Development of laser comb to calibrate high dispersion spectrograph for radial velocity measurement

Hajime Inaba

National Institute of Advanced Industrial Science and Technology

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Direct detection of cosmic acceleration of the universe



For the detection of cosmic acceleration,

We need a measurement precision of the radial velocity at the cm/s level.



Wavelength standard for astronomical spectrographs



- 1. Wavelength precision
- 2. Narrow spectral linewidths
- 3. Appropriate and uniform spectral density
- 4. High wavelength coverage
- 5. High and uniform spectral intensity



Optical frequency comb as a wavelength standard



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Wavelength standard for astronomical spectrographs



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Comb sources for astronomical wavelength references



Setup of our developed laser comb (2nd gen.)





Novelty #1:Newly developed optical cavity design





Novelty #1:Newly developed optical cavity design

Fabry-Perot resonator for filtering comb-modes: resonator length can b changed while maintaining optical axis alignment (Pat. Pend.)

Robust optical fiber coupling

Setup of our developed laser comb





Novelty #2: A novel wavelength conversion scheme





Novelty #2: Novel wavelength conversion scheme



AIST

Wavelength / nm

Other novelties



in visible region

 \checkmark

 \checkmark

 \checkmark

[Novelties]

- **Robust Er-doped fiber laser** \checkmark
- \checkmark **Optical cavities for comb-mode** filtering(特願2019-147825)
- Mirrors with a low reflectivity for \checkmark stable mode-filtering
- **Optical cavity stabilization using** \checkmark a CW laser
- \checkmark Putting the final-stage cavity after the boost amplifier for high contrast
- Broadband comb generation in the optical wavelength (特願2019-034525)

WG-chirped

PPLN

Self-referencing comb referencing microwave frequency reference Realized 200 fs optical pulse by chirped pulse amplification etc...

Strong points

1. Broad spectrum with 30 GHz frequency interval (Wavelength coverage > 50 % in the optical region)



2. High contrast > 40 dB in the optical region



3. Practicality

Robust, reliable, use neither super-high-power amplifier nor short lifetime item



Strong point #1: Broad spectrum with 30 GHz frequency interval





Strong point #2: High contrast





Strong point #3: Practicality

1. Durability

Do NOT use short-life device (Tapered fiber, super-high-power amplifier, special mode-locker such as SESAM)

2. Robustness

Use an erbium-doped fiber-based mode-locked fiber laser as the comb source, Almost-all-polarization-maintain fiber optics

3. User-friendliness

Easy to alignment (for specialist), auto-relock system, remote controllable,



Laser combs for astronomical observation in the world

Increasing the comb mode interval		Spectral broadening / wavelength conversion of the comb			Blue : Optical Red : Infrarec
Group	Spectrometer	Comb source	Scheme	Spacing	Coverage
Menlo Systems	HARPS (R = 115,000)	Yb fiber (~1030 nm)	Mode filtering SHG/PCF	18 GHz	450-600 nm (35 %)
Harvard	HARPS-N (R = 115,000)	Ti:Sapphire (~945 nm)	PCF Mode filtering	16 GHz	500-620 nm (25 %)
	TRES		BBO Mode filtering	50 GHz	400-420 nm 780- <mark>880</mark> nm (13 %)
AIST #2	HIDES-F (R = 50,000)	Er fiber (~1560 nm)	Mode filtering HNLF, WG-PPLN	30 GHz	350-405 nm 455-535 nm (57%) 670-830 nm
NIST	Pathfinder (R = 50000)	Er fiber (~1560 nm)	Mode filtering HNLF	25 GHz	1450-1700 nm
	GIANO-B (R = 50,000)	Microresonator-based comb with a wide mode-spacing		23.7 GHz	1480-1550 nm 1570-1650 nm
	HPF (R = 50,000)	EOM-based comb with	λ conversion, Broadening	30 GHz	680-1650 nm (17 %)
TAT	IRD (R = 70,000)	a wide mode- spacing	λ conversion, Broadening	12.5 GHz	1040-1750 nm



The second-generation laser comb @Okayama



A laser comb spectrum observed with HIDES-F

Appropriate wavelength interval: 30 GHz High contrast: 40 dB High wavelength coverage: 57 % of the optical region



Remote operation is available



The laser comb is almost continuously operating for about 2-month so far.



Current operation status



The laser comb is available as a wavelength reference of HIDES-F!



Outcomes from Minoshima-group@UEC

Setup of AIST laser comb



C : Optical bandpass filter (1542 nm)

) : Circulator $(1 \rightarrow 2 \rightarrow 3)$

Polarizing beam splitter

----- : Electronics

----- , ----- , and : Non-PM fiber

— and ——— :PM fiber

Outcomes from Minoshima-group are for... Higher rep-rate laser for wider mode spacing More robust optical cavity



Low noise 750-MHz Yb:fiber frequency comb

> Why 750-MHz Yb: fiber frequency comb ?

DC-YD

- Large mode spacing
- Higher output power scalability



Broad SC generation

Launched power into the tapered-PCF : ~ 2 W

- High power for every comb mode
- Visible region SC generation

Oscillator, mode-locked Yb:fiber laser

- Repetition rate: 750 MHz
 - ✓ Highest in fiber based low noise combs
- Mode Locking: NPR
- Output power: >600 mW
- Pulse duration: ~70 fs
- Direct detection of f_{ceo} (SNR >40dB)

Low noise amplifier

- Output power: 10 W
- Pulse duration: ~90fs







Y. Nakajima, A. Nishiyama, K. Minoshima, Opt. Express 26 (2018)

All-fiber-based mode-filtering technique



Conclusion

1. We have developed the world's best comb.

- 2. The comb is now available in Okayama.
- 3. The comb is still evolving.



Thanks



Usuda Aoki









Minoshima

Kashiwagi



Hong





Tsutsui



Thank you for your attention !!



Astro-Comb First Performance Tests

★ Comb was observed right from the start with HIDES with high intensity.
Intensity was so strong that ND-filter was necessary

wavelength

* Due to the setup of the HIDES spectrograph at the time of commissioning and the use of additional band-pass filters, we could not observe the entire comb spectrum at the blue and red end

★ Wavelength coverage for the green region is essentially as expected. We observe the comb spectrum as low as 437nm

* Only few seconds exposure time needed compared to several minutes for the 1st Gen Comb



Red 670-830nm

Green 450-530nm

Blue 350-405nm

★ The comb was running with a stable mode lock over several hours taking alternate spectra from Th-Ar lamp and the comb itself

* We observed some breathing but with little effect on the absolute calibration over short time scales (several min)







* We fit the comb spectrum with single Gaussian models and compute the wavelength solution using the known position of each comb mode. We achieve an rms as low as 0.00020A for several well exposed orders which corresponds to an systematic uncertainty for the absolute calibration of 12 m/s



* The Astro comb has much better line coverage compared to the very few ThAr lines and we typically achieved a factor 10 better rms in the wavelength solution

* The relative accuracy over hour time scale is even better <10m/s for small wavelength regions

Summary

- First tests show very bright Astro-comb spectrum
- Expected wavelength coverage achieved
- Absolute wavelength calibration ~12 m/s systematic uncertainty in green region
- Relative accuracy <10m/s over hour time scale for small wavelength regions

Improvements: * HIDES Echelle grating not in optimal setup for the comb during test, HIDES is also not an Ultra-stable spectrograph

* absolute Wavelength calibration varies factor 2-4 from blue to red depending on available comb modes and their intensity