

Summary without Summary

Random Thoughts on PBHs

Focus Week on Primordial Black Holes

November 13-17, 2023

Misao Sasaki

21 fantastic talks!

1. Anne Green: Stellar microlensing constraints on PBH dark matter
2. Florian Kühnel: Positive Evidence for PBHs
3. Sergey V. Ketov: Production of PBHs as a probe of high-scale inflation and SUSY (SUGRA)
4. Jason Kristiano: One-loop correction in PBH formation from single-field inflation
5. Xinpeng Wang: PBHs from R^2 gravity theory with a nonminimally coupled scalar field

6. Marcos M. Flores: Early structure formation and PBHs
7. Guillem Domenech: Early universe cosmology of Yukawa interactions and PBHs
8. Michael Zantedeschi: PBHs from confinement
9. Albert Escrivà: Formation of trapped vacuum bubbles during inflation, and consequences for PBH
10. Yuber Perez-Gonzalez: Spin Properties of Evaporating PBHs from a Neutrino Perspective
11. Elenna Capote: Advanced LIGO and LIGO Detector Commissioning for O4 (Special seminar)

12. Kazunori Kohri: Importance of Subsolar-Mass PBHs
13. Tomohiro Harada: Revisiting compaction functions
14. Aleksander Kusenko: Newest ideas regarding the oldest black holes (APEC seminar)
15. Joe Silk: Black Holes in the Cosmos (Colloquium)

16. Sachiko Kuroyanagi: Searching for planetary-mass PBHs
17. Ravi Sheth: Predicting the abundance of PBHs
18. Ryodai Kawaguchi: Highly asymmetric PDF from a finite-width upward step during
19. Mitsunari Takahashi: PBH evaporation searches with very-high-energy gamma-ray telescopes

20. Jessica Turner: PBHs and the matter-antimatter asymmetry
21. Volodymyr Takhistov: PBHs: new signatures and new dark matter

what we have discussed

- PBH as DM, or part of DM
- probe for early universe physics
- GWs associated with PBHs
- formation mechanism: during or after inflation
- formation criterion
- ...

and

observational evidence!

PBH as (part of) DM

- evaporating PBHs: $M \lesssim 10^{15} \text{g}$, $f_{\text{PBH}} \ll 1$

- tiny PBHs:

$$M \sim 10^{22} \text{g}, f_{\text{PBH}} = 1 ?$$

- planetary mass:

$$M \sim 10^{-5} M_{\odot}, f_{\text{PBH}} \sim 10^{-2} ?$$

- subsolar mass:

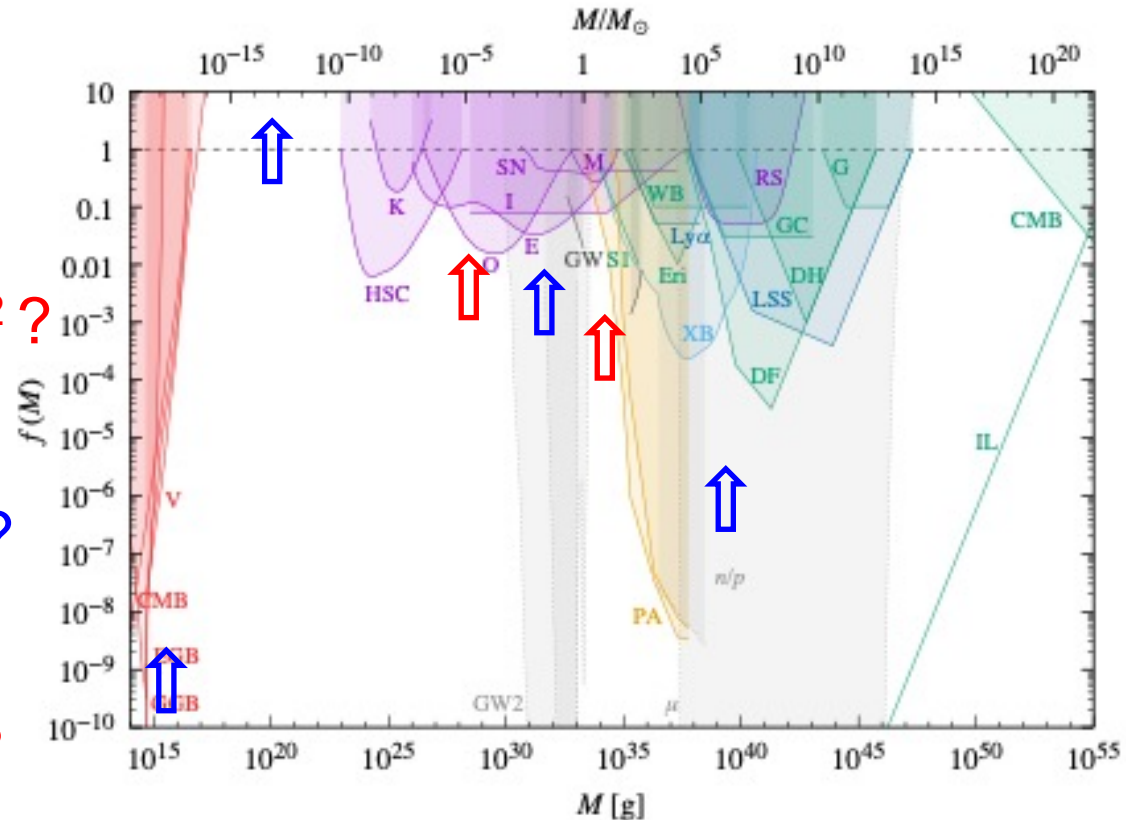
$$M \sim 0.1 M_{\odot}, f_{\text{PBH}} \sim 10^{-2} ?$$

- massive PBHs:

$$M \sim 10 M_{\odot}, f_{\text{PBH}} \sim 10^{-3} ?$$

- supermassive PBHs:

$$M \sim 10^5 M_{\odot}, f_{\text{PBH}} \sim 10^{-6} ?$$



Carr, Kohri, Sendouda, Yokoyama, 2021.12278

probe for early universe physics

- inflation

spectrum & non-Gaussianity

↔ potential features, multiverse, confinement?

- phase transitions, EOS, Yukawa force

oscillons, PBH formation + GWs?

- matter-antimatter asymmetry from PBH

PBH jets/evaporation: non-equil + C & CP violation
modified gravity?

GWs associated with PBHs

- **GWs from PBH formation**
strong force collapse, string-induced, bubble collision, early MD, ...
- **GWs from PBH binaries**
LVK binaries, subsolar mass binaries, SMBH,...
- **scalar-induced GWs**
curvature and/or isocurvature, **non-Gaussianity?**
- **GWs from evaporating PBHs**
poltergeist GWs, isocurvature, ...
- **parity violation signatures?**

Formation Mechanisms

-during or after inflation-

- enhanced curvature perturbation
formation during RD, MD, wD, ...
- type I (normal) vs type II (wormhole-like?)
small mass PBHs for $>O(1)$ amplitude perturbations?
- trapped during inflation/quantum tunneling
proving multiverse, need more quantitative studies
- strong force collapse
heating/cooling, fermion stars? $M_{ch} \sim m_{\psi} y^{-3} \quad (y \ll 1)$
- PBH clustering
non-Gaussian curvature/isocurvature perturbation, ...

Formation Criterion

- Press-Schechter formalism
perhaps not reliable any more...
- Window Function dependence?
observables shouldn't depend on WF

- Peaks theory
inclusion of non-sphericity

- Critical behavior
small mass tail wouldn't reflect reality

$$M = kM_H(\delta - \delta_c)^\gamma$$

- Compaction Function
legitimate C_{CMC} vs universal C_{ss}

$$F(\overset{\uparrow}{\partial^2\psi}, \partial\psi)$$

$$F(\overset{\nearrow}{\partial\psi})$$

$$R^{(3)} \ni \partial^2\psi, \partial\psi$$

✓ why doesn't the criterion contain 2nd derivatives?

✓ spherical symmetry artifact?

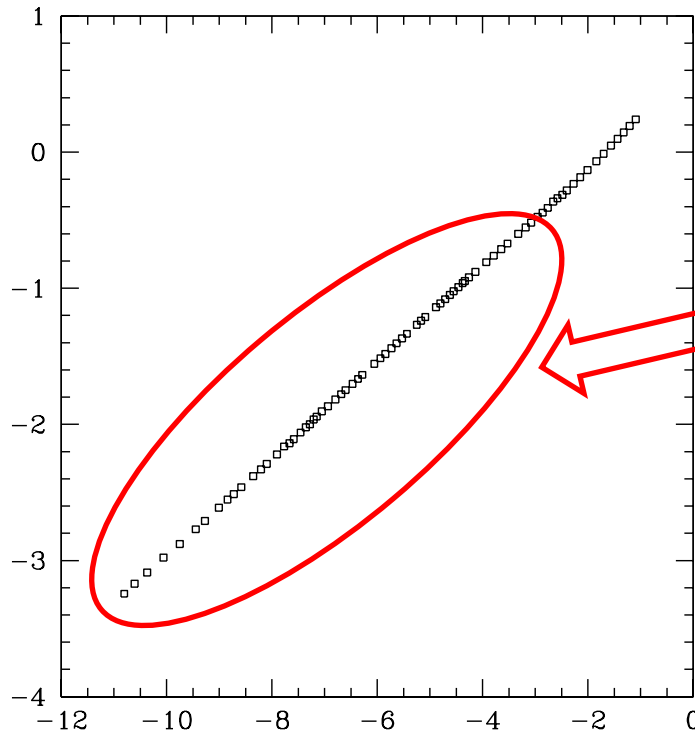
Critical behavior

(borrowed from Anne's slides)

$$M = kM_H(\delta - \delta_c)^\gamma$$

$$\gamma \simeq 0.36$$

$$\log\left(\frac{M_{\text{BH}}}{M_H}\right)$$



tiny mass BHs are produced just above the threshold δ_c

does this really happen?

Musco, Miller & Polnarev $\log(\delta - \delta_c)$

non-sphericity will kill critical behavior

温故知新 = learning from the past

THE ASTROPHYSICAL JOURNAL, **304**:15–61, 1986 May 1

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BBKS

THE STATISTICS OF PEAKS OF GAUSSIAN RANDOM FIELDS

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Received 1985 July 25; accepted 1985 October 9

BBKS Sec. 7

a) *The Triaxial Ellipsoid Approximation*

In the immediate neighborhood of a peak, the density profile is given by the Taylor expansion

$$F(r) = F(0) - \sum \lambda_i r_i^2 / 2 .$$

$$\Rightarrow F(r) \approx v\sigma_0 - x\sigma_2 \frac{r^2}{2} [1 + A(e, p)]$$

eigen values of $F_{,ij}$

$$x = -\nabla^2 F / \sigma_2 \quad e = \frac{\lambda_1 - \lambda_3}{2 \sum \lambda_i}, \quad p = \frac{\lambda_1 - 2\lambda_2 + \lambda_3}{2 \sum \lambda_i} .$$

$$A(e, p) = 3e[1 - \sin^2 \theta(1 + \sin^2 \phi)] + p(1 - 3 \sin^2 \theta \cos^2 \phi) .$$

$$\lambda_2 = \lambda_3 : \text{oblate spheroid} \Leftrightarrow e = p$$

$$\lambda_1 = \lambda_2 : \text{prolate spheroid} \Leftrightarrow e = -p$$

but the probability to be axi-symmetric is very low

High ν peaks are neither oblate nor prolate, but they are definitely triaxially asymmetric,

$$P_{ep}(e, p) \approx P_{ep}(e_m, p_m) \exp \left[-\frac{(e - e_m)^2}{2\sigma_e^2} - \frac{(p - p_m)^2}{2\sigma_p^2} \right],$$

$$e_m = \frac{1}{\sqrt{5x[1 + 6/(5x^2)]^{1/2}}}, \quad \sigma_e = \frac{e_m}{\sqrt{6}}, \quad p_m = \frac{6}{5x^4[1 + 6/(5x^2)]^2}, \quad \sigma_p = \frac{e_m}{\sqrt{3}}.$$

$p \sim 0$, but

$$e \rightarrow e_m \sim \nu^{-1}$$

(typically $\gtrsim 0.1$)



conjecture!

critical behavior is killed

for

$$(\delta - \delta_c) \lesssim e_m \delta_c \sim 0.1 \delta_c$$

non-Gaussianity!?

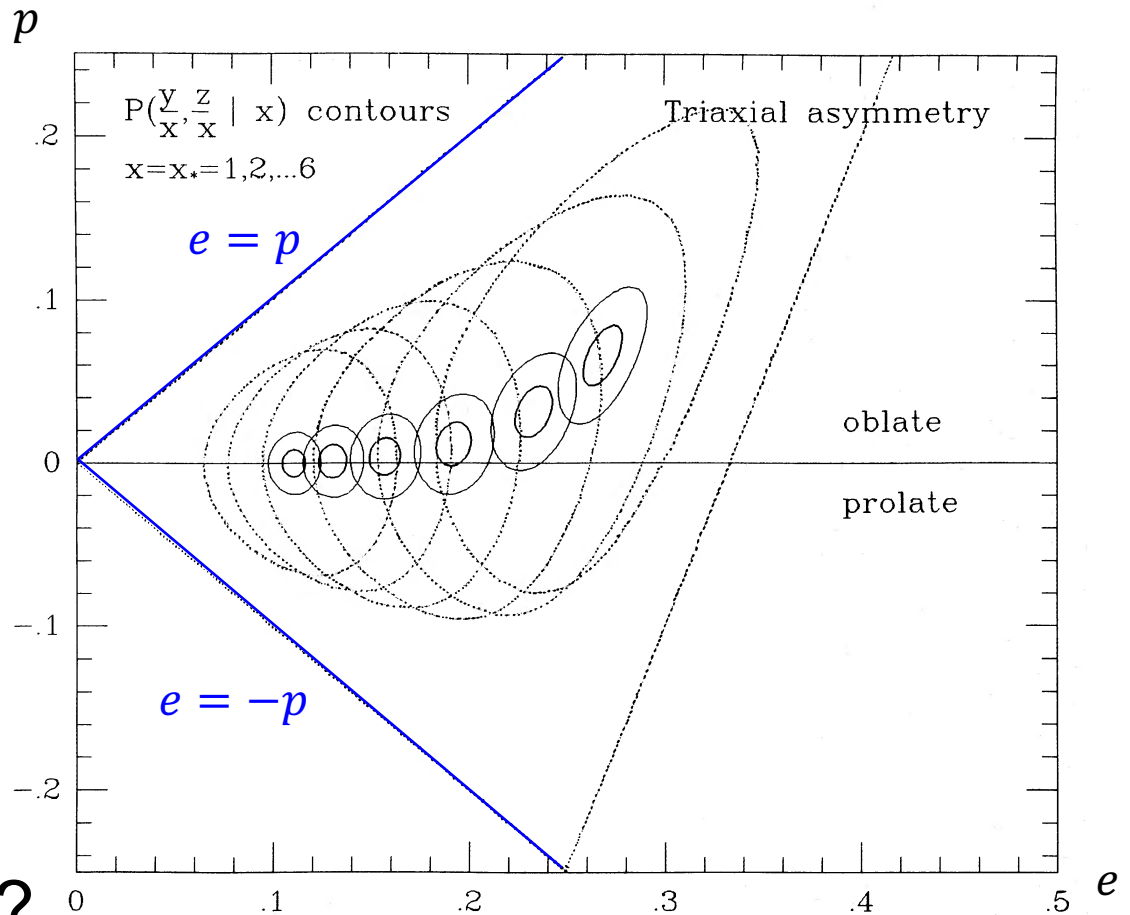


FIG. 7.—The 95%, 90%, and 50% contours

Summary of Summary without Summary

- PBHs have started to play an important role in astrophysics/cosmology/gravity/particle physics or in **fundamental physics**
- They may (have already?) become **a leading character.**
- There are a lot of **fascinating issues associated with/related to PBHs** waiting for us to be solved.



So, whether Florian wins the bet or not,

bottoms up! Kanpai! 乾杯!