

"3D-HST/CANDELS: The structural evolution of galaxies as a function of stellar mass since $z=3$ "

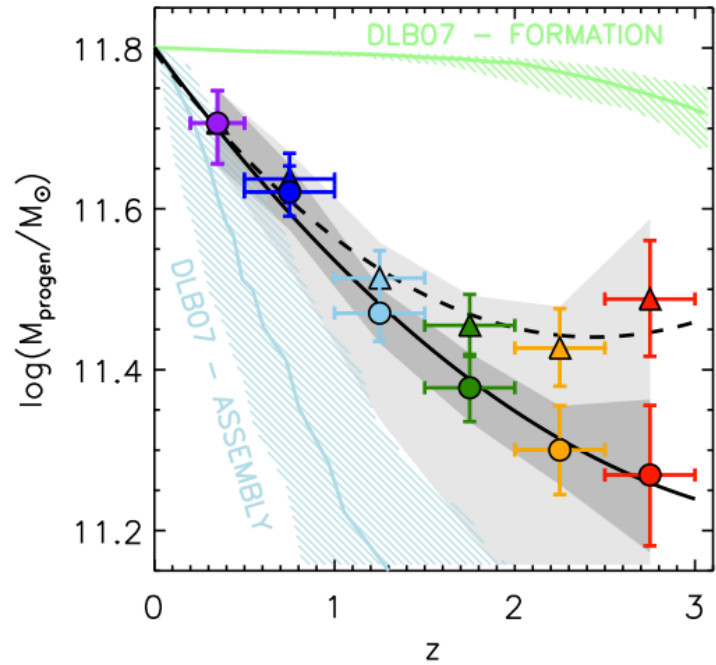
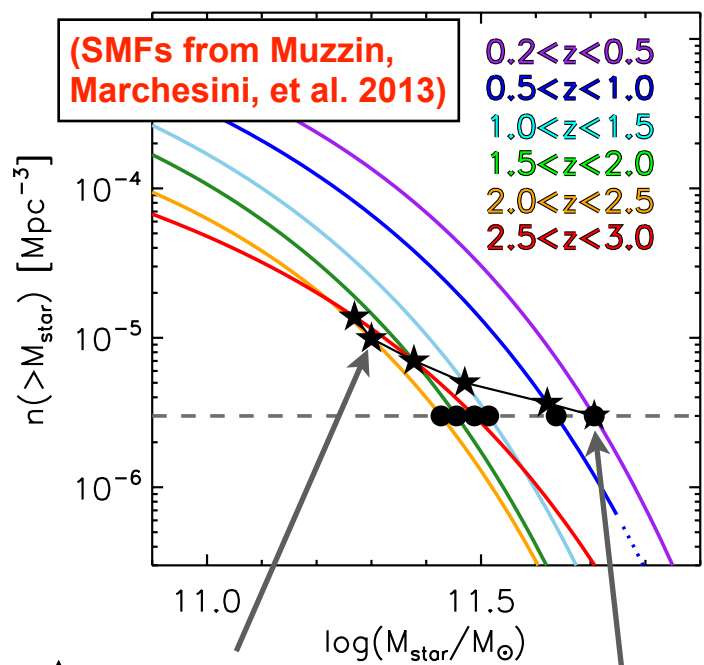
DANILO MARCHESINI (Tufts University)

and the 3D-HST collaboration

(Gabe Brammer, Ros Skelton, Kate Whitaker, Arjen van der Wel)



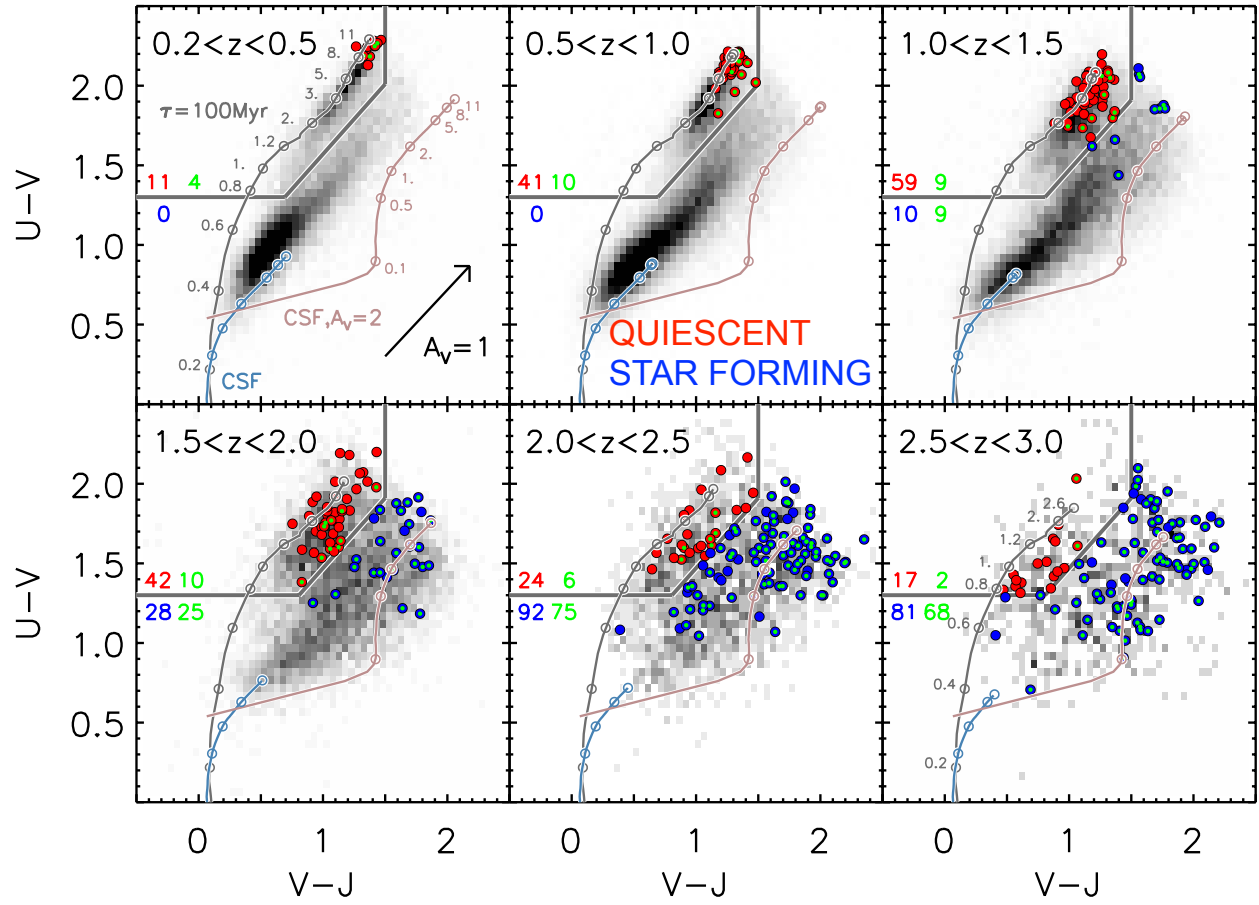
Getting a Grip on Galactic Girths
February 2-6, 2015 - IPMU (Tokyo)



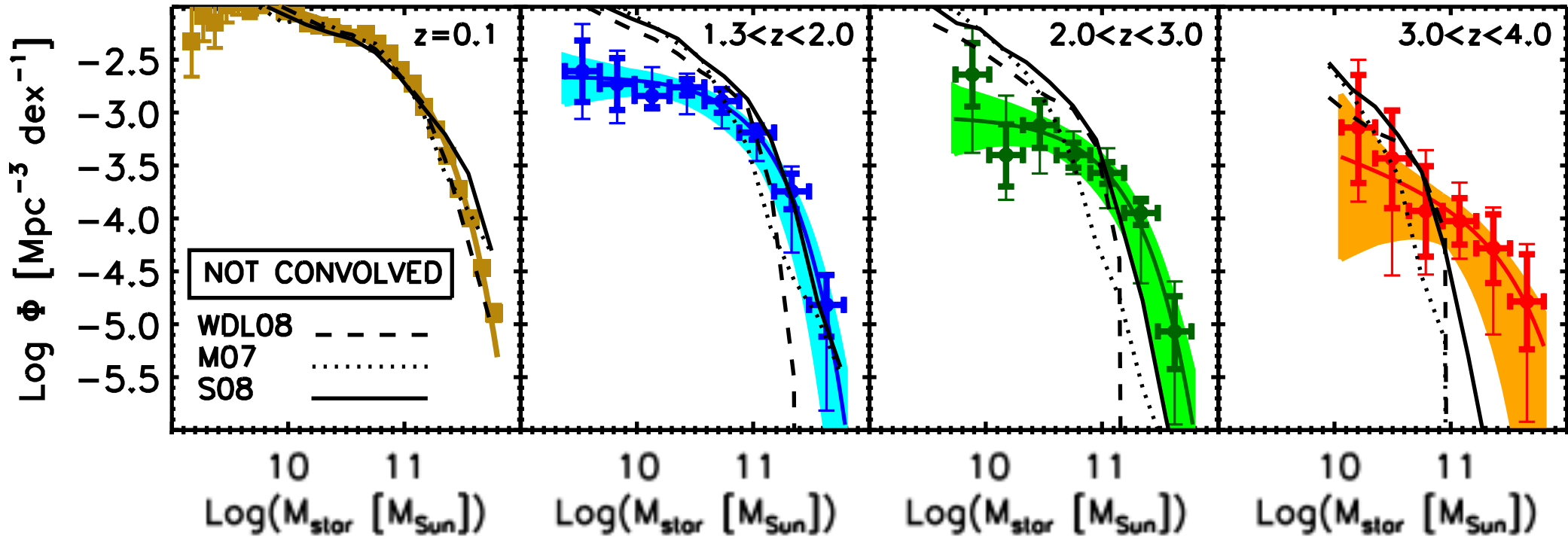
(Marchesini et al. 2014)

★ Abundance matching
(Behroozi, Marchesini, et al. 2014)

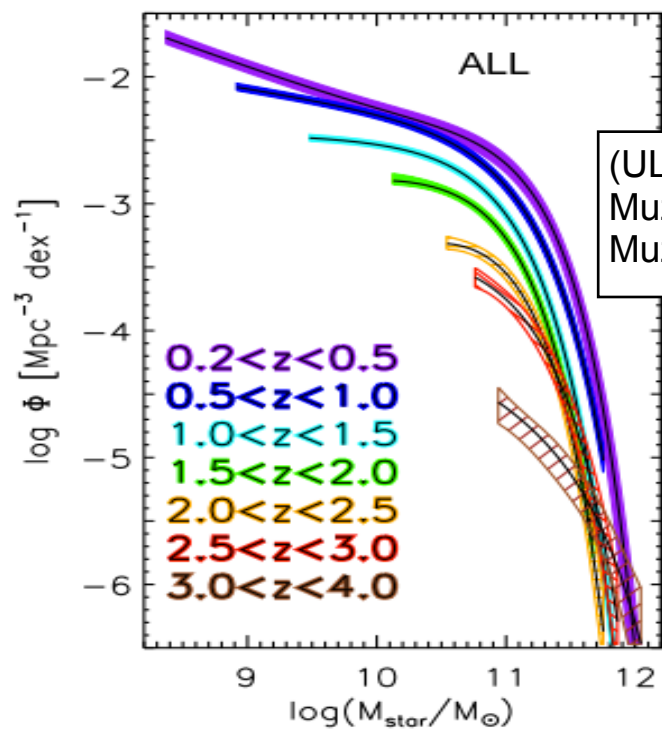
● Fixed cumulative number density
(e.g., Brammer, DM, et al. 2011; Leja et al. 2013)



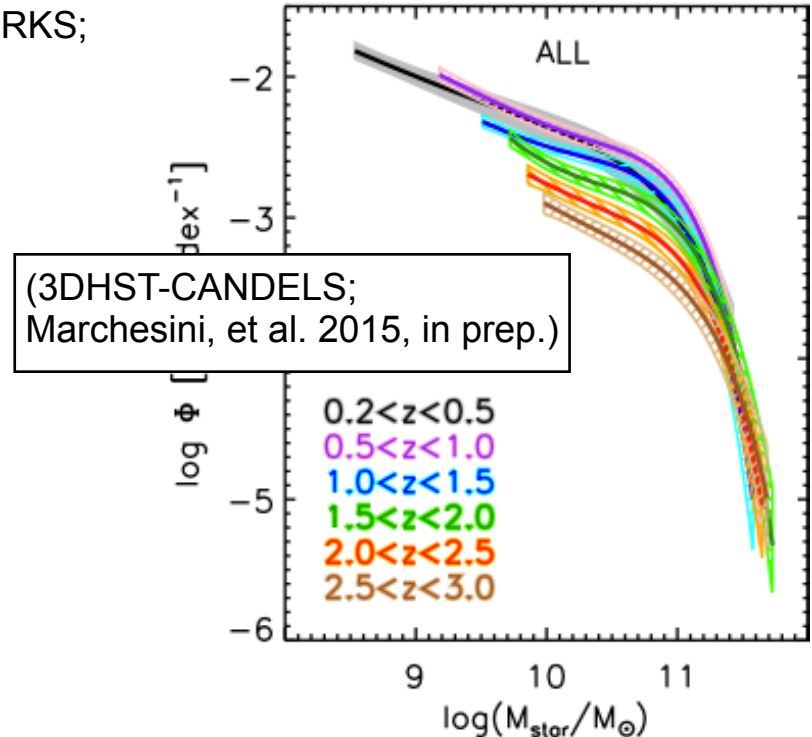
Evolution of the Stellar Mass Function (SMF)



(MUSYC+FIRES+FIREWORKS;
Marchesini et al., 2009)

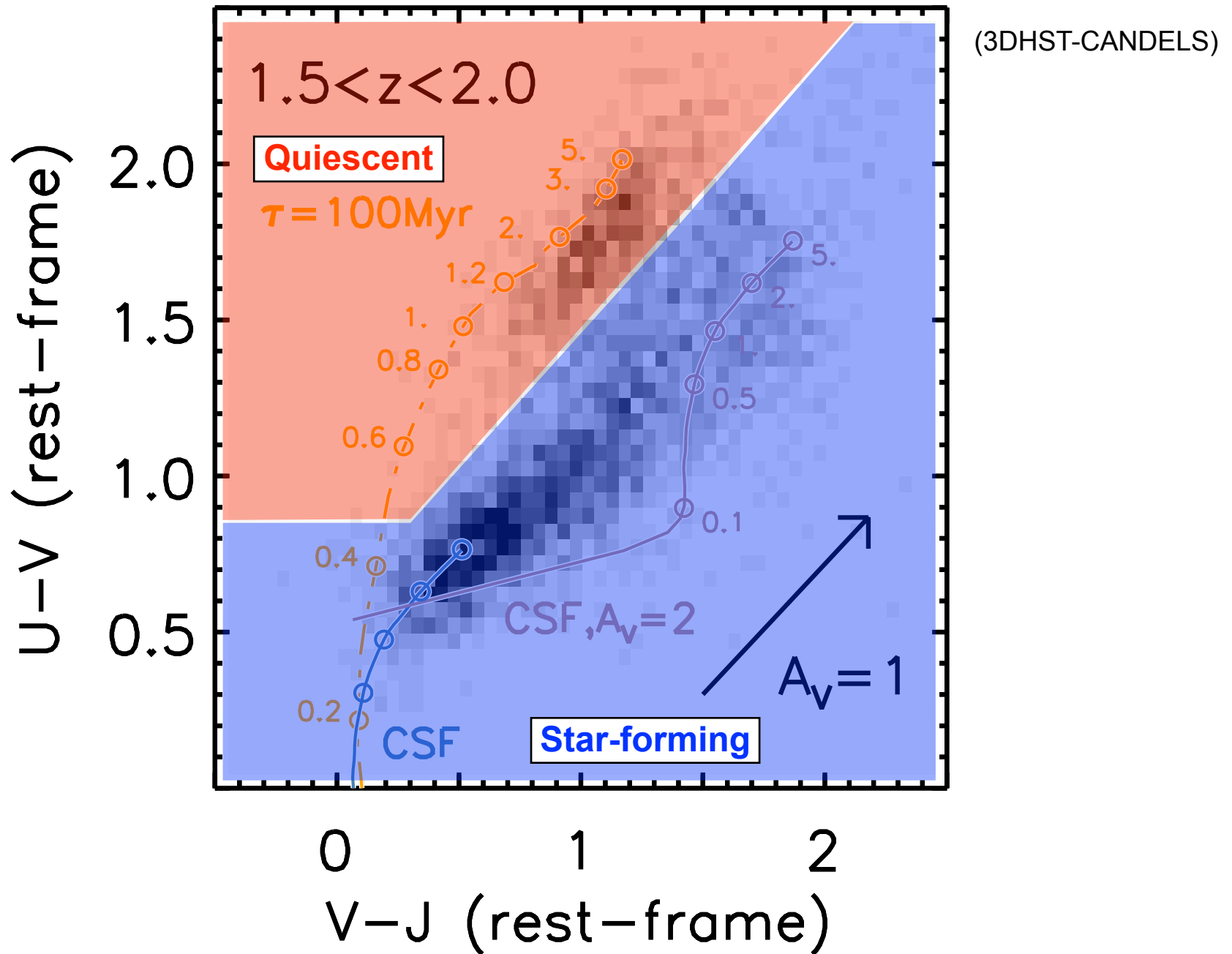


(ULTRAVISTA DR1;
Muzzin, Marchesini, et al. 2013a;
Muzzin, Marchesini, et al. 2013b)

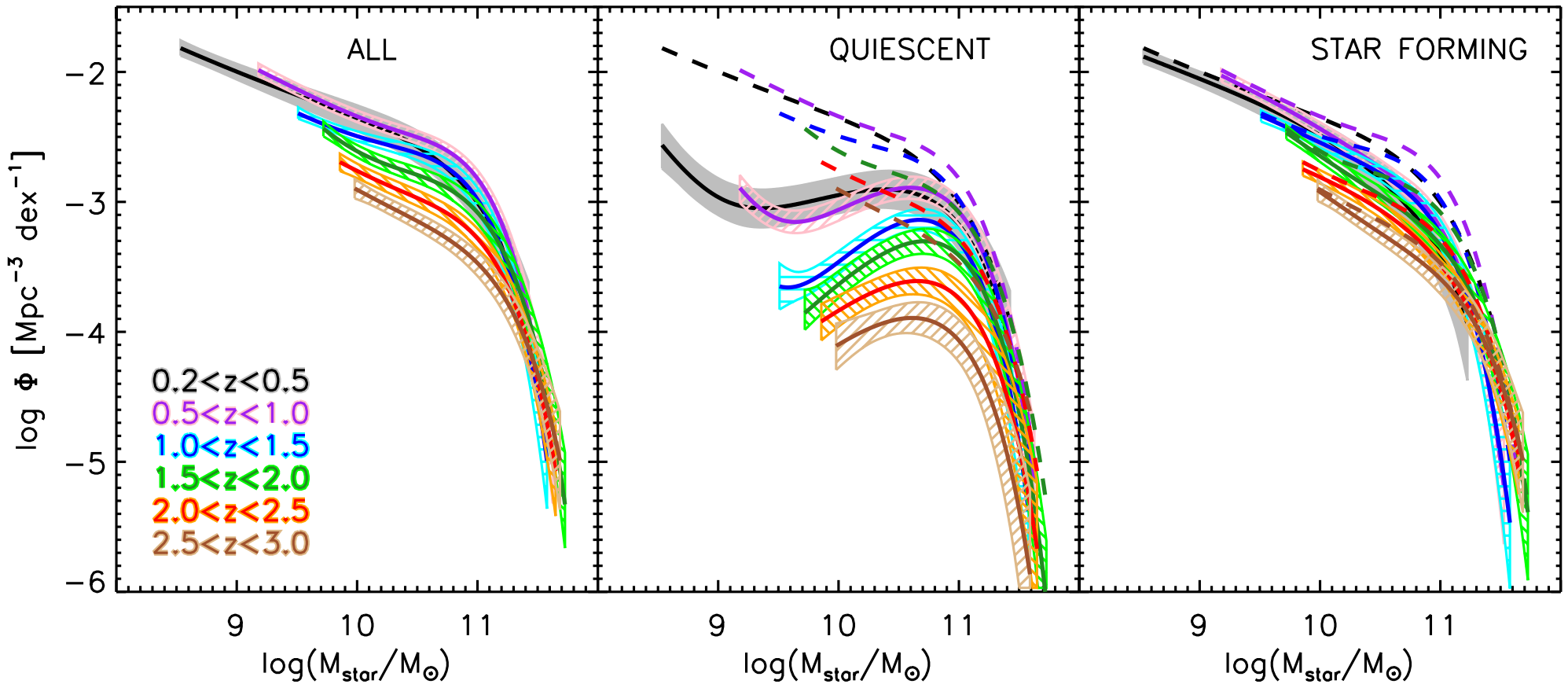


(3DHST-CANDELS;
Marchesini, et al. 2015, in prep.)

The UVJ diagram to separate Quiescent and Star-forming Galaxies



Evolution of the Stellar Mass Function (SMF)

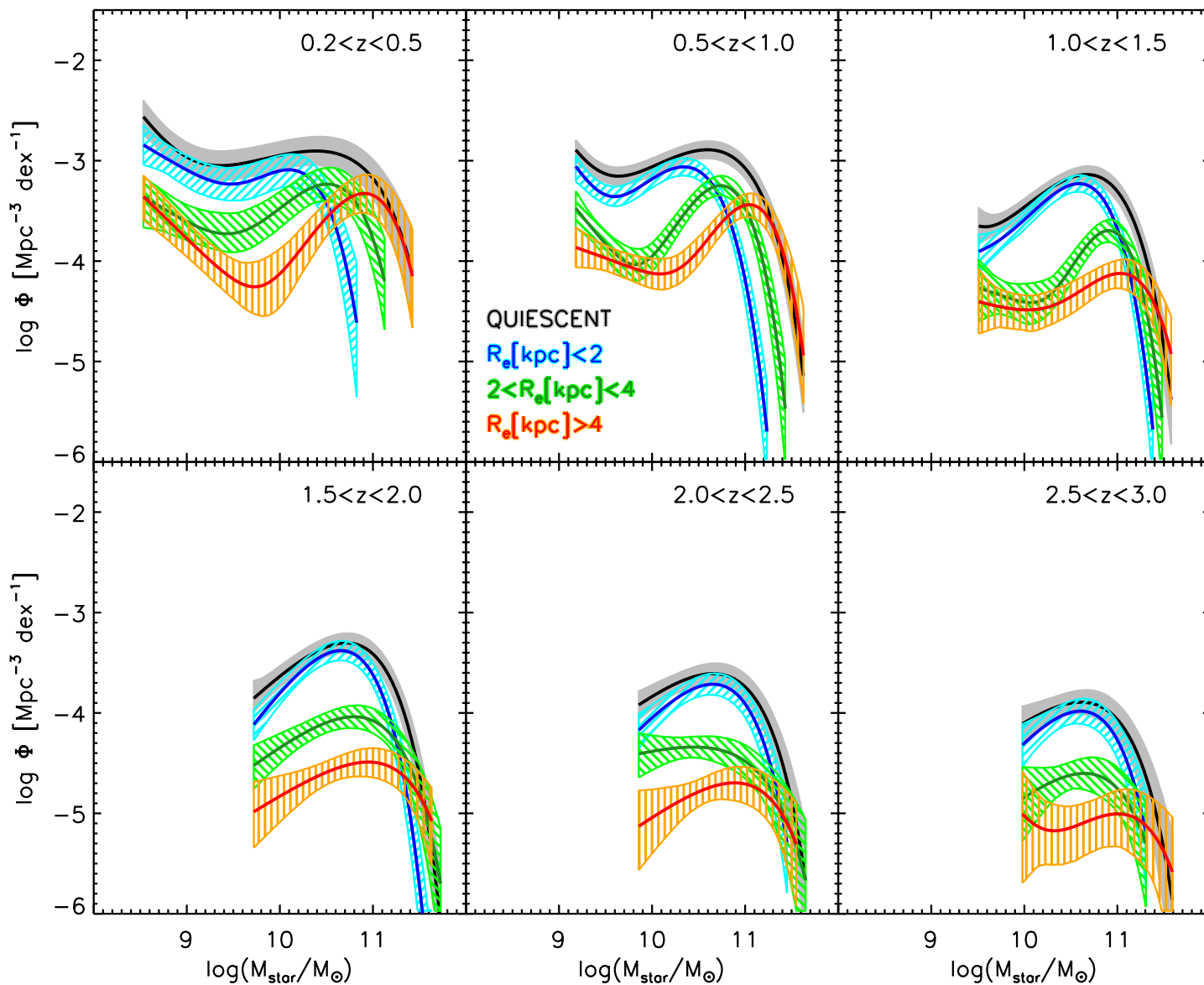


(Marchesini et al. 2015, in prep.; see also Muzzin, Marchesini, et al., 2013; and many many others)

3D-HST + CANDELS

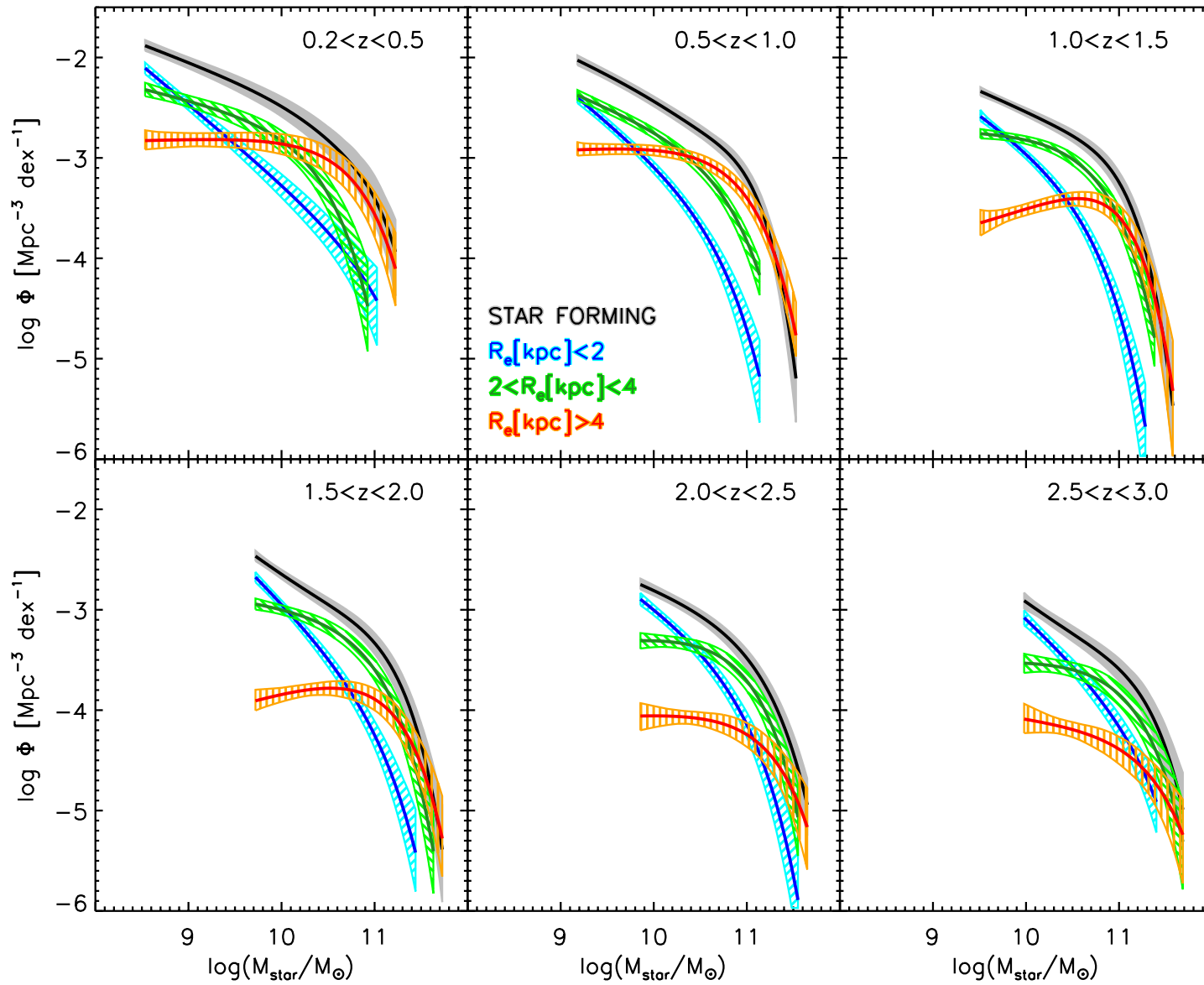
- Multi-wavelength photometry in AEGIS, COSMOS, UDS, and GOODS-N/S produced by the 3D-HST collaboration (Skelton et al. 2014)
- Grism redshifts from 3D-HST
- Effective surveyed area: **891.1 arcmin²** over **5 independent fields**
- **21717 galaxies** at $0.2 < z < 3$ with $H_{160} < 25.1$ AB (completeness far better than 90%)
- **16.5% with z_{spec}** , **39.1% with reliable z_{grism}** , 44.4% with z_{phot}
- **Standard definition of M_{star}** adopted (not the integral of SFH)
- Structural parameters (size and Sersic index) from van der Wel et al. (2012, 2014)
- Six redshift intervals: **$0.2 < z < 0.5$** (2437 galaxies), **$0.5 < z < 1.0$** (6799 galaxies), **$1.0 < z < 1.5$** (5229 galaxies), **$1.5 < z < 2.0$** (3745 galaxies), **$2.0 < z < 2.5$** (2228 galaxies), **$2.5 < z < 3.0$** (1279 galaxies).

SMF of **quiescent** galaxies as a function of **SIZE**



- At $z < 1$, large galaxies dominate high-mass end of the SMF of quiescent galaxies; small galaxies dominate the low-mass end.
- Increasing importance with redshift of small galaxies at high-mass end; large galaxies non negligible only for most massive galaxies.
- At $2.5 < z < 3$, quiescent galaxies mostly small, except for most massive.
- At all redshifts, the most massive galaxies are large ($R_e > 4$ kpc).

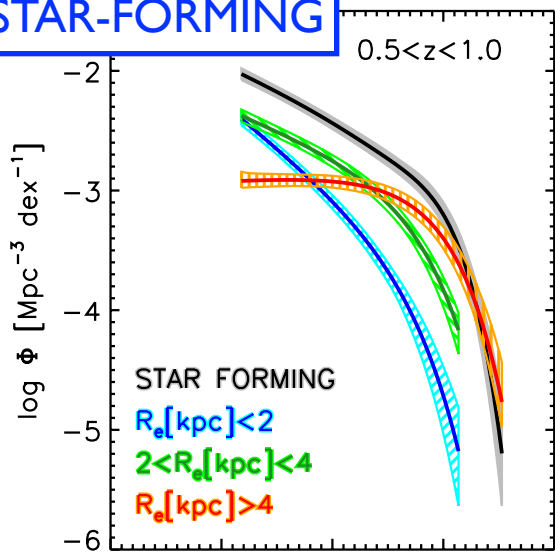
SMF of **star-forming** galaxies as a function of **SIZE**



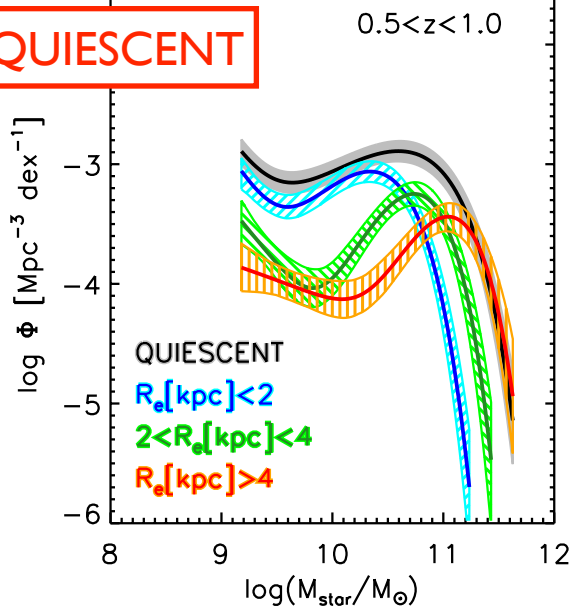
- Very different evolution compared to the quiescent galaxies.
- At high redshift, typical SF galaxies have intermediate sizes; large SF galaxies are rare, and small galaxies contribute significantly only at $\log M < 11$.
- With cosmic time, the relative contribution of small galaxies decreases, while that of large galaxies increases, dominating the high-mass end at $z < 1.5$.

Evolution with redshift of SMFs of **quiescent** and **star-forming** galaxies as a function of **SIZE**

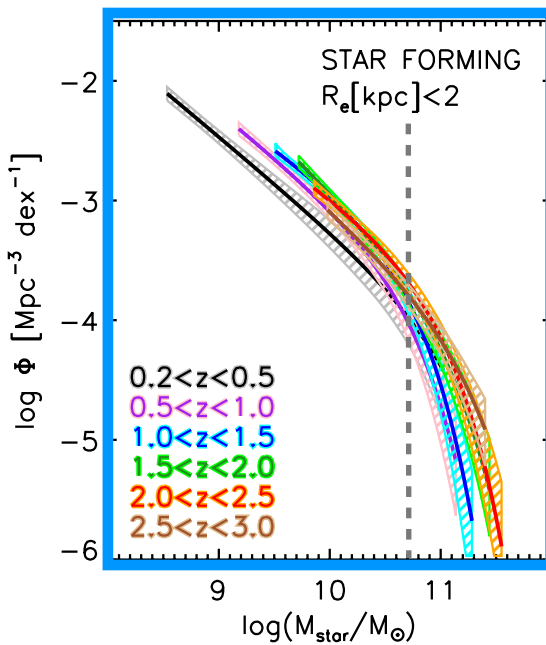
STAR-FORMING



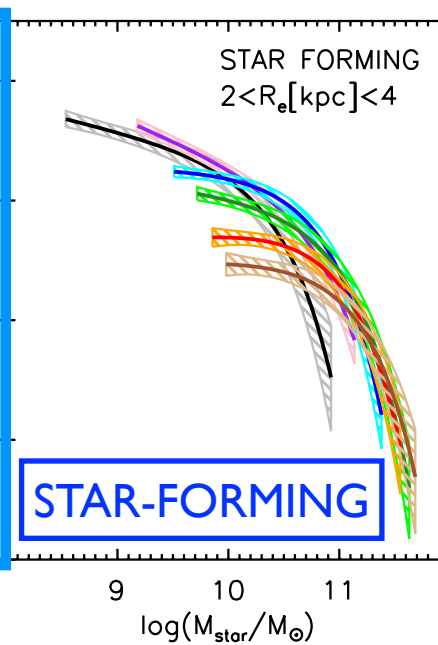
QUIESCENT



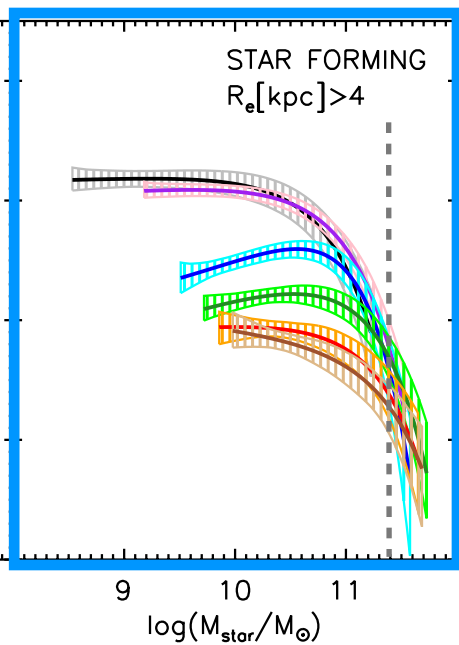
SMALL



INTERMEDIATE



LARGE

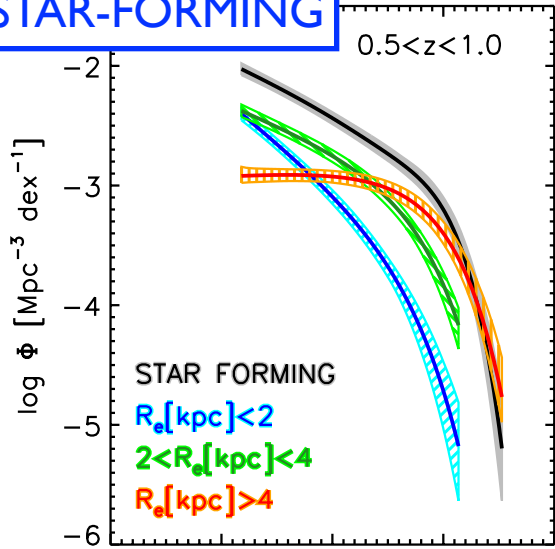


■ Star-forming galaxies become progressively larger with cosmic time

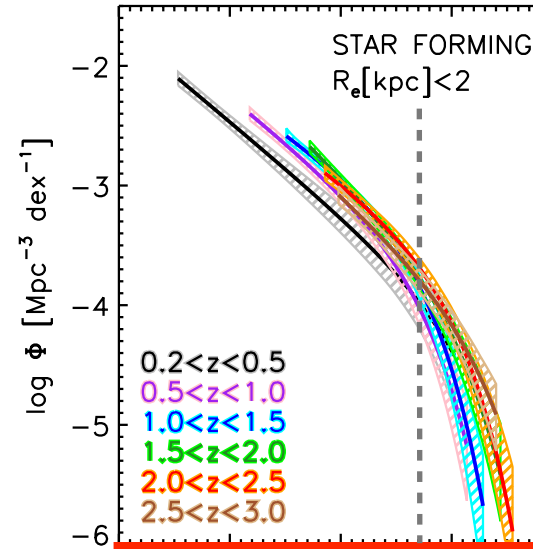
Q1: Why are small SF galaxies not evolving much in number at $\text{lg}M < 10.7$, whereas dramatic growth of the number of intermediate and large SF galaxies with $\text{lg}M < 11$? How do we interpret this?

Evolution with redshift of SMFs of **quiescent** and **star-forming** galaxies as a function of **SIZE**

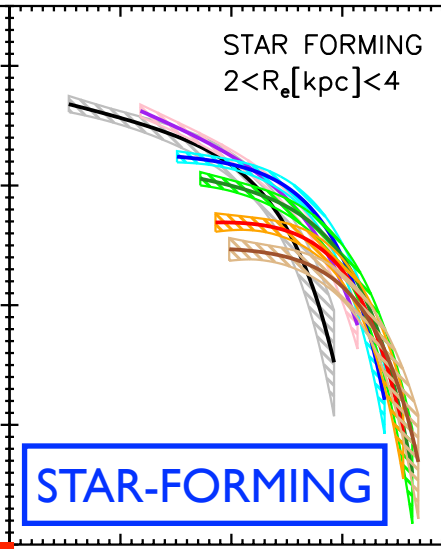
STAR-FORMING



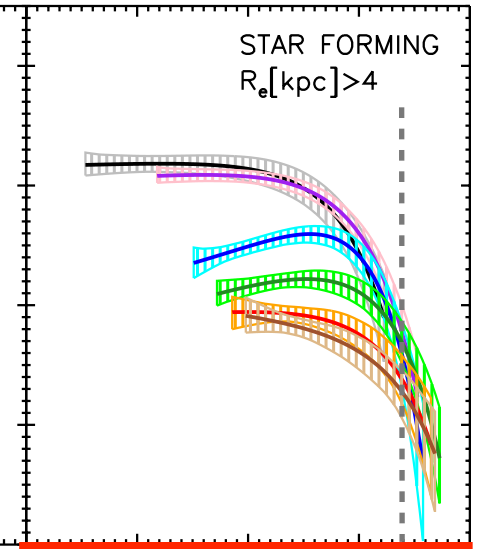
SMALL



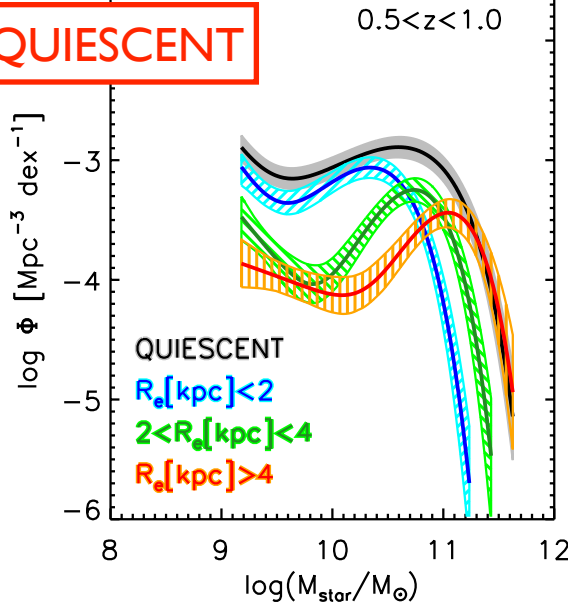
INTERMEDIATE



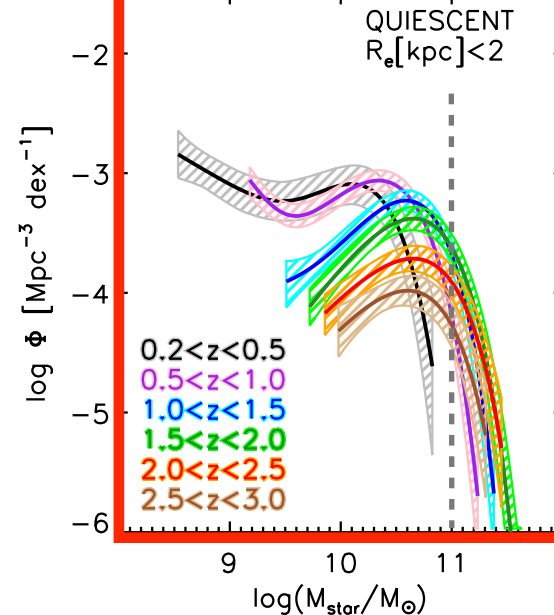
LARGE



QUIESCENT

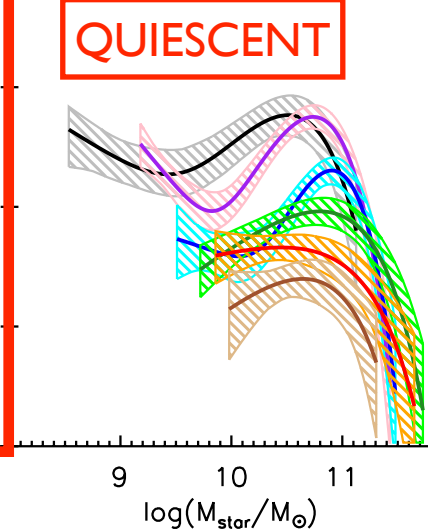


QUIESCENT
 $R_e[\text{kpc}] < 2$

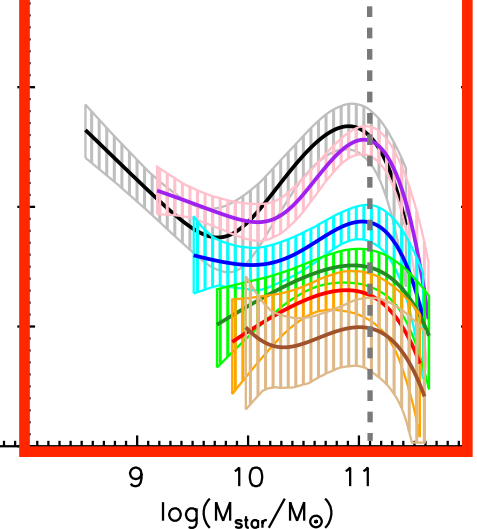


QUIESCENT
 $2 < R_e[\text{kpc}] < 4$

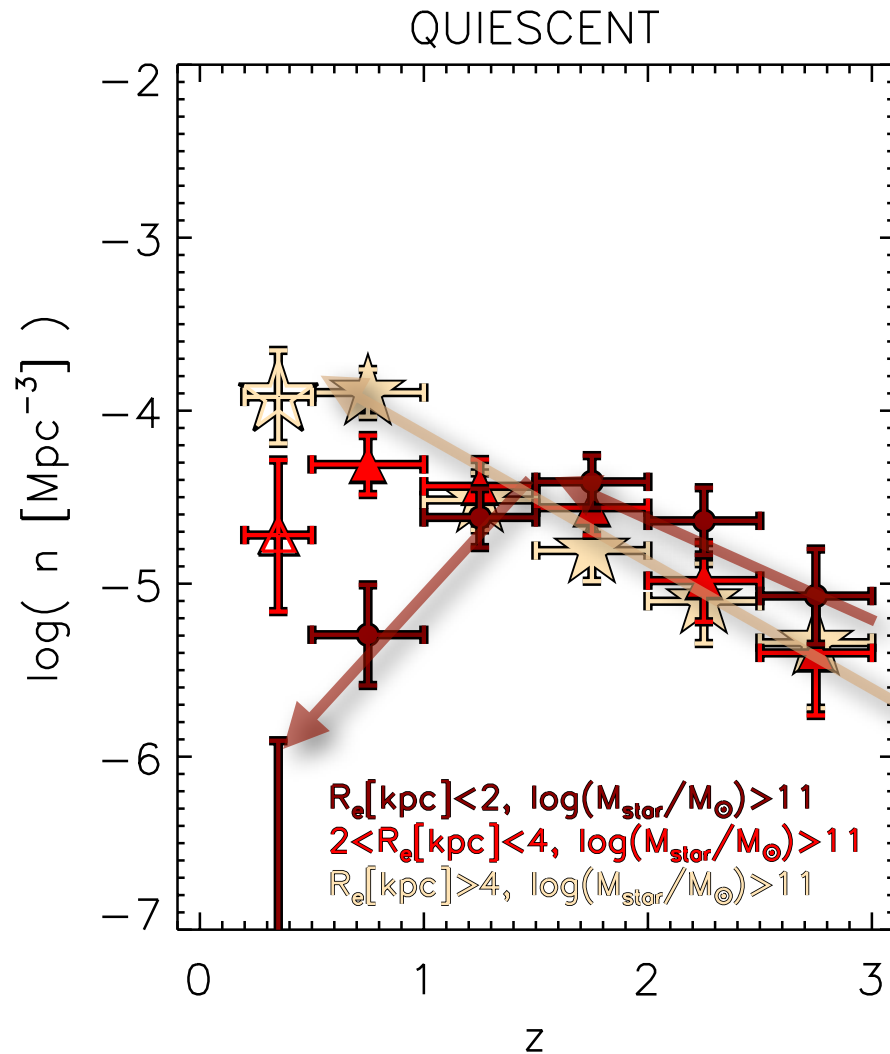
QUIESCENT



QUIESCENT
 $R_e[\text{kpc}] > 4$

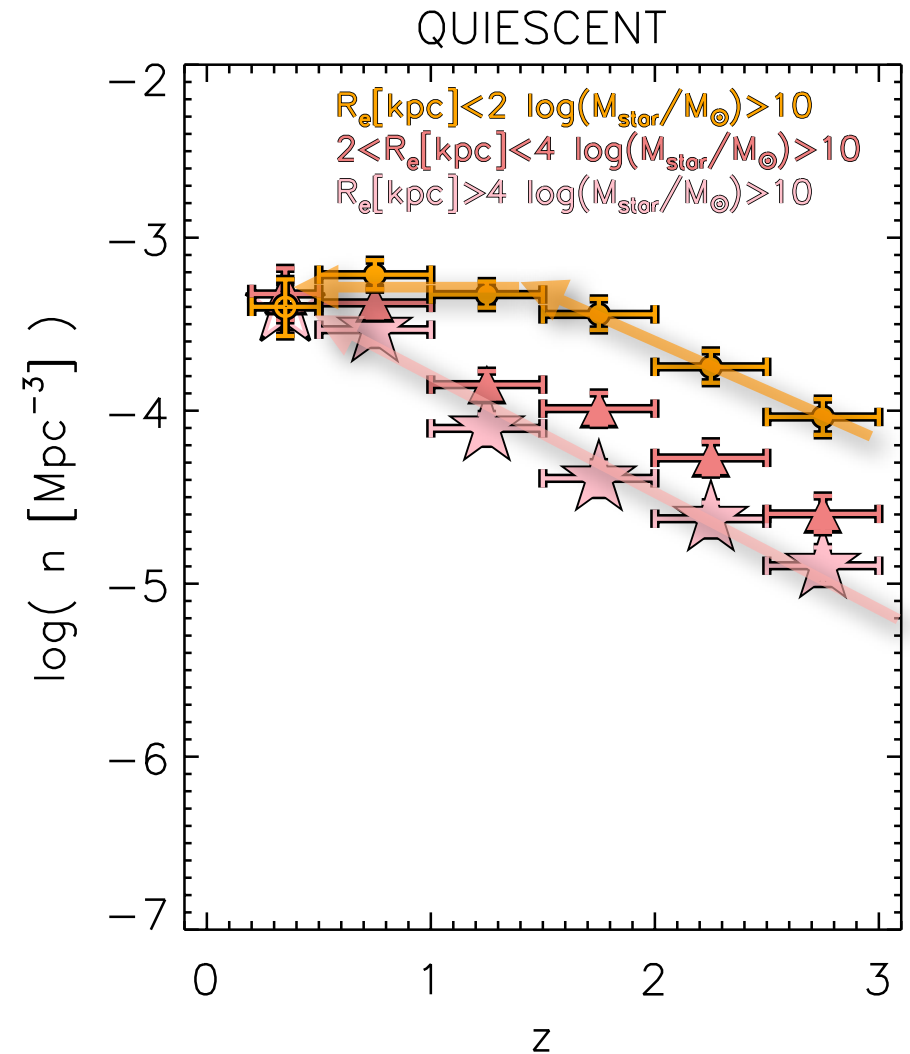
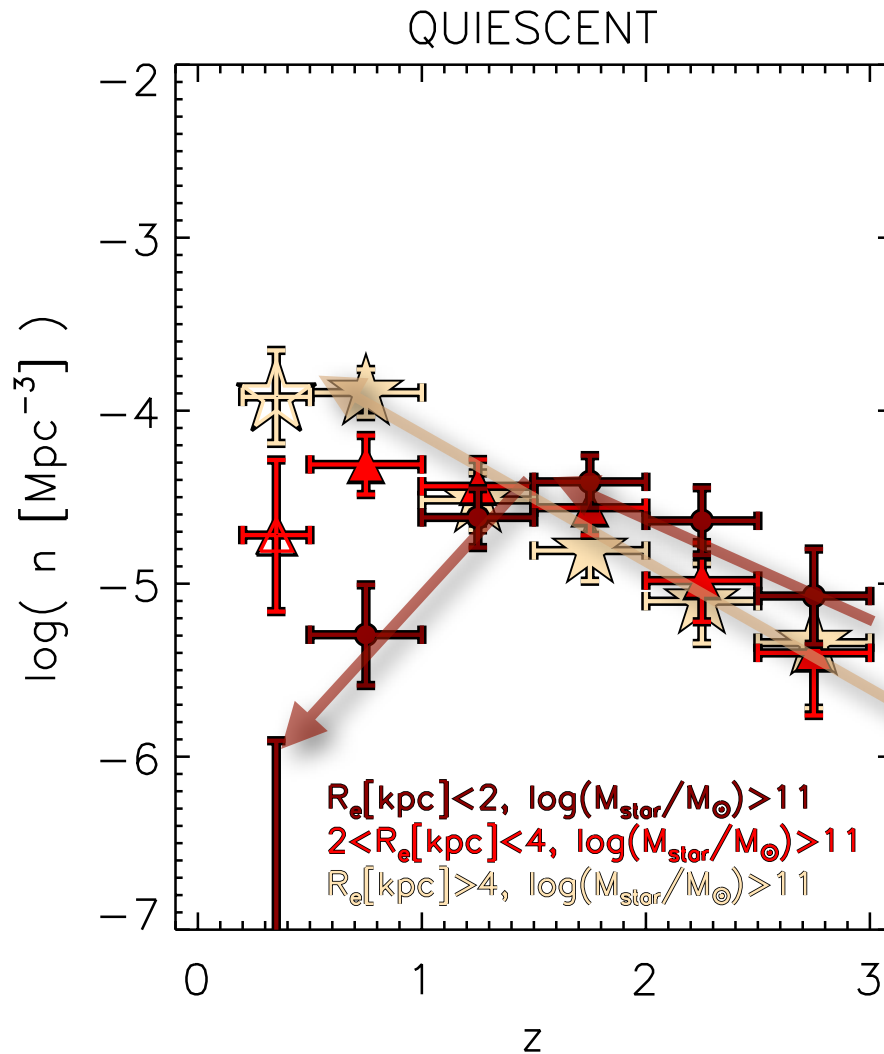


Evolution with redshift of the number density of **quiescent** galaxies of different **SIZES**



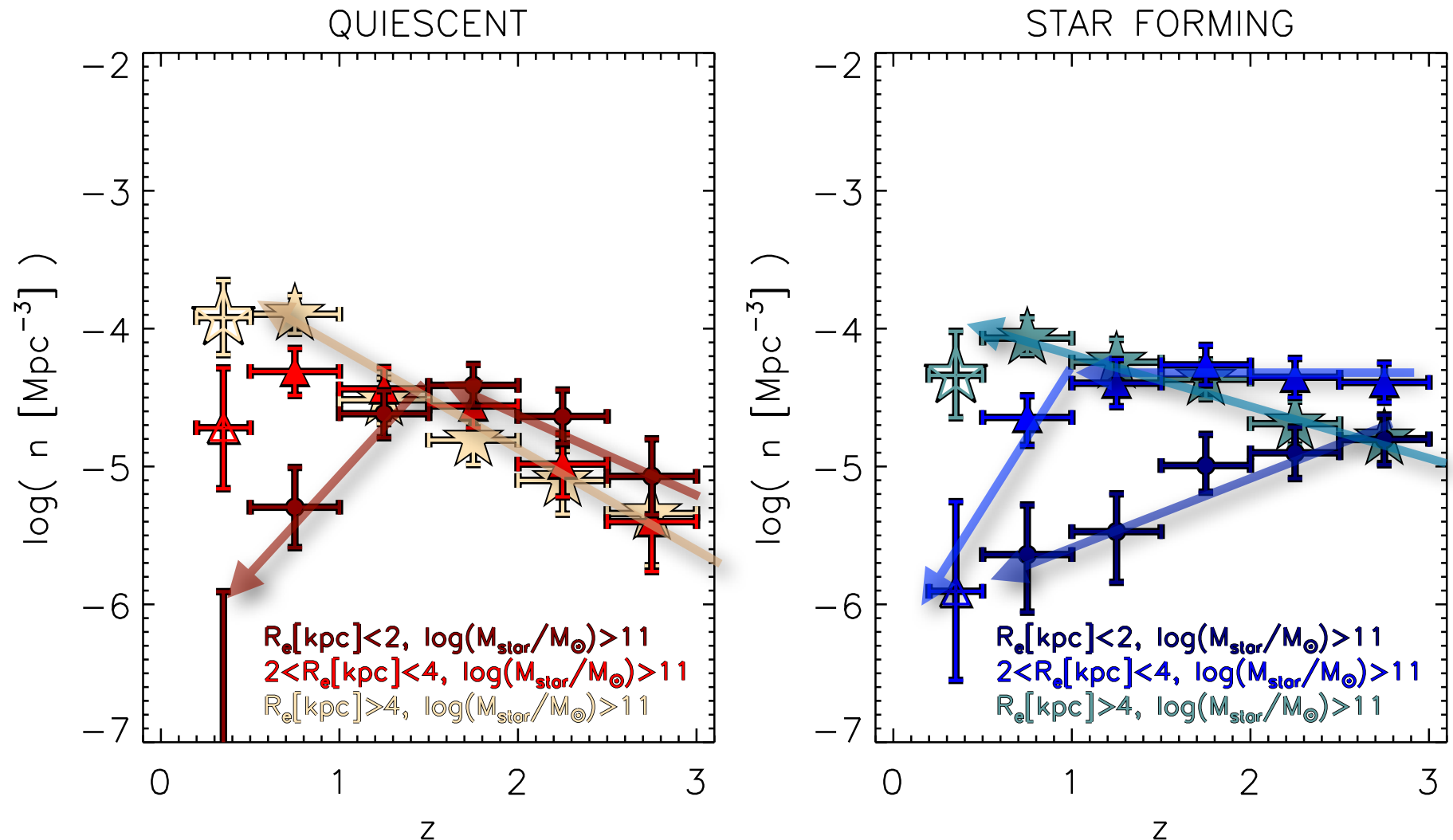
- Small galaxies dominate the population of massive quiescent galaxies at high z . Their number density grows by a factor of ~ 7 from $z=3$ to $z=1.5$, and then decreases rapidly by $10\times$ from $z=1.5$ to $z=0.5$. At $0.2 < z < 0.5$, small massive quiescent galaxies are very rare.
- Large increase ($\sim 30\times$) of the number density of massive large quiescent galaxies from $z=3$ to $z=0.5$. Large ($R_e > 4$ kpc) galaxies dominate the massive quiescent population at $z < 1$

Evolution with redshift of the number density of **quiescent** galaxies of different **SIZES**



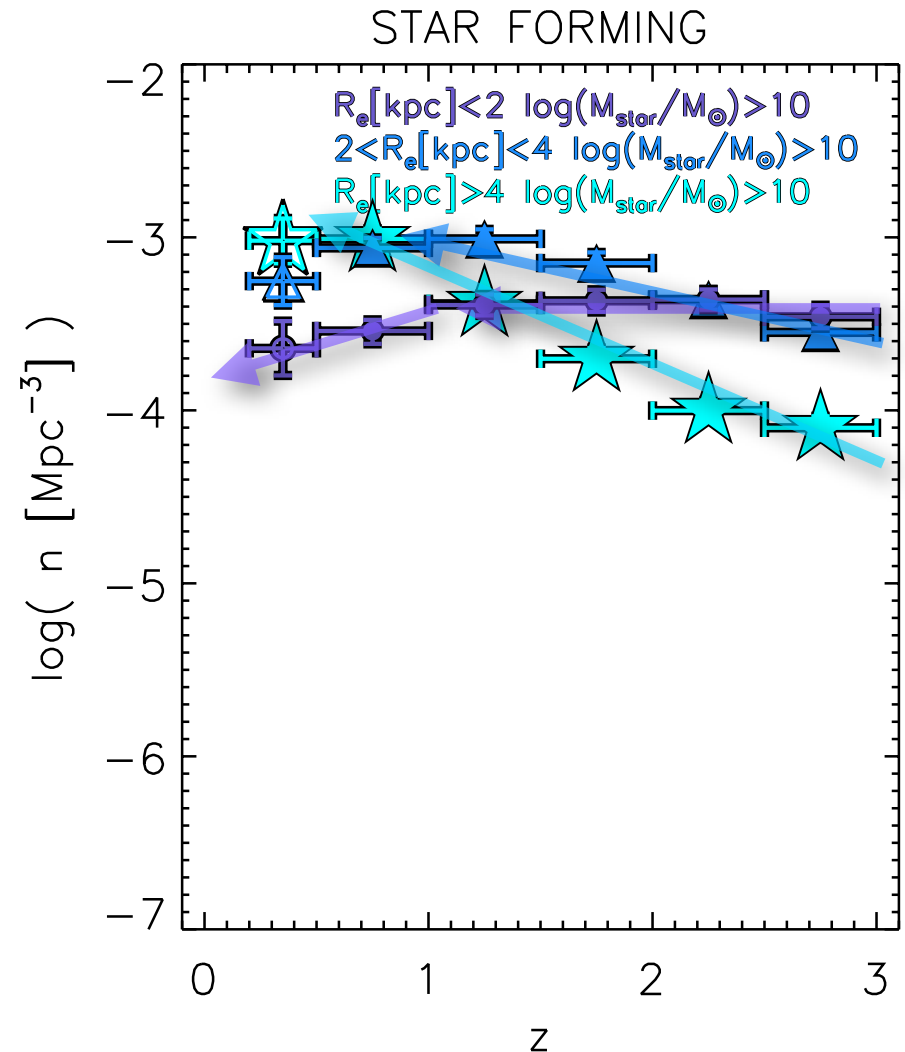
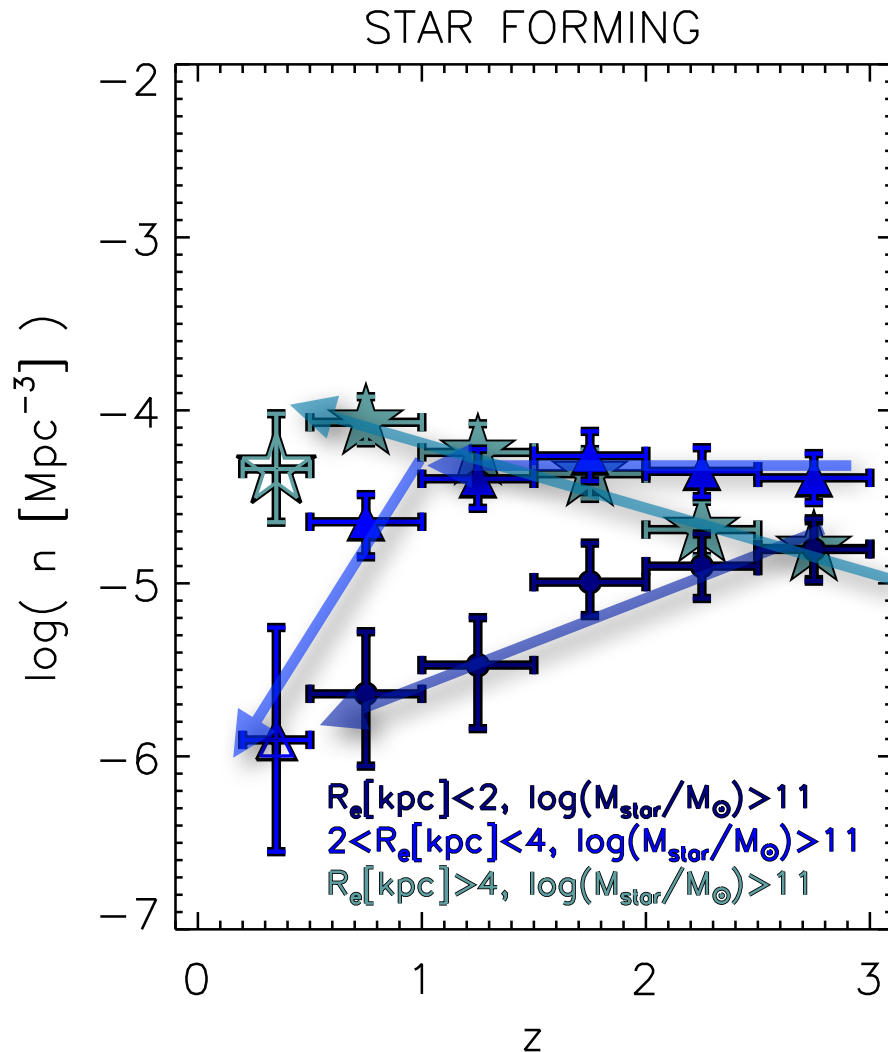
- Small galaxies dominate the population of $\lg M = 10-11$ quiescent galaxies at $z > 1$. The number density of small, quiescent galaxies with $\lg M > 10$ increases with time down to $z \sim 1.5$, and then it remains roughly constant (differently from $\lg M > 11$ small quiescent galaxies)
- The number density of large quiescent galaxies with $\lg M > 10$ increases steadily with time by $\sim 30x$ from $z = 3$ to $z = 0.5$.

Evolution with redshift of the number density of **star-forming** galaxies of different **SIZES**



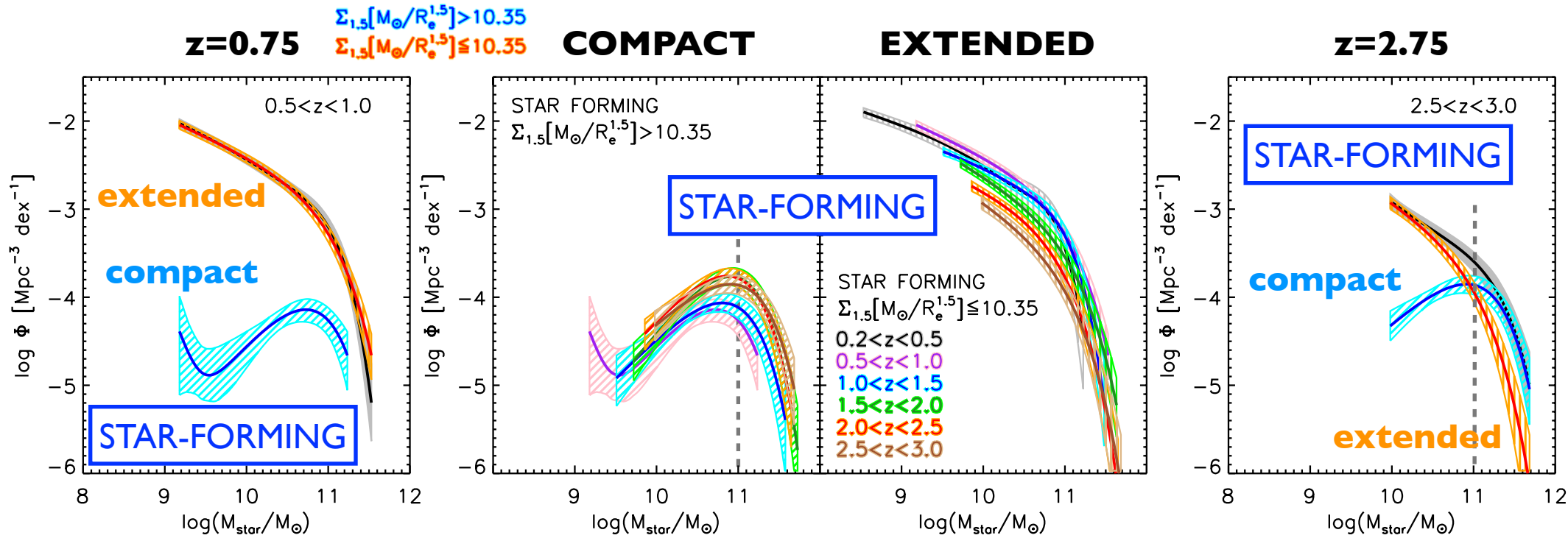
- The number density of small, massive SF galaxies decreases steadily from $z=3$ to $z=0.5$ by a factor of ~ 8 , and more than $10\times$ down to $z=0.2$.
- Intermediate size galaxies are dominant type among massive SF population at high z , with density roughly constant down to $z \sim 1$, then decreasing rapidly.
- Density of large, massive SF galaxies increases with time, becoming dominant at $z < 1.5$

Evolution with redshift of the number density of **star-forming** galaxies of different **SIZES**

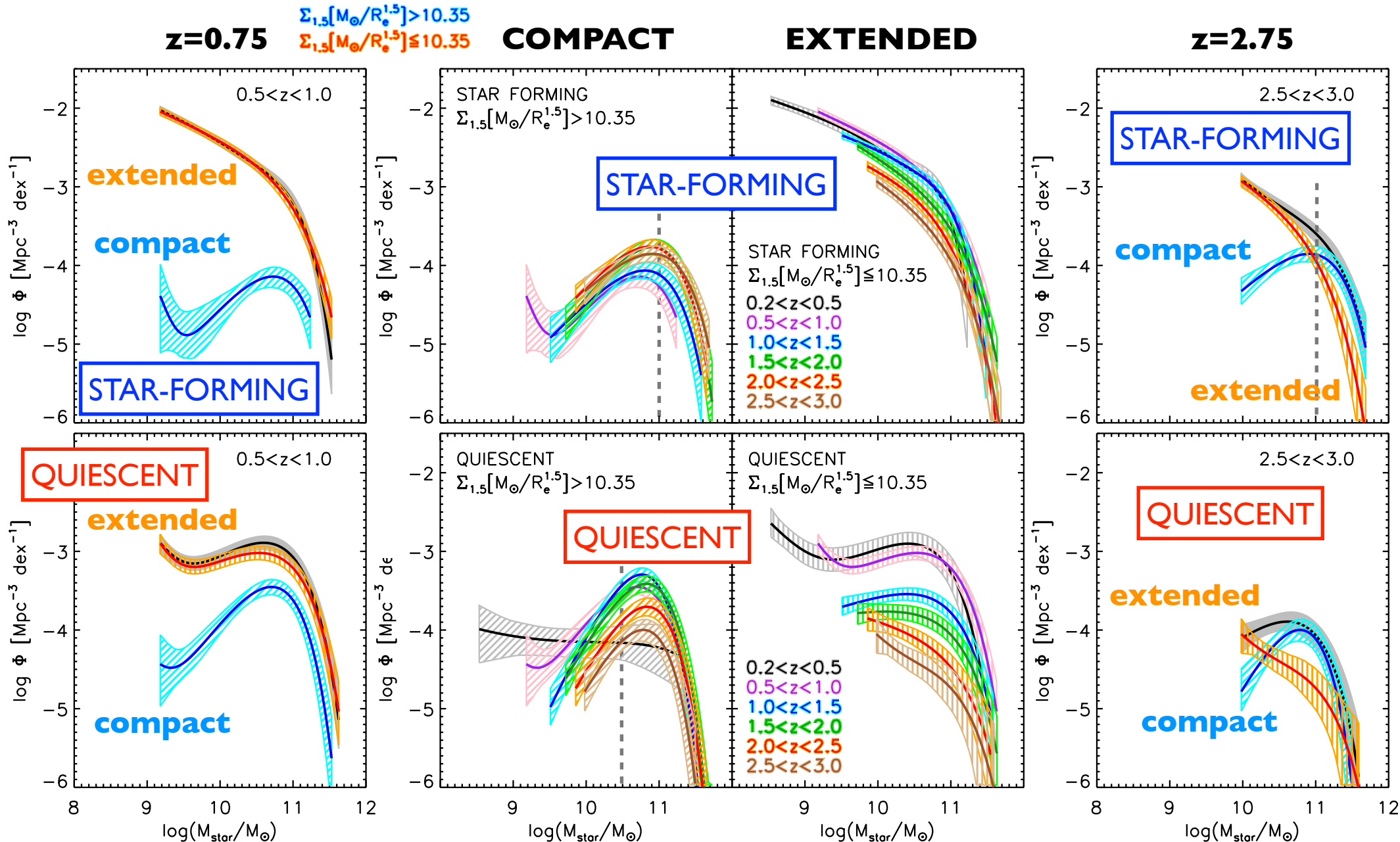


- The number density of small, SF lower-mass galaxies remains roughly constant with time, with slight decrease at $z < 1$
- The number density of intermediate size, lower-mass SF galaxies grows slowly with time; fast growth of the number density of large SF Galaxies.

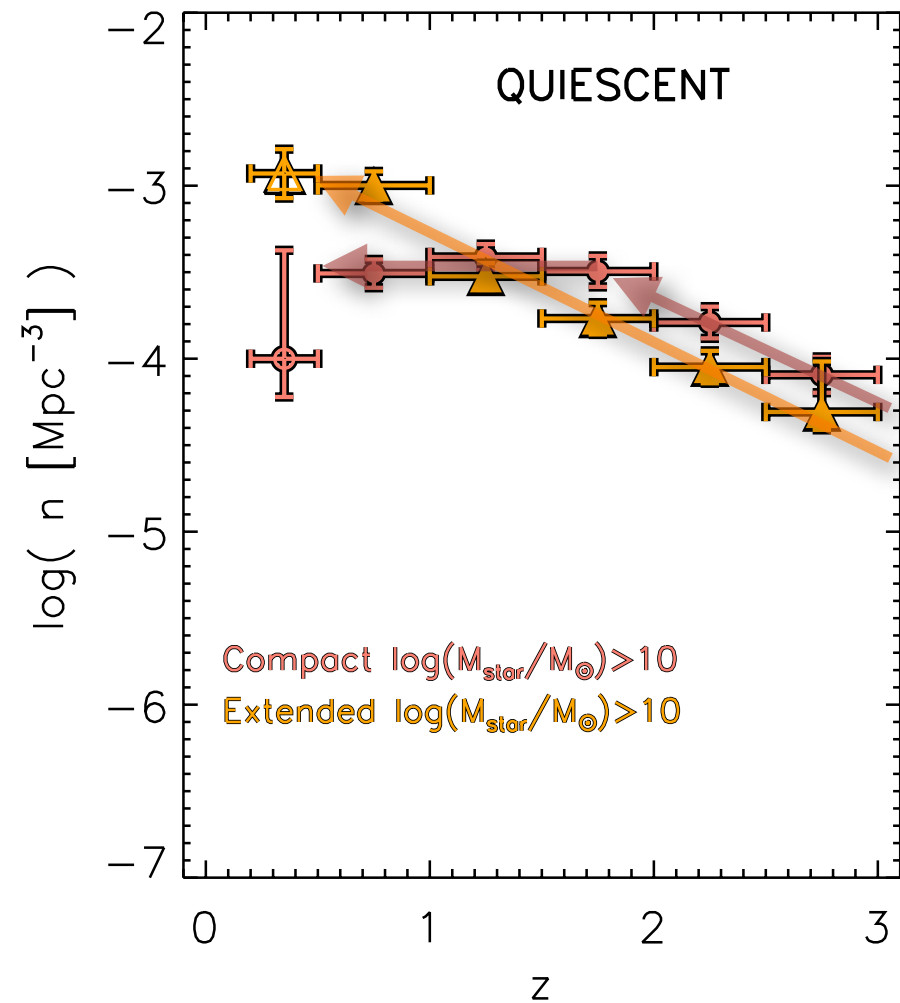
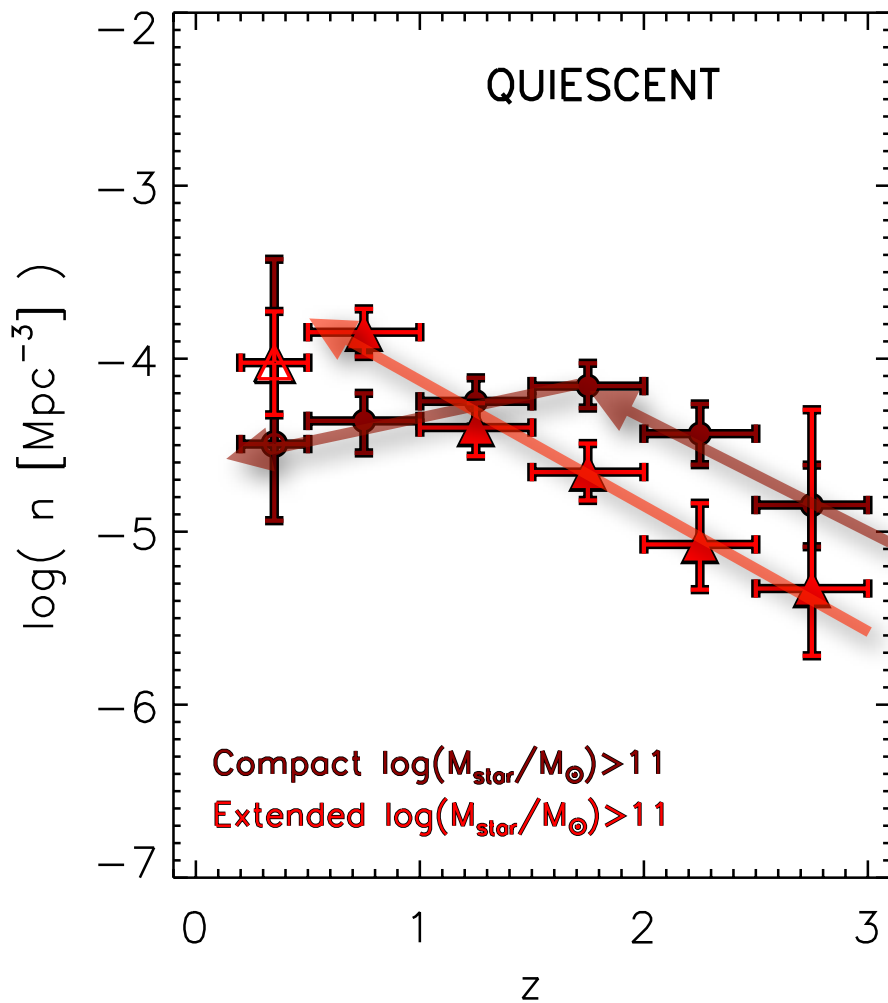
Evolution with z of the number density of **quiescent** and **star-forming** gals. of different **COMPACTNESS**



Evolution with z of the number density of **quiescent** and **star-forming** gals. of different **COMPACTNESS**

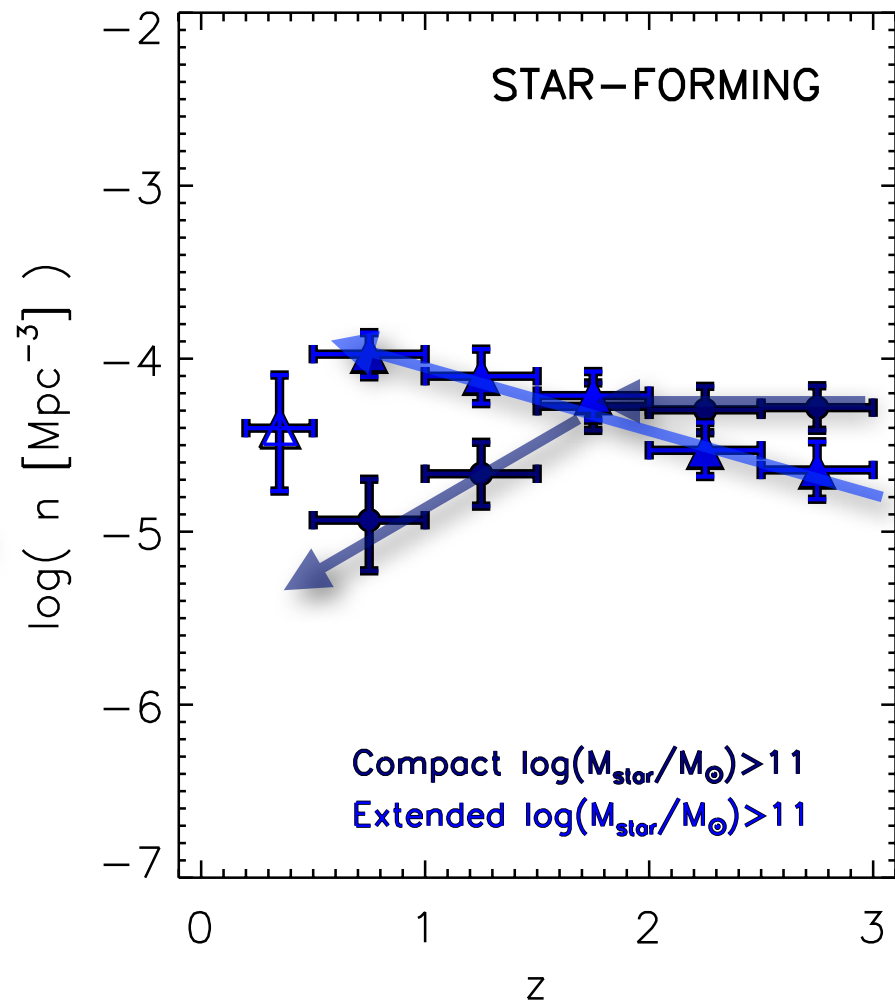
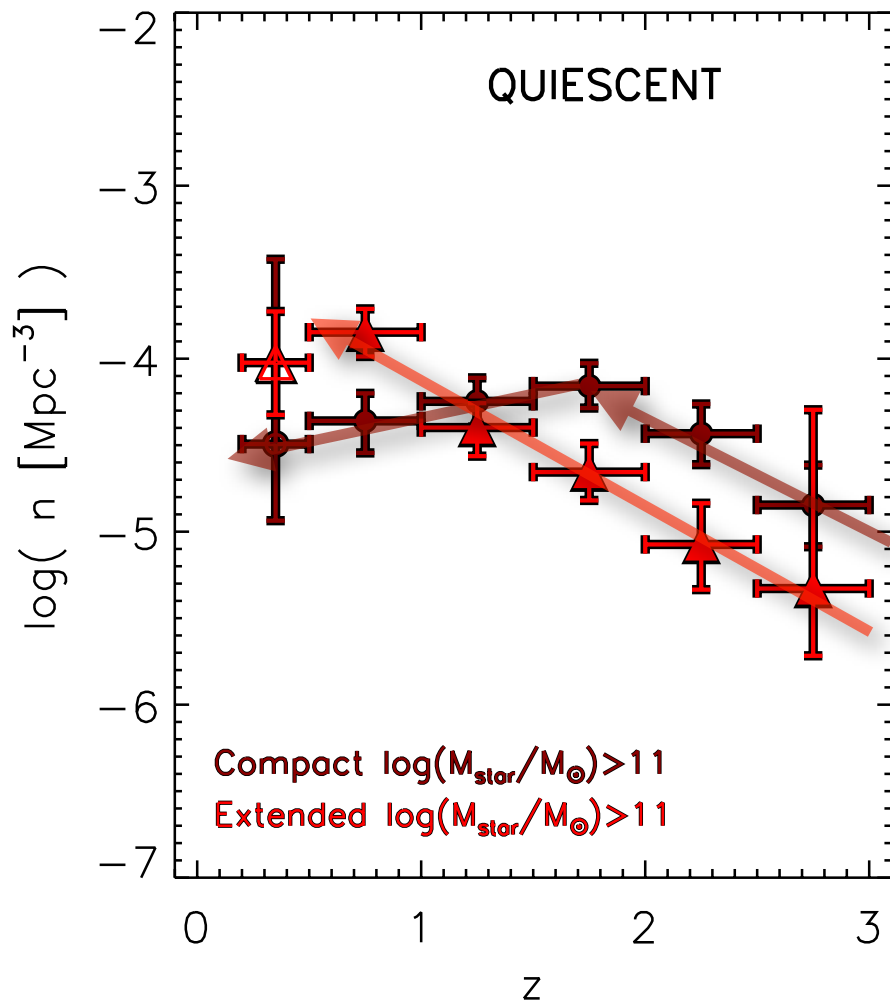


Evolution with z of the number density of **quiescent** galaxies of different **COMPACTNESS**



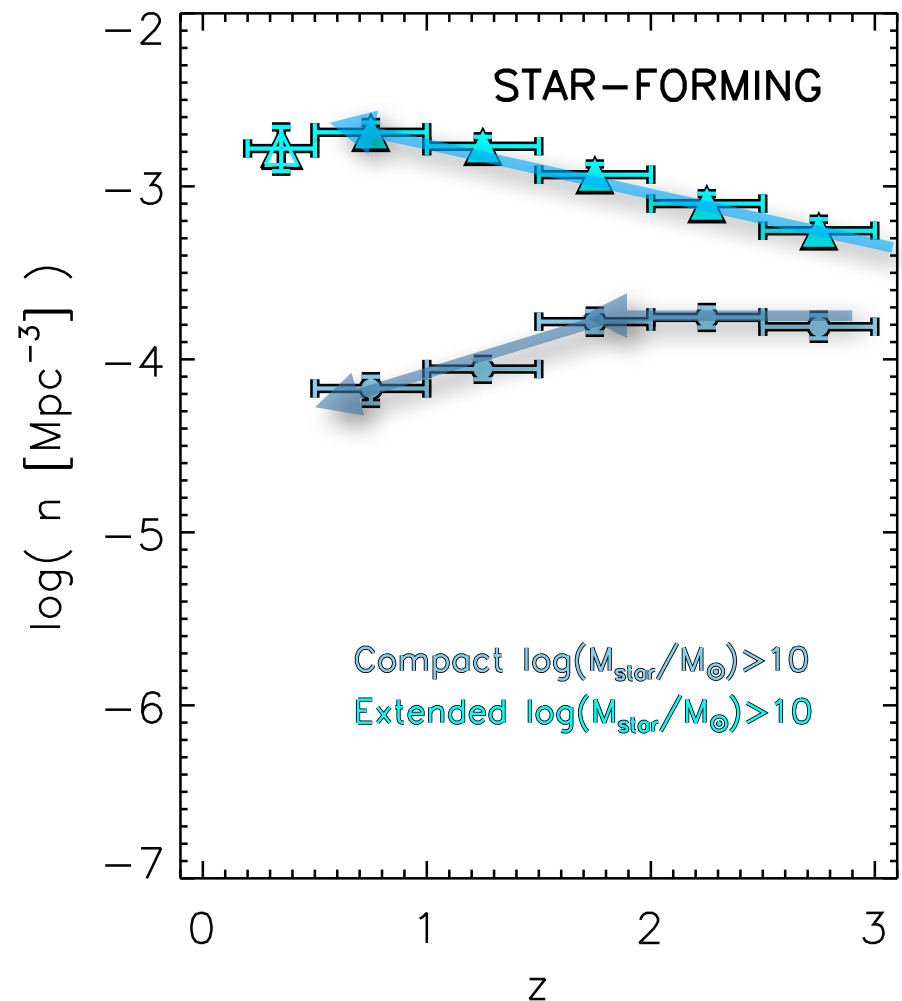
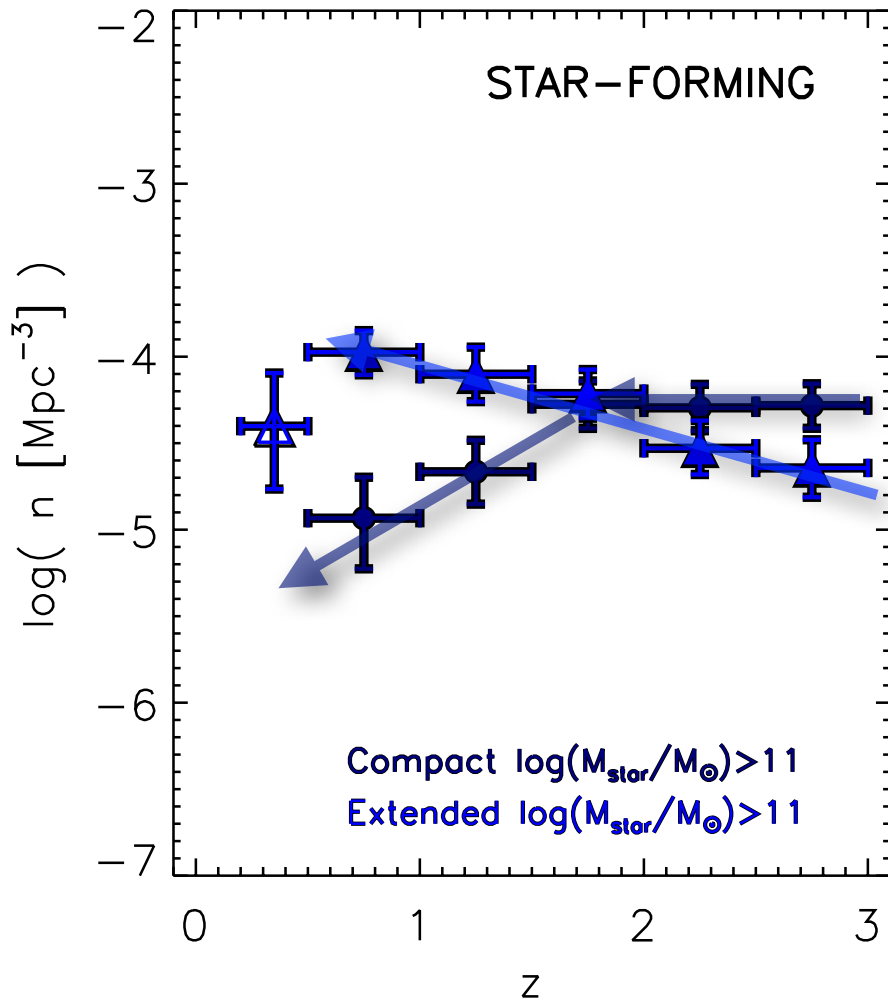
- Number density of compact massive ($\lg M > 11$) quiescent galaxies increases rapidly with time from $z=3$ to $z=1.5$, and then decreases slowly. The number density of extended massive quiescent galaxies increases steadily from $z=3$ to $z=0.5$ by 30x.
- Similar trends for lower-mass galaxies ($\lg M > 10$), but number density of compact quiescent galaxies consistent with remaining constant at $z < 2$

Evolution with z of the number density of **quiescent** galaxies of different **COMPACTNESS**



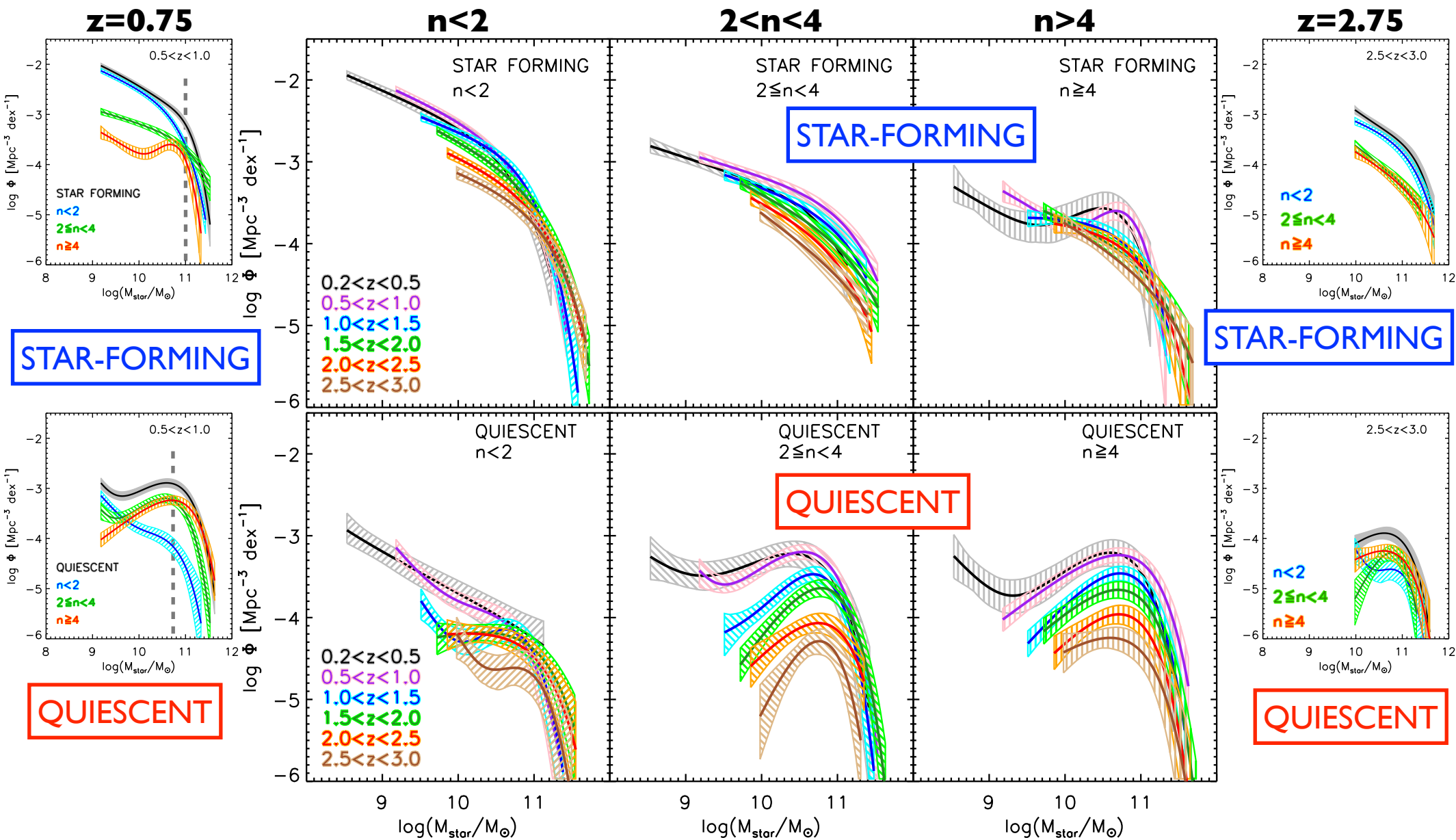
- Very different behavior for massive SF galaxies. The number density of compact SF galaxies remains constant from $z=3$ to $z \sim 1.75$, and then it decreases by 4x down to $z=0.5$.
- The number density of extended massive SF galaxies steadily increases with time by 7x from $z=3$ to $z=0.5$.

Evolution with z of the number density of **star-forming** galaxies of different **COMPACTNESS**

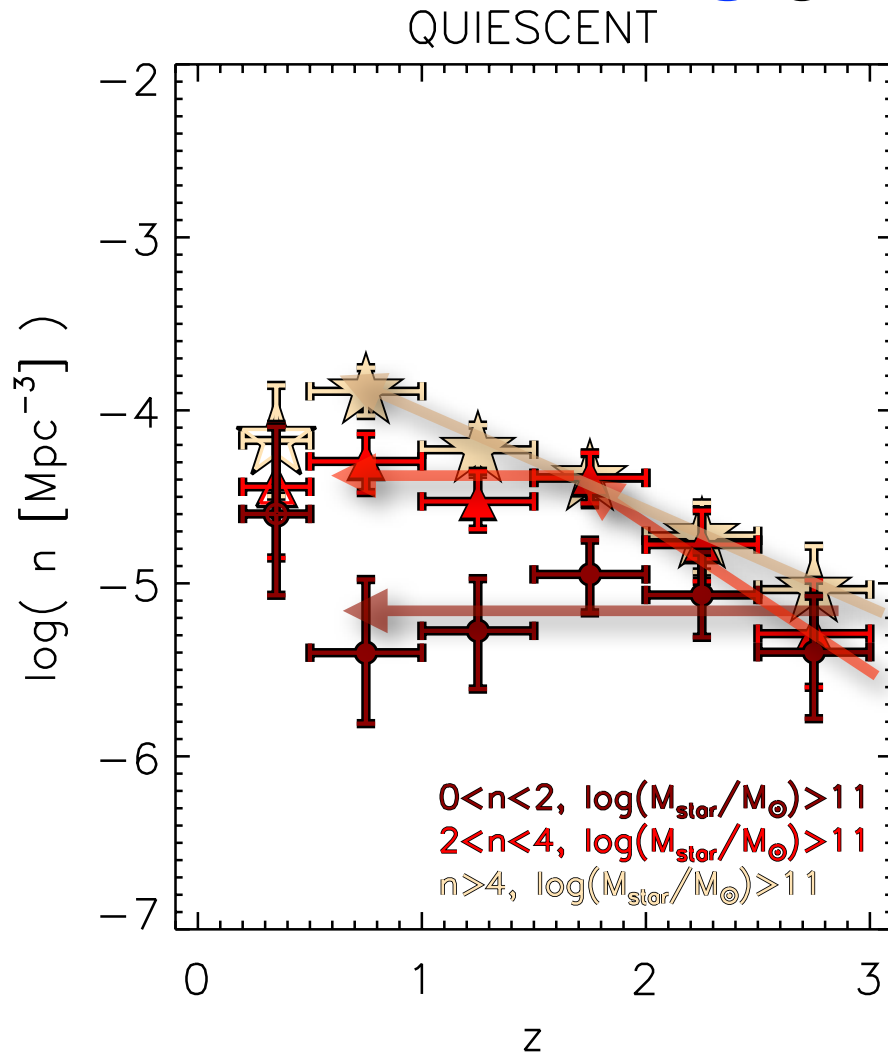


- SF galaxies with $\lg M = 10-11$ dominate the number density
- A similar trend is seen for lower mass SF galaxies, but extended lower-mass galaxies dominate the number density at all redshift, and their relative contribution to the overall SF population increases with time

Evolution with z of SMFs of **quiescent** and **star-forming** galaxies as a function of **SERSIC**

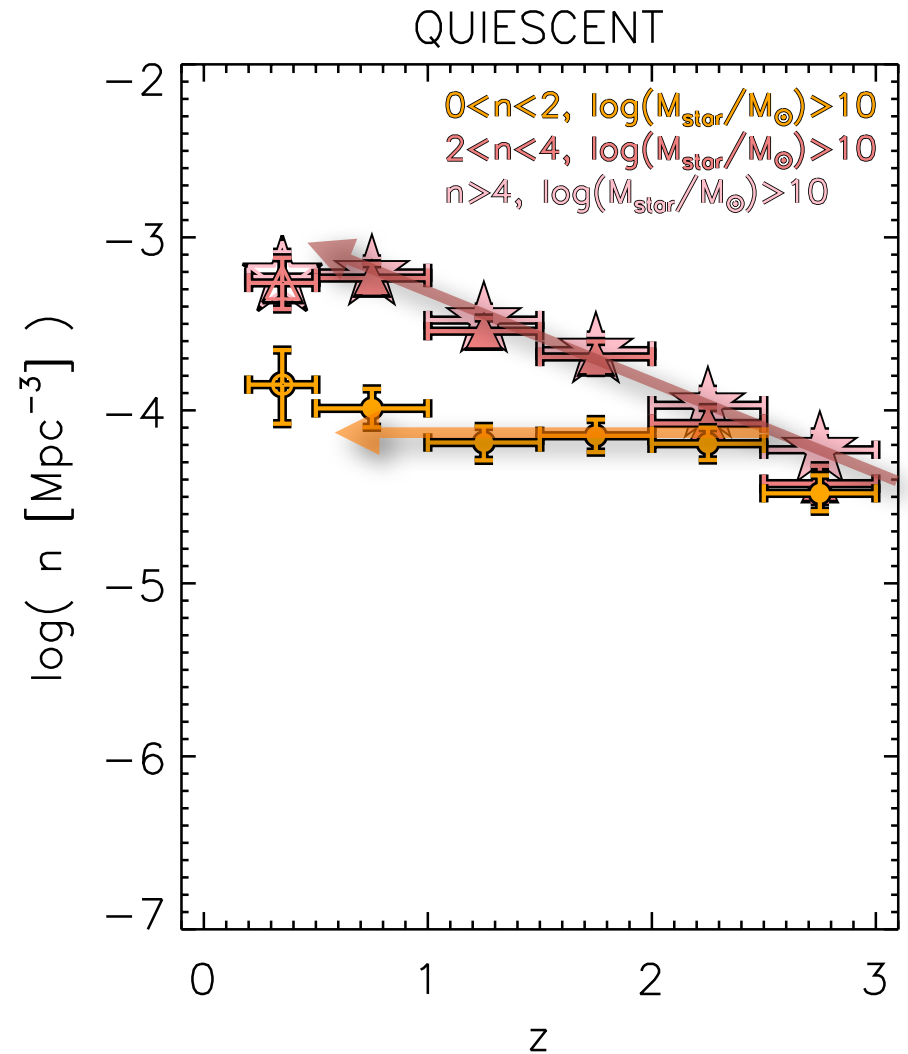
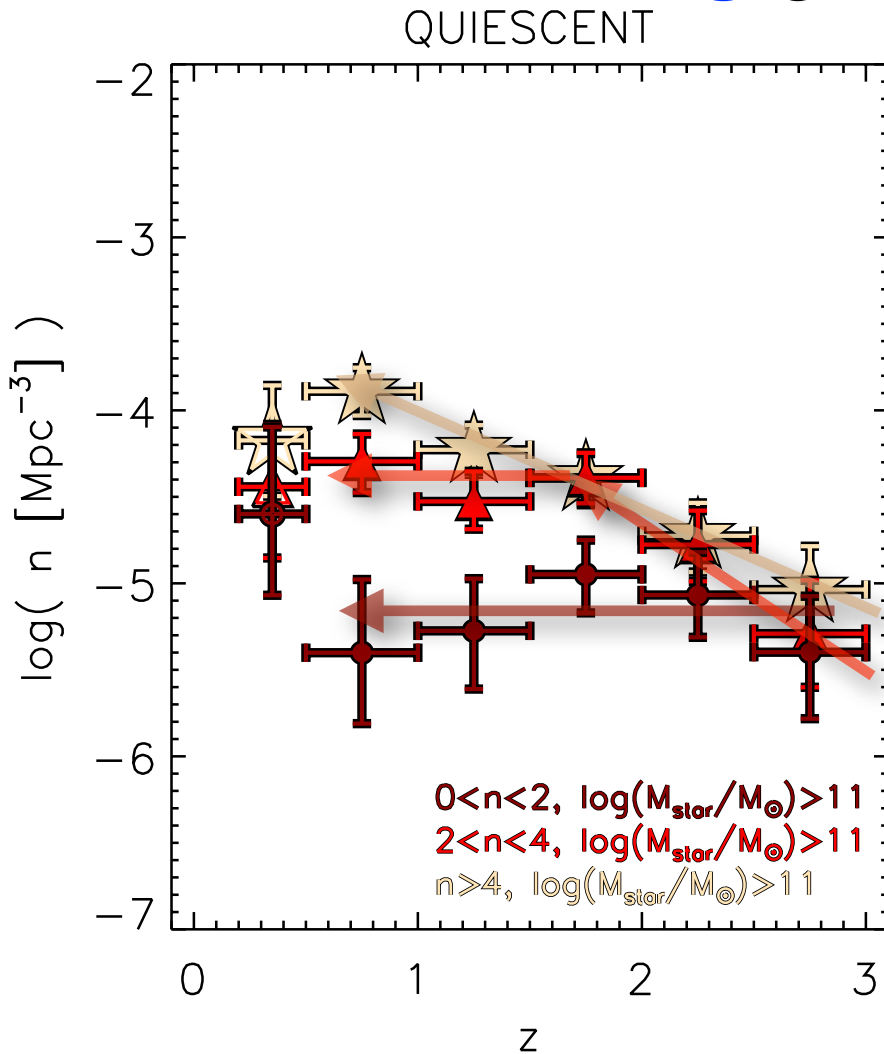


Evolution with of the number density of **quiescent** and **star-forming** gals. of different **SERSIC**



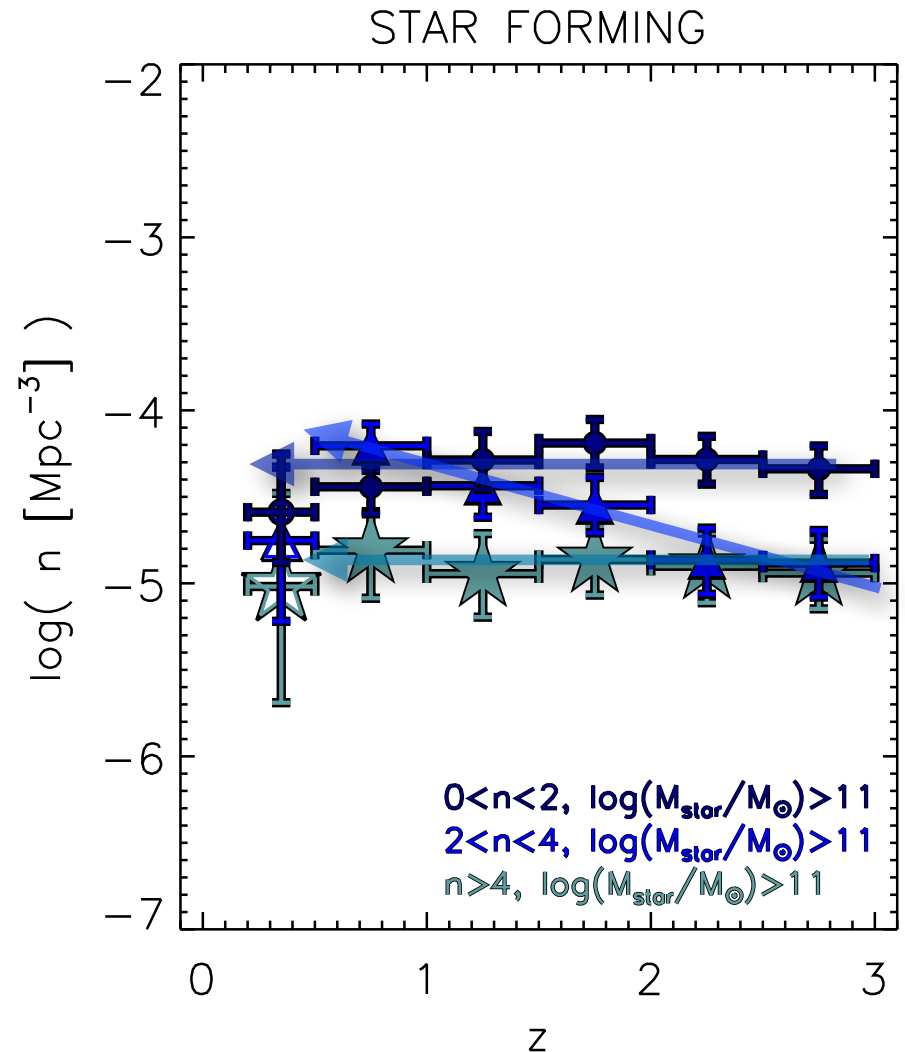
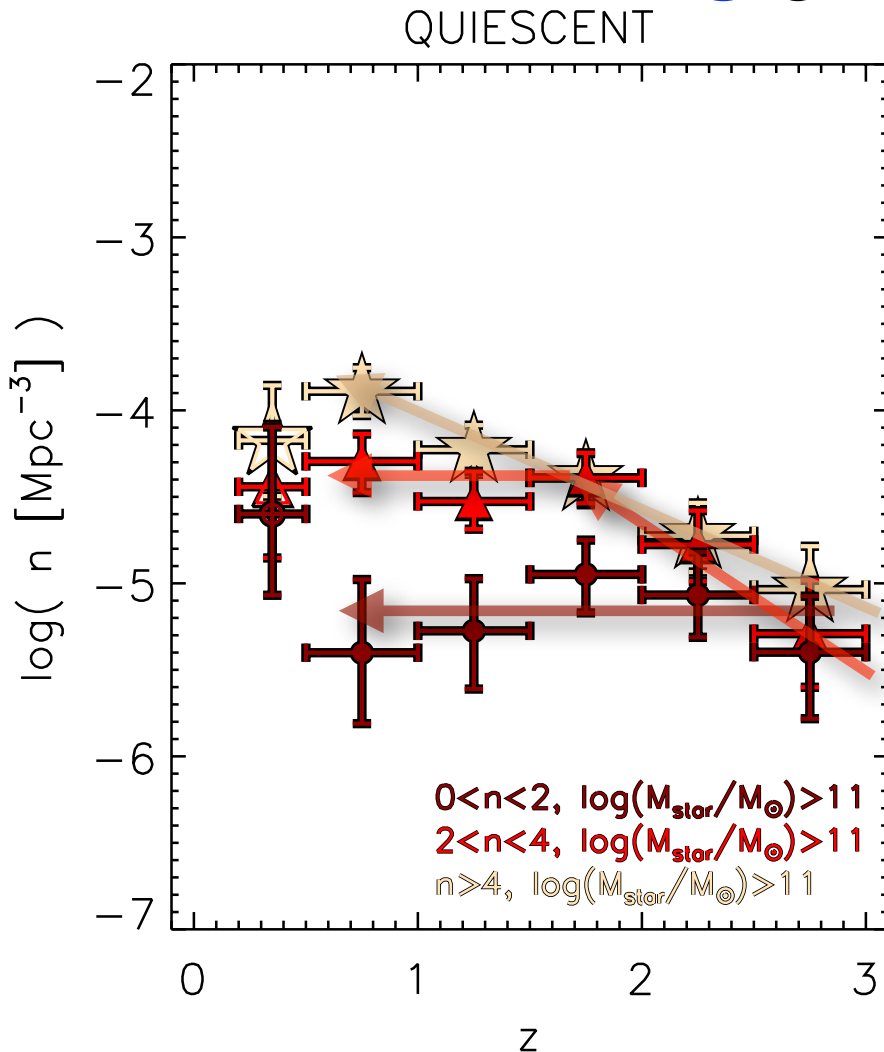
- For massive ($\lg M > 11$) quiescent galaxies, low- n galaxies contribute significantly to the number density of quiescent galaxies only at $2.5 < z < 3$; their number density remains roughly constant.
- The contribution of large and intermediate n galaxies increases rapidly with cosmic time.
- At $z > 1.5$, large and intermediate n galaxies contribute equally, while at $z < 1.5$ galaxies with $n > 4$ dominate the number density of massive, quiescent galaxies.

Evolution with of the number density of **quiescent** and **star-forming** gals. of different **SERSIC**



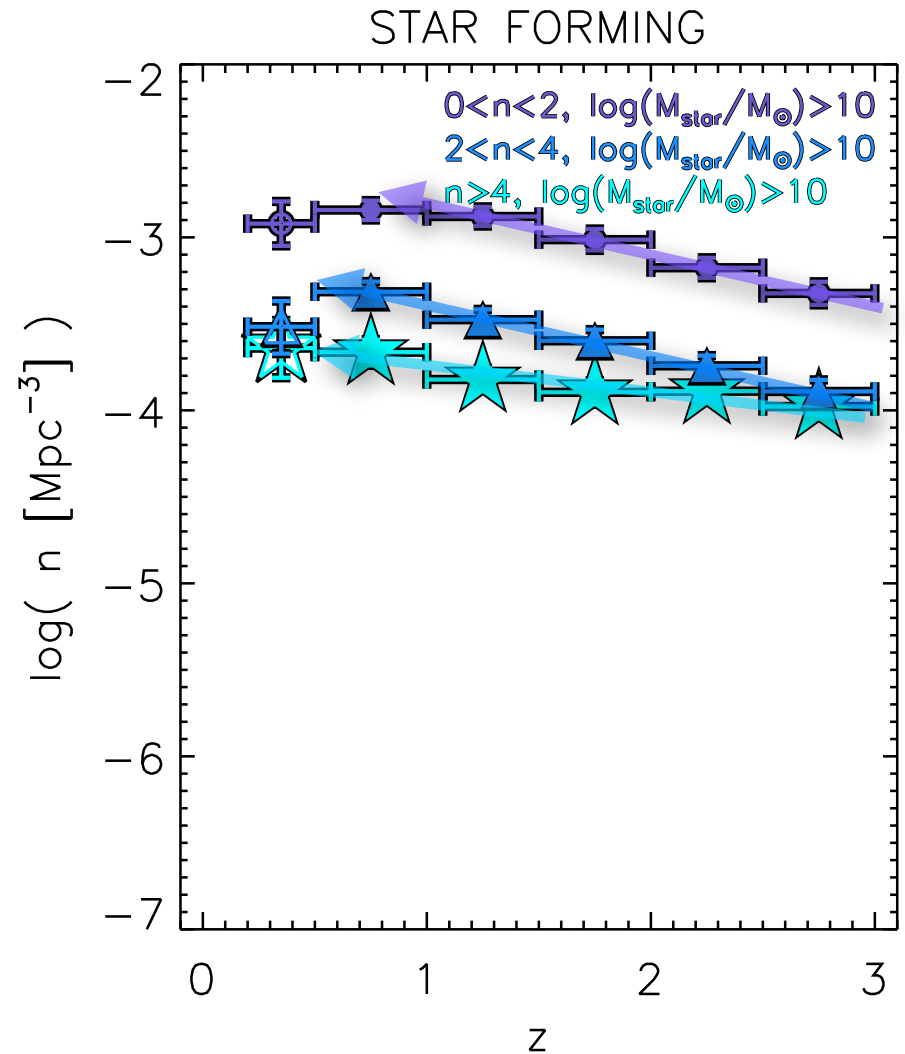
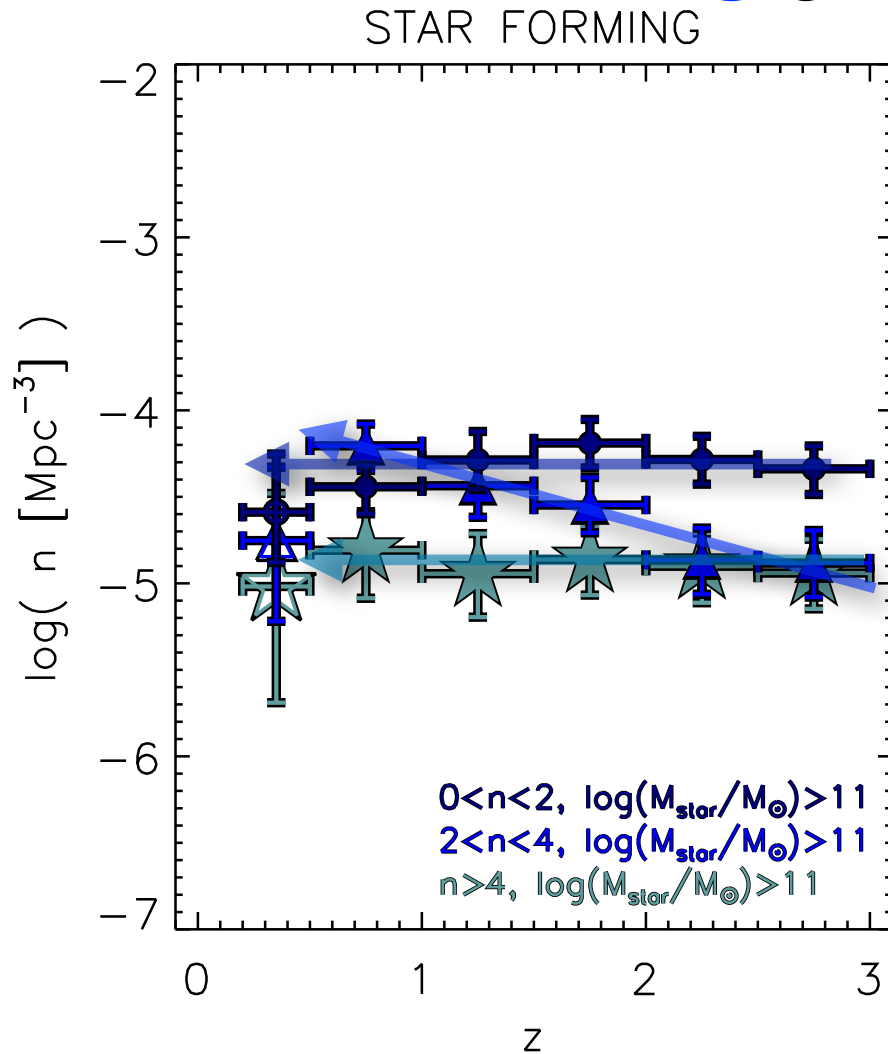
- For lower-mass quiescent galaxies, similar trends: significant growth (>10x) in the number density of galaxies with $n > 2$ from $z=3$ to $z=0.5$.
- Contribution of lower-mass quiescent galaxies with $n < 2$ negligible at $z < 2$, and relatively numerous only at high redshift.

Evolution with of the number density of **quiescent** and **star-forming** gals. of different **SERSIC**



- The number density of massive SF galaxies is dominated by low- n objects; their number density remains roughly constant with redshift
- SF galaxies with large n are rare and their number density does not evolve with redshift
- At high- z , SF galaxies with intermediate n are not numerous, but their contribution increases with time, becoming dominant at $z < 1$ at the high-mass end (e.g., bulge building up with time?)

Evolution with of the number density of **quiescent** and **star-forming** gals. of different **SERSIC**



- The number density of lower-mass star-forming galaxies is dominated by galaxies with $n < 2$ (3x more numerous than star-forming galaxies with $n > 2$)
- The number density of SF galaxies with $n < 2$ increases with time from $z=3$ to $z=0.5$ by 3x.
- The number density of SF galaxies with $n > 4$ grows much less with time, roughly constant from $z=3$ to $z=1$.

QUIESCENT GALAXIES:

- ☑ At high z , quiescent galaxies mostly small and compact, with very wide range in Sersic index; at $z < 1$, large/extended/ $n > 4$ galaxies dominate the massive end [$\lg(M_{\text{star}}/M_{\text{Sun}}) > 11$].
- ☑ At any redshift, most massive quiescent galaxies appear large/extended.
- ☑ Number density of small massive ($\lg(M_{\text{star}}/M_{\text{Sun}}) > 11$) galaxies grows from $z=3$ to $z=1.5$, and then decreases to lower z . Similar for compact massive galaxies.
- ☑ Dramatic build-up of large/extended galaxies with time. The number density of $n > 2$ galaxies increase with time rapidly, and by $z \sim 1.5$, $n > 4$ galaxies become dominant at $\lg(M_{\text{star}}/M_{\text{Sun}}) > 11$.

STAR-FORMING GALAXIES:

- ☑ At high redshift, SF galaxies mostly intermediate size and $n < 2$ objects; compact galaxies dominate the massive end. The SMF of small galaxies does not evolve at $M_{\text{star}} < 5 \times 10^{10} M_{\text{Sun}}$, whereas number density of large galaxies increases rapidly, especially at low-mass end.
- ☑ At low redshift, large galaxies with $2 < n < 4$ dominate the high-mass end; SF galaxies are extended at all masses. Low-mass galaxies have $n < 2$.
- ☑ **Q1:** are we seeing the build-up of the bulge component in disk galaxies?

- ☑ Galaxies grow bigger over time. Small quiescent galaxies disappear at high masses (i.e., they grow in size with time), BUT increase in number at low masses; yet, the number density of small star-forming galaxies remains roughly constant at the low-mass end.
- ☑ Small SF galaxies quench, and the number of small quiescent galaxies grows. IF MASSIVE, the small quiescent galaxies grow in size very efficiently; IF LOW MASS, small quiescent galaxies do not grow much in size.
- ☑ **Q2:** is this the effect of central/satellite? i.e., massive small galaxies grow in size by “eating” satellites, whereas the small, low-mass quiescent galaxies are the satellites...
- ☑ **Q3:** What is the best approach to interpret the observed evolution of the SMFs of quiescent and star-forming galaxies with different structural parameters?
This is a complementary approach to following progenitors/descendants...