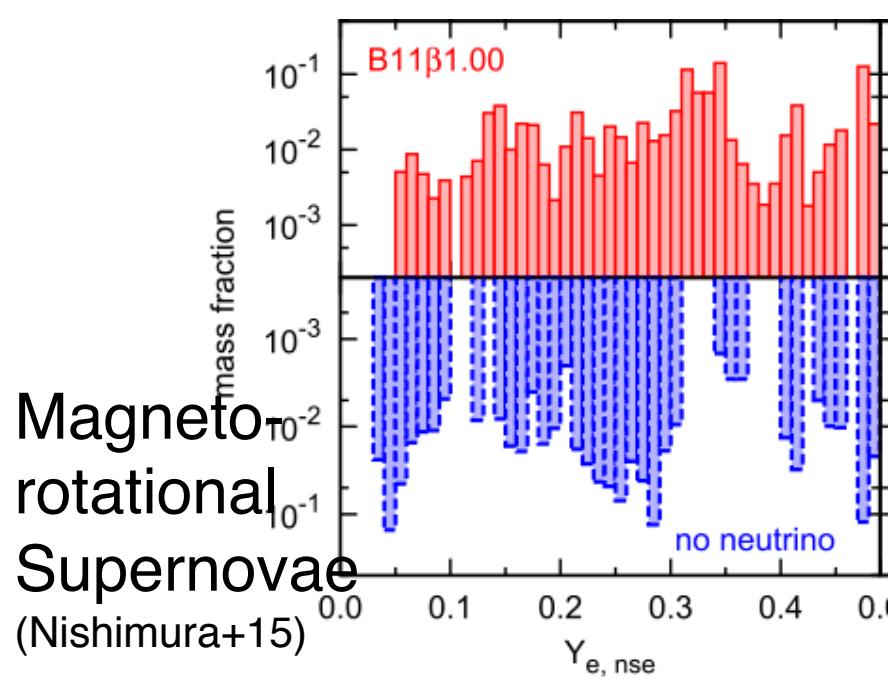
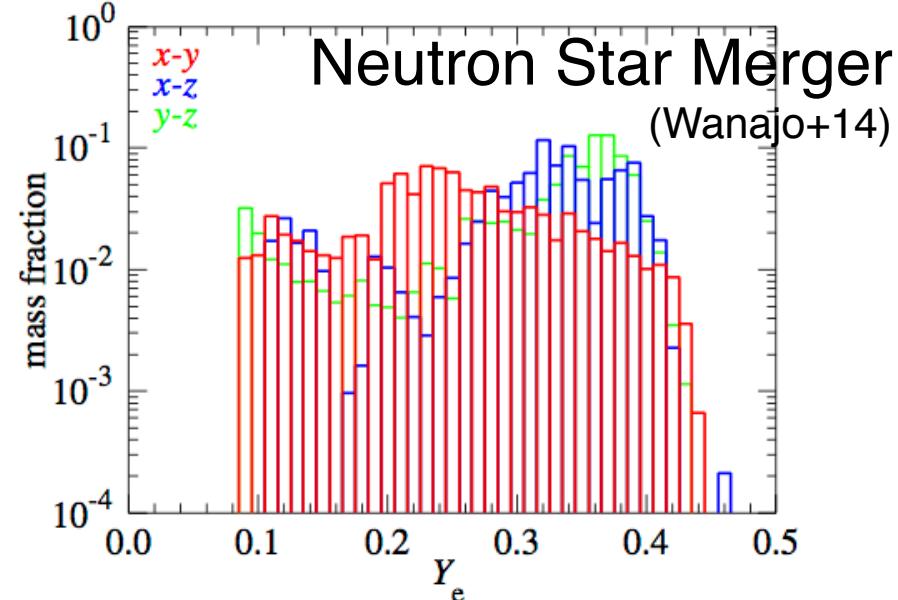
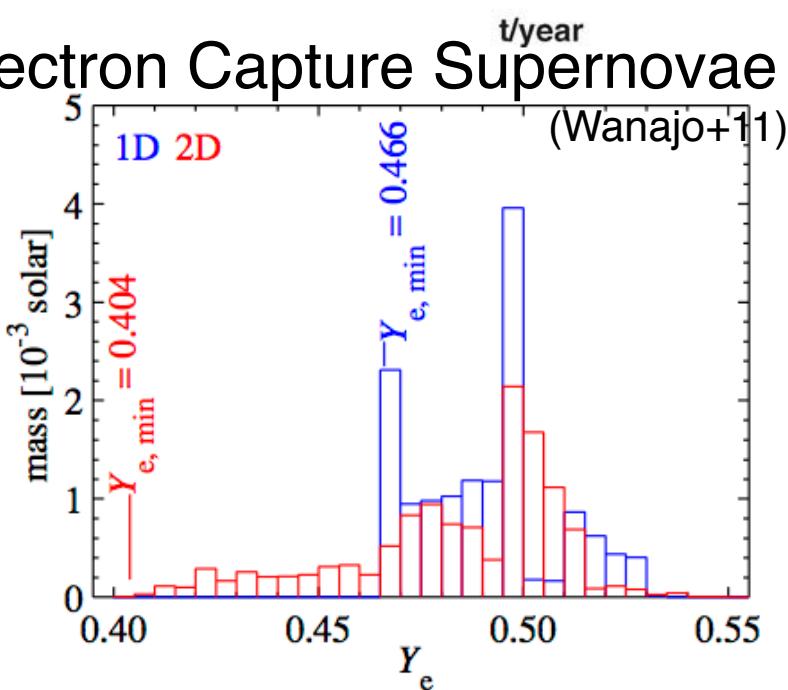
$$\beta = \frac{P_{\text{gas}}}{P_{\text{mag}}}$$

# Neutron-capture processes

AGB star: slow (s) process

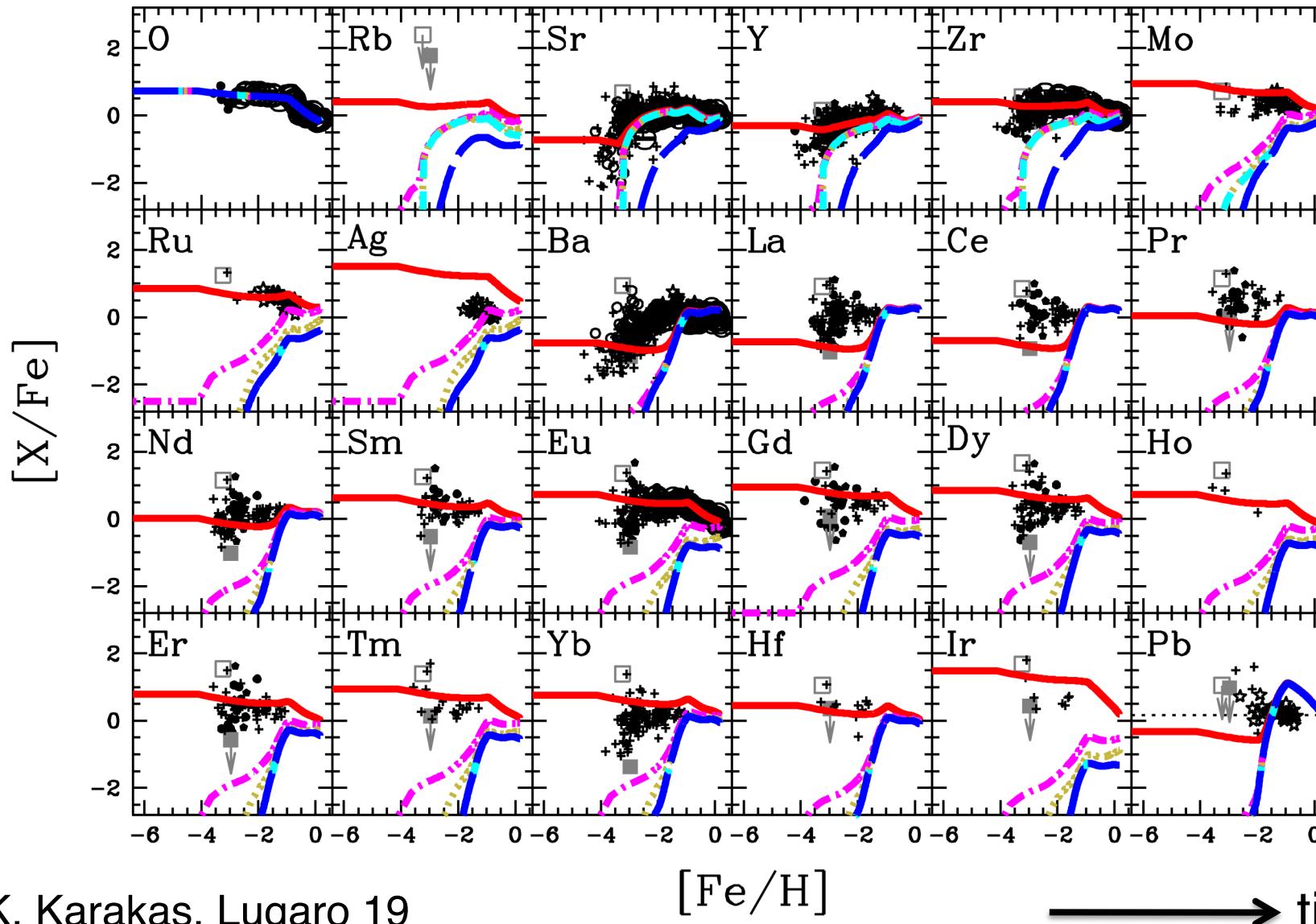


Electron Capture Supernovae



# Neutron-capture elements

s-process, ECSN,  $\nu$ -winds, NS+NS&NS+BHM, MRSN



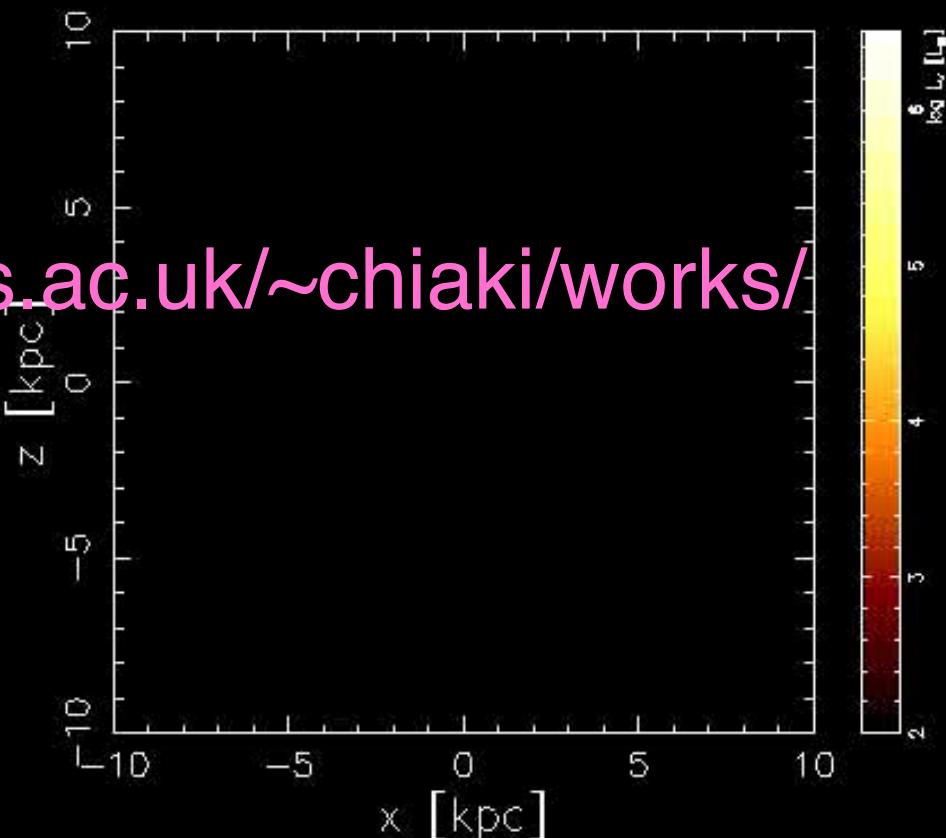
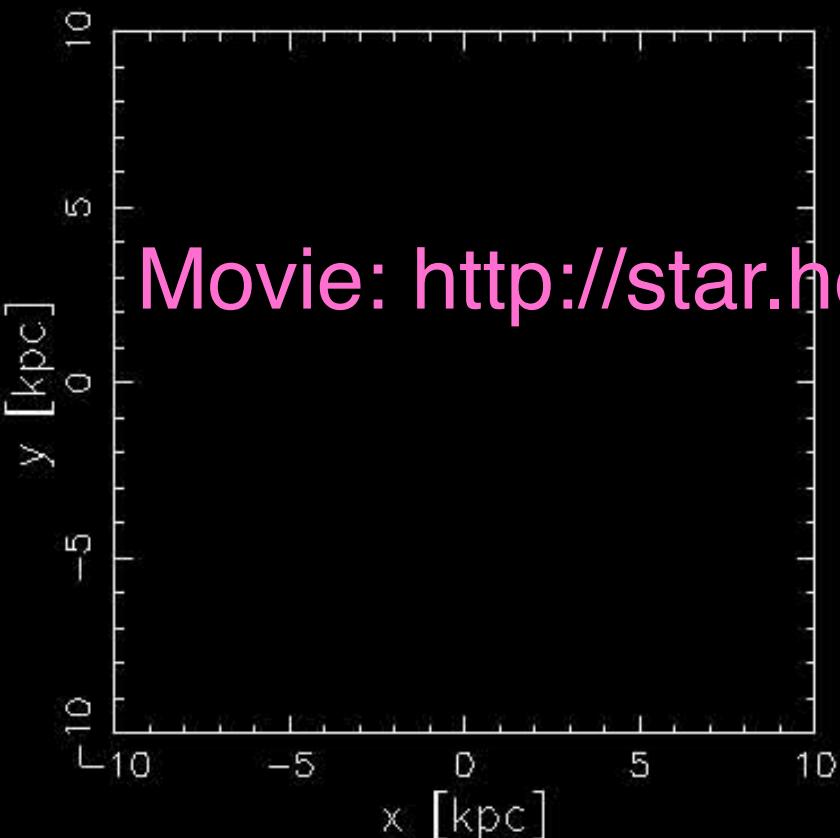
# Milky Way-type simulations

Initial Condition:  $\lambda$ CDM fluctuated sphere with  $\lambda \sim 0.1$ ,  $r \sim 3$ Mpc,  
 $M_{\text{tot}} \sim 10^{12} M_{\odot}$ ,  $N_{\text{tot}} \sim 120.000$ ,  $M_{\text{gas}} \sim 10^6 M_{\odot}$ ,  $M_{\text{DM}} \sim 10^7 M_{\odot}$   
(CK & Nakasato 2011, ApJ, 729, 16)

Face on

$t = 0.00$  Gyr,  $z = 23.69$

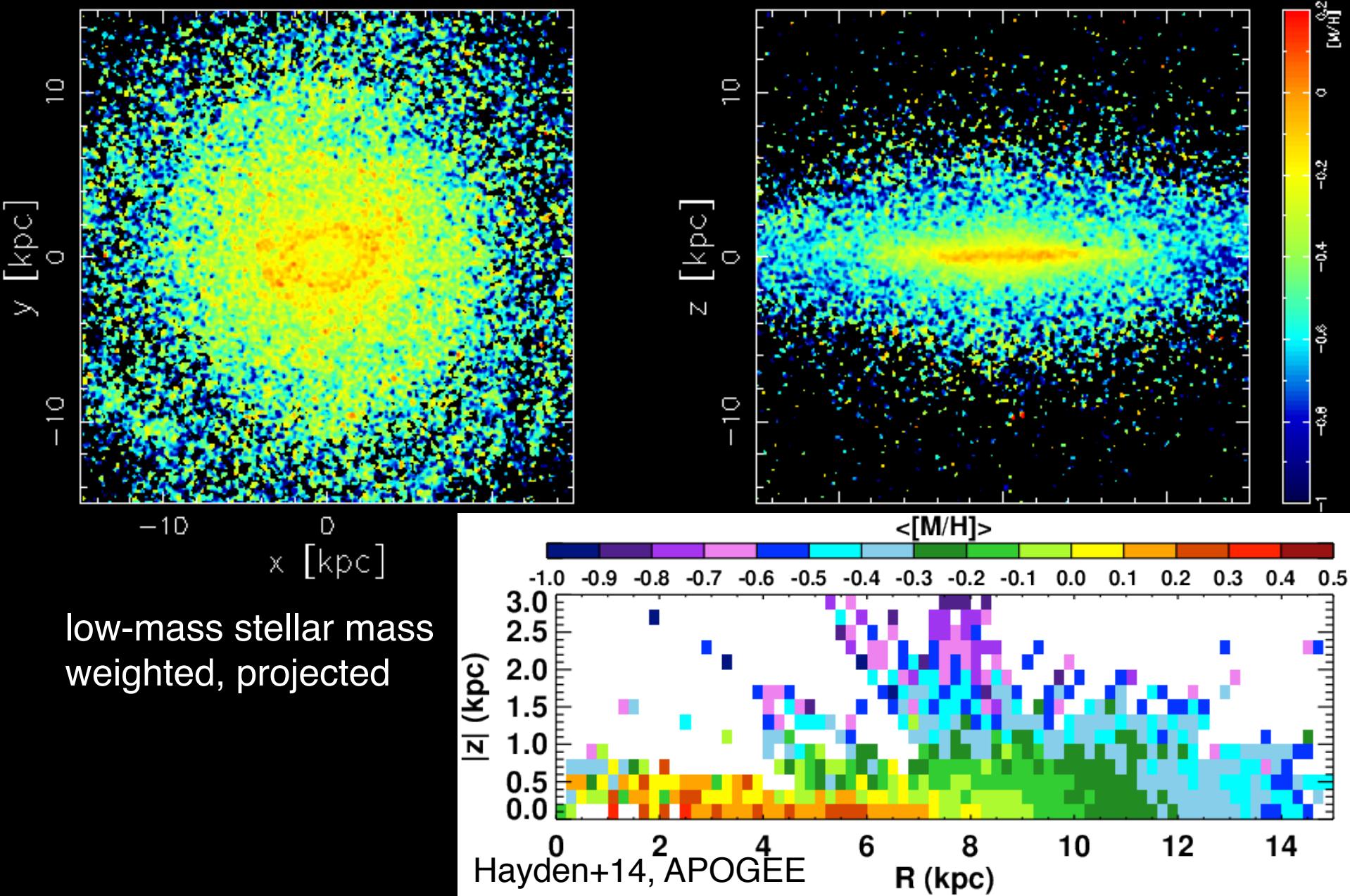
Edge on



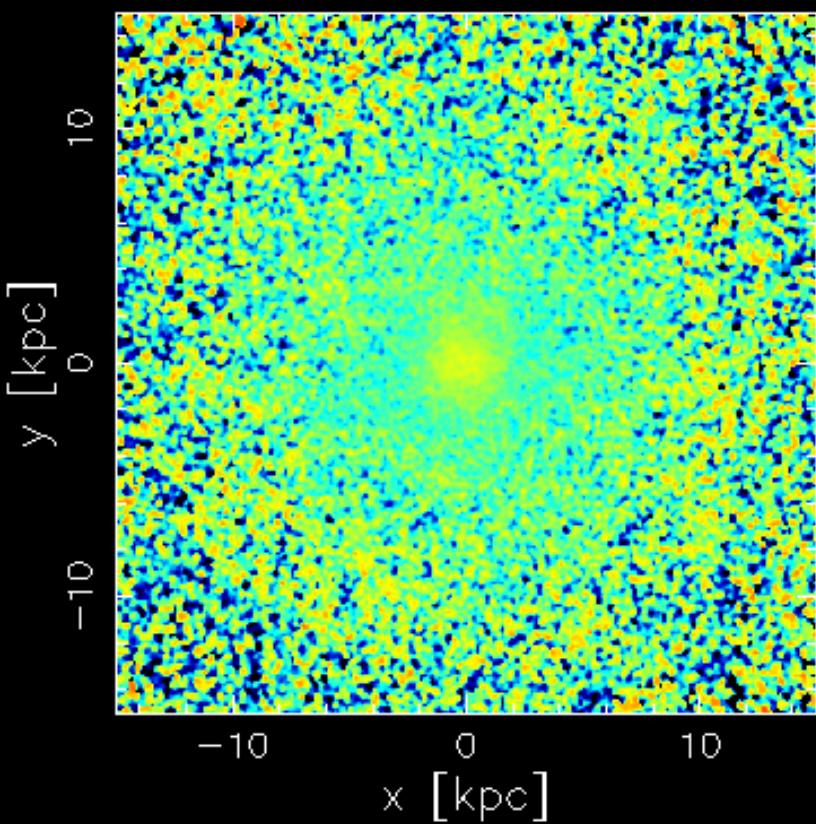
Movie: <http://star.herts.ac.uk/~chiaki/works/>

Similar results obtained also with Aquarius Initial Condition (CK 2015).

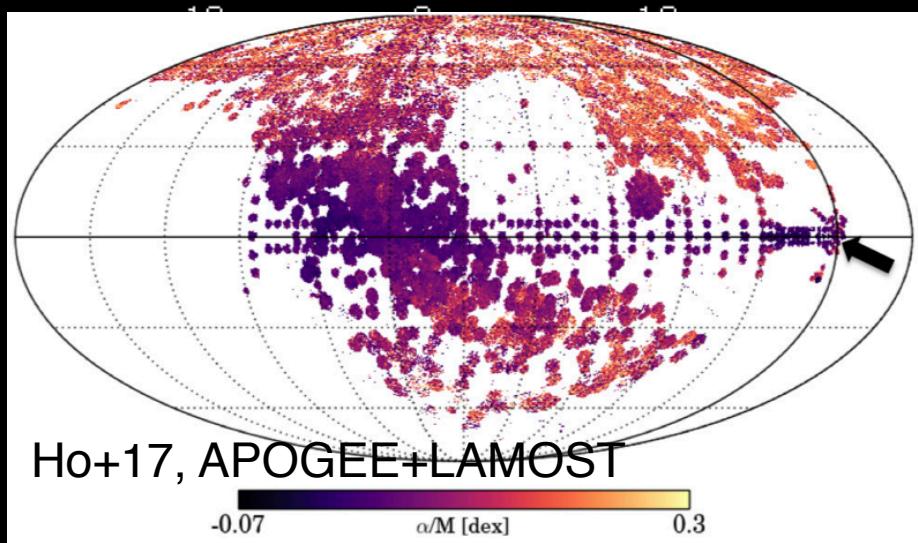
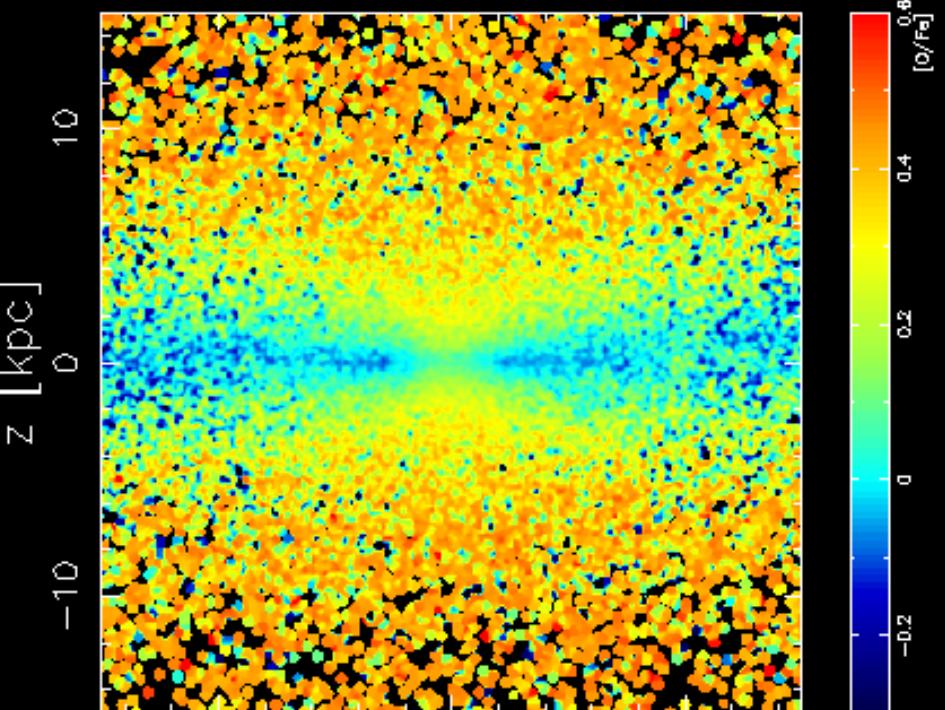
# Metallicity Map



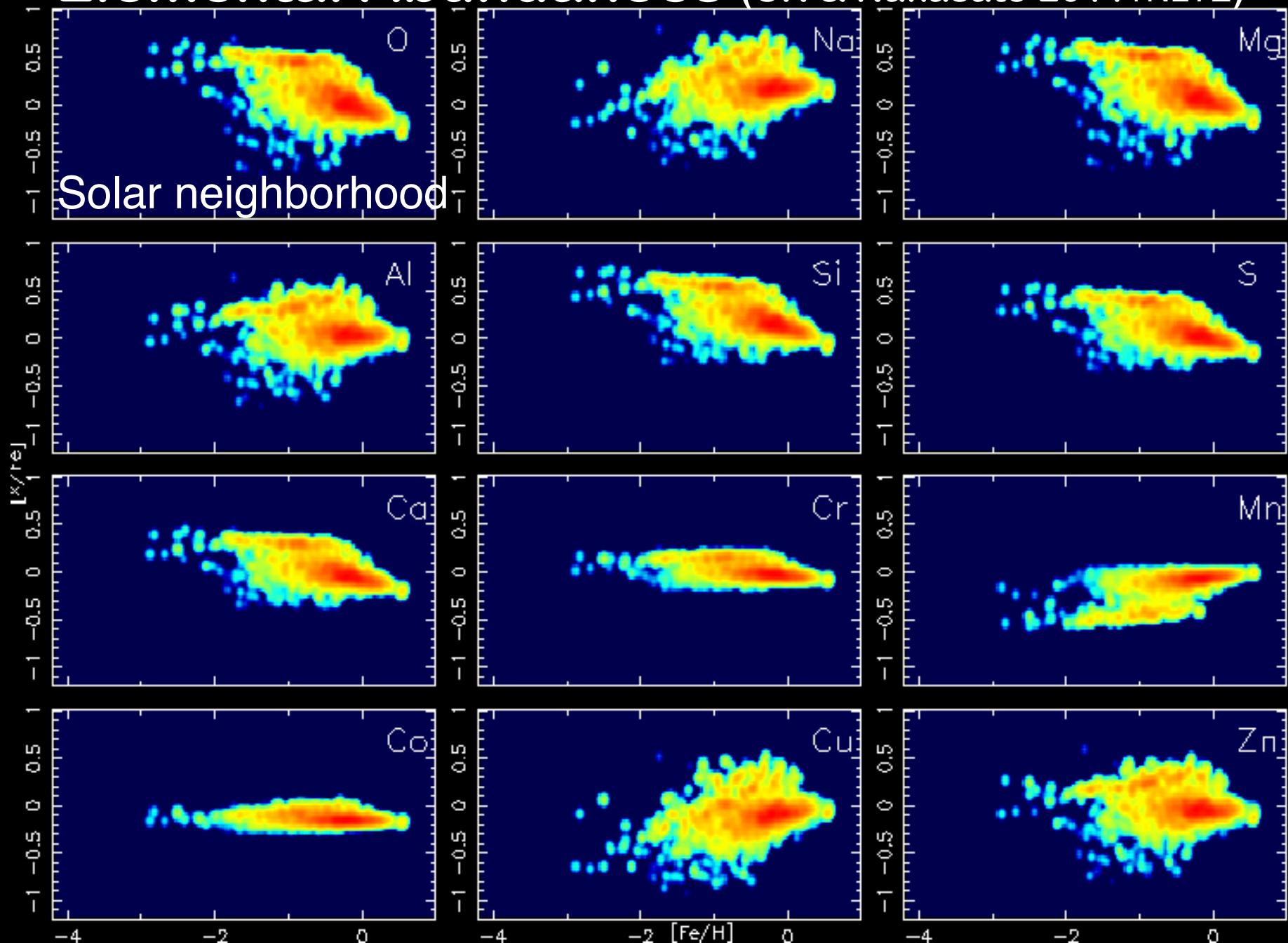
# [O/Fe] Map



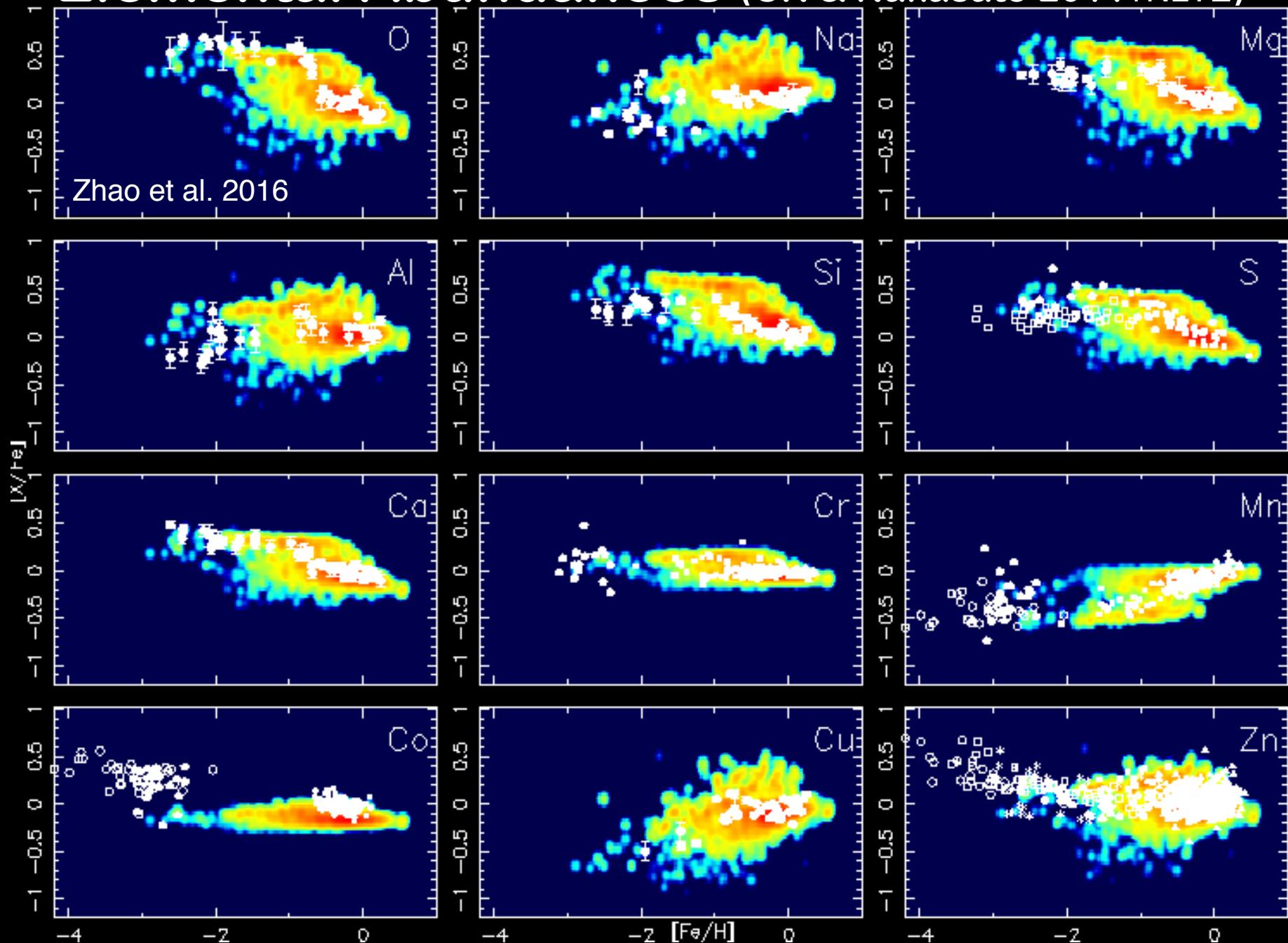
low-mass stellar mass  
weighted, projected



# Elemental Abundances (CK & Nakasato 2011+NLTE)

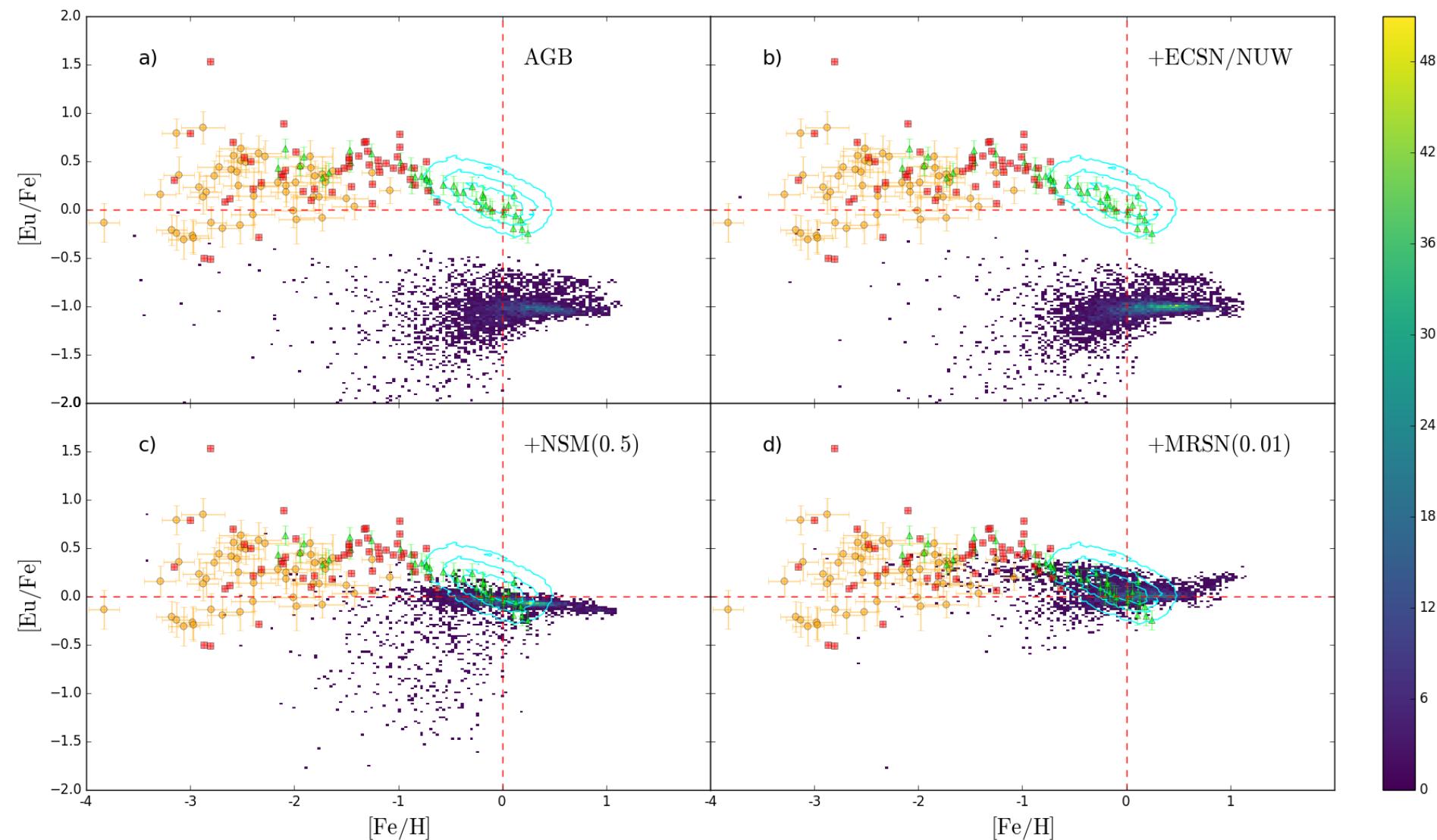


# Elemental Abundances (CK & Nakasato 2011+NLTE)



# [Eu/Fe]-[Fe/H]

Chemo-hydrodynamical Simulation  
Chris Haynes & CK 2019

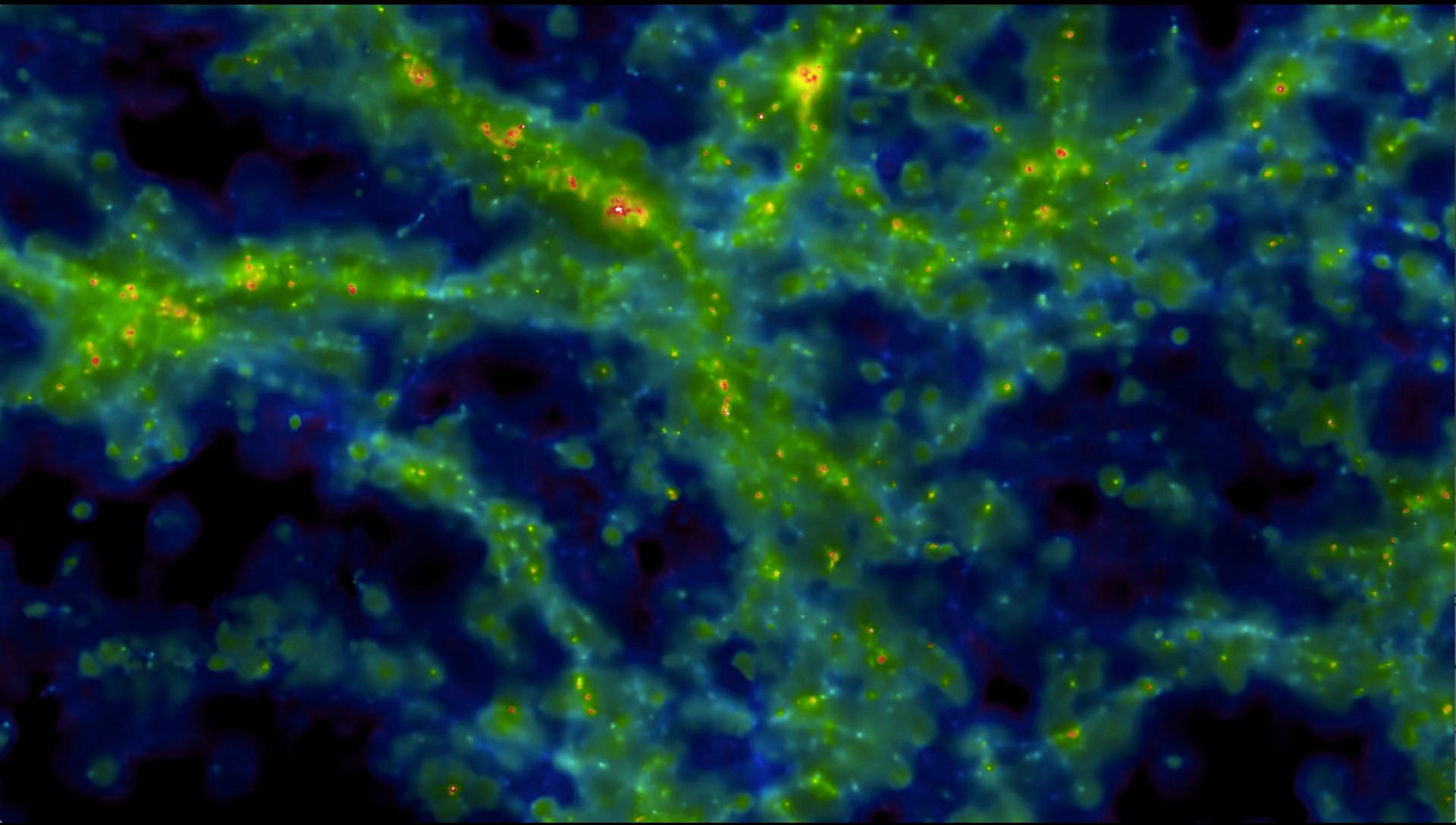


Neutron star mergers alone cannot reproduce the observations.

Hansen+17; Roederer+16; NLTE Zhao+16; HERMES-GALAH

# Cosmological Simulations

from z=5 to 0

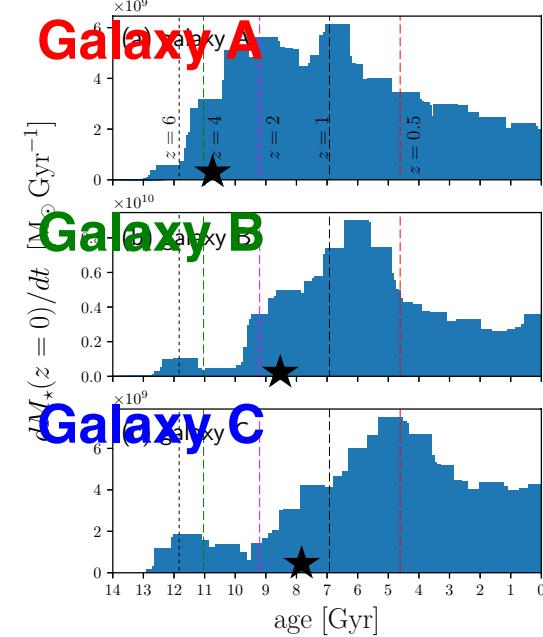
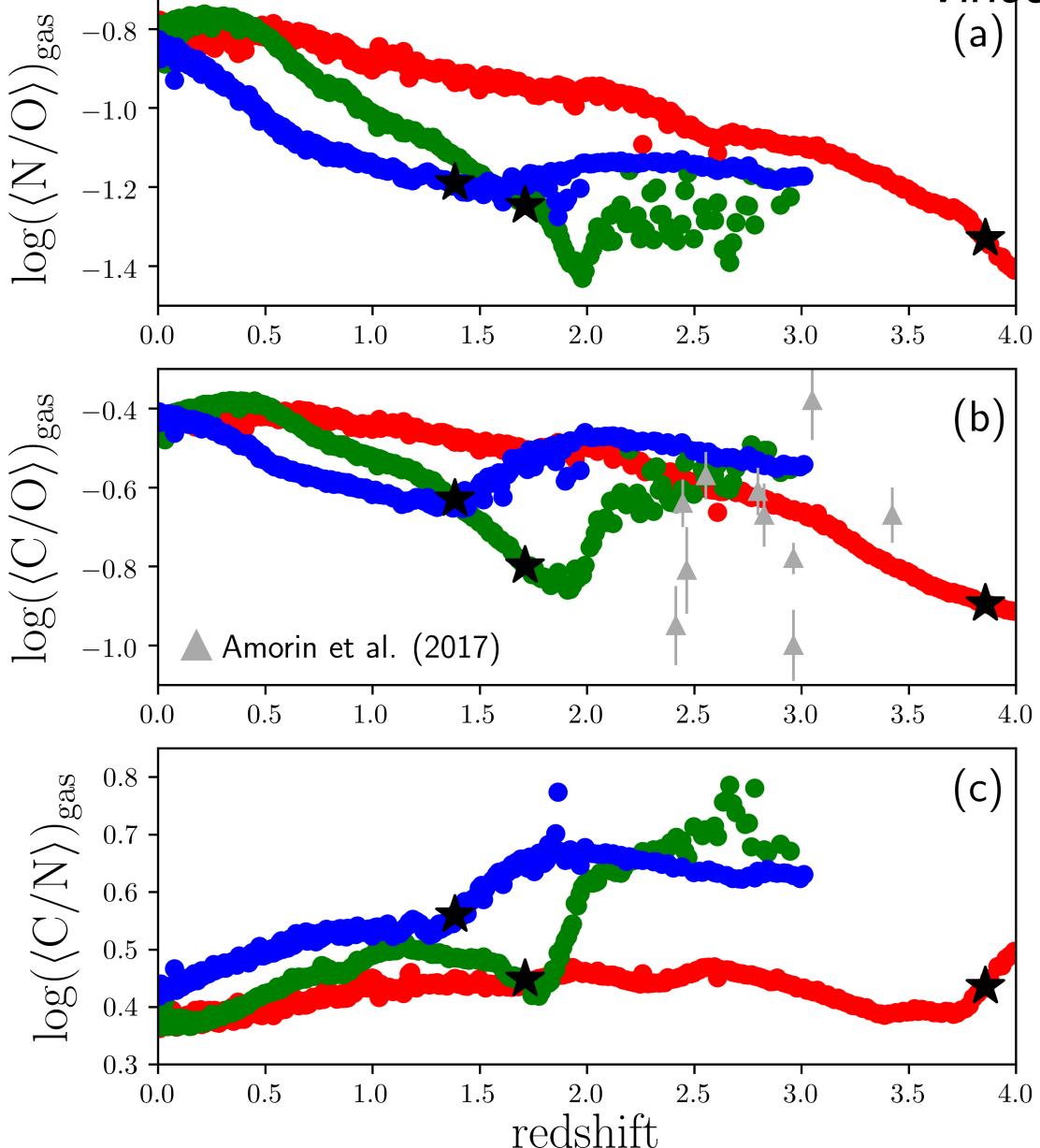


$[\text{O}/\text{H}] = -5$  (blue) to  $-1$  (red);  $> -1$  (white)

Philip Taylor, <https://www.youtube.com/watch?v=jk5bLrVI8Tw>

# Redshift evolution of CNO ratios

Vincenzo & CK 2018a, A&A, 610, 16



C: low-mass AGB,  $< 4 M_{\odot}$   
 N: massive AGB,  $> 4 M_{\odot}$   
 O: core-collapse SNe

Currently, N/O ( $z < 2.5$ ),  
 C/O ( $z > 2$ ), but C/N is  
 possible with JWST!

# Extra-galactic Archaeology

- \* Internal structures, i.e., kinematics and 2D map of gas, stars, chemical abundances, are measured with Integral Field Units (SAURON, CALIFA, MaNGA, SAMI, Hector, MUSE, KMOS...)



- \* Chemodynamical simulations can predict their redshift evolution for JWST, and explain the physical origins to test the galaxy formation theory.