Towards global S-matrix constraints on dark matter physics

Katsuki Aoki, YITP, Kyoto University

KA 2212.07659 (to appear in PRD) + ongoing discussions.

What do we aim for?

Consistency with UV completion?

Fundamental principles such as unitarity and causality are already enough to find strong constraints on low-energy physics. A. Adams et al. 2006 and many.

Noumi-san's talk:

Dark sector may not be very dark? Mass may not be very light?

There are still theoretical subtleties but the predictions are testable!

We need to

MHz GHz Earth 10^{-2} 10^{-3} 10 COBE/FIRA 10mixing, 10^{-1} DPDM heating 10^{-8} 10^{-9} DPDM 10^{-10} **Xinetic** 10^{-1} 10^{-14} Dark 10^{-15} 10^{-16} 10^{-17} $10^{-10} - 10^{-10} - 10^{-10} - 10^{-1} 0^{-1} 0^{-1} 0^{-1} 0^{-1} 0^{-1} 0^{-1} 0^{-1} 0^{-1} 0^{-1} 0^{-1} 0^{0} 10^{0} 10^{0} 10^{2} 10^{3} 10^{4} 10^{5}$ Dark photon mass, $m_{A'}$ [eV]

From Noumi, Sato, and Tokuda, 2022.

✓ resolve the theoretical subtleties and get rigorous&stronger bounds,
✓ understand the bounds systematically and their pheno. consequences,
✓ ...

The bounds are well established for 2-to-2 scattering of the lightest state in the gapped system.



Nice properties (analytic structure, high energy behaviour) are known. Unitarity gives a simple positivity constraint $\text{Im}\mathcal{M}|_{t=0} > 0$.

 \Rightarrow Positivity bounds on coupling constants of EFT. A. Adams et al. 2006 and many.

□ However, our world is more complicated!

There are massless particles (photon, graviton)

 \rightarrow Graviton may give non-trivial constraints a la swampland.

Cf. Noumi-san's talk.

There are many massive particles.

- The bounds are well established for 2-to-2 scattering of the lightest state in the gapped system.
 - \rightarrow What are the bounds arising from all possible 2-to-2 scatterings?



Of course, this is not a new idea but must be crucial, e.g. Four-fermi interaction \rightarrow W boson is required W boson scattering \rightarrow Higgs is required.

- The bounds are well established for 2-to-2 scattering of the lightest state in the gapped system.
 - \rightarrow What are the bounds arising from all possible 2-to-2 scatterings?

Technically cumbersome but straightforward?

$$I = \Lambda$$

- + *M*: the heavy particle *A* + *µ*: the lightest particle φ

 $2\mu^2 < M^2 < 4\mu^2$: S-matrix has new singularities (anomalous thresholds). $4\mu^2 < M^2$: A can decay to $\varphi \varphi$.

 \rightarrow A does not appear in the asymptotic state. Veltman 1963.

Then, what is the S-matrix of A? If exists, what is the consequence of unitarity? ...

□ Physically, there must be unitarity constraints! $\text{Im}\mathcal{M}|_{t=0} > 0$??? But, when is this naïve intuition correct and when is it wrong?

Current status:

There indeed exist unitarity constraints and "optical theorem" even for unstable-particle amplitude.

$$\text{Im}\mathcal{M}|_{t \simeq [\text{Im}M^2]^2/s} > 0$$
 KA, 2212.07659.

(Optimistic) expectation:

Positivity bounds on unstable particles might be similar to the stable case if there is no t-channel exchange of a spin-J ($J \ge 2$) particle.

 \rightarrow Graviton is spin-2! May be crucial to swampland (and more?).

Unstable-particle amplitudes

The unstable particles do not appear in the asymptotic states.
Its existence is only seen in internal states (resonance in physical process).

The unstable particles have a complex mass and the residue is factorized. The amplitude for $\mathcal{A} \to \varphi \varphi$ can be defined by the residue.

See e.g. The analytic S-matrix, R. J. Eden et al, 1966

$$\mathcal{M}_{\mathcal{A}\to\varphi\varphi} = \frac{p_2}{p_3} \mathcal{N}_p p_1$$

The "on-shell" conditions are understood as

$$p_2^2 = p_3^2 = -\mu^2$$
, $p_1^2 = -M^2 \in \mathbb{C}$ p_1 is decaying mode.

Unstable-particle amplitudes

□ There is also a growing solution (complex conjugate).



Decaying $(+i\varepsilon)$ and growing $(-i\varepsilon)$ are denoted by + and -.

Unstable-particle amplitudes



Unitarity of unstable-particle amplitudes



Unstable case: we want to keep (Amplitude) × (Conjugate) > 0 on RHS.
Expected unitarity constraint

$$\operatorname{Im} \mathcal{M}_{\mathcal{A}\varphi \to \mathcal{A}\varphi} = \underbrace{\uparrow}_{+} \underbrace{\downarrow}_{-} \underbrace{\downarrow}_{+} \underbrace{\downarrow}_{+} \underbrace{\downarrow}_{+} \underbrace{\downarrow}_{+} \underbrace{\downarrow}_{-} \underbrace{\downarrow}_{+} \underbrace{\downarrow}_{-} \underbrace{\downarrow}_{+} \underbrace{\downarrow}_{+} \underbrace{\downarrow}_{-} \underbrace{\downarrow}_{+} \underbrace{\downarrow}_{-} \underbrace{\downarrow}_{+} \underbrace{\downarrow}_{-} \underbrace{\downarrow}_{+} \underbrace{\downarrow}_{-} \underbrace{\downarrow}_{-}$$

Unitarity of unstable-particle amplitudes

Assumptions:

<u>KA</u>, 2212.07659.

1. Lorentz invariance, 2. Unitarity, 3. Analyticity of higher pt. amplitudes.

 \Rightarrow Unitarity equations for $\mathcal{M}_{\mathcal{A}\varphi \rightarrow \mathcal{A}\varphi}$ indeed arise from unitarity of higher pt.



 $\text{Im}\mathcal{M}_{\mathcal{A}\varphi \to \mathcal{A}\varphi}(s,t) > 0$ holds in $(\text{Im}M^2)^2/s < t < t_*$

The positivity may be extended up to the nearest singularity t_* in the t-plane. e.g. $t_{\rm tri} \simeq 4\mu^2 ({\rm Im}M^2)^2/({\rm Re}M^2)^2$ (narrow width)

Unitarity equation for $(+)^{\sim}$ can be also derived although complicated.

Analyticity of unstable-particle amplitudes

Analyticity is important:

singularities determine properties of complex functions \rightarrow positivity bounds

D Analyticity of $\varphi \varphi \rightarrow \varphi \varphi$ (in s-plane).

All singularities are normal thresholds (= easily predicted by unitarity)



Analyticity of unstable-particle amplitudes

Analyticity is important:

singularities determine properties of complex functions \rightarrow positivity bounds

$$\Box$$
 Analyticity of $\mathcal{A}\varphi \to \mathcal{A}\varphi$ (in s-plane).

Anomalous thresholds appear (= not immediately followed from unitarity) But consistent with and can be derived from unitarity



The simulation is more complicated when $\text{Re}M^2 > 4\mu^2$.

Analyticity of unstable-particle amplitudes



We only need the information in $R_* < |s| < \Lambda^2$ for positivity bounds. If correct, it is not so difficult to derive positivity bounds for general masses!

t-channel anomalous threshold

The s,u-channel anomalous thresholds could be handled while we still have t-channel anomalous thresholds!

Roughly, positivity bounds are expected to give

 $\frac{d^{2n}}{ds^{2n}} [\mathcal{M}_{\mathcal{A}\varphi \to \mathcal{A}\varphi}(s, t) - \text{subtractions}]_{t \simeq 0} > 0, \quad n \ge 1$ $\mathcal{M}_{\mathcal{A}\varphi \to \mathcal{A}\varphi} = \underbrace{\mathcal{M}_{\mathcal{A}\varphi \to \mathcal{A}\varphi}(s, t) - \text{subtractions}}_{\text{Spin-J}} \propto s^{J}/(m_{J}^{2} - t) + \cdots$

The t-channel triangle has a singularity near $t\simeq 0!$ $t_{\rm tri}\simeq 4\mu^2({\rm Im}M^2)^2/({\rm Re}M^2)^2~~{\rm (narrow~width)}$

⇒ If the theory contains a spin-J ($J \ge 2$) state, $\overbrace{}$ may give a new contribution that is not present in stable case.

Relevant to gravitational EFTs (graviton) or quantum gravity (Regge states)?

Summary and Discussions

Global constraints

= constraints arising from all possible 2-to-2 scatterings



This requires general knowledge about unstable particles.

- □ There are unitarity constraints and maybe positivity bounds as well. $AA \rightarrow AA$ is technically more complicated but is possible.
- □ Whether stable or unstable may be important if a spin-J ($J \ge 2$) exists. Swampland and DM, Gravitational EFT, Quantum gravity, QCD?
- □ They are still very preliminary and further investigations are needed.