

2105.04559



Detecting new forces in the gravitational wave background

Benjamin V. Lehmann [blehmann@ucsc.edu]

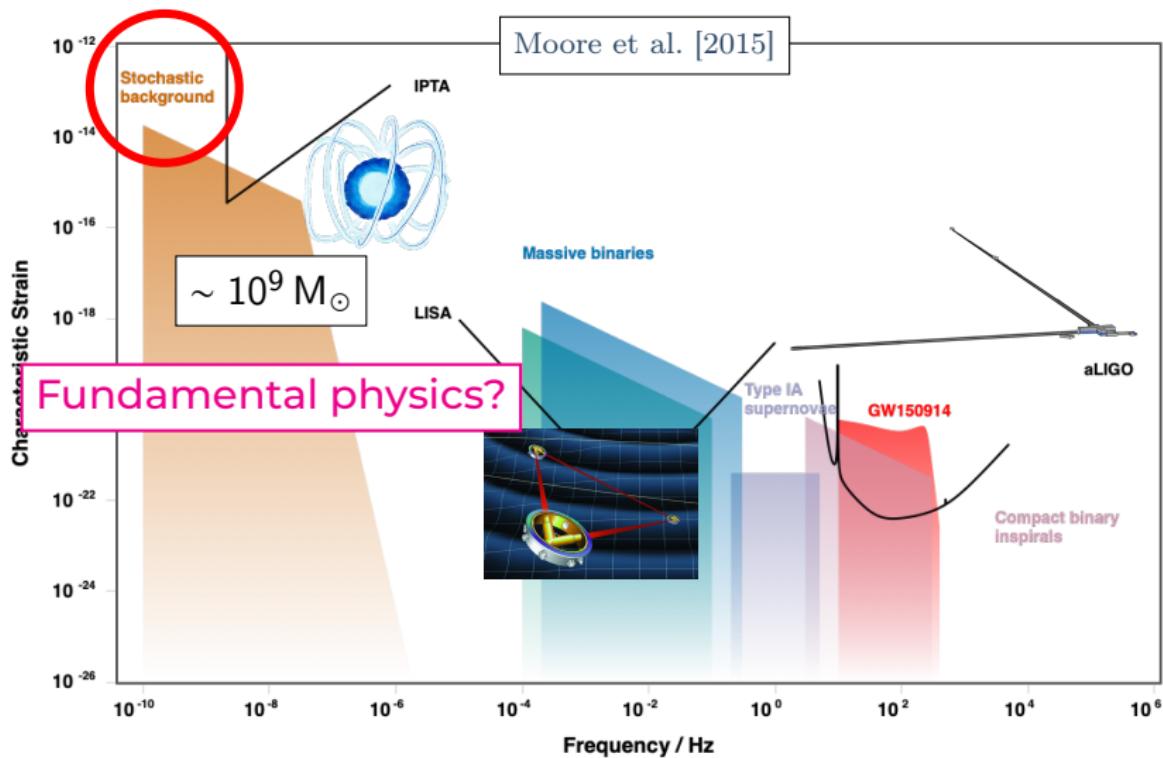
+Jeff A. Dror, Hiren H. Patel, & Stefano Profumo

UC SANTA CRUZ

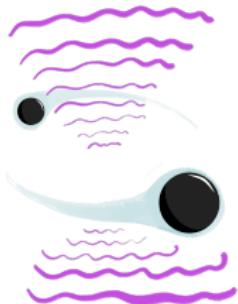
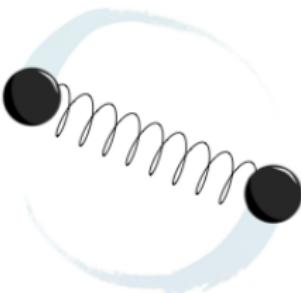


SCIPP
SANTA CRUZ INSTITUTE
for PARTICLE PHYSICS
UC SANTA CRUZ

The era of gravitational waves

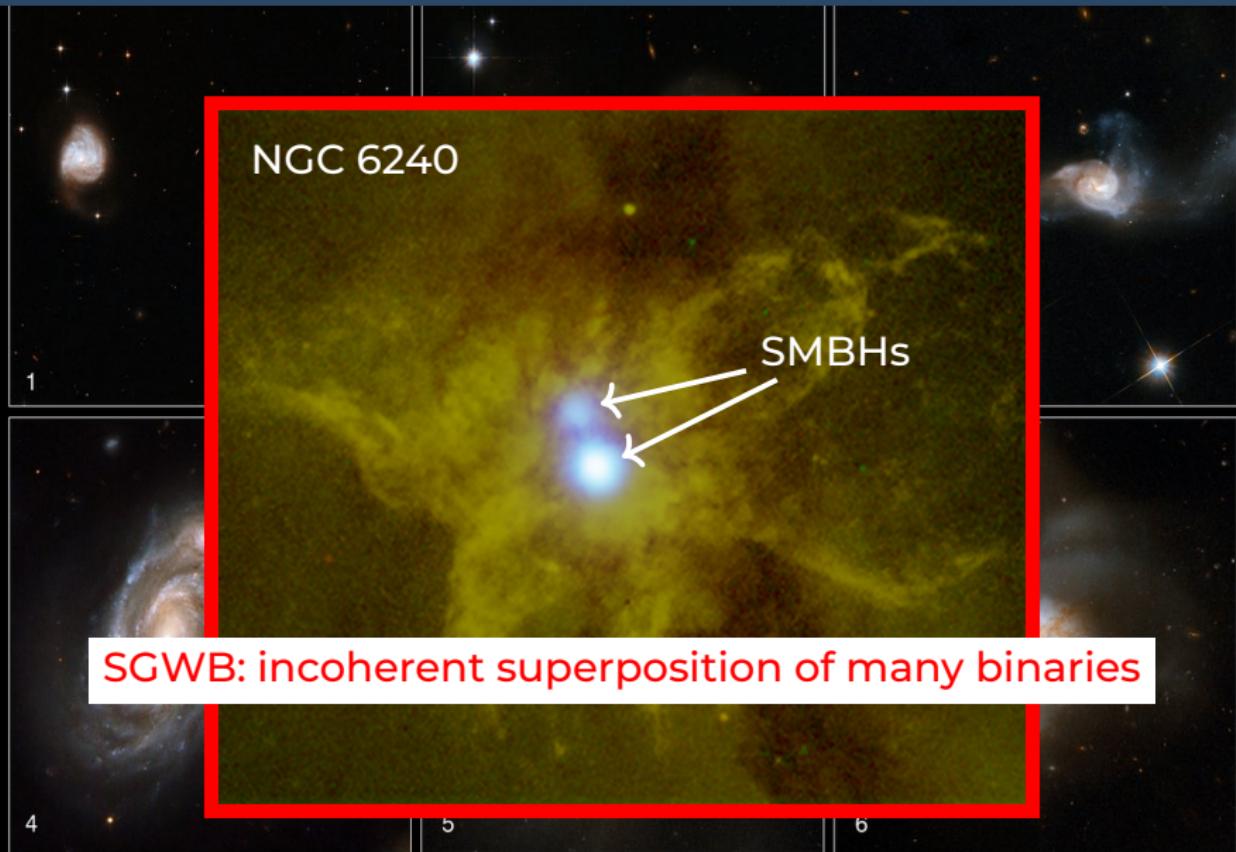


This talk in one slide

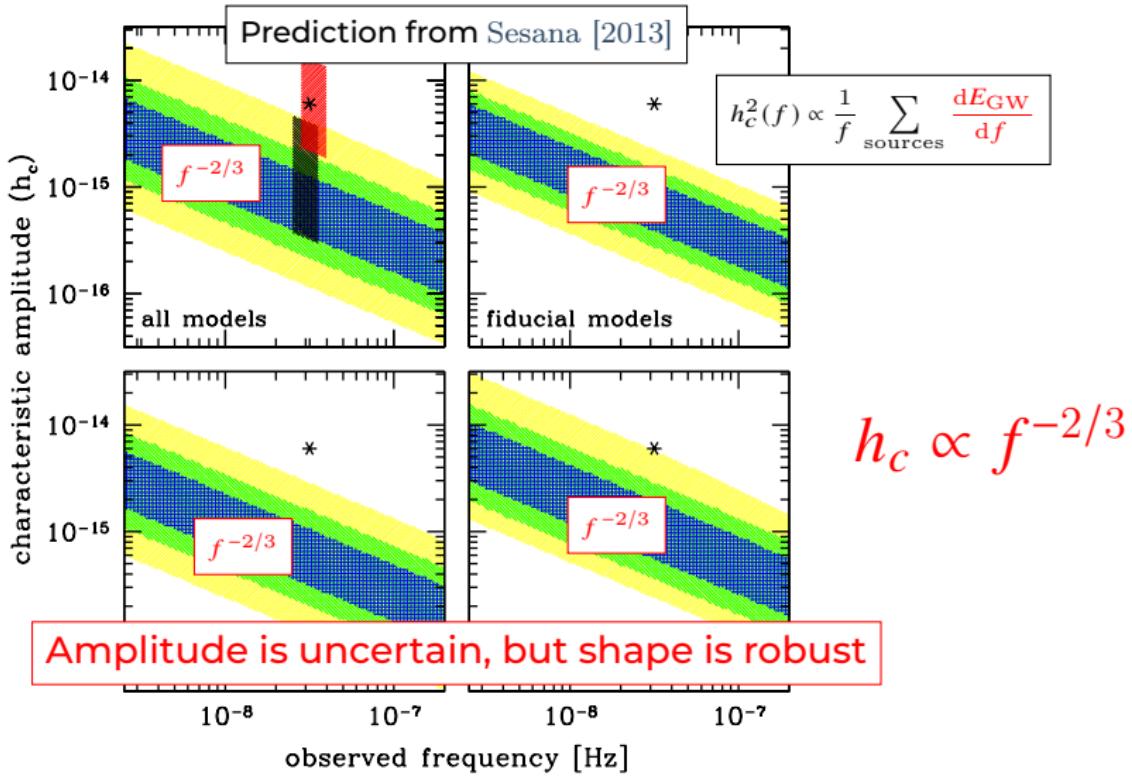


1. SMBH GW background is a guaranteed discovery
2. Long range forces can *detectably* modify spectrum
3. SMBH GWs potentially probe many BSM scenarios

The stochastic GW background

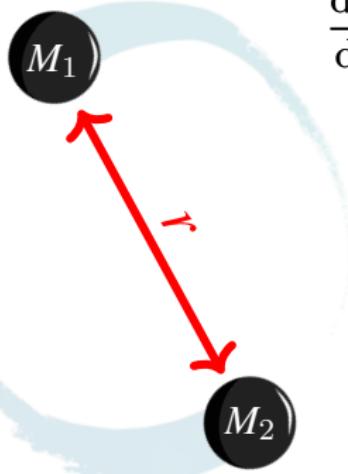


Stochastic background spectrum



Why is the index $-2/3$?

Gravitational waves drive the evolution of the binary



$$\frac{dE_{\text{GW}}}{df_{\text{GW}}} = -\pi^2 \mu r^2 f_{\text{GW}} \left(\frac{2f_{\text{GW}}}{r} \frac{dr}{df_{\text{GW}}} + 1 \right)$$
$$f_{\text{orbit}}(r) = \left(\frac{G(M_1 + M_2)}{4\pi^2 r^3} \right)^{1/2}$$

(Kepler's third law)

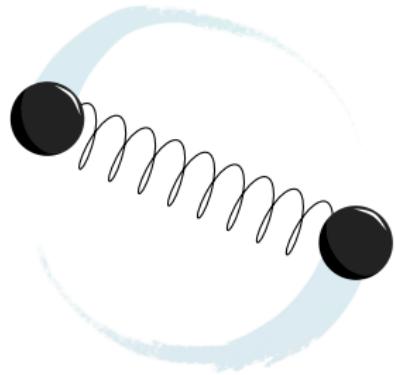
$$\frac{dE_{\text{GW}}}{df_{\text{GW}}} \propto f^{-1/3} \implies \frac{dh_c}{df_{\text{GW}}} \propto f^{-2/3}$$

[Phinney, 2001]

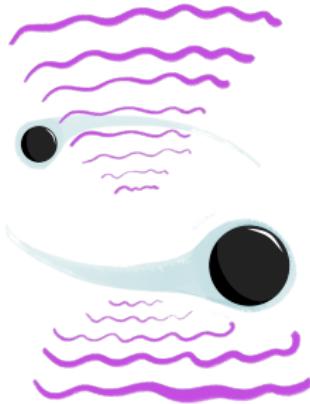
New physics can break this prediction

Assumptions are made to be broken

- ① $f \leftrightarrow r$ relation
Kepler's law



- ② All energy loss is gravitational

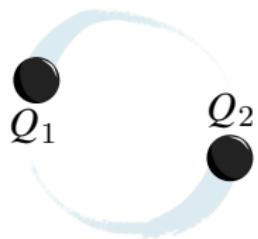


Benchmark model

Additional dynamics spoil the $-2/3$

Toy model: charge BHs* under a new long-range force
*or their surroundings

- ① New force changes Kepler's law
- ② New radiation takes energy

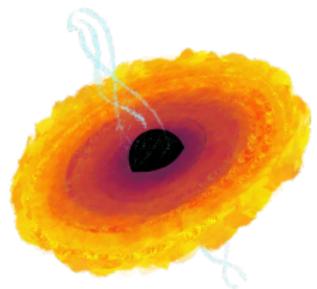


Charge parameters

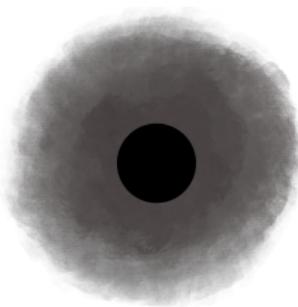
$$\underbrace{\alpha = \frac{Q_1 Q_2}{G M_1 M_2}}_{\text{Force}}$$

$$\underbrace{\gamma^2 = \frac{1}{G} \left(\frac{Q_1}{M_1} - \frac{Q_2}{M_2} \right)^2}_{\text{Radiation}}$$

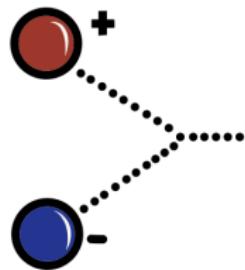
Whence charge?



Particle production



Accretion



Separation

Toy model assumption:

charge is pointlike relative to binary separation

Modified single-source spectrum

New force and radiation modify the spectrum

$$\frac{dE_{\text{GW}}}{df_{\text{GW}}} = -\pi^2 \mu r^2 f_{\text{GW}} \left(\frac{2f_{\text{GW}}}{r} \frac{dr}{df_{\text{GW}}} + 1 \right) \underbrace{\frac{P_{\text{GW}}}{P_{\text{GW}} + P_{\text{new}}}}_{\text{radiation}}$$

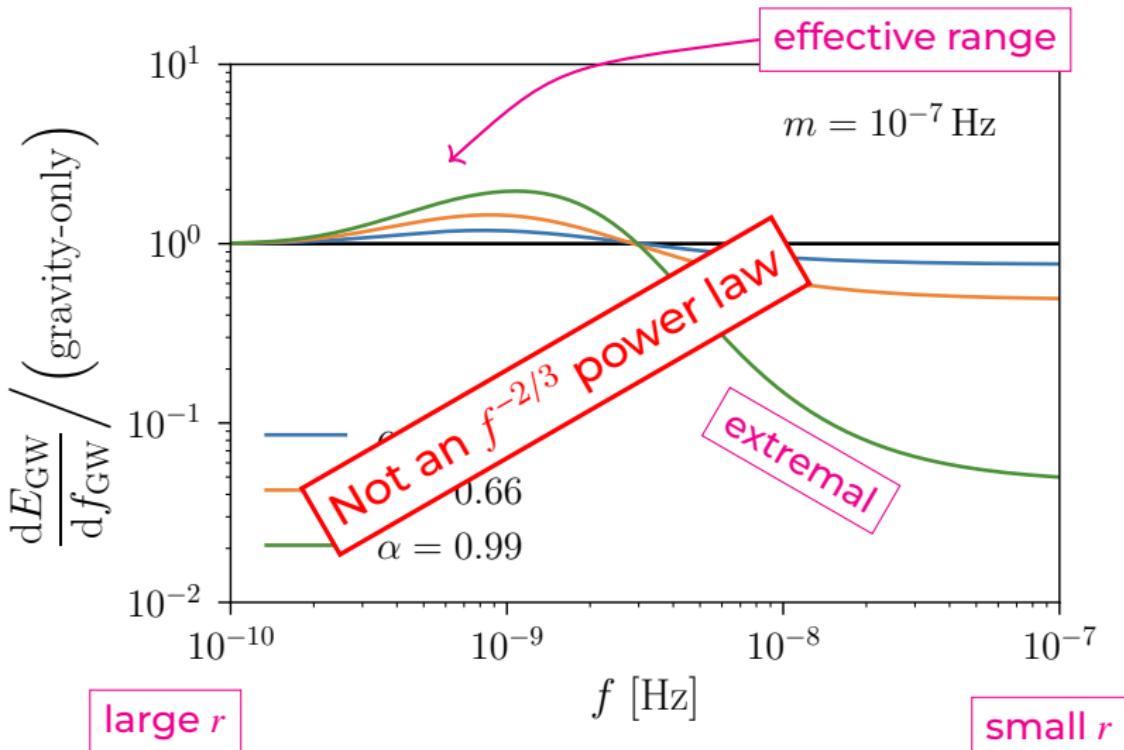
m = mediator mass

$\omega = 2\pi f_{\text{orbit}}$

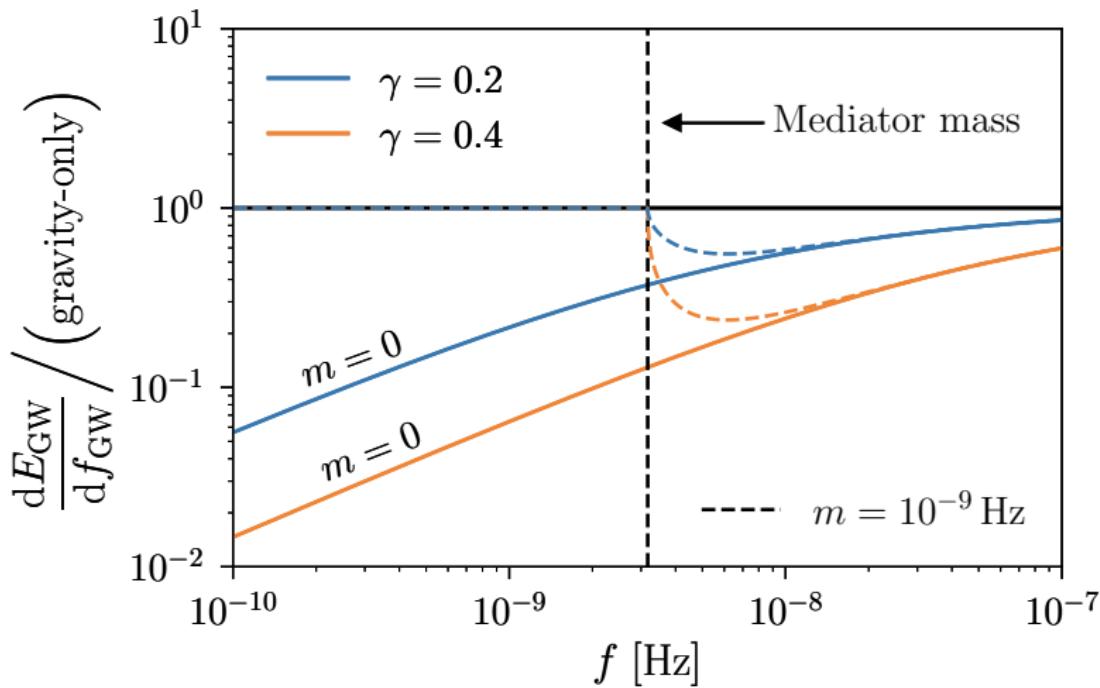
$$F = \frac{GM_1M_2}{r^2} \left(1 - \underbrace{\alpha e^{-mr} (1 + mr)}_{\text{new force}} \right)$$

$$P_{\text{new}} = \frac{1}{3} G \gamma^2 \mu^2 r^2 \omega^4 \operatorname{Re} \left[\sqrt{1 - \frac{m^2}{\omega^2}} \right] \begin{cases} \left(1 - \frac{m^2}{2\omega^2} \right) & (\text{scalar}) \\ 2 \left(1 + \frac{m^2}{2\omega^2} \right) & (\text{vector}) \end{cases}$$

Modifying the force law ($\alpha \neq 0$)

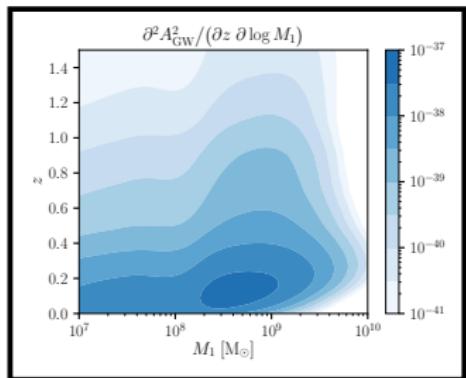


New dipole radiation ($\gamma \neq 0$)

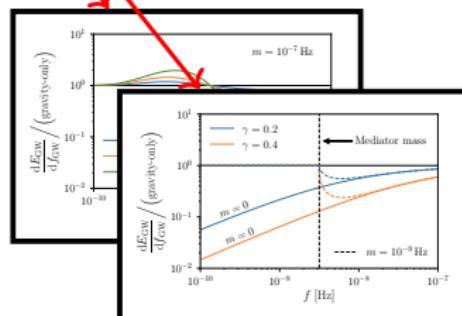


From single sources to h_c

$$h_c^2(f) = \int dz dM_1 dM_2 \frac{dn_G}{dz dM_1 dM_2} \frac{f_s}{1+z} \frac{dE_{\text{GW}}}{df_s} \frac{3H_0^2}{2\pi^2 \rho_c^2}$$



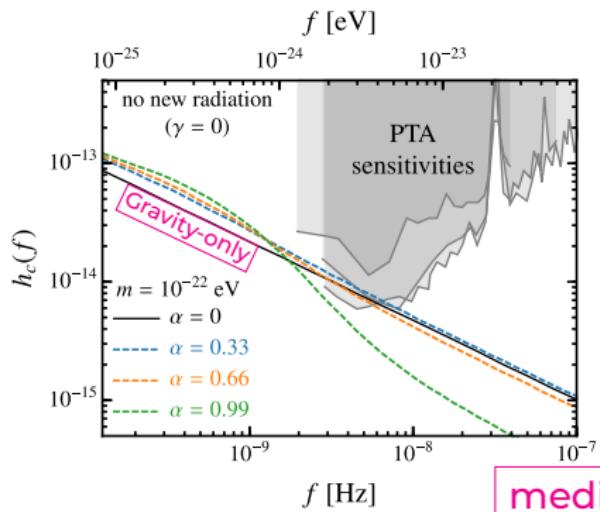
Source distribution



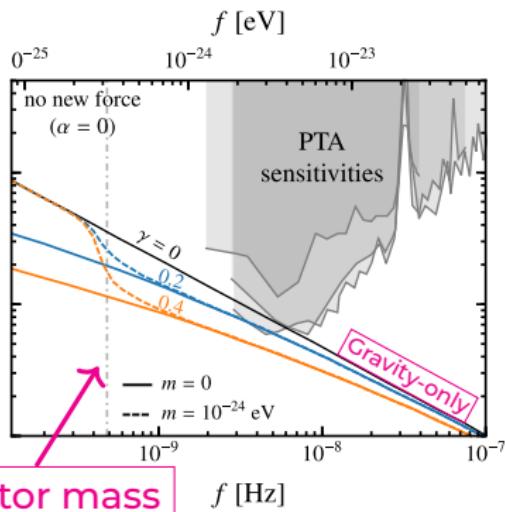
Single-source spectrum

Observables

Force law ($|\alpha| > 0$)



Dipole radiation ($|\gamma| > 0$)

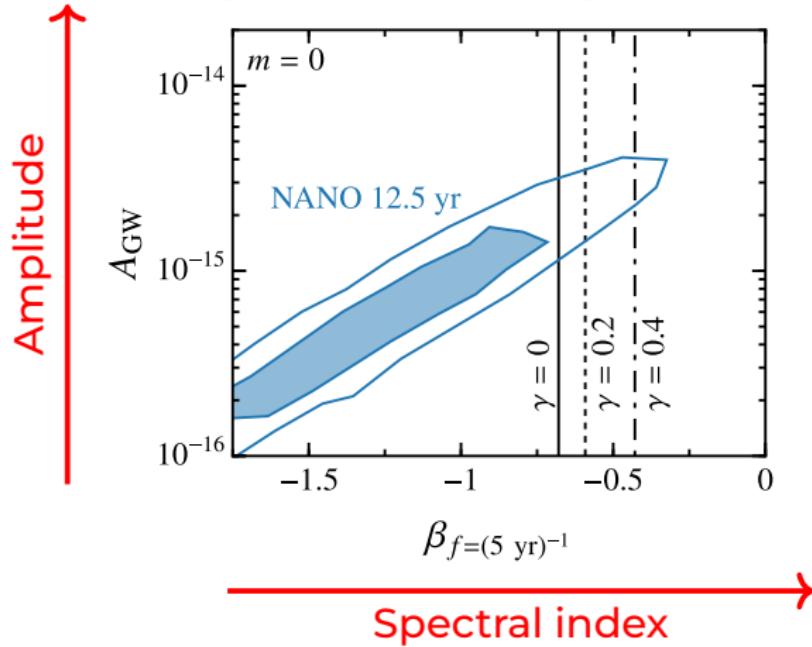


- ① Single-source features are intact
- ② Both modified slope and novel features observable
- ③ Sensitivity curves: this is happening **NOW**

Current data

Interpret the **NANOGrav 12.5-yr result** in this framework

[Arzoumanian et al., 2020]



Beyond the benchmark

Probe **any** new physics that affects binary dynamics

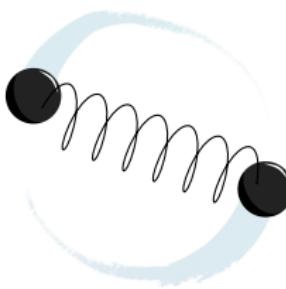
- ① Charged clouds
- ② Dark matter spikes
- ③ Superradiance

Conclusions

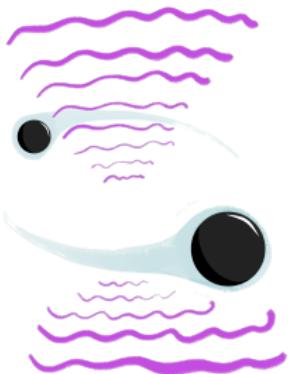
Supermassive black holes are our new laboratories



SGWB discovery
is imminent



Long-range forces
are detectable



SMBHs can probe
many NP scenarios

Data is on the way!

References I

- Z. Arzoumanian et al. The NANOGrav 12.5 yr Data Set: Search for an Isotropic Stochastic Gravitational-wave Background. *Astrophys. J. Lett.*, 905(2):L34, 2020. doi: 10.3847/2041-8213/abd401.
- C. J. Moore, R. H. Cole, and C. P. L. Berry. Gravitational-wave sensitivity curves. *Class. Quant. Grav.*, 32(1):015014, 2015. doi: 10.1088/0264-9381/32/1/015014.
- E. S. Phinney. A Practical theorem on gravitational wave backgrounds. 7 2001.
- A. Sesana. Systematic investigation of the expected gravitational wave signal from supermassive black hole binaries in the pulsar timing band. *Mon. Not. Roy. Astron. Soc.*, 433:1, 2013. doi: 10.1093/mnrasl/slt034.