

Big Present from Hitoshi

-----Beyond the Seesaw-----

IPMU 2024/12

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Anomaly Mediation

Hitoshi et al. (2000)

Supersymmetry

- Discovery of the supersymmetry --- *1971-1974* ---

I studied it from the papers by Abdus Salam
(I was a graduate student)

I was shocked by unification of fermion and boson!!!

$$\{Q_\alpha, \bar{Q}_{\dot{\beta}}\} = 2(\sigma^\mu)_{\alpha\dot{\beta}} P_\mu$$

- **Discovery of supergravity --- 1976 ---**

Further shock attacked me,
since they extended the space time by adding new
fermionic coordinates:

$$(x,y,z,t) \rightarrow (x,y,z,t, \theta)$$

Conceptual Big Change !!!

But I ignored such a theoretical discovery

1973; Neutral currents were discovered at CERN
1973; Asymptotic freedom was discovered
1974; Charm quark was discovered
1974; SU(5) GUT was proposed
1974; Lattice QCD was proposed
1975; Instanton solution in QCD was discovered
1975-1977; All anomalies in neutral currents were excluded
---- the SM established<---- 1978; Sakurai's talk
1977; Peccei-Quinn mechanism and Axion
1979; Glashow, Salam and Weinberg were awarded the Nobel prize
1979; Seesaw mechanism was discovered
1981; Inflation universe was proposed
1980-1981; Supersymmetric SM

- However, it brought about a big change in our physics
~1980

The quadratic divergence in the Higgs boson mass
is cancelled out between boson and fermion loops

The light Higgs boson of ~ 100 GeV
is technically natural !!!

- **SUSY standard model became very popular after ~1980**

The paradigm shifted to SUSY !!!

and

The SUSY paradigm lasts at present

- However, a lot of problems have been found

1. Gravitino problem; The decay of gravitino destroys the BBN Weinberg (1982)

→ $m(3/2) > 30-100 \text{ TeV}$

2. FCNC problem; We have too large FCNC decay of mesons and muon

→ $m(\text{squarks, sleptons}) > 100 \text{ TeV}$

3. Polonyi Problem; The decay of the SUSY breaking field Z destroys the BBN

$$\rightarrow m(Z) \sim m(3/2) > 100 \text{ TeV}$$

All above problems suggested the large SUSY breaking scale $> O(100) \text{ TeV}$ *~1980-1990*

But all DM candidates need to be lighter than $O(1) \text{ TeV}$

Dynamical SUSY Breaking

No Polonyi field and hence no Polonyi problem !!!

BUT

It predicts very light gauginos

$$m(\text{gauginos}) < \mathcal{O}(1) \text{ GeV}$$

No hope for the Dark Matter !!!

BUT

Hitoshi gave us a Big Present

at 2000

Anomaly Mediation

Quantum gravity effects generate the gaugino masses !!!

$$m(\text{gauginos}) = \mathcal{O}(1) \text{ TeV}$$

A gaugino can be a Dark Matter !!!

The anomaly mediation;

$$m(\text{gaugino}) = \alpha \frac{\Lambda^2}{M_{PL}}$$

The seesaw:

$$m(\text{neutrino}) = f^2 \frac{\langle H \rangle^2}{M_R}$$

“Pure gravity mediation” was proposed

Ibe, Moroi, et al. (2005)

$m_{3/2}=30-100\text{TeV}$

$m_H=120-130\text{ GeV}$

$m(\text{wino DM})=1-2\text{ TeV}$

Anomaly Mediation Saved Us

*The **wino DM** will be discovered in direct and indirect detection experiments*

*The Hubble constant H_i must be $> 10^{13}$ GeV
The tensor/scalar ratio **$r > 0.001$***

Harigaya, Ibe ... (2015)

Dear all participants

Let us thank Hitoshi for his big present

The bright our future's coming soon!!!