



UNIVERSITÉ
DE GENÈVE

FACULTÉ DES SCIENCES



vPRISM 2nd workshop

vPRISM in antineutrino mode

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Introduction

Now that the mixing angle and the magnitude of mass differences of the PMNS matrix are accurately known, what remains to be determined is the mass hierarchy and the CP-violating phase δ_{CP} .

T2K has a good sensitivity to constrain δ_{CP} and ran in anti- ν mode last June.

The anti- ν mode is mandatory to determine the CP violation by observing a difference of oscillating events between ν and anti- ν :

$$A = P(\nu_\mu \rightarrow \nu_e) - P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)$$

$$A = \cos(\theta_{13}) \sin(2\theta_{12}) \sin(2\theta_{23}) \sin(2\theta_{13}) \sin(\delta) \sin^2\left(\frac{\Delta m_{31}^2 L}{4E_\nu}\right) \sin\left(\frac{\Delta m_{21}^2 L}{4E_\nu}\right)$$

If we build vPRISM, it will need to run in anti- ν mode : a good characterisation of the flux is mandatory to extract the physics quantities of our interest !

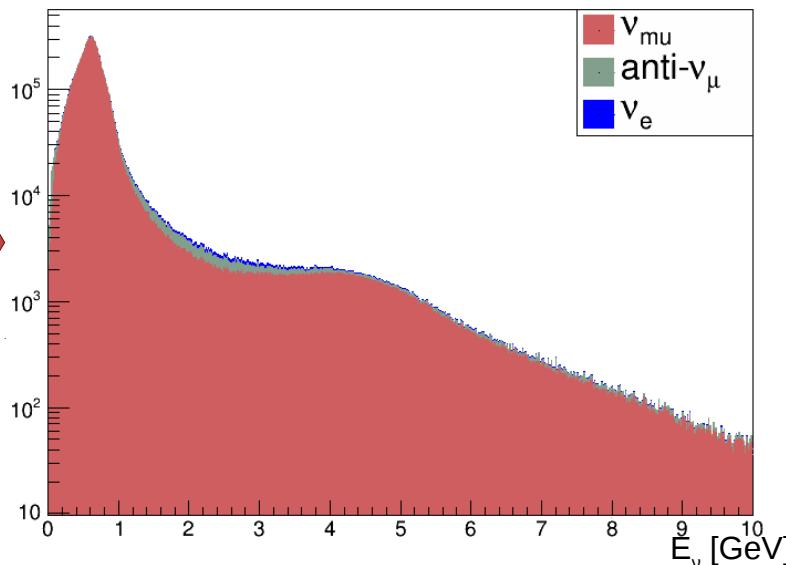
disclaimer : the plots I will show are done processing files make by M. Scott and M. Hartz. More a “what I think I will do” than a “what I have done” talk.

ν PRISM VS Super-Kamiokande flux

Super-K flux



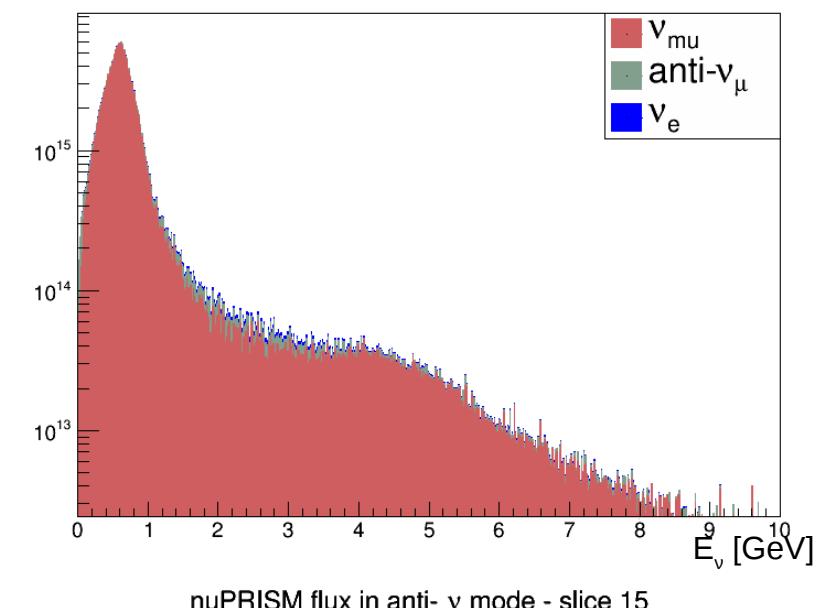
SK flux in ν mode



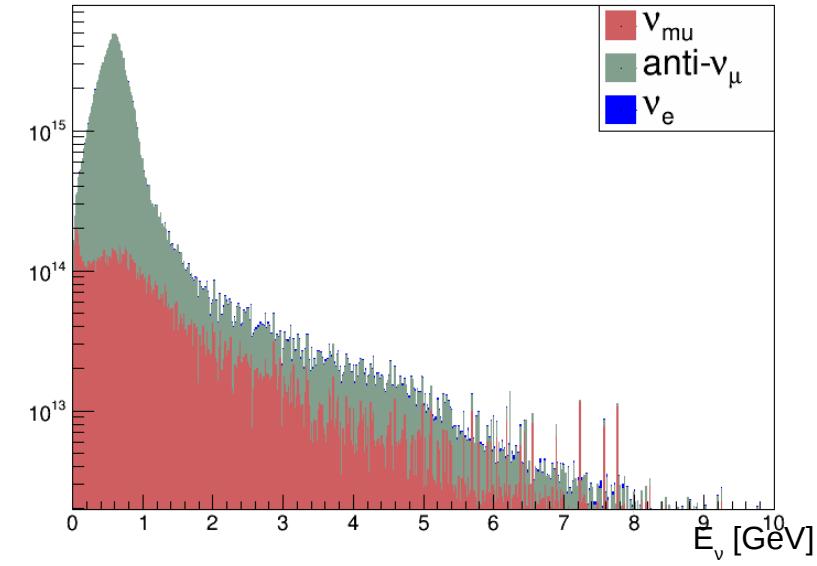
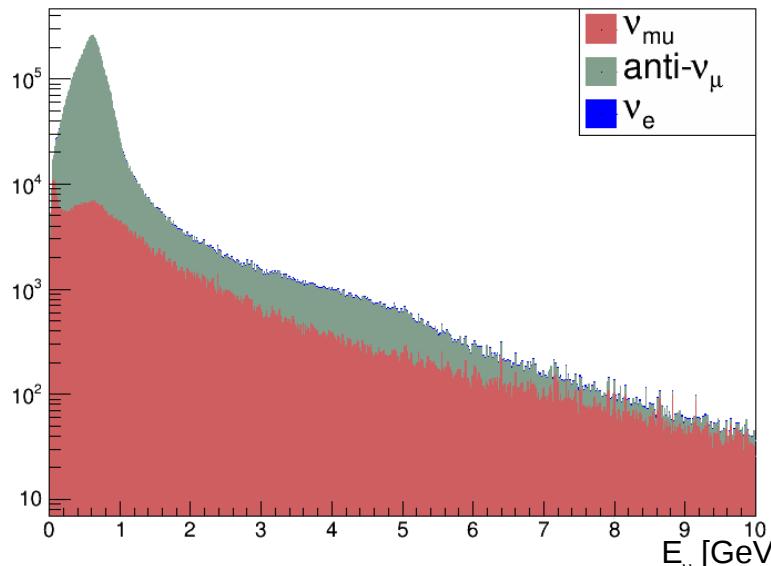
ν PRISM flux (2.4°)



nuprism flux in ν mode - slice 15



SK flux in anti- ν mode



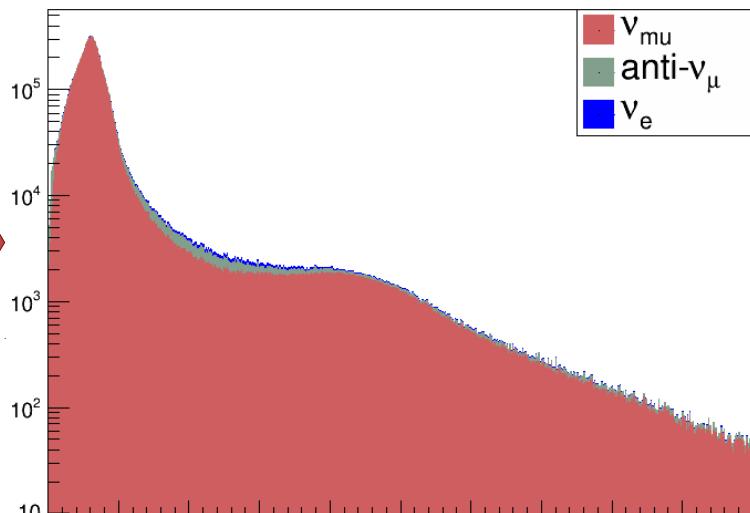
(flux stored in the nuprism_spectra.root file)

ν PRISM VS Super-Kamiokande flux

Super-K flux



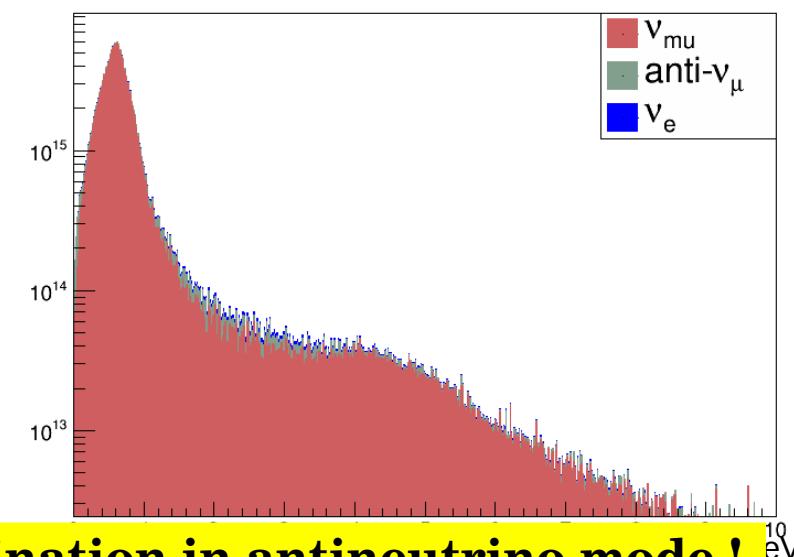
SK flux in ν mode



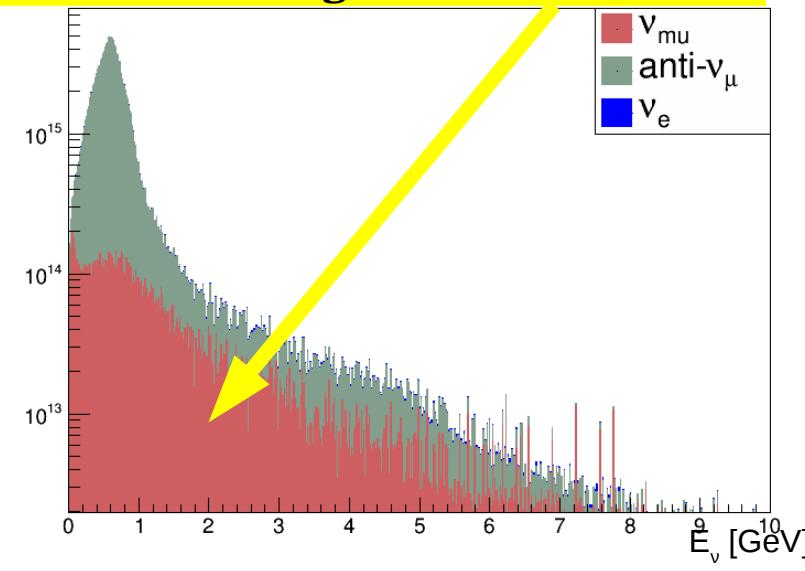
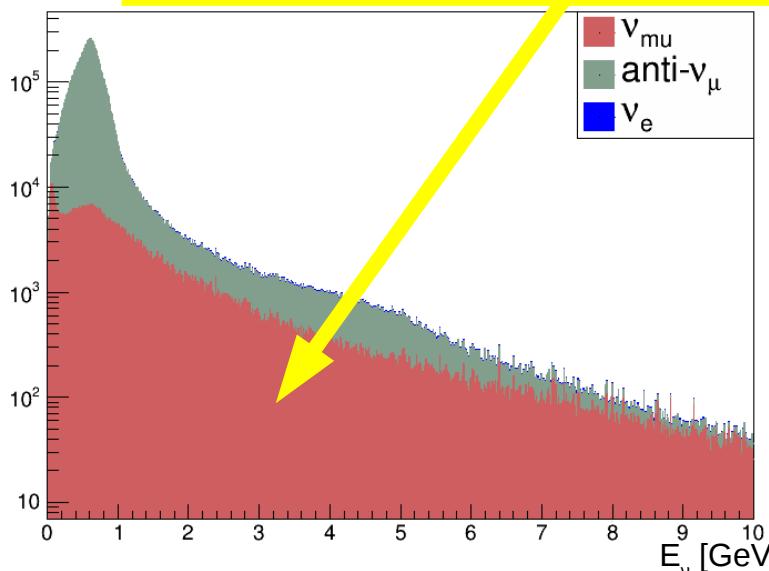
ν PRISM flux (2.4°)



nuPRISM flux in ν mode - slice 15



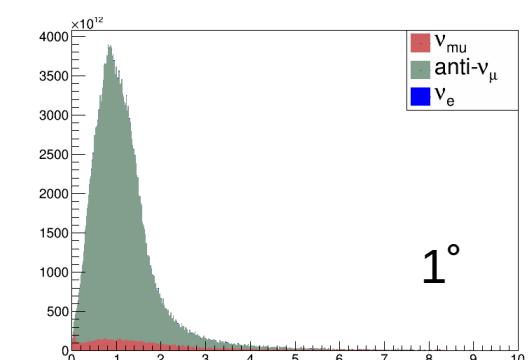
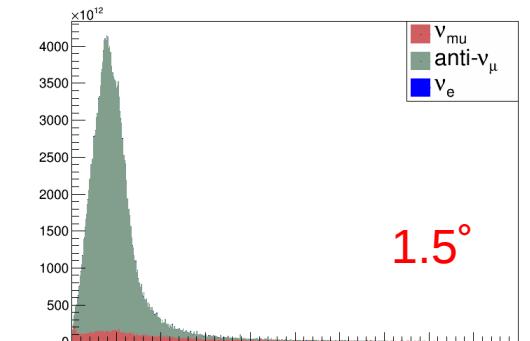
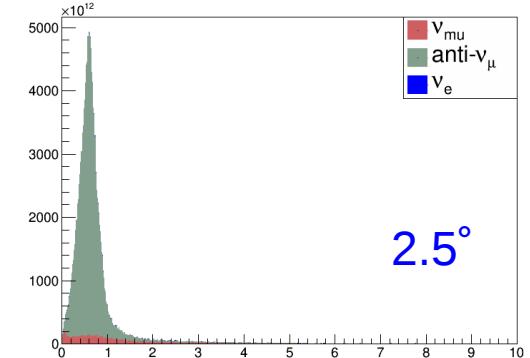
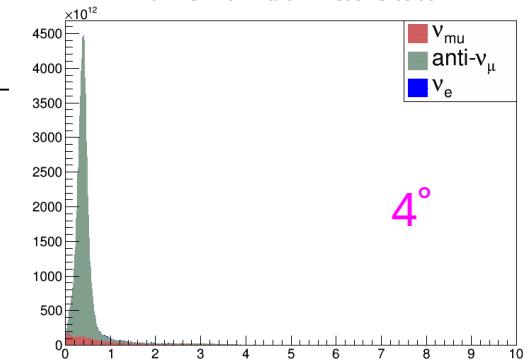
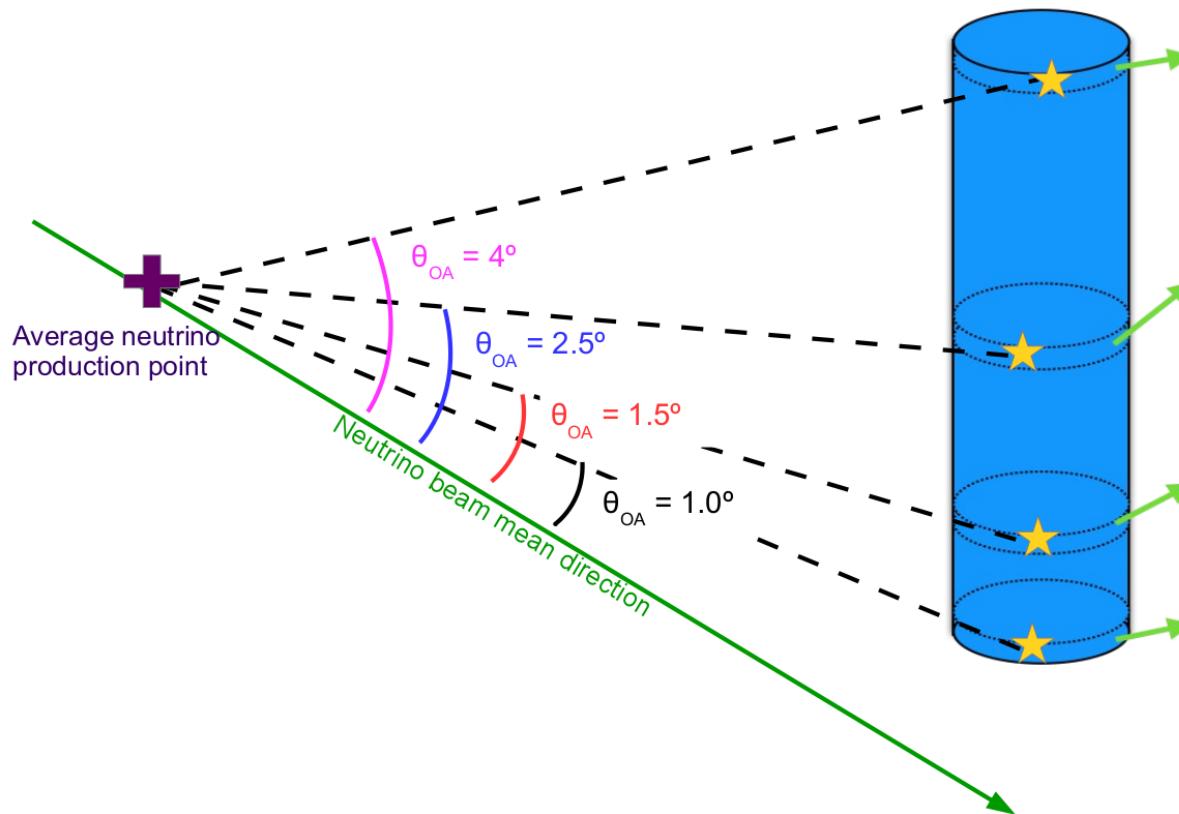
**much more “wrong sign” contamination in antineutrino mode !
(mind the log scale in Y)**



(flux stored in the nuprism_spectra.root file)

vPRISM in anti- ν mode

I only saw those plots for ν mode so here there are for the anti- ν mode.

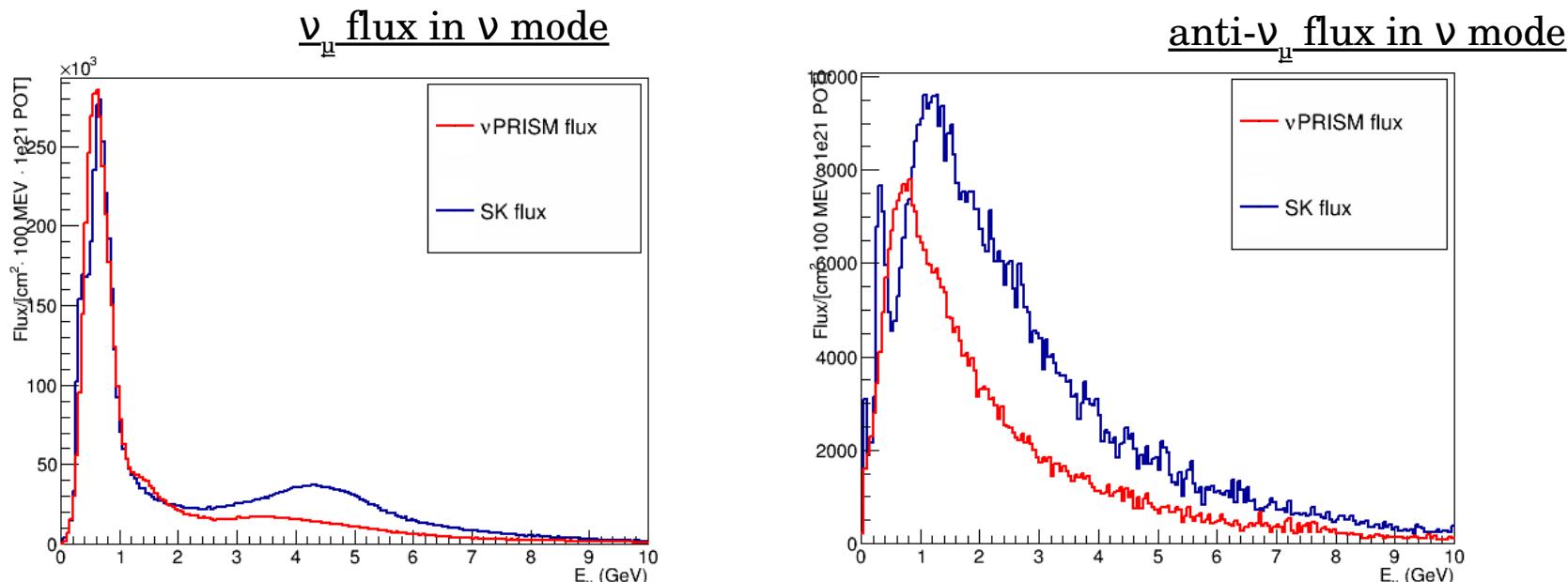


fitting the anti- ν mode flux

So now we need to fit 30 angles slices (spanning 3° from 1° to 4°) and assign a coefficient to each off-axis angle “bin” in order to model the Super-K (oscillated) flux.

Contrarily to what has been done for the flux in ν mode, we cannot neglect the background in anti- ν mode : we need to subtract the ν_μ content before we fit the anti- ν_μ vPRISM flux.

A parenthesis : here is the output I get from the macro fitting the ν_μ content of the beam in ν mode and applies the determined coefficients to the ν_μ and the anti- ν_μ vPRISM flux bins. I am not sure we can say it works well for both, neither do I get why the oscillation peak only appear for the anti- ν_μ .



fitting the anti- ν mode flux

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Contrarily to what has been done for the flux in ν mode, we cannot neglect the background in anti- ν mode : we need to subtract the ν_μ content of the beam before we fit the anti- ν_μ vPRISM flux.

In steps :

1. fit of the vPRISM ν_μ contamination in anti- ν mode with a linear combination of the vPRISM ν_μ content in forward-current mode (ν mode)
2. fit of the Super-K anti- ν_μ content in reversed-current mode with a linear combination of the vPRISM anti- ν_μ content in reversed-current mode (anti- ν mode)
3. fit of the Super-K ν_μ contamination in reversed-current mode with a linear combination of the vPRISM ν_μ content in forward-current mode

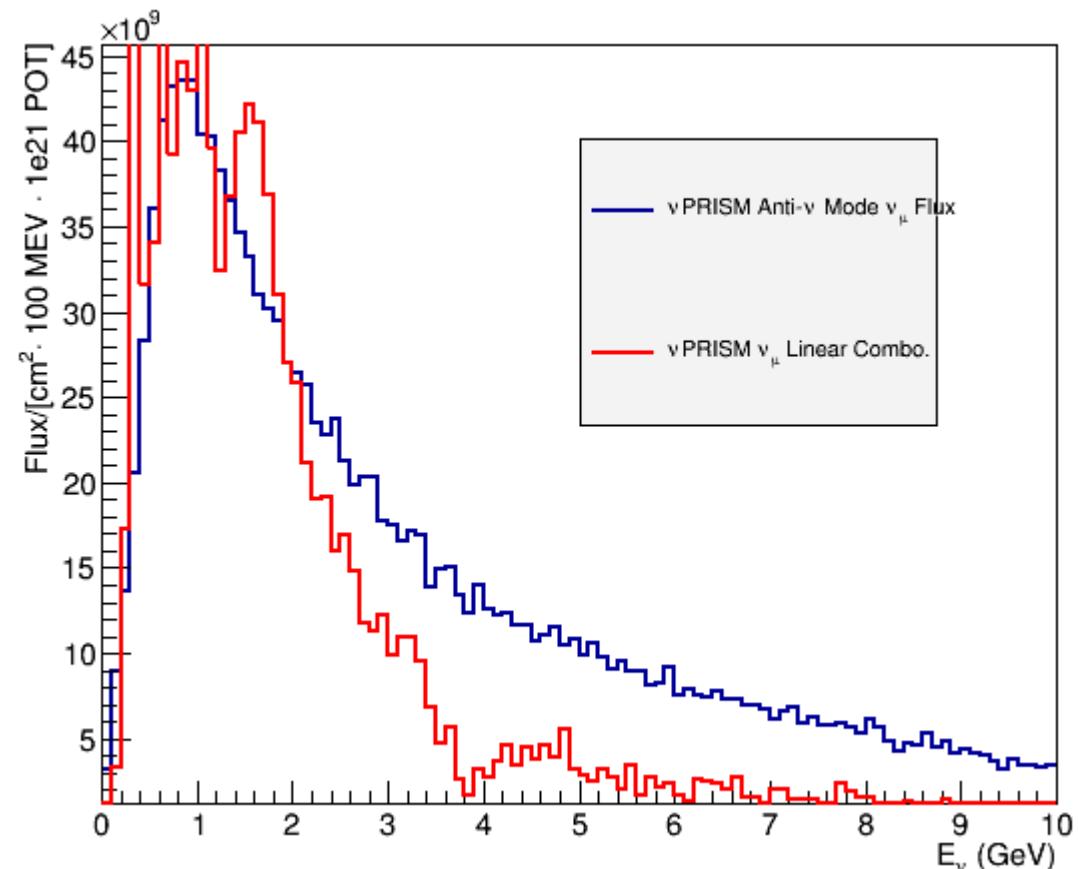
fitting the anti- ν mode flux

1. fit of the vPRISM ν_μ contamination in reversed-current mode with a linear combination of the vPRISM ν_μ content in forward current mode
2. fit of the Super-K anti- ν_μ content in reversed-current mode with a linear combination of the vPRISM anti- ν_μ content in reversed-current mode
3. fit of the Super-K ν_μ contamination in reversed-current mode with a linear combination of the vPRISM ν_μ content in forward-current mode



on the plot, output of the macro :
fit_spectrum_Wrongsign.cc

Can we make it better ? Should we
use ND280 as an input ?



fitting the anti- ν mode flux

1. fit of the vPRISM ν_{μ} contamination in reversed-current mode with a linear combination of the vPRISM ν_{μ} content in forward-current mode
2. fit of the Super-K anti- ν_{μ} content in reversed-current mode with a linear combination of the vPRISM anti- ν_{μ} content in reversed-current mode
3. fit of the Super-K ν_{μ} contamination in reversed-current mode with a linear combination of the vPRISM ν_{μ} content in forward-current mode

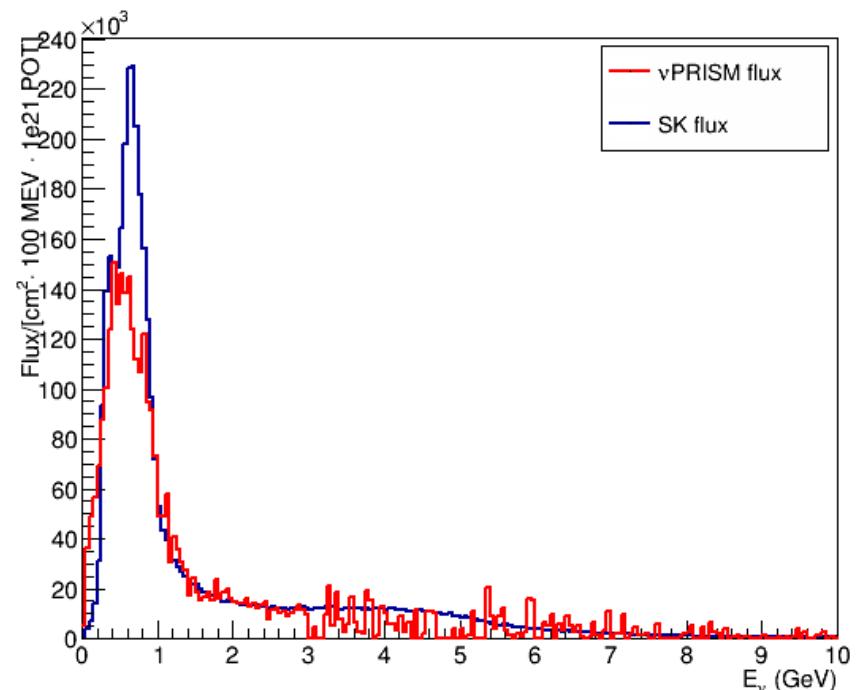


I managed to get some results, but the fit fails to correctly recover the peak.

The distribution is not smooth, but the introduction of a penalty term in the χ^2 makes the fit totally unable to resolve the peak.

I need to improve the fit.

anti- ν_{μ} flux in reversed-current mode



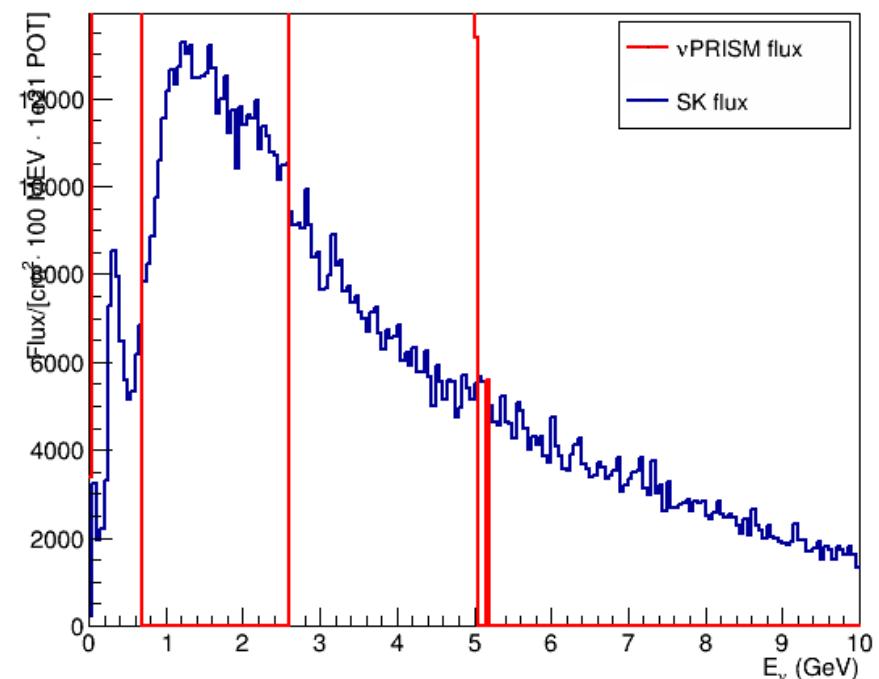
fitting the anti- ν mode flux

1. fit of the vPRISM ν_{μ} contamination in reversed-current mode with a linear combination of the vPRISM ν_{μ} content in forward-current mode
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A parenthesis : it really does not work to fit the ν_{μ} flux content with the same parameters.

ν_{μ} flux in reversed-current mode



fitting the anti- ν mode flux

1. fit of the vPRISM ν_{μ} contamination in reversed-current mode with a linear combination of the vPRISM ν_{μ} content in forward-current mode
2. fit of the Super-K anti- ν_{μ} content in reversed-current mode with a linear combination of the vPRISM anti- ν_{μ} content in reversed-current mode
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Not done yet... May need to think how to scale the ν_{μ} content determined in ν mode to the ν_{μ} content in anti- ν mode ? Or will it happens naturally in the fit ?

To be tried soon...

and after...

Summary of the fit of the vPRISM flux in anti-v mode :

1. fit of the vPRISM $\bar{\nu}_\mu$ contamination in anti-v mode with a linear combination of the vPRISM $\bar{\nu}_\mu$ content in forward-current mode → done
2. fit of the Super-K anti- $\bar{\nu}_\mu$ content in reversed-current mode with a linear combination of the vPRISM anti- $\bar{\nu}_\mu$ content in reversed-current mode → in progress
3. fit of the Super-K $\bar{\nu}_\mu$ contamination in reversed-current mode with a linear combination of the vPRISM $\bar{\nu}_\mu$ content in forward-current mode → to do

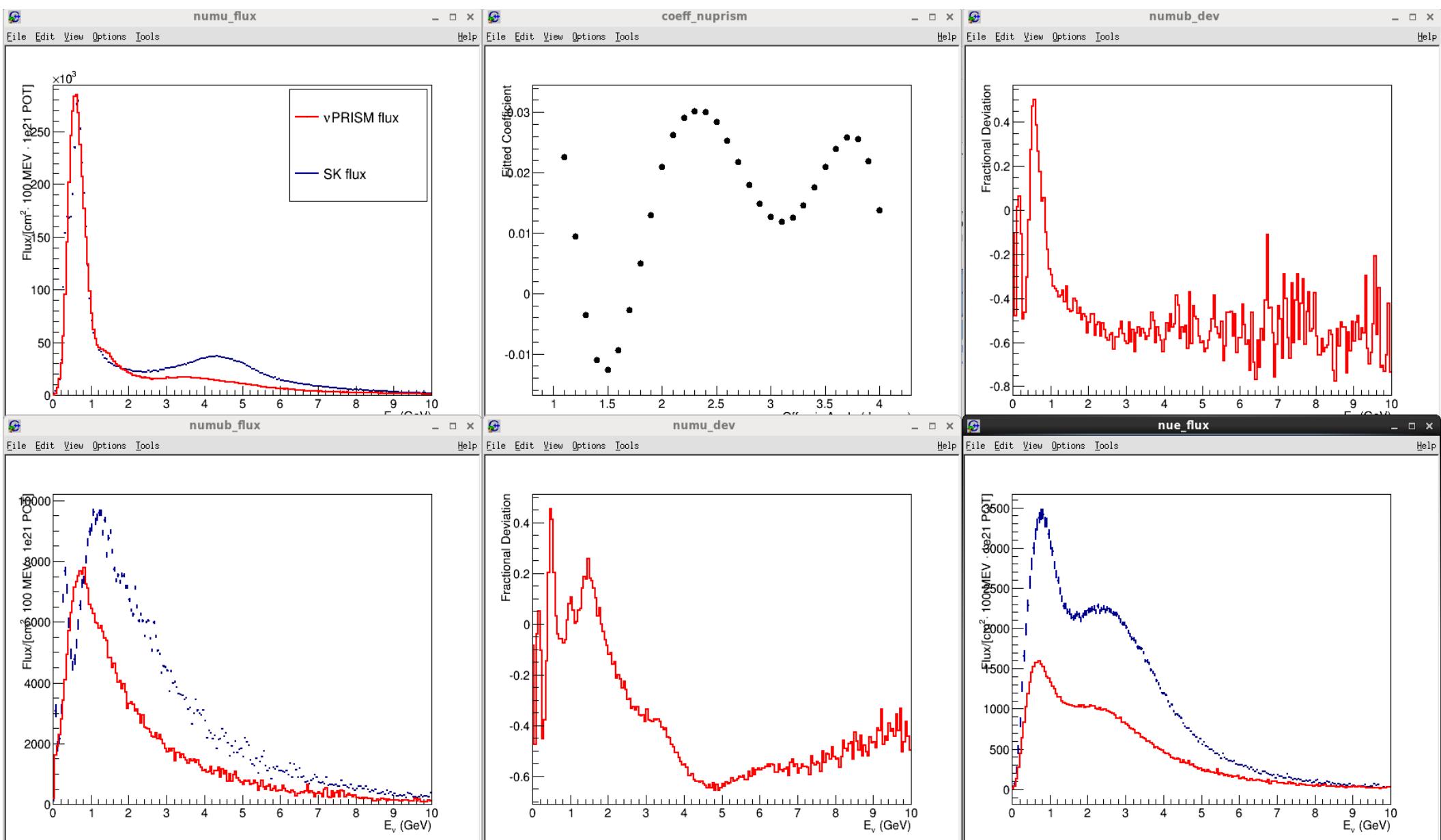
Then there will be many studies similar to the v mode that will probably be needed :

- e+ background sources
- cross-section ratio $\frac{\sigma(\bar{\nu}_e)}{\sigma(\bar{\nu}_\mu)}$
- cross section on water (CCQE, NC...)

Stay tuned, and thanks for your attention !

backup

ν mode flux fit output



anti- ν mode flux fit output

