





Strong lensing at sub-mm wavelengths: results from Herschel follow up programs

Julie Wardlow



Herschel/SPIRE: 250, 350, 500µm

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Calanog et al. 2014

Why survey the far-IR?



Strong lensing from Herschel follow up programs

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HerMES data: GOODS-N



Lensed galaxies are readily identifiable in the far-IR



Strong lensing from Herschel follow up programs

HerMES lens selection (1): S₅₀₀>100mJy



Wardlow et al. 2013

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HerMES lens selection (2) S500>100mJy & no blazars or local spirals



Strong lensing from Herschel follow up programs

See also Negrello et al. 2013 Julie Wardlow

At many λ lensed SMGs easily separated from lenses



N(z): typically z=2–4 SMGs lensed by z<1.5 galaxies



A sample of Herschel lens systems



Strong lensing from Herschel follow up programs

Lens candidates have ~35–75% fidelity

Blazars & spirals removed



Wardlow et al. 2013

HBoötes02: a radio-loud AGN in a lensed SMG

edge-on spiral

z=0.41



z=2.80

Lensed: HATLAS12-00 @ z=3.3



Observed

Fu et al. 2012

Strong lensing from Herschel follow up programs

FIR magnification factors are ~a few



Adapted from Bussmann et al. 2013

Dust regions in lensed SMGs are ~0.5–3 kpc radius



The submm emission is typically more magnified & smaller than the NIR



Summary

Wide-area, submm surveys can efficiently identify strongly lensed high-redshift galaxies by simply selecting the brightest sources.



Herschel selected galaxy-galaxy lenses are proving useful in probing obscured star-formation.



Strong lensing from Herschel follow up programs