

# Update from group C02: Cosmic structure formation

**Shin'ichiro Ando**

University of Amsterdam / University of Tokyo

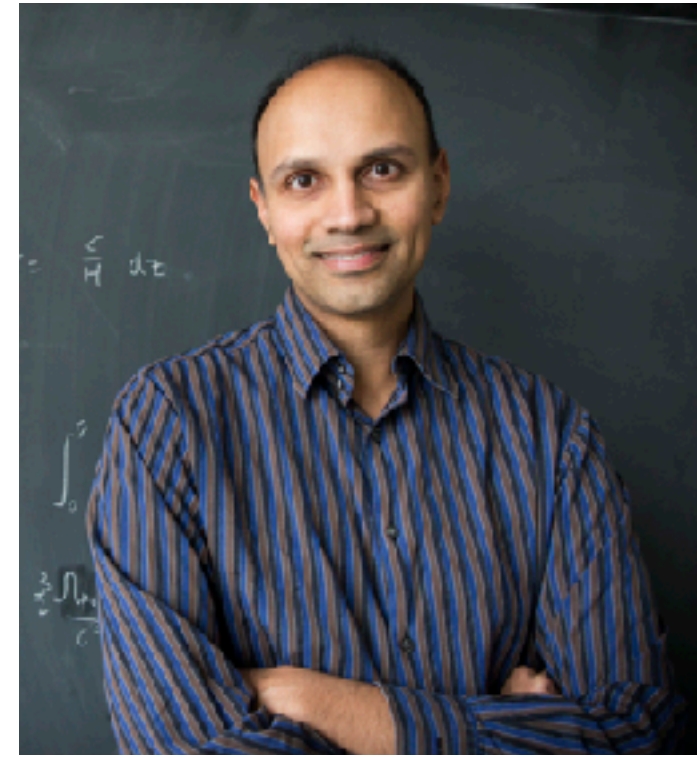


# C02: Members

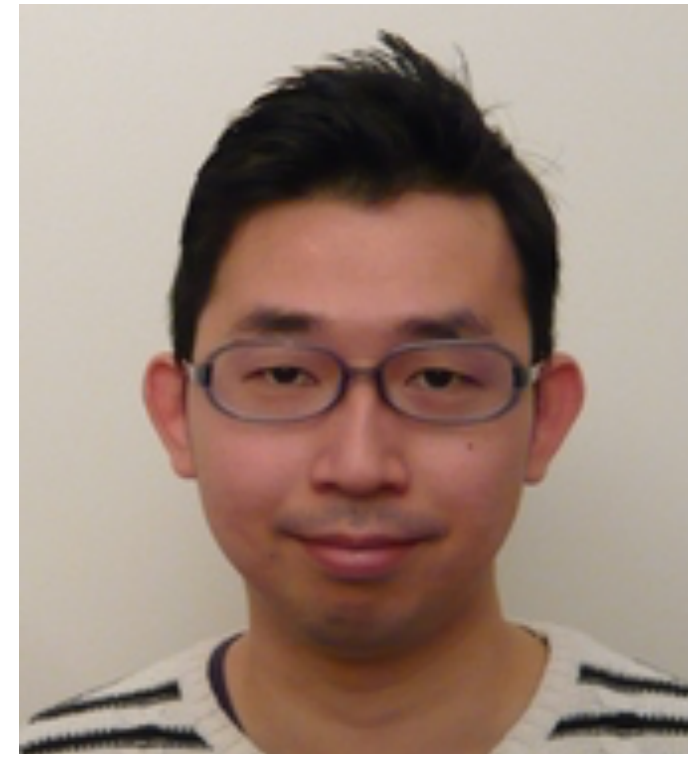
## PI + co-I



**Shin'ichiro Ando**



**Neal Dalal**



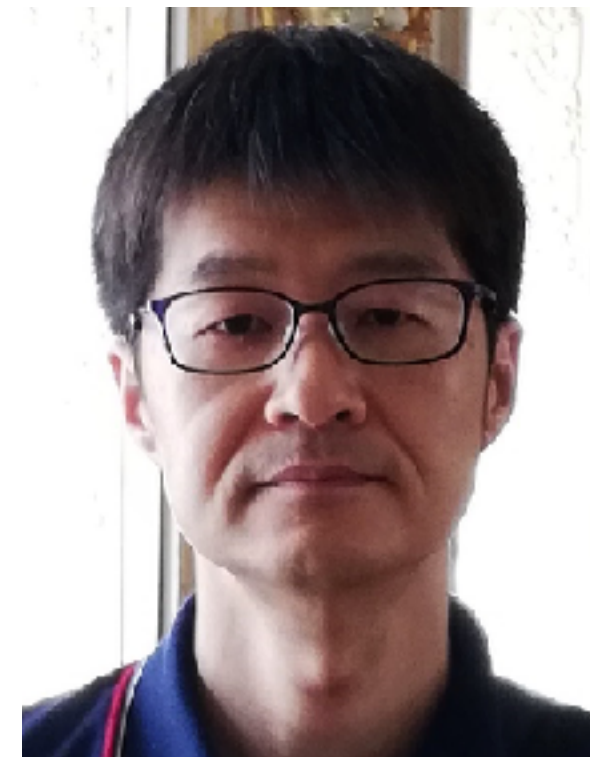
**Takahiro Nishimichi**



**Takashi Okamoto**



**Masato Shirasaki**

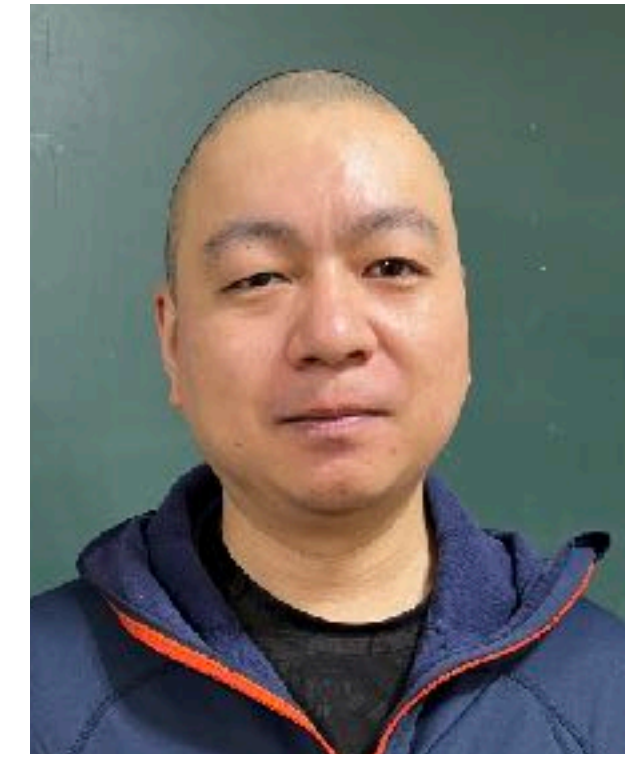


**Atsushi Taruya**

## Project postdocs

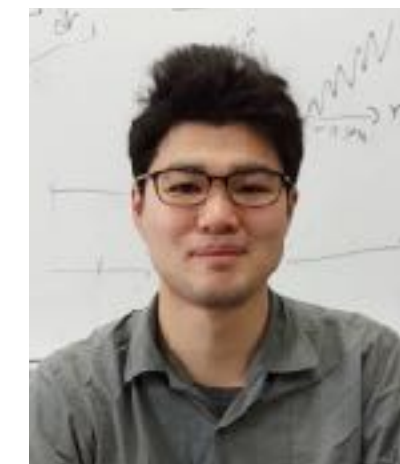


**Shunichi Horigome**

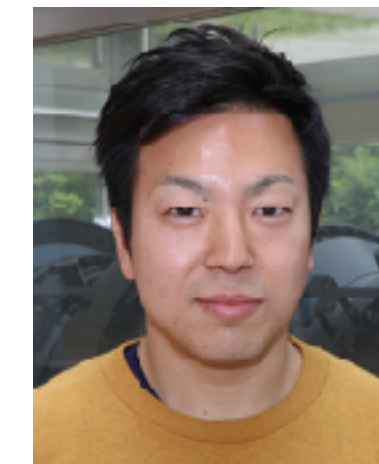


**Satoshi Tanka**

## Collaborators



**Y. Enomoto**



**K. Hayashi**



**M. Ono**

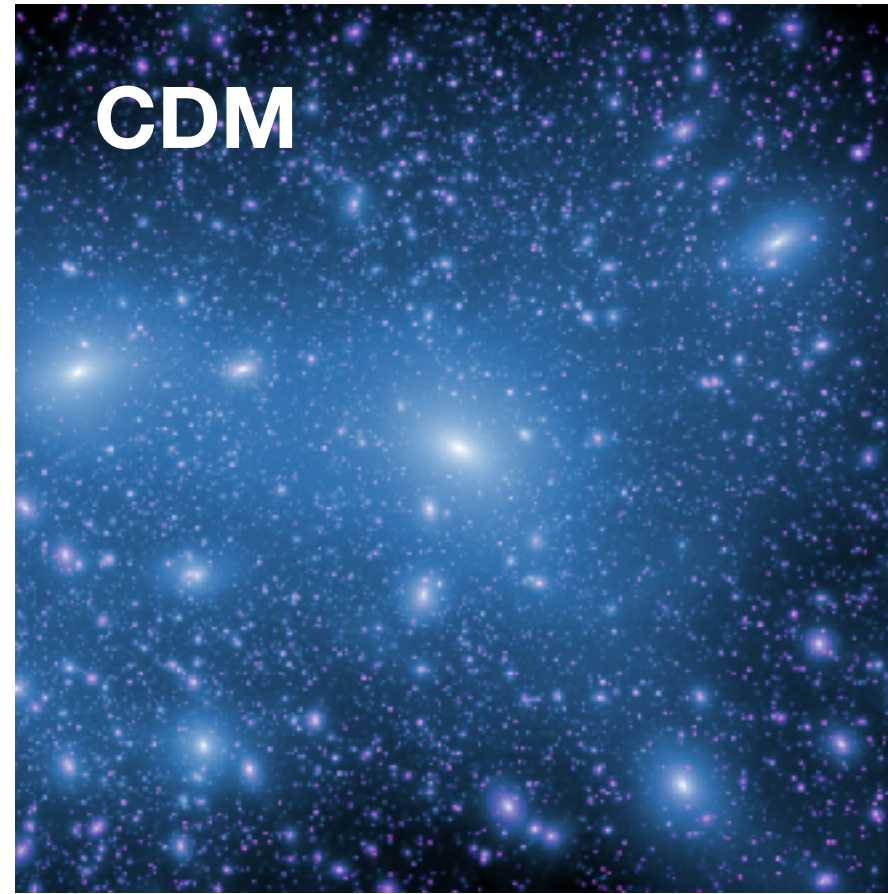


**N. Hiroshima**



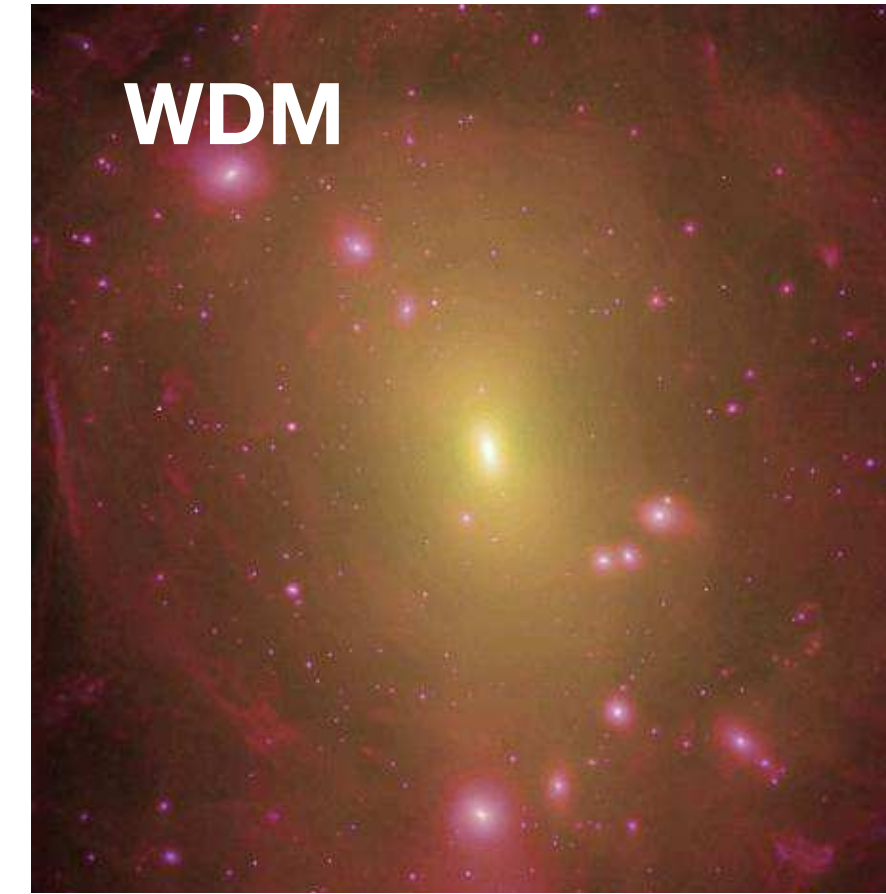
**T. Ishiyama**

# Goals: Small scale structure



**WIMPs, axions,  
ALPs, PBHs**

- Cusps in density profiles
- Very many small (sub)structures



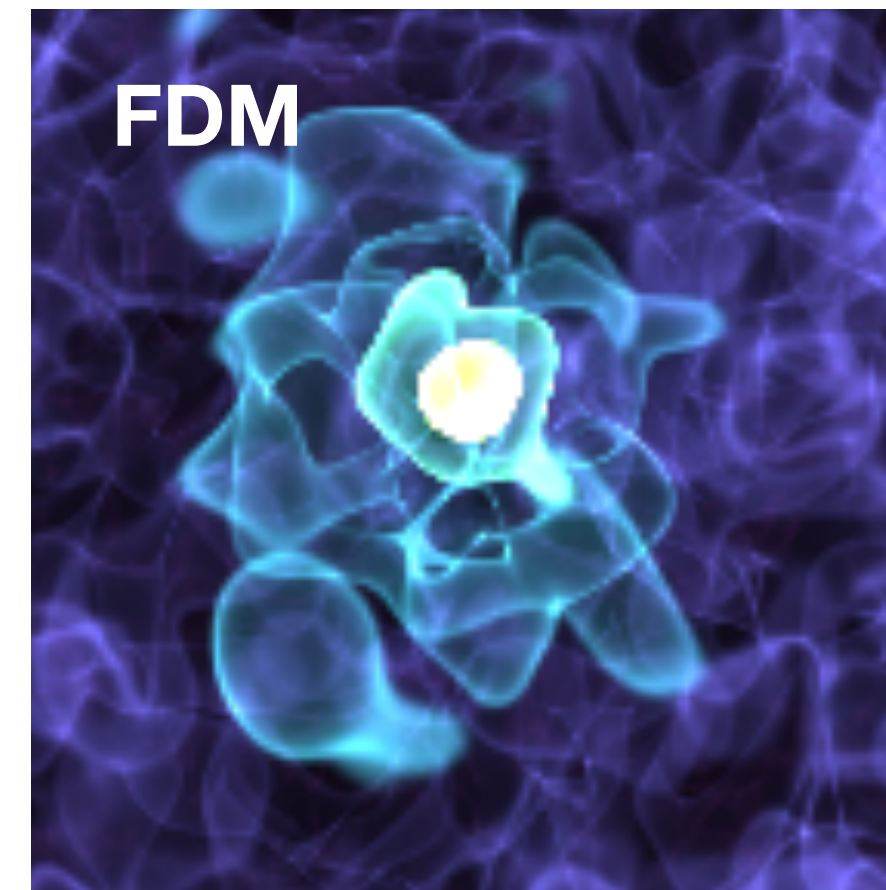
**Sterile neutrinos**

- Cutoff at sub-galaxy scale in the power spectrum



**SIMPs, dark atoms**

- Cores in density profiles induced by self scattering



**Ultralight bosons**

- Pattern induced by de Broglie length at sub-galactic scales

# Output: Small scale structure

✓ Release of public codes (Ando et al.)

✓ Dwarf galaxies (Horigome et al.)

✓ Galaxies and dark matter structure classification (Inoue, Okamoto et al.)

✓ Phase-space structure (Enomoto, Nishimichi, Taruya)

• **Primordial power (Tanaka, Nishimichi, Hiroshima)**

✓ Simulation of SIDM subhalos falling onto the SIDM halo (Shirasaki, Okamoto et al.)

• **Semi-analytical models of SIDM subhalos (Horigome, Ando)**

• **SIDM modeling against cosmological simulations (Shirasaki, Horigome, Ando)**

✓ Numerical simulations of WDM halos and subhalos (Okamoto, Inoue et al.)

✓ Developing semi-analytical models (Ando et al.)

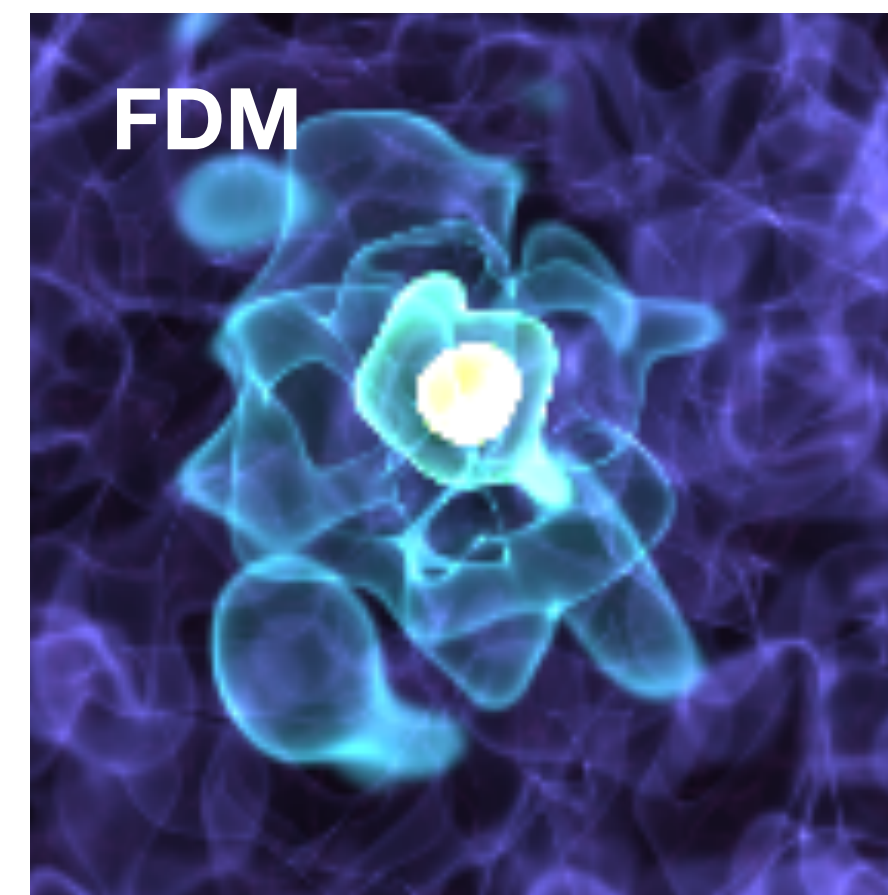
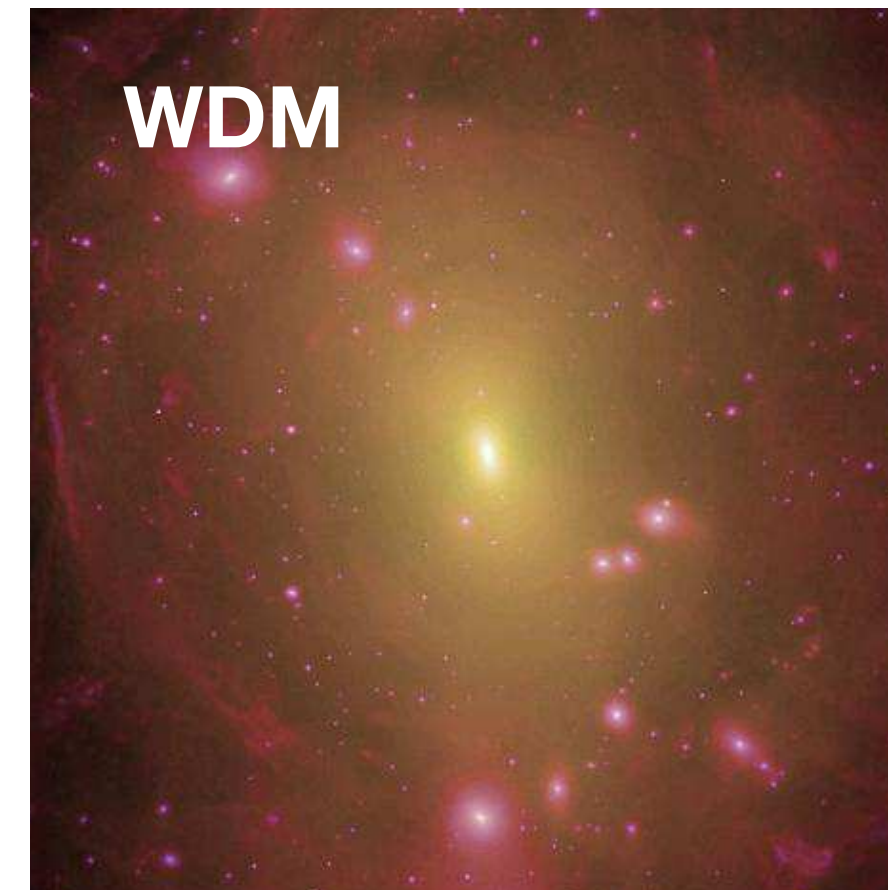
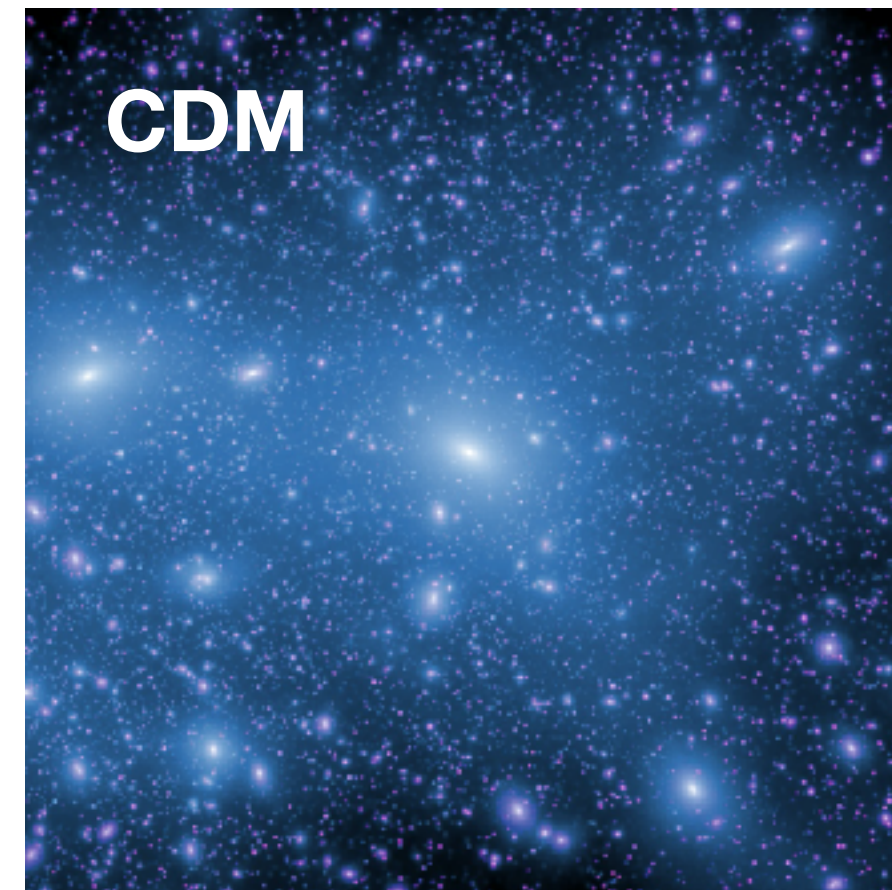
• **Phase-space structure (Enomoto, Nishimichi, Taruya)**

• **Calibration of semi-analytical models using numerical simulations (Ono, Okamoto, Ando)**

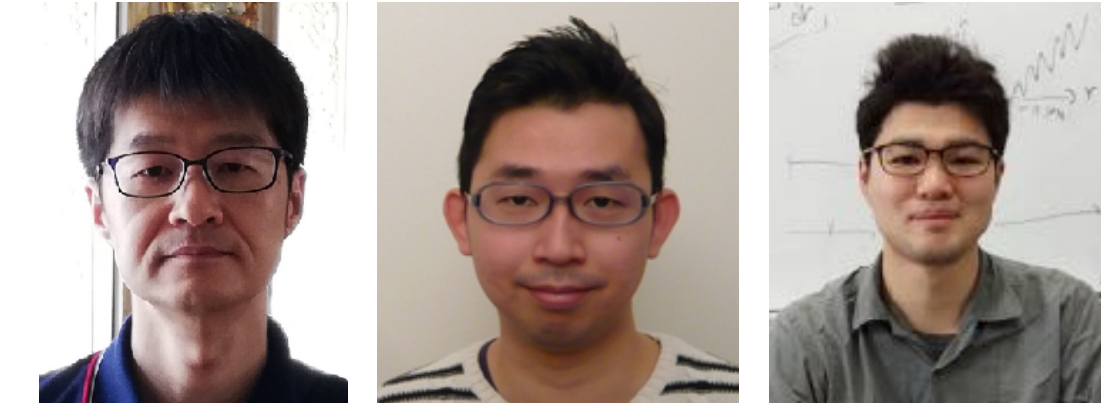
• **Novel approach using Vlasov equations (Tanaka, Taruya, Nishimichi)**

✓ Tight constraints using stellar motion in ultrafaint dwarf galaxies (Dalal)

• **Estimates of density profiles of dwarf galaxies (Horigome, Ando)**

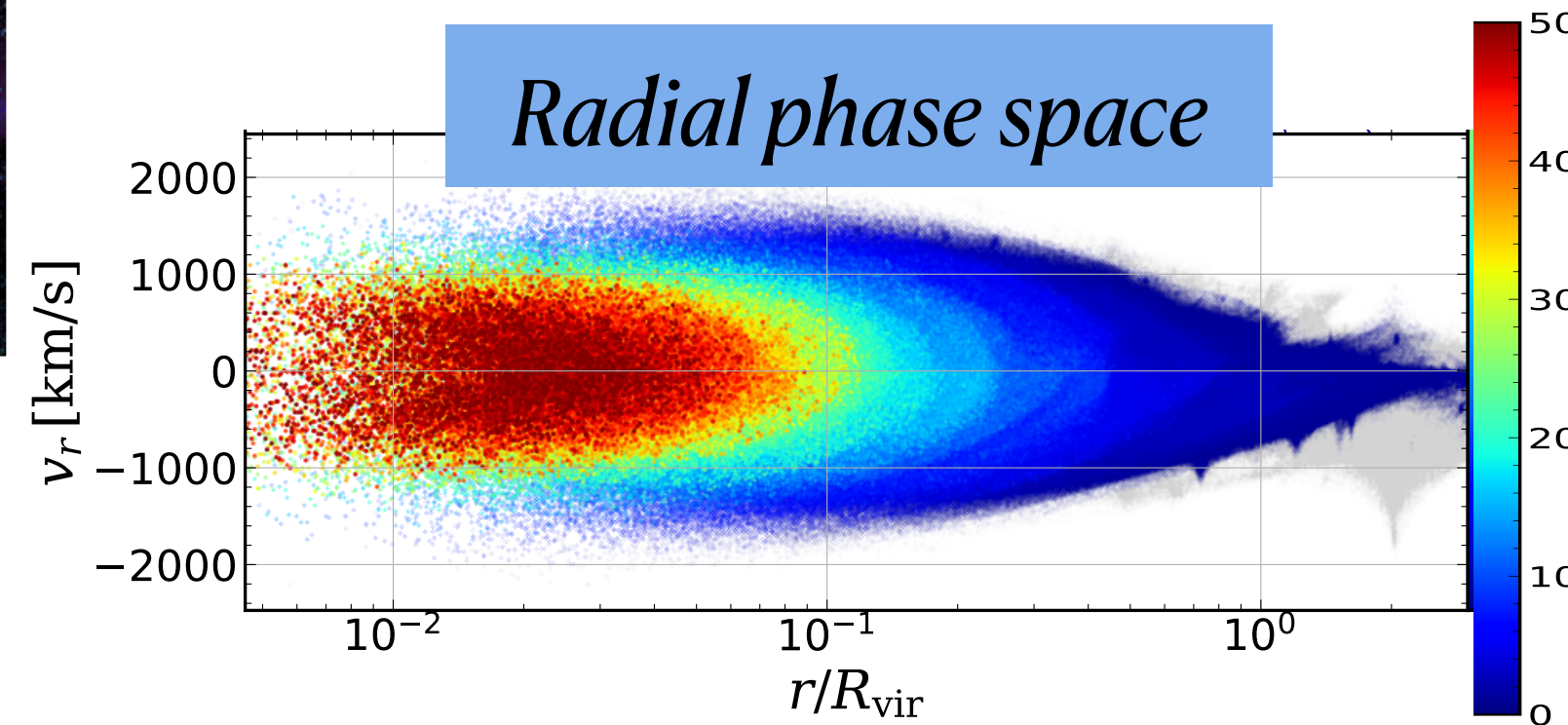


# Multi-stream radial structure of CDM halos



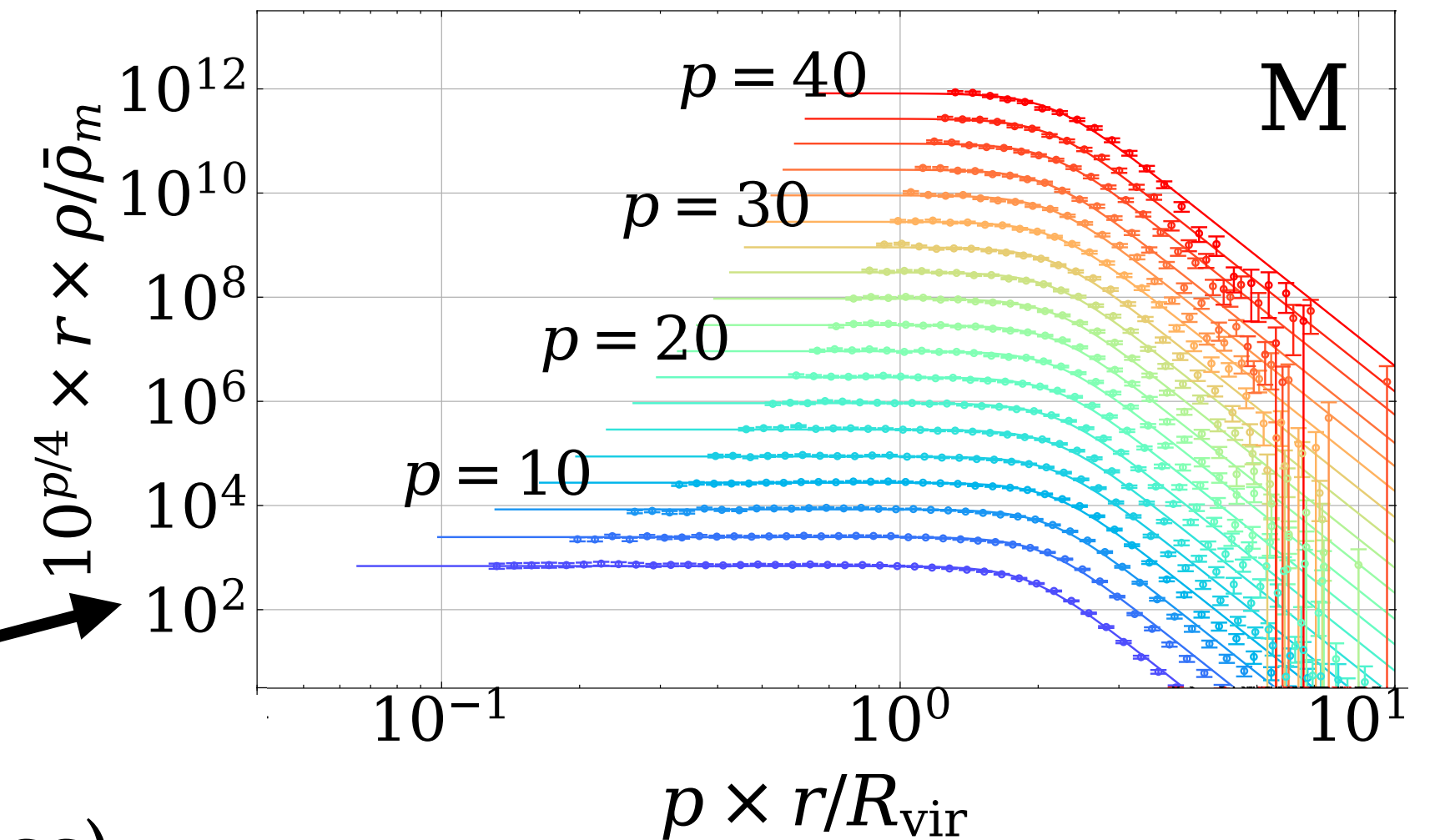
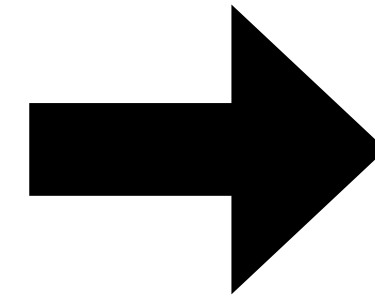
Enomoto, Nishimichi & Taruya, ApJ Lett. 950, L13 ('23), arXiv:2302.01531

MNRAS 527, 7523 ('24), arXiv:2309.13560



# of apocenter passages

$P \equiv$



Universal double power-law nature found here

also seems to persist in warm dark matter (in progress)

## Searching for axion DM from PEM data of gravitational wave (GW) detectors

Physical Environment Measurement (Work in progress)

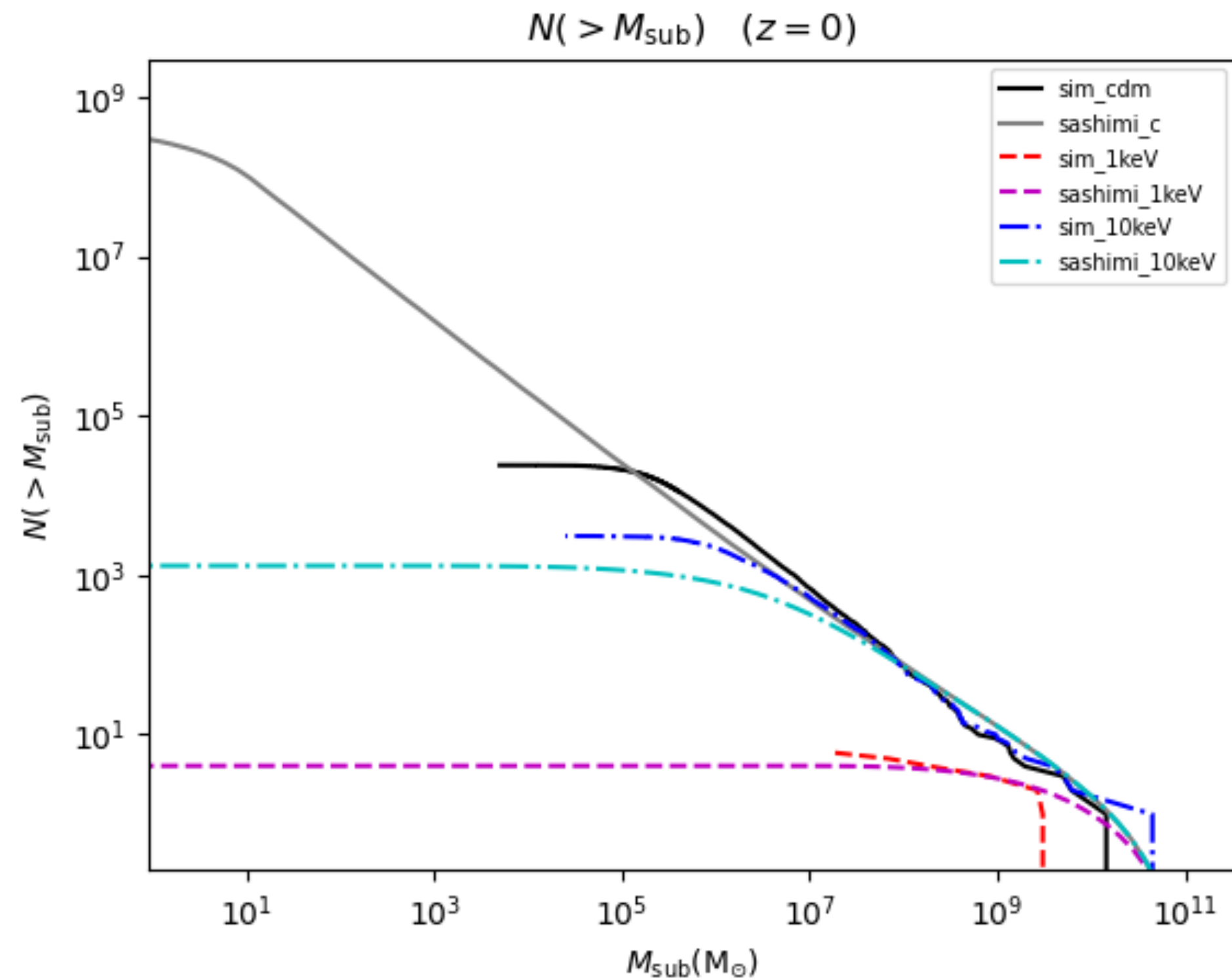
Axion DM of  $m_{\text{DM}} \sim 10^{-13}$  eV can produce low-freq. B-fields ( $f \sim 10$  Hz) in the presence of geomagnetic field  $\rightarrow$  PEM data of GW detectors helps to constrain axion-photon coupling (would be tighter than the CAST experiment)

# WDM simulations v.s. semi-analytic modelling

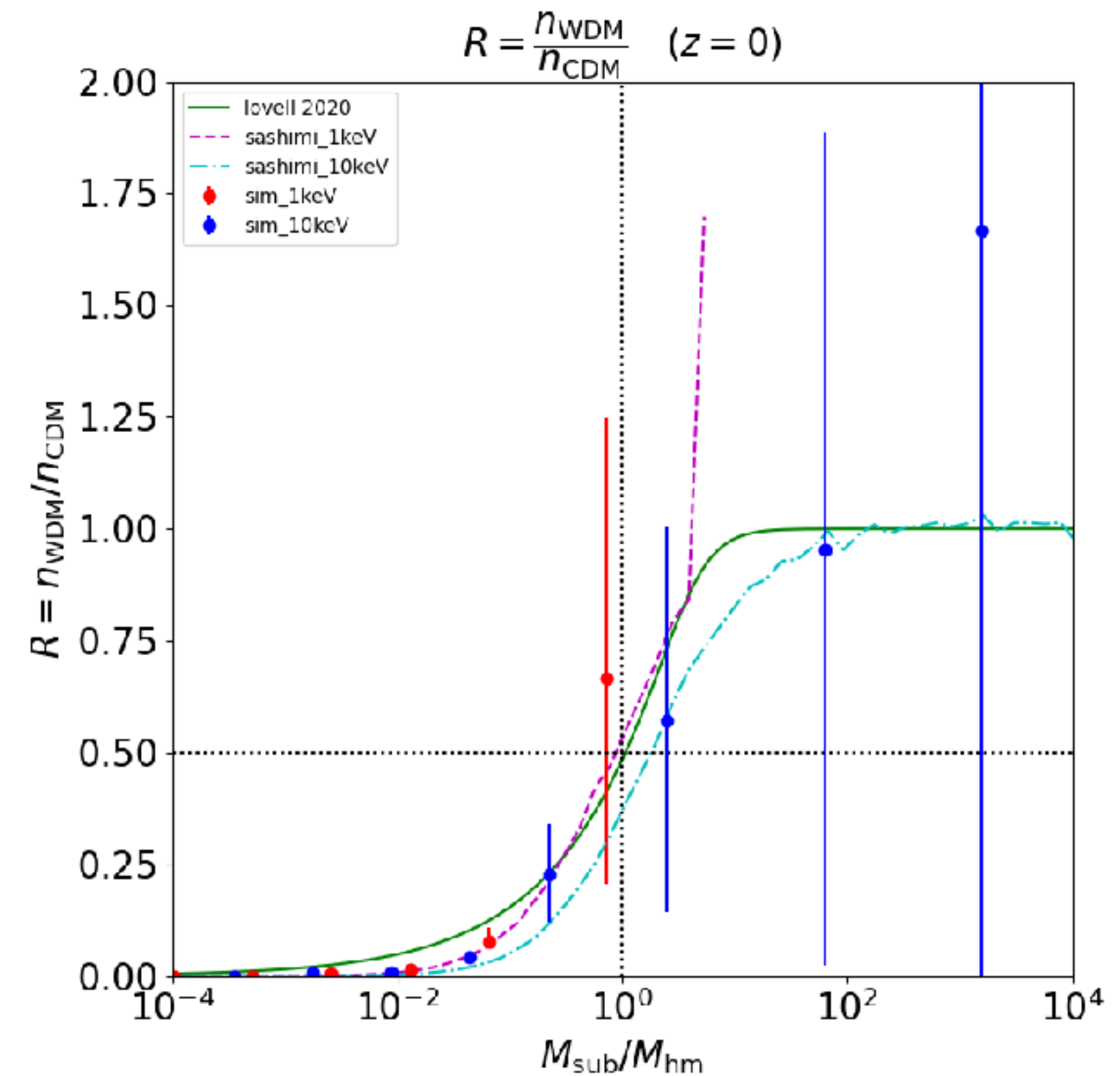
Mizuki Ono, Takashi Okamoto, Shin'ichiro Ando



Cumulative subhalo mass functions



Ratio of differential mass functions of WDM to that of CDM



# Vlasov-Poisson Simulation for Warm Dark Matter Halos

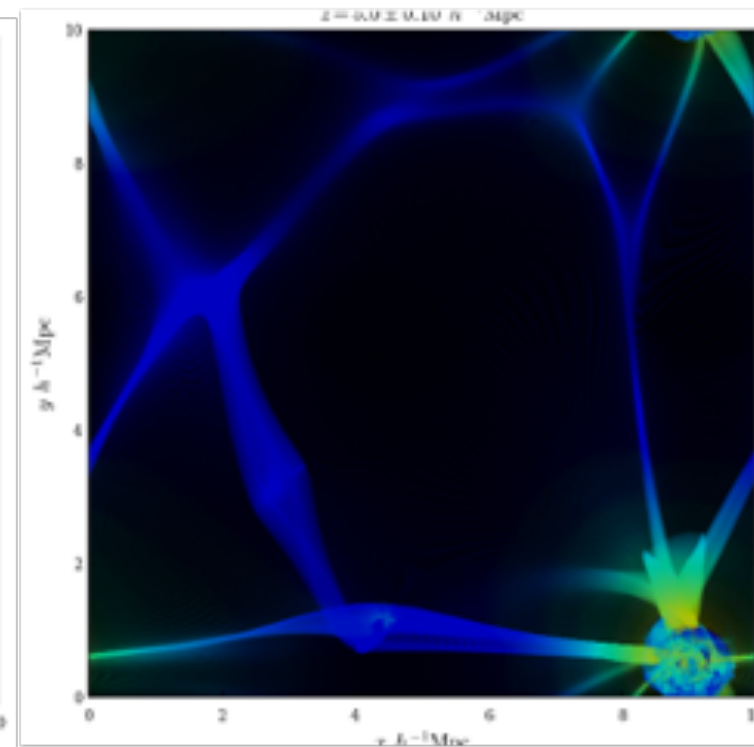
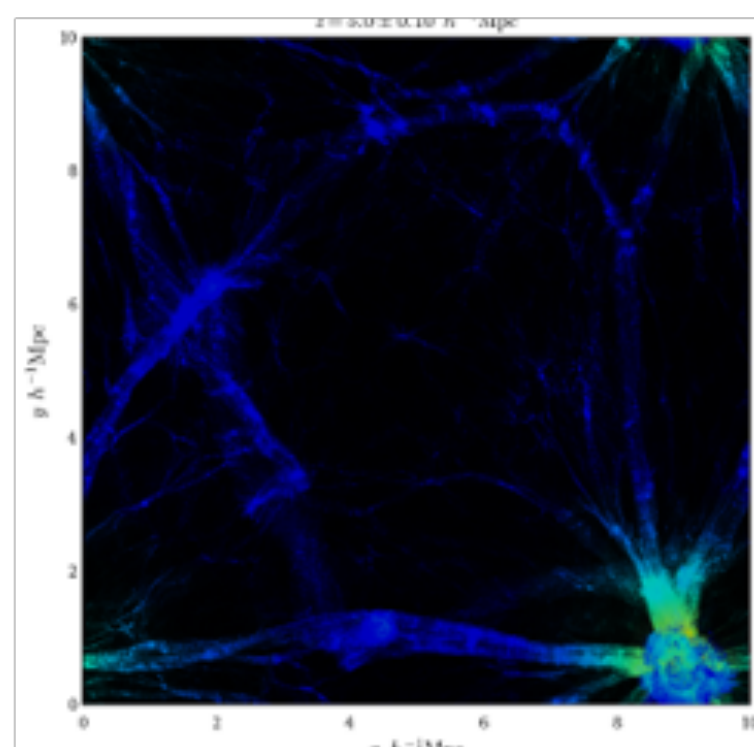
## Confirmation with N-body simulation

Satoshi Tanaka, Atsushi Taruya,  
Yohsuke Enomoto, Takahiro Nishimichi

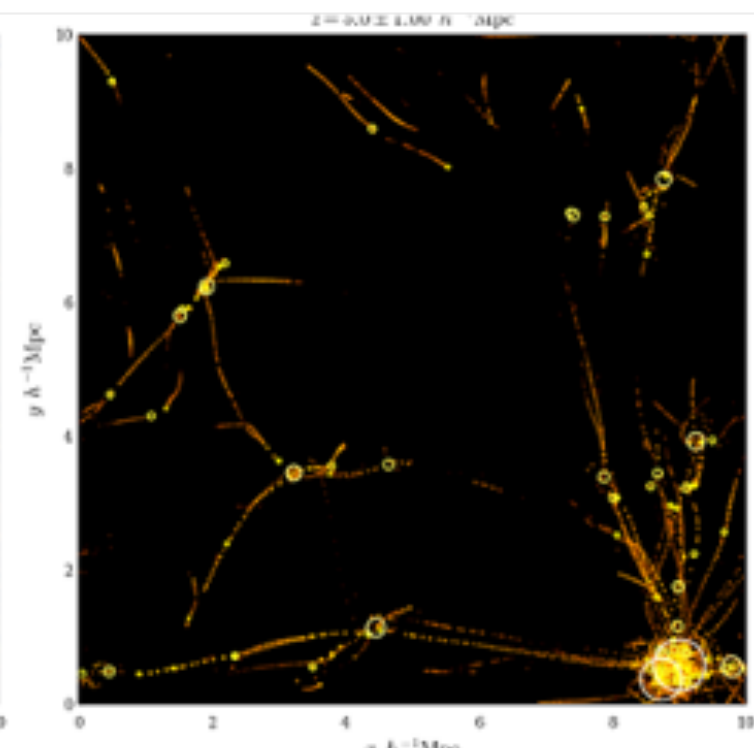
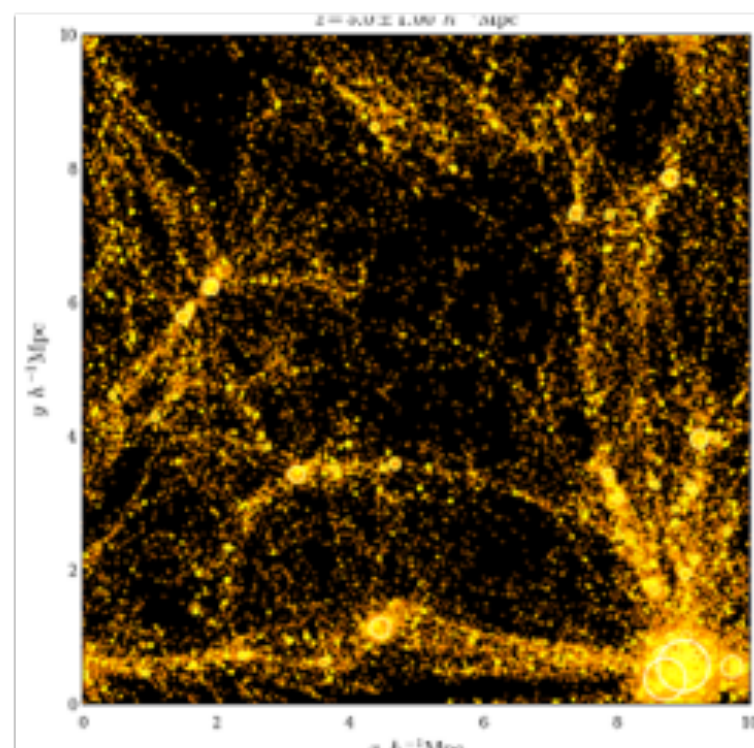
DM particles

CDM

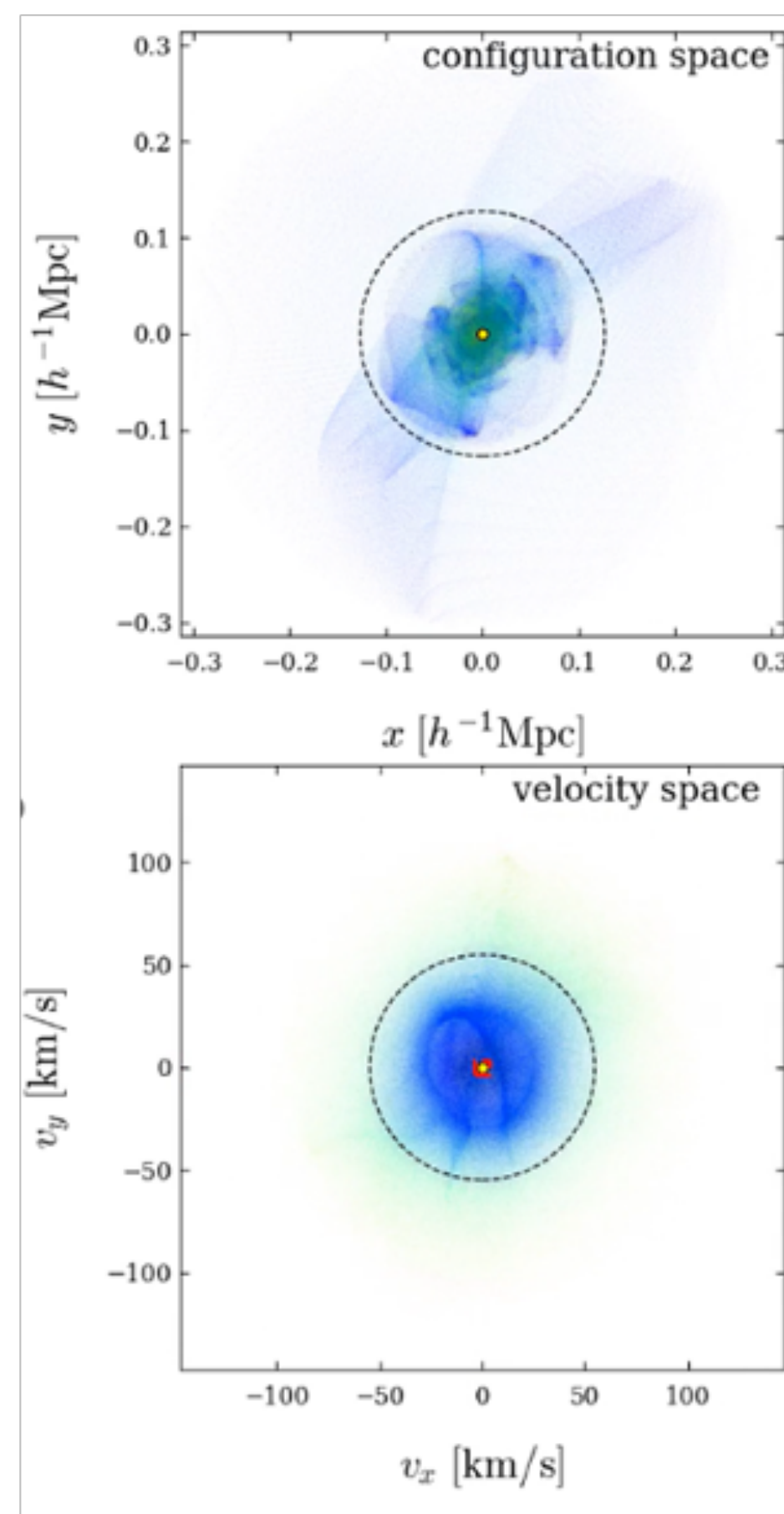
WDM 1keV



Halos

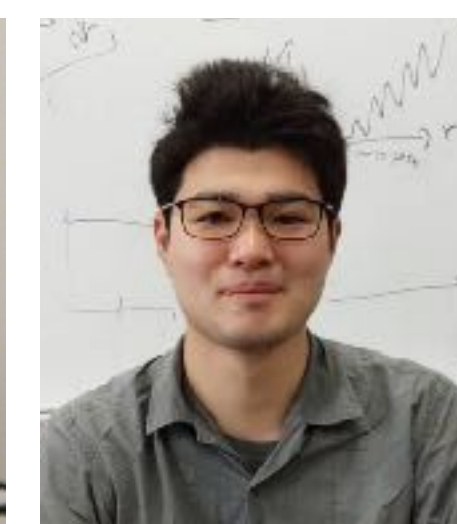
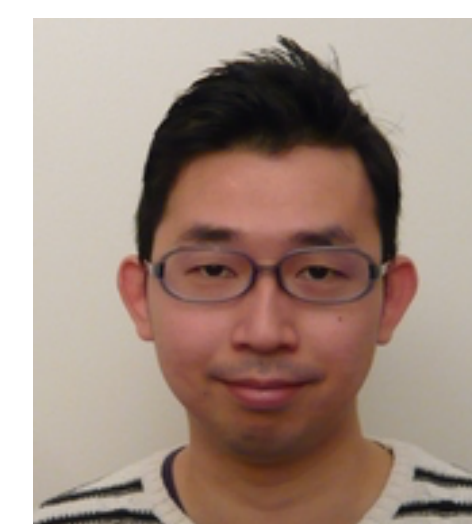
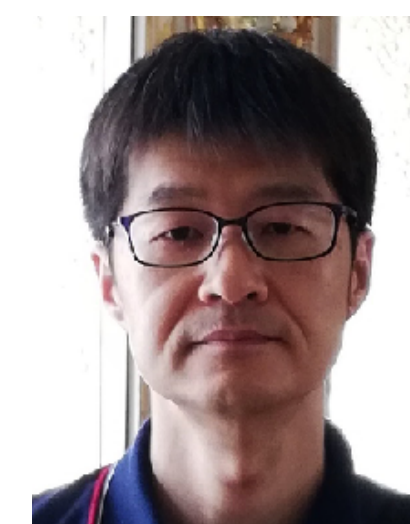


N-body simulation



Phase space of WDM halo

- Difficult to calculate N-body simulations to account for WDM velocity dispersion.
- We investigated conditions suitable for Vlasov simulations by the phase space from N-body simulations.
- Vlasov simulation may be able to handle the effects of WDM velocity dispersion until the halo prompt cusp is formed.

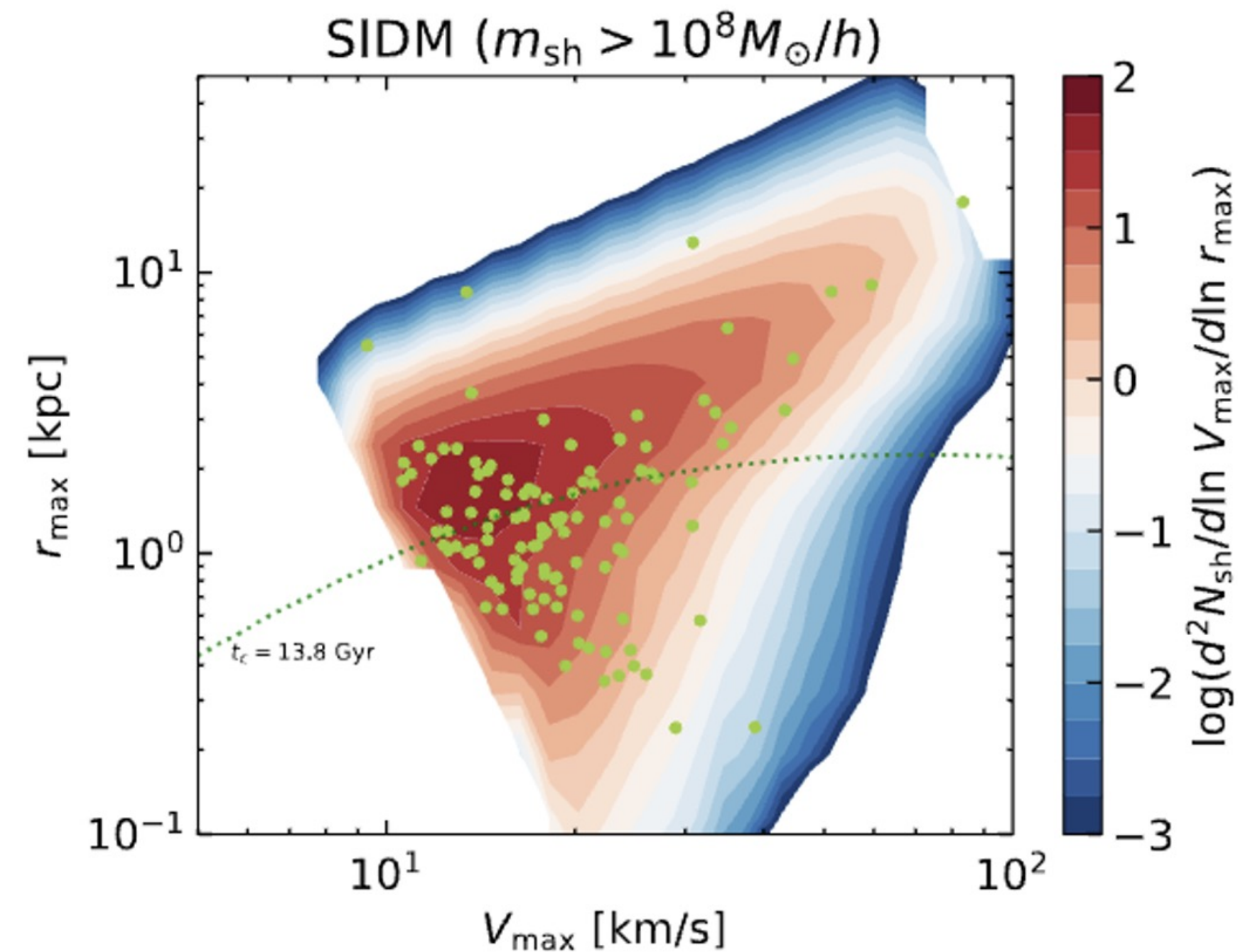
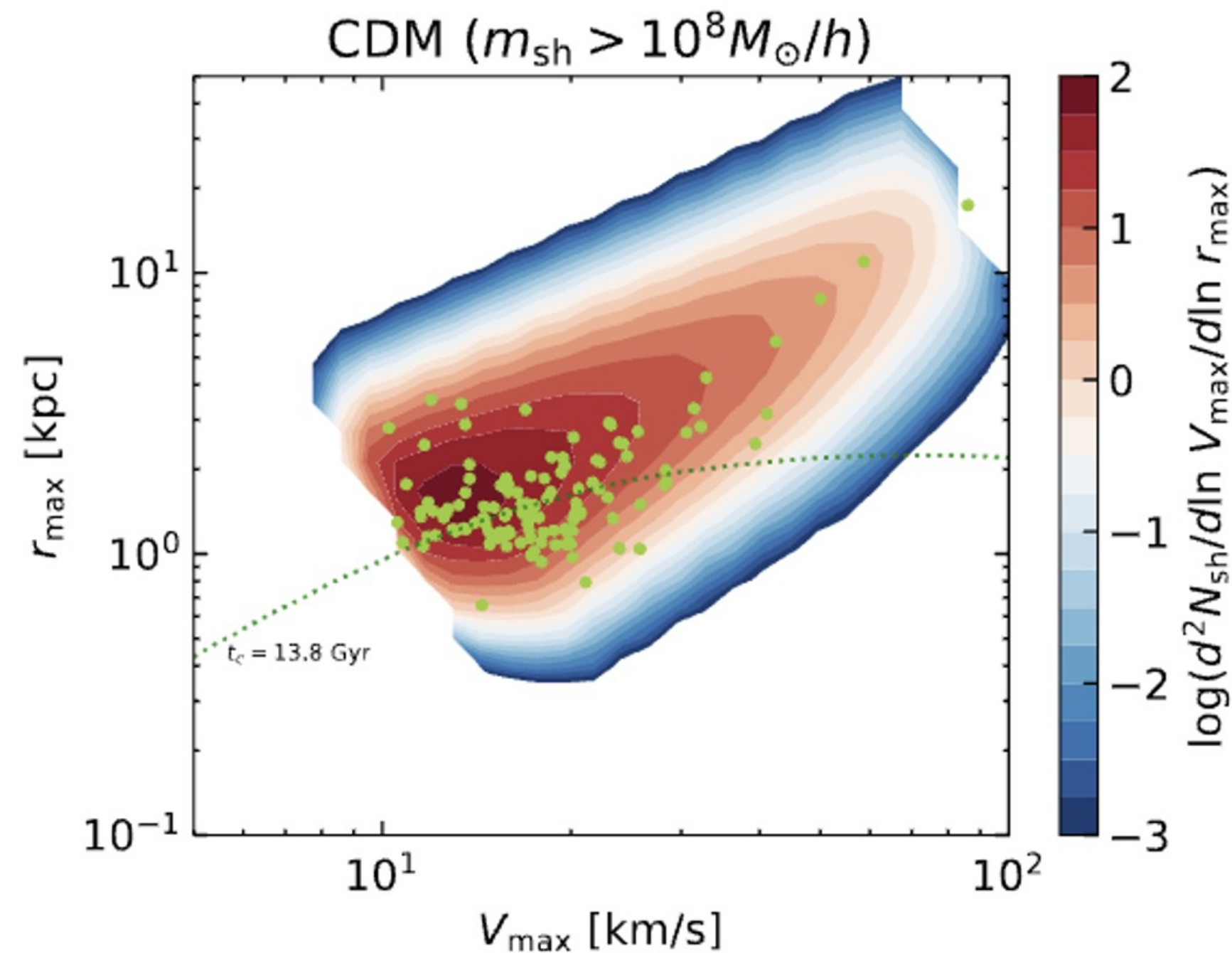


# SIDM halo structure and evolution: Semi-analytical and effective models

S. Ando, S. Horigome, E. O. Nadler,  
D. Yang, H.-B. Yu [in prep.]

SASHIMI-SIDM: Semi-analytical approach to simulate SIDM models

- Quick calculation of SIDM halo properties and subhalo mass functions





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