

# Searching for Dark Matter with Galileo atomic clocks

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# Outline

- Introduction
- Theoretical approach
- Statistical analysis
- First theoretical results

# Introduction

- Atomic clocks are a good tool to search a uniform-in-time drifts of fundamental constants
  - => Such transient in time changes of fundamental constants can be induced by Dark Matter objects
  - => Network of correlated atomic clocks (GPS) can be used to search DM objects

$$\frac{\delta\omega(r, t)}{\omega_c} = \sum_X K_X \frac{\delta X(r, t)}{X}$$

- Scalar interaction ( $\varphi$ ) between DM and atoms of clocks induces shifts in effective values of fundamental constants:
  - fine structure constant  $\alpha$ ,
  - the ratio of the light quark mass to the quantum chromodynamics  $m_q/\Lambda_{\text{QCD}}$ ,
  - the electron and proton (fermion) masses  $m_e$  and  $m_p$

$$\frac{\delta X(r, t)}{X} = \Gamma_X \varphi^2(r, t)$$

# Model



- 24 Galileo satellites with atomic clocks (H-maser) on boards
- 3 months of data
- Clocks give time by counting number of oscillations and multiplying them by the period
- Experimentally relevant quantity is the total phase accumulated by the quantum oscillator  $\varphi_0(t) = \int_0^t \omega_0 dt'$
- Interaction with DW leads to a shift in the oscillator frequency:

$$\varphi(t) = \int_0^t (\omega_0 + \delta\omega(t')) dt'$$

# Comparison signal from two clocks

Before DW arrival, clocks are synchronized  $\Rightarrow \Delta\varphi(t) = 0$

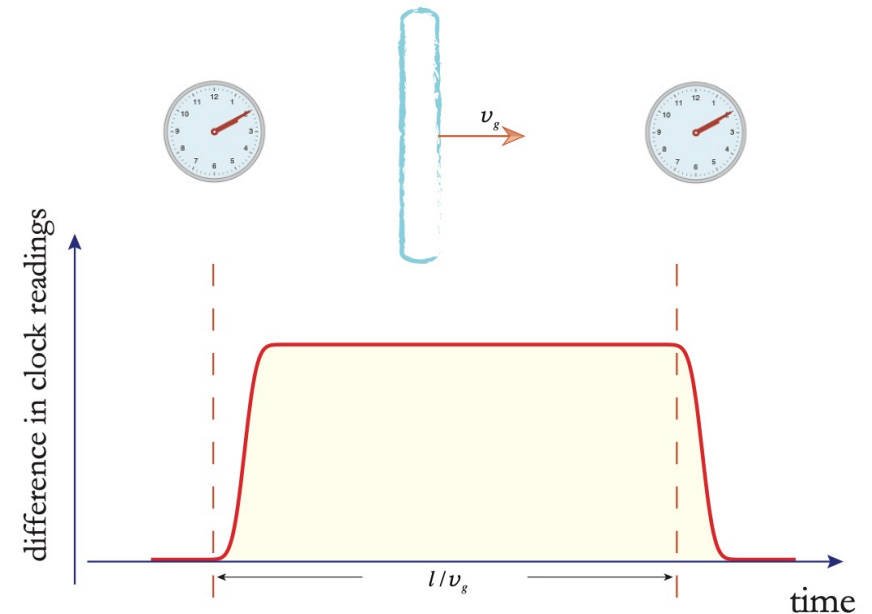
- DW passing the first clock introduces a phase difference  
 $\Delta\varphi(t) \sim d/v_g$

DW propagates through the clock network with  $v_g$ :

$$\Delta\varphi(t) \sim \int_{-\infty}^t \left( f\left(t' - \frac{l}{v_g}\right) - f(t') \right) dt' = \omega_0 \Delta t(t)$$

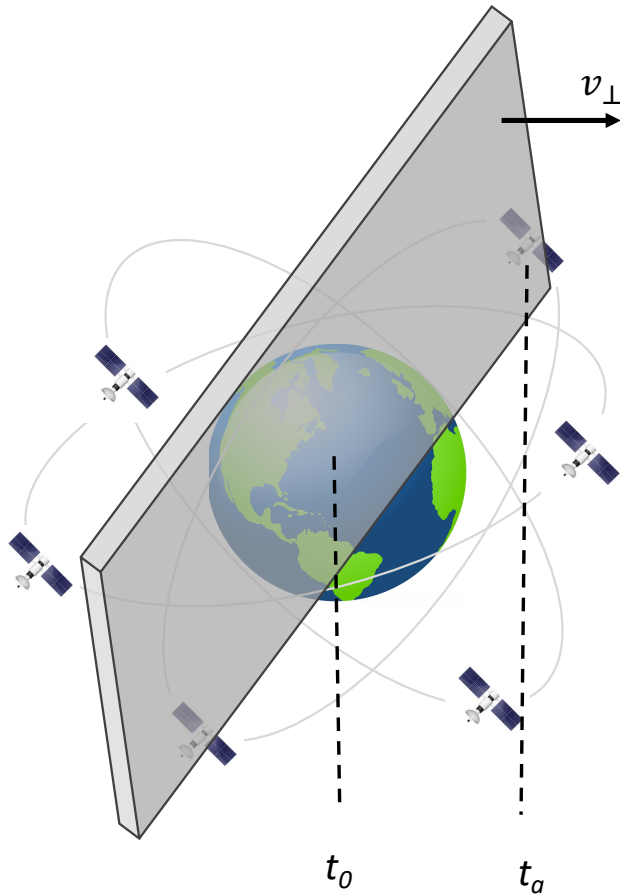
- After passing the second clock, the phase difference vanishes

$\Rightarrow$  Monitoring correlated time difference  $\Delta t(t)$  between two clocks, one can search for DM



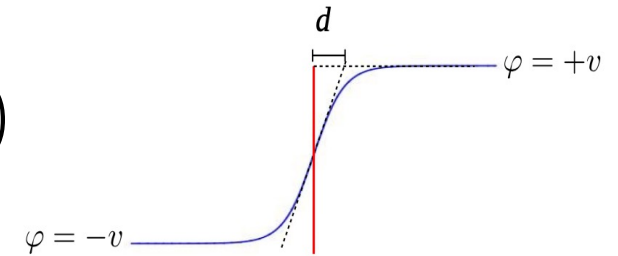
A. Derevianko, M. Pospelov, Nat. Phys. **10**, (2014)

# Signal simulation



DM signal:  $\varphi(z, V, d) = V \cdot \tanh\left(\frac{z}{2d}\right)$

$$V^2 = \frac{3}{2} v_{gal} \Im d \rho_{DM}$$



The distance between DW and an aircraft  $a$ :

$$z_a(t, v_{\perp}, t_0) = -v_{\perp}(t - t_0) - \mathbf{r}_a(t) \mathbf{n}_{\perp}$$

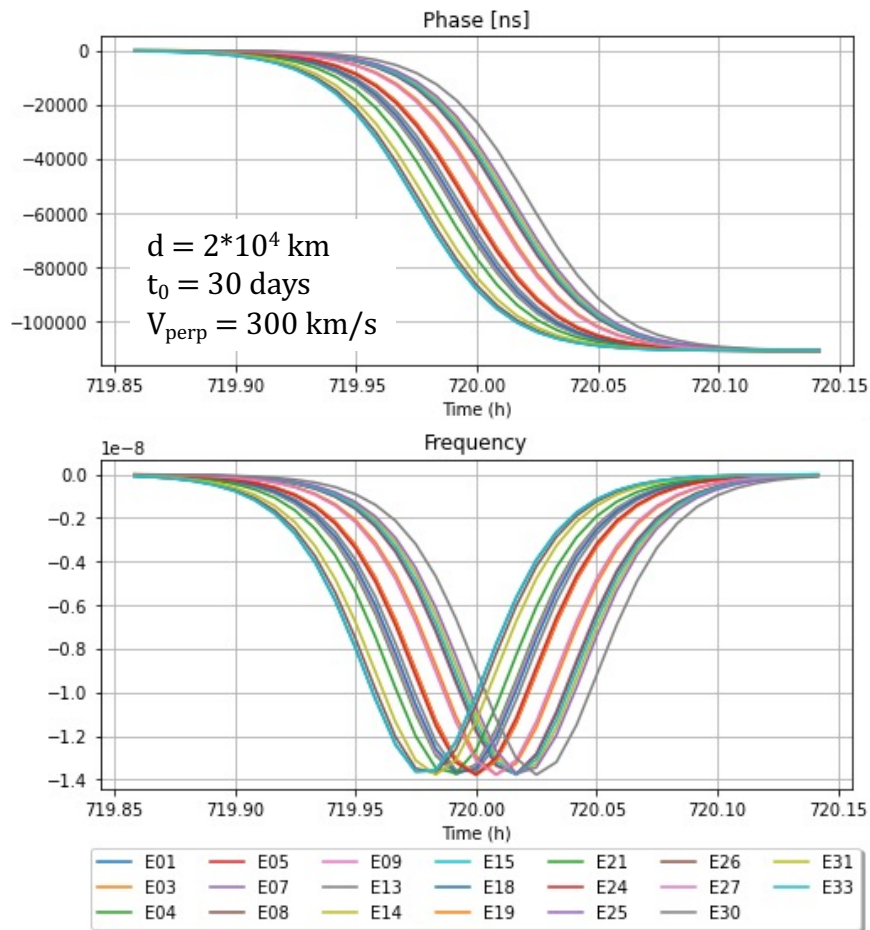
Signal from DW:  $S_a(t) = \int_{-\infty}^t \Gamma_a^{\text{eff}}(\varphi^2(z_a(t', v_{\perp}, t_0), V, d) - V^2) dt'$

The first-order difference:

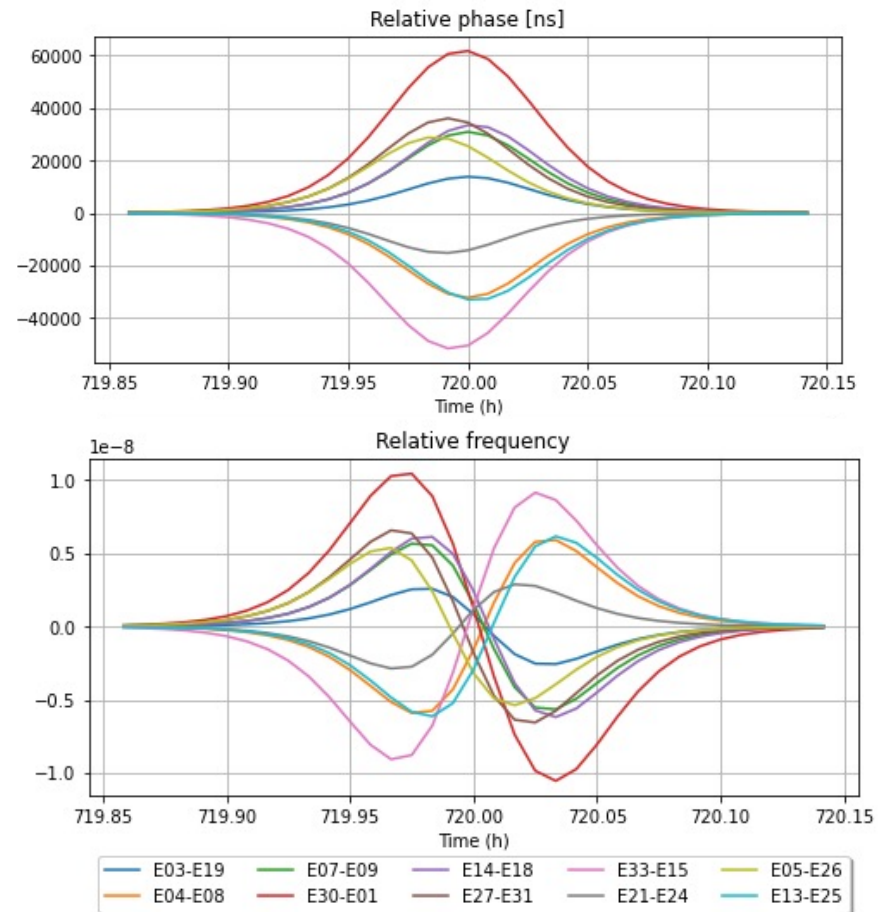
$$\Delta S(t) = S(t) - S(t - \Delta t) = \int_{t-\Delta t}^t \Gamma_a^{\text{eff}}(\varphi^2(z_a(t', v_{\perp}, t_0), V, d) - V^2) dt'$$

# Signal simulation

- Thin DW ( $d < 10^4$  km)

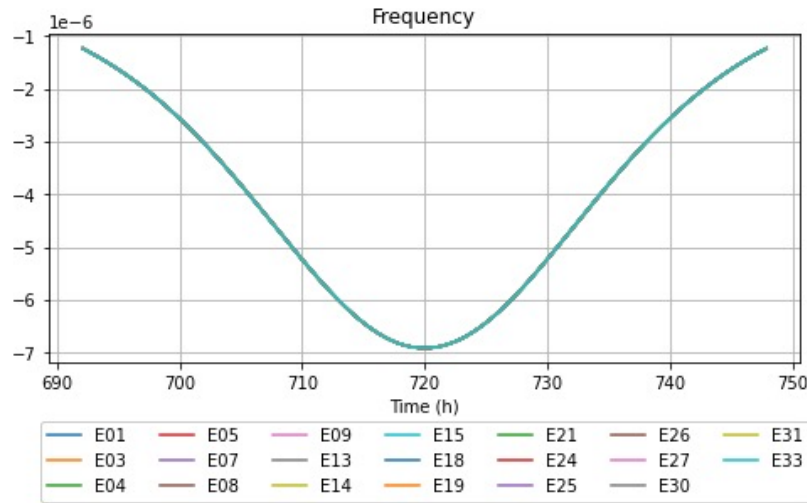
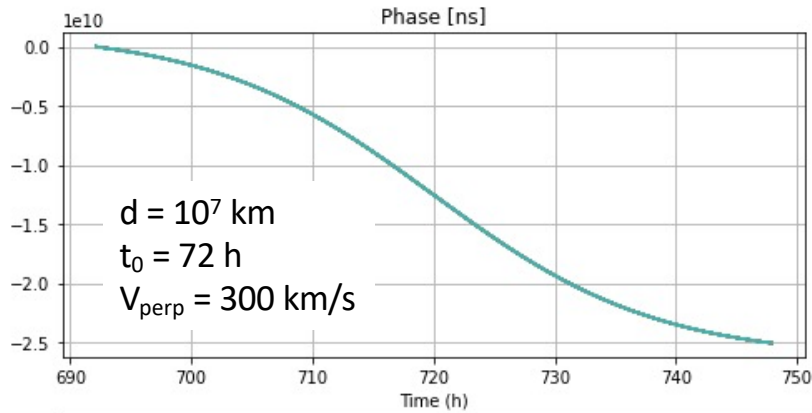


## Relative differences

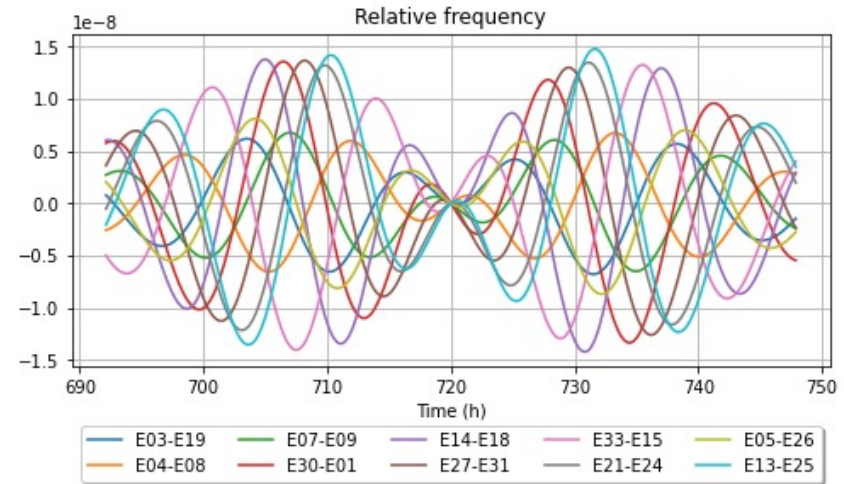
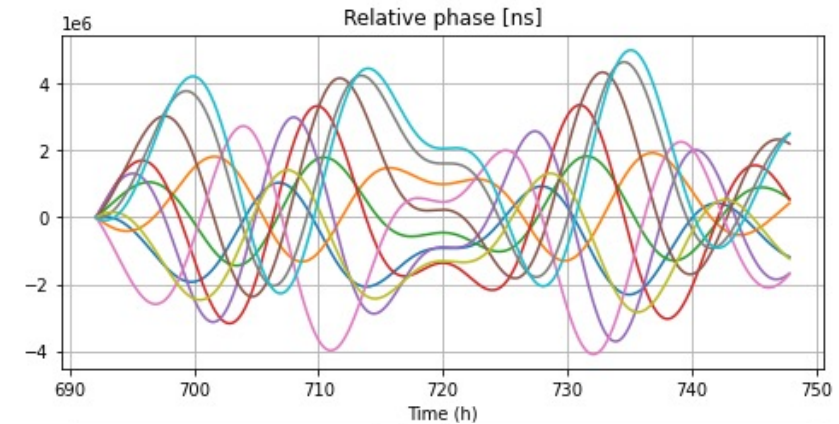


# Signal simulation

- Thick DW ( $d > 10^4$  km)

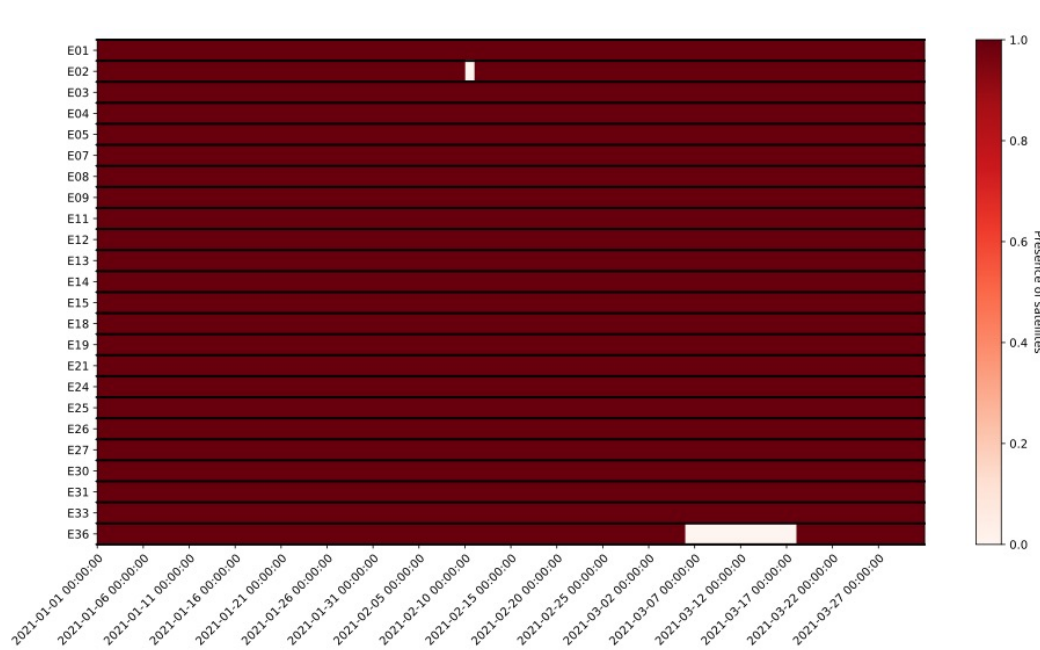


## Relative differences

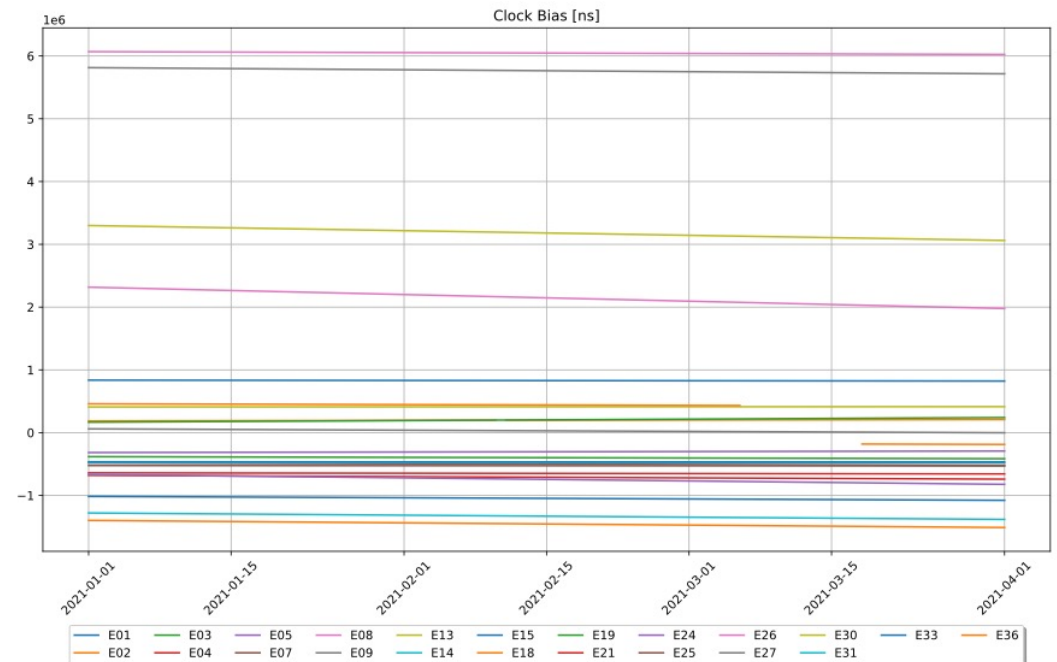




# Presence of data



- 2 satellites have missing data  
=> We removed them from the analysis



- Different linear drift for each clocks  
=> should be removed

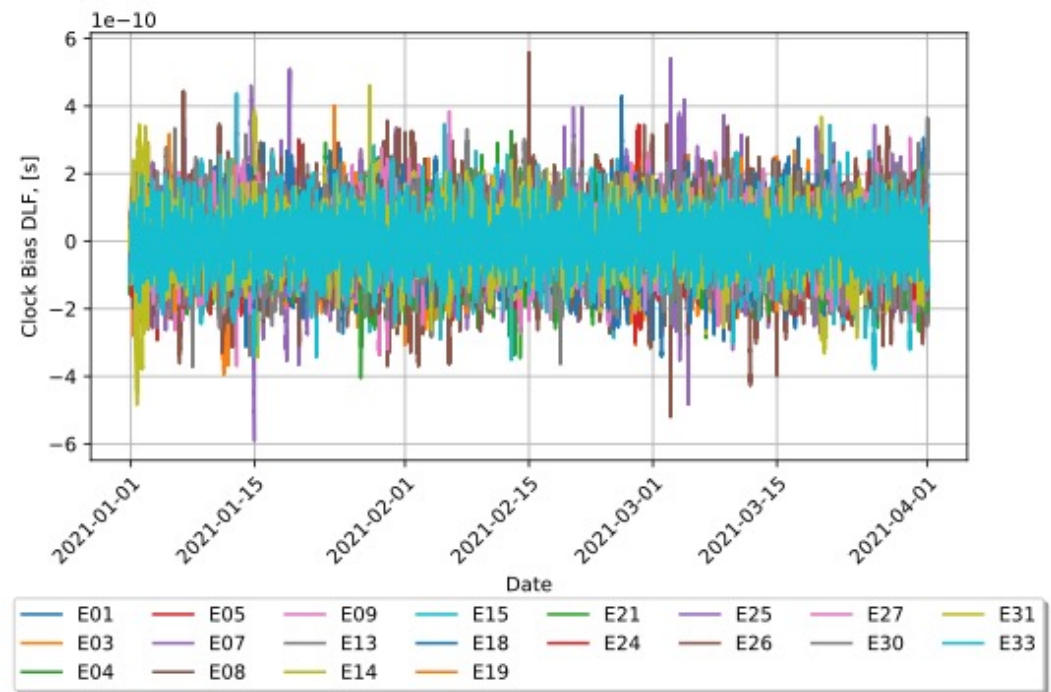
# Pre-process

From Clock Bias, we remove daily linear fit



2 clocks have problems:

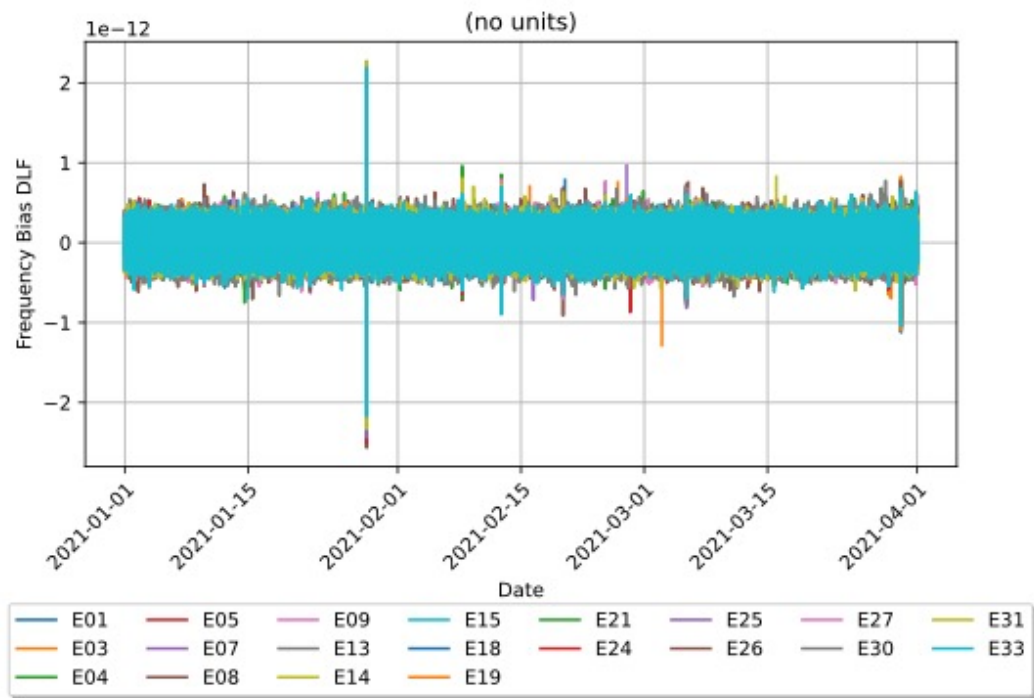
- E11 works with Rb clock,
- E12 has a big jump



Clock bias with removed daily linear fit after removing E02, E11, E12, E36 clocks

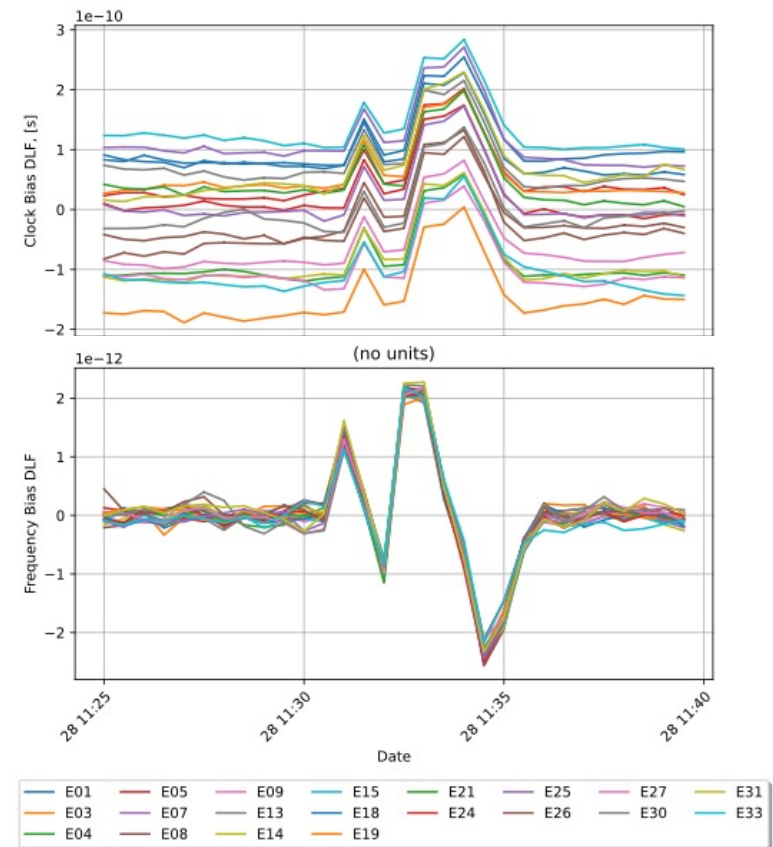
# Pre-process

From Clock Bias, we also can calculate a derivative, which corresponds to frequency



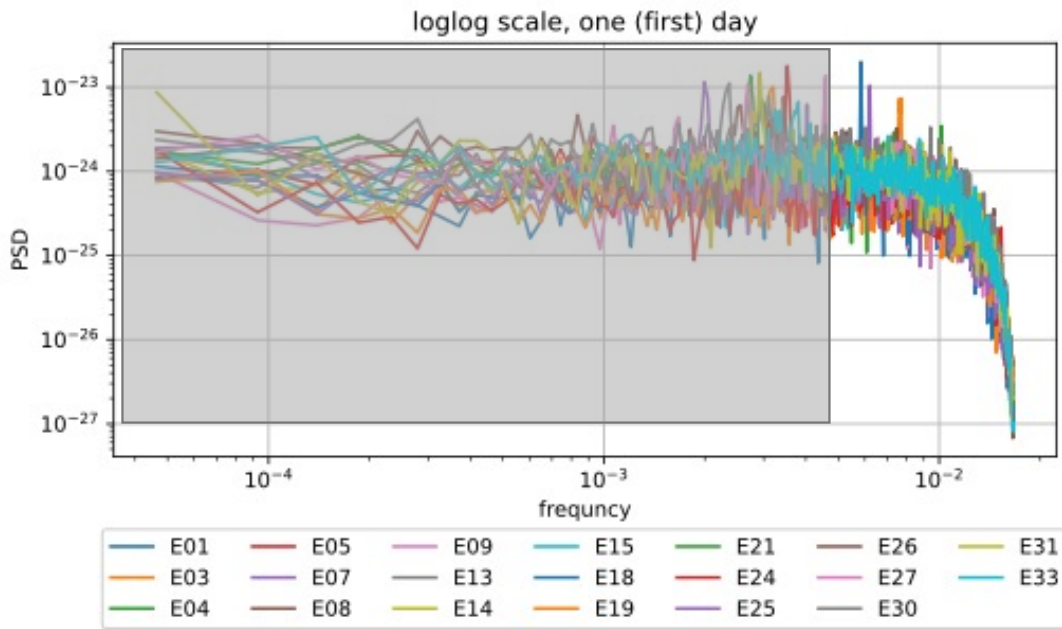
Frequency bias with removed daily linear fit after removing E02, E11, E12, E36 clocks

Jump from a ground reference clock



# Noise characterization

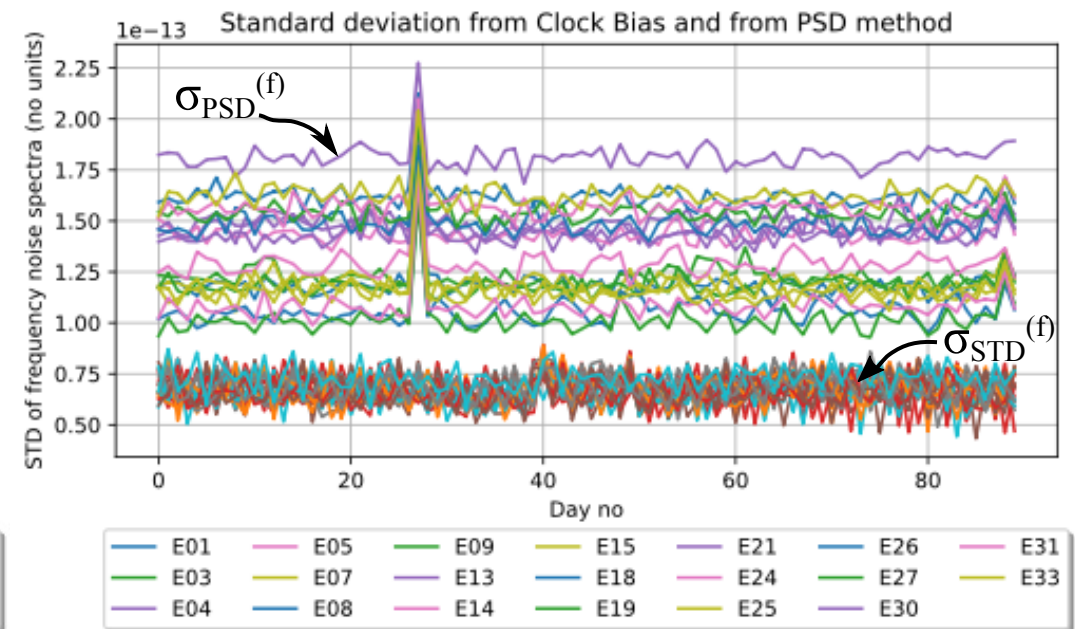
PSD (Welch method) for one day



=> white noise

$$\sigma_{\text{PSD}} = \sqrt{\frac{b_0 \cdot f_s}{2}} \quad f_s = 1/30$$

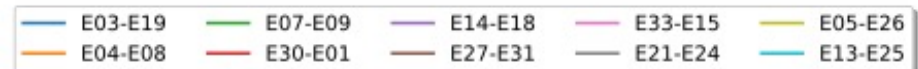
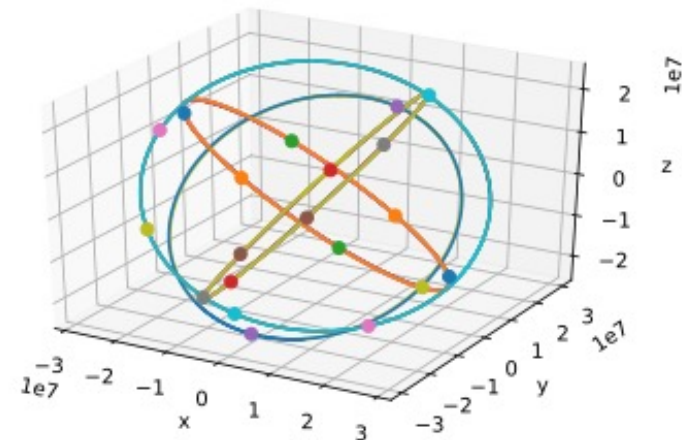
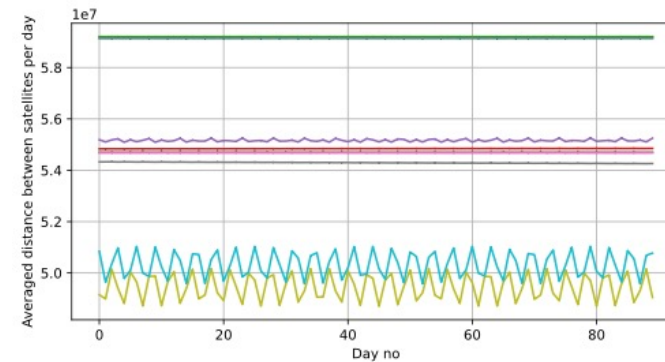
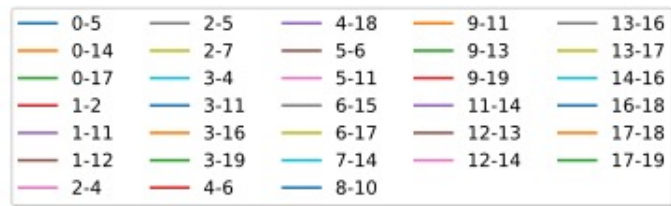
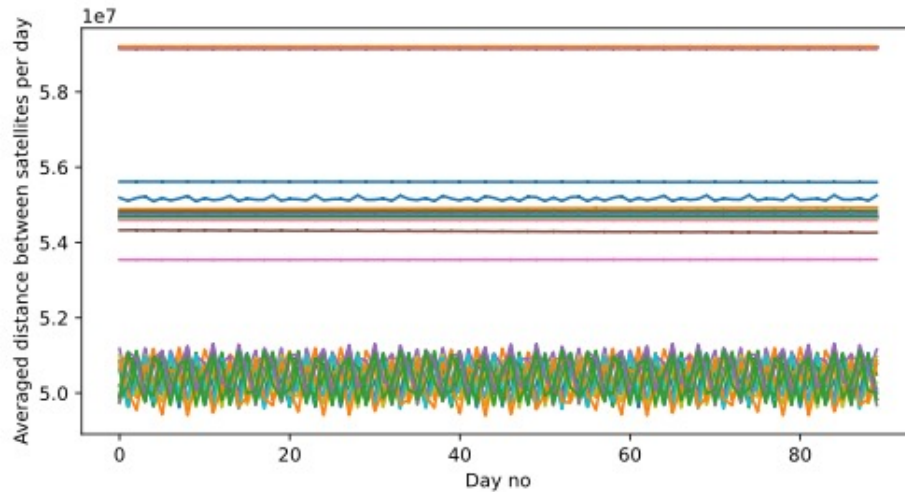
$\sigma$  for linear region for each satellite for every day



Comparison with sigma from clock files

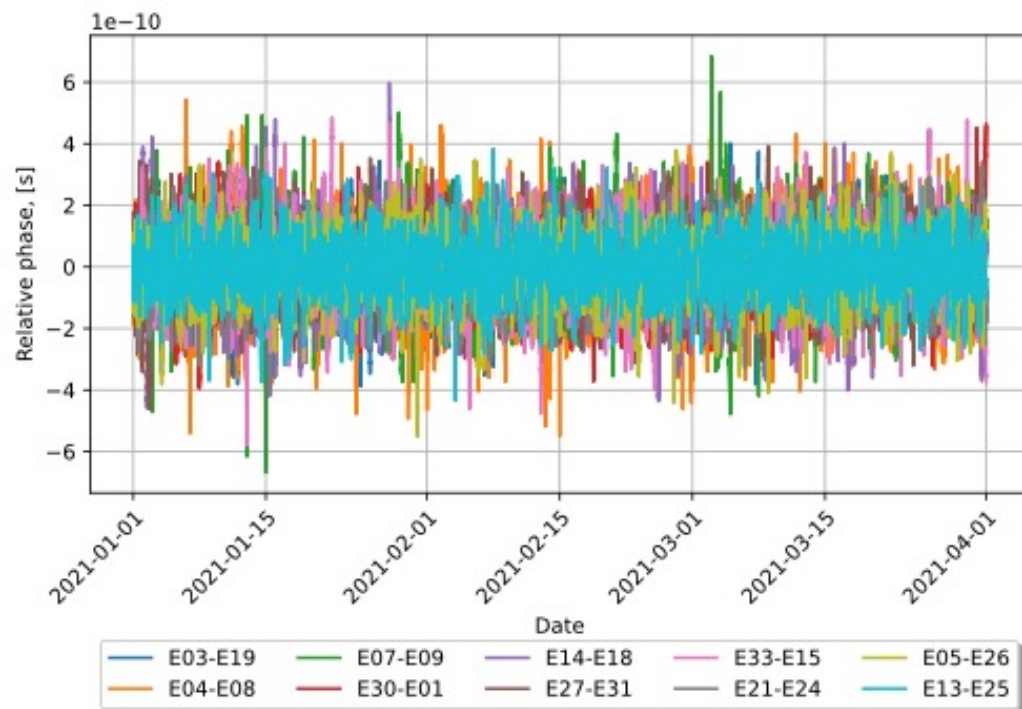
$$\sigma_{\text{STD}}^{(f)} = \frac{\sigma_{\text{STD}}^{(t)}}{\Delta T}$$

# Choosing independent pairs





# Relative phase and frequency differences



# Data Analysis: frequentist approach

1. Posterior probability  $p \equiv p(D_t|\varphi(\zeta)) = Cp(\zeta)\exp\left(-\frac{1}{2}[d-s]^T\mathbf{C}^{-1}[d-s]\right)$

$\zeta$  defines DW parameters: incident direction, velocity, encounter time

$\mathbf{C}$  is the covariance matrix

$d$  – data

$s \sim hs(t_{0j}, d_j, \mathbf{v}_{\text{perp},j})$  – simulated signal

2. We work with independent pairs

$$\begin{aligned} d_i &= d_a - d_b \\ s_i &= s_a - s_b \end{aligned} \Rightarrow C_{ii} = \sqrt{\sigma_{\text{PSD},a}^2 + \sigma_{\text{PSD},b}^2}, i \text{ is pair of clocks } a \text{ and } b$$

3. Posterior on the signal amplitude ( $D_t$ : templates with  $\theta, \varphi, \mathbf{v}_{\text{perp}}, d, t_0$ )

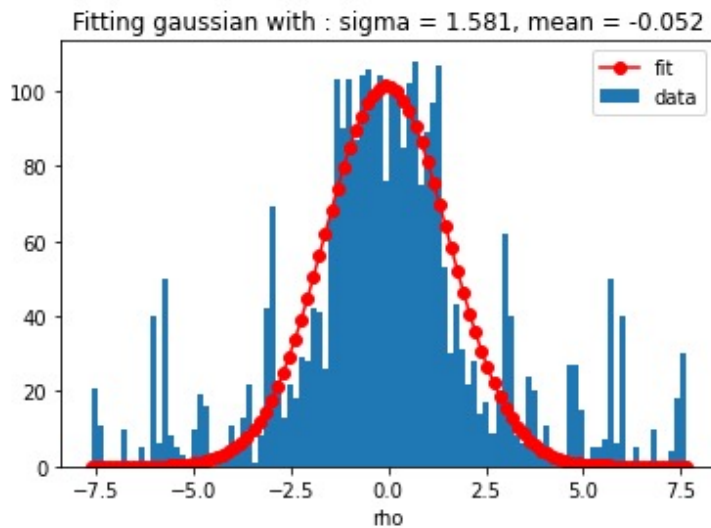
$$h_\zeta = \frac{\mathbf{d}^T \mathbf{C}^{-1} \mathbf{s}(\zeta)}{\mathbf{s}^T(\zeta) \mathbf{C}^{-1} \mathbf{s}(\zeta)} = \frac{\sum_{i,k}^{Ncp} \sum_{j,l}^{D_t} d_j^i (C^{-1})_j s_l^k(\zeta)}{\sum_{i,k}^{Ncp} \sum_{j,l}^{D_t} s_j^i(\zeta) d_j^i (C^{-1})_j s_l^k(\zeta)} \quad \sigma_{h_\zeta}^2 = \frac{1}{\mathbf{s}^T(\zeta) \mathbf{C}^{-1} \mathbf{s}(\zeta)}$$

$$\text{Signal to Noise Ratio} \quad \text{SNR} = \frac{h_\zeta}{\sigma_{h_\zeta}} = \frac{\mathbf{d}^T \mathbf{C}^{-1} \boldsymbol{\varphi}(\zeta)}{\sqrt{\boldsymbol{\varphi}^T(\zeta) \mathbf{C}^{-1} \boldsymbol{\varphi}(\zeta)}}$$

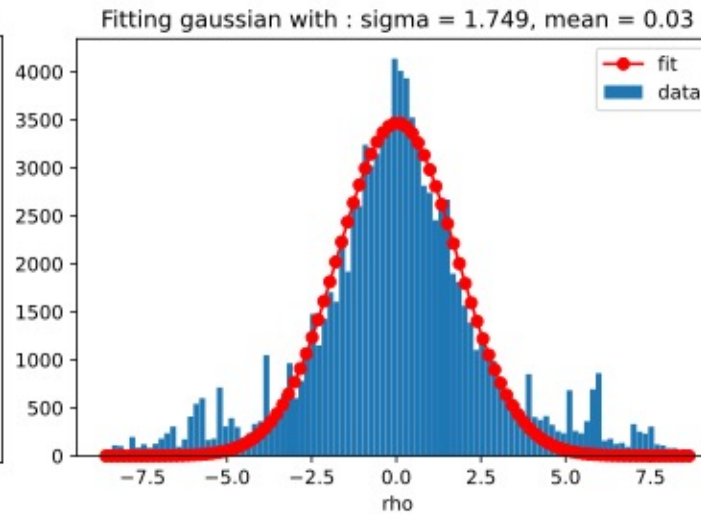
# Data Analysis: frequentist approach

1. Forming templates for 5 parameters:  $(t_0, d, v, \theta, \varphi) \Rightarrow$  calculation SNR on event free simulated data

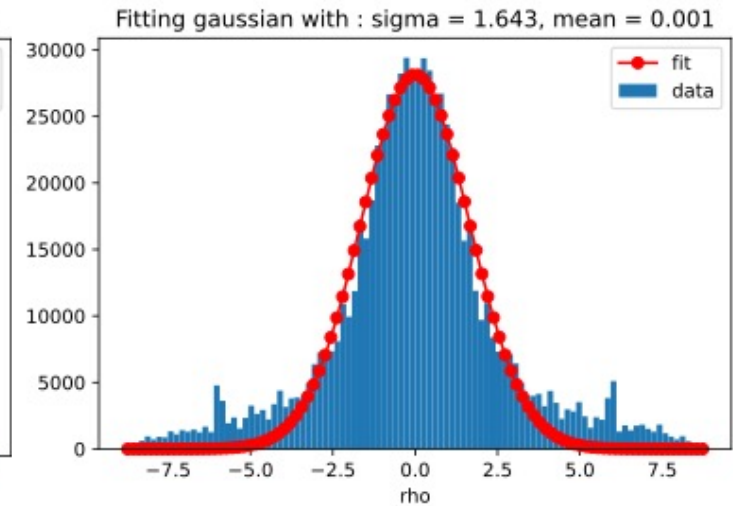
$5^5$  templates



$10^5$  templates



$15^5$  templates



$\Rightarrow$  Gaussian distribution



# Still need to be done...

- Estimate a threshold value

$$1 - \left[ \operatorname{erf} \left( \frac{\rho_{\text{thresh}}}{\sqrt{2}} \right) \right]^{N_{\text{templates}}} = 0.05$$

- SLR campaign has provide high quality of the clock data with very small residual with respect to daily linear fits .
- In the case of a positive event of DW signal, we will need to exclude the signal from an orbital error. This will be possible by analysing SLR residual. This comparison can be done.

# Conclusions

- Our analysis can be applied to any model describing the interaction of the usual Maxwell Lagrangian with scalar fields beyond the Standard Model
- The tangential profile of DM field gives more information about interaction process
- Our model allows to consider thick domain walls which was not considered before

Thank you for your attention!