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Innermost structure of dark matter halos finding a clue of dark matter

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Its properties reflect nature of dark matter (DM)

ure & inner structures sometimes exhibit universal features \rightarrow a unique channel to access nature of DM

Cold DM (CDM)

Warm DM

Self-Interacting DM







Here, we focus on "universal" properties in CDM & FDM

Enomoto, Nishimichi & AT, arXiv:2302.01531 AT & Saga, arXiv:2208.06562 (in PRD)



Cold dark matter (CDM) halo

Baseline DM in the concordant cosmological model (ACDM)

Cuspy structure

Radial density p (Navarro, Frenk & White '96)

CDM

Pseudo-phase space density

(Taylor & Navarro '01)

A more profound & universal property as a distinct feature of CDM?

Studied extensively by N-body simulations

orofile
$$\rho(r) \propto r^{\alpha} r^{-\alpha}$$
 ($\alpha = 1 - 1.5$)

(c.f. prompt cusp of $\rho \propto r^{-3/2}$ of first halos) (Ishiyama et al. '10; Delos & White'22)

 $Q(r) \equiv \rho(r) / \{\sigma_{\rm v}(r)\}^3$ $\propto r^{-\alpha_Q} \quad (\alpha_O = 1.875)$







nature of CDM halos

- rovides a distinctive feature in CDM halos
- persion at an early time
 - Through accretion/merger processes

Onion-like structure

- Multi-stream structures with an outer sharp boundary (=Splashback radius) (e.g., Diemer & Kravtsov '14; Adhikari et al. '14)
 - A fundamental universal feature may be hidden in phase space?









Using 1,001 snapshot data of cosmological N-body simulations over z=0-5

Keep track of apocenter passage(s) for particle trajectories (Sugiura et al. '20) and count the number of <u>apocenter passages</u>, *p*, for each particle See also Diemer ('17)





 $\Lambda \text{CDM}, L_{\text{box}} = 41 \, h^{-1} \text{Mpc} \& N_{\text{dm}} = 500^3$



φ





Stacked multi-stream radial profiles Enomoto, Nishimichi & AT ('23) $[5.71, 24.2] \times 10^{11} h^{-1} M.$ $[24.2, 132] \times 10^{1} h^{-1} M_{\odot}$ p = 40p = 40p = 30p = 30p = 20p = 20p = 10 10^{-1} $M_{10} \equiv M_{\rm vir} / (10^{10} h^{-1} M_{\odot})$ -0.11910gm/M $-3.89 + 0.243 \log_{10}(M_{\odot})$ $= 2.46 - 0.0474 \log_{10}(M_{10}) + \{-2.29 - 0.0639 \log_{10}(M_{10})\} p^{1/8}$ $\log_{10}{S(p)}$





Fuzzy dark matter (FDM) halo

Flat dense core = soliton

Mass: M_c

Radius: $r_{\rm c}$

Fuzzy DM

Core-halo relations

 \rightarrow unique feature to constrain mass of fuzzy DM (e.g., Safarzadeh & Spergel '20; Hayashi & Obata '20)

However,

some contradiction as well as a diversity of their relations reported (Schwabe et al. '16, Du et al. '17, Mocz et al. '17, Nori & Baldi '21, Mina et al. '22, Chan et al. '22)

Alternative DM candidate having a ultralight mass ($\sim 10^{-22}$ eV)

found by numerical simulations of Schrödinger-Poisson (S-P) equation

Schive et al. ('14) $M_{\rm h}$: Halo mass $r_{\rm c} \propto m_{\phi}^{-1} M_{\rm h}^{-1/3} \qquad M_{\rm c} \propto m_{\phi}^{-1} M_{\rm h}^{1/3}$ m_{ϕ} : Mass of FDM

X(\= ///s] XX(==(1/5g)

soliton eigenstate $(\mathscr{E}_{1,0}/\alpha)$

Analytical approach to core-halo structure

AT & Saga ('22)

A missing factor largely changes core-halo relations, showing extra cosmological dependence

Present work

adopting the concentration-mass relation $c_{\rm vir}(M_{\rm h})$ of FDM/CDM models

Non power-law behaviors

Non-negligible amount of scatter

c.f.) recent claim by Zagorac et al. ('23) "There is no universal core-halo relationship"

Summary

To be or not to be... (non-)universal features of innermost structure of dark matter halos based on analytical & numerical study

Cold dark matter (CDM)

Enomoto, Nishimichi & AT, arXiv:2302.01531

A new remarkable feature found in multi-stream structures

Fuzzy dark matter (FDM)

A missing factor in core-halo relations found analytically

Radial multi-stream profiles $\rho_{\text{stream}}(r;p) = \frac{A(p)}{x(1+x^7)}; \quad x \equiv \frac{r}{S(p)}$ With A(p) & S(p) described by a simple fitting form

AT & Saga, arXiv:2208.06562 (in PRD)

Non power-law feature of core-halo relation dependent on <u>cosmology</u> (concentration-halo mass relation) \rightarrow no universal relation ?

