

Adaptive Semi-linear Inversion: Sophisticated Lens Modeling of Large Strong-Lensing Data Sets

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Kavli IPMU: Galaxies and Cosmology in Light of Strong Lensing



Introduction

- Scientific Motivations
- Square and Adaptive Semi-linear Inversion
 - Overview of methods.
 - Comparison of methods.
 - Data discretization Unique and Unrelated Data Discretization.
- Strong Lens Analysis Pipeline
 - Multi-phase analysis Example.



Adaptive SLI - Overview

• Adaptive SLI has two primary goals:

1) Accurately perform modeling of degenerate lens models with just strong long data.

2) Do this in a fast, efficient and streamlined manner allowing analysis to be performed on the largest data sets possible.



Square Semi-linear Inversion (SLI)

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Simulated Image 1



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Step 1 – Trace Image Pixels to Source Plane

Simulated Image 1 Source Plane -2.2 0.07 0.2 0.06 0.15 0.05 -1.1 0.1 0.04 ,0.0 0 (arcsec) 0.0- V 0.05 0.03 y (arcsec) 0.02 0.01 -0.1 0 1.1 -0.15 -0.01 -0.2 -0.02 2.2 -1.1 -2.2 2.2 1.1 -0.2 -0.1 0.1 0.2 0 0 x (arcsec) x (arcsec)



Step 2 – Overlay Square Grid





Step 3 – Image / Source Pixel Mappings





Step 3 – Reconstruct Source via Matrix Inversion





Step 4 – Compare Observed and Reconstructed Images





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Step 2 – Input Into h-means Clustering Algorithm





Step 3 – Image / Source Pixel Mappings

Simulated Image 1 Adaptive Source Plane -2.2 0.07 0.2 0.06 0.15 0.05 -1.1 0.1 0.04 ,0.0 0 (arcsec) 0.0- V 0.05 y (arcsec) 0.03 0.02 0.01 -0.1 0 1.1 -0.15 -0.01 -0.2 -0.02 2.2 -2.2 -1.1 2.2 1.1 -0.1 0.1 0.2 -0.2 0 0 x (arcsec) x (arcsec)



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Additional Inversion Caveats

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- Nearest-Neighbor regularization scheme used which penalizes non-constant solutions.
 - Forces a physical source reconstruction.
 - Prevent over fitting to image noise.
- Maximize Bayesian Evidence (Suyu et al. 2006) to set level of Regularization for every lens model.



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• Every trial lens model returns a reconstructed image

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 - Converts χ^2 and regularization to a global evidence ϵ .
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 - Highest evidence ϵ -> most probable lens model.
- Standard non-linear search problem.

- Find combinations of lens parameters (x, y, $\sigma,$ q, $\Phi,$ $\alpha)$ which maximize $\epsilon.$

- MultiNest (Feroz et al. 2009) – based on nested sampling.



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 - Fewer source pixels required -> faster inversion.
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- Can make source plane of arbitrarily large size
 - No source plane set up required -> streamline inversion.
- Discretization the source plane in a *Unique* way for every lens model.
 - Data discretization free of any geometric prescription.
 - Every Pixelization is derived from a random initalization \rightarrow Unrelated data discretization.
 - Data discretization is different for every single inversion.



Fitting a Power-Law (PL) Lens Model Heavy degeneracy between σ and α .

Input model: $\alpha = 2.0$



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Fitting a Power-Law (PL) Lens Model Heavy degeneracy between σ and α .

Reconstructed Image 1: σ = 276 km/s, q = 0.76, α = 2.1 Reconstructed Image 1: σ = 297 km/s, q = 0.84, α = 1.9 -2.2 -2.2 0.06 0.06 0.05 0.05 -1.1 -1.1 0.04 0.04 y (arcsec) y (arcsec) 0.03 0.03 0 0 0.02 0.02 1.1 1.1 0.01 0.01 2.2 2.2 -2.2 -1.1 1.1 2.2 -2.2 -1.1 1.1 2.2 0 0 x (arcsec) x (arcsec)

α = 1.9

α = 2.1

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Data Discreteness Bias

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- With the NDOF fixed, Square SLI could not accurately fit a PL lens model.
 - Different lens model calculated if inversion setup is changed.
 - Results varied with source plane resolution, size and position.
- Investigation -> why do results change with setup?
 - Fit 9 PL lens models to simulated image shown previously.
 - All 9 inversions identical, except source plane position.
 - Phase shift each grid by a fractional pixel interval.





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Phase Shifts – α Estimation

PDF of α for Image 1 using 9 Phase Shifted Square Grids





Phase Shifts – α Estimation

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Simulated Image 2





Phase Shifts – α Estimation

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Phase Shifts – α Estimation

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Data Discreteness Bias

- What is the same in every phase shifted run? Nearly everything:
 - Observations Image, Noise Map, PSF,
 - Inversion Setup Source Resolution, Image Subgridding, etc
 - Method Inversion process, Regularization scheme etc.



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- What is actually changing?
 - Data Discretization
 - Each phase shifts changes the accessible allocations of image pixels to source pixels.



Adaptive Source Plane 1 - σ = 284990 m/s



 5 lens model – all parameters fixed except σ which is changed by 5 m/s in each.

$$\epsilon_1 = -1599.773$$



Adaptive Source Plane 1 - σ = 284995 m/s



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Adaptive Source Plane 1 - σ = 285000 m/s



- 5 lens model all parameters fixed except σ which is changed by 5 m/s in each.
 - $-\epsilon_1 = -1599.773$
 - $-\epsilon_2 = -1606.381$
 - $\epsilon_{3} = -1598.001$



Adaptive Source Plane 1 - σ = 285005 m/s



- 5 lens model all parameters fixed except σ which is changed by 5 m/s in each.
 - $-\epsilon_1 = -1599.773$
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 - ε₄ = -1604.852



Adaptive Source Plane 1 - σ = 285010 m/s



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 - $-\epsilon_4 = -1604.852$
 - $\epsilon_{5} = -1599.654$



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 - Random initialization -> every pixelization completely different.
 - Other methods have a changing pixelization, however all are derived from the lens model \rightarrow all are 'related'.
 - Tested using adaptive SLI with fixed intialization -> discreteness biases return (and amplify?)



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 - Tested using adaptive SLI with fixed intialization -> discreteness biases return (and amplify?)
- Adaptive SLI naturally explores the underlying systematics.
 - More natural error sampling.



Adaptive SLI – α Estimation

PDF of α for Image 1 using 9 Phase Shifted Square Grids





Adaptive SLI – α Estimation

PDF of α for Image 2 using 9 Phase Shifted Square Grids



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Strong Lensing Analysis Pipeline (Preliminary Work)

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Pipeline Development – 3 Phase Inversion

- Phase 1 Fast inversion used to explore parameter space and map out degeneracy.
 - Use parametric source inversion.
 - Maps out lens model and degeneracies -> Initializes phase 2.



Pipeline Development – Phase 1





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 - Maps out lens model and degeneracies \rightarrow Initializes phase 2.
- Phase 2 Setup Adaptive SLI with 'fast' settings.
 - Lower Source resolution.
 - Lower Image Subgridding.
 - Faster evaluation of Bayesian evidence.
 - Further maps out lens model \rightarrow Initializes phase 3.



Pipeline Development – Phase 2





Pipeline Development – 3 Phase Inversion

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- Phase 2 Setup Adaptive SLI with 'fast' settings.
 - Lower Source resolution.
 - Lower Image Subgridding.
 - Faster evaluation of Bayesian evidence.
 - Further maps out lens model \rightarrow Initializes phase 3.
- Phase 3 Full Adaptive SLI
 - Use Adaptive SLI as presented here.
 - Accurate lens model and fully sampled errors.



Pipeline Development – Phase 3





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- Must rigorously test on both observational and simulated data.
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 - Compare with other modeling \rightarrow learn more about systematics.
- How streamed lined can we go Depends on what you want?
 - Reliable source reconstructions?
 - Accurate fitting of degenerate lens models?
 - Extremely precise modeling for cosmology or substructure?
- When will modeler input be required?



Summary

 Adaptive SLI – A new method for the modeling of strong lens data.

- Creates an **unrelated** pixelization in a **unique** way for every lens model.

- Removes data discretization biases.
- More natural error sampling.
- Strong lens analysis pipeline.
 - Modeling process can be streamlined...
 - But how streamlined and for what purpose?

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Additional Slides



Example – Poor Lens Model





Fixing NDOF



Bias 1 – Varying NDOF

• Degeneracy – Gives isotropic source scaling



- Source expands -> Fewer image pixels trace into source plane.
- Solutions biased to those which minimize the number of traced image pixels -> Must fix NDOF.



Fix NDOF

• NDOF must be fixed for every possible lens model.



- Large source plane, narrow priors and tightened image mask.
- inefficient resolution, pixels constrained by just regularization and varying effective resolution.


Fix NDOF

• NDOF must be fixed for every possible lens model.



- Can set Adaptive SLI source plane to arbritarily large size.
- Circumvents all previous problems



Noisey / Fluctuating Parameter Space



Calculation of Input Lens Model for Square and Adaptive SLI with σ iterated over steps of 1 m/s





Calculation of Input Lens Model for Fixed Initialization Adaptive grid and σ iterated over steps of 1 m/s





Discretization Bias With Related Pixelization









PDF of α for Image 1 using 9 Fixed Initialization Adaptive Grids







Square / Adaptive Discretization Comparison





$$\epsilon_1 = -1599.877$$





$$\epsilon_1 = -1599.877$$





$$-\epsilon_1 = -1599.877$$

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- ε₂ = -1599.026
- $\epsilon_{3} = -1599.000$
- ε₄ = -1599.390
- ε₅ = -1599.705



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- Pixelization adapts to lens magnification.
 - Lower source resolution required -> Faster and more efficient.
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 - Minimize user setup pre-inversion -> streamline process.
 - Can use large priors without worry of reducing NDOF.
- Naturally changes data discretization, removing alignment biases.
 - Accurately fits degenerate lens models like a PL.
 - Errors determined purely by fit of model and observation.