

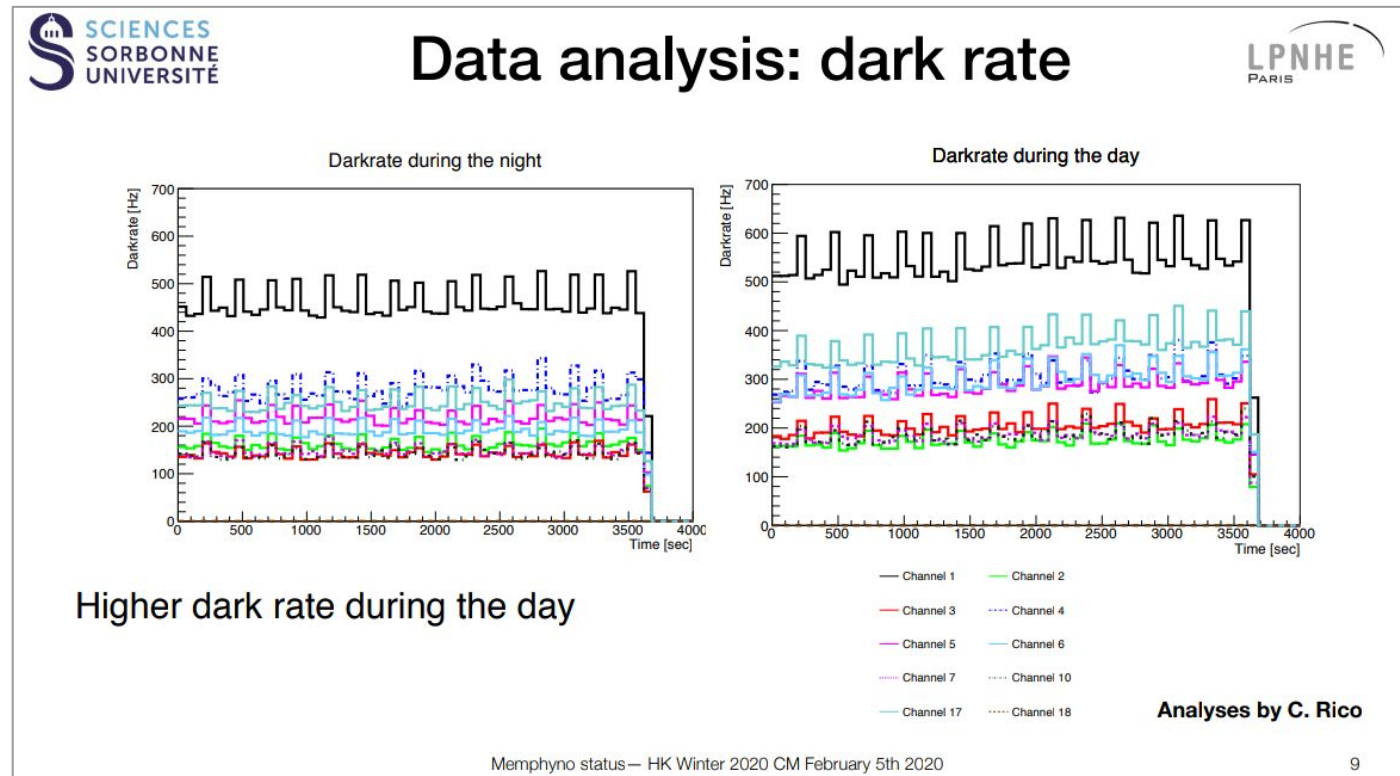
# Status Report

## MEMPHYNO darkrate measurements

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24 Apr. 2020  
mPMT Japan meeting

# Introduction

- Fixed light leak of the MEMPHYNO's water tank at APC and took the data for 2 weeks w/o any light source (only darkhits and cosmic muon)
- Analyzed the data in order to understand about the mPMT
  - Stability of the darkrate



# Analysis

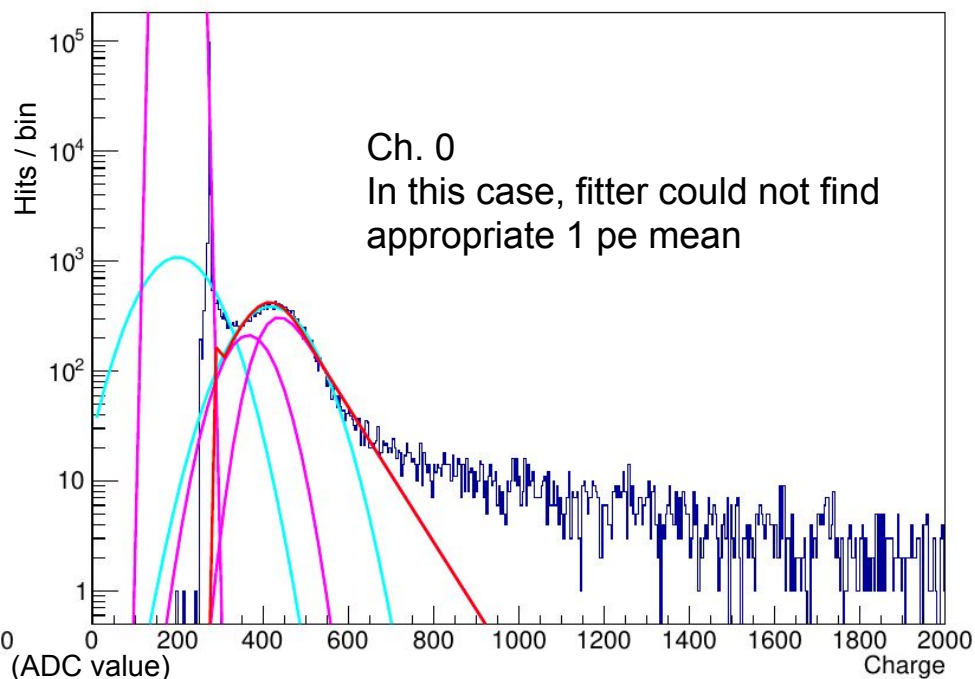
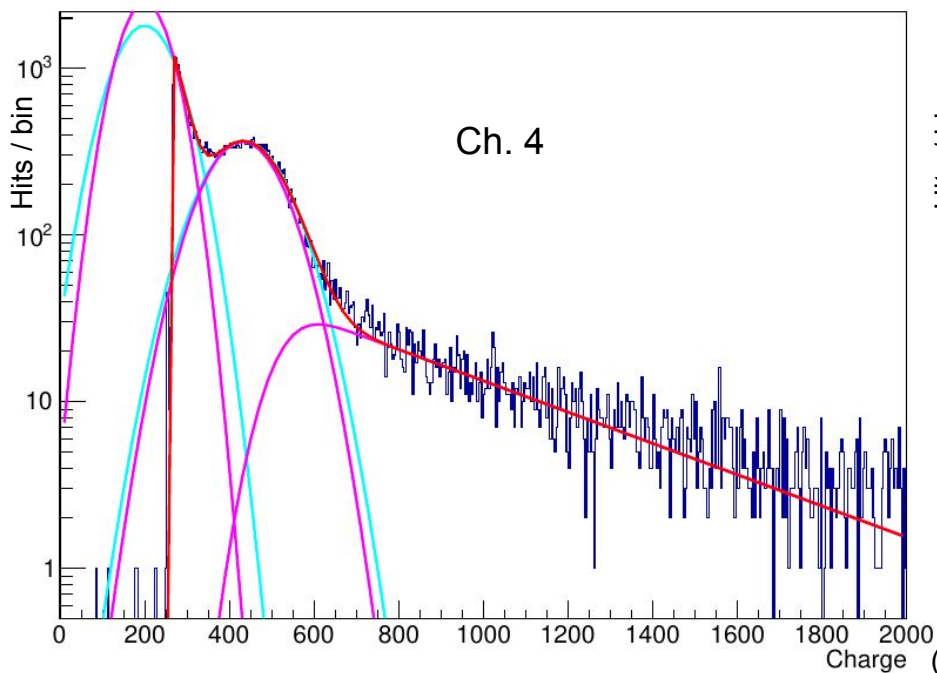
1. Divided all hits-data by the period = 100 sec
2. Made charge-distributions for the divided time-range
3. Found 1 p.e. mean with the fitting
4. Calculated the darkrate over all run-time
  - Defined the threshold as (1 p.e. mean) x 0.8

# Charge-distributions

- Fit with red function: like this

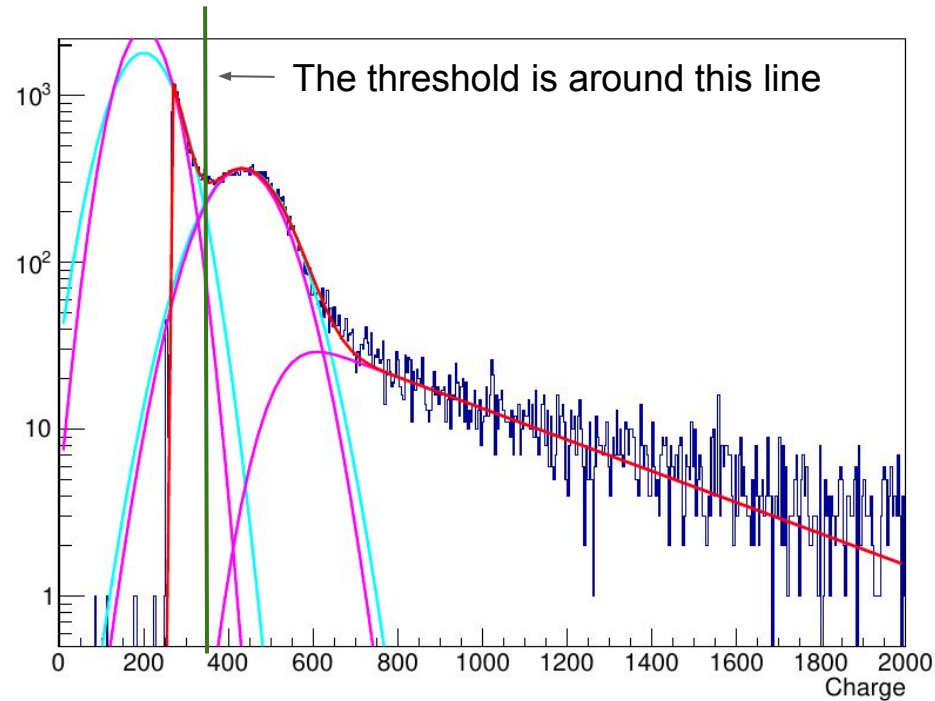
$$F(q) = \left( p_0 \times \text{gaus}_{\text{ped.}}(q) + p_1 \times \text{gaus}_{1\text{pe}}(q) + \frac{1}{2} (\text{Erf}(q/\sigma_0 - \alpha) + 1) \exp(q/\tau - q_0) \right) \times \frac{1}{2} \{ \text{Erf}(q/q_1 - \beta) + 1 \}$$

- Magenta plots show each component: pedestal, 1 p.e., exponential noise
  - Skyblue plots show the result of fitting with only one gaussian.
    - This fitting was done with a range that I defined.
  - Here I fixed the mean value of pedestal-gaussian because the fitter could not find this parameter
- However the fittings failed with some channels like right plots



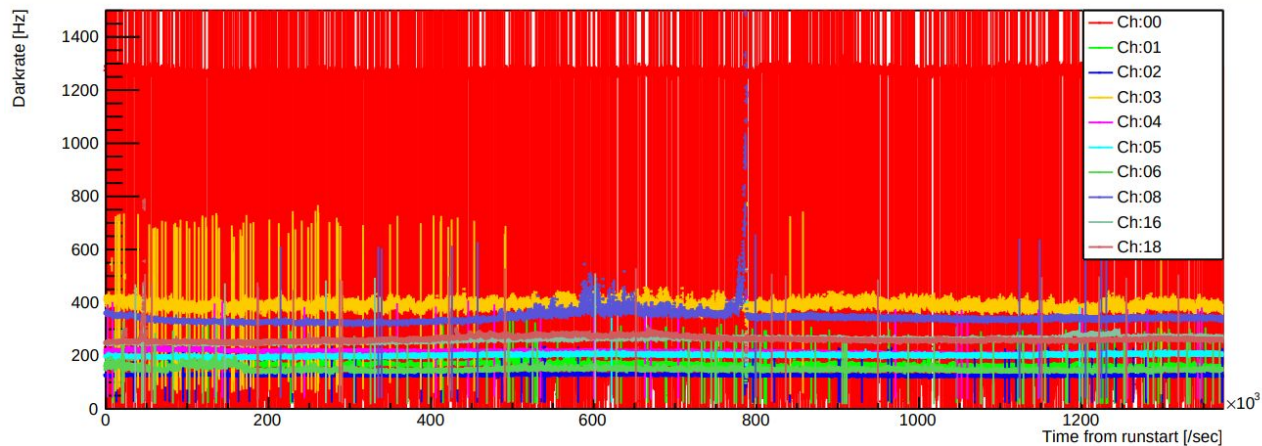
# Calculation of the darkrate

- Got the 1 p.e. mean and its error from the fitting
  - I tuned the HV so that 1 p.e. is around 430
- Defined the darkrate as:  $\text{darkrate} := \frac{(\text{total hits above 1pe mean} \times \alpha)}{\text{time}}$ 
  - In this slide, I set  $\alpha=0.8$  (~ 340 in ADC value)
- Added two errors
  - Stat.:  $\text{Sqrt}(\text{\# of hits}) / \text{time}$
  - Sys.: Difference when I changed 1 pe mean value by the fitting error
  - Typical contributions is O(1)% (stat.) and O(1)% (sys.)

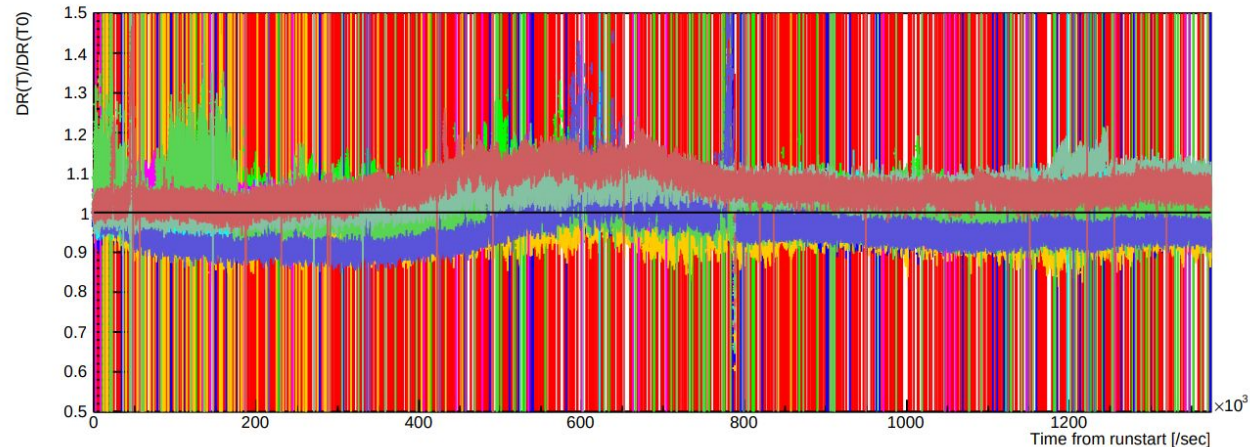


# Time trend of the darkrate

- Drew the trend for 2 weeks (  $\sim O(1e6)$  sec)
  - This covered large error bar (in particular red channel). I should remove these data points
  - Top plots show the absolute value [Hz] and bottoms show the ratio of run-start

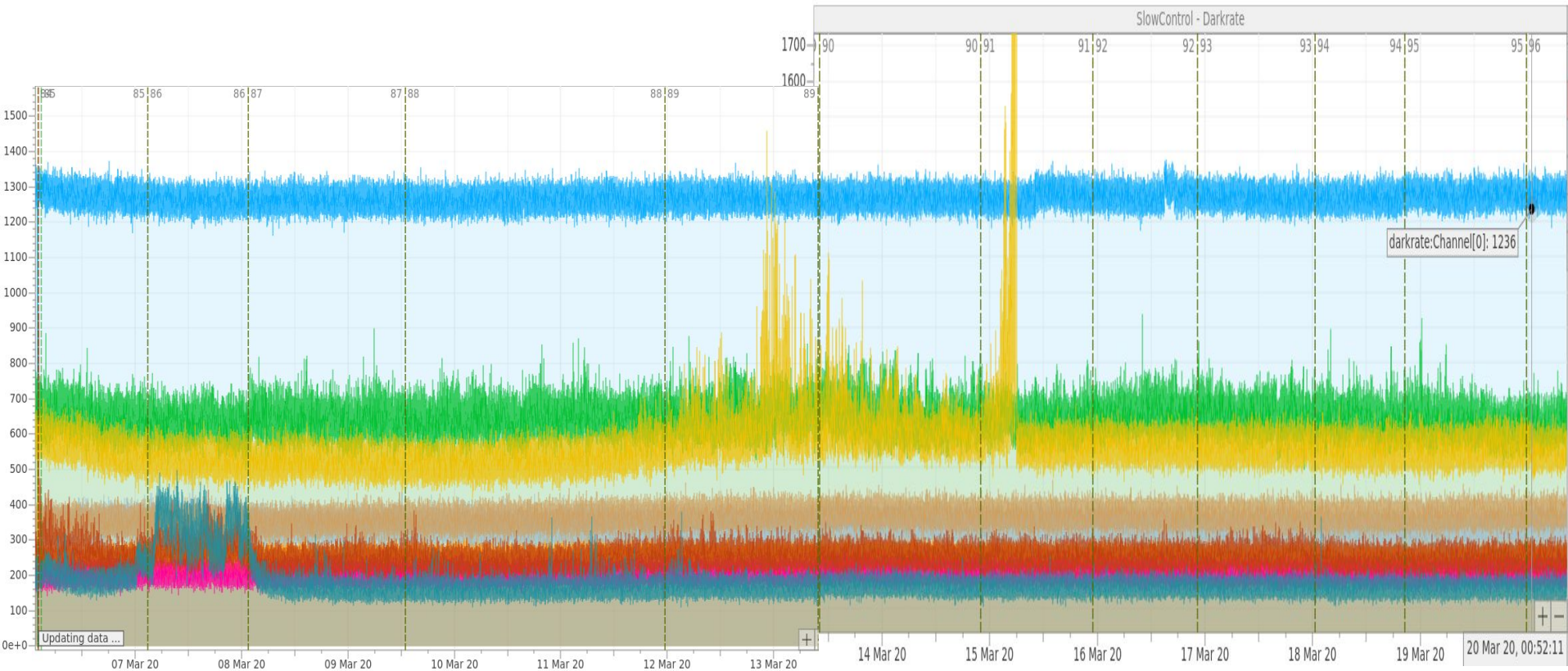


hDarkrateRatio\_0



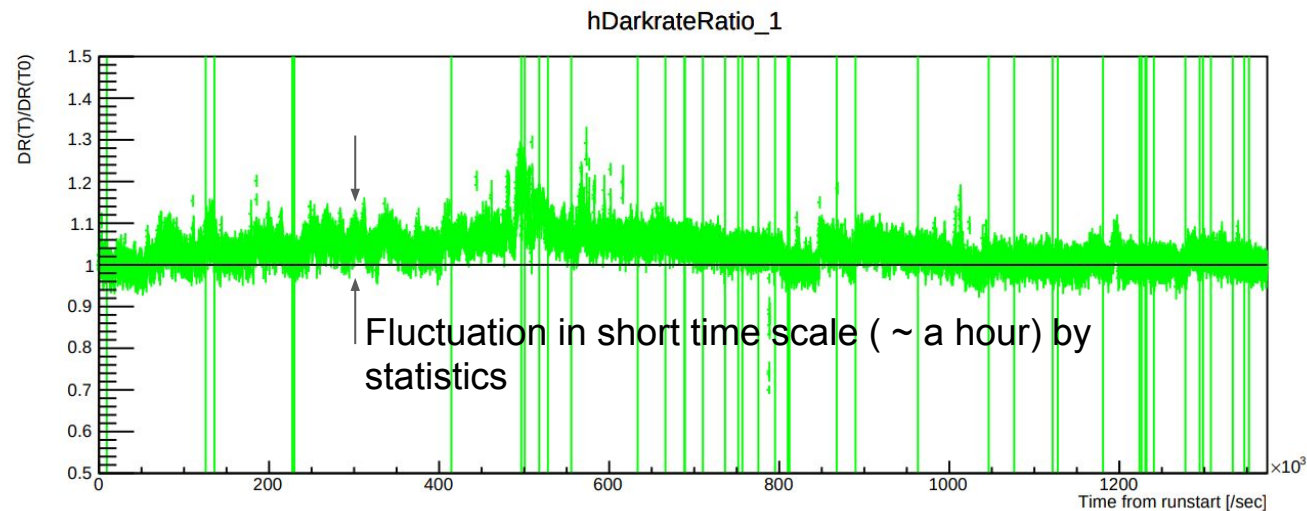
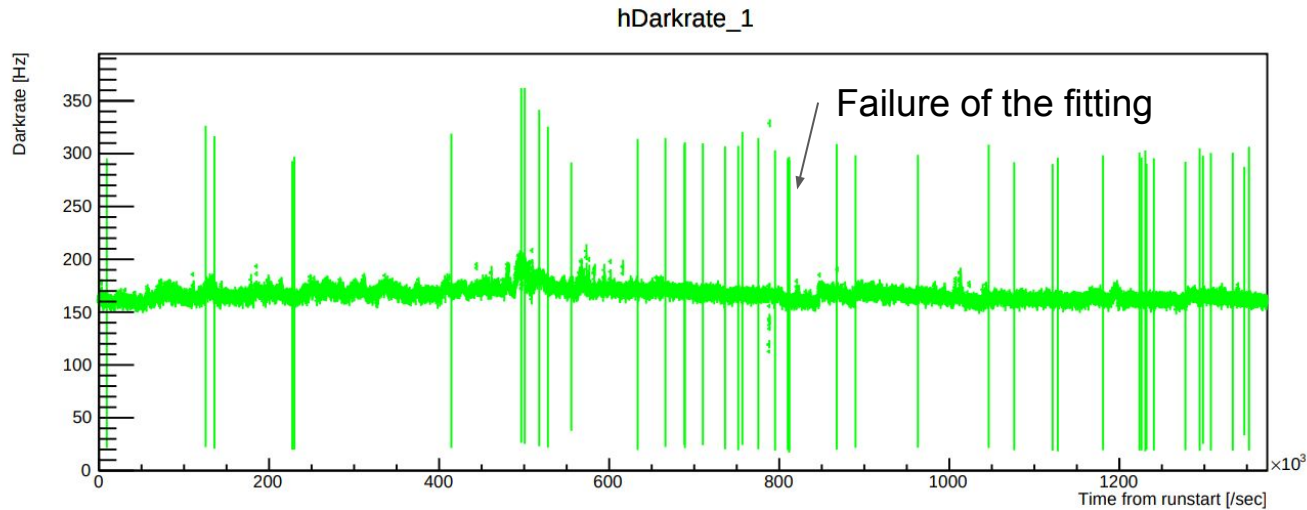
# Comparison with slow control monitor

- Main structures are consistent (the color annotation is not same with my plots)



# Time trend of the darkrate (ch 1)

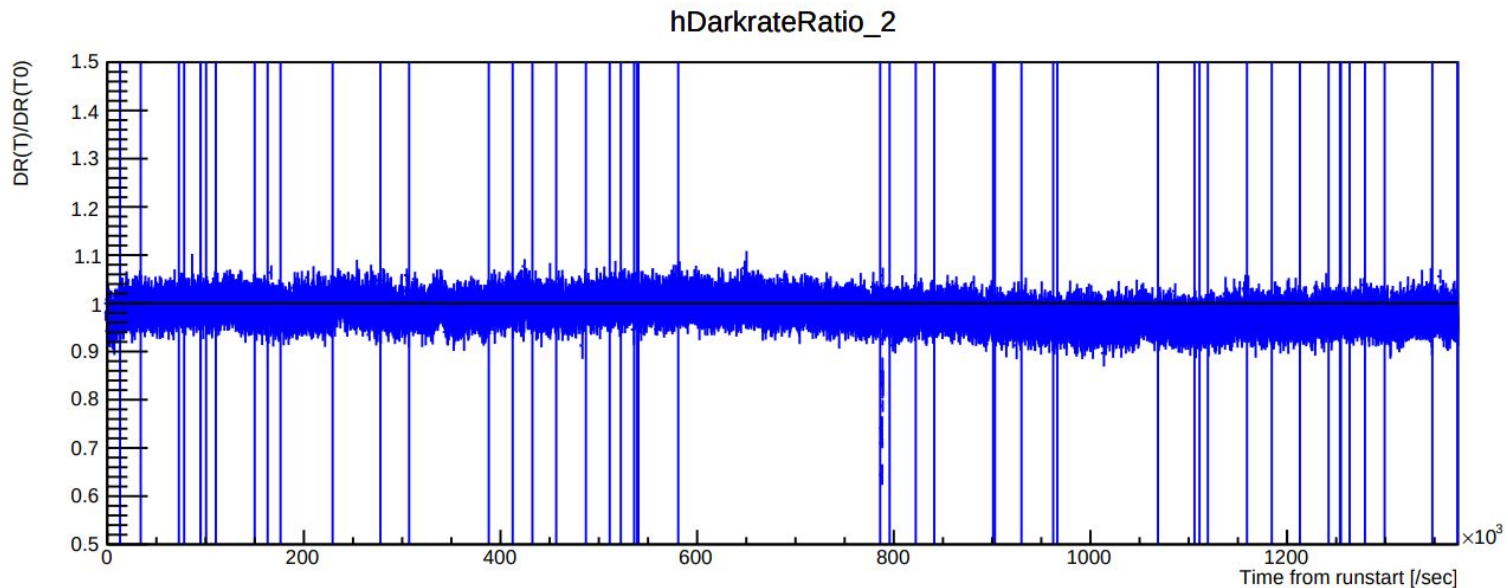
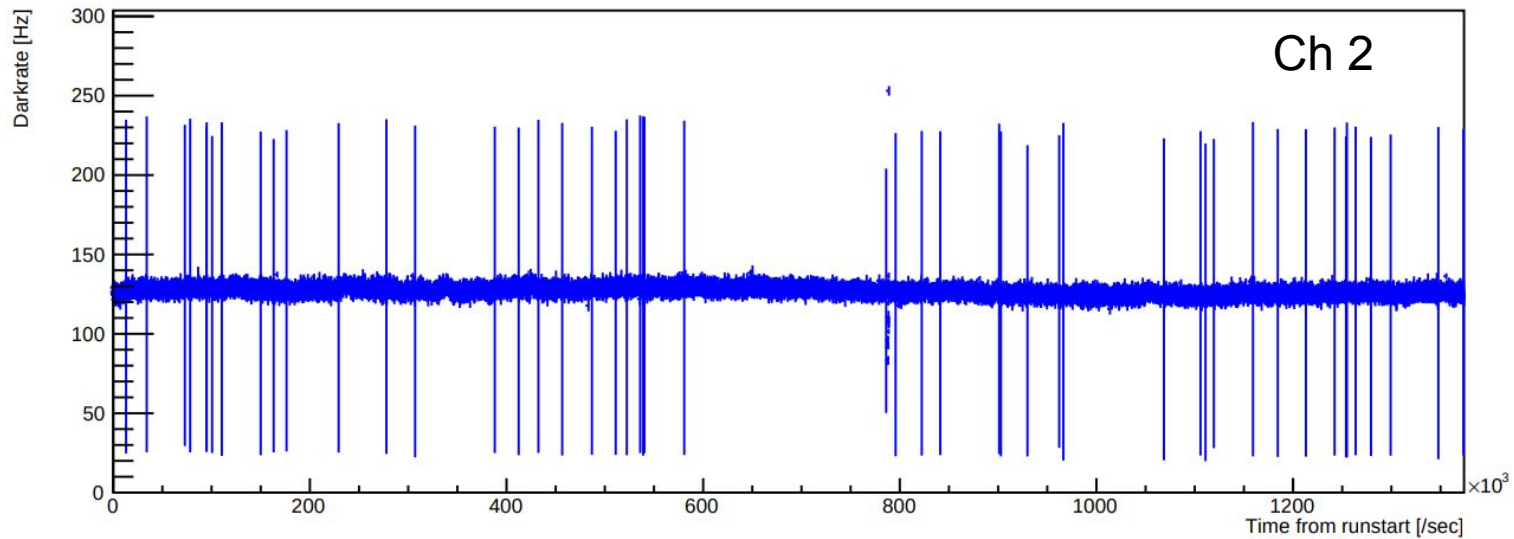
- Vertical lines, which appear sometimes: large error due to failure the fitting



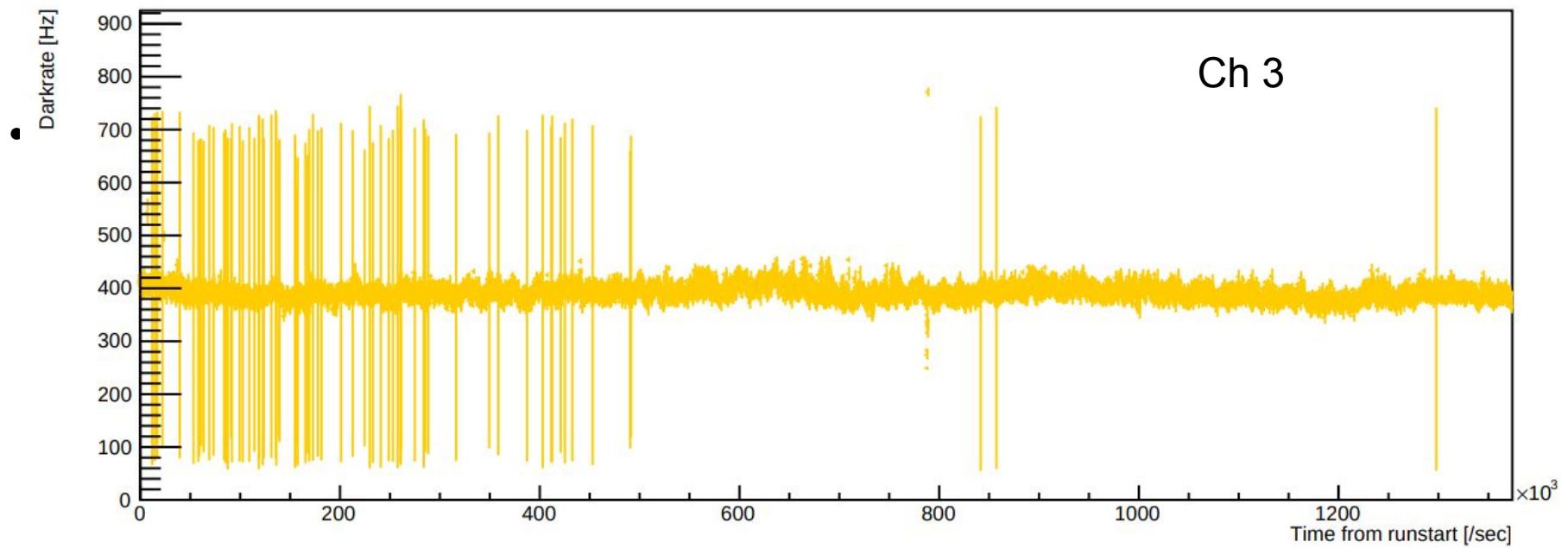


# Time trend of the darkrate (ch 2)

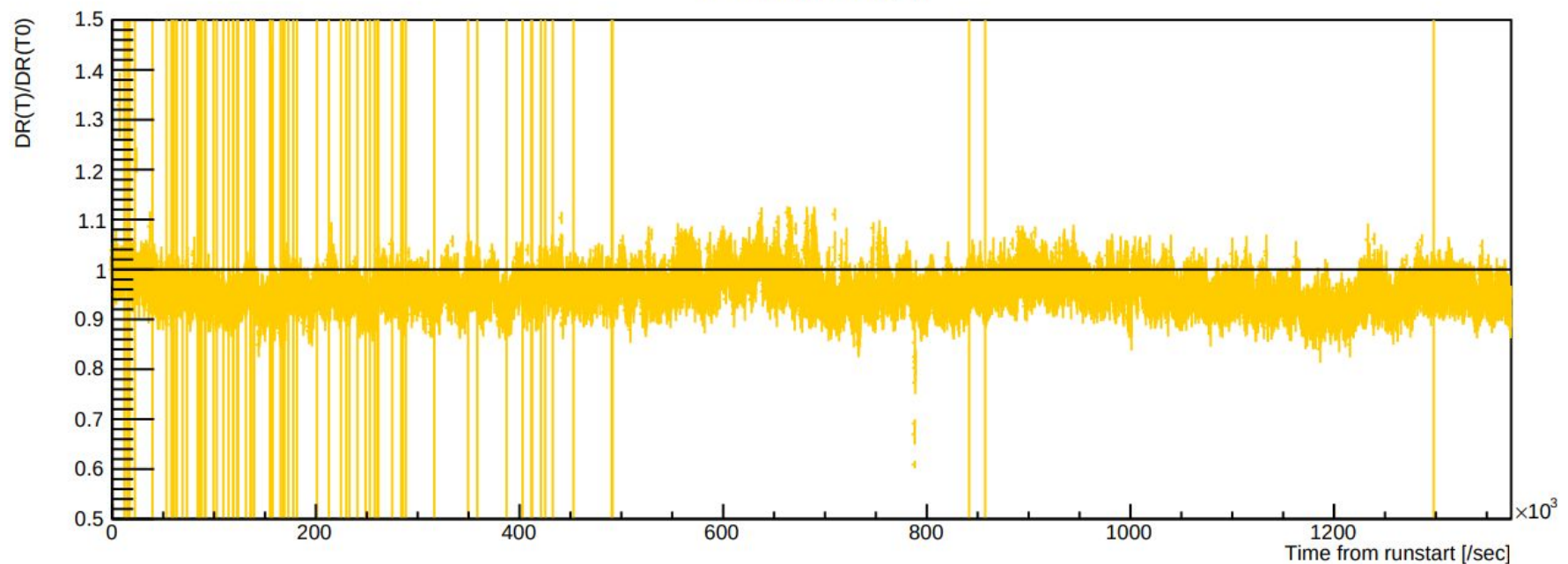
- No day/night difference, relatively stable channel



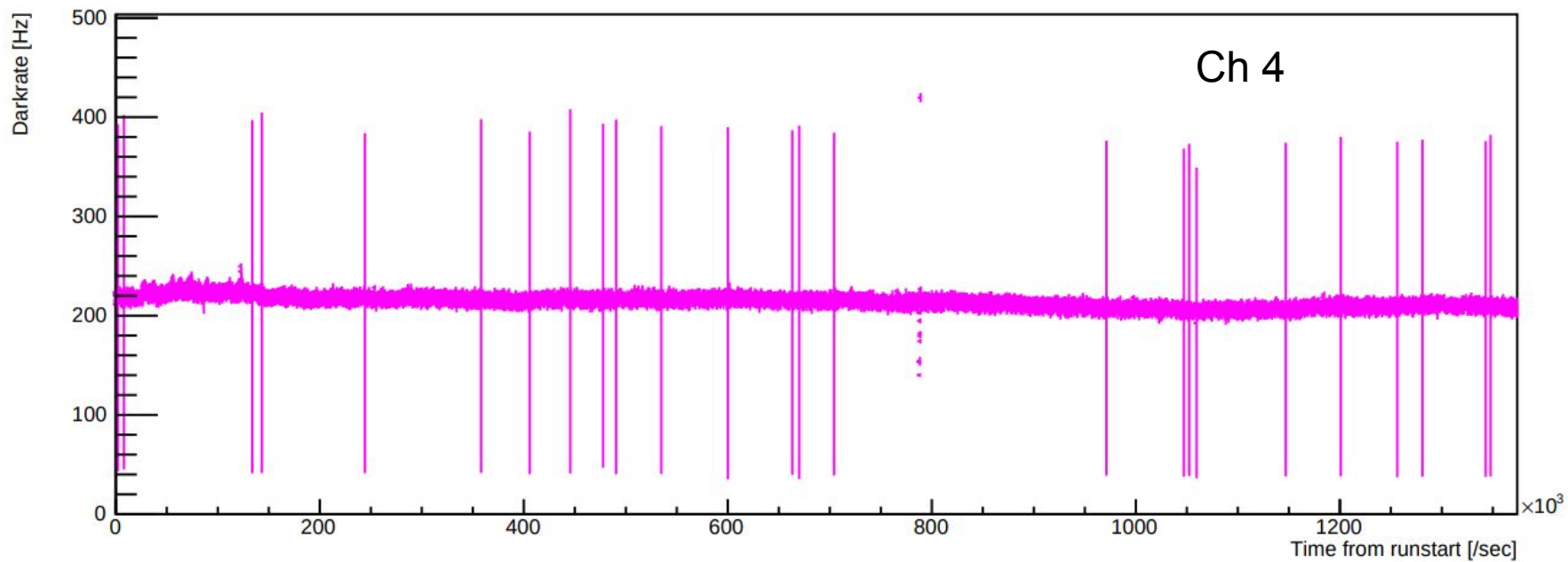
hDarkrate\_3



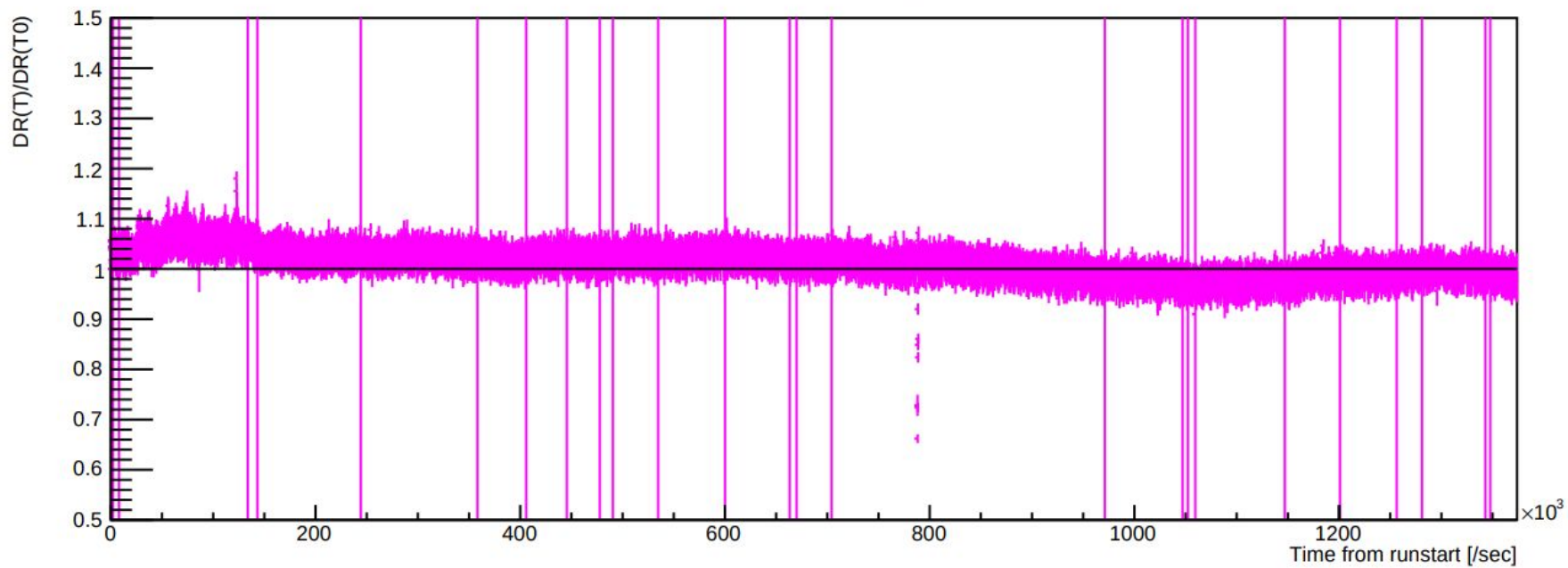
hDarkrateRatio\_3



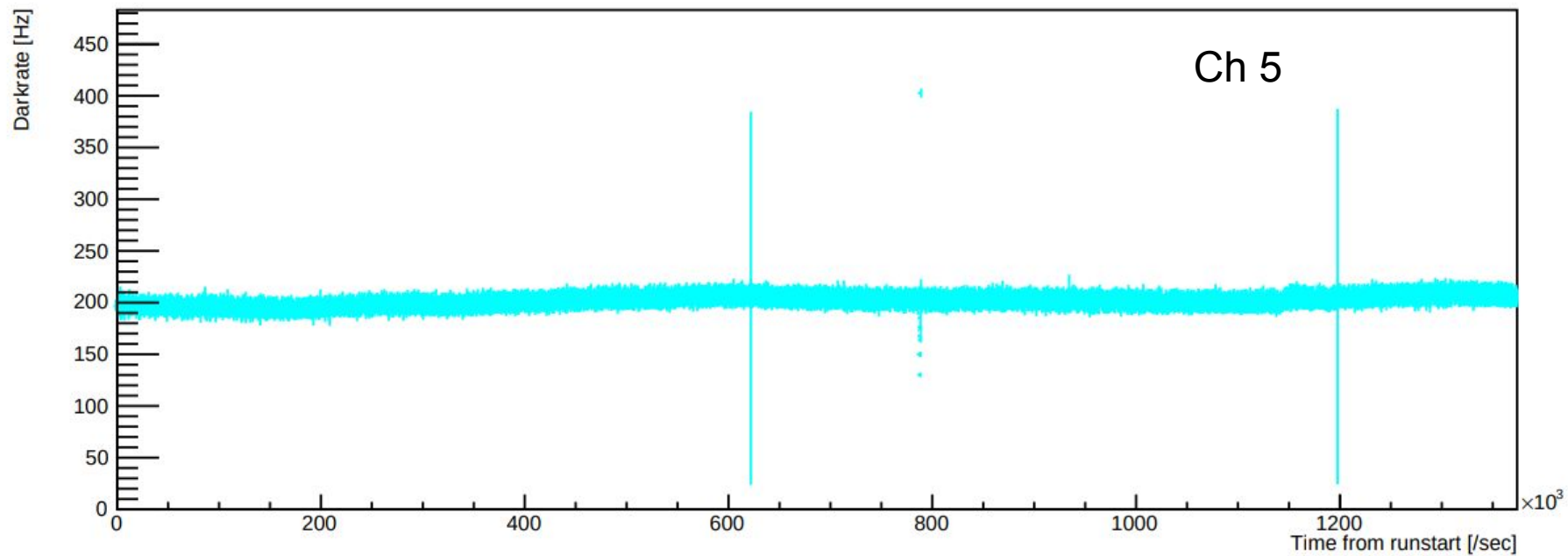
hDarkrate\_4



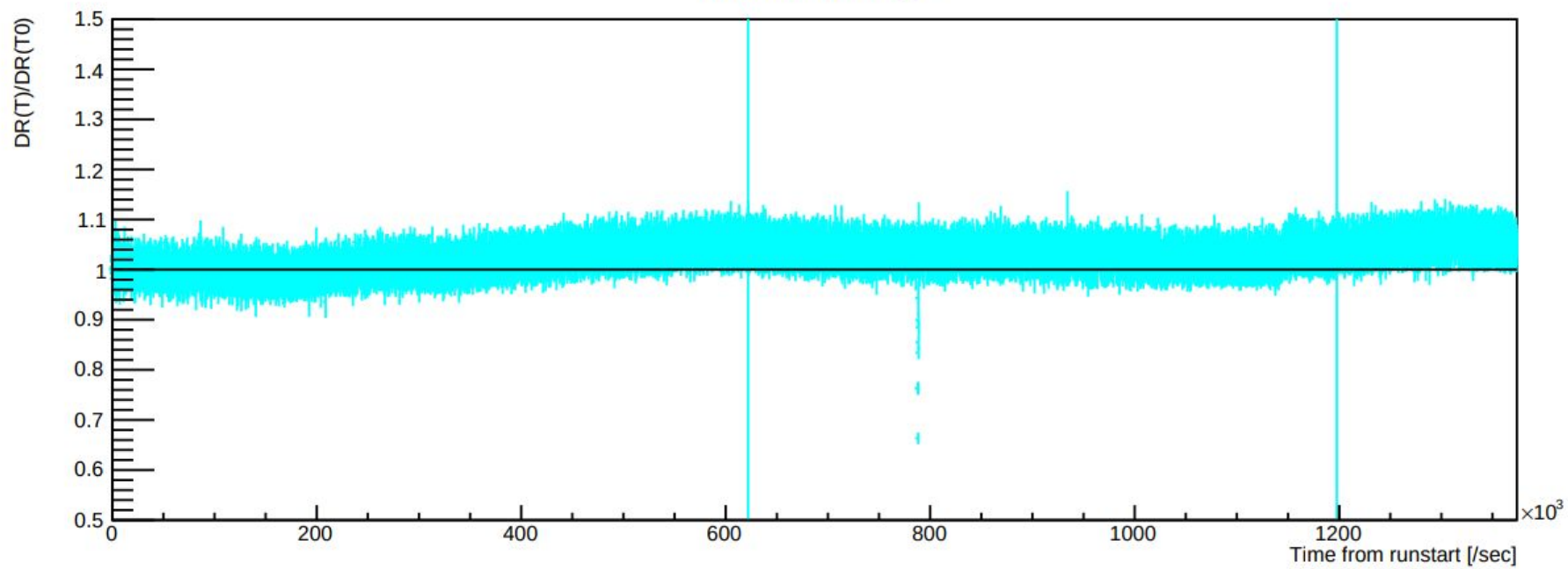
hDarkrateRatio\_4



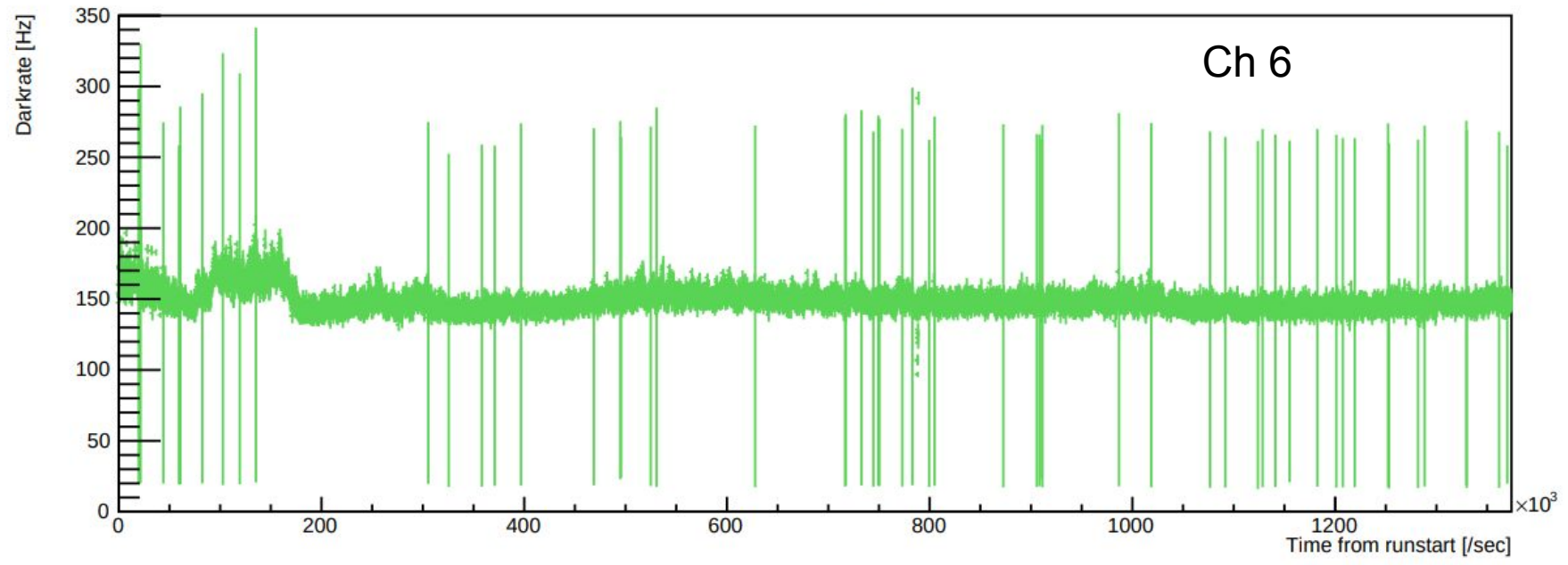
hDarkrate\_5



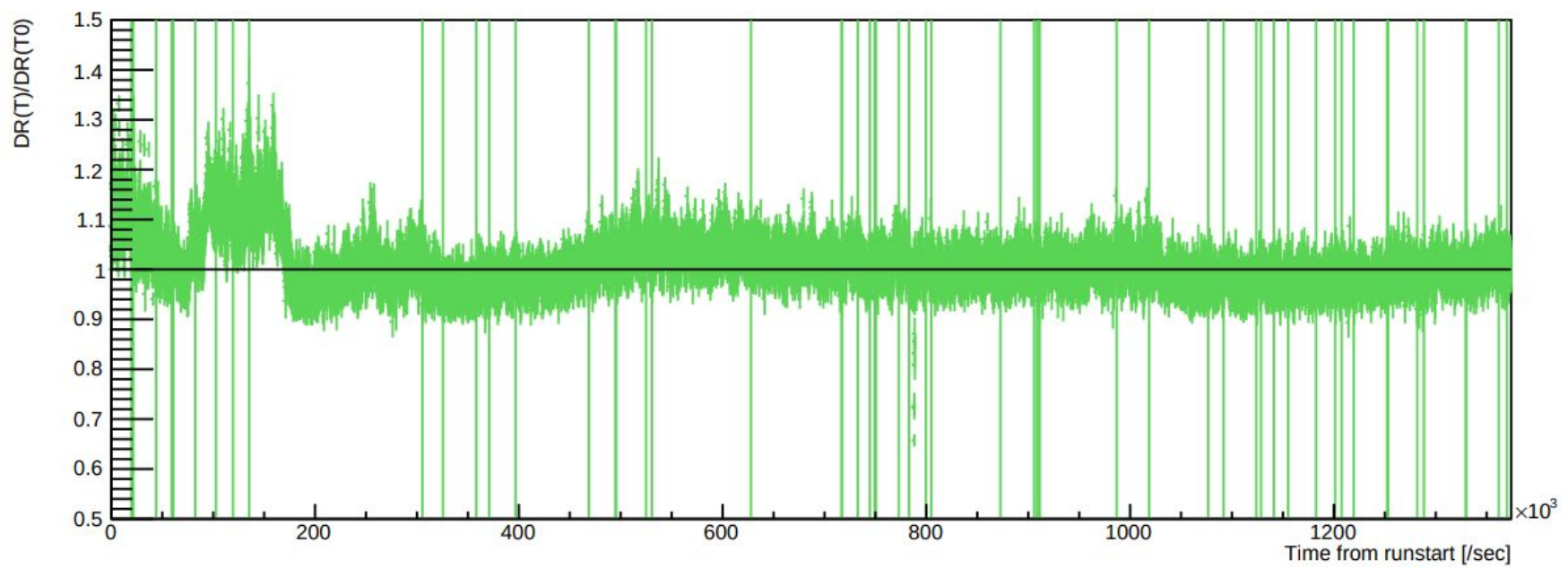
hDarkrateRatio\_5

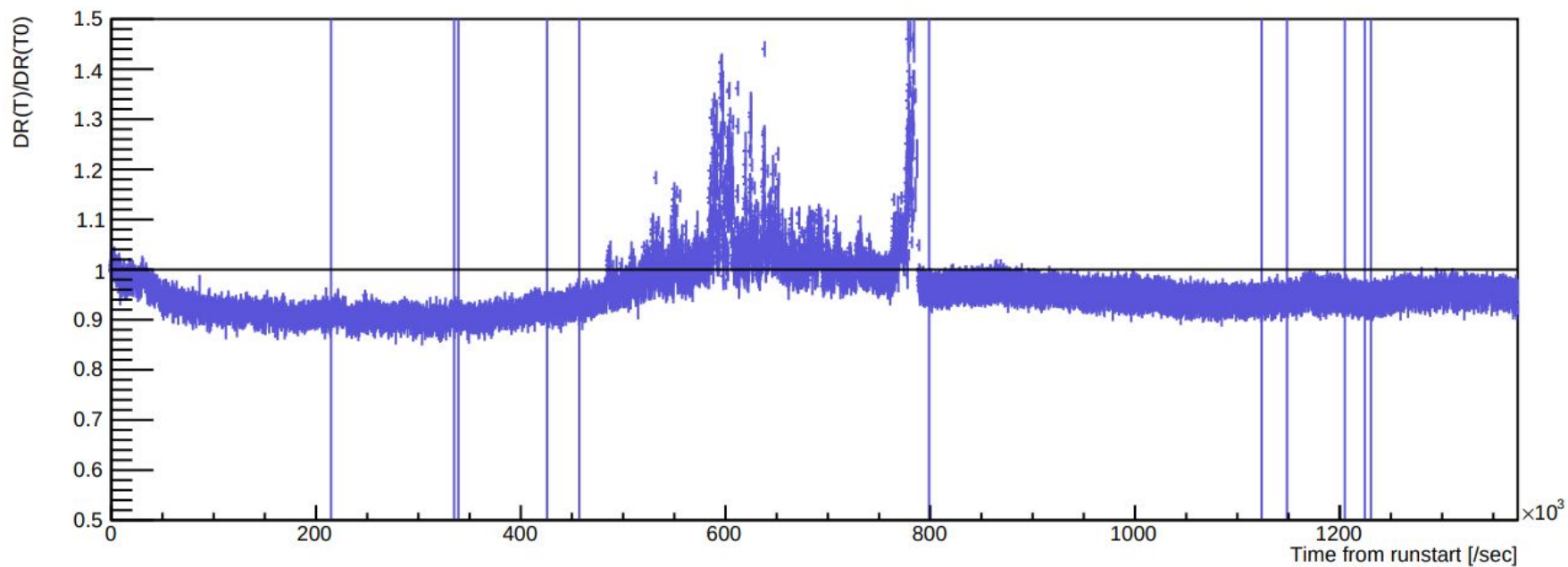
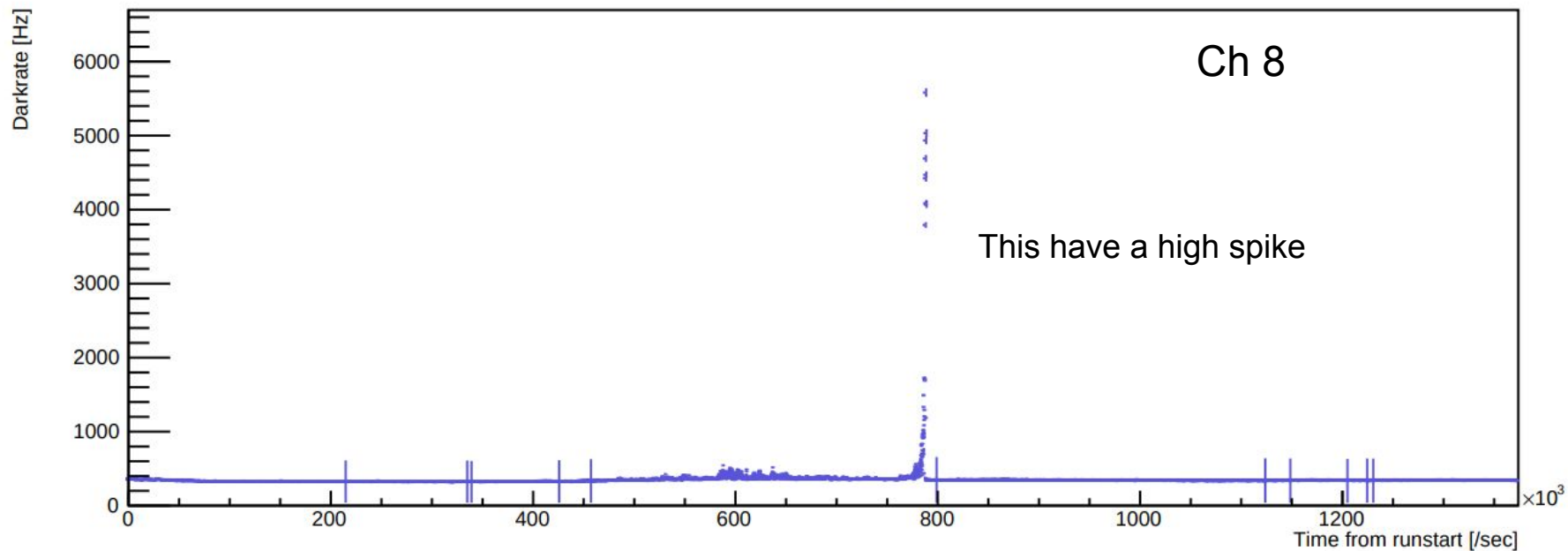


hDarkrate\_6

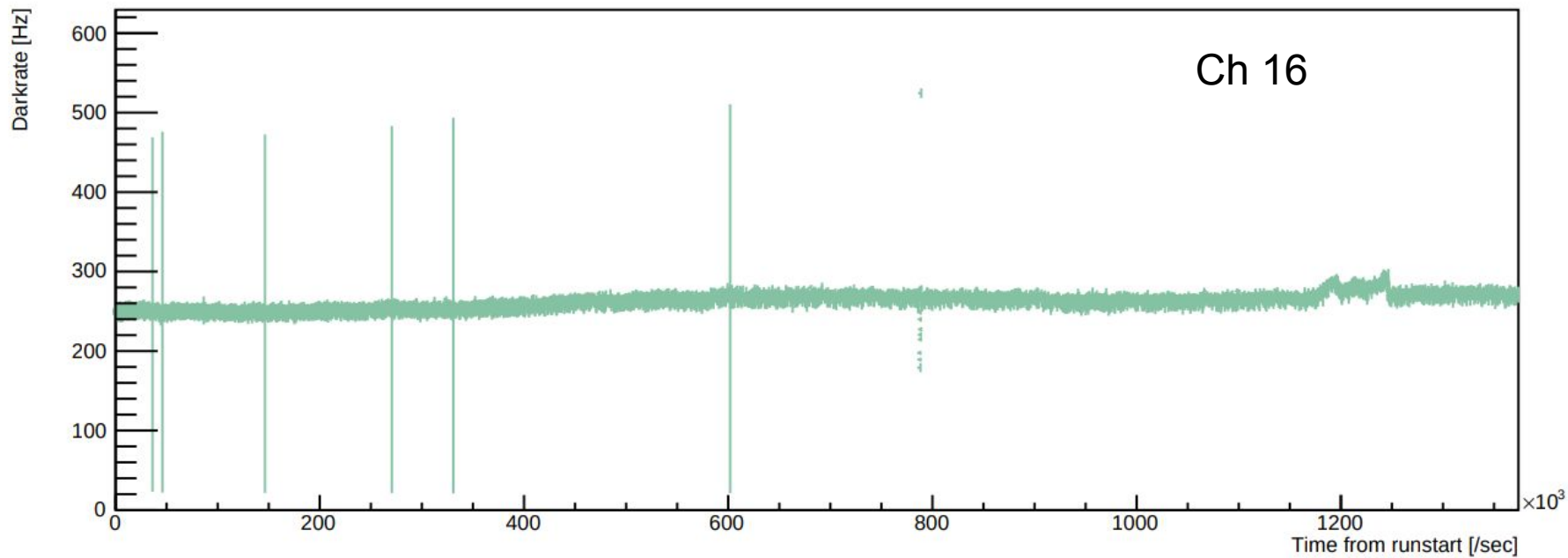


hDarkrateRatio\_6

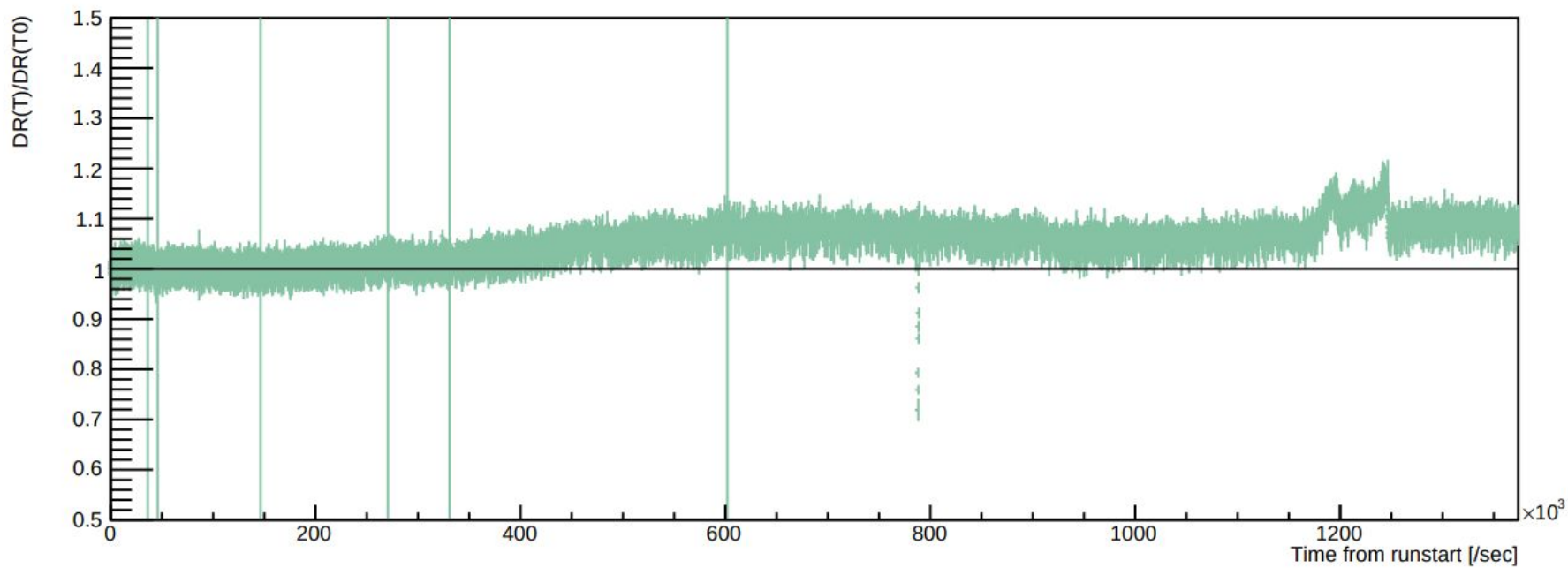


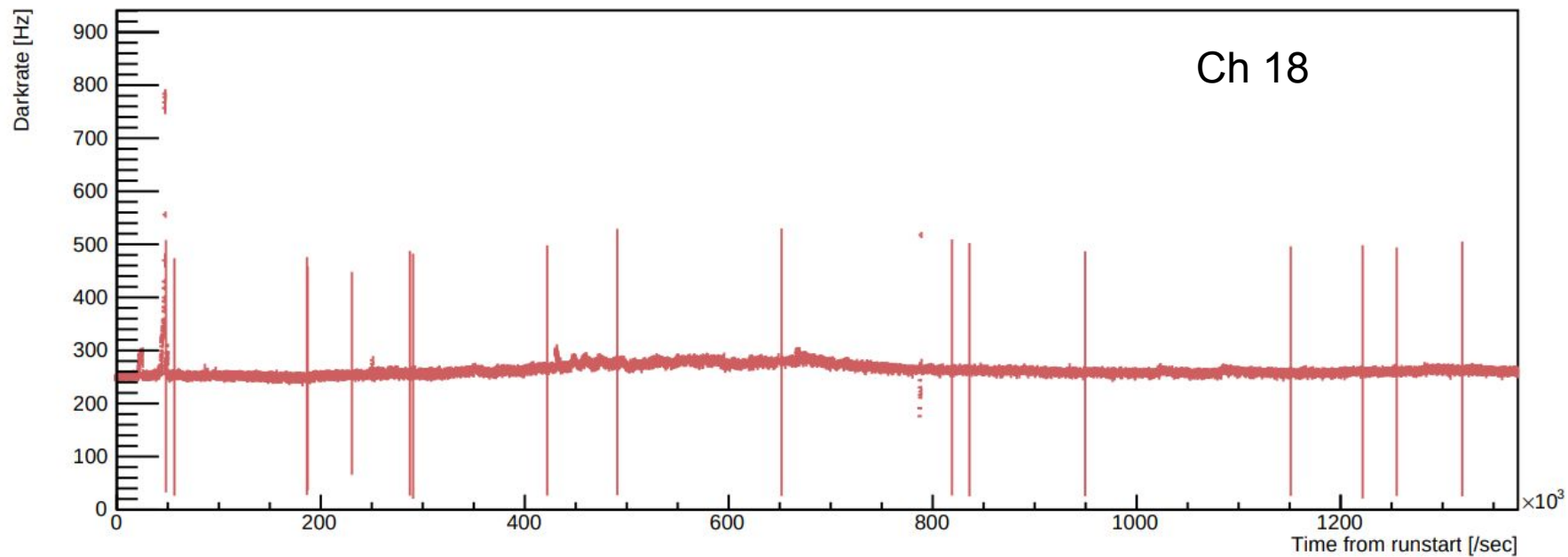


hDarkrate\_16

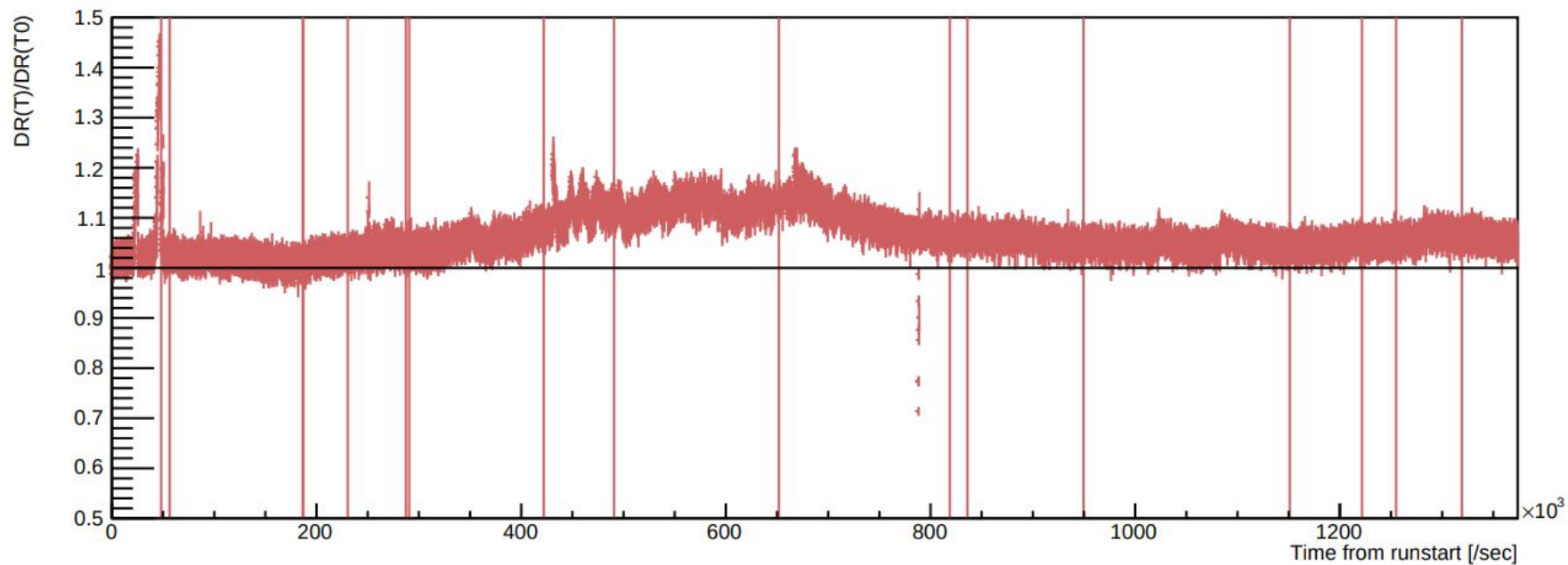


hDarkrateRatio\_16





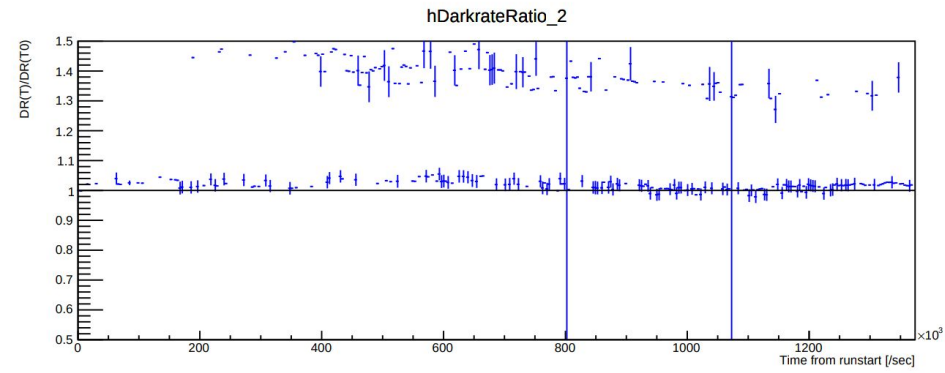
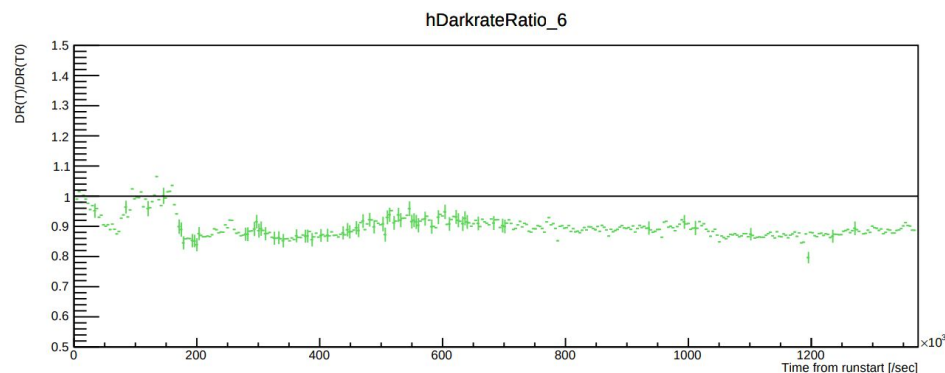
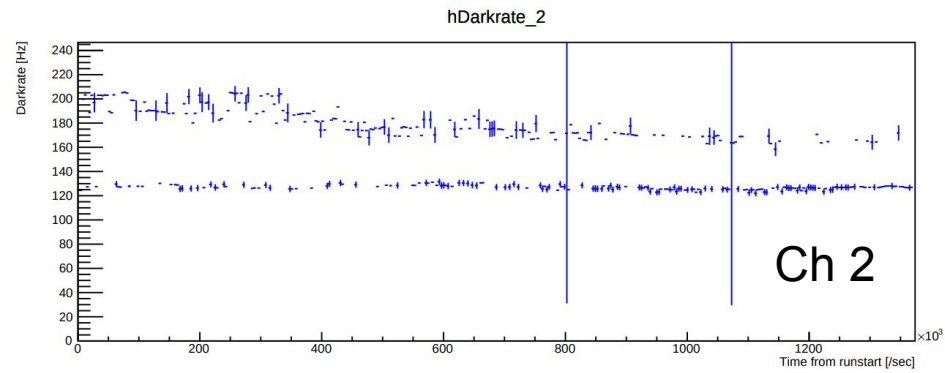
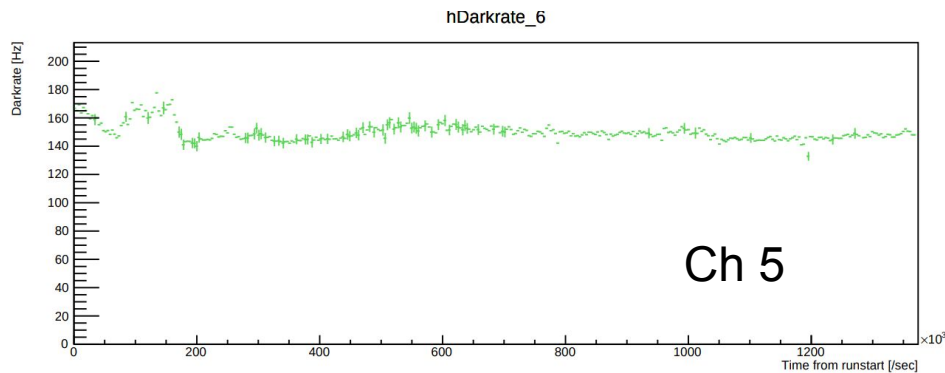
hDarkrateRatio\_18





# Changed time-period

- Changed the time period of the first step from 100 s to 1 h
- Some channels still had higher errors (like right plot)
  - I am now investigating and I have to be sure the fitter result



# Summary

- Plotted the time trends of the darkrate
- They are not so stable
  - We have many unknown structures (spikes, step etc.)

## Plan

- Make sure the fitter result
  - Now I checked only fitting error of the parameter (1 pe mean)
- Check the HV (I recorded the HV when changing the run)
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