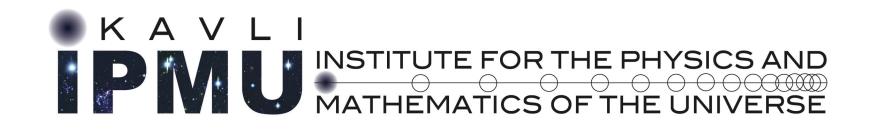
Axion detection using Helium-3

Berkeley Week, 13 March 2024

arXiv:2309.09160 In collaboration with **So Chigusa** (Berkeley), **Dan Kondo** (IPMU), Hitoshi Murayama (Berkeley, IPMU), Hiroyuki Sudo (ISSP)





Risshin Okabe (Kavli IPMU, U. Tokyo)



Strong CP problem **Dark Matter**



What we have shown

Axion-neutron interaction can be probed Evaluate application of +via a ferromagnetic phase of superfluid ³He quantum measurement

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Axion detection using superfluid Helium-3

Hypothetical particle Axion



Interacts with nucleons (e.g. neutrons, protons)

Berkeley Week 2024 (13 Mar. 2024)



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- Mixing between magnon and cavity photon
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Constraint on axion-neutron coupling by our method

Axion

Strong CP problem

- - Spontaneous symmetry breaking of $U(1)_{PO}$ generates a (pseudo-)Nambu-Goldstone boson = Axion
- $m_a = O(1 10) \mu eV$ is favored for DM axion

• Quantum chromodynamics (QCD) has a problem of unnaturalness

• This can be solved by introducing an additional symmetry $U(1)_{PO}$

Axion is also a candidate for Dark Matter (DM) because it has mass m_a

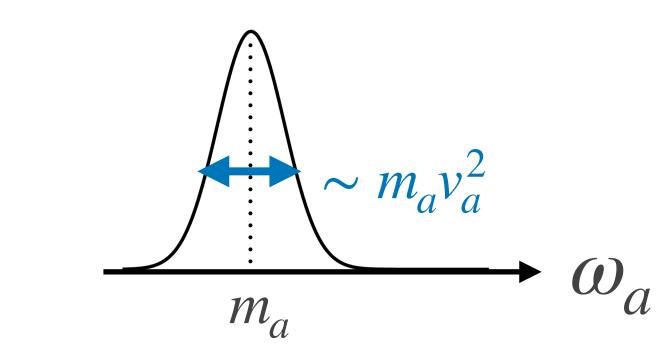
Axion as Dark Matter

- DM axion is non-relativistic and
 - behaves like classical waves oscillating coherently

$$a(t, \vec{x}) = A \sin[\omega_a(t - \vec{v}_a \cdot \vec{x}) + \varphi]$$

- Angular momentum: $\omega_a \simeq m_a$
- Coherence length $\simeq (m_a v_a)^{-1} \sim \text{km} \gg \text{typical laboratory size}$
- Coherence time $\simeq (m_a v_a^2)^{-1} \sim \text{msec}$

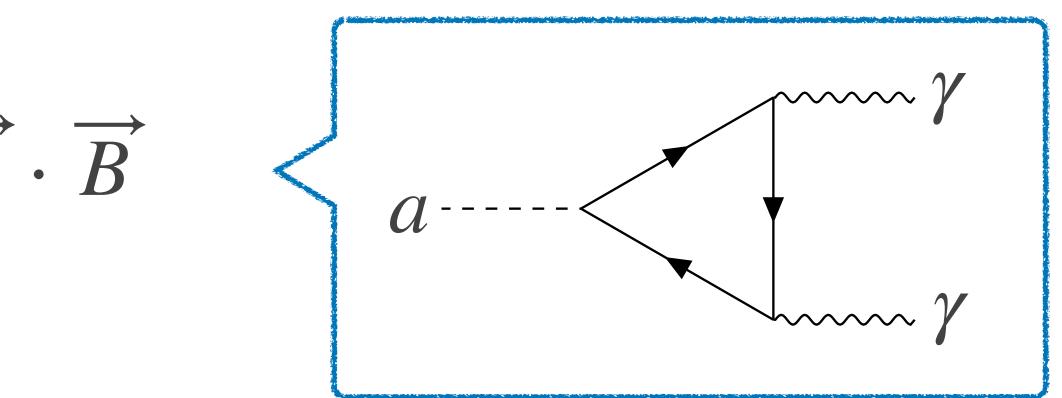
$$v_a + m_a v_a^2 / 2 \simeq m_a$$
 (:: $v_a \sim 10^{-3}$)

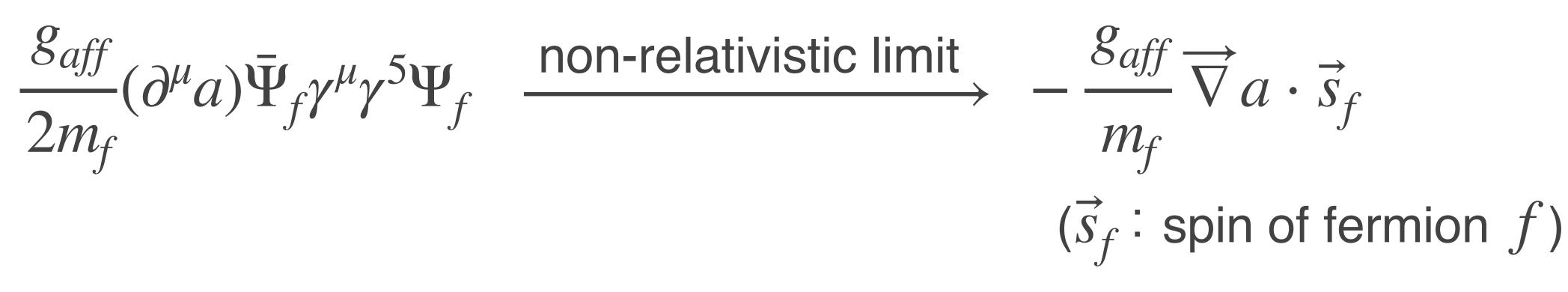


Interaction between axion and Standard Model particles Photon

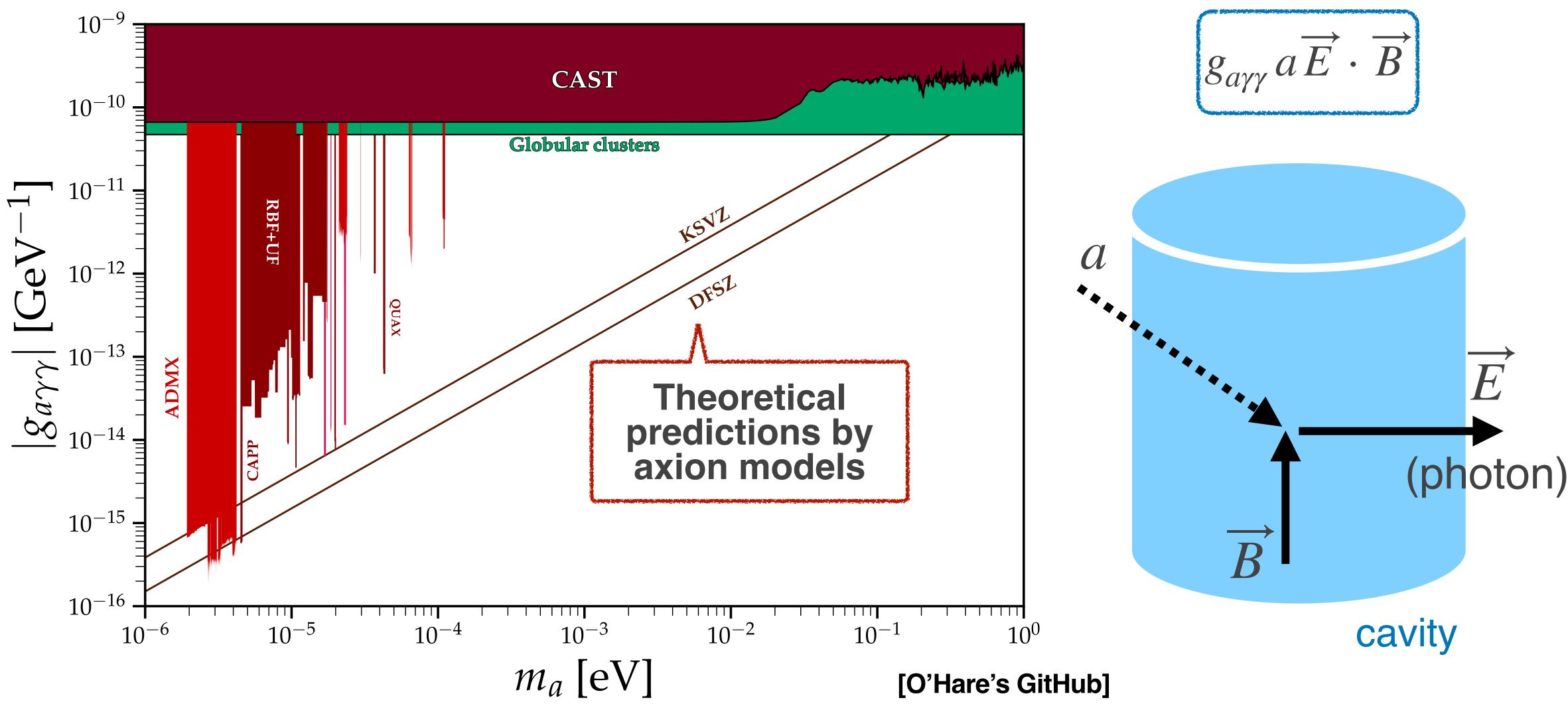
$$-\frac{1}{4}g_{a\gamma\gamma}a\tilde{F}^{\mu\nu}F_{\mu\nu} = g_{a\gamma\gamma}a\vec{E}$$

• Fermion





Constraints on axion-photon coupling

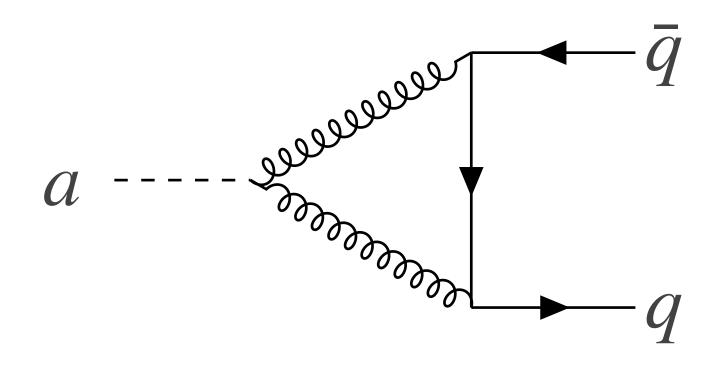


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Axion detection using superfluid Helium-3

Axion-nucleon coupling

Axion-nucleon coupling is an important property of axion

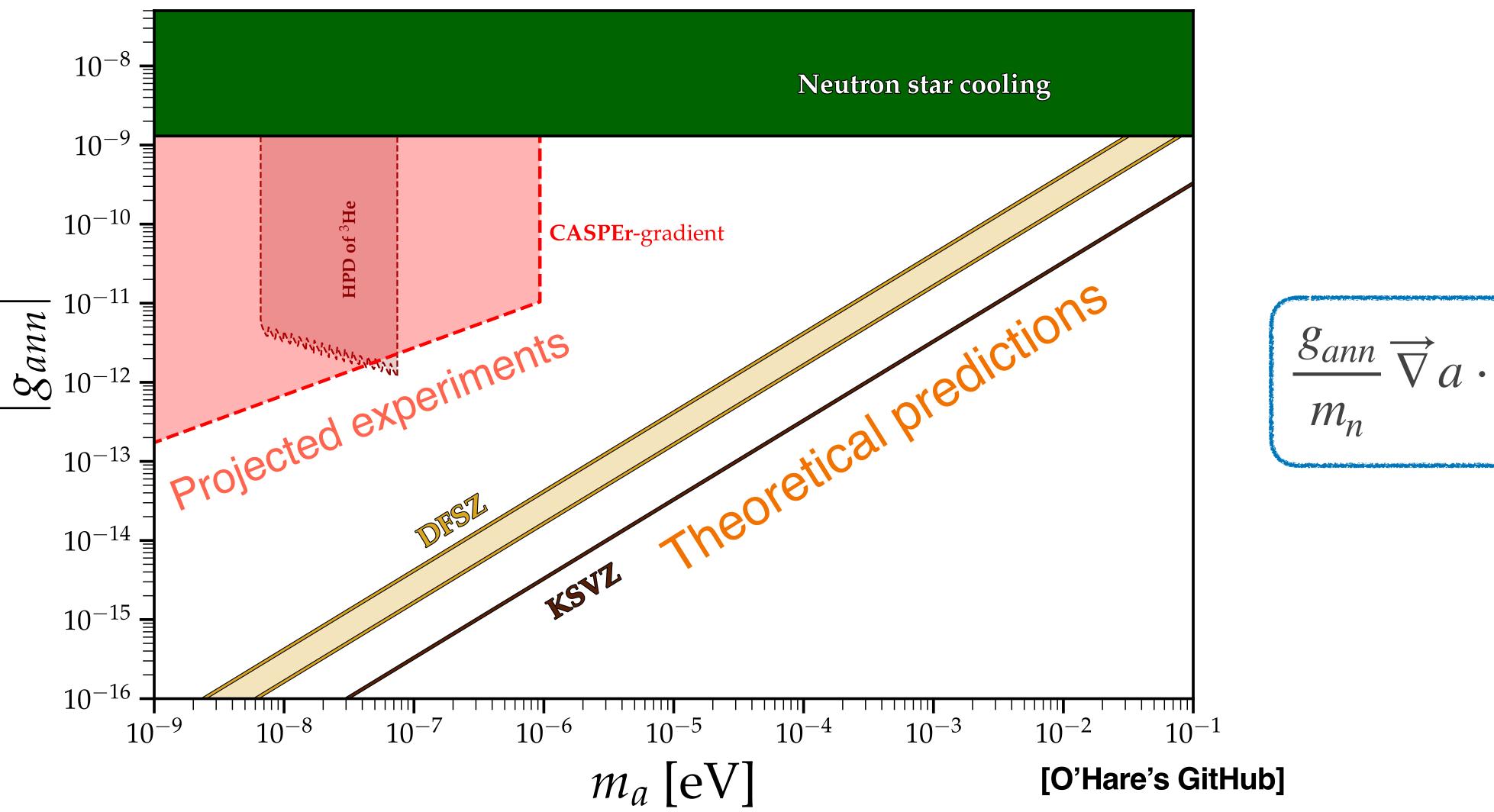


- Interaction with nucleons is crucial for axions, which is directly related to strong interaction
- photons or electrons
- However, axion-nucleon coupling has not been well probed

$$\frac{g_{aNN}}{m_N} \overrightarrow{\nabla} a \cdot \overrightarrow{s}_N$$

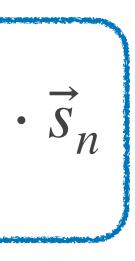
This coupling is less model-dependent than other couplings with

Constraints on axion-neutron coupling



Axion detection using superfluid Helium-3

Berkeley Week 2024 (13 Mar. 2024)



We propose a new experiment to explore axion-neutron coupling using superfluid Helium-3

Axion detection using superfluid Helium-3

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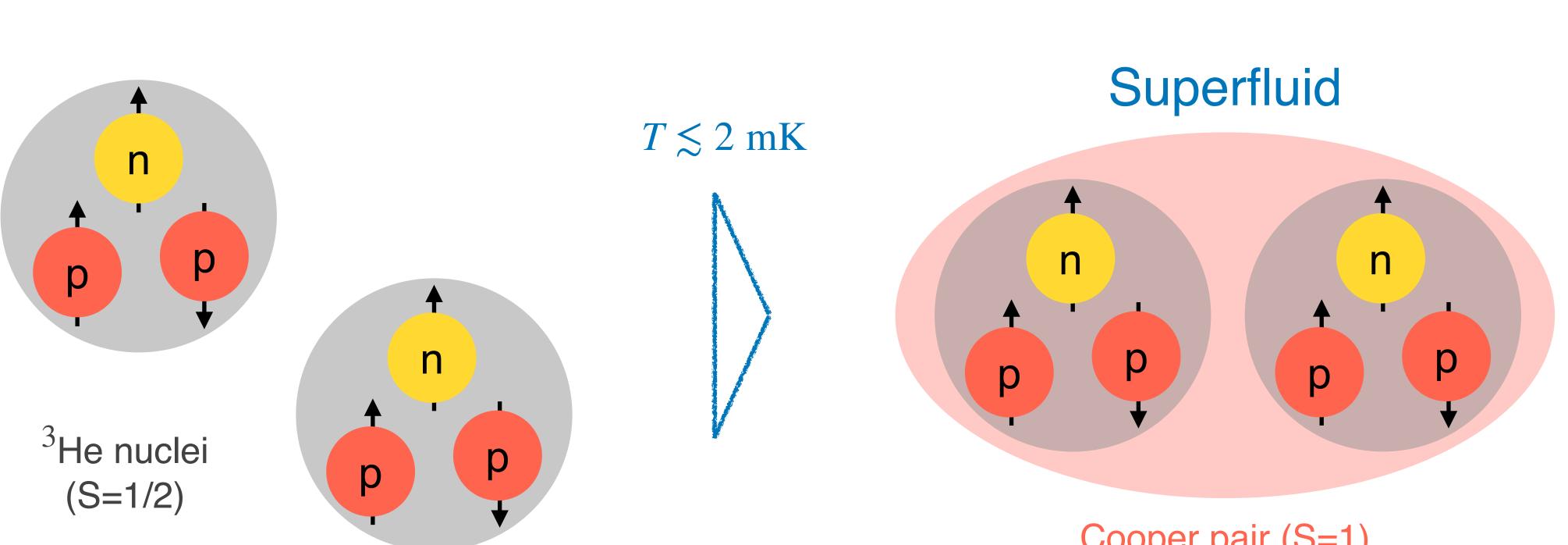
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Superfluid Helium-3

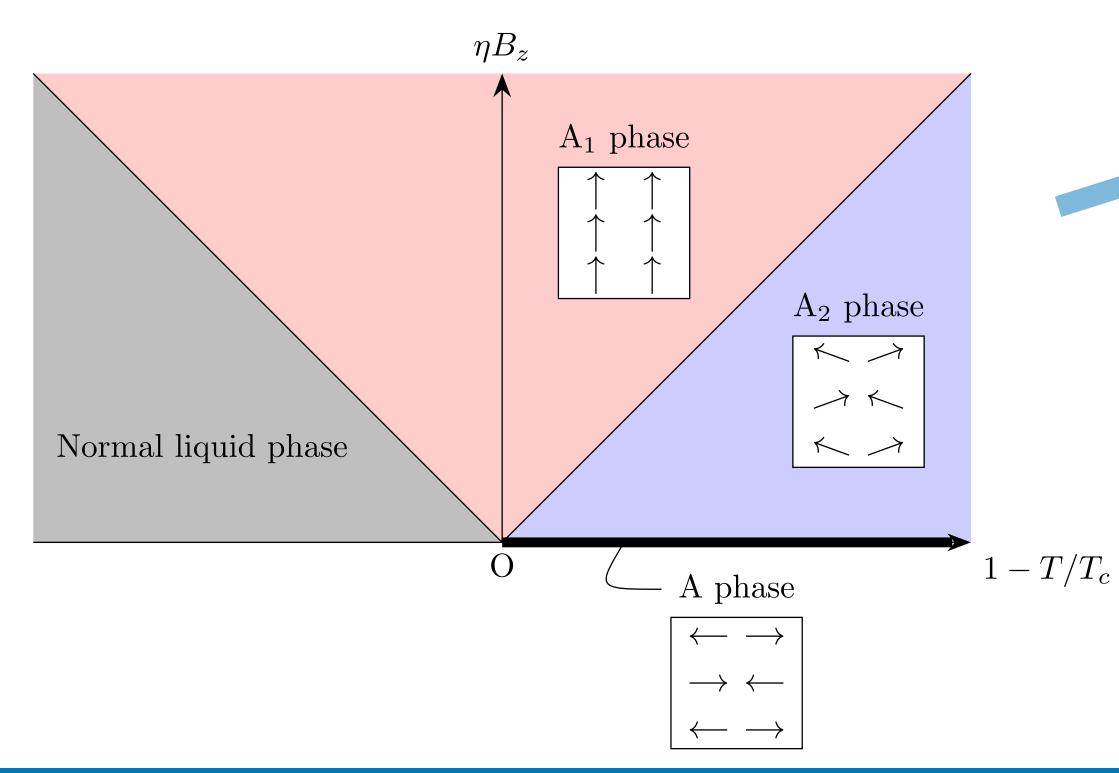


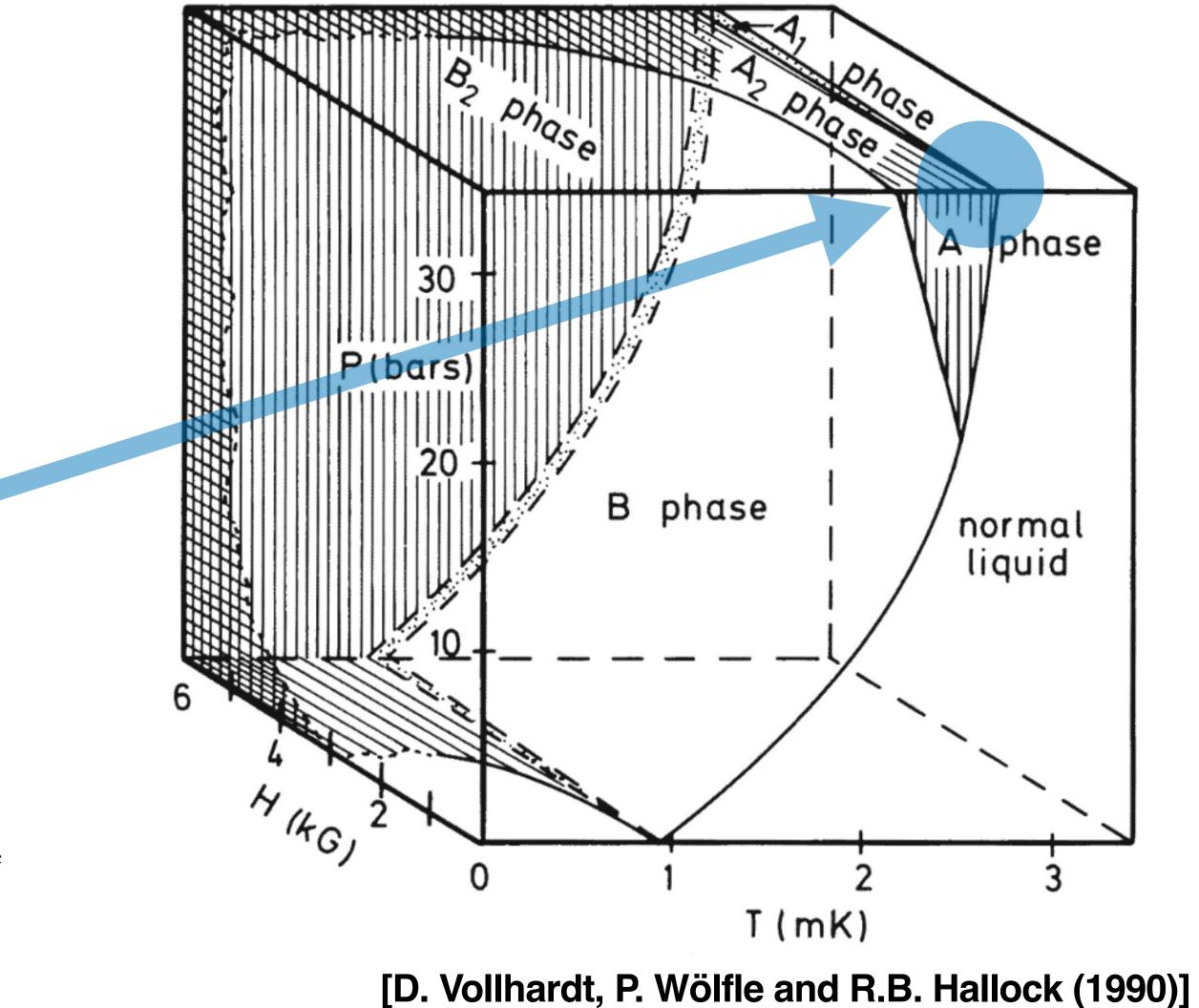
Spins of protons and electrons are canceled Interaction between axion & 3 He \simeq interaction between axion & **neutron**

Cooper pair (S=1)

A1 phase of superfluid ³He

- Ferromagnetic for nuclear spin
 - Spin polarization ~ few %
- has a magnon mode





Magnon

- Quasiparticle (collective excitation) of the spin structure
- Example: ferromagnet
 - Ground state $\langle S_7 \rangle \neq 0$
 - Final Two broken generators S_{r} and S
 - One "Type-B" Nambu-Goldstone mode = gapless magnon mode

\Leftrightarrow spin-rotation symmetries along x, y axes are spontaneously broken

$$S_{y}: \langle [S_{x}, S_{y}] \rangle \propto \langle S_{z} \rangle \neq 0$$

[Watanabe & Murayama (2012), Hidaka (2013)]

• External magnetic fields induce energy gap $\omega_L \equiv \gamma B_{\gamma}$ (Larmor frequency)



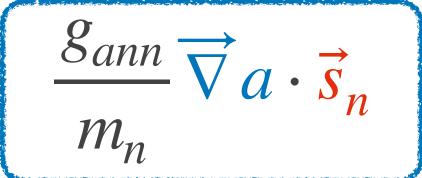


Magnon excitation by axions

- - Resonantly enhanced when $m_a = magnon energy gap$
- Larmor frequency of ³He : $\omega_L \simeq 1.3 \,\mu \text{eV} \left(\frac{B_z}{10 \,\text{T}}\right)$
 - By scanning magnetic fields of O(10) T, axions with μeV mass can be probed using the resonance

Axion-neutron interaction induces (homogeneous) magnon modes

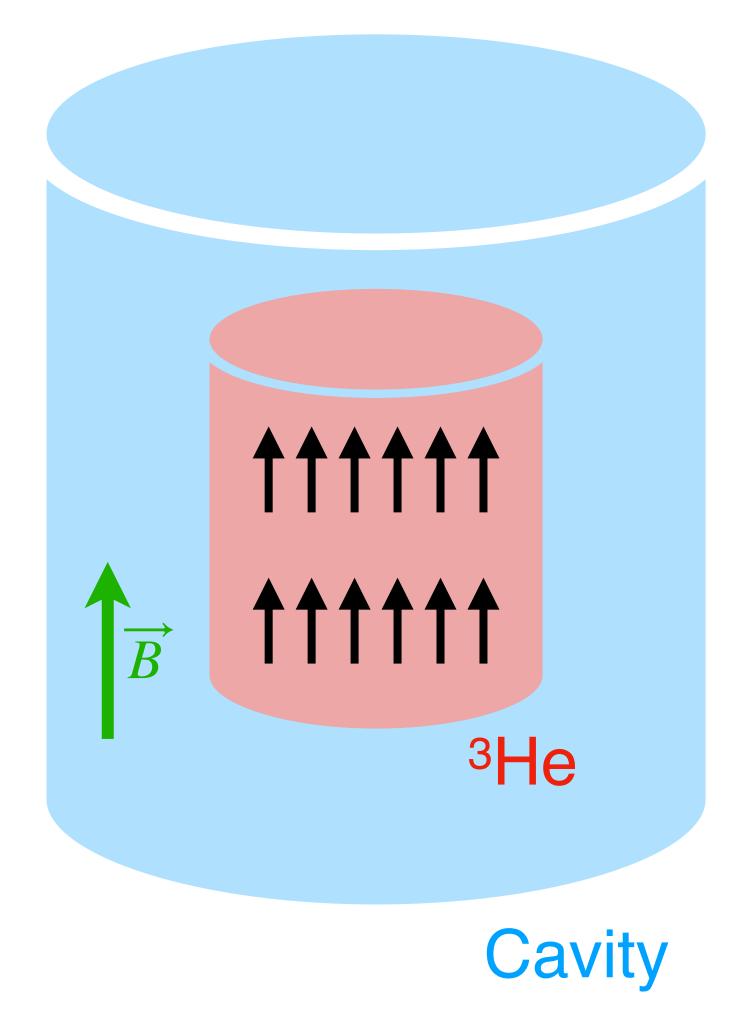








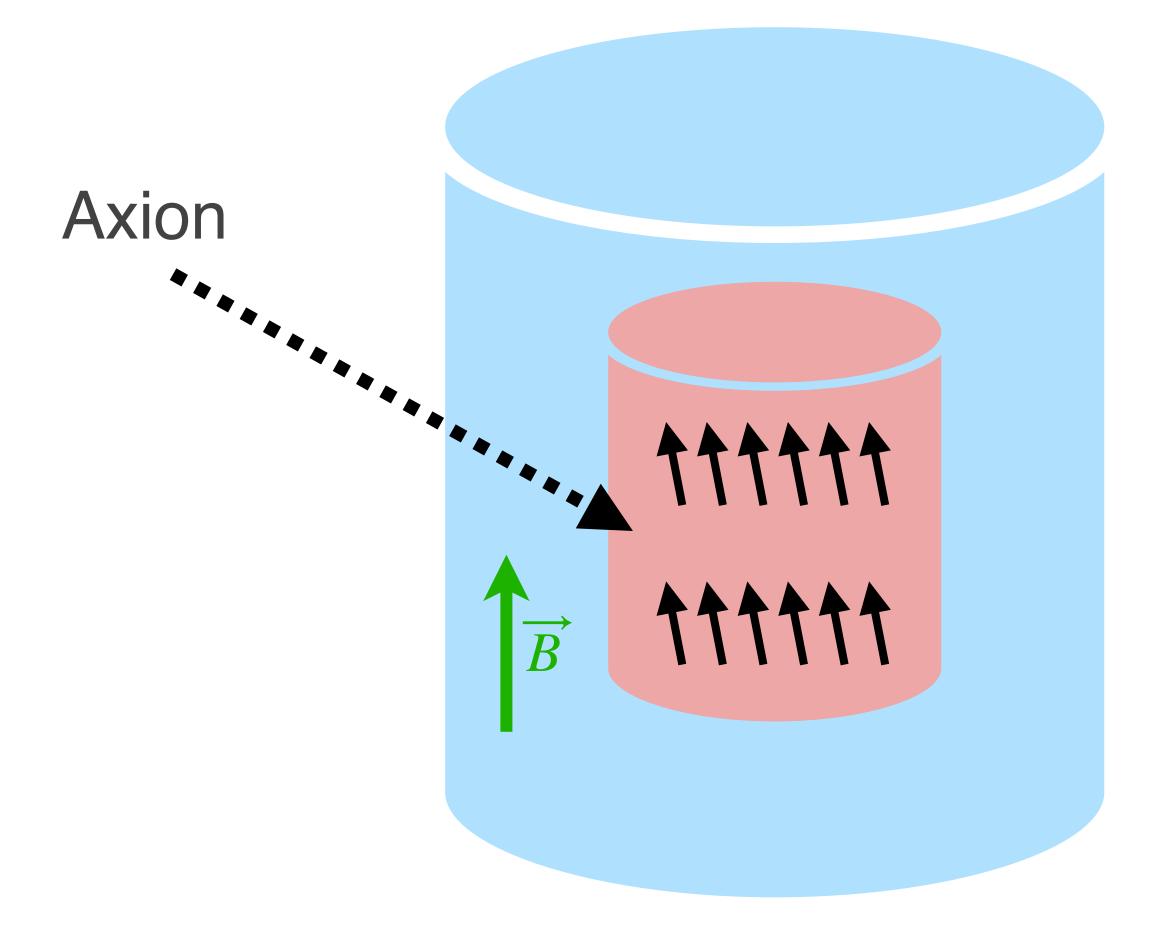
Experimental setup



- 1. Put a superfluid ³He sample in a cavity
- 2. Apply a magnetic field to align nuclear spins



Experimental setup



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Put a superfluid ³He sample in a cavity

2. Apply a magnetic field to align the nuclear spins

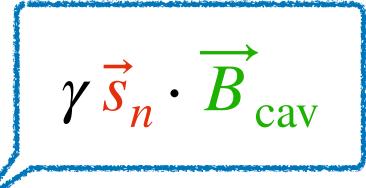
3. Axion excites "magnon" modes

How to detect magnons

- Magnon modes mix with cavity photons
 - "magnon polariton"
- Mixing is maximized when

Larmor frequency $\omega_L = cavity$ frequency ω_{cav} Scan ω_L and ω_{cav} with $\omega_L = \omega_{cav} = \mathcal{O}(1) \, \mu eV$

- Typical size of cavity

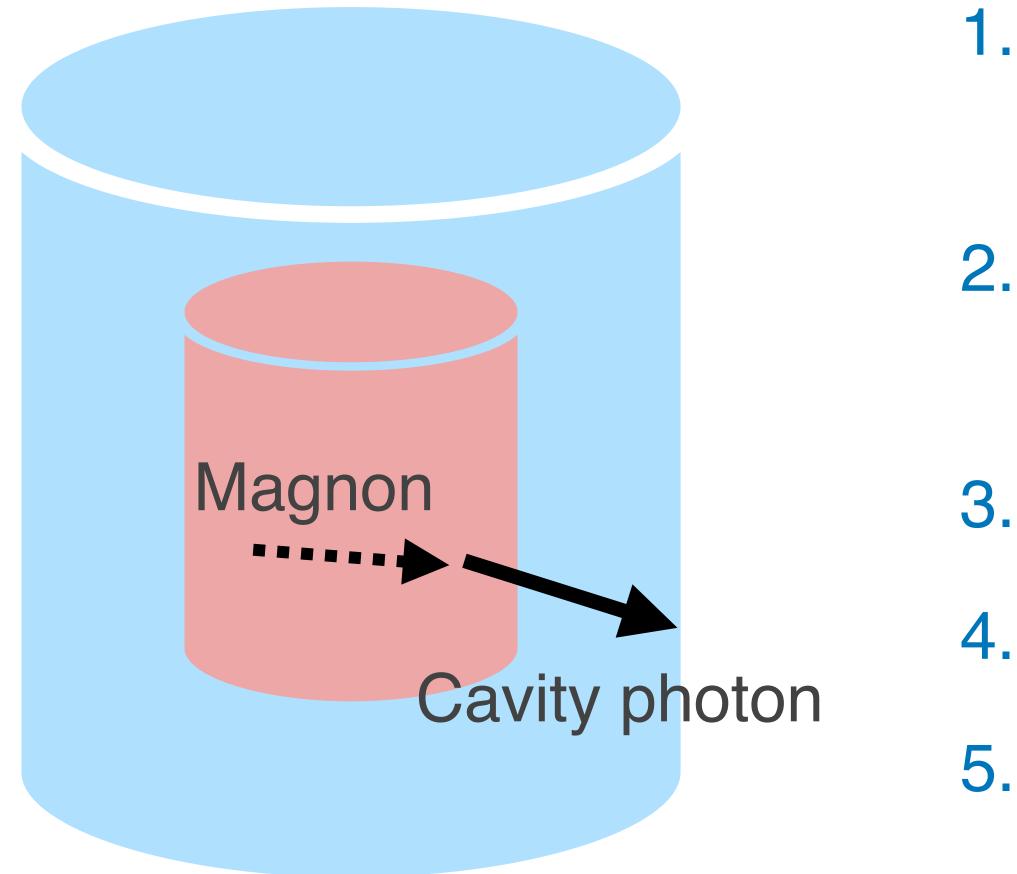


through interaction between nuclear spins and magnetic fields

 $2\pi/\omega_{cav} \simeq 1.2 \,\mathrm{m} \,(m_a/\mu \mathrm{eV})^{-1}$



Experimental setup



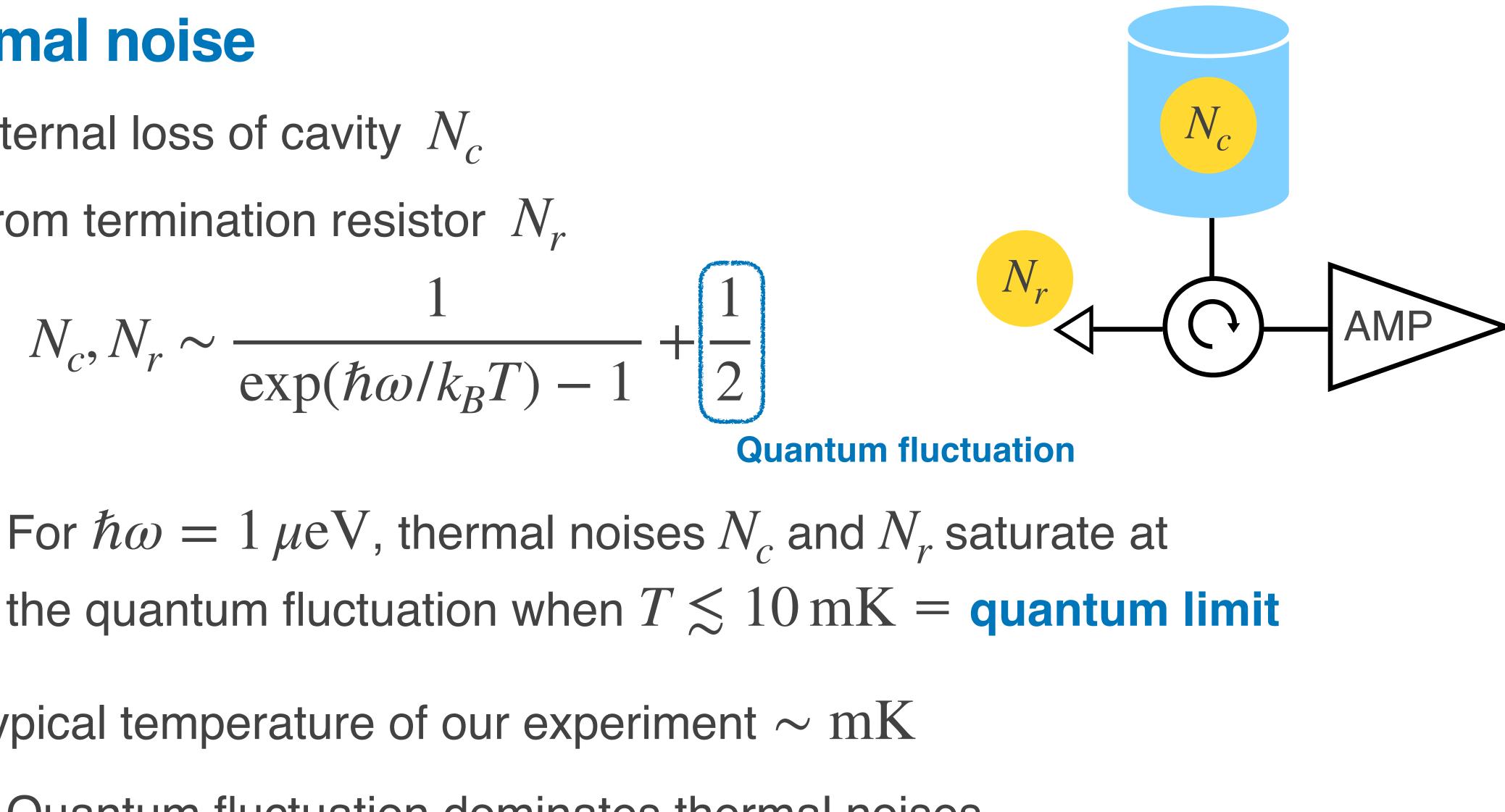
- 1. Put a superfluid ³He sample of the A1 phase in a cavity
- 2. Apply a magnetic field to align the nuclear spins
- **3.** Axion excites "magnon" modes
- 4. Magnons mix with cavity photons
- 5. Amplify and measure the signal (cavity photons)

Thermal noise

- Internal loss of cavity N_c
- From termination resistor N_r

$$N_c, N_r \sim \frac{1}{\exp(\hbar\omega/k_B T) - 1}$$

- For $\hbar\omega = 1 \,\mu eV$, thermal noises N_c and N_r saturate at
- Typical temperature of our experiment $\sim mK$
 - Quantum fluctuation dominates thermal noises



Quantum measurement techniques

- Circumvent quantum noise by quantum measurement techniques
- Quadratures

$$\hat{X} = \frac{\hat{a} + \hat{a}^{\dagger}}{\sqrt{2}}, \quad \hat{Y} = \frac{\hat{a} - \hat{a}^{\dagger}}{\sqrt{2}i}$$

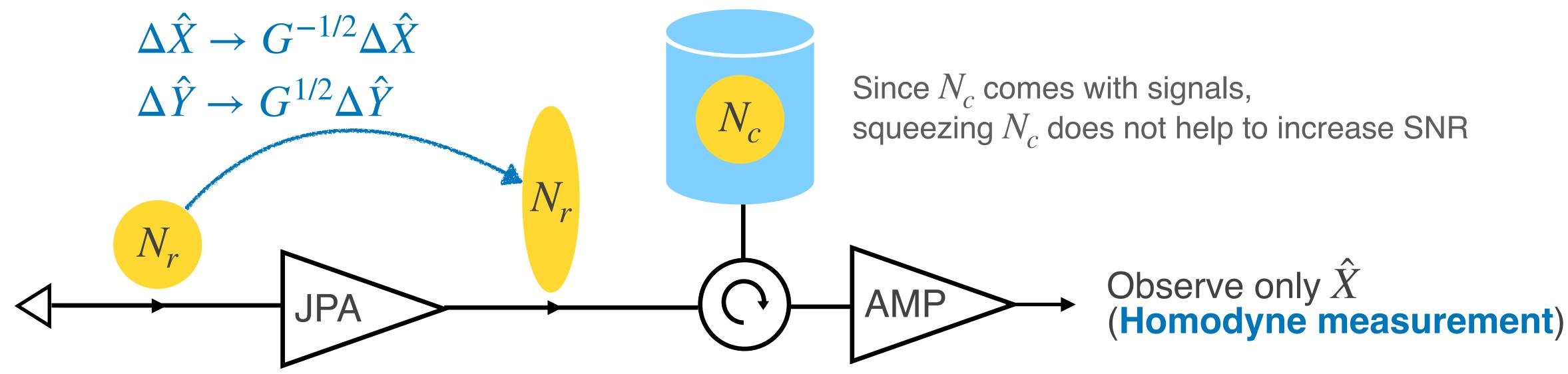
- Usually, we measure #photons by observing \hat{X} and \hat{Y} simultaneously
- Quantum noise $\Delta \hat{X} \sim \Delta \hat{Y} \sim 1/2$ becomes dominant in low temperature
- By imposing quantum fluctuation to \hat{Y} and observing only \hat{X} , we can measure photon signals with quantum fluctuation less than 1/2

$$\hat{[X, \hat{Y}]} = i \quad \hat{[\Delta \hat{X}]}(\Delta \hat{Y}) \ge \frac{1}{4}$$

Squeezing

Josephson Parametric Amplifier (JPA)

uncertainty relationship



can reduce quantum fluctuation of one quadrature while maintaining

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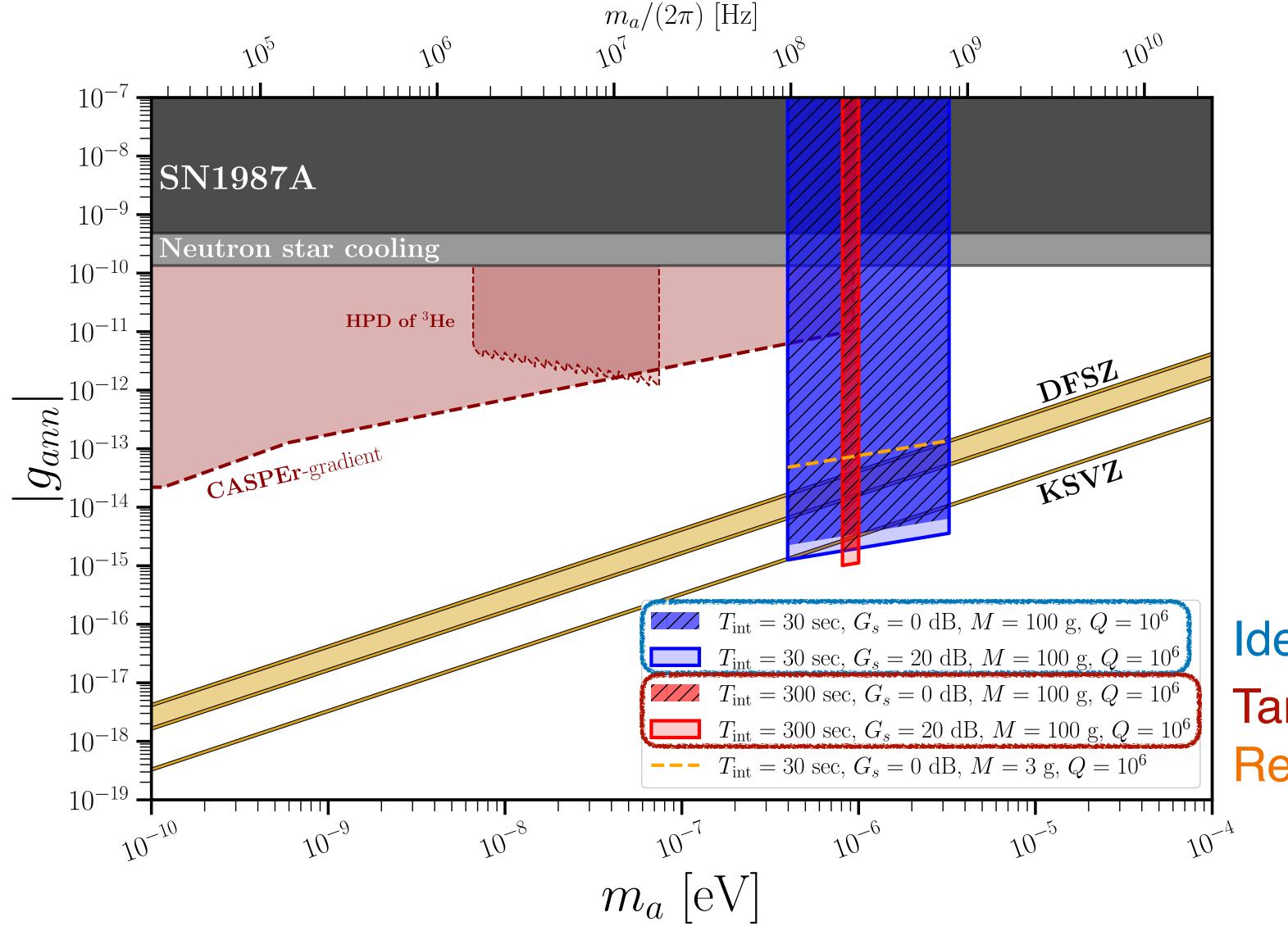
Bench mark setups

- Total observation time: 2 years
- Cavity quality factor: $Q = 10^6$

	Integration time for each scan T_{int}	Magnetic field	Squeezing parameter G_s	Helium mass $M_{^{3}\mathrm{He}}$
Ideal	30 sec	3.1 - 25 T	0 or 20 dB	100 g
Targeting	300 sec	6.2 - 7.7 T	0 or 20 dB	100g
Realistic	30 sec	3.1 - 25 T	0 dB	3 g

Axion is 100% Dark Matter

Result: 95% exclusion limit



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Ideal w/o & w/ squeezing Targeting w/o & w/ squeezing Realistic





Comparison with other axion-neutron coupling experiments

- CASPEr-gradient
 - detects spin precession of liquid Xenon-129 induced by axion

$$\gamma_{129Xe} \simeq 0.37 \gamma_{^{3}He} > \text{needs}$$

• Homogeneous Precession Domain of superfluid ³He

- uses B2 phase, which exists for $B \leq 0.5 \,\mathrm{T}$ • sensitive only to axion with $m_a \lesssim 0.07 \,\mu \text{eV} < \omega_L \simeq 1.3 \,\mu \text{eV} \left(\frac{B_z}{10 \,\text{T}}\right)$
- By using A1 phase of ³He, we can detect μeV axion with realistic magnetic fields!

- larger magnetic fields than ³He case
- [Gao et al. (2022)]





Summary

- Axion is a fascinating particle since it can solve the strong CP problem and also be a candidate for Dark Matter
- Axion-photon coupling has been well explored, but not for nucleon coupling
- We have shown that axion-neutron coupling can be probed by using magnons in A1 phase of superfluid helium-3
 - sensitive to $\mu e V$ axion, which is favored for Dark Matter axion
 - can explore heavier mass regions than experiments which use ¹²⁹Xe or B2 phase of ³He
- We also quantitatively evaluated enhancement of sensitivity by squeezing



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Back up

Model dependence of axion-photon coupling

Axion-photon coupling

$$g_{a\gamma\gamma} = \frac{\alpha}{2\pi f_a} \left(\frac{E}{N} - \frac{2}{3} \frac{4m_d + m_u}{m_d + m_u} \right)$$

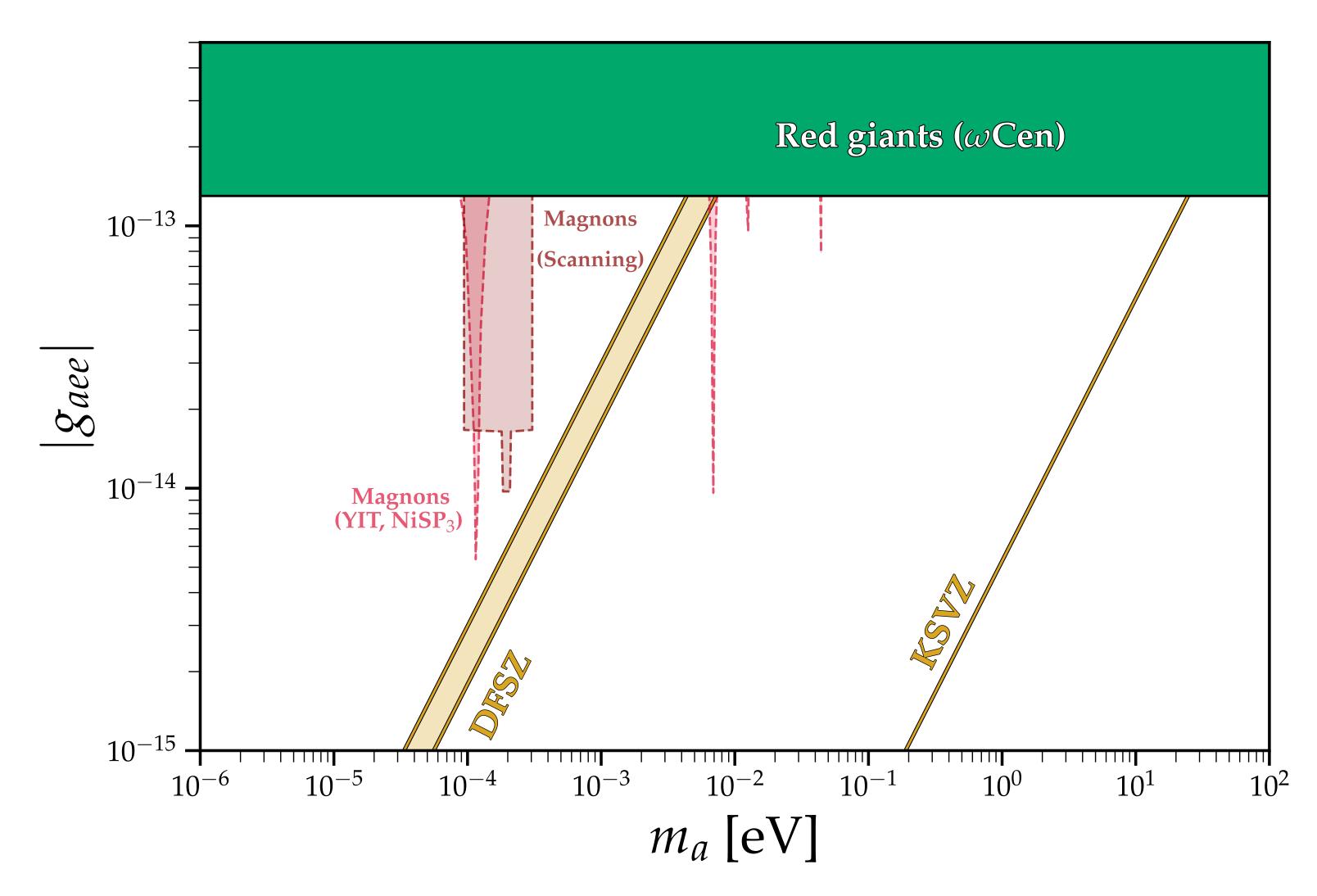
• Highly depends on electromagnetic (color) anomaly coefficient E(N)

• KSVZ: E/N = 0

- DFSZ: E/N = 8/3
- E/N can take arbitral value in general

Axion detection using superfluid Helium-3

Axion-electron coupling



Axion detection using superfluid Helium-3

ADMX

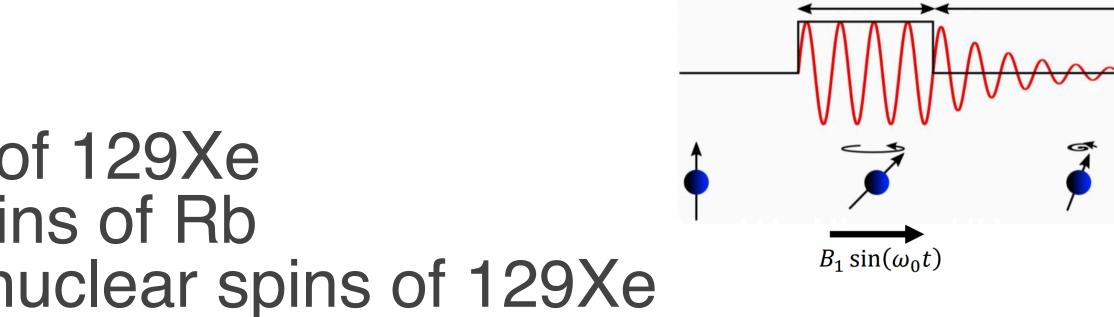
- Temperature inside cavity: 0.1 K
- Superconducting magnet > 7.5 T
- Cavity volume: 140 L
- Cavity quality factor: $\sim 10^5$
- Tune cavity frequency by moving "rods" inside cavity

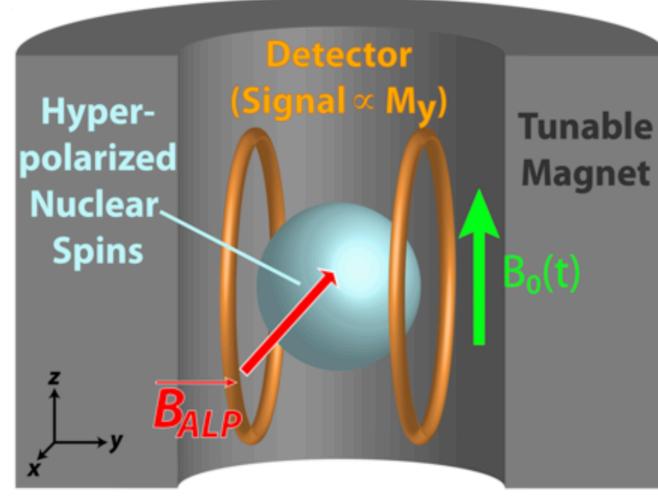


CAPSEr-gradient

- Hyperpolarization
 - Increase spin polarization of 129Xe by exchanging electron spins of Rb excited by laser light with nuclear spins of 129Xe
 - Spin polarization ~ 10⁻⁵ at thermal equilibrium (1 T, 300 K) can be increased to ~ 50%
 - Currently ~11% polarization achieved under 0.2 T magnetic field

Spin precession induced by axion is measured by magnetic sensor





pulse

[Budger group]

free induction decay (FID)





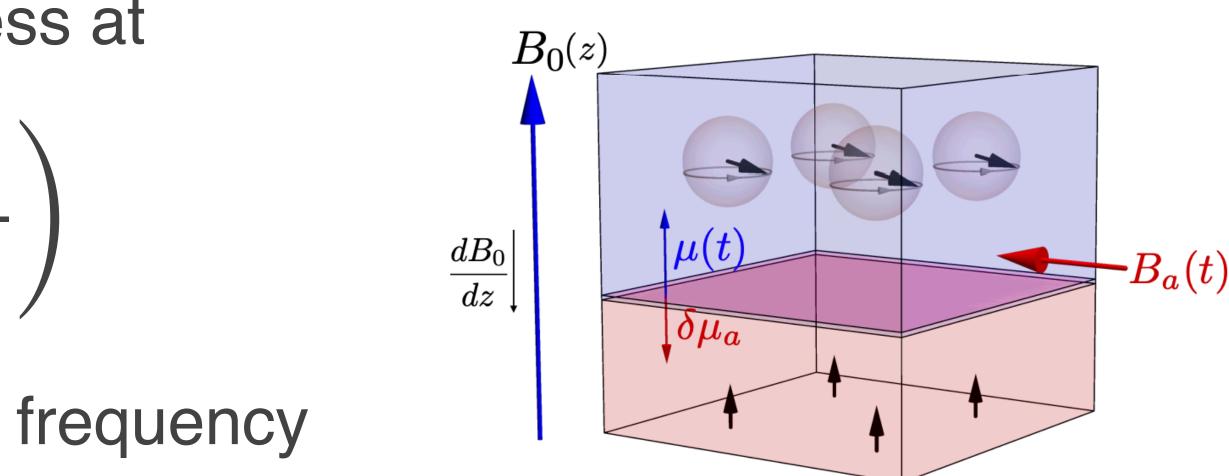
HPD of Helium-3 [Gao et al. (2022)]

- - Nuclear spins uniformly precess at

$$\omega_L \simeq 1.3 \,\mu \mathrm{eV} \,\left(\frac{B_z}{10 \,\mathrm{T}} \right)$$

- Measure a shift of precession frequency induced by axion
- Since the B2 phase exist only under $B_Z \lesssim 0.55$ T,

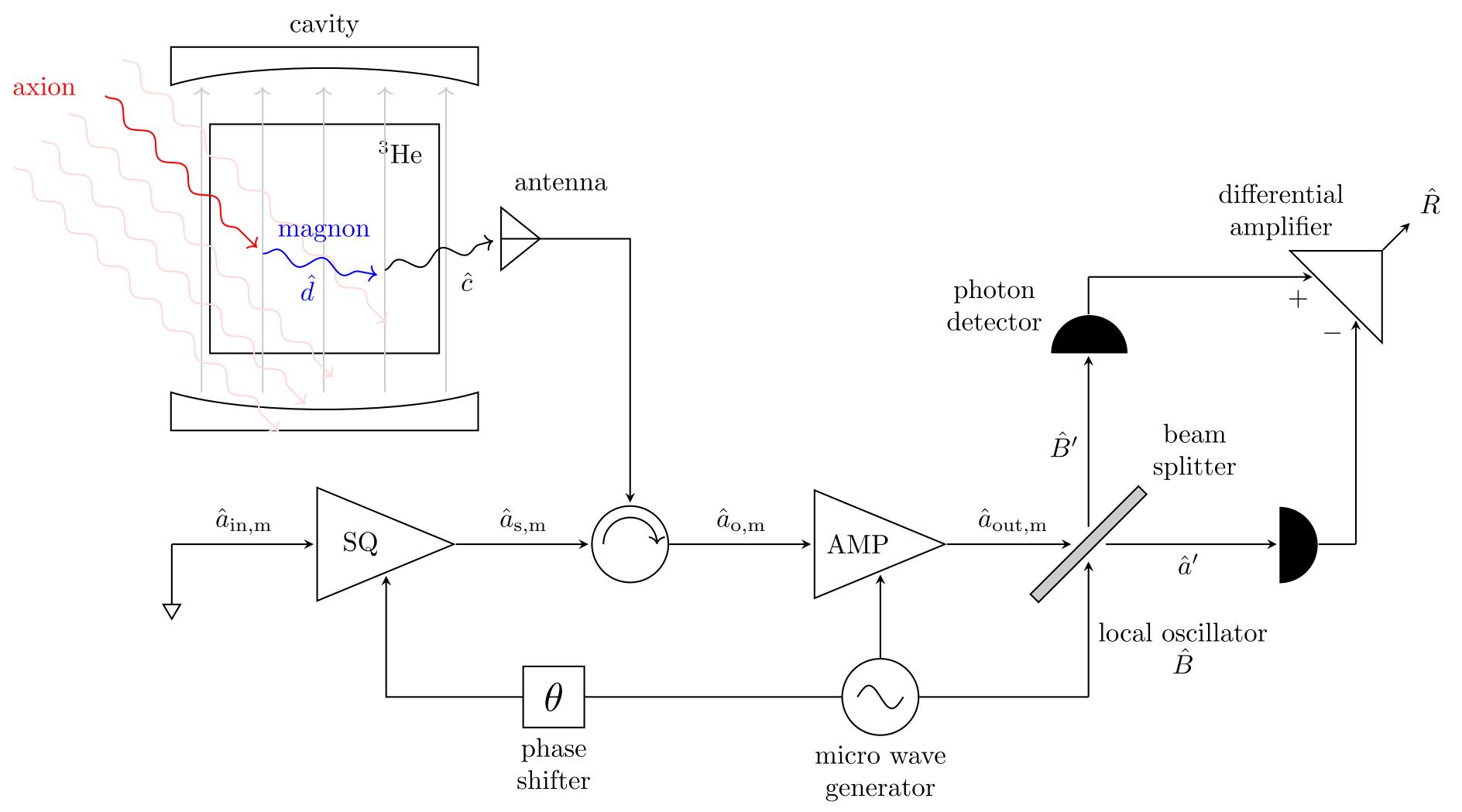
• B2 phase of superfluid 3 He = Homogeneous Precession Domain (HPD)



this experiment is sensitive to axion with mass $m_a \lesssim 0.07 \,\mu eV$

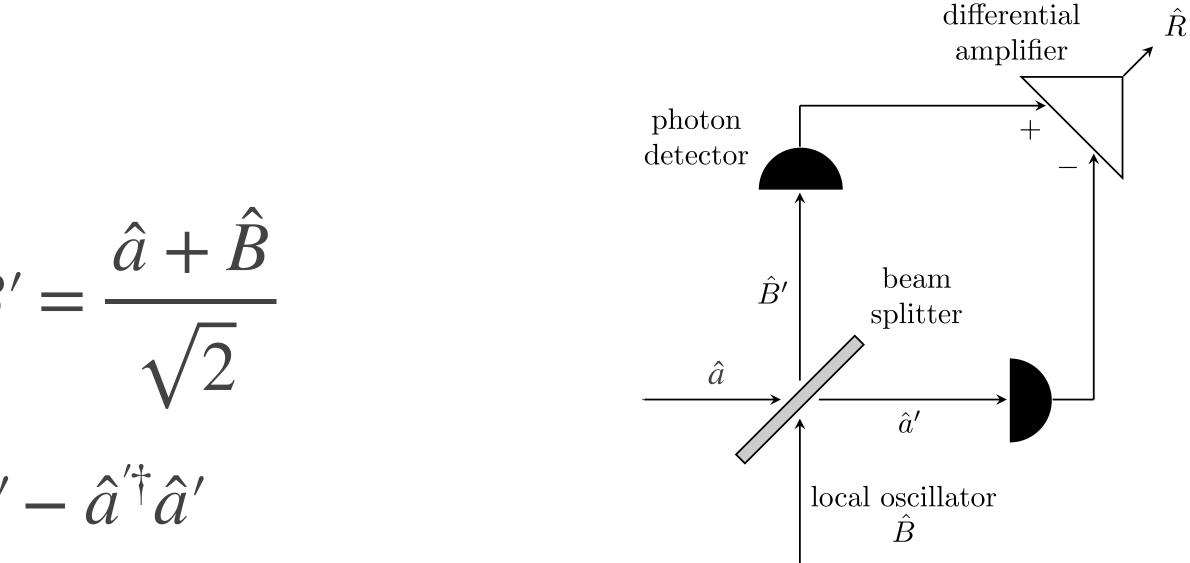


Experimental configuration



Homodyne measurement

- 1.
- Beam splitter $\hat{a}' = \frac{\hat{a} \hat{B}}{\sqrt{2}}, \quad \hat{B}' = \frac{\hat{a} + \hat{B}}{\sqrt{2}}$
- $\hat{R} = \hat{R}^{\dagger \dagger} \hat{R}^{\prime} \hat{a}^{\prime \dagger} \hat{a}^{\prime}$ Differential amplifier 2.
- Assuming that initial state of local oscillator is a coherent state $|\beta\rangle$ ($\beta = |\beta|e^{i\theta}$),
 - $\langle \hat{R} \rangle = \sqrt{2} |\beta| \langle \hat{X} \cos \theta + \hat{Y} \sin \theta \rangle$
 - In particular, $\langle \hat{R}/\sqrt{2} |\beta| \rangle = \langle \hat{X} \rangle$ for $\theta = 0$
- The error is proportional to $|\beta|^{-2} = \langle \hat{B}^{\dagger} \hat{B} \rangle^{-2}$



• We can measure only \hat{X} precisely in the limit of large input of local oscillator

Magnon lifetime

- We identify magnon lifetime as **spin relaxation time**
- Two types of spin relaxation time
 - Longitudinal spin relaxation time T_1 : relaxation of S_7
 - **Transverse** spin relaxation time T_2 : relaxation of S_x , S_y
- Since axion couples with transverse components, we should use T_2
- But, due to experimental difficulties, there is no measured value of intrinsic T_2 in A1 phase
- T_1 has been measured: $T_1 = O(1 10)$ sec
- Since $T_2 \lesssim T_1$ typically, we used $T_2 = 1$ sec as magnon lifetime

Log-likelihood ratio test

• Dataset: $\mathcal{D} = \{X_n^{(\mathcal{D})} \equiv X^{(\mathcal{D})}(n\Delta t)\}_{n \in \{0, \dots\}}$

• Fourier transformation: $X_k^{(\mathcal{D})} \equiv \sum_{k=1}^{N-1} X_n^{(\mathcal{D})} \epsilon$

• Power spectral density: $P_k^{(\mathcal{D})} = \frac{\Delta t^2}{T_{\text{int}}} |X_k^{(\mathcal{D})}|^2$

• Log-likelihood: $L(\mathcal{D} | P_k) = \prod_{k=1}^{N-1} \frac{1}{P_k} e^{-P_k^{(\mathcal{D})}/P_k}$ ($P_k = S_k + B_k$: model of signal and noise output)

Test statistic:

$$q = 2\sum_{k=1}^{N-1} \left[\ln L(\mathscr{D}_{\text{null}} | S_k + B_k) - \ln L(\mathscr{D}_{\text{null}} | B_k) \right] \simeq -\frac{T_{\text{int}}}{2\pi} \int_0^\infty d\omega \left(\frac{S(\omega)}{B(\omega)} \right)^2$$

$$q = -2.71 \text{ determines 95\% exclusion limit for null axion signal } (\mathscr{D}_{\text{null}})$$

$$e^{-i2\pi kn/N}$$
 $(k \in \{0, \dots, N-1\})$

- Amount of Helium-3
 - From natural gas ~ 2 kg / year
 - From decay of Tritium ~ 20 kg / year
- 100g Helium-3 $\simeq 10^3 \,\mathrm{cm}^3 \sim \$3,000,000$
- Cooling of meter-sized cavity
 - kept it for 15 days
- CAPP-25T
- Current highest level of squeezing ~ 15 dB
 - 20 dB will be achieved in a few years

• CUORE experiment ($0\nu 2\beta$ search) successfully cooled a ~ 1 m detector to 6 mK and

- Cavity size 5 L with a magnetic field of 25 T (high- T_c superconducting magnet) planned

Discussion

- To explore smaller- g_{ann} regions...

 $g_{ann}^{95\%} \propto m_a^{1/2} T_{\rm int}^{-1/2}$

- Longer T_{int} ? ... Explored regions become narrower
- Higher Q ? ... State-of-the-art Q is ~ 10^8 under 8 T [Ahn et al. (2024)]
- More ³He ? ... Typical amount $\sim 1 \mod \simeq 3 \text{ g}$
- There is no technology that can cool a meter-sized cavity to $\sim 2 \text{ mK}$
 - An adiabatic demagnetization refrigerator capable of continuously maintaining sub-mK has recently been developed [Toda et al. (2018)]

• Larger magnetic fields (> 25 T) are needed to explore heavier mass regions

$$^{4}Q^{-3/8}M_{^{3}\text{He}}^{-7/8}G_{s}^{-1/8}$$

