

Gravitationally lensed supernovae

Masamune Oguri

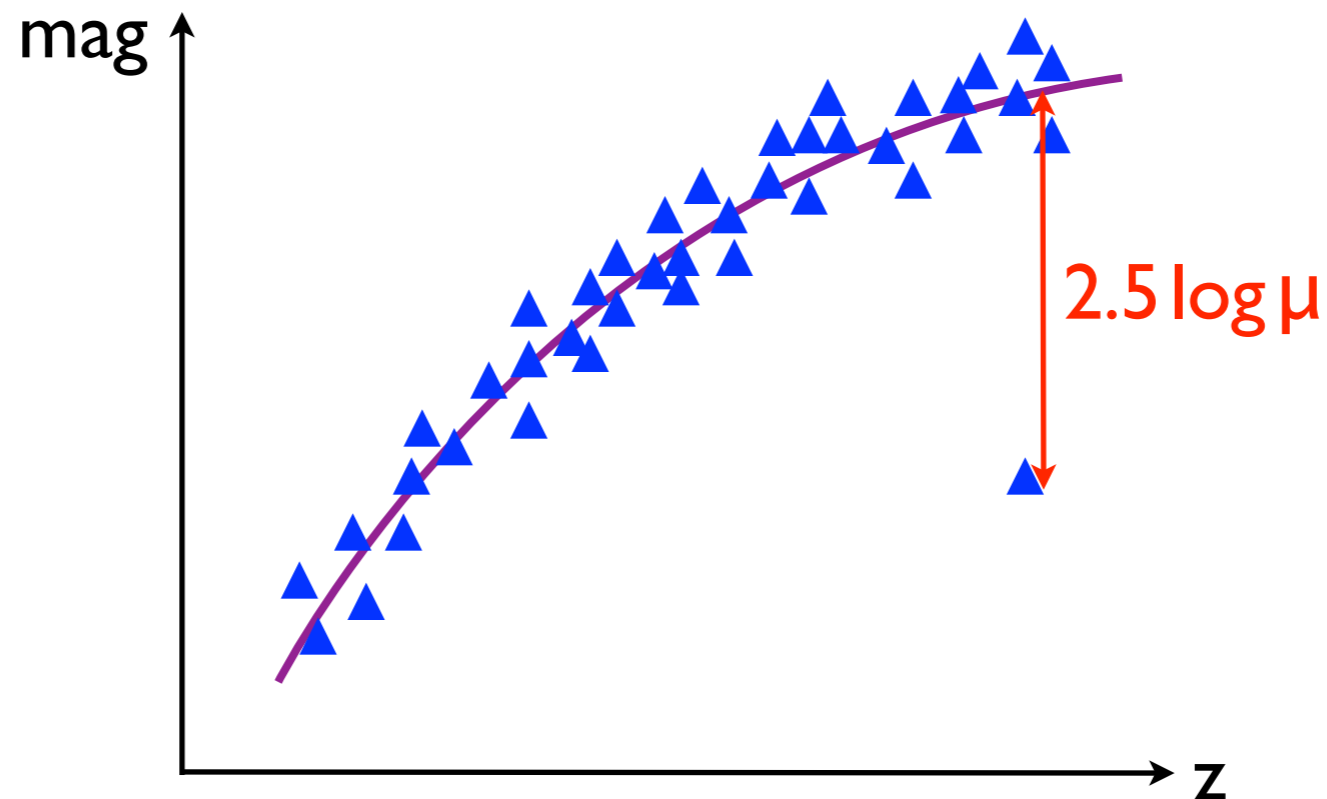
University of Tokyo/Kavli IPMU



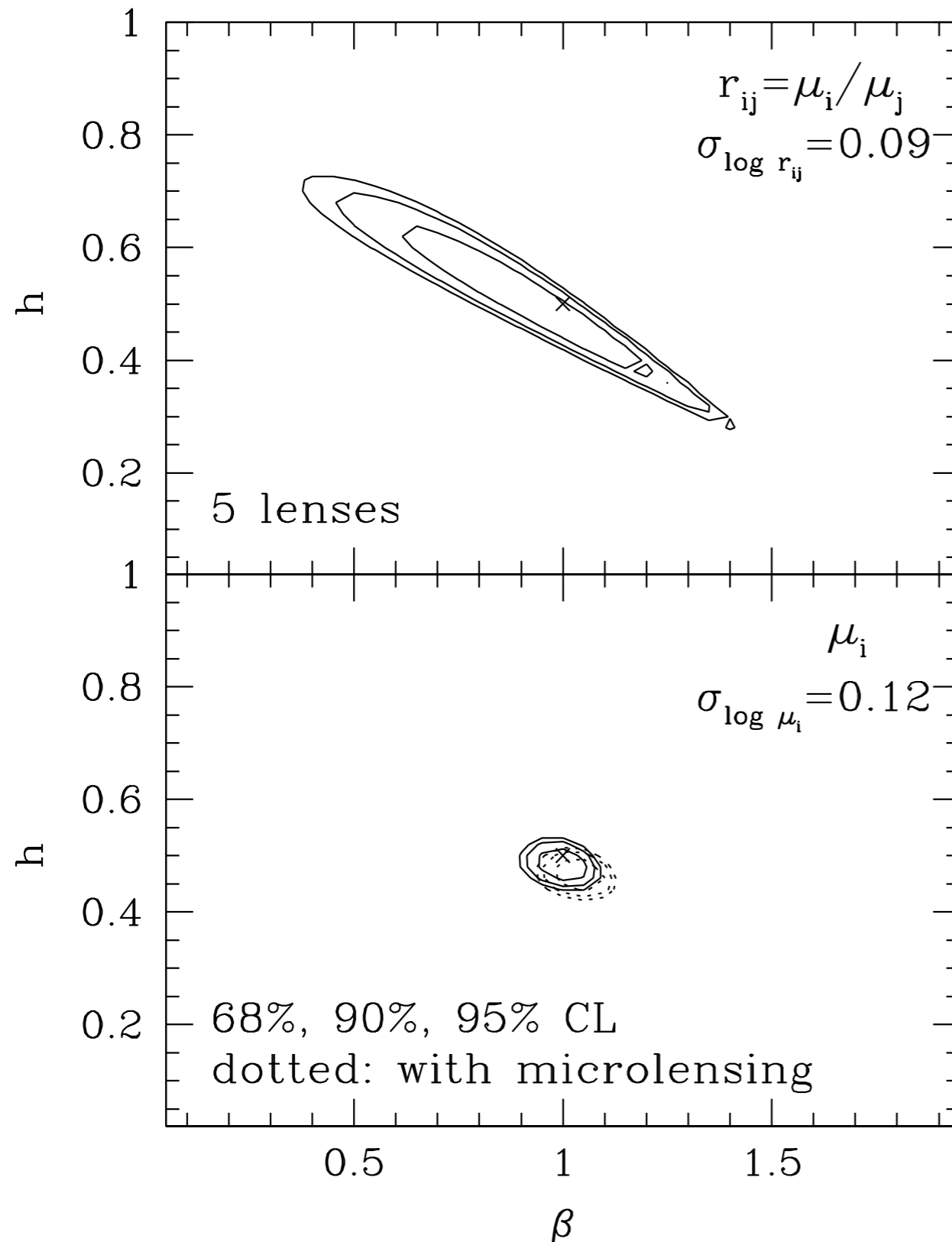
with **Robert Quimby**, Marcus Werner,
Anupreeta More, Surhud More,
Takashi Moriya, Masayuki Tanaka,
Gaston Folatelli, Melina Bersten,
Keiichi Maeda, Ken'ichi Nomoto

Why is lensed SN interesting? (I)

- **standard candle**
direct measurement of the magnification factor for lensed type Ia supernovae, breaking various (e.g., mass-sheet) degeneracies



Breaking the H_0 -slope degeneracy

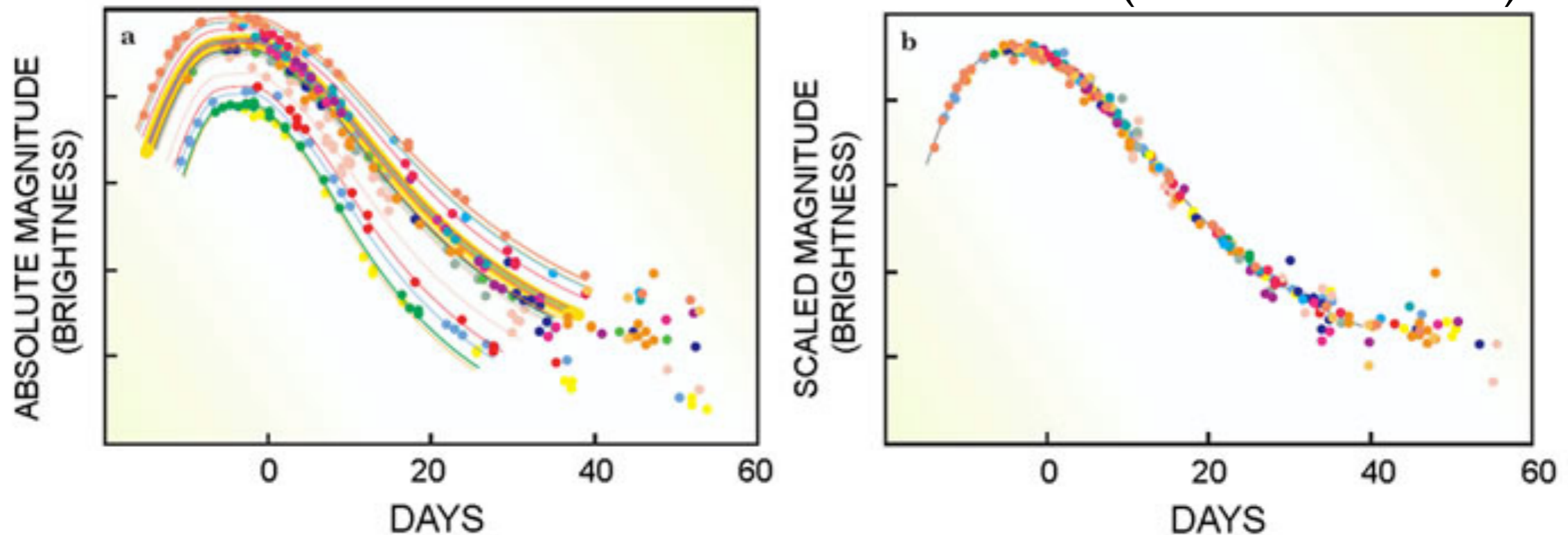


- the use of total magnifications μ breaks degeneracy btw H_0 and β [$\rho(r) \propto r^{\beta-3}$]

Why is lensed SN interesting? (II)

- **known light curves**
we have template light-curves of SNe
→ accurate and robust time delay measurements
(but microlensing can be important; Dobler & Keeton 2006)

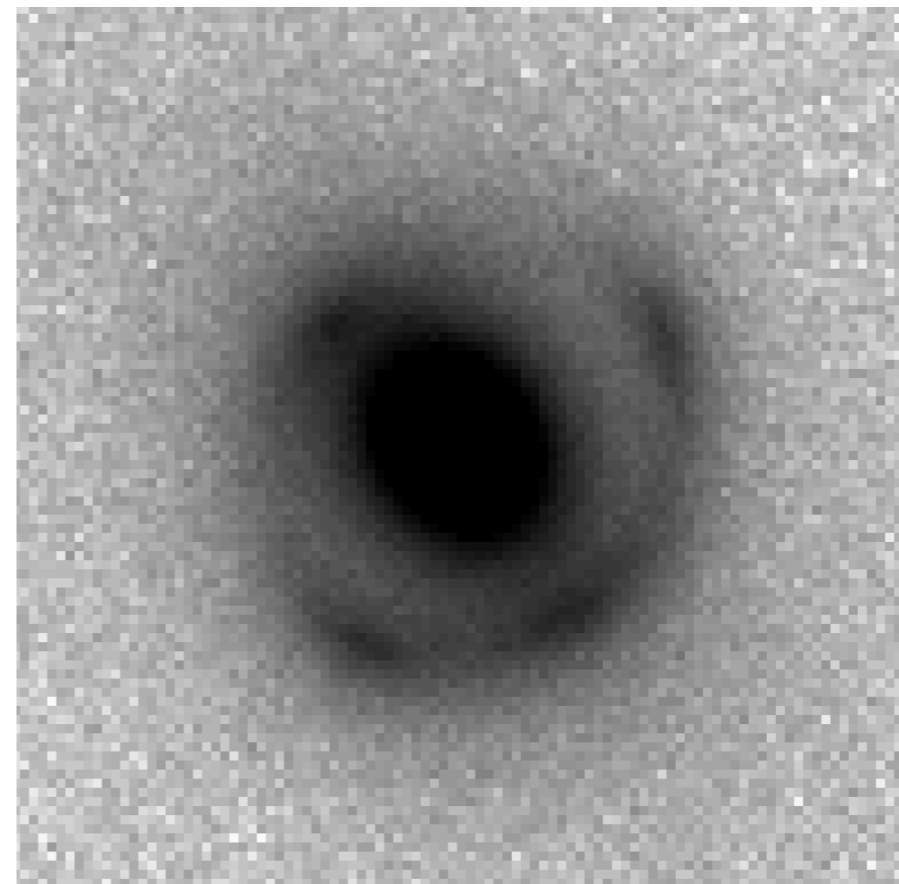
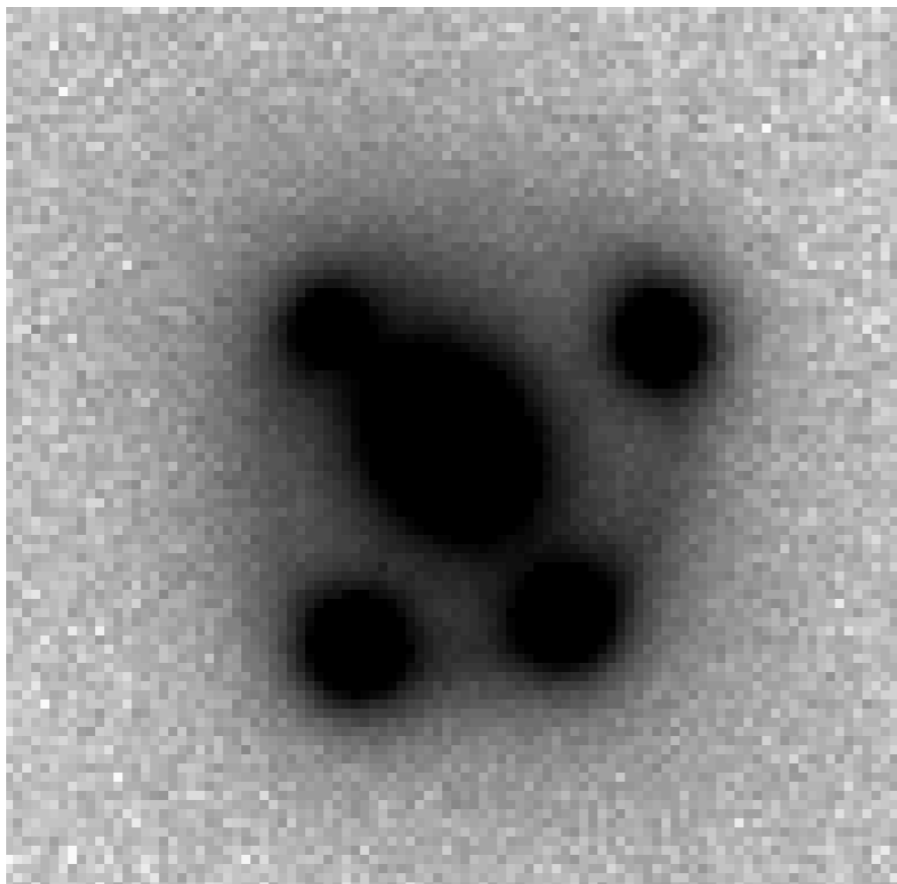
(from LBL website)



Why is lensed SN interesting? (III)

- **better use of host galaxy**
better measurement of detailed morphology of lensed host galaxy after SN fades away
→ better constraints of the lens potential

w/ SN
images



w/o SN
images

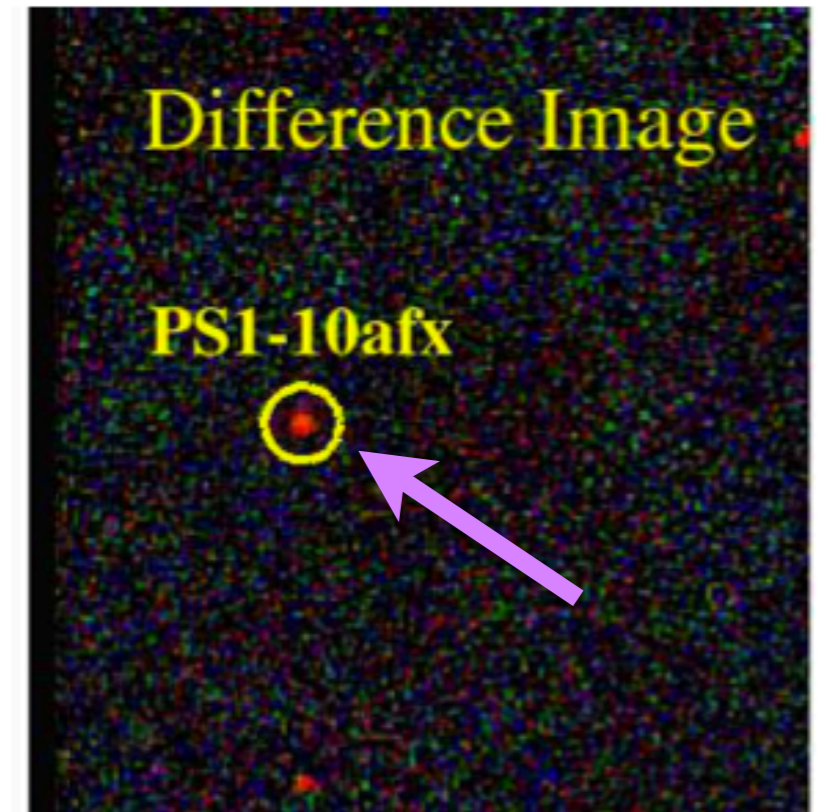
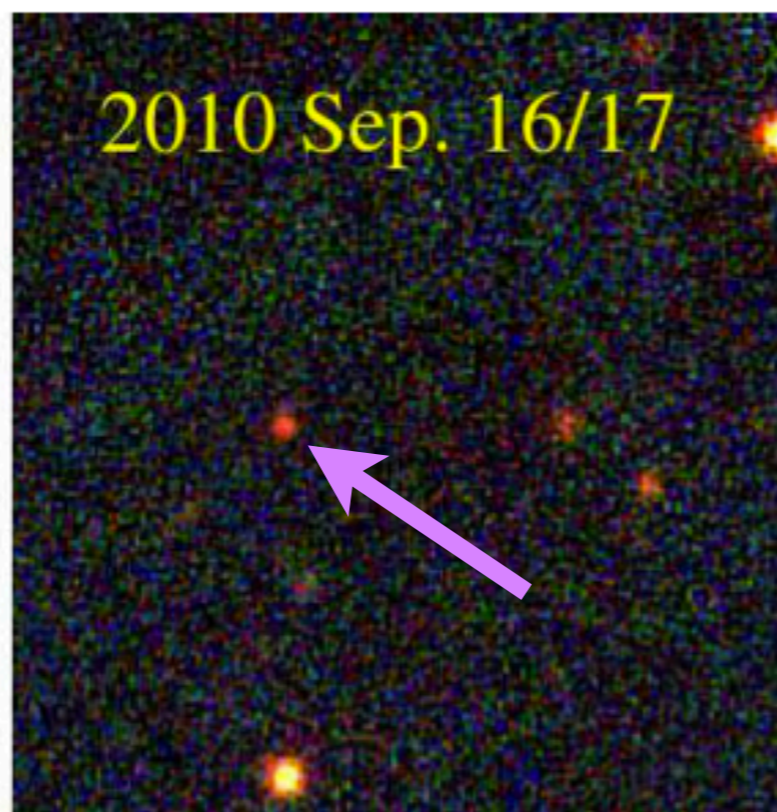
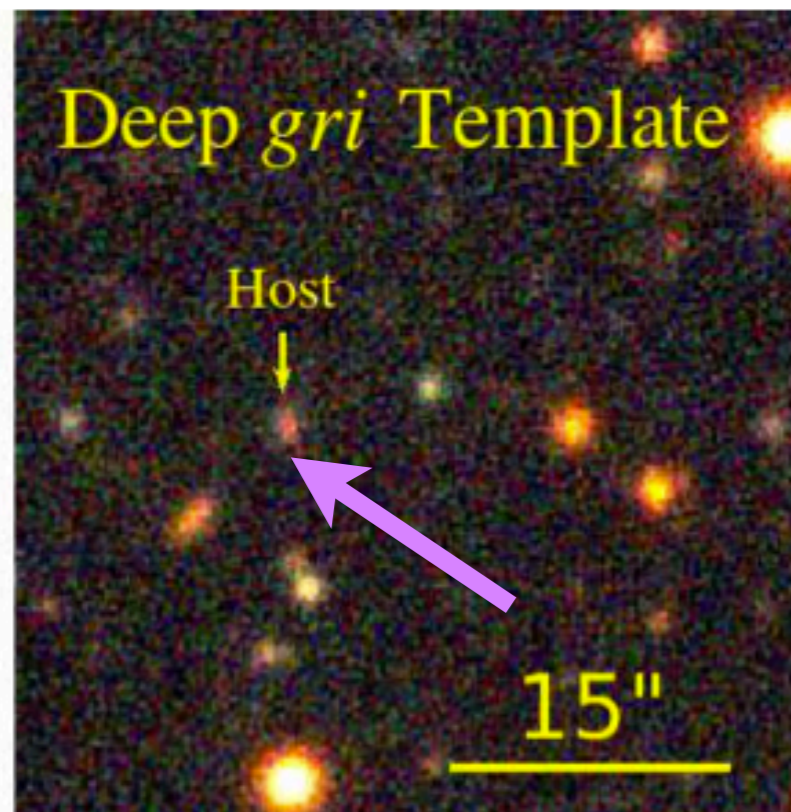
simulated by *glafic*

Expected number of lensed SNe

- rare, wide-field time-domain surveys are needed to find them
- $O(0.1)$ lensed SNe expected in [Pan-STARRS I \(PSI\)](#), can be the first survey to find lensed SNe
- $O(100)$ lensed SNe expected in [Large Synoptic Survey Telescope \(LSST\)](#)

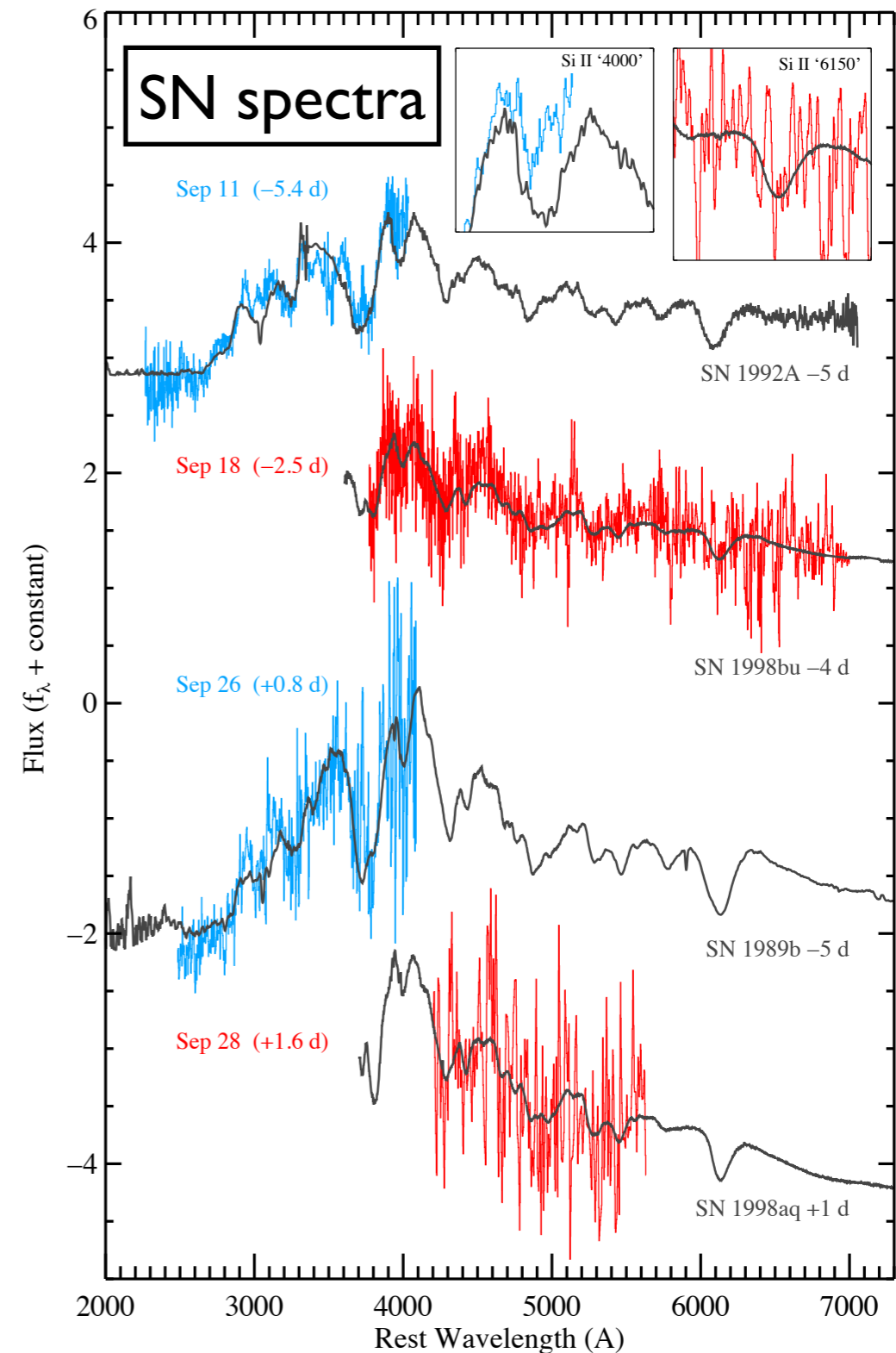
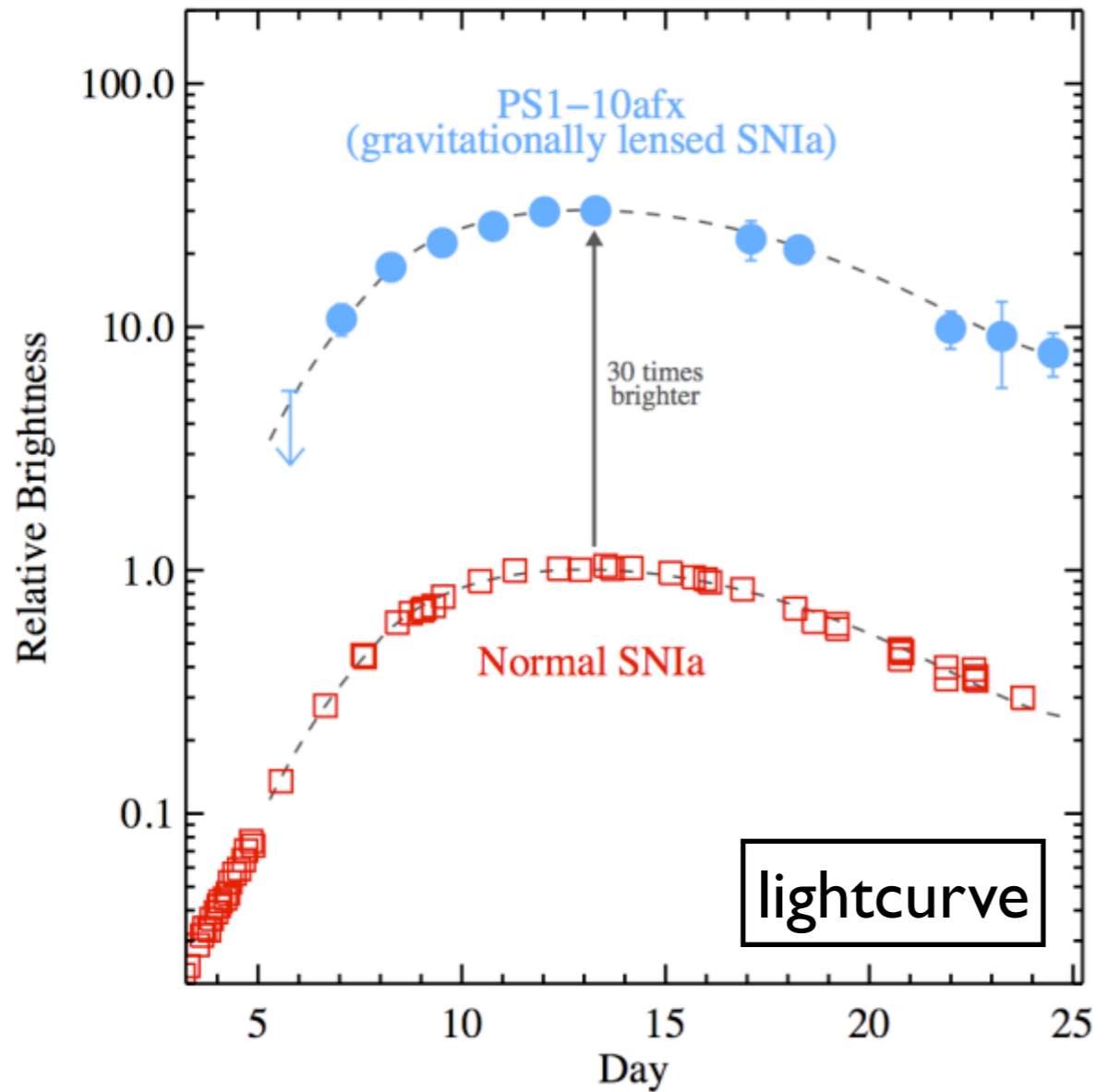
Discovery of PS1-10afx

- unusually red transient at $z=1.388$ found on Aug 31, 2010 in the PS1 Medium Deep Survey (MDS)
- PS1 team concluded that it is a new class of super-luminous supernova (SLSN), but no physical model of SLSN can explain this event



Chornock et al. (2013)

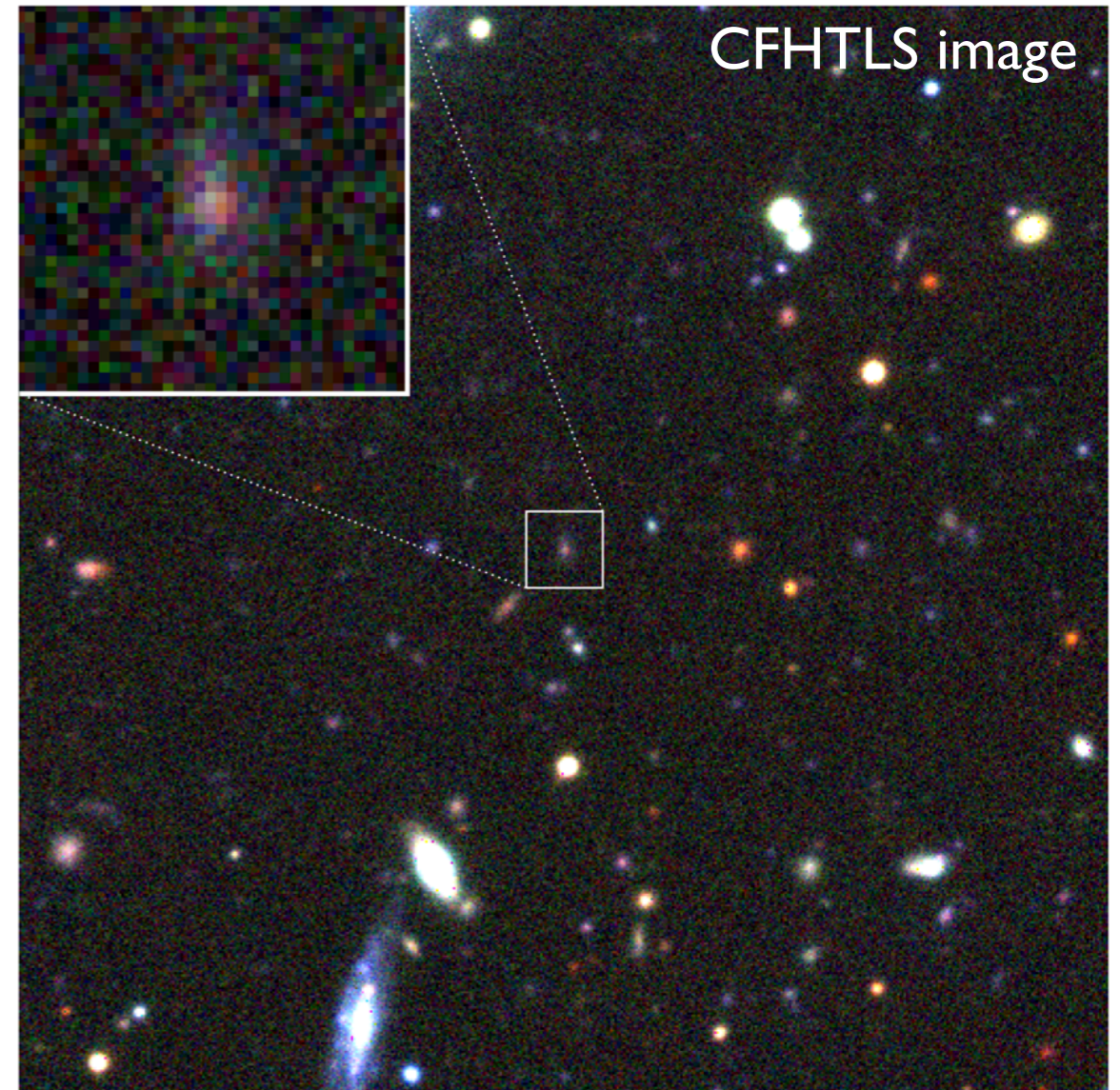
Type Ia interpretation of PS1-10afx



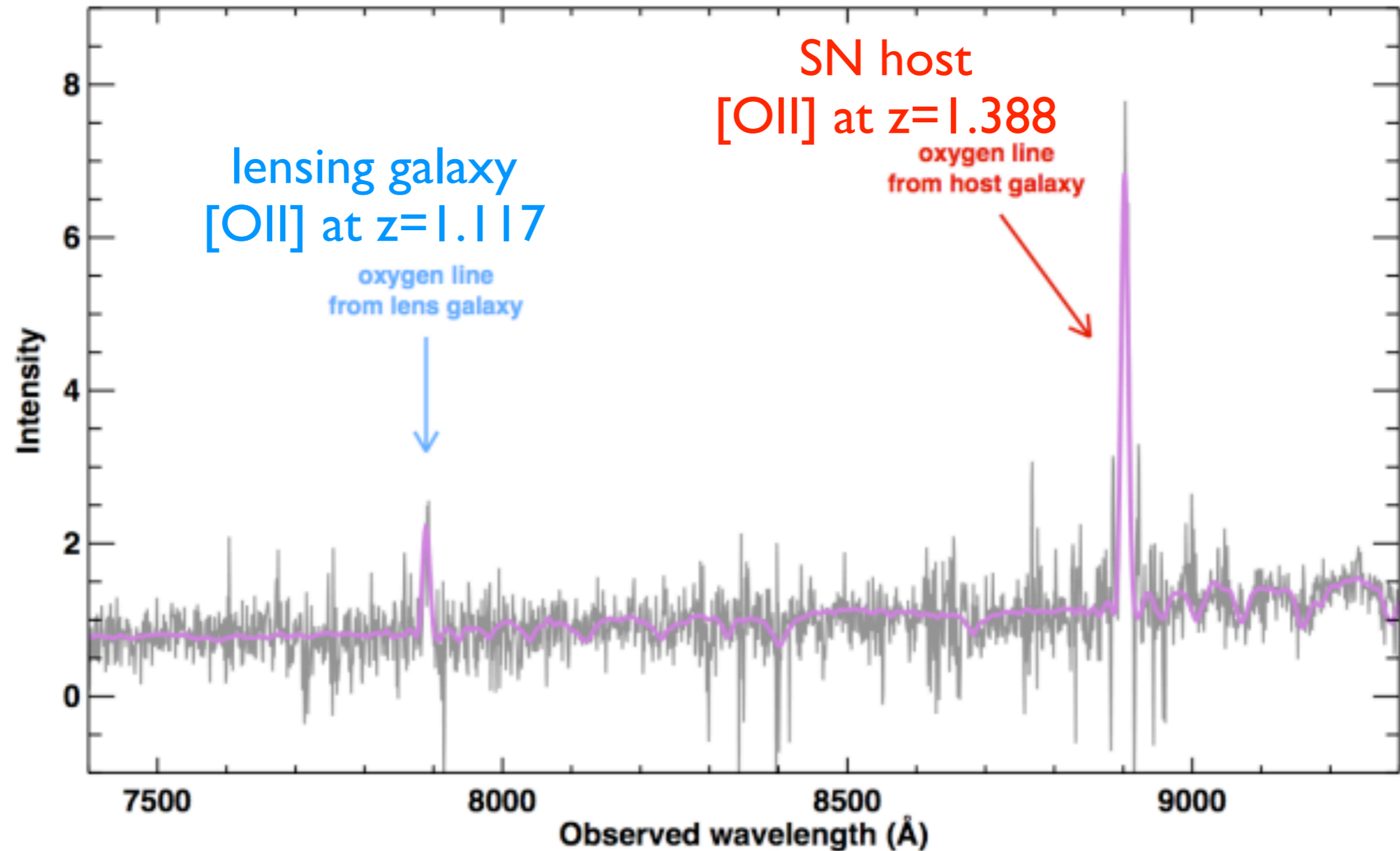
lightcurve and spectra
consistent with 30x
magnified type Ia SN

Where is the lens?

- image before SN exploded shows only one galaxy
- we speculated that this is in fact a superposition of two galaxies, SN host and foreground lens
- 6.5 hr **Keck spectroscopy** on Sep 7, 2013 to find out true nature of this object

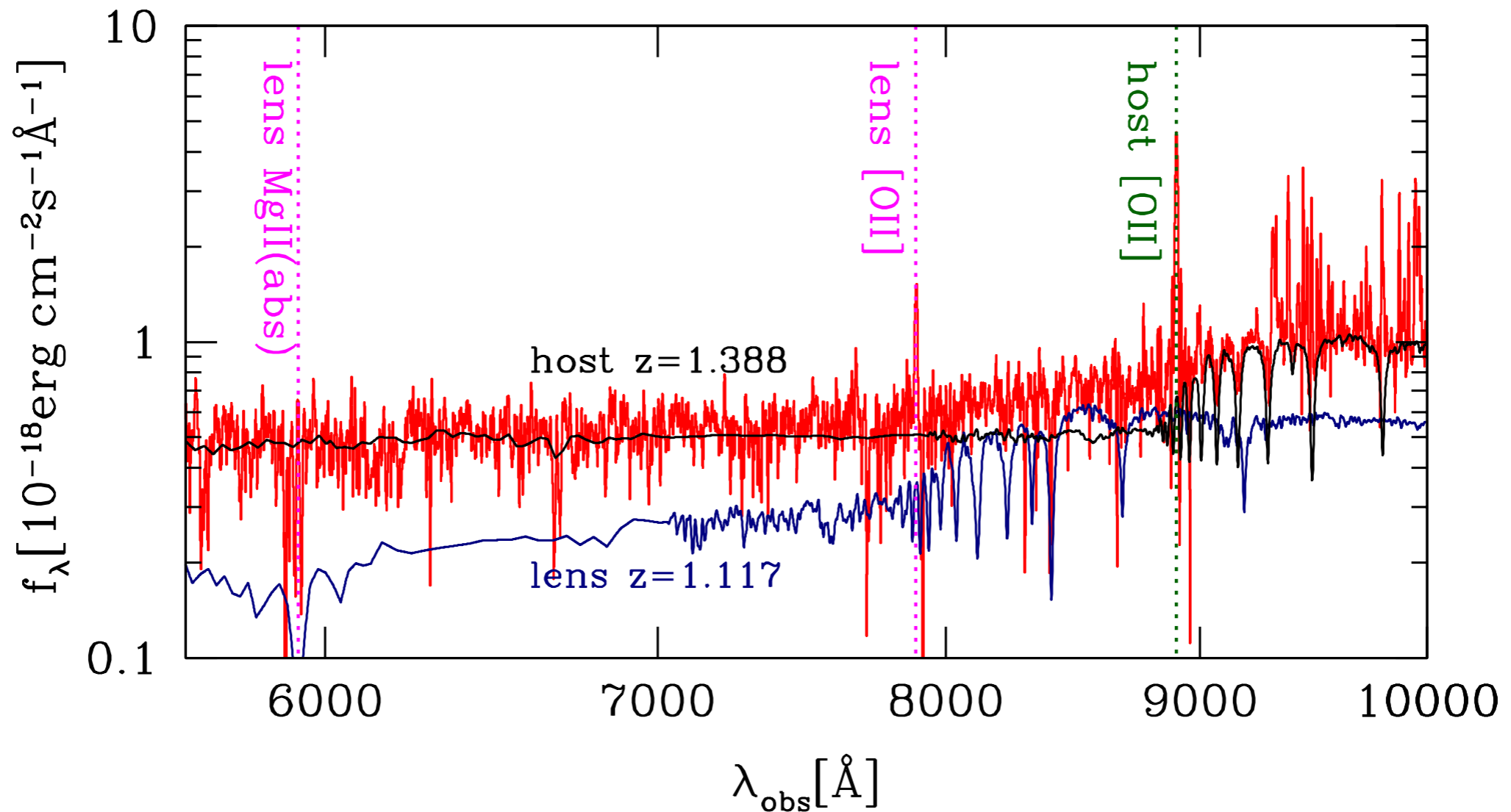


Detection of the lensing galaxy



- foreground lensing galaxy at $z=1.117$ discovered !

Property of the lensing galaxy

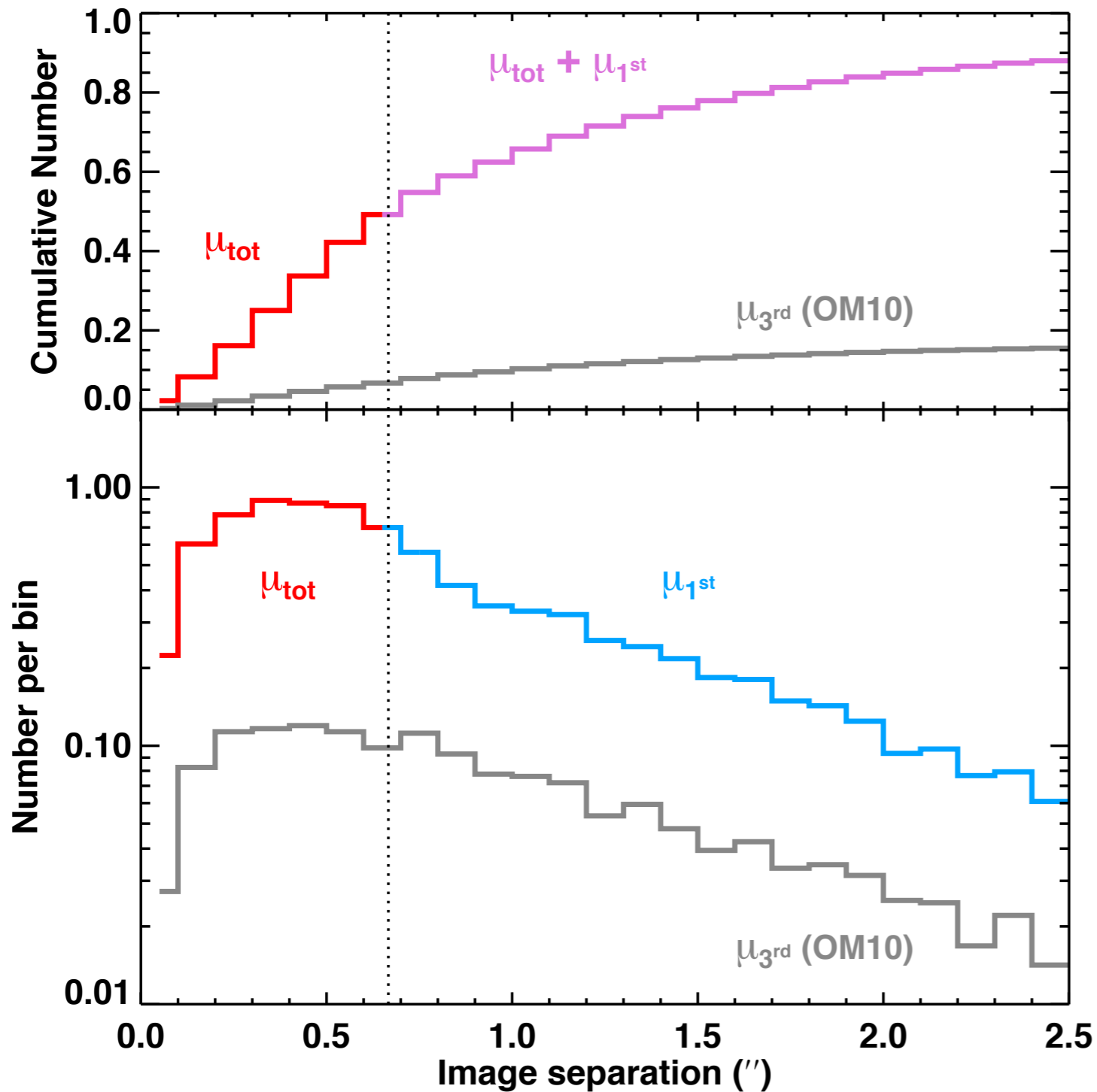


- best-fit stellar mass $M_* \sim 9 \times 10^9 M_\odot$, age $\sim 1 \text{ Gyr}$
- low lens mass suggests small image separation and time delay, which are consistent with obs

Consistent with expectation?

- Oguri & Marshall (2010) predicted ~ 0.1 lensed SNIa in PS1-MDS
- however, it assumed multiple images be resolved and detected, unlike PS1-I0afx
- updated calculation indicates ~ 1 lensed SNIa in PS1-MDS, quite consistent with the discovery!

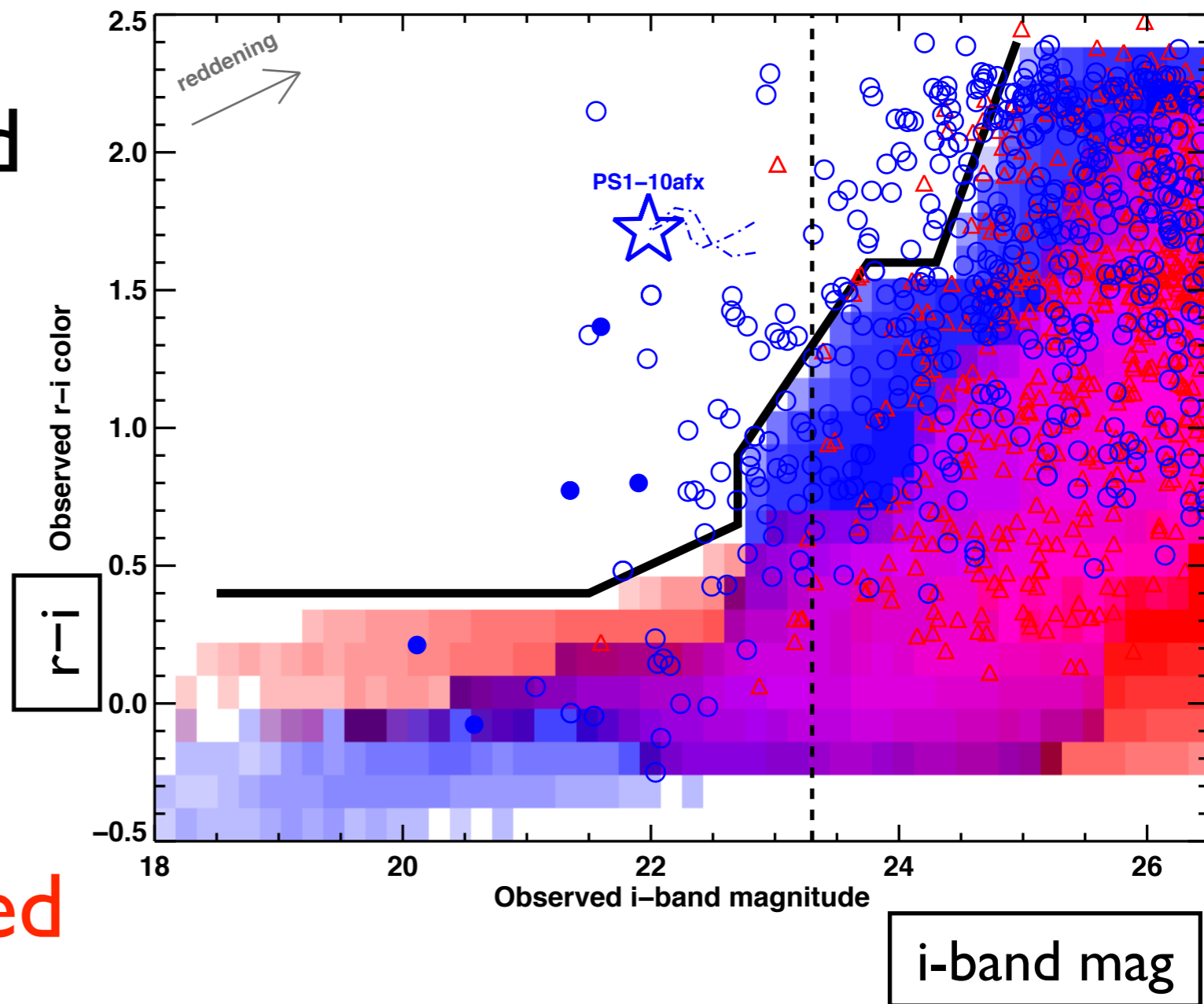
Expected number distribution



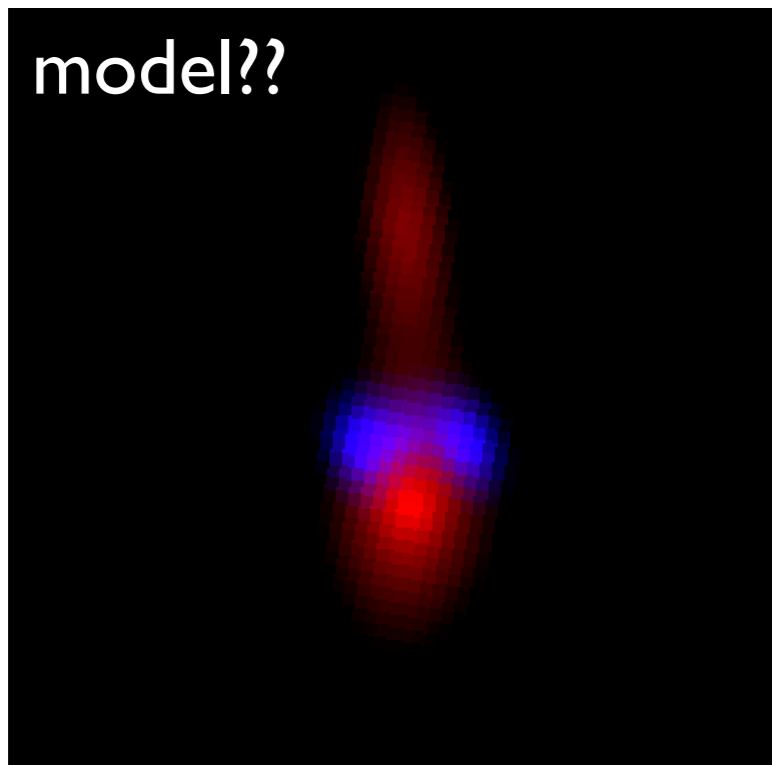
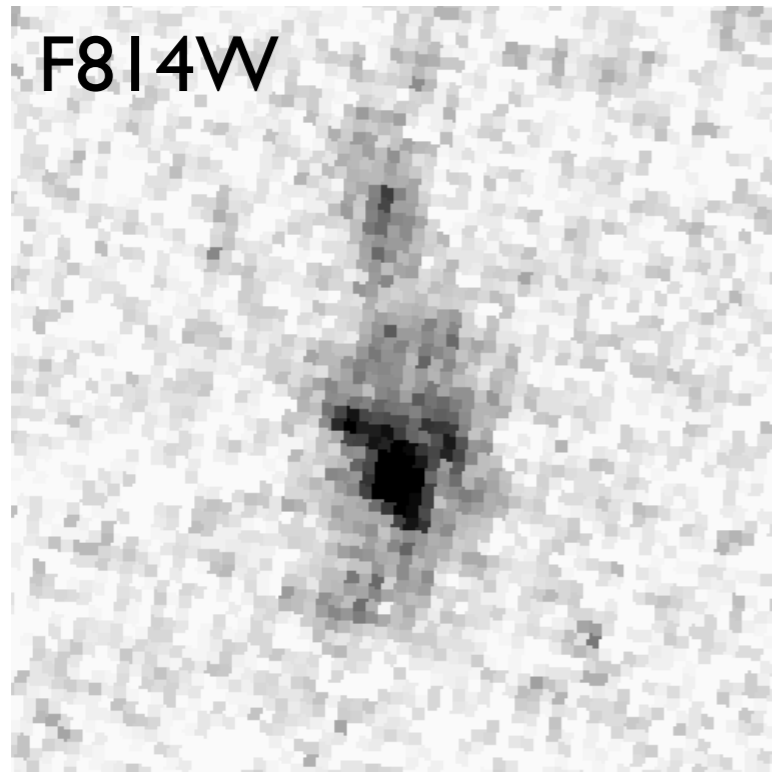
allow unresolved and/or
single image detection
enhance the expected
number by **an order of
magnitude**

New approach to find lensed SNe

- red, bright SNe are almost always lensed SNe!
- find unresolved lens events in surveys, quick follow-up to get multiple images and time delays
- can find **~1000 lensed SNe** in LSST!



HST image of PS1-10afx



- F814W HST image taken by Pan-STARRS team (after SN faded away)
- reveal its complex structure, further supporting lensing scenario
- need more observations to robustly disentangle lens and host galaxies

Summary

- gravitationally lensed supernovae are rare but very interesting phenomena
- in particular the standard candle nature of SNIa enables direct measurement of magnification factor, which contains rich info on the lens
- we show that the very unusual SN in PS1-MDS, PS1-10afx, is in fact the normal type Ia SN magnified by a factor of 30, representing the first discovery of strongly lensed SN
- in LSST, we can discover of order 1000 lensed SNe by identifying them in the color-mag space