BLACK HOLES IN THE COSMOS

Joseph Silk IAP/JHU/Oxford

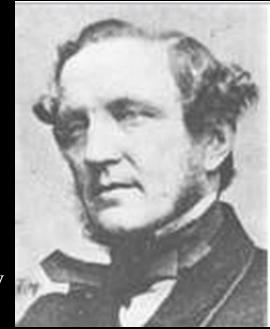
IPMU Nov 15, 2023

The black holes of nature are the most perfect macroscopic objects there are in the universe: the only elements in their construction are our concepts of space and time. Subrahmanyan Chandrasekhar



PREDICTION

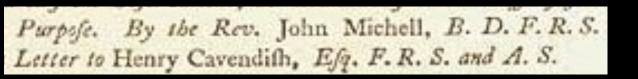
If the semi-diameter of a sphere of the same density as the Sun were to exceed that of the Sun in the proportion of 500 to 1, a body falling from an infinite height towards it would have acquired at its surface greater velocity than that of light, and consequently supposing light to be attracted by the same force in proportion to its vis inertiae, with other bodies, all light emitted from such a body would be made to return towards it by its own proper gravity.



John Michell 1783 Professor of Geology at Cambridge, then a Yorkshire vicar John Michell 1724-1793

Escape velocity from star satisfies $\frac{1}{2}$ v² = GM/r v < speed of light c If a star is too compact, light cannot escape!

The critical size is $r=2GM/c^2$



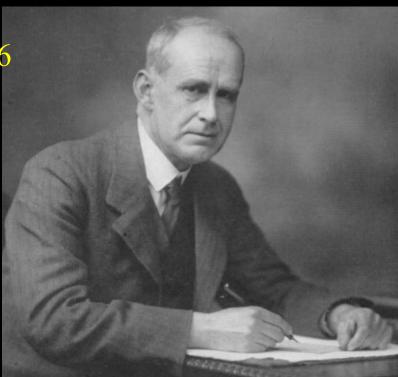


Arthur Eddington 1882-1946

imagine a physicist calculating on a cloud-bound planet and ending with the dramatic conclusion, "What 'happens' is the stars."

In 1913, Einstein wrote to many astronomers urging them tosearch for his predicted shift of 0.87"

In 1916, he changed his mind: 0.87" was the newtonian prediction, his was double this!



Eddington avoided conscription as a conscientious objector by agreeing to lead an expedition to Principe, West Africa in 1919 to measure the deflection of light during a total eclipse. Another expedition went to Sobral, Brazil.

The results:

Einstein's prediction: 1.74", Newton's theory: 0.87"

Sobral: 1.98+-0.16"; Principe: 1.61+-0.40"

LIGHTS ALL ASKEW IN THE HEAVENS

Men of Science More or Less Agog Over Results of Eclipse Observations.

EINSTEIN THEORY TRIUMPHS

Stars Not Where They Seemed or Were Calculated to be, but Nobody Need Worry.

A BOOK FOR 12 WISE MEN

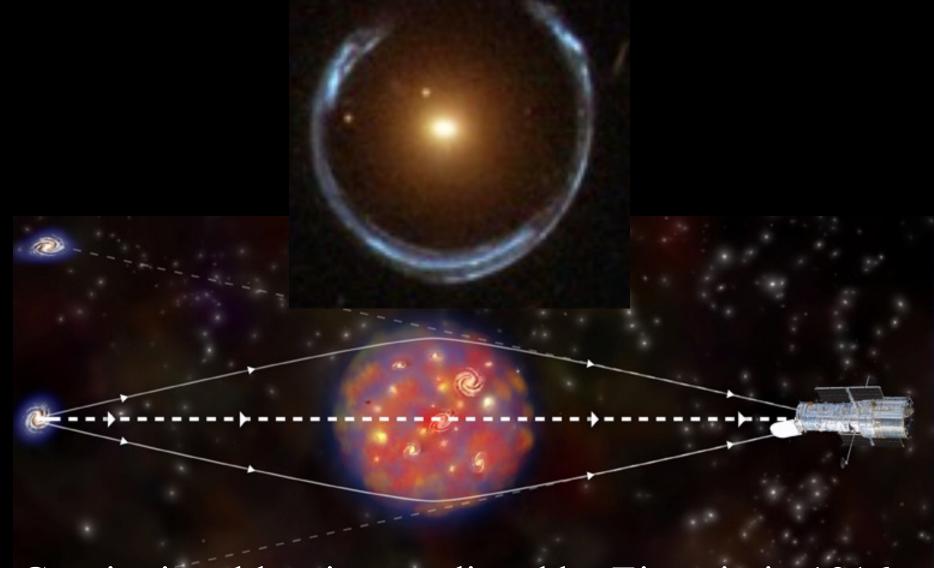
No More in All the World Could Comprehend It, Said Einstein When His Daring Publishers Accepted It.



Albert Einstein 1879-1955

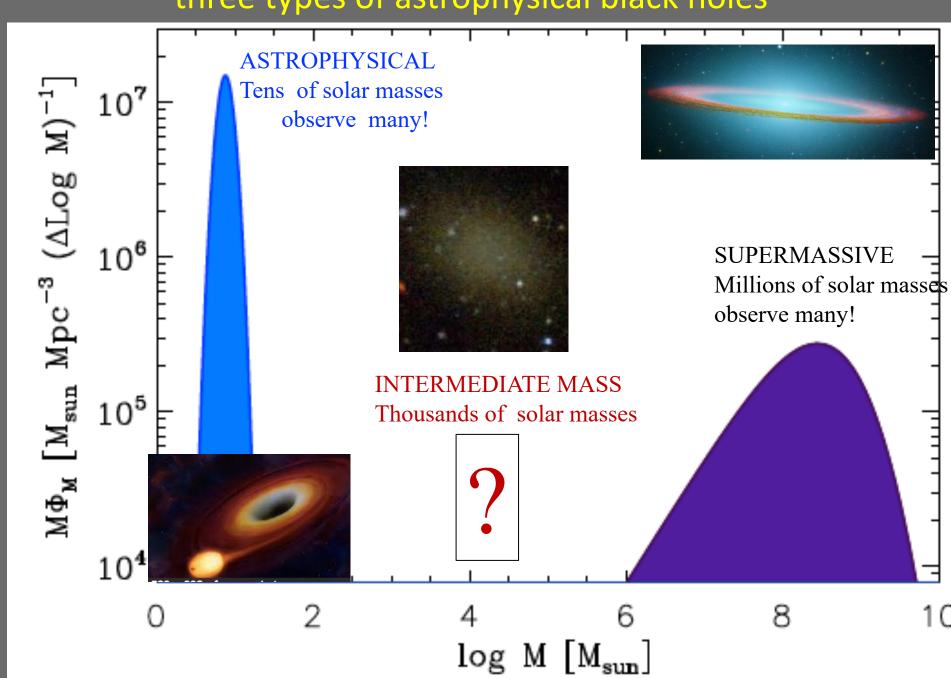
In 1916 used general relativity correct result of 4GM/rc² (time dilation + space curvature)

Gravitational lens observed

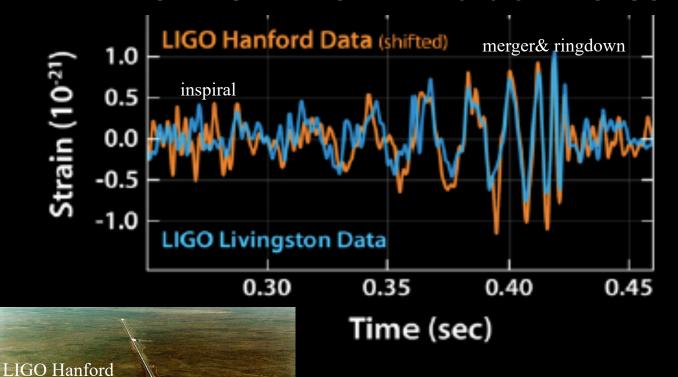


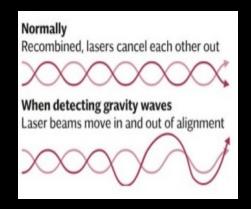
Gravitational lensing predicted by Einstein in 1916

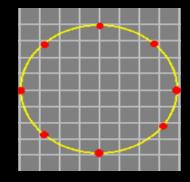
three types of astrophysical black holes

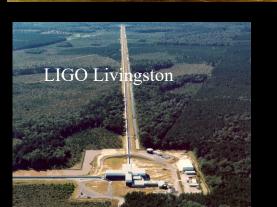


ASTROPHYSIVAL black holes observed



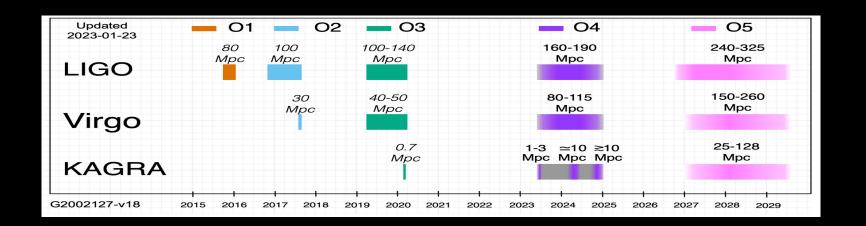




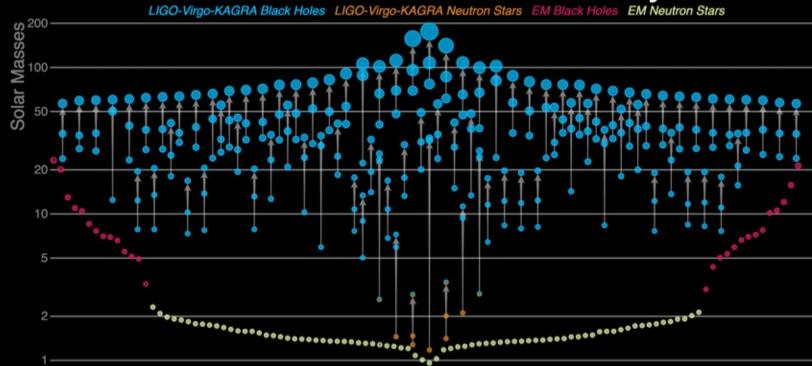








Masses in the Stellar Graveyard



LIGO-Virgo-KAGRA | Aaron Geller | Northwestern

Orbits of stars tell us black hole mass

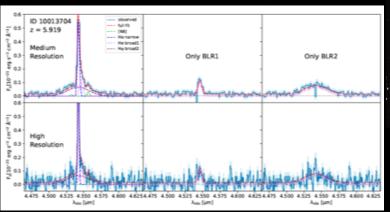
Centre of our galaxy 4.10⁶ Msun



1995.5

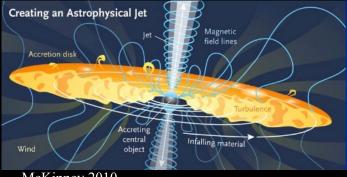
S0-24
S0-16
S0-2
S0-53
S0-3
S0-3
S0-1
S0-4
S0-4
S0-45
S0-49
S0-38
S0-23
S0-20
S0-5
S0-17
S0-6
S0-18
Keck/UCLA Galactic Center Group





Hercules A

Radio galaxy and jet

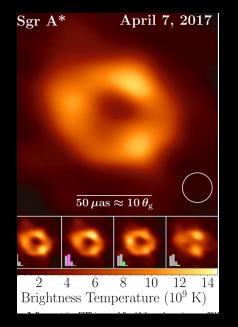


McKinney 2010

The central engine

Imaging the BH

 $4.10^6 \, M_{sun}$

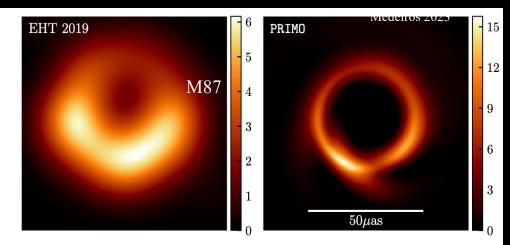


Black hole shadow observed

Event Horizon Telescope Resolve black hole horizon scale GM/c^2 at $\sim 5 \mu$ arcsec

M87 distance 2000 x GC but M_{RH} 1500 x SagA*

6.10⁹ Msun





















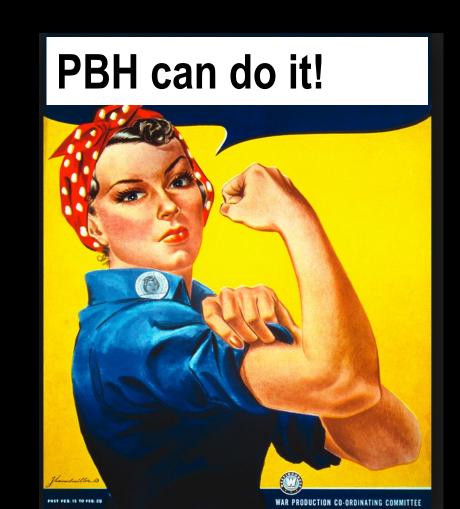




PRIMORDIAL BLACK HOLES

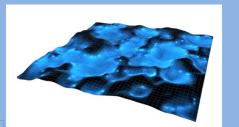
A plausible candidate for the dark matter

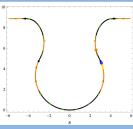
primordial black holes form by known physics



Formation Mechnisms of Primordial Black Holes

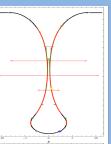
Large density perturbations (inflation)



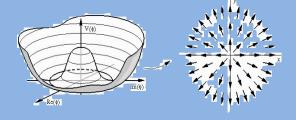


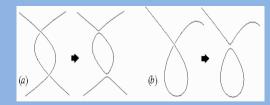
Pressure reduction

Kopp 2010

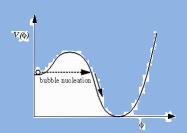


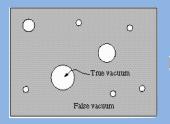
★ Cosmic string loops

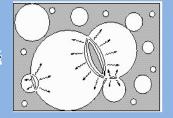




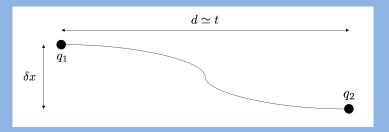
★ Bubble collisions







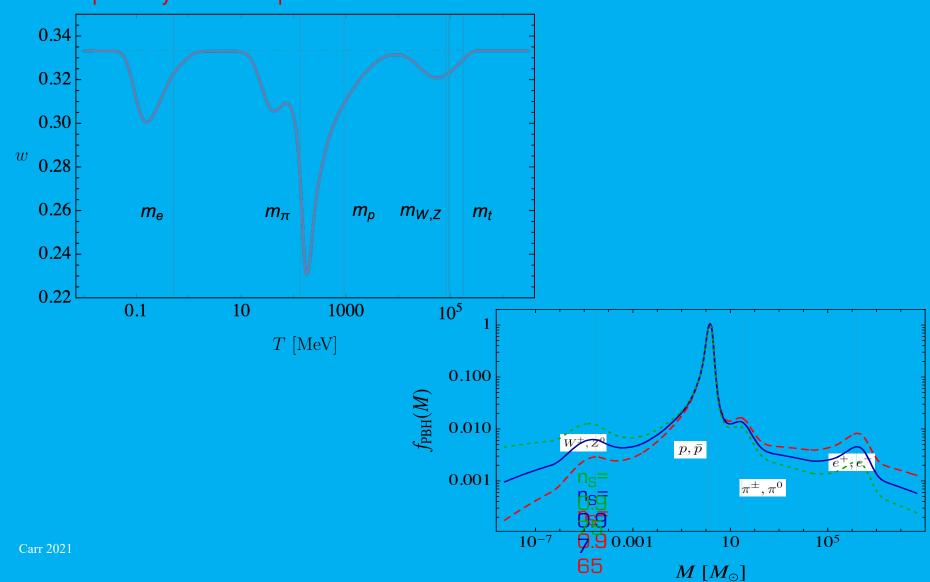
Quark confinement



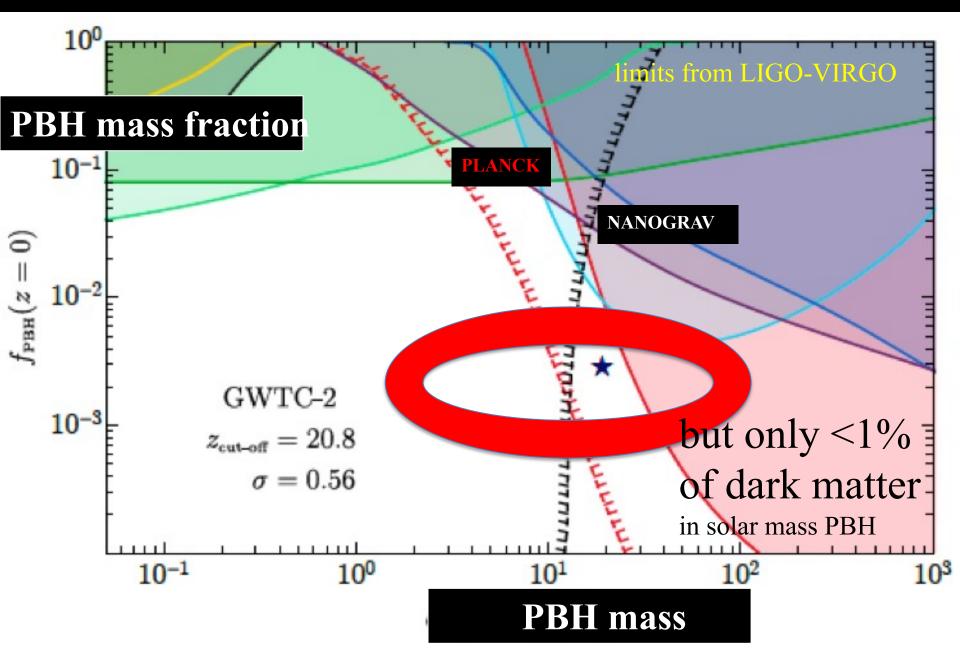
Thermal History of the Universe

Changes in equation-of-state parameter $\,w=p/
ho\,$ are destabilizing

Especially at QCD epoch

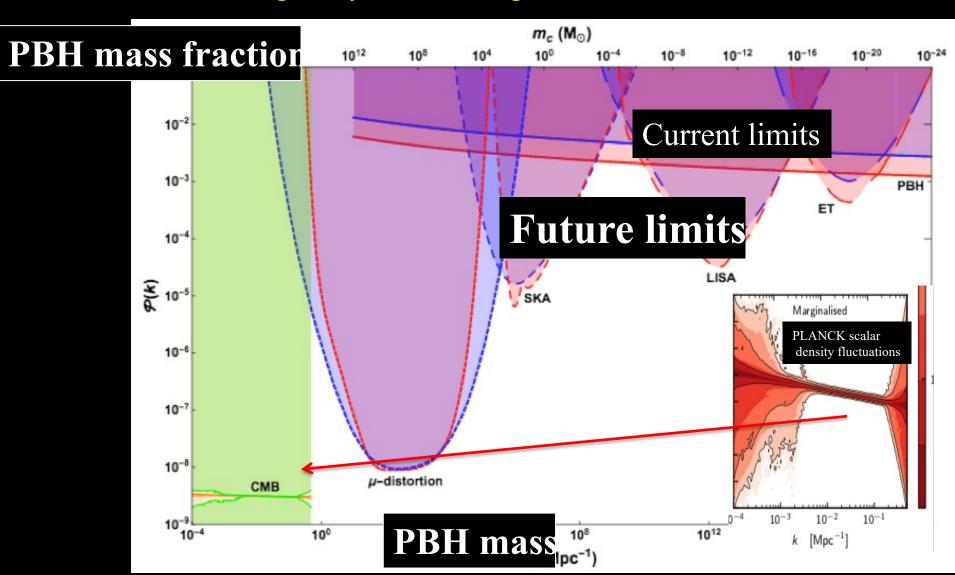


Can solar massPBH be the dark matter?

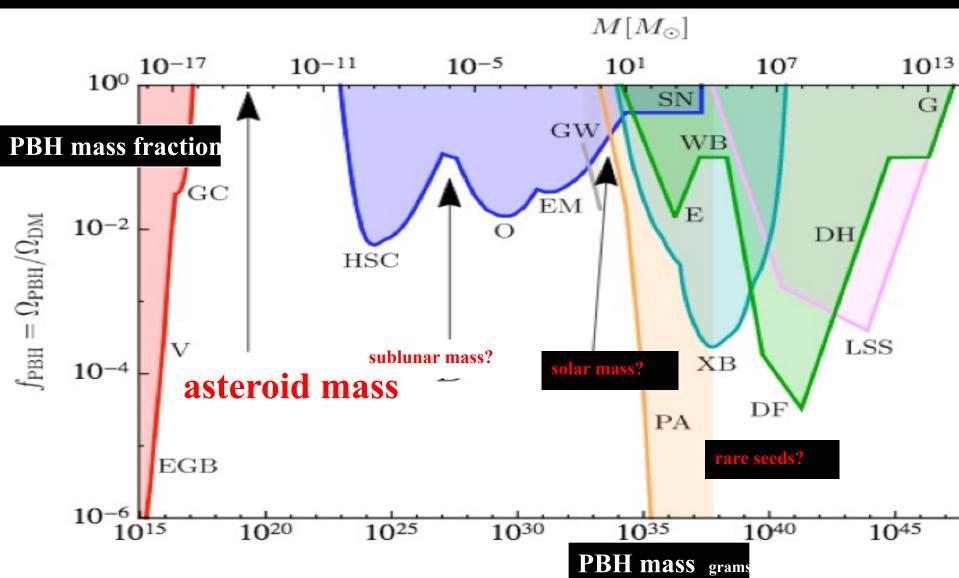


How to constrain PBH!

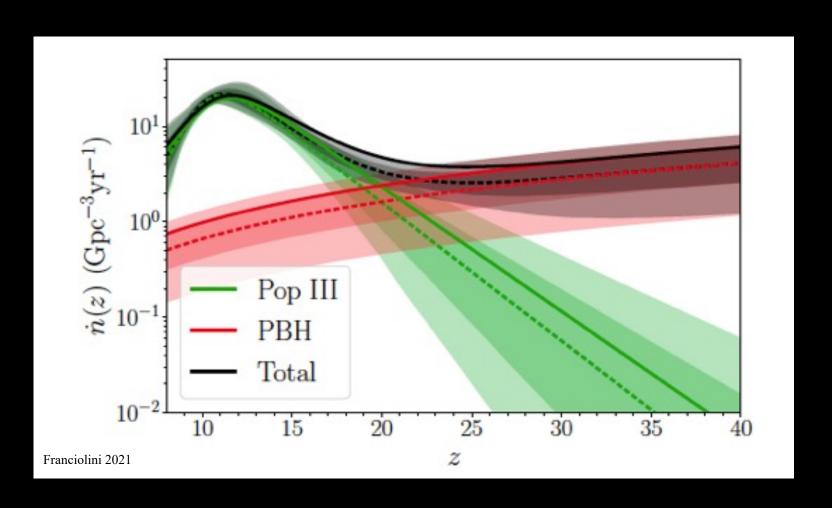
Predict stochastic gravity wave background at formation



PBH mass windows of interest but only one is a 100% dark matter option

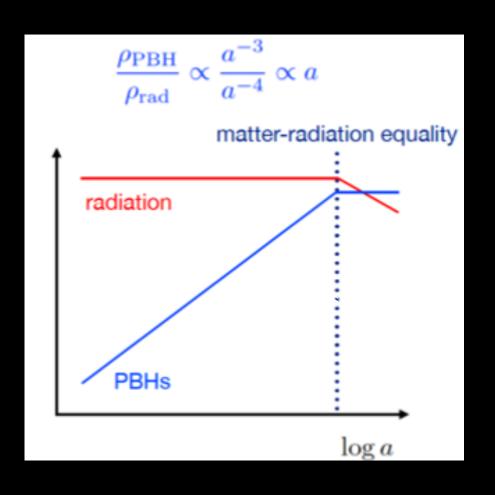


How to distinguish primordial from astrophysical BH



PBHs are rare at formation

a tiny birth fraction dominates DM at late times



Green 2001

PBH EVAPORATION

Black holes radiate thermally with temperature

$$T = \sim 10^{-7} (Msun/M) K$$

=> evaporate completely in time $t_{evap} \sim 10^{64}$ (M/Msun

 $M \sim 10^{15}g =>$ final explosion today. Multi-messenger signals

How to slow evaporation: go to higher dimension gravity!

higher dimensional black hole has T ---> 0 as Schwarzschild radius ----> dimension scale

Horizon < dimensional scale R < ~1 µm ----> 5d BH

$$\varepsilon_{\text{eff}}^{-4} = 3 \frac{(d-3)}{(d-1)} \frac{S_4}{S_d} \left(\frac{2\pi R}{r_s} \right)^{2(d-4)}$$

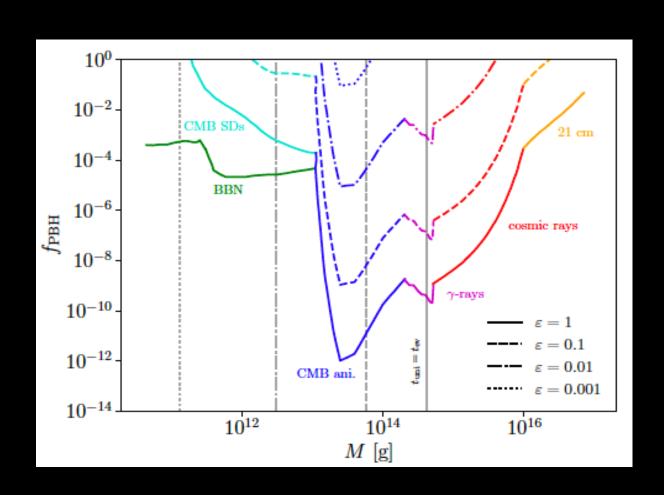
 $r_s >> R$ evaporates in 4d...until $r_s \sim R$.

$$1 - \frac{a^2}{M^2} = \varepsilon^2,$$

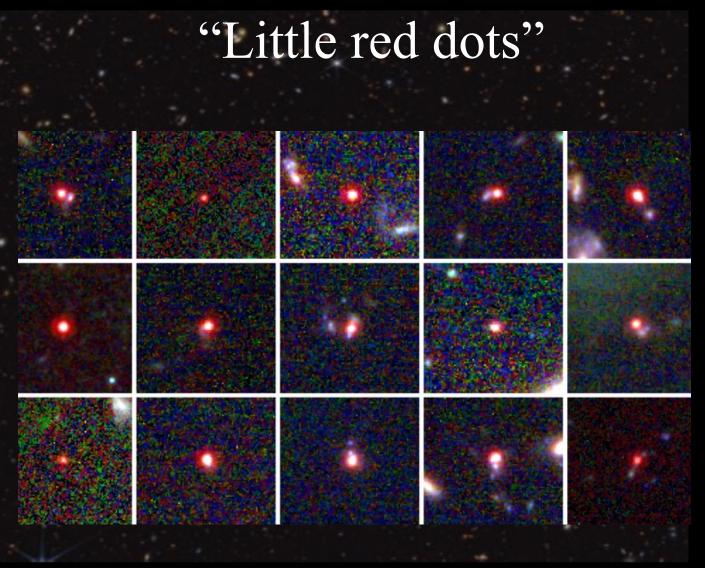
$$T = \frac{1}{4\pi r_{+}} \left(1 - \frac{Q^{2}}{r_{+}^{2}} \right) = \frac{1}{8\pi M} \frac{2^{2} \varepsilon}{(1 + \varepsilon)^{2}} \qquad L = A\sigma T^{4} \propto r_{+}^{2} T^{4} \propto \frac{2^{6} \varepsilon^{4}}{(1 + \varepsilon)^{6} M^{2}}$$

$$1 - \frac{Q^2}{M^2} = \varepsilon^2 \ll 1$$

New limits on PBH masses and on extremality



SMBH: Whats new from JWST



many compact red galaxiess

SMBH $10^7 - 10^9 M_{sun}$

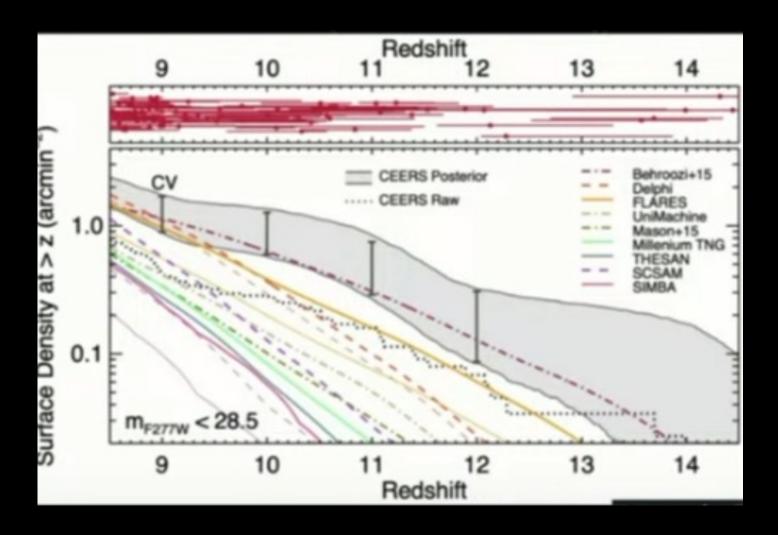
Ultracompact: R_{eff} = 150pc

z > 7

100 x frequency of quasars

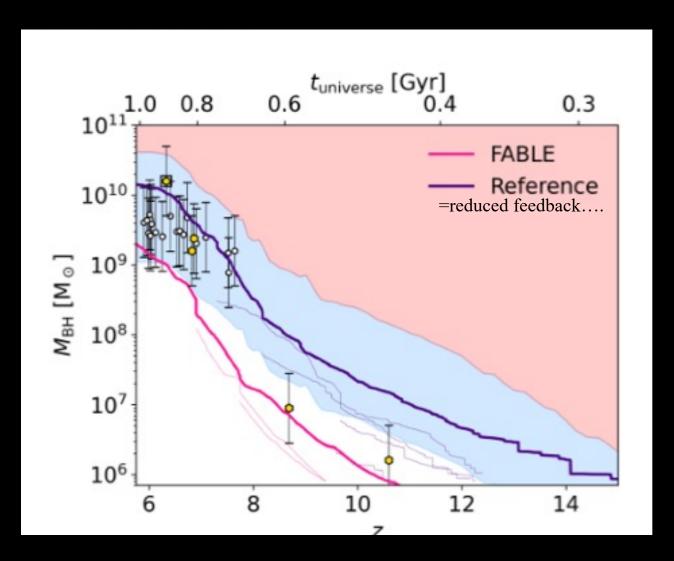
10% of "normal" galaxies

Too many galaxies, too early, too luminous



Do SMBH form first?

The missing seeds?



Origin of the seeds

a). PBH

$$f_{seed} = M_{seed}/M_{BH} \times M_{BH}/M^* \times M^*/M_{halo}$$

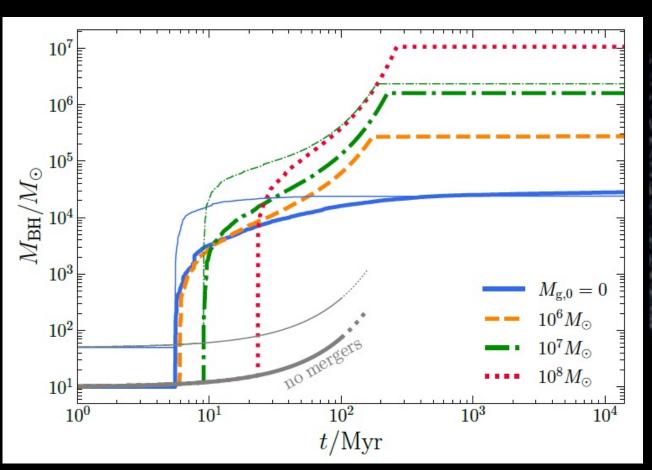
 $\sim 0.001 \times 0.01 \times 0.1 \sim 10^{-6}$

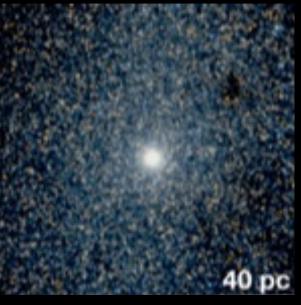
b). Early mergers in ultradense star clusters

Astrophysical black holes sink to centre and undergo runaway mergers

Growing astrophysical BH seeds in nuclear star clusters

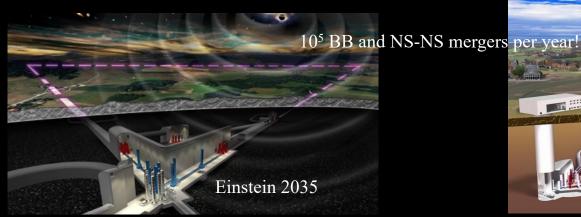
The densest star clusters in the unverse

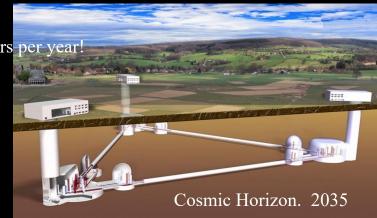


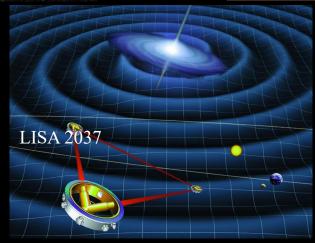


Runaway growth

The future







ありがとう