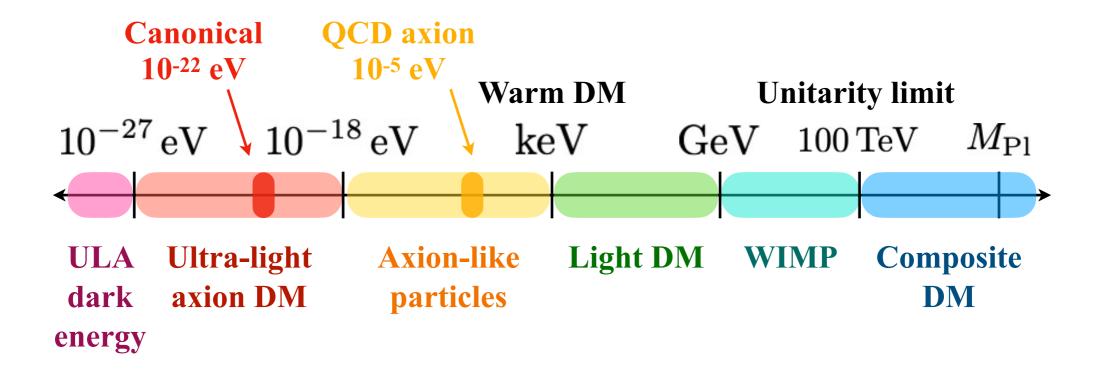
# COSMIC PROBES OF DARK MATTER PHYSICS

Keir K. Rogers

Dunlap Fellow, Dunlap Institute for Astronomy & Astrophysics,

University of Toronto

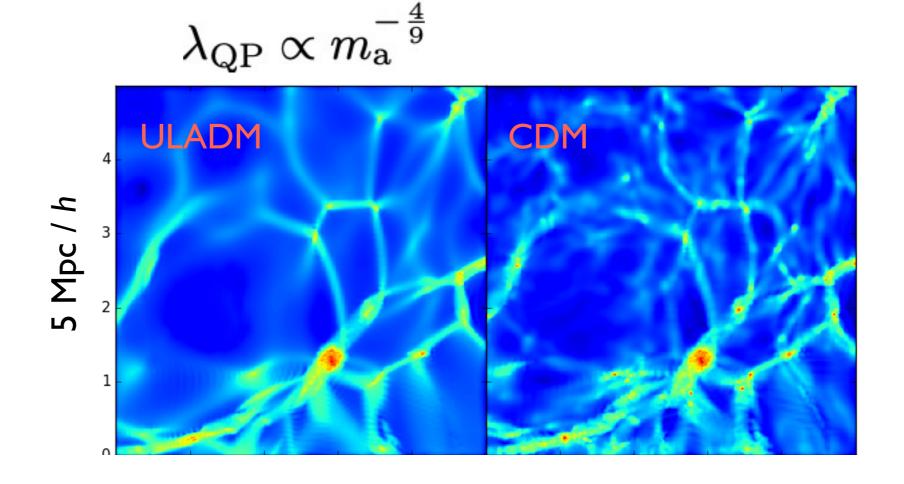
### Dark matter candidates well-motivated at many masses



#### Ultra-light axion is a compelling dark matter candidate

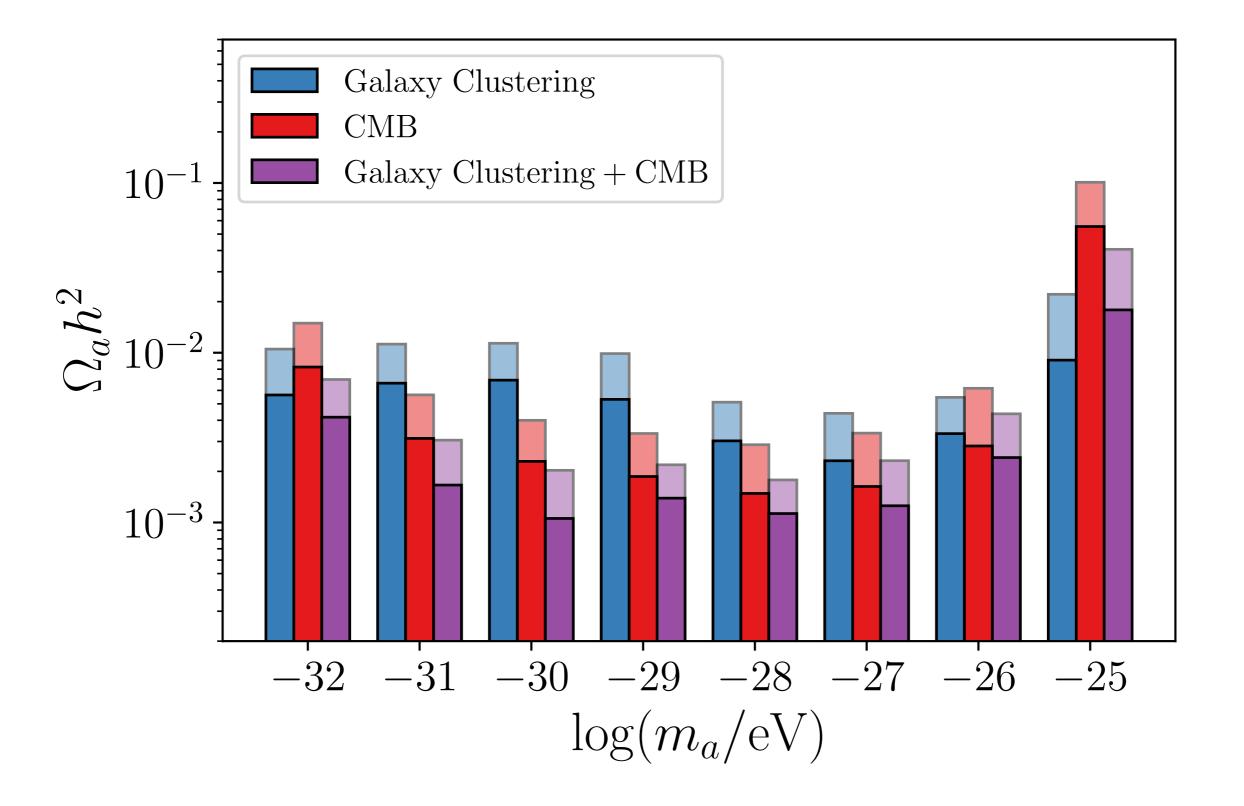
- Axion-like particles are lowmass bosons abundantly produced in BSM theories, e.g., string theory axiverse
- New physics to resolve cosmological model tensions

$$10^{-27} \,\mathrm{eV} < m < 10^{-18} \,\mathrm{eV}$$



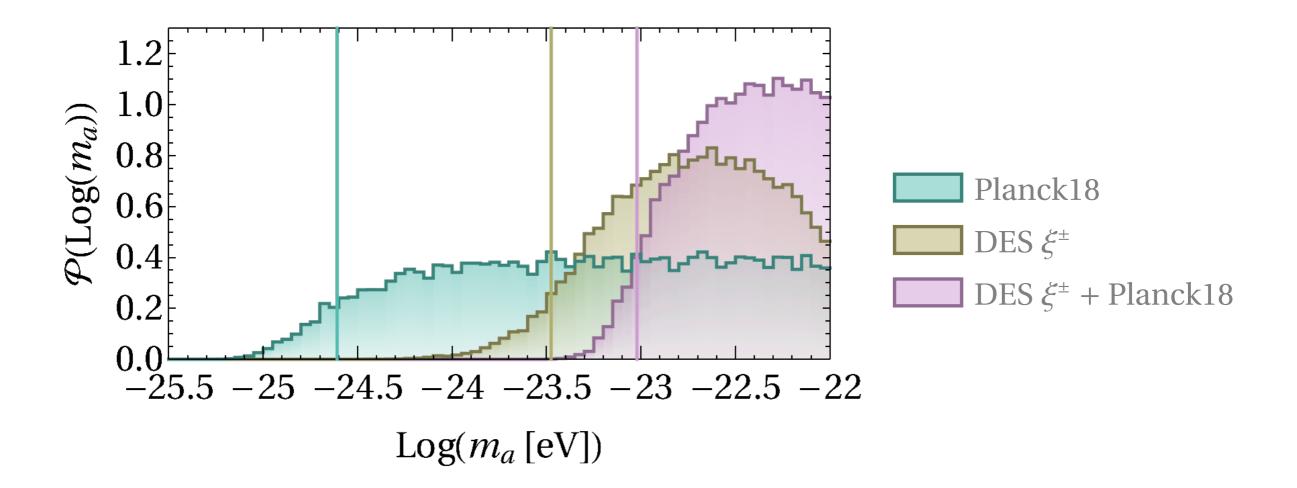
Hui et al. (2017); Allali et al. (2021); Blum & Teodori (2021); Armengaud et al. (2017)

#### Combining galaxy clustering from BOSS with *Planck* CMB improves bound by up to 4.5 x



Laguë, Bond, Hložek, Rogers, Marsh, Grin (2021, in review)

#### Joint CMB & galaxy weak lensing bounds



### Ultra-light axions are invoked to resolve so-called cold dark matter "small-scale crisis"



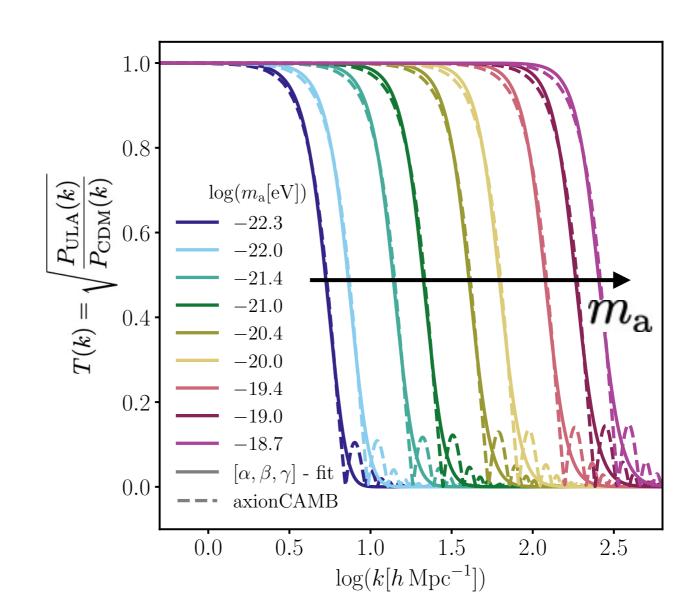
$$\lambda_{\mathrm{dB}} \sim \mathrm{kpc}$$
 $\lambda_{\mathrm{QP}} \sim \mathrm{Mpc}$ 

CDM "small-scale crisis" prefers
DM mass scale ~ 10-22 eV

### Lyman-alpha forest traces ultra-light axion dark matter cut-off

 $k_{rac{1}{2}} \propto m_{
m a}^{rac{4}{9}}$ 

Hu et al. (2000)



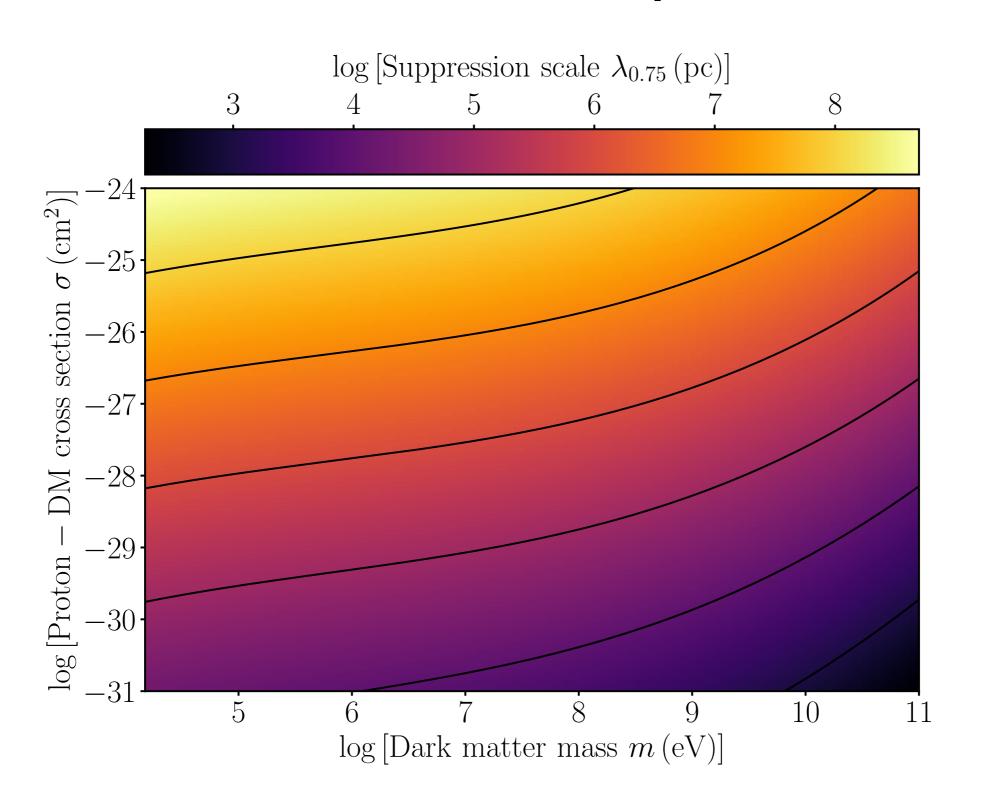
• Ly-alpha forest traces linear, high-redshift ( $z \sim 5$ ), small-scale density perturbations



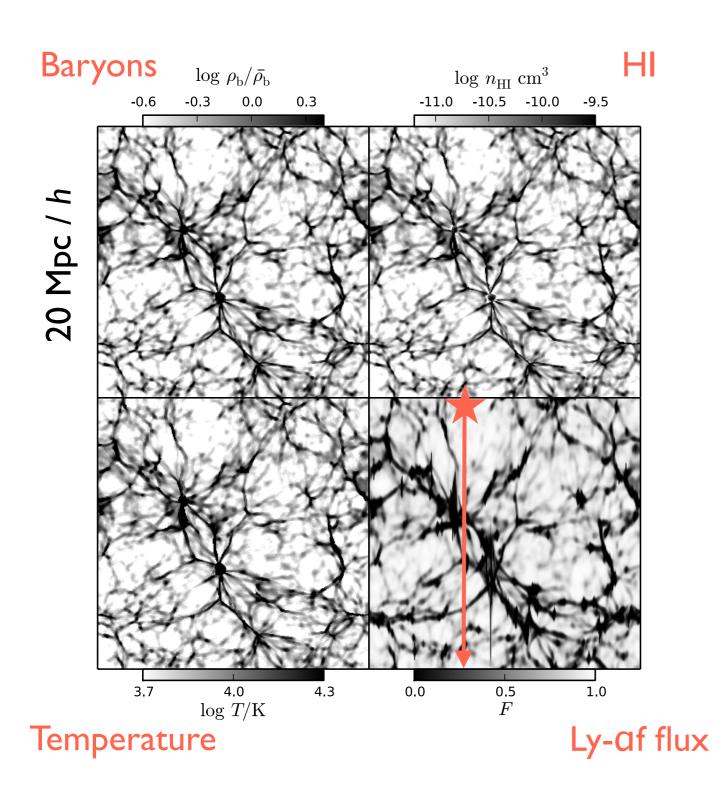




## Smaller cosmological scales probe light DM that cannot be directly detected



#### Dark matter bounds must marginalise astrophysics



- Ly-alpha forest traces DM & intergalactic medium astrophysics
- ~ 3000 CPU-hours per simulation in I2-D parameter space
- ⇒ need ML-accelerated "emulator"



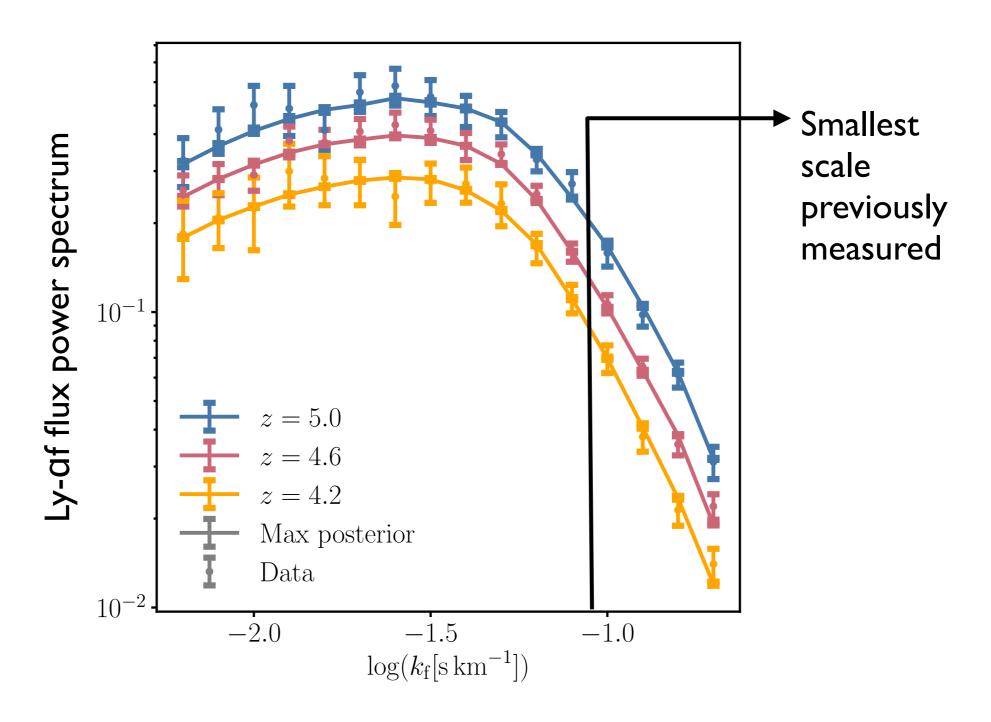




# STRONG BOUND ON ULTRA-LIGHT AXION DARK MATTER

Phys. Rev. Lett., 126, 071302, 2021 Phys. Rev. D, 103, 043526, 2021 with Peiris

#### Dark matter bound driven by new small-scale data



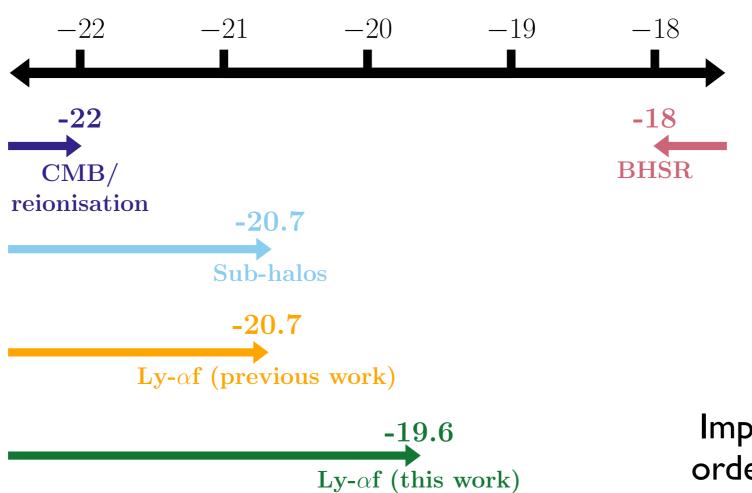






### "Canonical" 10-22 - 10-21 eV ULA DM is strongly disfavoured by new bound

Axion dark matter mass [log(eV)]



Improve bound by order of magnitude

$$m_{\rm a} > 2 \times 10^{-20} \, {\rm eV}$$







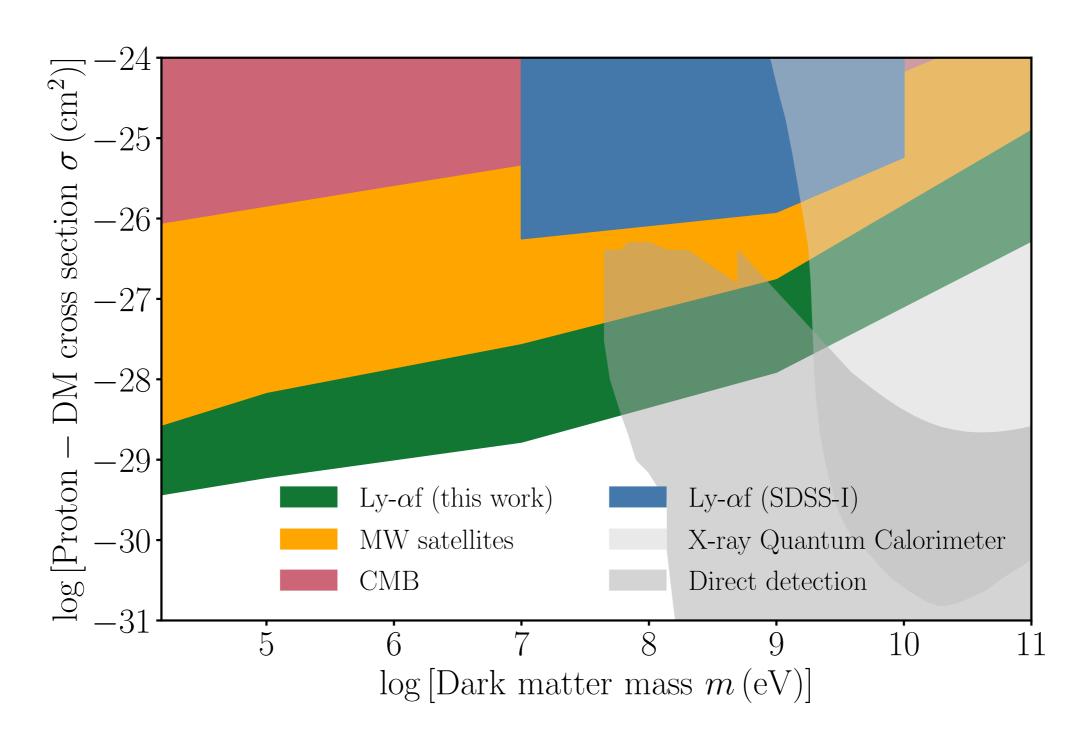


#### NEW LIMITS ON LIGHT DARK MATTER

arXiv:2111.10386

with Dvorkin, Peiris

## Cosmology probes well-motivated light dark matter that cannot be directly detected





#### Multi-probe approach to reach GUT-scale physics

