

# *H0LiCOW*

$H_0$  Lenses in COSMOGRAIL Wellspring

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Galaxies and Cosmology in Light of Strong Lensing  
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# H0LiCOWers

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**Chiara Spiniello** (MPA, Garching)

**Malte Tewes** (AlfA, Bonn)

**Tommaso Treu** (UCLA)

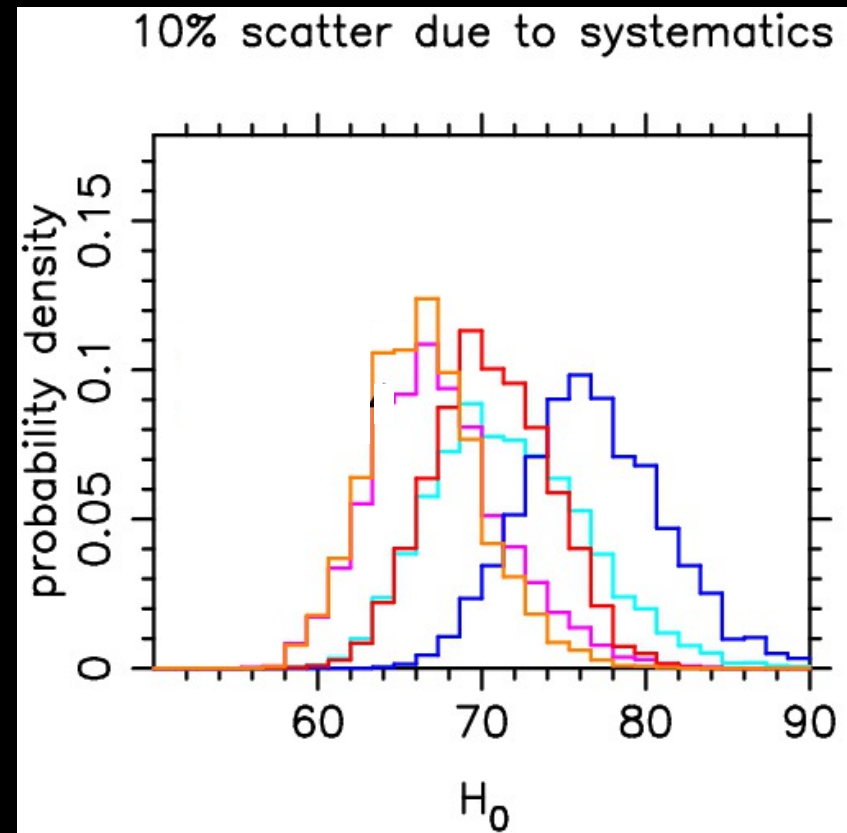
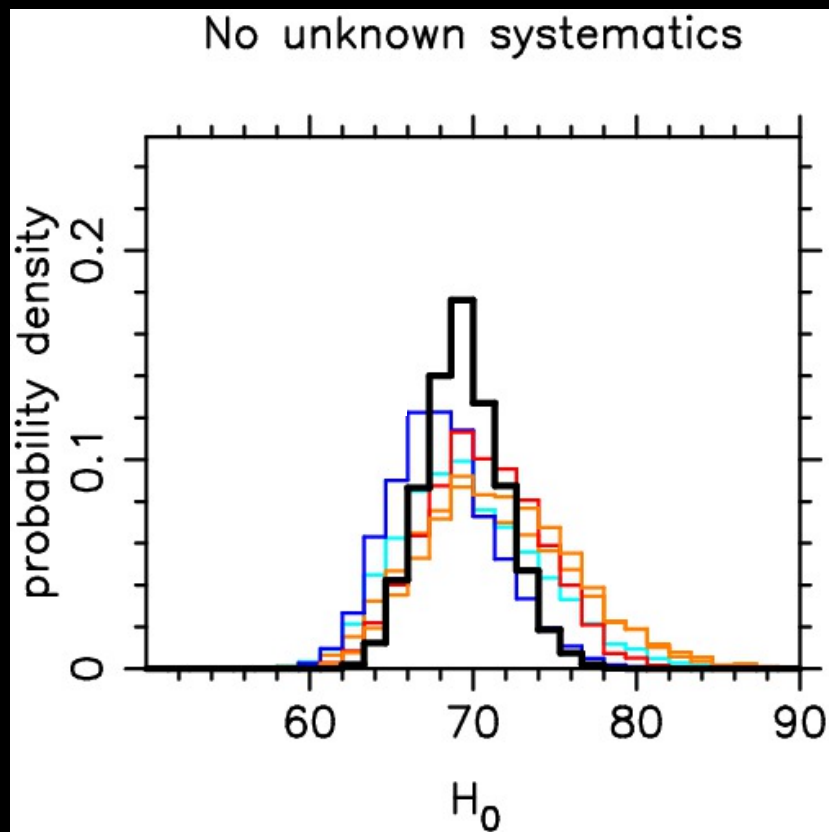
**Simona Vegetti** (MPA, Garching)

**Kenneth Wong** (ASIAA, Taipei)

# Goal: $H_0$ with $<4\%$ uncertainty

H0LiCOW: Pilot sample of 5 lenses with excellent ancillary data to

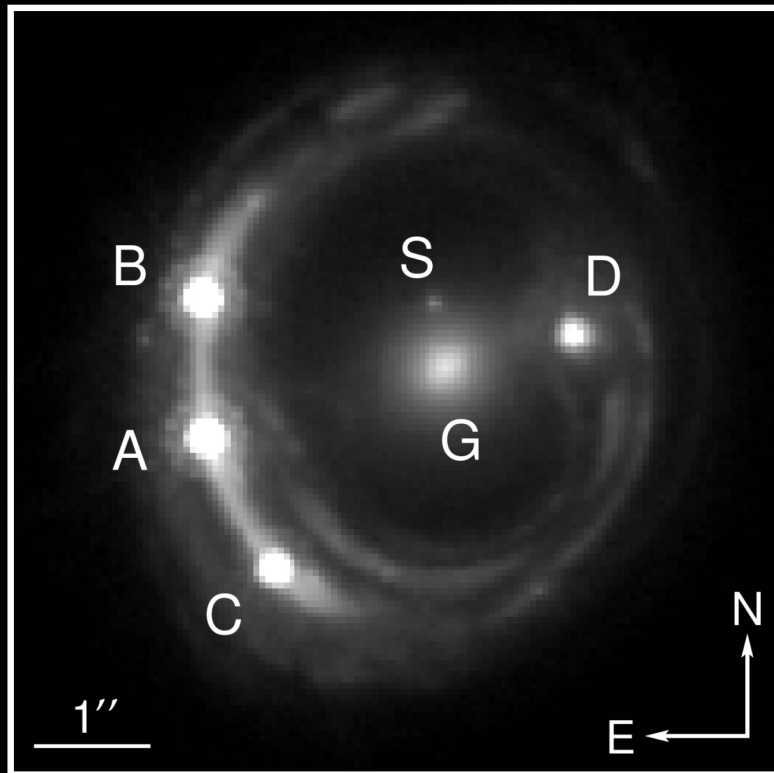
- gain in statistical precision
- test for systematics, particularly the unknown unknowns





# Cosmography with time-delay lenses

RXJ1131-1231



Time delay:

$$t = \frac{1}{c} D_{\Delta t} \phi_{\text{lens}}$$

Time-delay distance:

$$D_{\Delta t} \propto \frac{1}{H_0}$$

Obtain from lens mass model

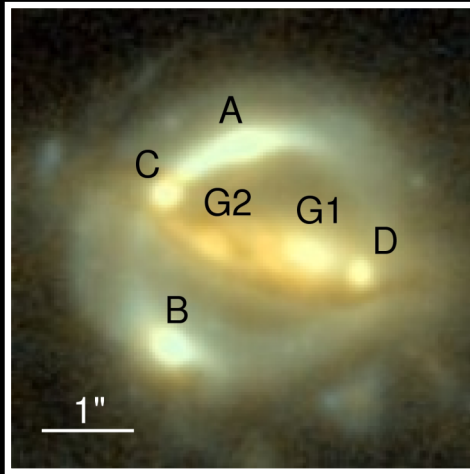
To measure  $D_{\Delta t}$ , need:

- (1) time delays
- (2) lens mass model
- (3) mass along line of sight

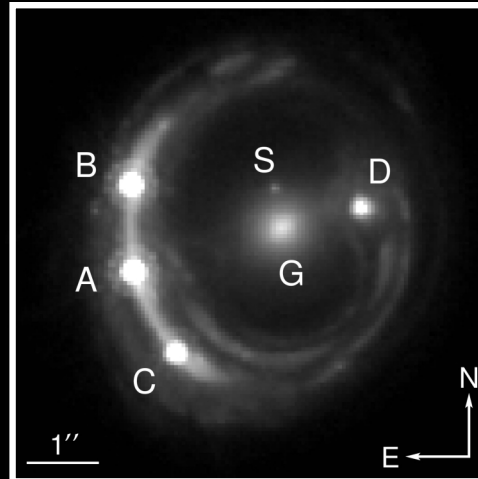


# H0LiCOW Sample

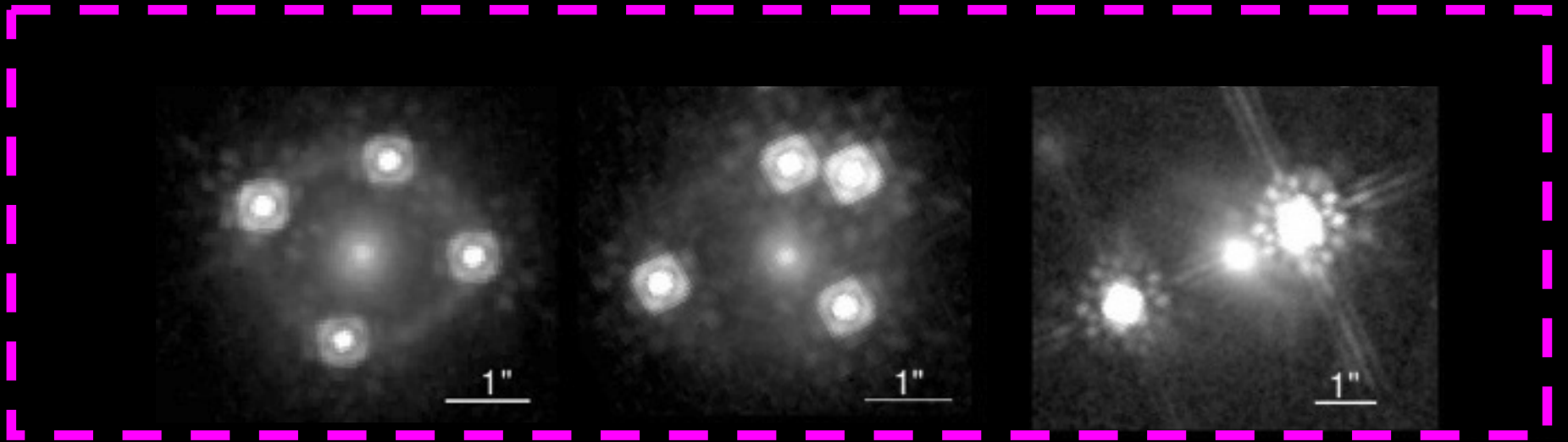
B1608+656



RXJ1131-1231



completed



S. Suyu, 2014/11/17

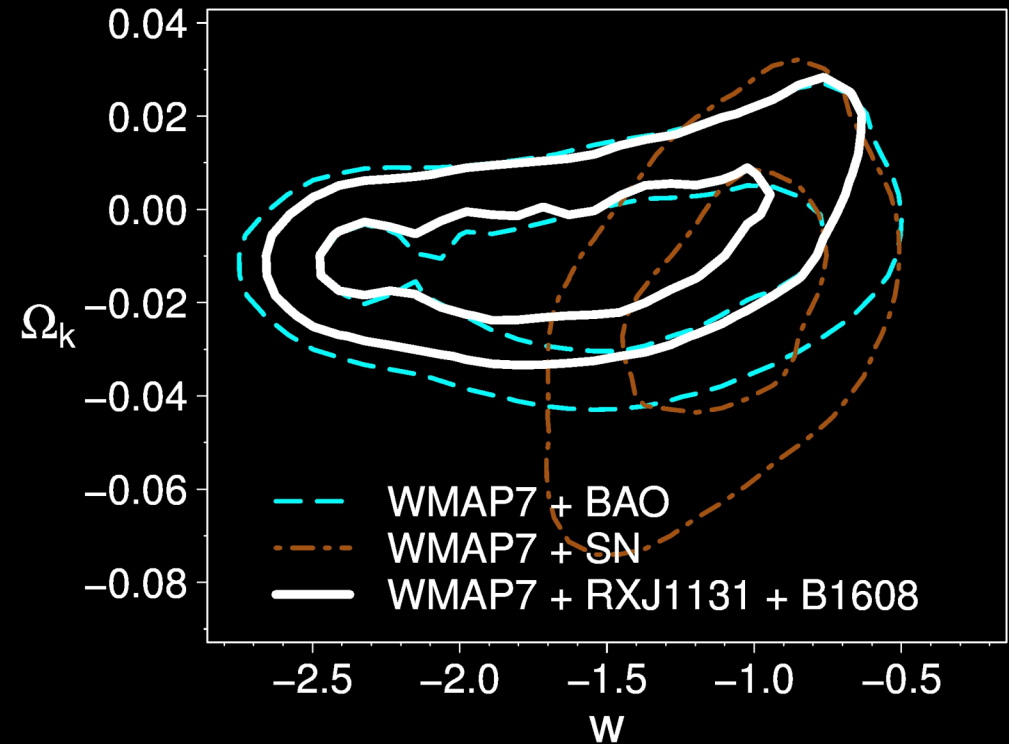
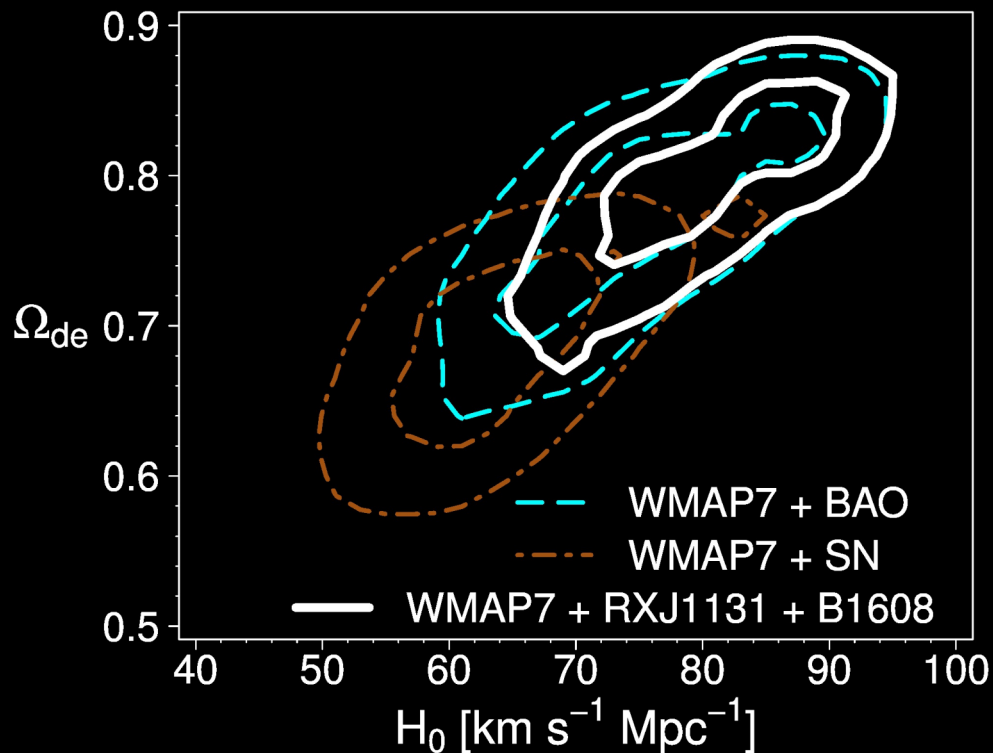
ongoing

5

# Cosmological Constraints from Two Time-delay Lenses

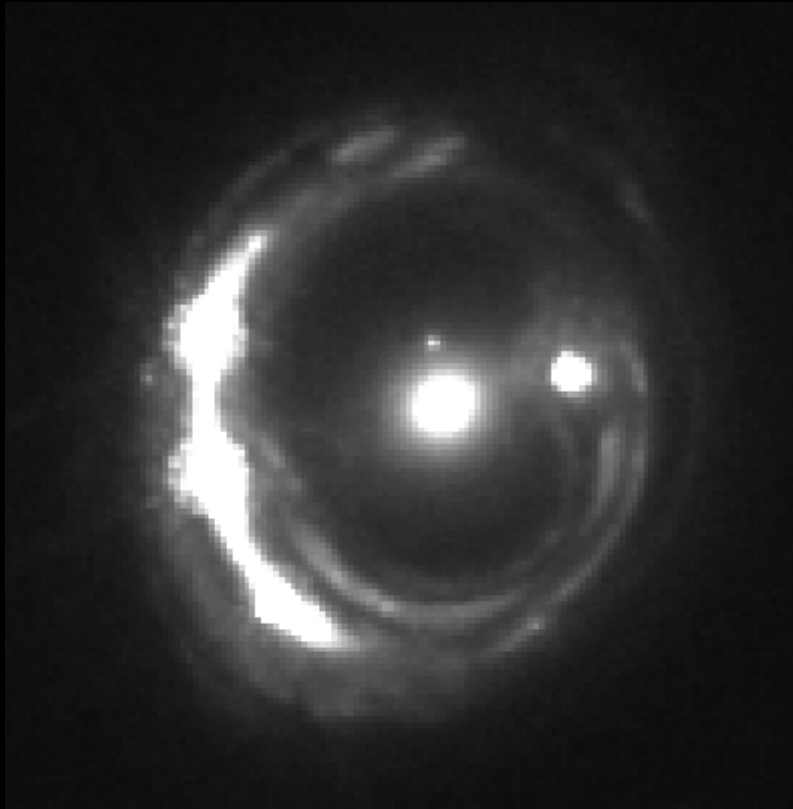
WMAP7 $\Lambda$ CDM prior

[Suyu et al. 2013]



- contour orientations are different: complementarity b/w probes
- contour sizes are similar: lensing is a competitive probe

# RXJ1131-1231



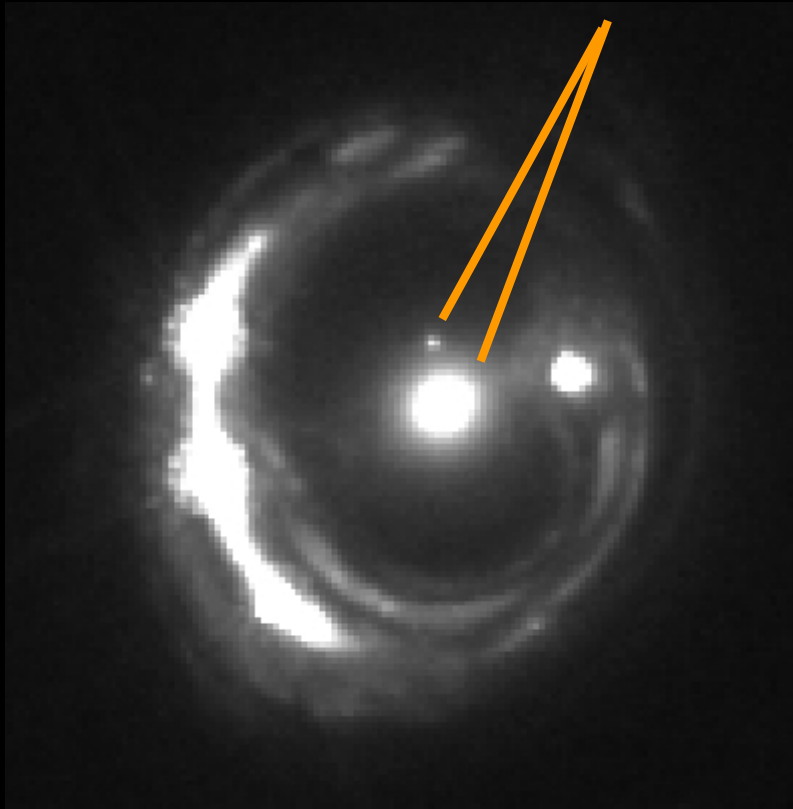
## Data:

- COSMOGRAIL time delays with 1.5% precision [Tewes et al. 2013]
- HST F814W imaging
- Lens velocity dispersion of  $323 \pm 20$  km/s from spectroscopy at Keck
- Relative galaxy number count of 1.4 within  $45''$  from the lens [Fassnacht et al. 2011]



# Lens Mass Model

mass distribution  
of lens



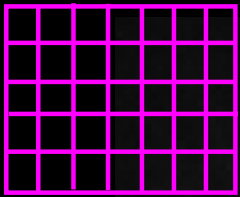
- Three lens mass models:
- (1) power-law profile
  - (2) cored power-law profile
  - (3) composite
    - baryons = light  $\times$  M/L
    - dark matter = NFW halo

[Suyu et al. 2013, 2014]

# Lens Mass Model

light distribution  
of extended source

mass distribution  
of lens

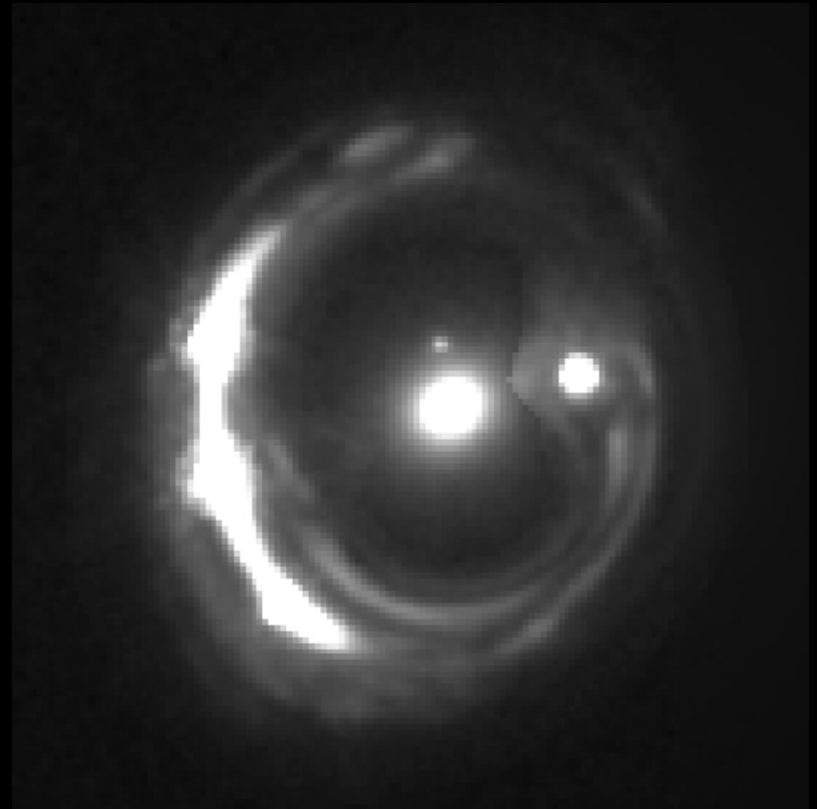
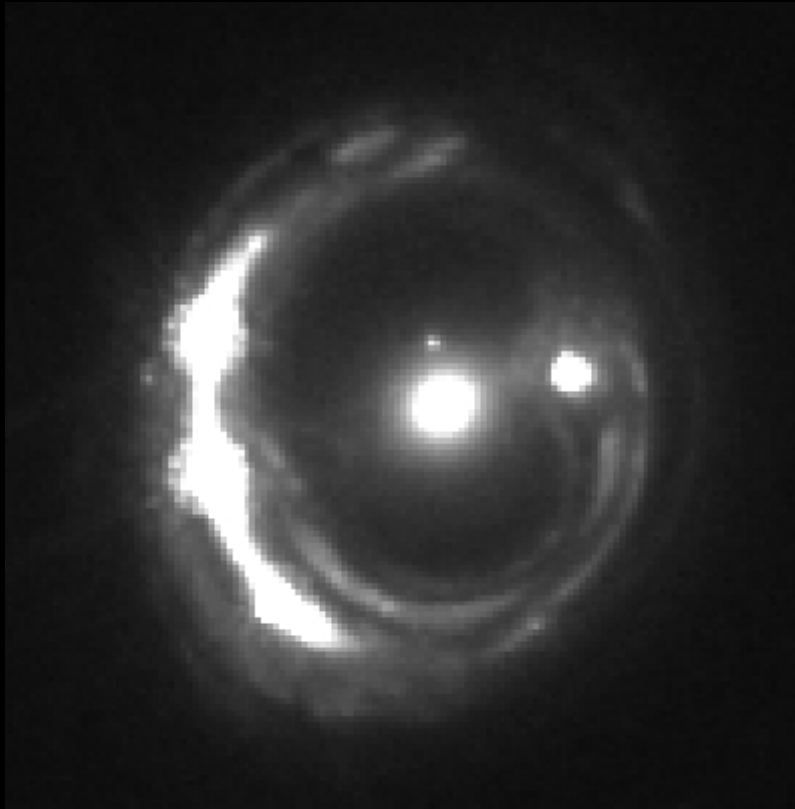


light of  
lensed  
AGN

+  
time delays

light of  
lens  
(Sersic)

# Lens Mass Model

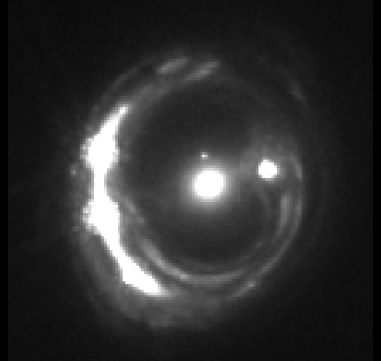


S. Suyu, 2014/11/17

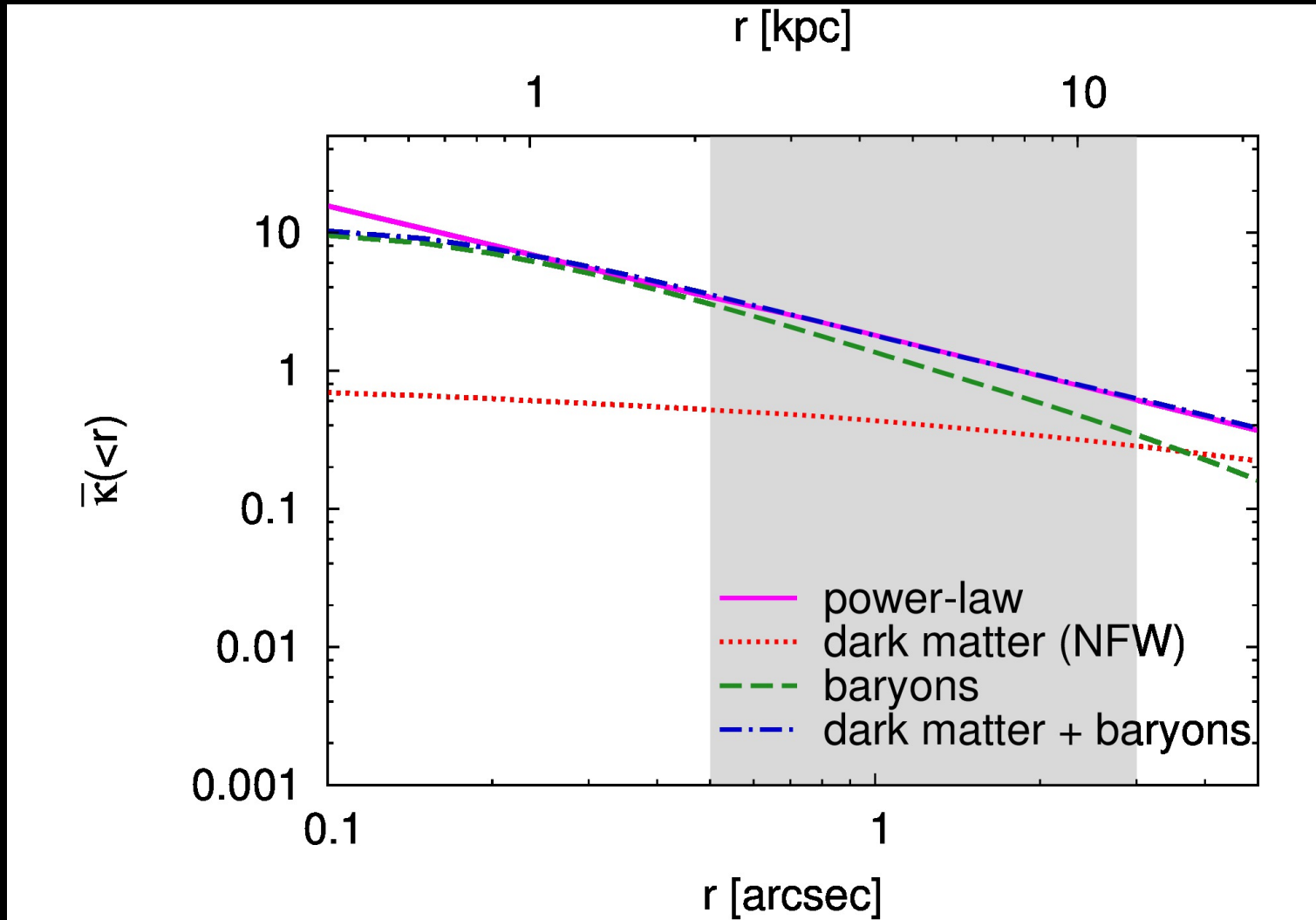
[Suyu et al. 2013] 10



# Lens mass profiles

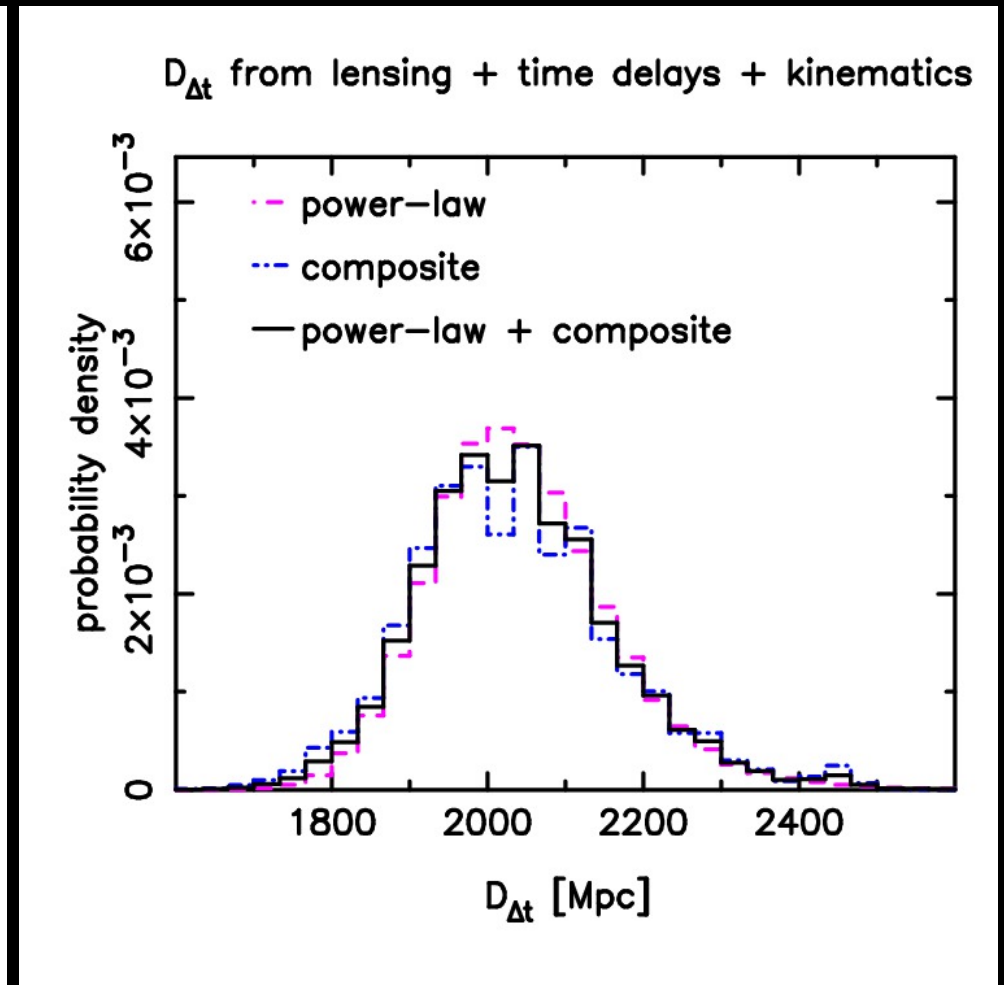
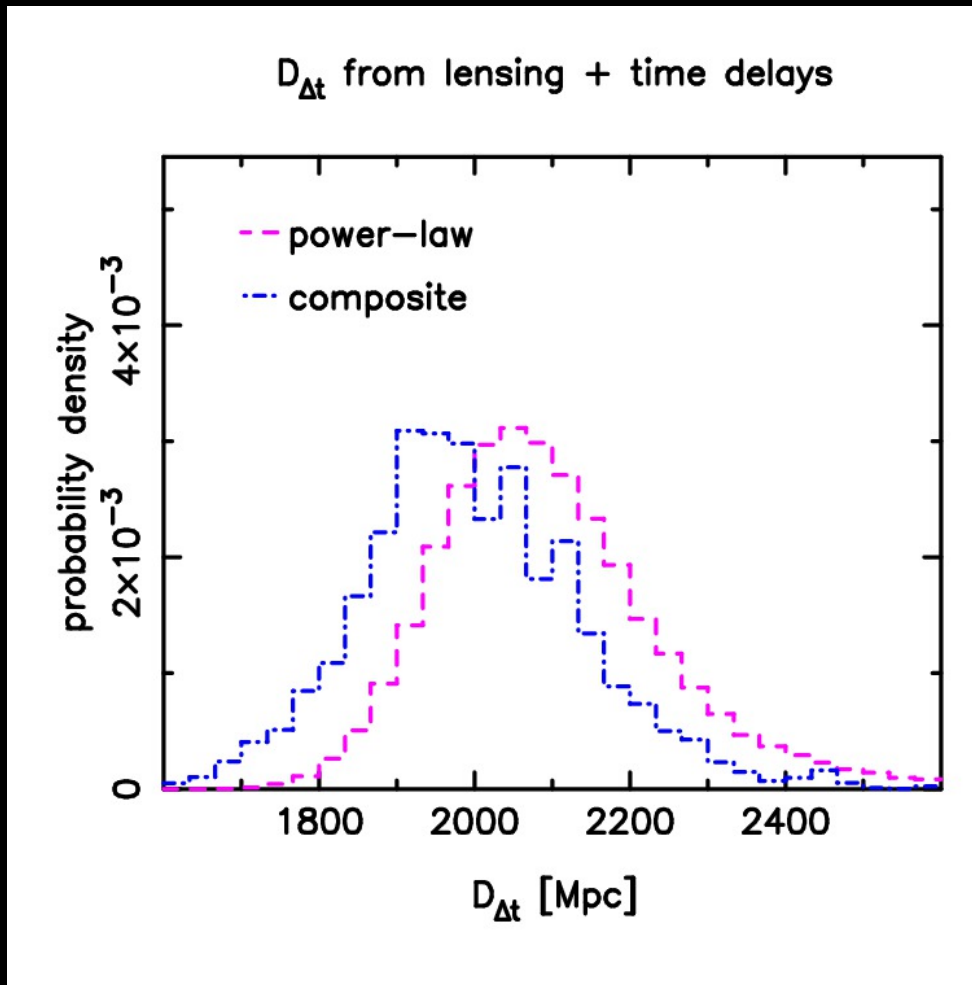


Comparison of power-law and composite model:



[Suyu et al. 2014]

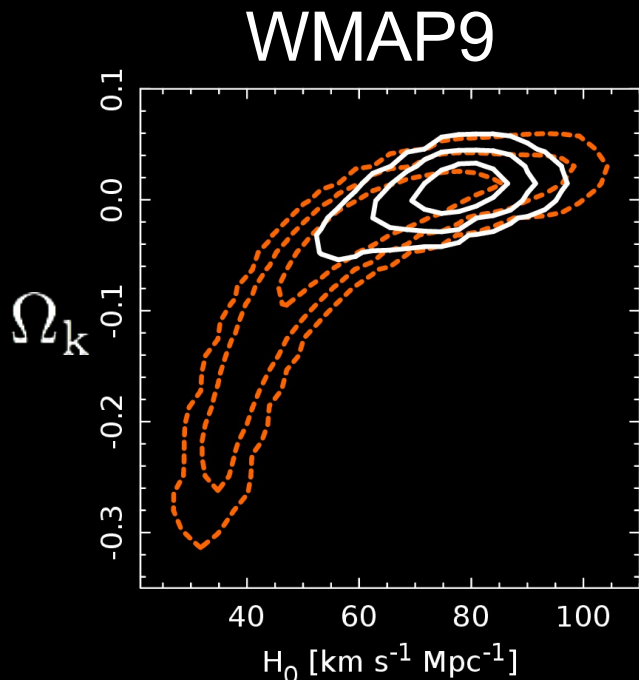
# Power of lensing + kinematics



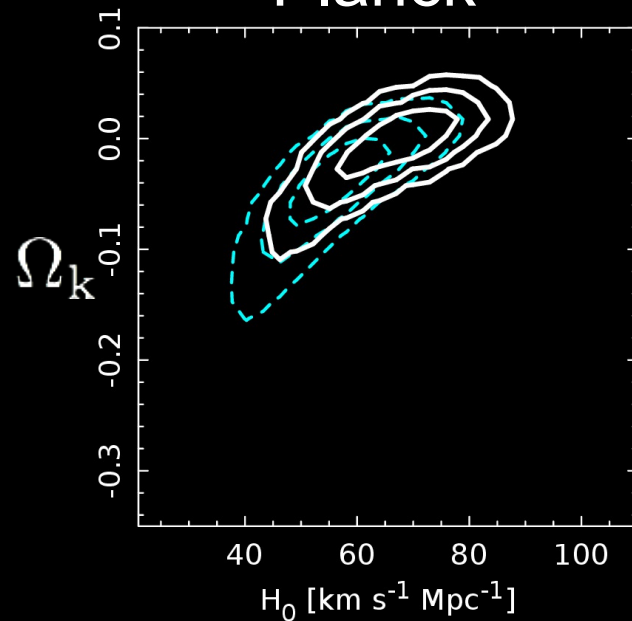
Velocity dispersion of the lens provide substantial information to break lens model degeneracies [Suyu et al. 2014]

# Update: WMAP9 and Planck

open  
 $\Lambda$ CDM

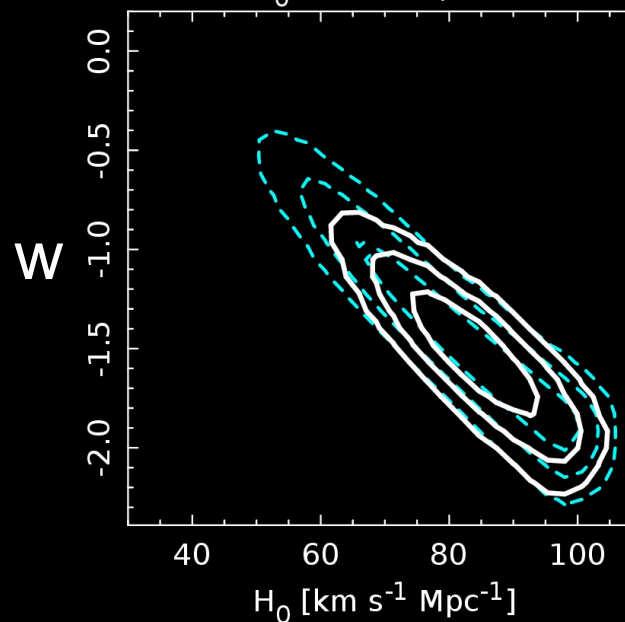
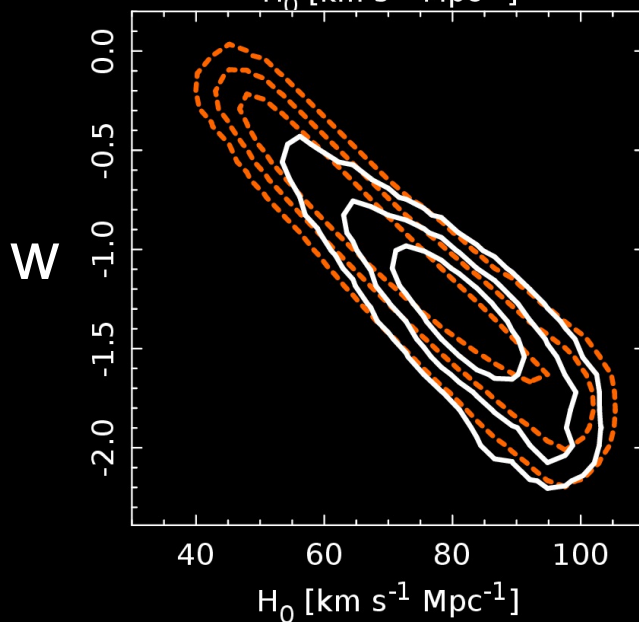


Planck



RXJ1131  
+  
CMB

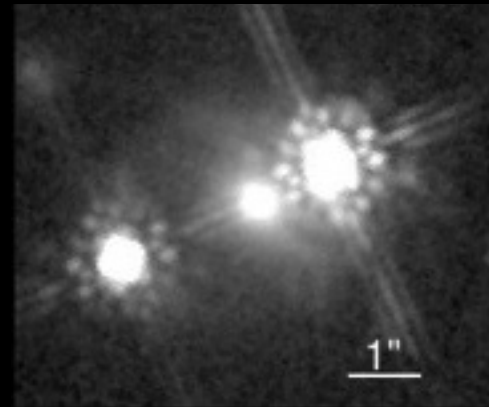
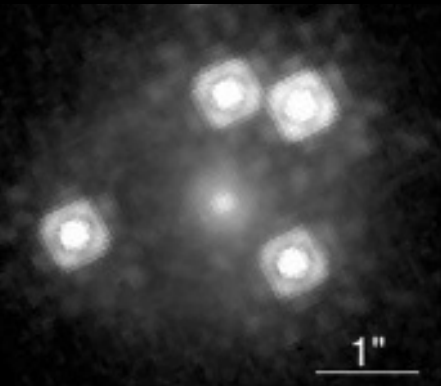
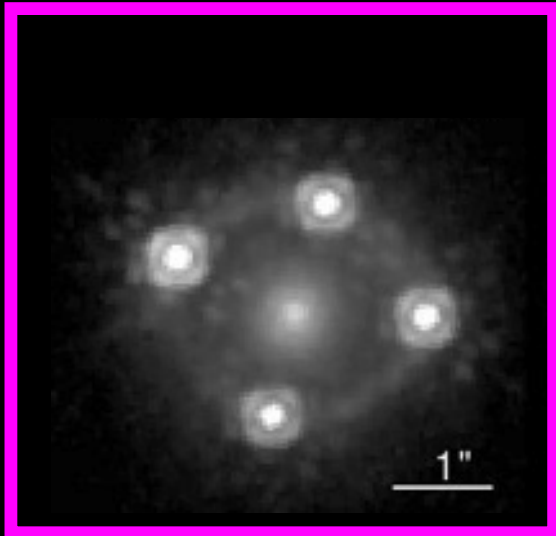
flat  
wCDM



[Suyu et al. 2014]



# Next 3 H0LiCOW Lenses

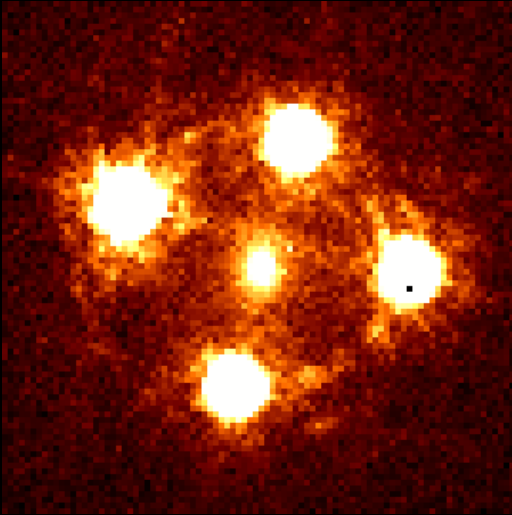


## *Follow ups on next 3 lenses:*

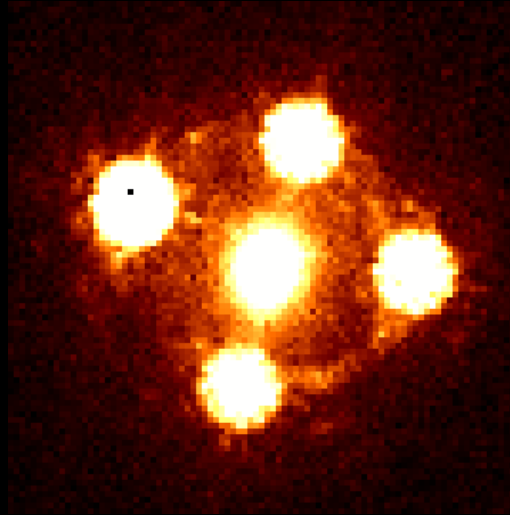
- delays with a few percent uncertainty from COSMOGRAIL
- *HST* deep imaging for modeling
- Lens galaxy spectroscopy pending
- Imaging and spectroscopy of the fields pending

# HE0435: modeling

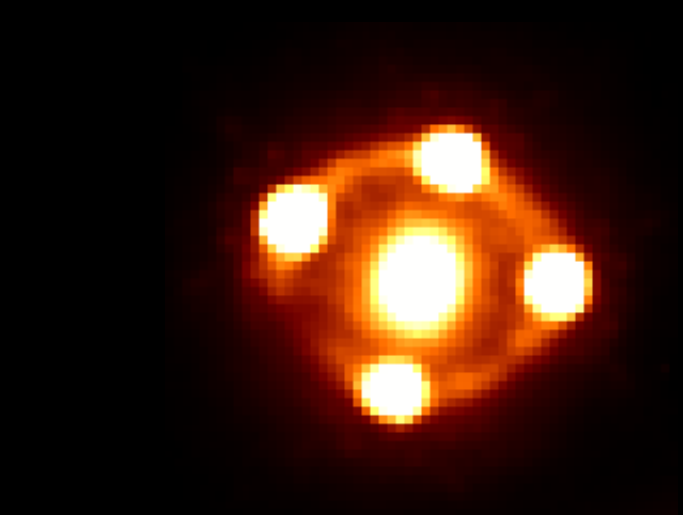
ACS F555W



ACS F814W



WFC3 F160W

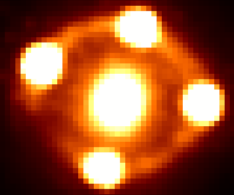


- HST imaging in 3 bands
- Modeling:
  - elliptical power-law profile
  - and external shear

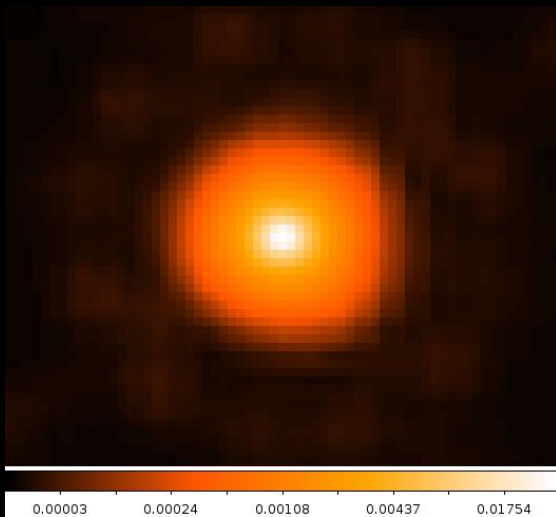


Ken Wong

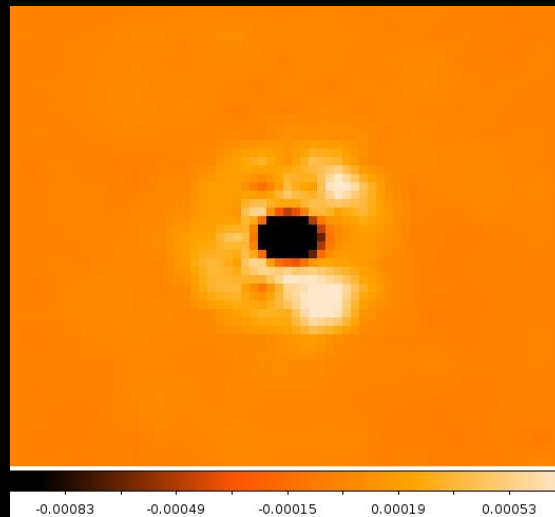
# HE0435: PSF (F160W) [Preliminary]



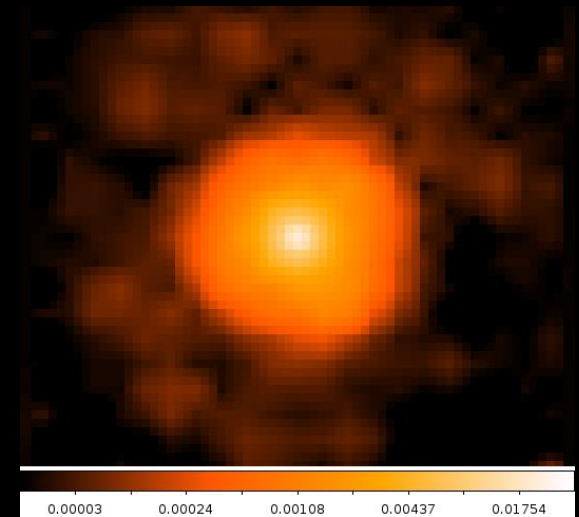
Initial PSF  
(from stars)



PSF correction  
(fractional)



Corrected PSF

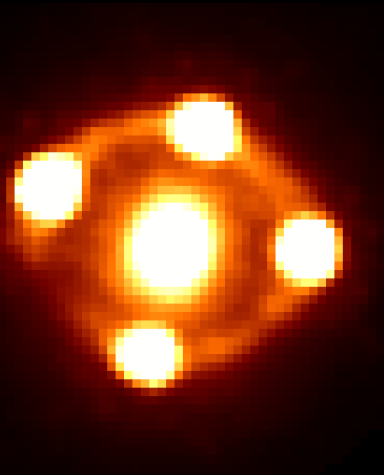


[figure material courtesy of Vivien Bonvin and Ken Wong]

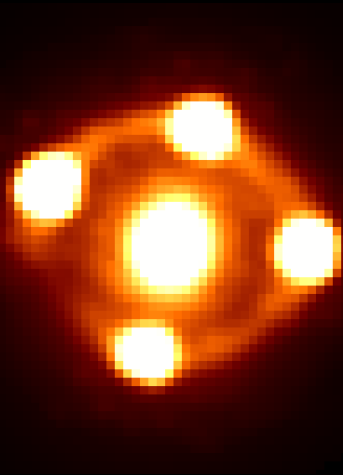
- AGNs are very bright, so need accurate PSF model to extract the ring (of AGN host galaxy) for lens mass modeling
- Initial PSF built from stars in field (by Vivien Bonvin & Frederic Courbin)
- Correct the PSF on a grid of pixels, simultaneously with the ring modeling

# HE0435: image reconstruction (F160W) [Preliminary]

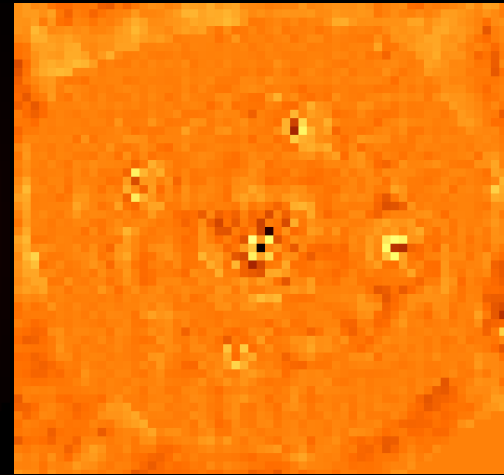
Data



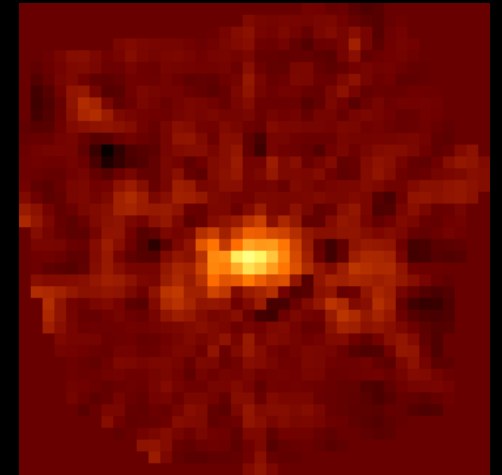
Model



Residuals



Source recon.



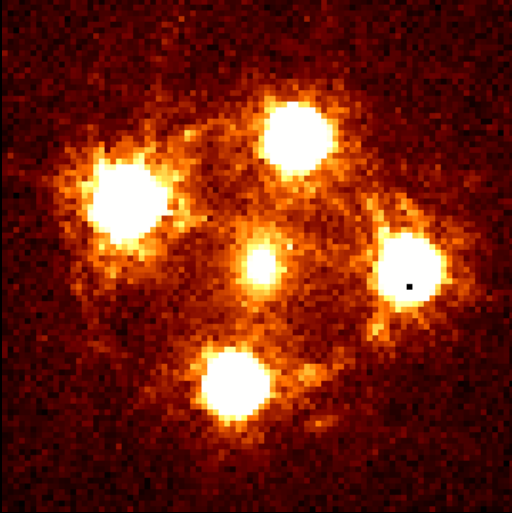
[figure material courtesy of Ken Wong]

Based on:

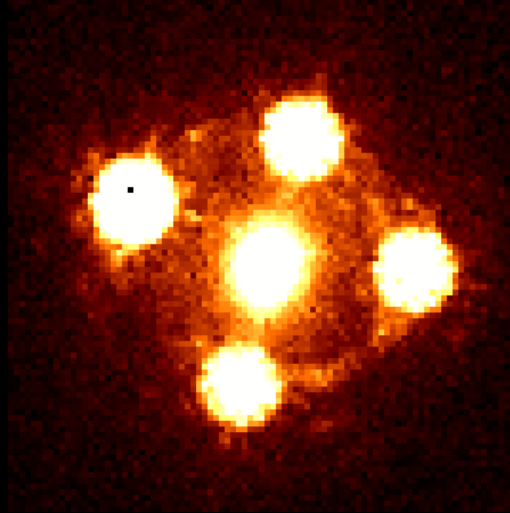
- power-law lens with external shear
- pixelated PSF correction

# HE0435: next steps

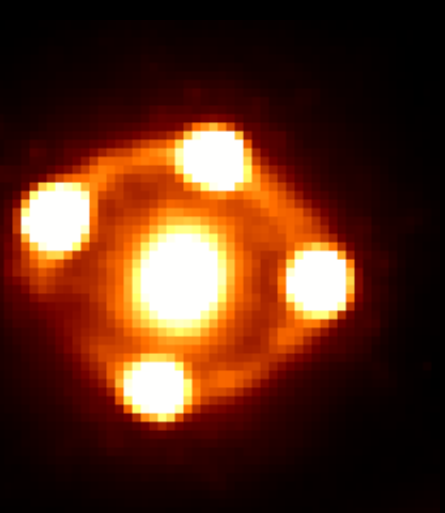
ACS F555W



ACS F814W



WFC3 F160W



- ACS bands modeling (ongoing)
- simultaneous modeling of all 3 bands
- use flexible mass models incorporating nearby perturbers
- characterize mass along the line of sight

# Cosmography and Beyond

- Exquisite imaging and spectroscopic data on H0LiCOW lenses enable a wide range of scientific studies
- Current projects based on H0LiCOW lenses:
  - Cosmography
  - Lens environment studies (number counts, weak lensing, galaxy group identification)
  - Interplay between baryons and dark matter
  - Stellar Initial Mass Function
  - Host galaxies of Active Galactic Nuclei
  - Data Challenge (simulations and real data)



# Summary

- H0LiCOW: pilot sample of 5 time-delay lenses with exquisite data to measure  $H_0$  to better than 4% (stat. + sys.)
- 2 lenses analyzed with cosmological results competitive and complementary to other probes
- spatially extended Einstein ring and the multiple time delays provide constraints on the local profile of lens mass distribution in RXJ1131-1231
- velocity dispersion further breaks lens model degeneracies
- analysis of HE0435 ongoing
- H0LiCOW lenses are also great for studying dark matter distributions, stellar IMF and supermassive black holes