# From Axion Searches

# **TES TECHNOLOGY AT**

**Christina Schwemmbauer**<sup>1</sup>, Katharina-Sophie Isleif<sup>2</sup>, Friederike Januschek<sup>1</sup>, Axel Lindner<sup>1</sup>, Manuel Meyer<sup>3</sup>, Gulden Othman<sup>4</sup>, Elmeri Rivasto<sup>3</sup>, José Alejandro Rubiera Gimeno<sup>2</sup>

75

<sup>1</sup>Deutsches Elektronen Synchrotron DESY, Hamburg, Germany <sup>2</sup>Helmut-Schmidt Universität, Hamburg, Germany <sup>3</sup>CP3-Origins, University of Southern Denmark, Odense, Denmark <sup>4</sup>Institut für Experimentalphysik, Universität Hamburg, Germany

## To Direct Detection

### -DM – ELECTRON SCATTERING

**DM-electron scattering** [3] is a promising approach for detecting WIMP-like light Dark Matter (DM) candidates with sub-GeV masses.

When probing these lower mass ranges, the maximum energy transferred in a scattering event is entire kinetic energy of incident DM particle:

$$E_{T_{\text{max}}} = E_{\text{kin}} \sim m_{\chi} v^2 \sim 10^{-6} m_{\chi}$$

where  $v \approx 10^{-3}$  is the DM halo velocity.



## THE ALPS II EXPERIMENT

ALPS II (Any Light Particle Search) is a currently running Light-Shining-through-Walls (LSW) experiment at DESY Hamburg, Germany. It exploits photon-ALP conversion in magnetic fields [1] to search for ALPs at an ALP-photon coupling down to  $g_{a\gamma\gamma} = 2 \cdot 10^{-11} \text{ GeV}^{-1}$ , independent of astrophysical and cosmological models.

- ALPS II consists of two  $L = 12 \times 8.8$  m magnet strings of B = 5.3 T dipole magnets
- A production cavity (PC i.e. Fabry-Perot resonator) increases probability for photon-ALP conversion of laser light in a magnetic field
- Generated ALPs can cross a light-tight wall between the cavities, which cannot be penetrated by the laser light
- ALPs are re-converted to photons by the inverse effect in the regeneration cavity (RC)



regenerated photon is











High system detection efficiency	Newest measurements: > 90% efficiency		<ul> <li>Future ALPS II upgrades may incorporate TES detectors for photon counting via ALP-photon conversion, with the system's intrinsic background and detection efficiency already meeting current</li> </ul>
Good energy resolution (at 1064 nm)	Enhanced analysis methods improve the (1064 nm) <b>energy resolution from 8% to 5.3%</b>		requirements
Low intrinsic background - no fiber (radioactive	Cryostat shielding and pulse shape analysis based background		extrinsic backgrounds for ALPS II
decays, cosmics, etc.)	reduction yield a <b>background of 6.9x10<sup>-6</sup> cps</b> [2] at high signal efficiency		<ul> <li>The TES system is being explored for the independent detection of ~MeV dark matter</li> </ul>
Low extrinsic background - with fiber (mainly black-body radiation)	Enhanced energy resolution improved the background reduction down to <b>10<sup>-5</sup> cps</b> . Additional efforts:	Background requirements: < 7.7 x 10 <sup>-6</sup> cps	<ul> <li>Linear TES energy response allows for energy-resolved dark matter searches down to sub-eV energies, distinguishing photon-like pulses from background through fit parameters</li> </ul>
	<ul> <li>Currently building a setup to employ optical bandpass filters in the cold including cryo adjustment</li> </ul>		O Analysis towards an energy spectrum with subsequent investigation of
	<ul> <li>Fiber curling to cutoff higher wavelengths</li> </ul>		uncharted terrestrial light Divi parameter space currently ongoing
SDU: SDU: SDU: SDU: SDU: SDU: SDU: SDU:	Universität Hamburg R FORSCHUNG   DER LEHRE   DER BILDUNG	References : [1] K. Ehret et al., <u>NIMA 612(1)83-960 (2009)</u> [2] R. Shah et al., <u>J Low Temp Phys 209, 355–362 (2022)</u> [3] Y. Hochberg et al., Physical Review Letters, 116 (1) (2 [4] C. Schwemmbauer et al., <u>PoS COSMICWISPers (2024</u> )	016) ), 055 <u>christina.schwemmbauer@desy.de</u>