

New Neutrino Interactions at COHERENT

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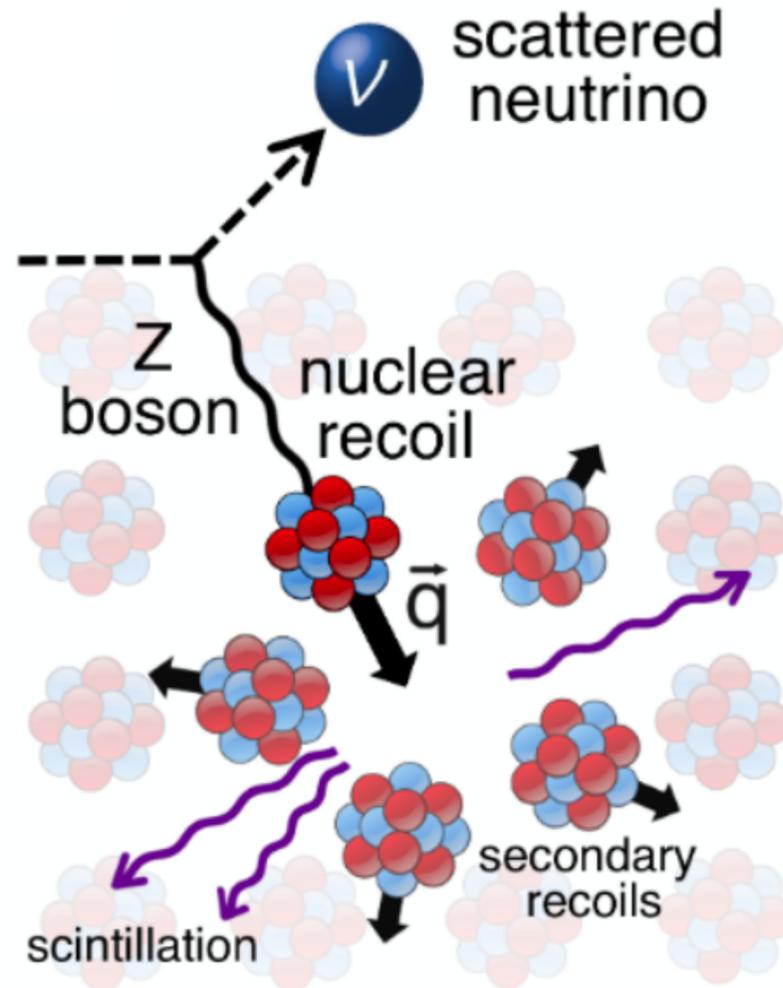


ASTRODARK-2021: Axions, Neutrinos, Black Holes, and
Gravitational Waves

7-10 December 2021, Tokyo

Based on: *JHEP* (2020), arXiv: 2002.12342, L. Flores, NN, E. Peinado,
JHEP (2021), arXiv: 2107.04037, de la Vega, L. Flores, NN, E. Peinado,
arXiv: 2112.0XXXX, L. Flores, NN, E. Peinado

CE ν NS: Coherent Elastic Neutrino-Nucleus Scattering



PHYSICAL REVIEW D

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1 MARCH 1974

Coherent effects of a weak neutral current

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(Received 15 October 1973; revised manuscript received 19 November 1973)

If there is a weak neutral current, then the elastic scattering process $\nu + A \rightarrow \nu + A$ should

$$E_\nu \lesssim 50 \text{ MeV}$$

$$q \cdot R \ll 1$$

3-momentum transfer

Nuclear radius

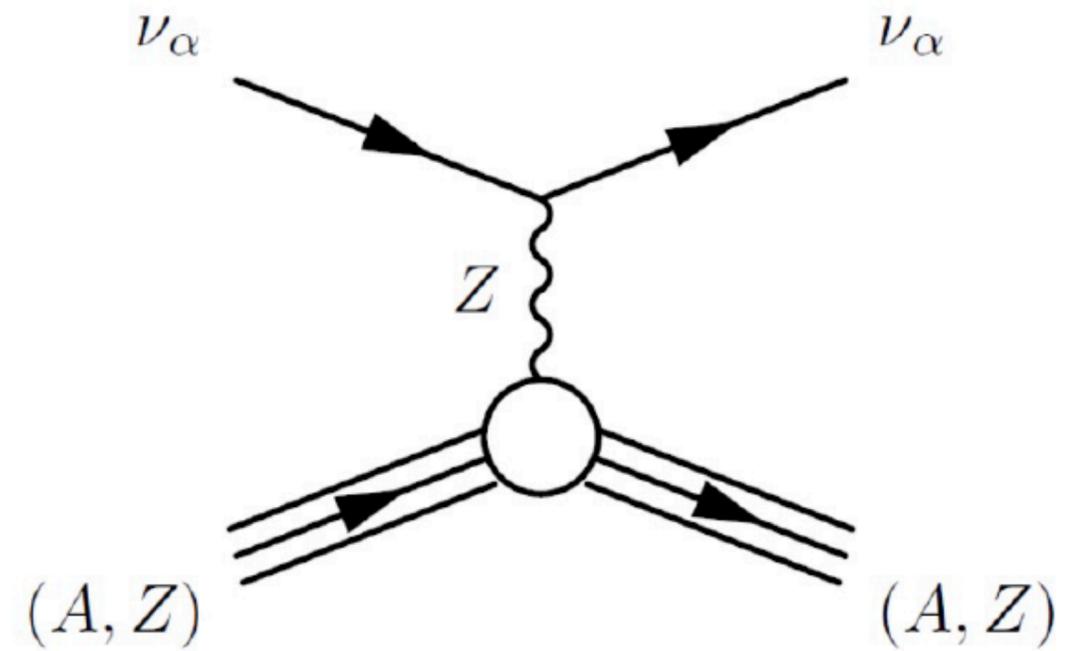
Cross Section:

$$\frac{d\sigma}{dT} = \frac{G_F^2}{2\pi} M_N Q_w^2 \left(2 - \frac{M_N T}{E_\nu^2} \right)$$

Weak Nuclear Charge:

$$Q_w^2 = \left[Z g_p^V F_Z(q^2) + N g_n^V F_N(q^2) \right]^2$$

$$g_p^V = 1/2 - 2 \sin^2 \theta_W, \quad g_n^V = -1/2$$

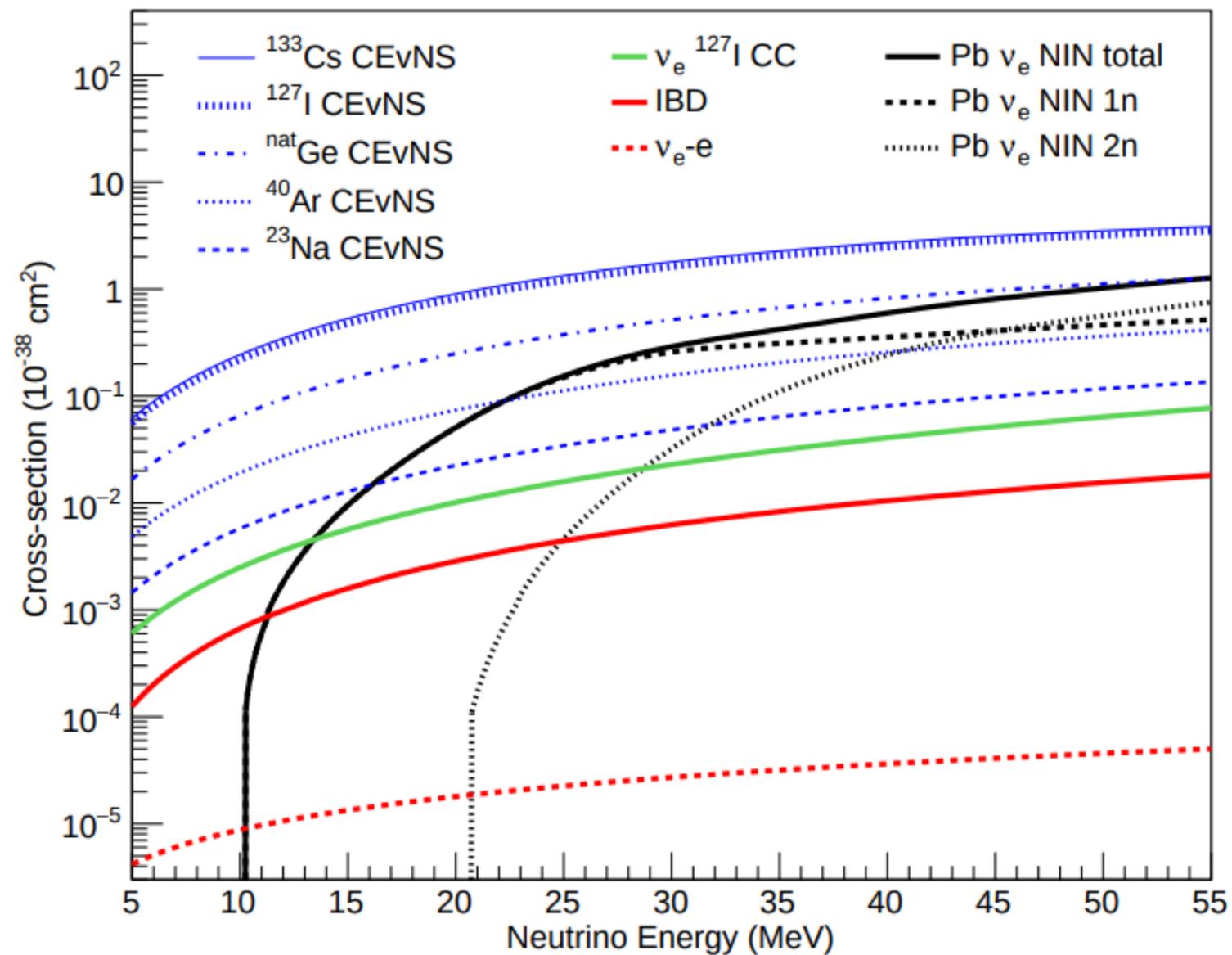


$$\frac{d\sigma}{dT} \propto N^2$$

Cont...

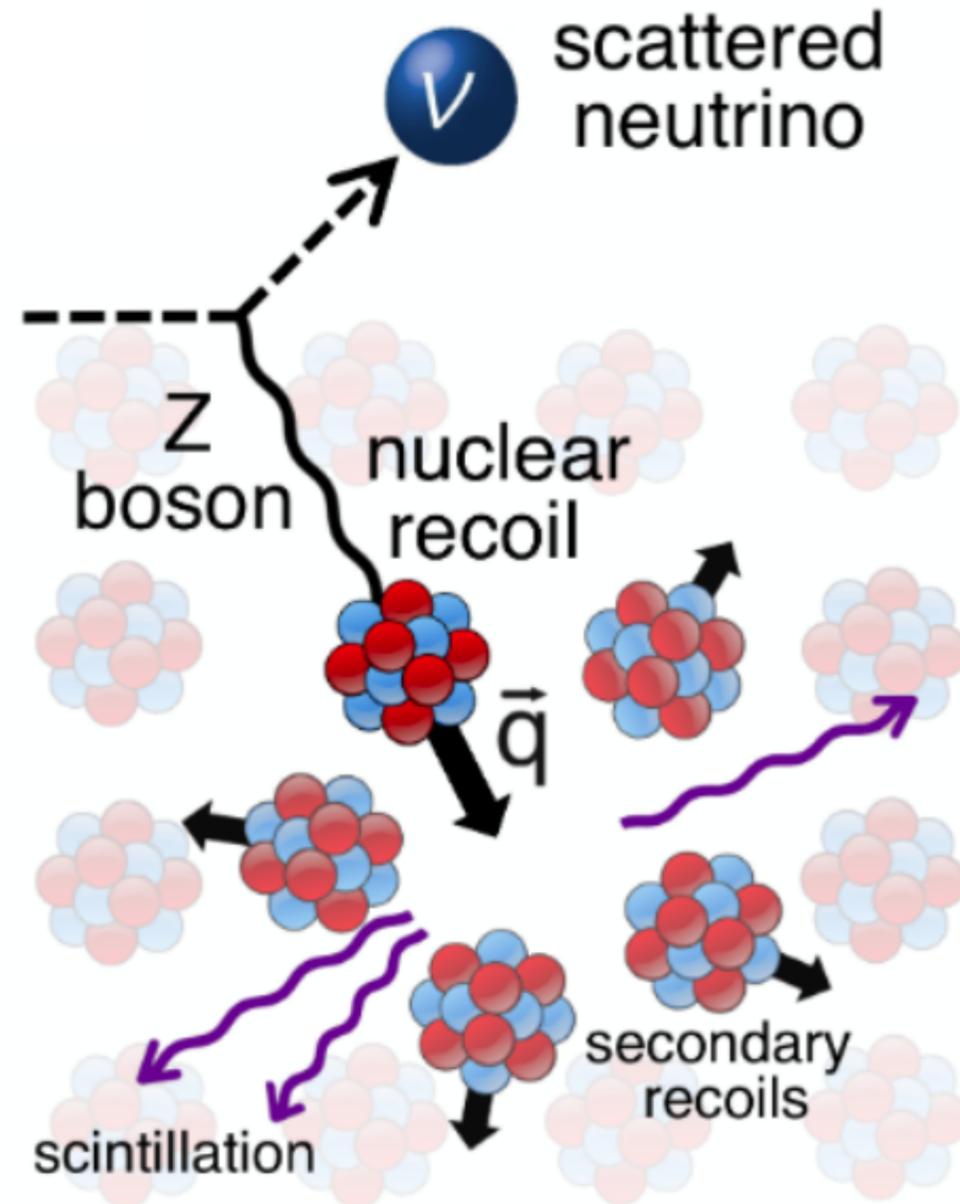
$$\sigma \propto N^2$$

Higher cross-section



COHERENT Collaboration, Science 357,1123 (2017)

Cont...



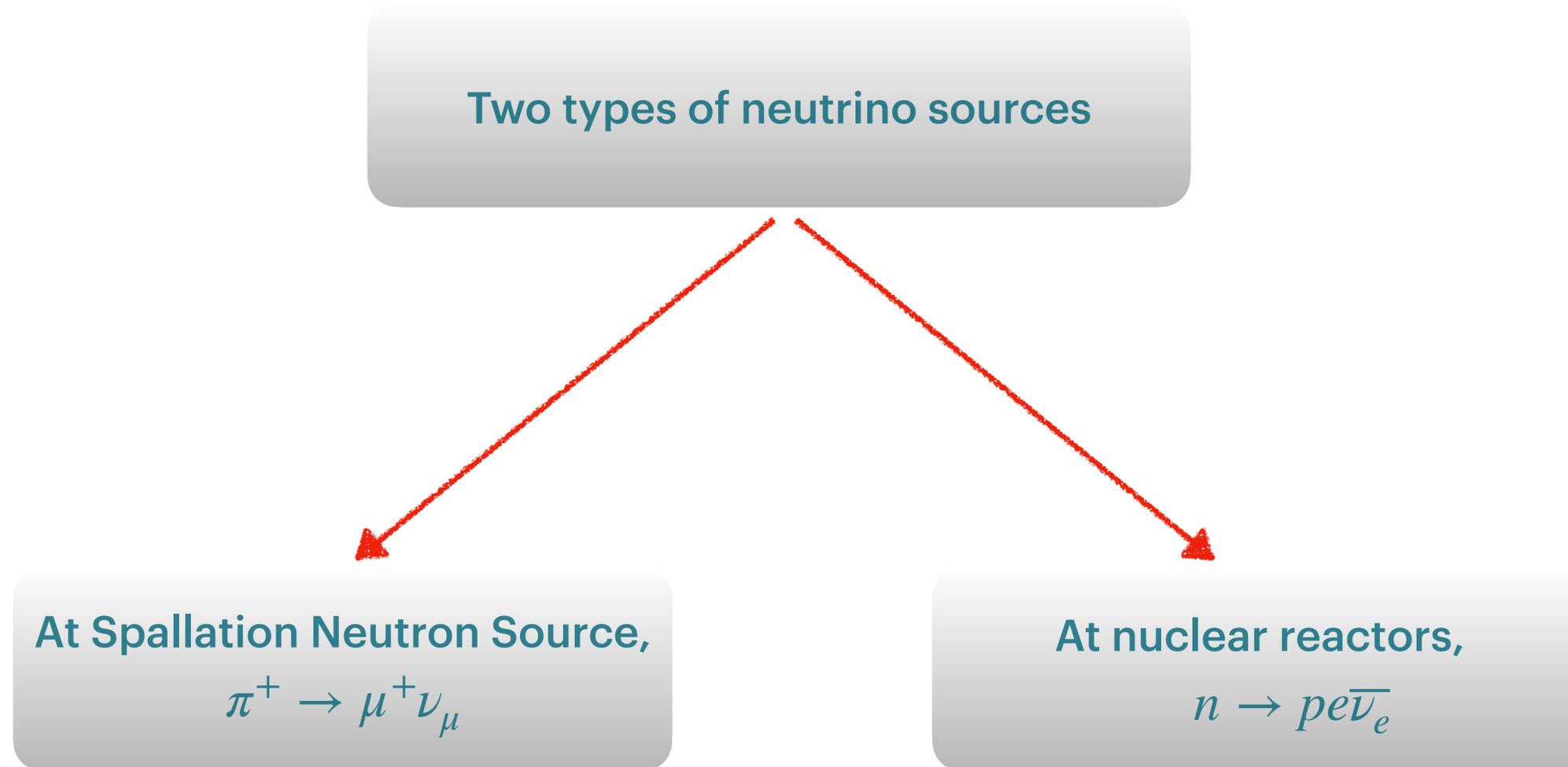
Main challenge to observe nuclear recoils
with a very small kinetic energy of a few keV

$$T_{\max}(E_{\nu}) = \frac{2E_{\nu}^2}{M + 2E_{\nu}}$$

For $E_{\nu} \sim 50\text{MeV}$, $T_{\max}(E_{\nu}) \approx 40\text{keV}$

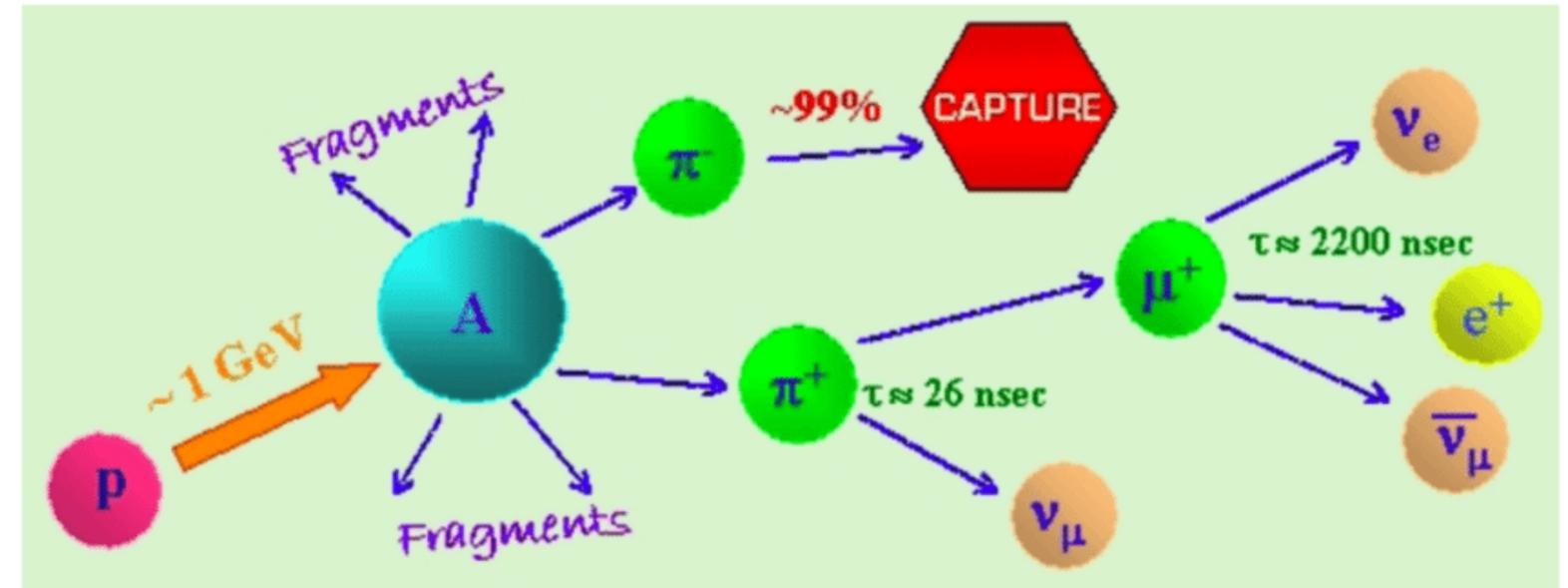
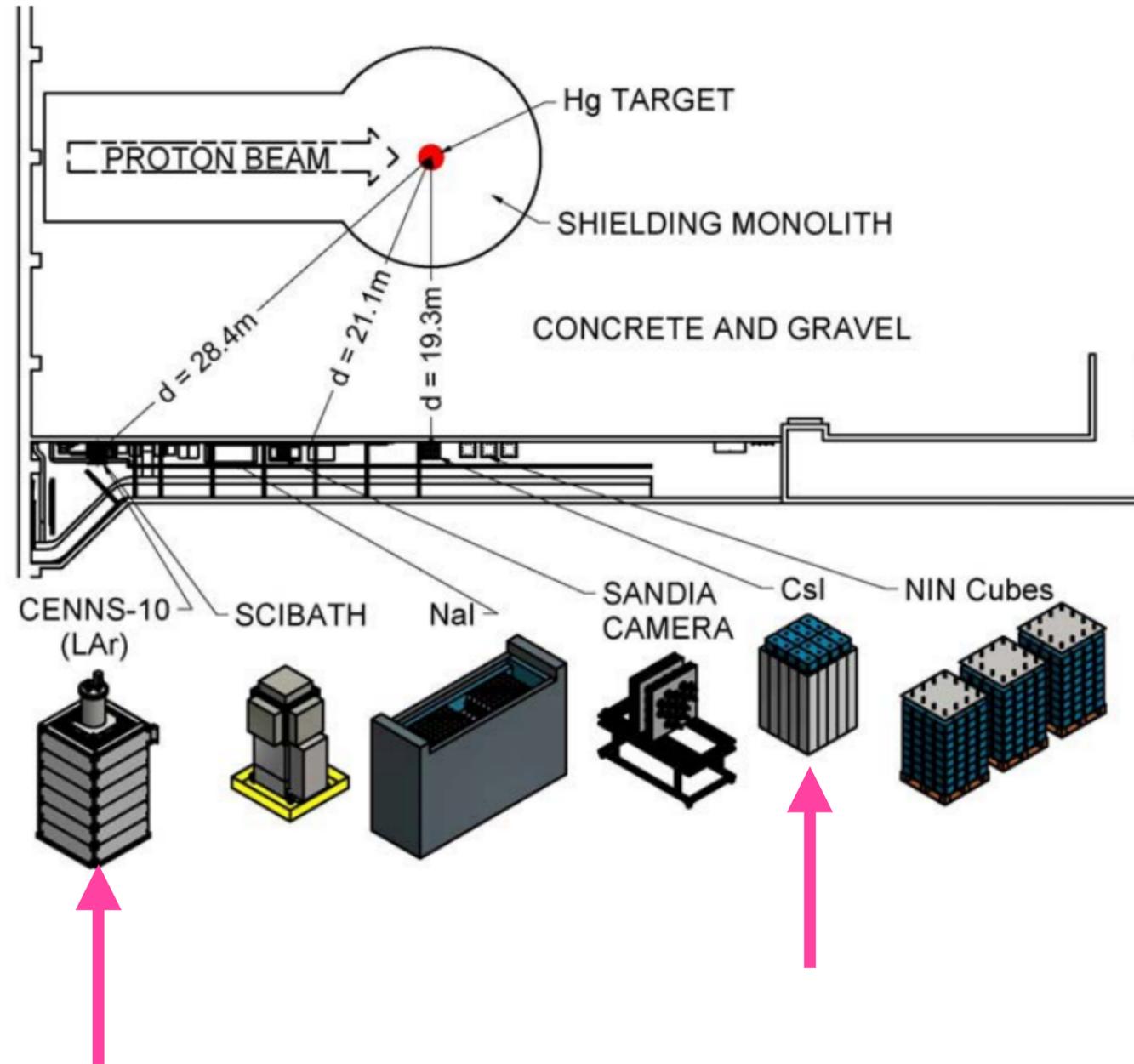
For CE ν NS:

- a low nuclear-recoil-energy threshold detector,
- a low-background environment,
- an intense neutrino flux.



COHERENT Collaboration:

Spallation Neutron Source (SNS) at Oak Ridge National Laboratory (ORNL)



Results:

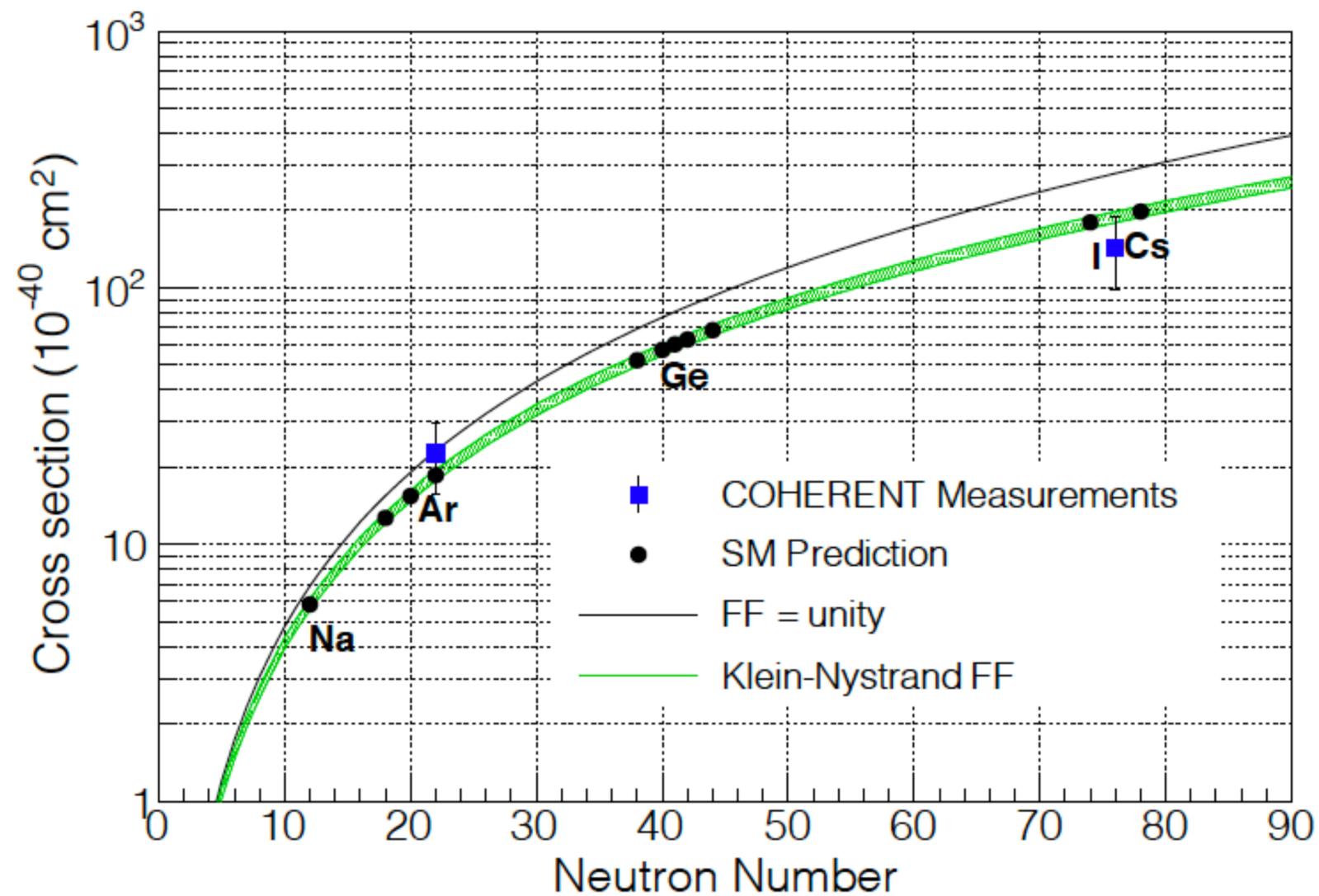
- 134 ± 22 events at 6.7σ CL (Csl)
- 306 ± 20 events at 11.6σ CL (Csl)
- 159 ± 43 and 121 ± 36 events at 3σ CL (LAr)

COHERENT Collaboration, Science 357,1123 (2017),
arXiv: 2110.07730 [hep-ph] , arXiv: 2006.12659 [nucl-ex]

Cont...

$$\sigma \propto N^2$$

Observed cross-section consistent with N^2 dependence

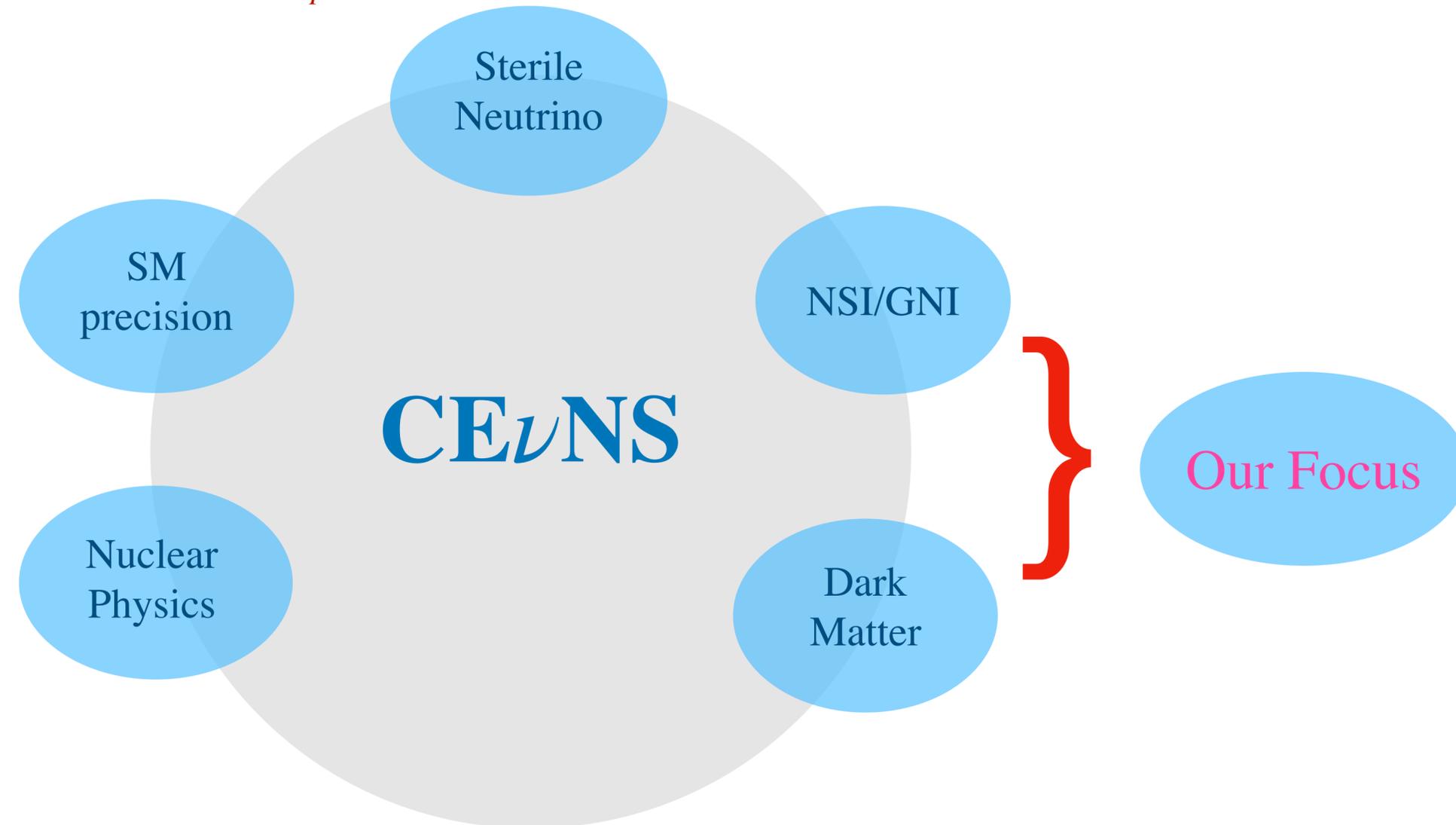


COHERENT Collaboration, arXiv: 2006.12659 [nucl-ex]

Importance: The SM differential cross-section:

$$\frac{d\sigma}{dT} = \frac{G_F^2}{2\pi} M_N \left[Z g_p^V F_Z(q^2) + N g_n^V F_N(q^2) \right]^2 \left(2 - \frac{M_N T}{E_\nu^2} \right) + \dots$$

$$g_p^V = 1/2 - 2 \sin^2 \theta_W, g_n^V = -1/2$$

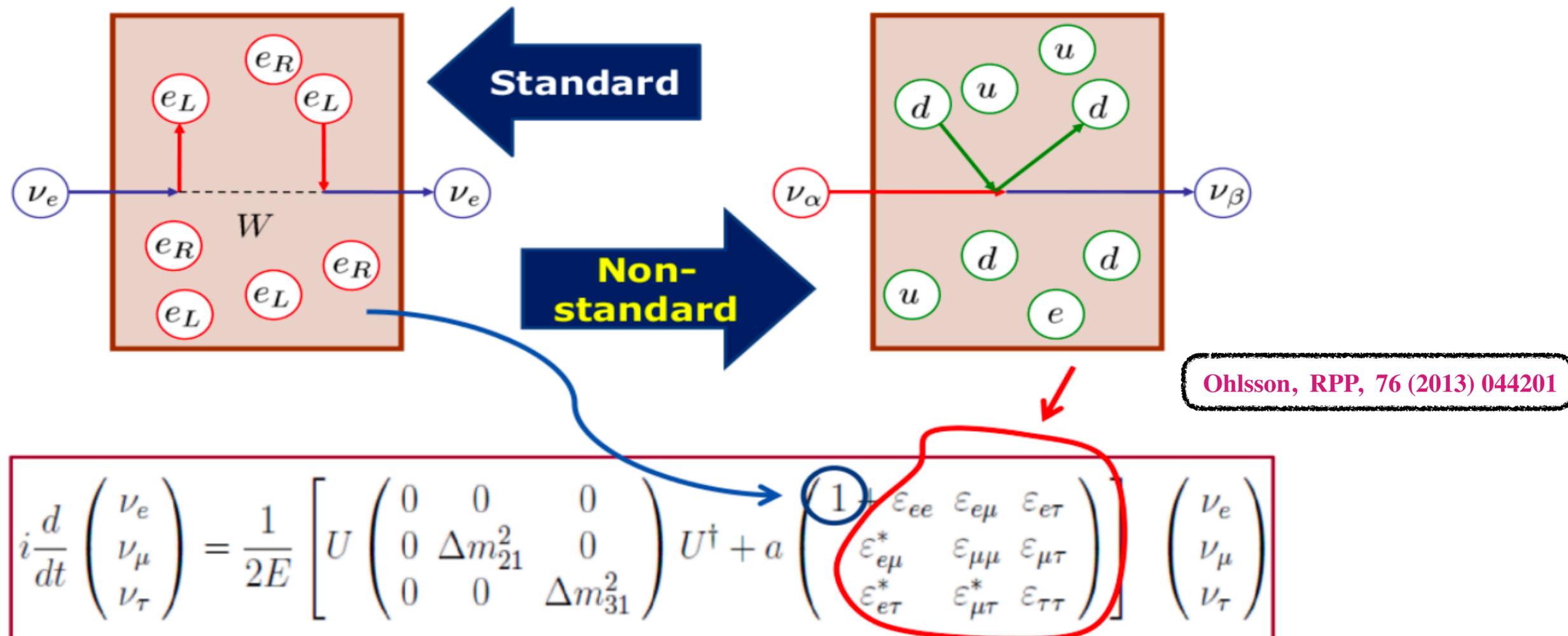


Non-standard interactions

- The effective Lagrangian: $\mathcal{L}^{eff} = \mathcal{L}_{SM} + \frac{1}{\Lambda} \delta \mathcal{L}^{d=5} + \frac{1}{\Lambda^2} \delta \mathcal{L}^{d=6} + \dots$

- The effective dimension-6 term: $\supset (\bar{\nu}_\alpha \gamma^\rho P_L \nu_\beta) (\bar{f} \gamma_\rho P_C f) 2\sqrt{2} G_F \epsilon_{\alpha\beta}^{fC} + h.c.$

[Wolfenstein, '78, Valle '87]



NSI @ CE ν NS:

The SM differential cross-section:

$$\frac{d\sigma}{dT} = \frac{G_F^2}{2\pi} M_N Q_w^2 \left(2 - \frac{M_N T}{E_\nu^2} \right)$$

In SM: $Q_w^2 = [Z g_p^V F_Z(q^2) + N g_n^V F_N(q^2)]^2$

NSI

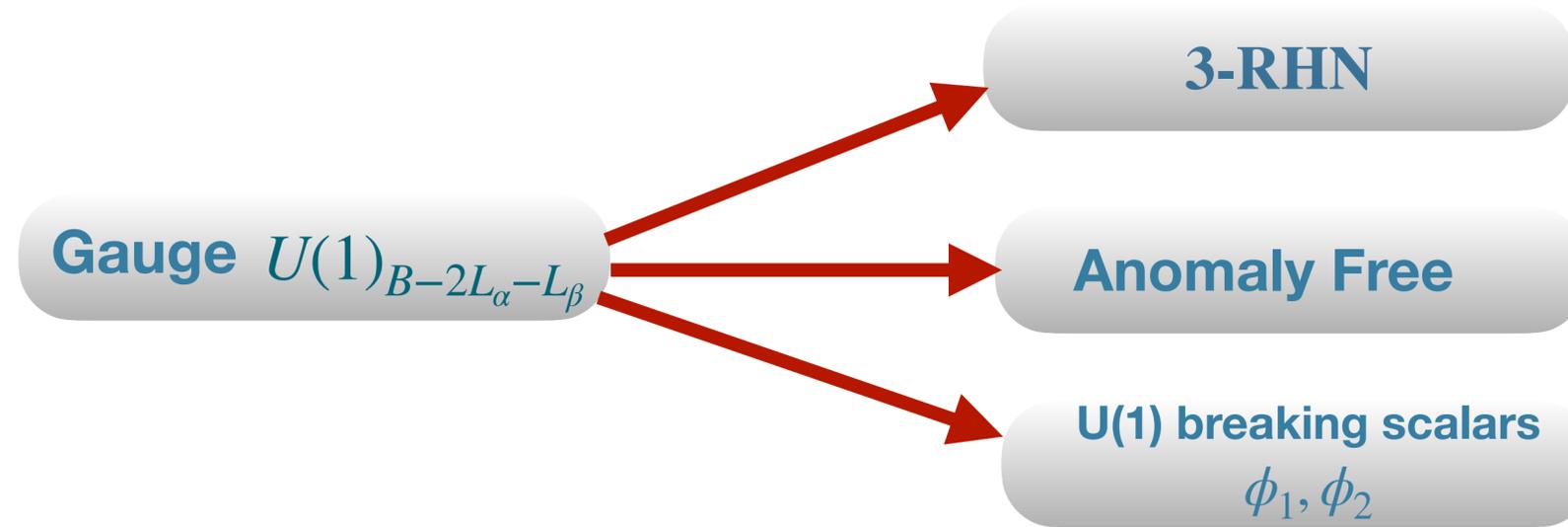
In presence of NSI:

$$Q_{W,\alpha\alpha}^2 = \left[Z(g_p^V + 2\varepsilon_{\alpha\alpha}^u + \varepsilon_{\alpha\alpha}^d) + N(g_n^V + 2\varepsilon_{\alpha\alpha}^d + \varepsilon_{\alpha\alpha}^u) \right]^2,$$

$$Q_{W,\alpha\beta}^2 = \sum_{\beta \neq \alpha} \left| Z(2\varepsilon_{\alpha\beta}^u + \varepsilon_{\alpha\beta}^d) + N(2\varepsilon_{\alpha\beta}^d + \varepsilon_{\alpha\beta}^u) \right|^2.$$

of events:
$$N_i = \int_{E_r^i}^{E_r^{i+1}} A(E_r) \sum_{\alpha} N_{\alpha} \int_{E_{\nu}^{\min}}^{E_{\nu}^{\max}} \phi_{\alpha}(E_{\nu}) \frac{d\sigma}{dE_r} dE_{\nu} dE_r$$

NSI in $U(1)'$ Models:



$$\mathcal{L}_{\text{eff}} = -\frac{g'^2}{Q^2 + M_{Z'}^2} \left[\sum_{\alpha} x_{\alpha} \bar{\nu}_{\alpha} \gamma^{\mu} P_L \nu_{\alpha} \right] \left[\sum_q x_q \bar{q} \gamma_{\mu} q \right]$$

leads to

$$\epsilon_{\alpha\alpha}^{qV} = \frac{g'^2 x_{\alpha} x_q}{\sqrt{2} G_F (Q^2 + M_{Z'}^2)}$$

Flores, NN, Peinado, JHEP 06 (2020) 045

Cont...

x_e	x_μ	x_τ	Neutrino mass matrix	Type	NSI parameters
0	-1	-2	$\begin{pmatrix} 0 & 0 & \times \\ 0 & \times & \times \\ \times & \times & \times \end{pmatrix}$	A_1	$\epsilon_{\mu\mu} \ \& \ \epsilon_{\tau\tau}$
0	-2	-1	$\begin{pmatrix} 0 & \times & 0 \\ \times & \times & \times \\ 0 & \times & \times \end{pmatrix}$	A_2	$\epsilon_{\mu\mu} \ \& \ \epsilon_{\tau\tau}$
-1	0	-2	$\begin{pmatrix} \times & 0 & \times \\ 0 & 0 & \times \\ \times & \times & \times \end{pmatrix}$	B_3	$\epsilon_{ee} \ \& \ \epsilon_{\tau\tau}$
-1	-2	0	$\begin{pmatrix} \times & \times & 0 \\ \times & \times & \times \\ 0 & \times & 0 \end{pmatrix}$	B_4	$\epsilon_{ee} \ \& \ \epsilon_{\mu\mu}$

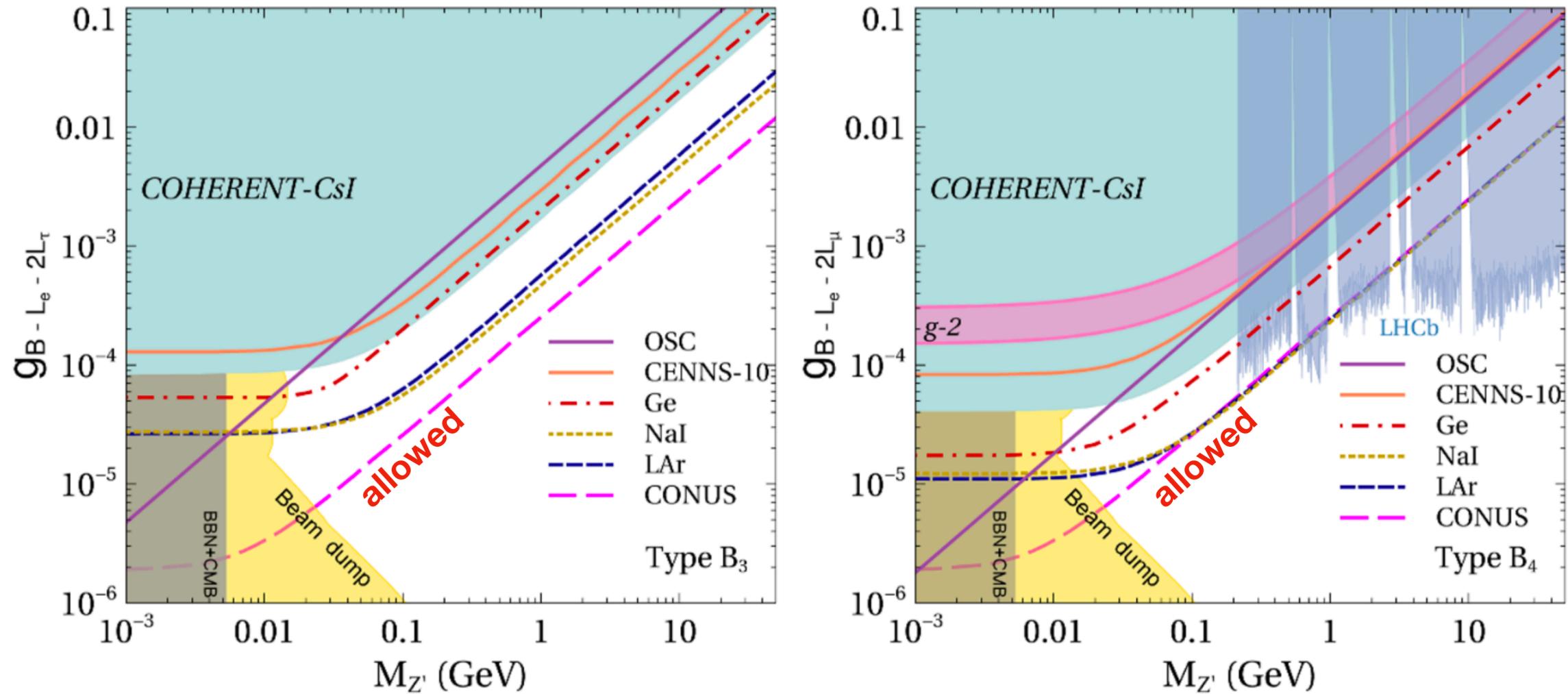
Consistent with neutrino phenomenology

Flores, NN, Peinado, JHEP 06 (2020) 045

Cont...

$$\frac{1}{2}M_{Z'}^2 = g'^2 \frac{1}{2}(v_1^2 + 4v_2^2)$$

$$M_{Z'} = 0.1 \text{ GeV} \quad g' = 2.8 \times 10^{-5} \quad v_1 \approx 3 \text{ TeV} \quad v_2 \approx 1 \text{ TeV}$$



Flores, NN, Peinado, JHEP 06 (2020) 045

CEvNS and DM searches in $U(1)'$ models

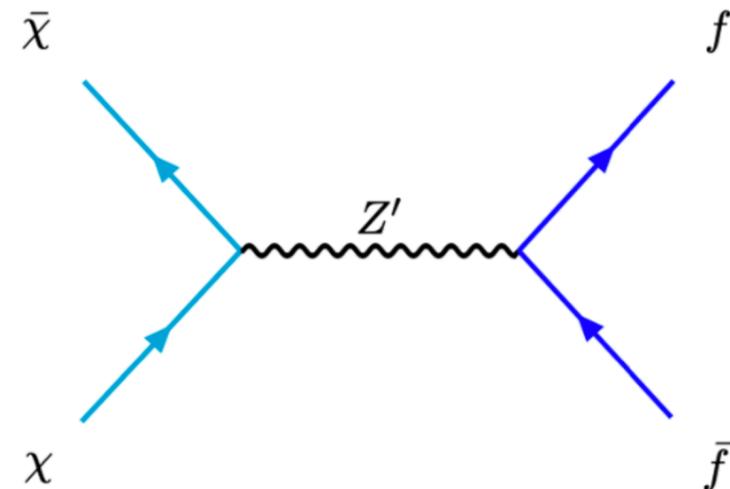
Extend the SM with a Dirac fermion χ with $Q_\chi = 1/3$

de la Vega, Flores, NN, Peinado, JHEP 09 (2021) 146

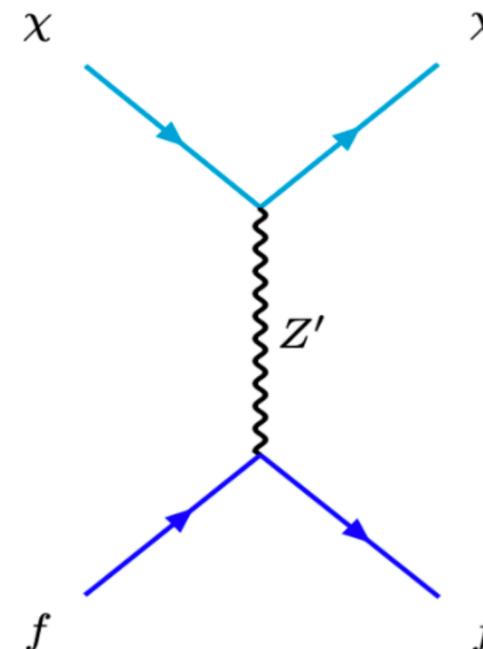
Z_3 symmetry stabilises χ

Resonant annihilation
 $M_\chi \simeq M_{Z'}/2$

$$\mathcal{L} \supset M_D \bar{\chi} \chi + \bar{\chi} \gamma^\mu (\partial_\mu + ig' Q_\chi Z'_\mu) \chi$$



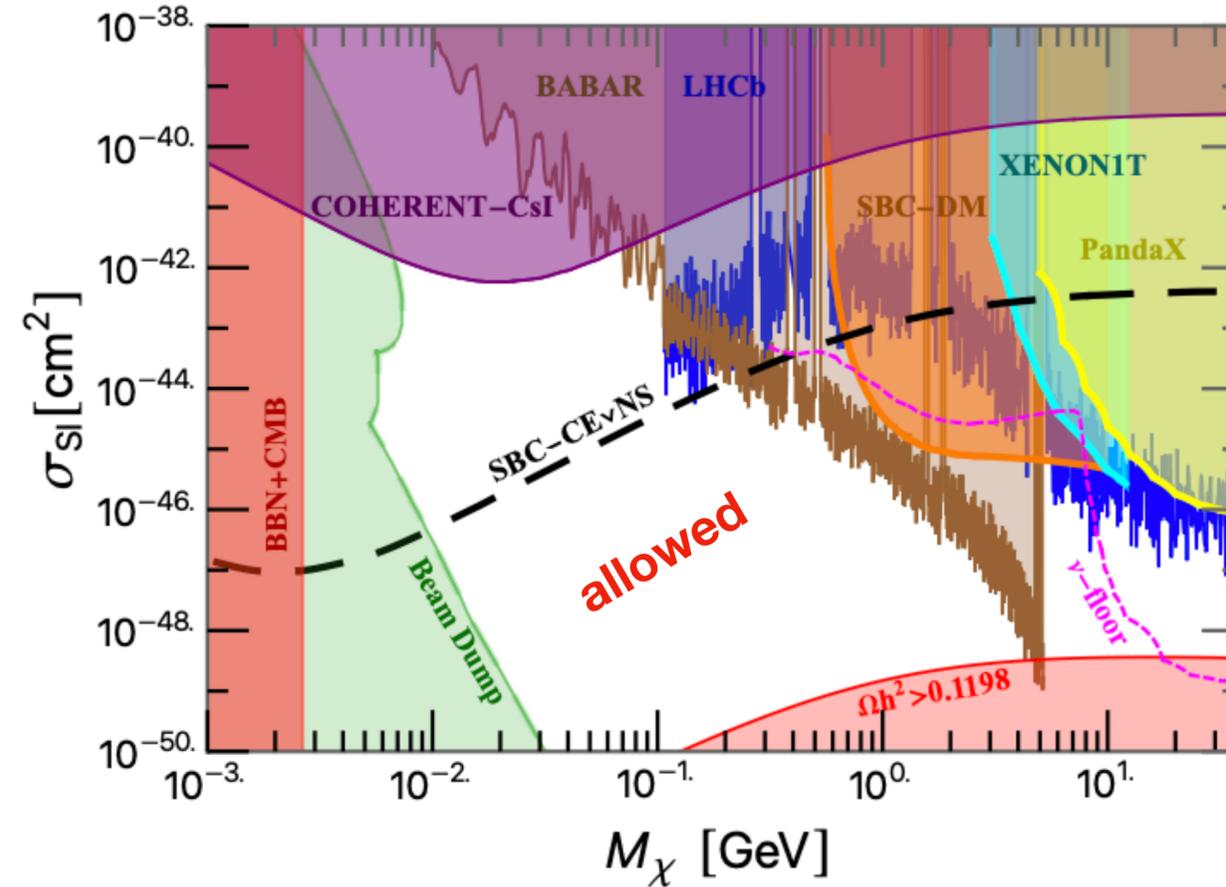
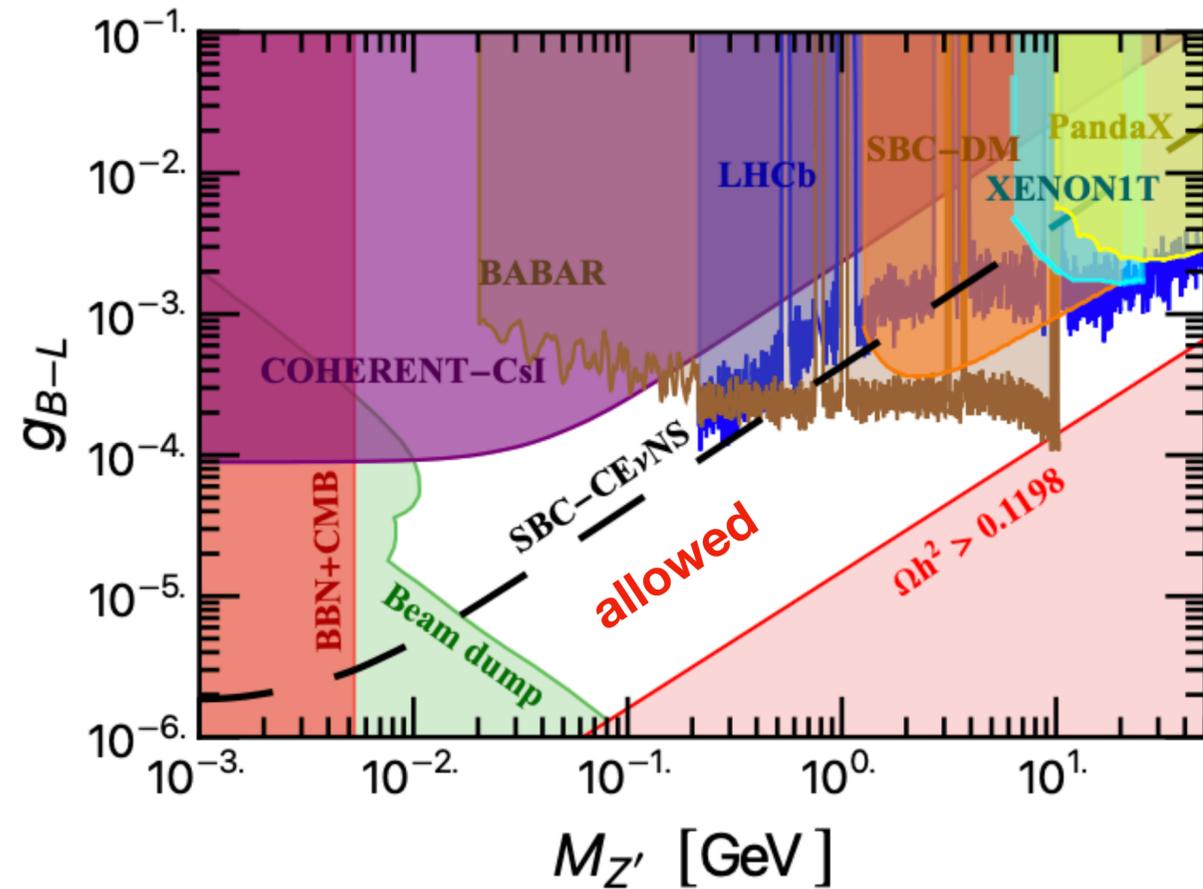
DM annihilation
freeze-out



DM
Direct-detection

Cont...

The dark matter-nucleus SI cross-section per nucleon: $\sigma_{\text{SI}} \approx \frac{\mu_{\chi n}^2}{\pi} \frac{g'^4}{9M_{Z'}^4}$.



SBC: Scintillating Bubble Chamber, a reactor-based CE ν NS experiment

de la Vega, Flores, NN, Peinado, JHEP 09 (2021) 146

Generalized Neutrino Interactions (GNI):

$$\mathcal{L}_{\text{NSI}} \supset 2\sqrt{2}G_F\epsilon_{\alpha\beta}^{fC} (\bar{\nu}_\alpha\gamma^\mu P_L\nu_\beta)(\bar{f}\gamma_\mu P_C f') \quad \text{Bergmann, Grossman, Nardi, PRD 60, 093008 (1999)}$$

j	$\tilde{\epsilon}^j$	\mathcal{O}_j	\mathcal{O}'_j
1	ϵ_L	$\gamma_\mu(1 - \gamma^5)$	$\gamma^\mu(1 - \gamma^5)$
2	$\tilde{\epsilon}_L$	$\gamma_\mu(1 + \gamma^5)$	$\gamma^\mu(1 - \gamma^5)$
3	ϵ_R	$\gamma_\mu(1 - \gamma^5)$	$\gamma^\mu(1 + \gamma^5)$
4	$\tilde{\epsilon}_R$	$\gamma_\mu(1 + \gamma^5)$	$\gamma^\mu(1 + \gamma^5)$
5	ϵ_S	$(1 - \gamma^5)$	1
6	$\tilde{\epsilon}_S$	$(1 + \gamma^5)$	1
7	$-\epsilon_P$	$(1 - \gamma^5)$	γ^5
8	$-\tilde{\epsilon}_P$	$(1 + \gamma^5)$	γ^5
9	ϵ_T	$\sigma_{\mu\nu}(1 - \gamma^5)$	$\sigma^{\mu\nu}(1 - \gamma^5)$
10	$\tilde{\epsilon}_T$	$\sigma_{\mu\nu}(1 + \gamma^5)$	$\sigma^{\mu\nu}(1 + \gamma^5)$

Another parametrization:

$$\mathcal{L}_{\text{GNI}} \supset \frac{G_F}{\sqrt{2}} \sum_{a=S,P,V,A,T} \bar{\nu} \Gamma^a \nu \left[\bar{q} \Gamma^a (C_a^{(q)} + \bar{D}_a^{(q)} i\gamma^5) q \right]; \Gamma^a = \{I, i\gamma^5, \gamma^\mu, \gamma^\mu\gamma^5, \sigma^{\mu\nu}\}; \sigma^{\mu\nu} \equiv \frac{i}{2}[\gamma^\mu, \gamma^\nu]$$

Generalized Neutrino Interactions (GNI):

$$\mathcal{L}_{\text{GNI}} \supset \frac{G_F}{\sqrt{2}} \sum_{j=1}^{10} \tilde{\varepsilon}_{\alpha\beta\gamma\delta}^{q,j} (\bar{\nu}_\alpha \mathcal{O}_j \nu_\beta) (\bar{q}_\gamma \mathcal{O}'_j q_\delta)$$

Bergmann, Grossman, Nardi, PRD 60, 093008 (1999)

j	$\tilde{\varepsilon}^j$	\mathcal{O}_j	\mathcal{O}'_j
1	ε_L	$\gamma_\mu(1 - \gamma^5)$	$\gamma^\mu(1 - \gamma^5)$
2	$\tilde{\varepsilon}_L$	$\gamma_\mu(1 + \gamma^5)$	$\gamma^\mu(1 - \gamma^5)$
3	ε_R	$\gamma_\mu(1 - \gamma^5)$	$\gamma^\mu(1 + \gamma^5)$
4	$\tilde{\varepsilon}_R$	$\gamma_\mu(1 + \gamma^5)$	$\gamma^\mu(1 + \gamma^5)$
5	ε_S	$(1 - \gamma^5)$	1
6	$\tilde{\varepsilon}_S$	$(1 + \gamma^5)$	1
7	$-\varepsilon_P$	$(1 - \gamma^5)$	γ^5
8	$-\tilde{\varepsilon}_P$	$(1 + \gamma^5)$	γ^5
9	ε_T	$\sigma_{\mu\nu}(1 - \gamma^5)$	$\sigma^{\mu\nu}(1 - \gamma^5)$
10	$\tilde{\varepsilon}_T$	$\sigma_{\mu\nu}(1 + \gamma^5)$	$\sigma^{\mu\nu}(1 + \gamma^5)$

} NSI

Another parametrization:

$$\mathcal{L}_{\text{GNI}} \supset \frac{G_F}{\sqrt{2}} \sum_{a=S,P,V,A,T} \bar{\nu} \Gamma^a \nu \left[\bar{q} \Gamma^a (C_a^{(q)} + \bar{D}_a^{(q)} i\gamma^5) q \right]; \Gamma^a = \{I, i\gamma^5, \gamma^\mu, \gamma^\mu \gamma^5, \sigma^{\mu\nu}\}; \sigma^{\mu\nu} \equiv \frac{i}{2} [\gamma^\mu, \gamma^\nu]$$

GNI @ CE ν NS:

- Differential cross-section:

$$\left(\frac{d\sigma}{dE_r}\right)^f = \frac{G_F^2}{4\pi} M_N N^2 F^2(Q^2) \times \left[\xi_S^{f2} \frac{E_r}{E_r^{\max}} + (\xi_V^f + A_{\text{SM}})^2 \left(1 - \frac{E_r}{E_r^{\max}} - \frac{E_r}{E_\nu}\right) \pm \right. \\ \left. 2\xi_V^f \xi_A^f \frac{E_r}{E_\nu} + \xi_A^{f2} \left(1 + \frac{E_r}{E_r^{\max}} - \frac{E_r}{E_\nu}\right) + \xi_T^{f2} \left(1 - \frac{E_r}{2E_r^{\max}} - \frac{E_r}{E_\nu}\right) \mp R \frac{E_r}{E_\nu} + \mathcal{O}\left(\frac{E_r^2}{E_\nu^2}\right) \right]$$

Lindner, Werner Rodejohann, Xu, *JHEP* 03 (2017) 097
Aristizabal Sierra, De Romeri, Rojas, *PRD* 98 (2018) 075018

where $A_{\text{SM}} = 1 - (1 - 4 \sin^2 \theta_w) Z/N$, and

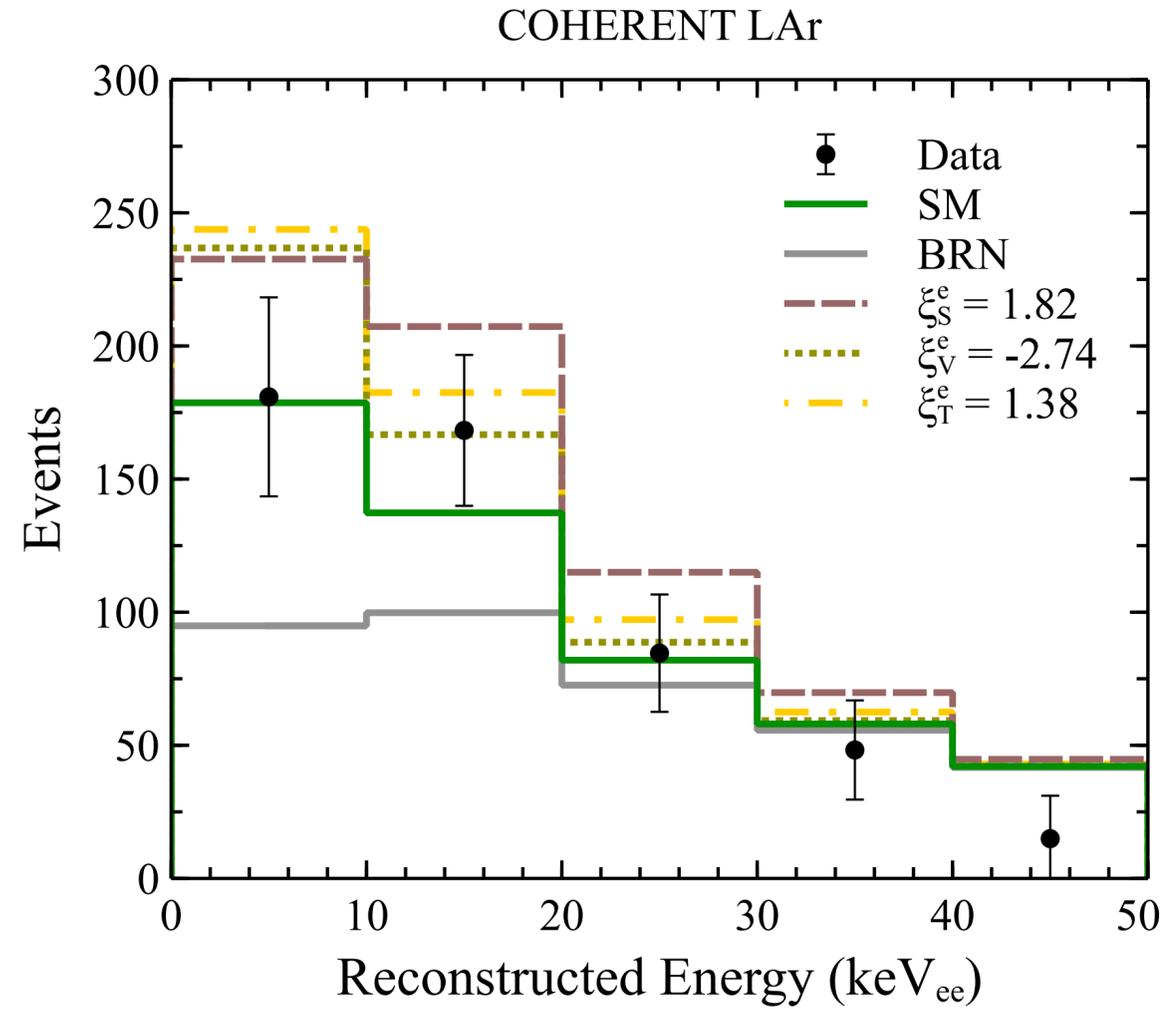
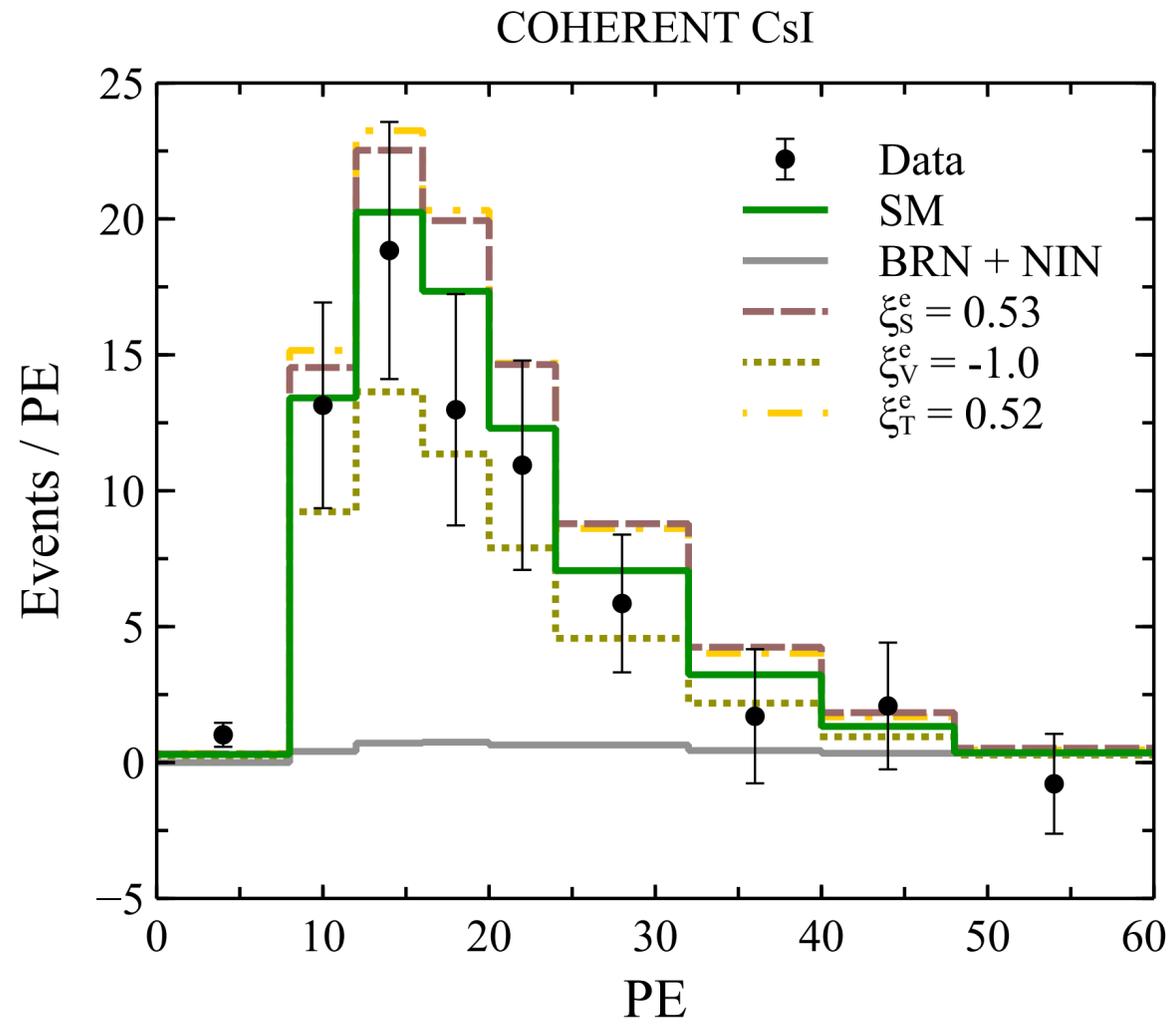
$$\xi_S^{f2} = \frac{1}{N^2} (C_S^2 + D_P^2), \quad \xi_V^f = \frac{1}{N} (C_V + D_A), \\ \xi_A^f = \frac{1}{N} (C_A + D_V), \quad \xi_T^{f2} = \frac{8}{N^2} (C_T^2 + D_T^2), \\ R = \frac{2}{N^2} (C_S C_T - C_P C_T + D_S D_T - D_P D_T).$$

- # of events:

$$N_i = \int_{E_r^i}^{E_r^{i+1}} A(E_r) \sum_{\alpha} N_{\alpha} \int_{E_{\nu}^{\min}}^{E_{\nu}^{\max}} \phi_{\alpha}(E_{\nu}) \frac{d\sigma}{dE_r} dE_{\nu} dE_r$$

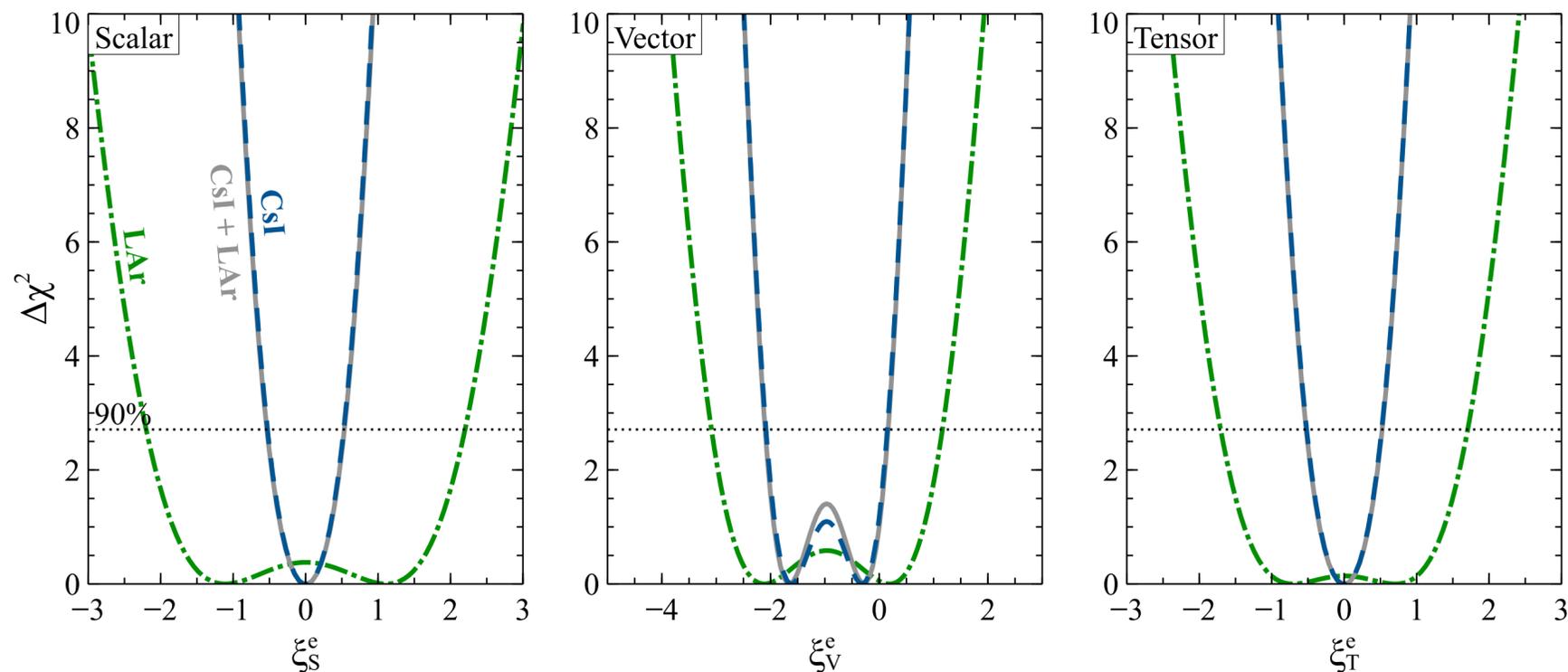
where $\alpha = \nu_{\mu}, \bar{\nu}_{\mu}, \nu_e$

• # of events:

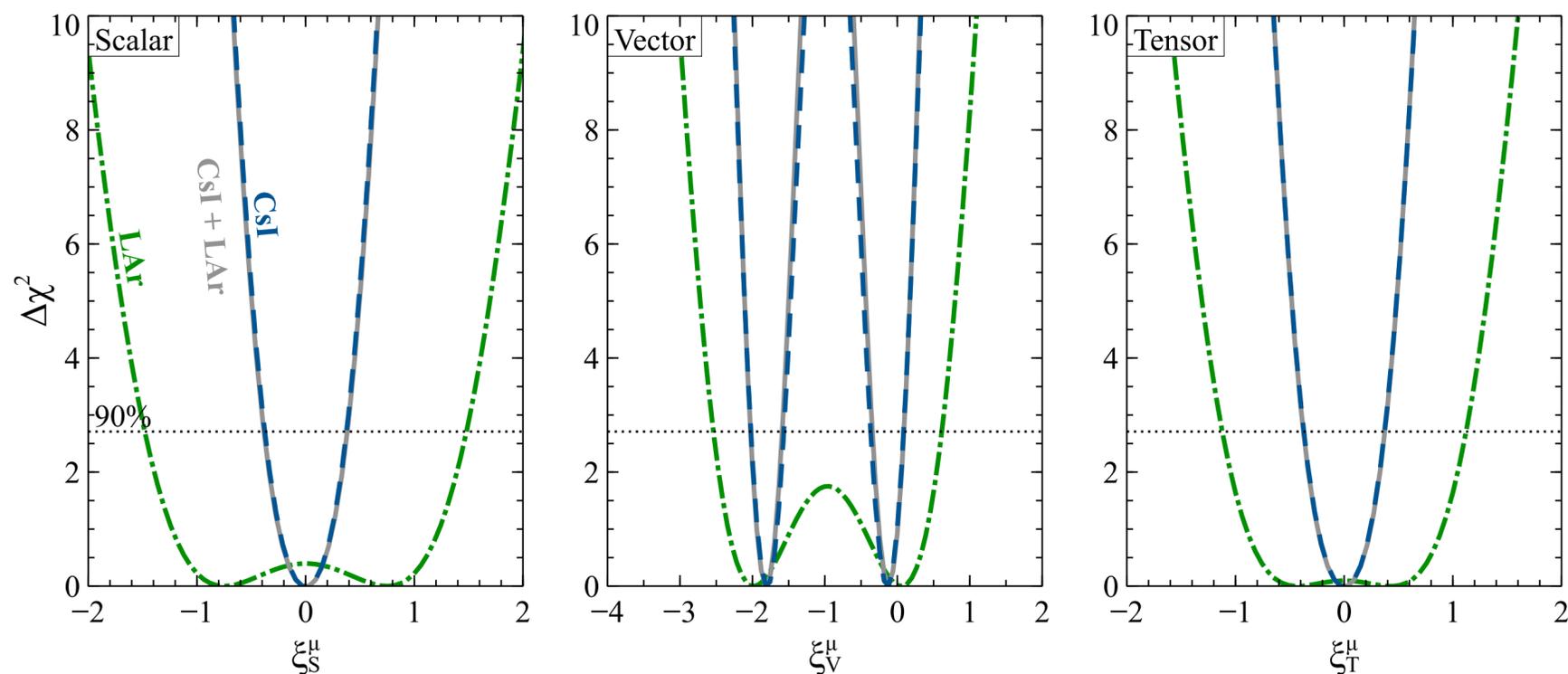


Flores, NN, Peinado, (in preparation)

One-parameter analysis:



$$\frac{d\sigma}{dE_r} \propto \xi_S^2 \frac{E_r}{E_r^{\max}} + (\xi_V + A_{\text{SM}})^2 \left(1 - \frac{E_r}{E_r^{\max}} - \frac{E_r}{E_\nu} \right) + \xi_T^2 \left(1 - \frac{E_r}{2E_r^{\max}} - \frac{E_r}{E_\nu} \right)$$

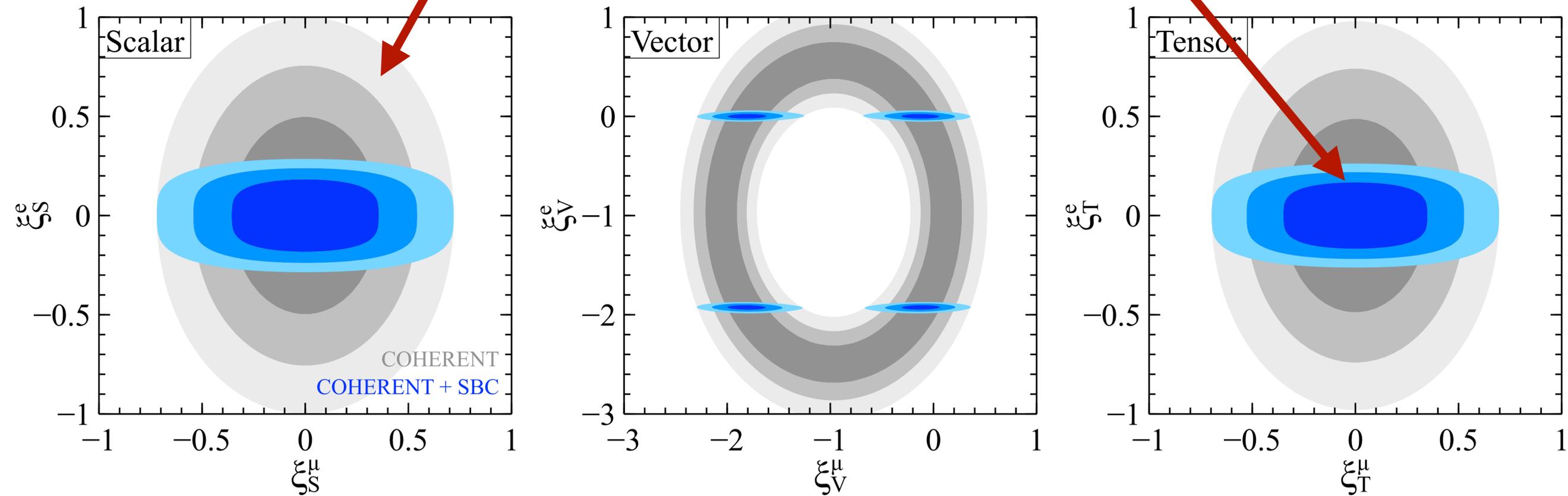


- ξ_e 's are less constrained than ξ_μ 's given the lower ν_e flux at the SNS
- Limits from CsI measurement are better than the LAr

Two-parameter analysis:

1, 2, 3 σ
COHERENT

1, 2, 3 σ
COHERENT+SBC



Note: combined analysis of COHERENT+SBC significantly constrained different ξ_e 's

Final Comments:

- NSI coming from $U(1)'$ models have been studied for the COHERENT data
- Complementarity between dark-matter direct searches and CE ν NS experiments has been explored
- Combined analysis using COHERENT and reactor data has been performed to constrain GNI parameters

