

MINERvA's Flux Prediction

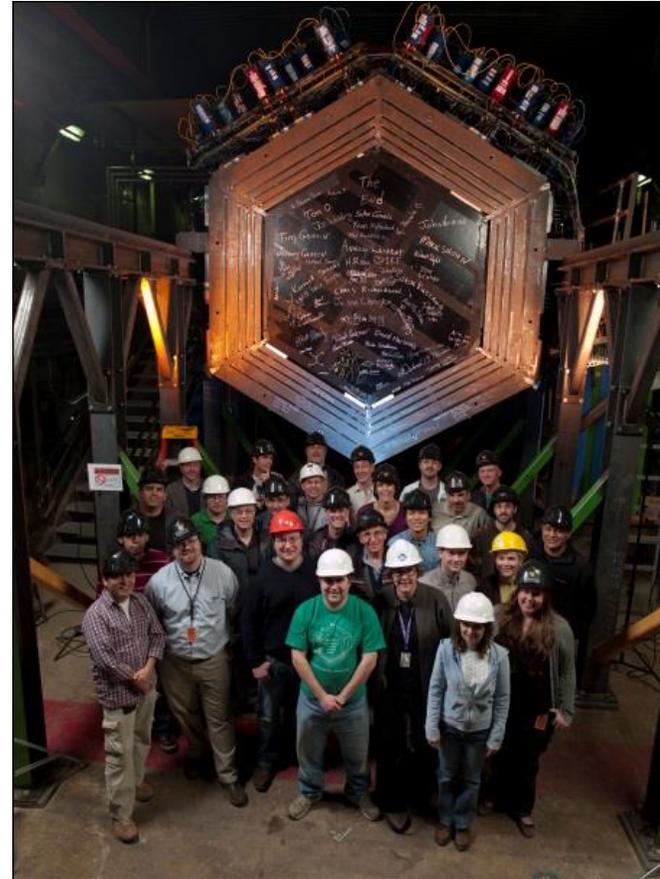
Tomasz Golan

(University of Rochester / Fermilab)

On behalf of MINERvA Collaboration

Main INjector ExpeRiment v-A aka MINERvA

- MINERvA is a neutrino scattering experiment at Fermilab
- Collaboration of about 65 nuclear and particle physicists from more than 20 institutions
- NuMI beam is used to measure cross section for neutrino-nucleus interactions
- Detector includes several different nuclear targets which can be used to measure nuclear effect

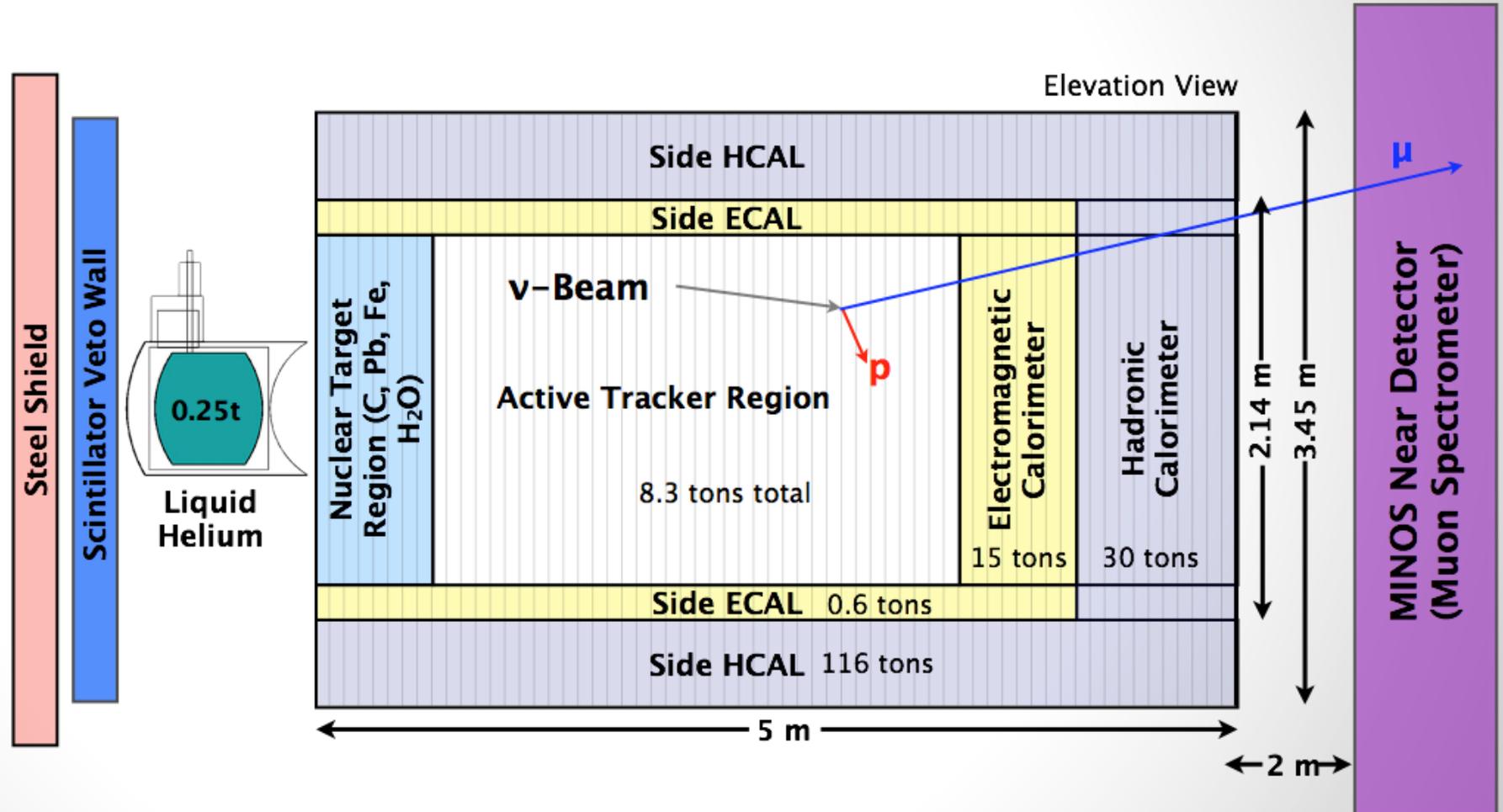


MINERvA Detector

120 modules of 4 types:

- nuclear targets
- scintillator
- ECal
- HCal

MINOS near detector as muon spectrometer



MINERvA's Flux Prediction

Previous analyses used flux prediction known as
Generation 0 and Generation 1

Flux	Analysis	Reference
Generation 0	ν_μ CCQE	PRL 111 (2013) 022502
Generation 0	anti ν_μ CCQE	PRL 111 (2013) 022501
Generation 1	CC target ratios	PRL 112 (2014) 231801
Generation 1	Coherent π	PRL 113 (2014) 261802
Generation 1	ν_μ muon+proton	PRD 91 (2015) 071301
Generation 1	anti ν_μ CC π^0	PLB 749 (2015) 130-136

Generation 2 includes many improvements:

- Geometry
- Focusing uncertainties
- Additional hadron production data
- Better accounting for beam attenuation
- New framework allows to use MIPP thick target data
- And allows predictions of the flux for NOvA and uBooNE

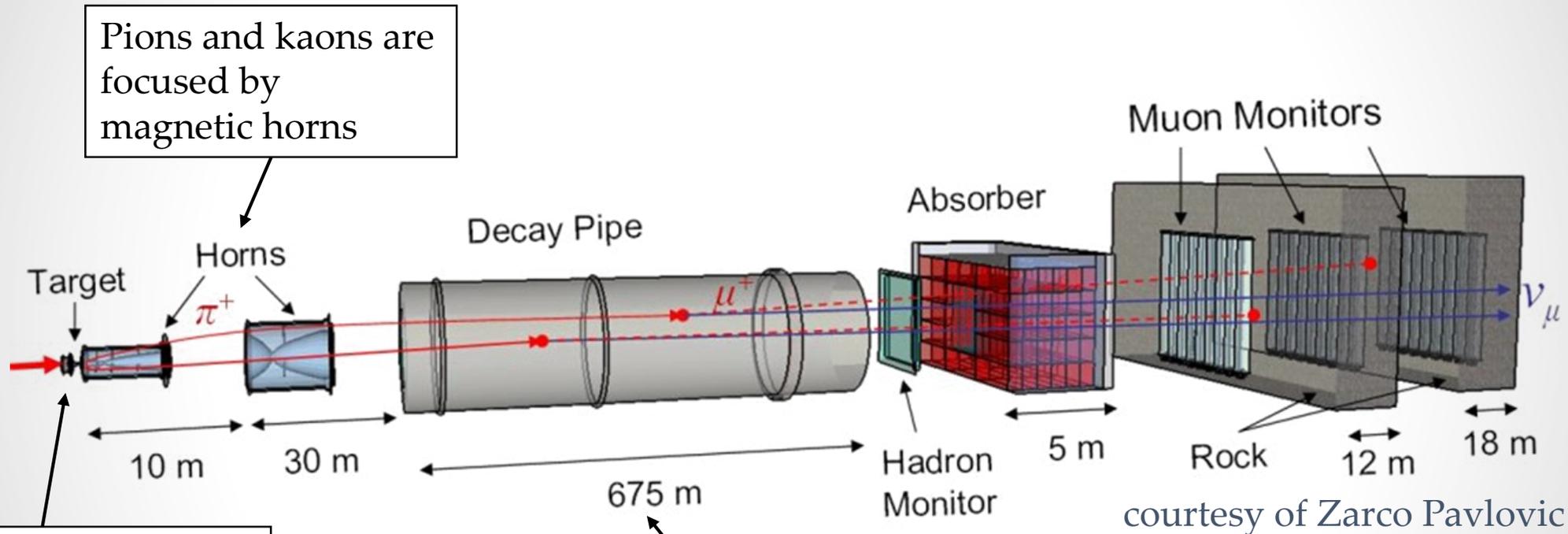
MINERvA @ NUINT15

There are many improvements in our understanding of NuMI beam. Do not miss other MINERvA's talk to see how it affects our results:

Talk	Speaker	Date	Flux
ν_e CCQE results	J. Wolcott	Tue, 14:05	gen1 with ν -e constraint
ν_μ CCQE update	M. Betancourt	Tue, 14:30	gen1 for muon + proton gen2 update on previous CCQE
Search for MEC	P. Rodrigues	Tue, 15:25	gen2
DIS in nuclear targets	J. Mousseau	Thu, 10:15	gen1

NuMI Beam

NIM A806 (2016) 279-306



Pions and kaons are focused by magnetic horns

120 GeV protons strike graphite target and meson mess is created

Pions and kaons decay (note, they can also interact!)

Need to understand each step since proton hits the target to neutrino production

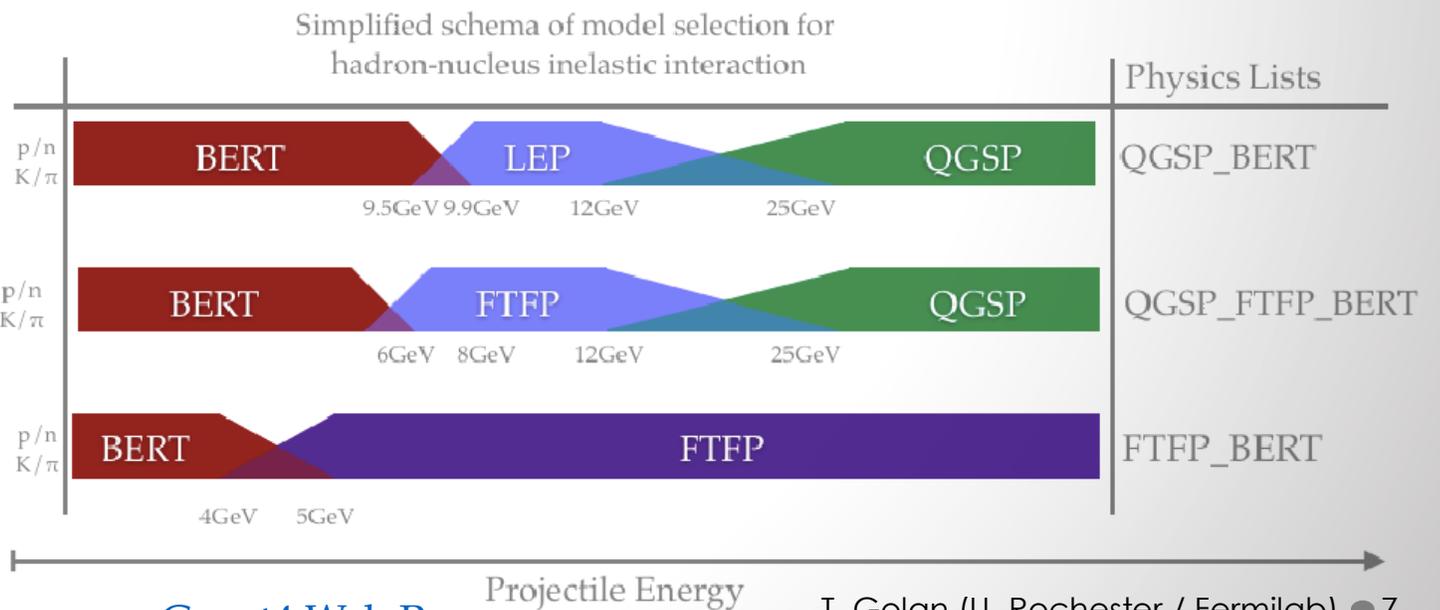
courtesy of Zarco Pavlovic

NuMI Beam Simulation

- Flux simulation starts with a Geant4 simulation of the NuMI beamline (G4NuMI)
- Geant 4.9.2p03 is used with FTFP_BERT hadronic physics list
- The results of the simulation are corrected by external data
- Similar approach to the one T2K used (Phys. Rev. D 87, 012001 (2013))

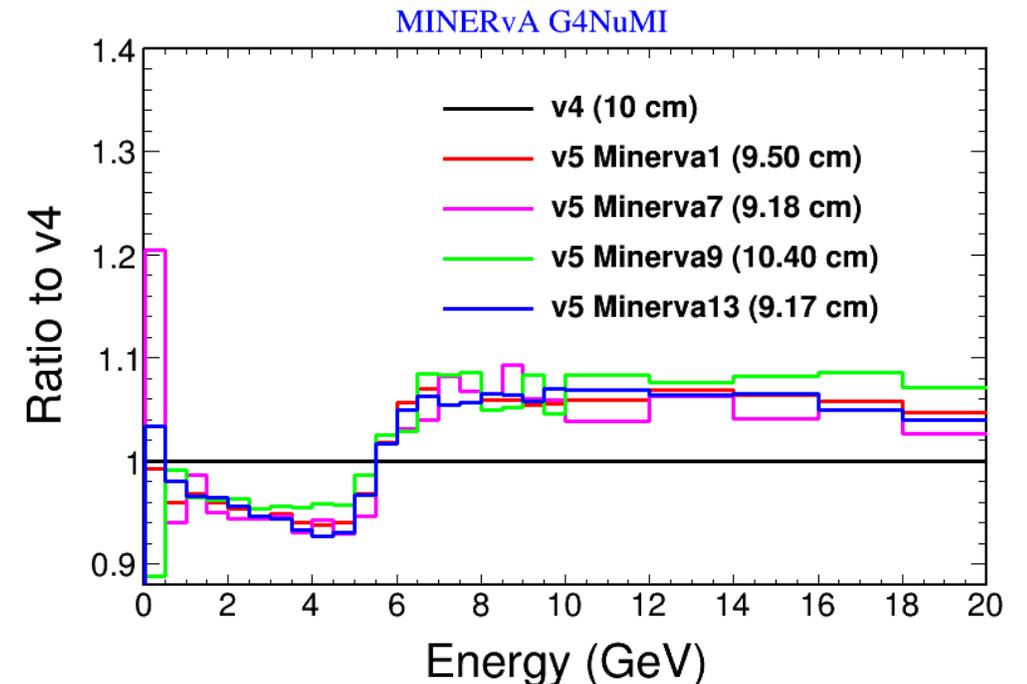
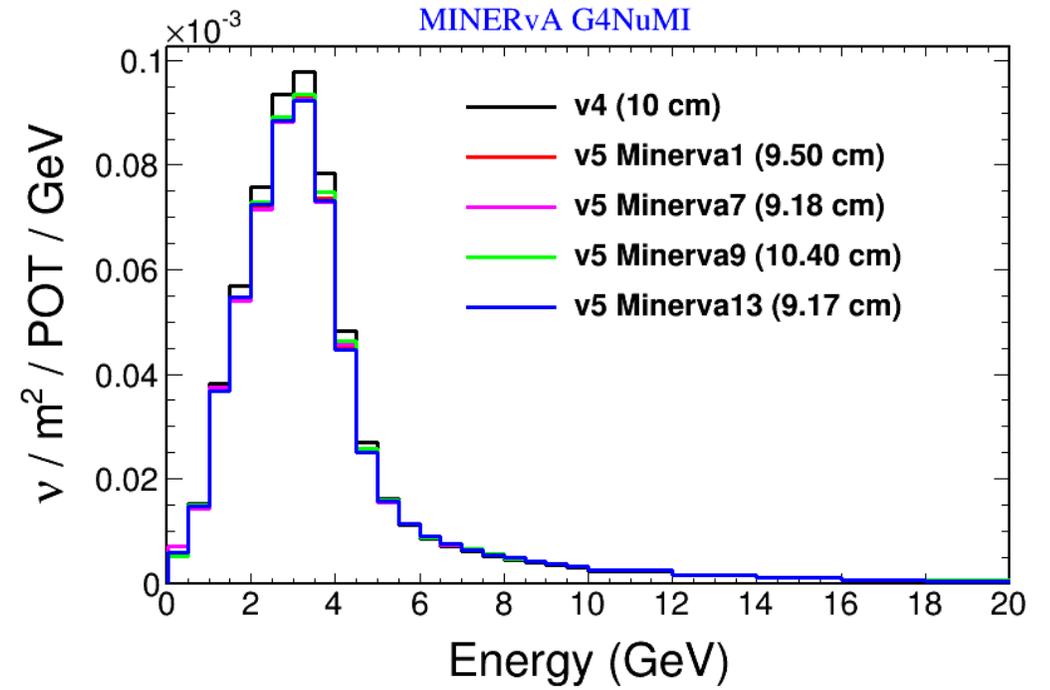
BERT – Bertini intranuclear cascade

FTFP – The FRITIOF Precompound



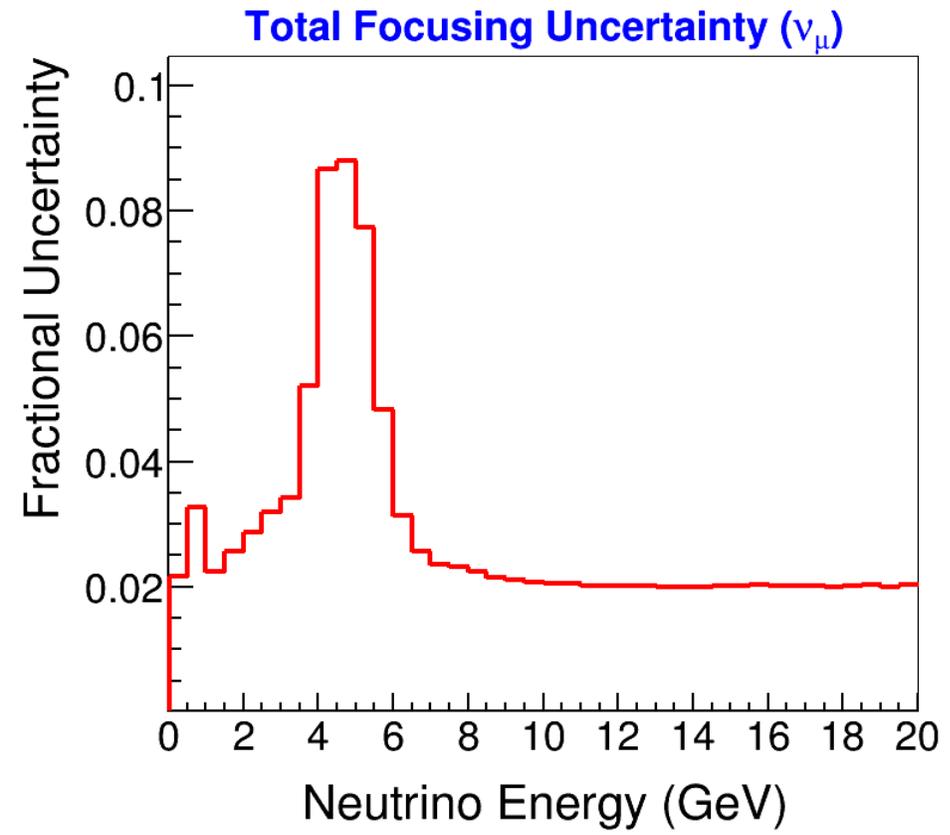
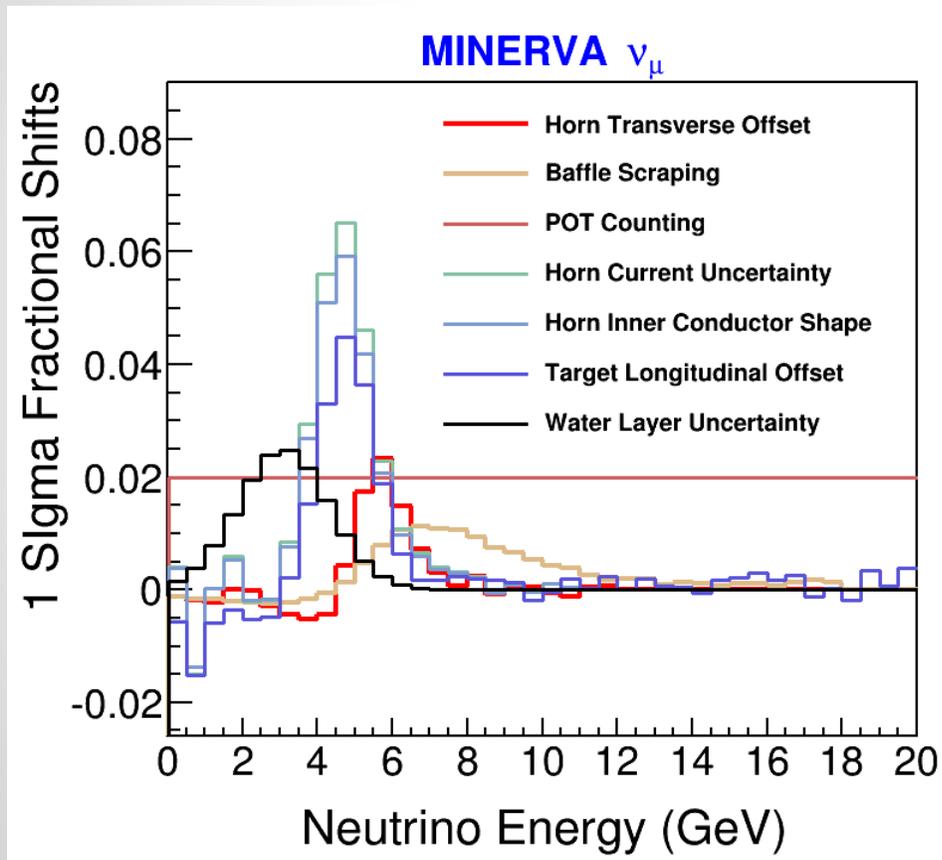
Geometry improvements

- Thin layer (1.0 +/- 0.5mm) of cooling water around the inner conductor
- Beryllium windows (0.5 mm each) at the entrance and at the exit of NuMI baffle
- Argon instead of Air inside the magnetic horns
- The longitudinal target position

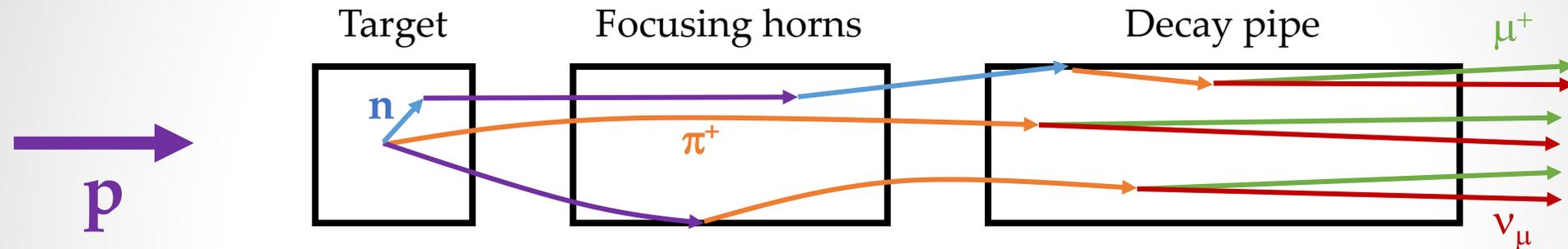


Focusing uncertainties

Horn geometry and water layer was (re)studied to evaluate focusing uncertainty



Hadron interactions

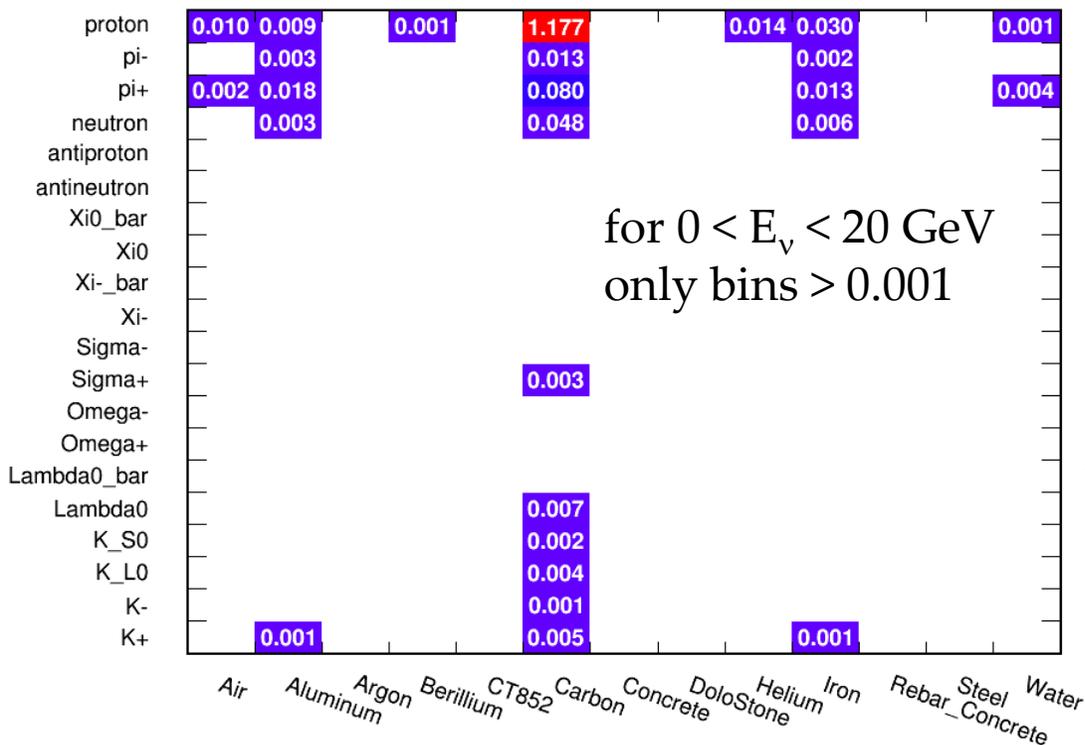


- All interactions leading to neutrino production are stored in ntuple
- Including whole kinematics and material
- This information is used to apply corresponding data for our reweighting procedure

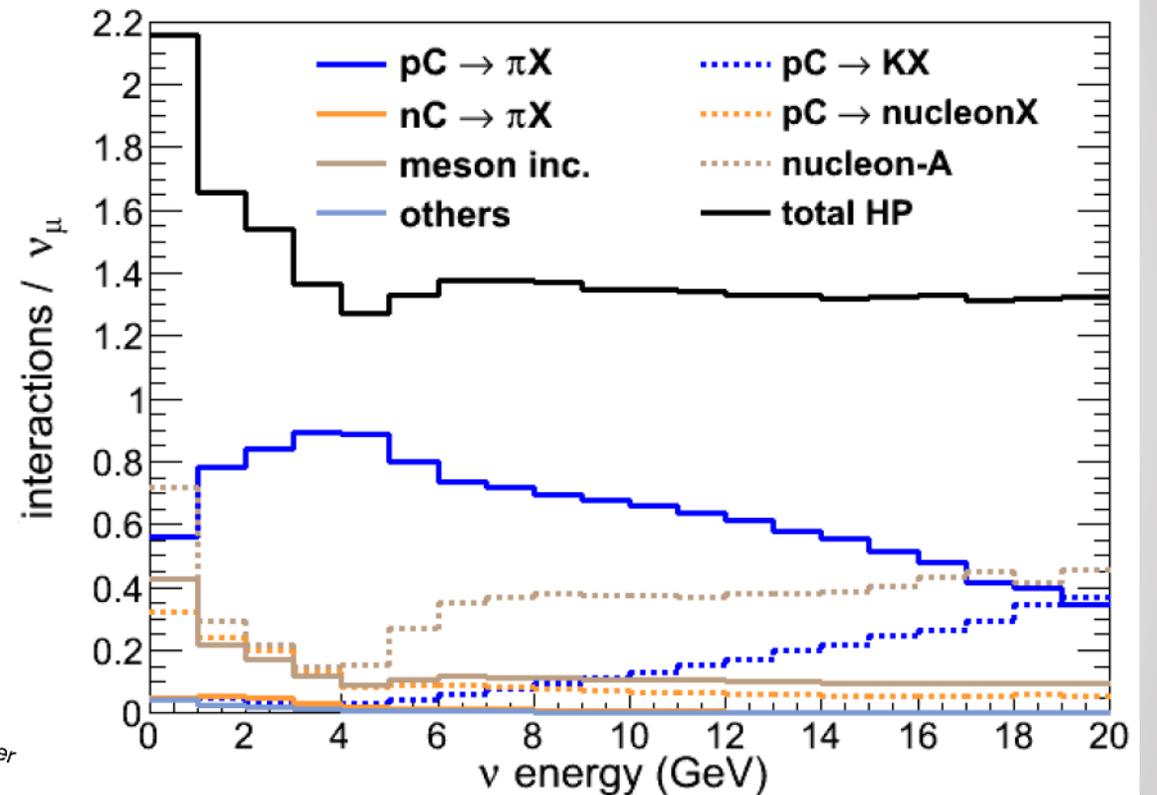
What interacts where

Number of interactions per neutrino at the detector

Projectile particle

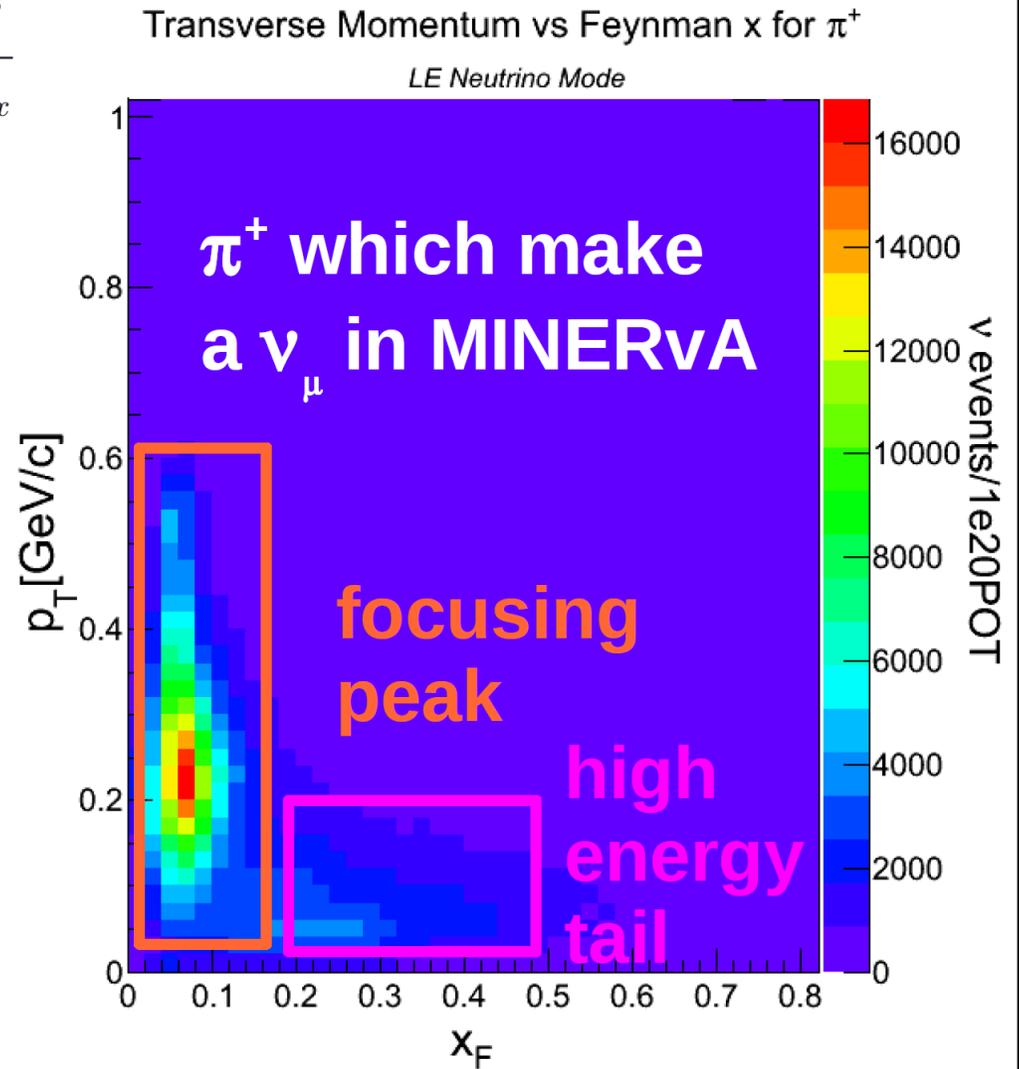
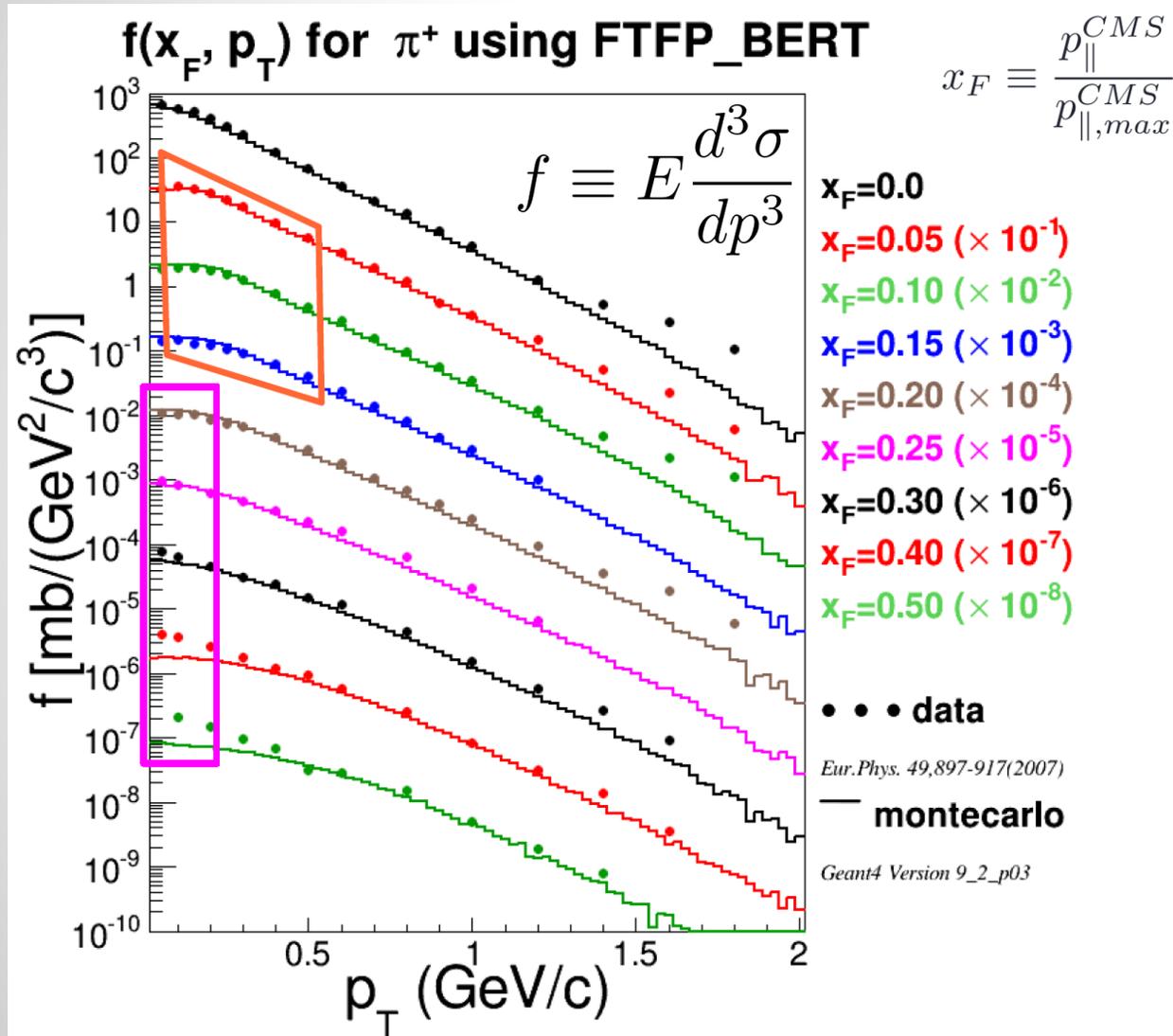


Target material



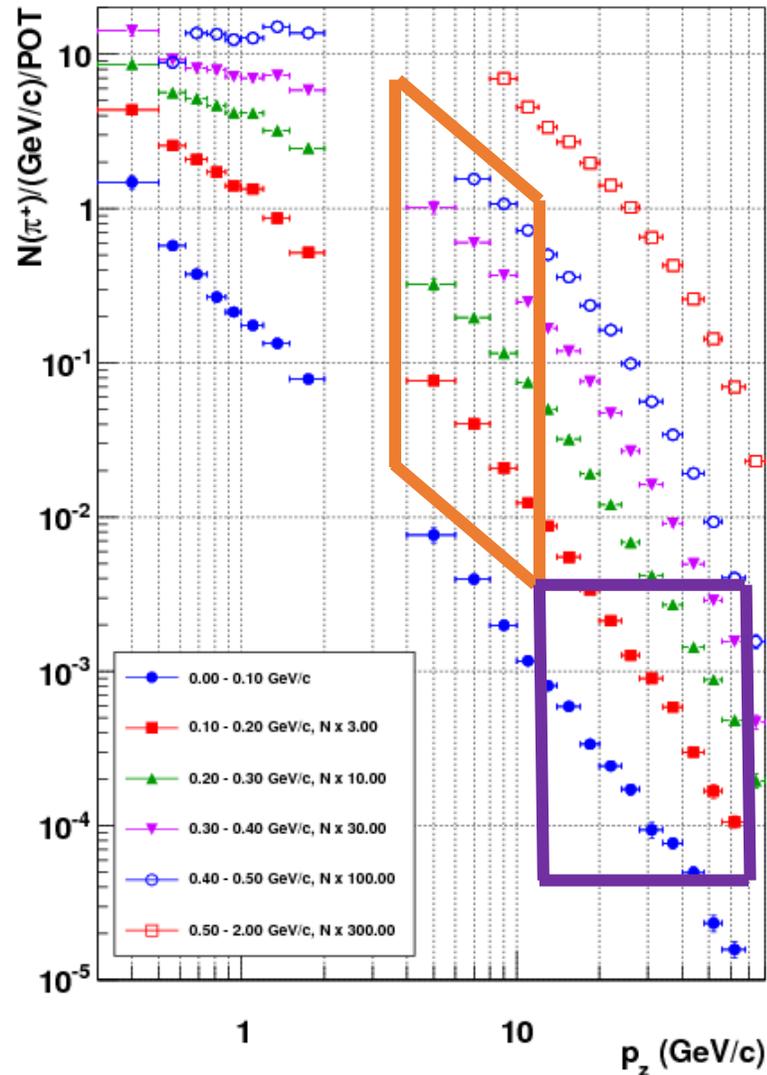
- As simulated by g4numi
- Most of it is covered by NA49 and MIPP data

NA49 data



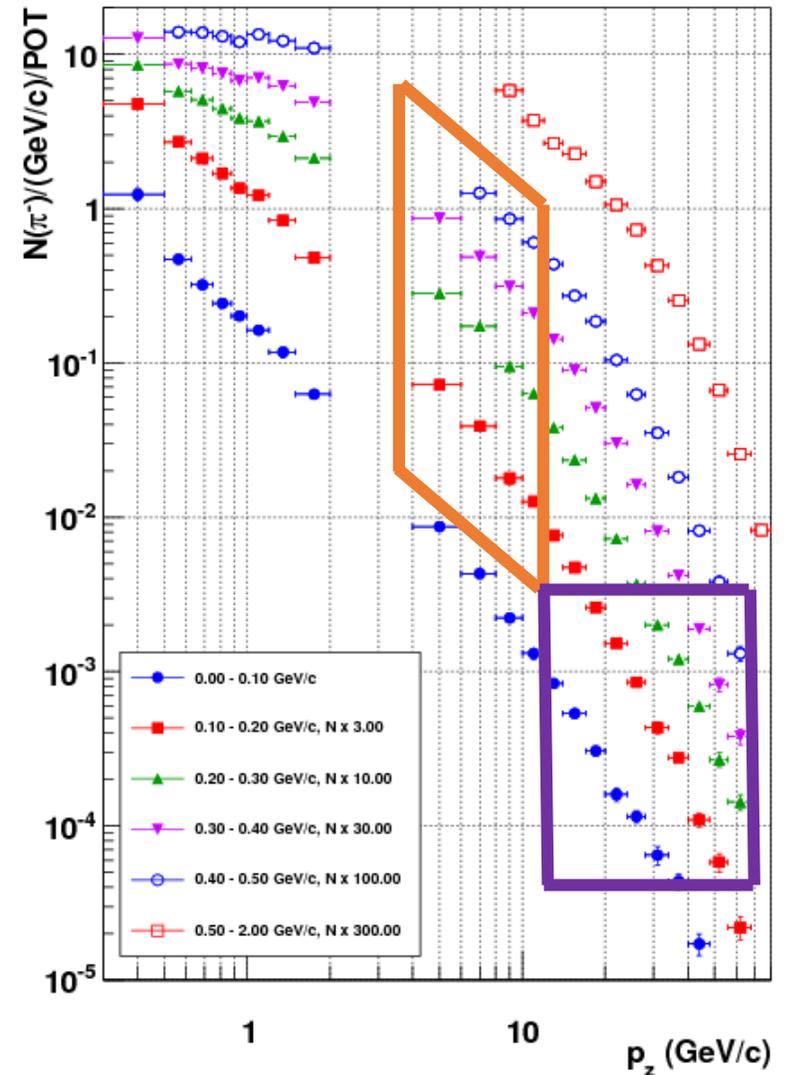
MIPP data

- MIPP is hadron production experiment located at Fermilab
- Pions yields off the NuMI target are measured
- Ratios of kaon/pion yields are measured on thin target and NuMI replica



Low Energy Focusing Peak

Low Energy High Energy Tail



Reweighting

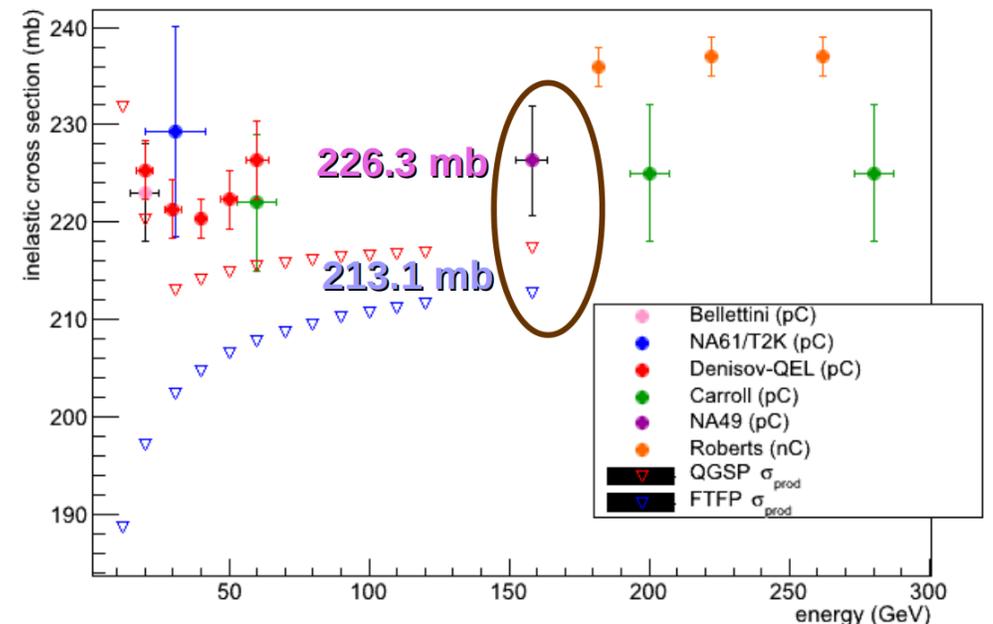
- Using external hadron production data events are weighted by:

$$w_{HP} = \frac{f_{data}(x_F, p_T, E)}{f_{MC}(x_F, p_T, E)}, \quad f \equiv \frac{1}{\sigma} E \frac{d^3\sigma}{dp^3}$$

- The beam attenuation (due to mismodeling of total inelastic cross section) is taken into account by applying a second weight (for each volume):

$$w_{att} = e^{-L\rho(\sigma_{data} - \sigma_{MC})}$$

Data of inelastic cross section p(n)C



Uncertainties

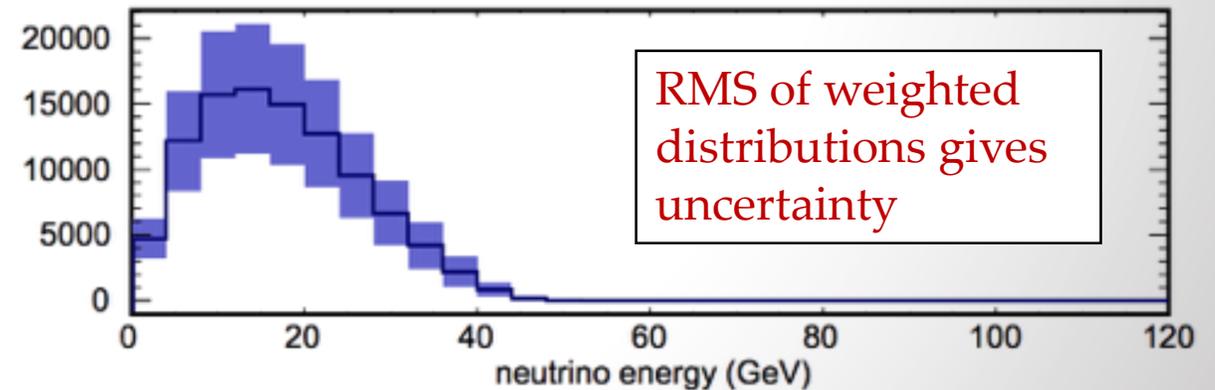
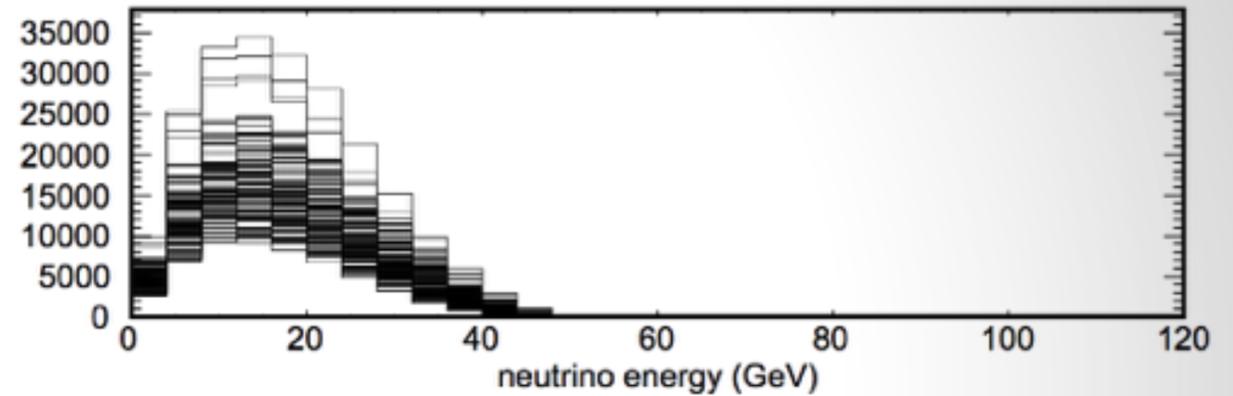
by M. Kordosky

- „Many-Universes” method is used to propagate external data uncertainties to our flux

$$w_u = w_{att,u} \cdot \prod_i w_{HP,u,i}$$

for each universe (u) data central value is randomly shifted according to uncertainty

product of HP weights for all reweightable interactions (i)

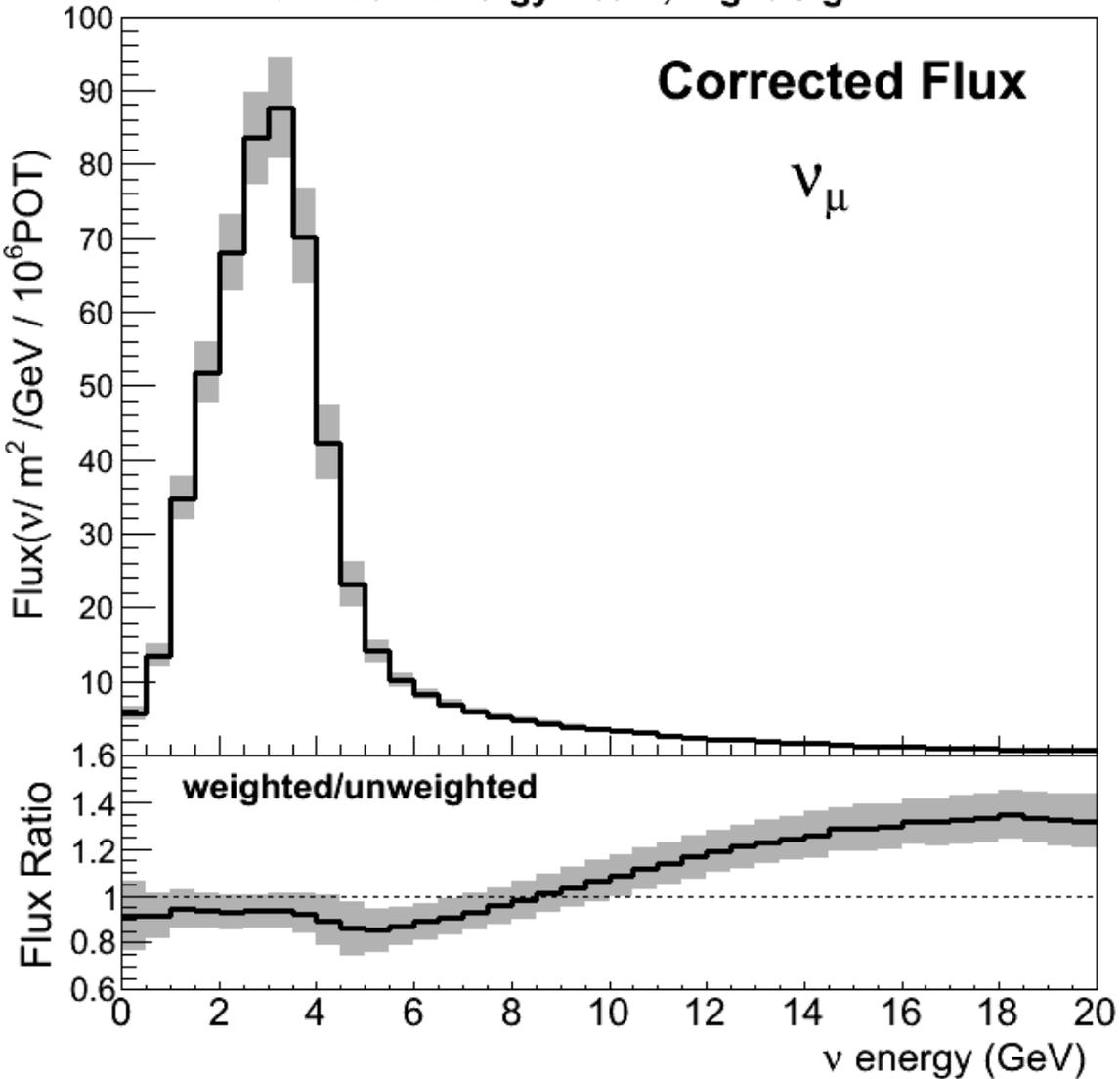


It is time for results

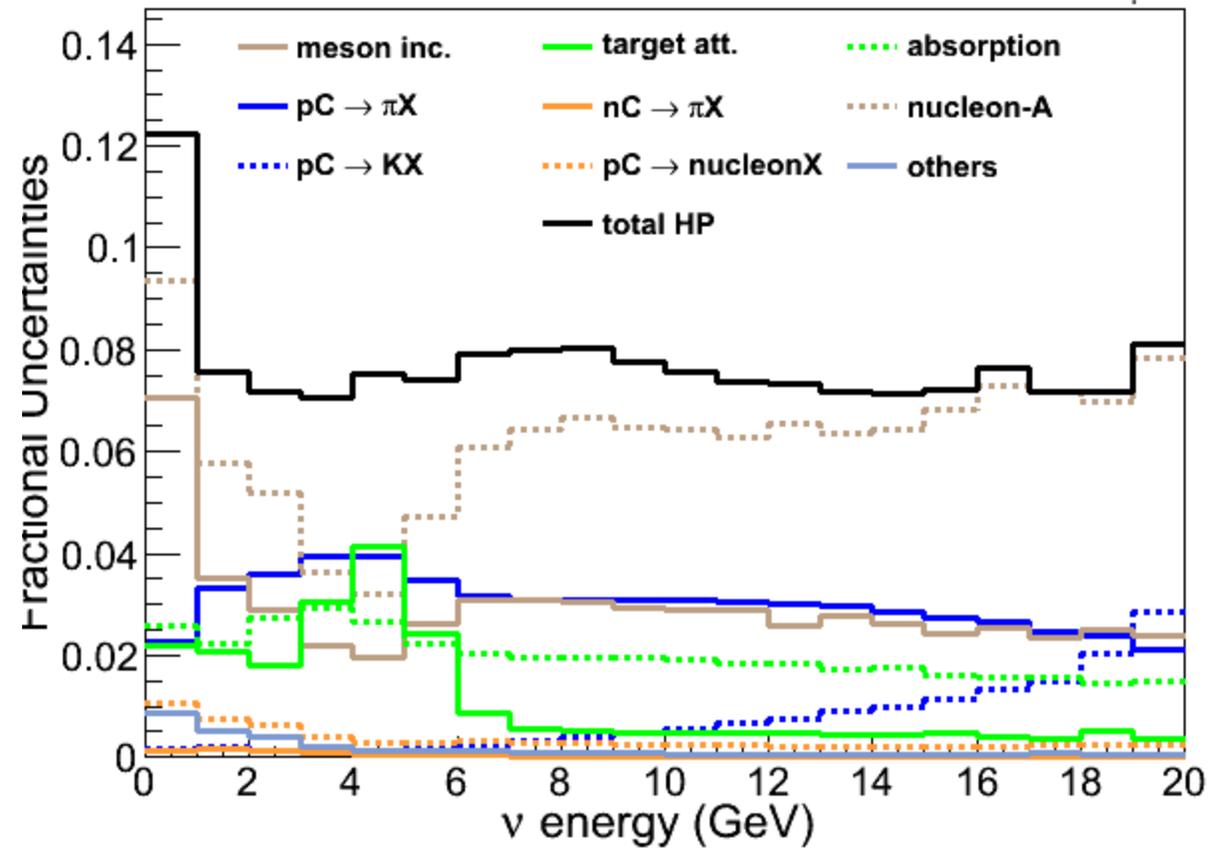
- Plots presented in this talk include all geometry improvements
- They include thin target hadron production data correction
- They do NOT include thick target hadron production correction
- The thick target data will be discussed at Joint Experimental-Theoretical Physics Seminar at Fermilab (aka Wine&Cheese) by L. Aliaga on Dec 17 (and in his PhD thesis!)

Muon Neutrino

NuMI Low Energy Beam, Right Sign

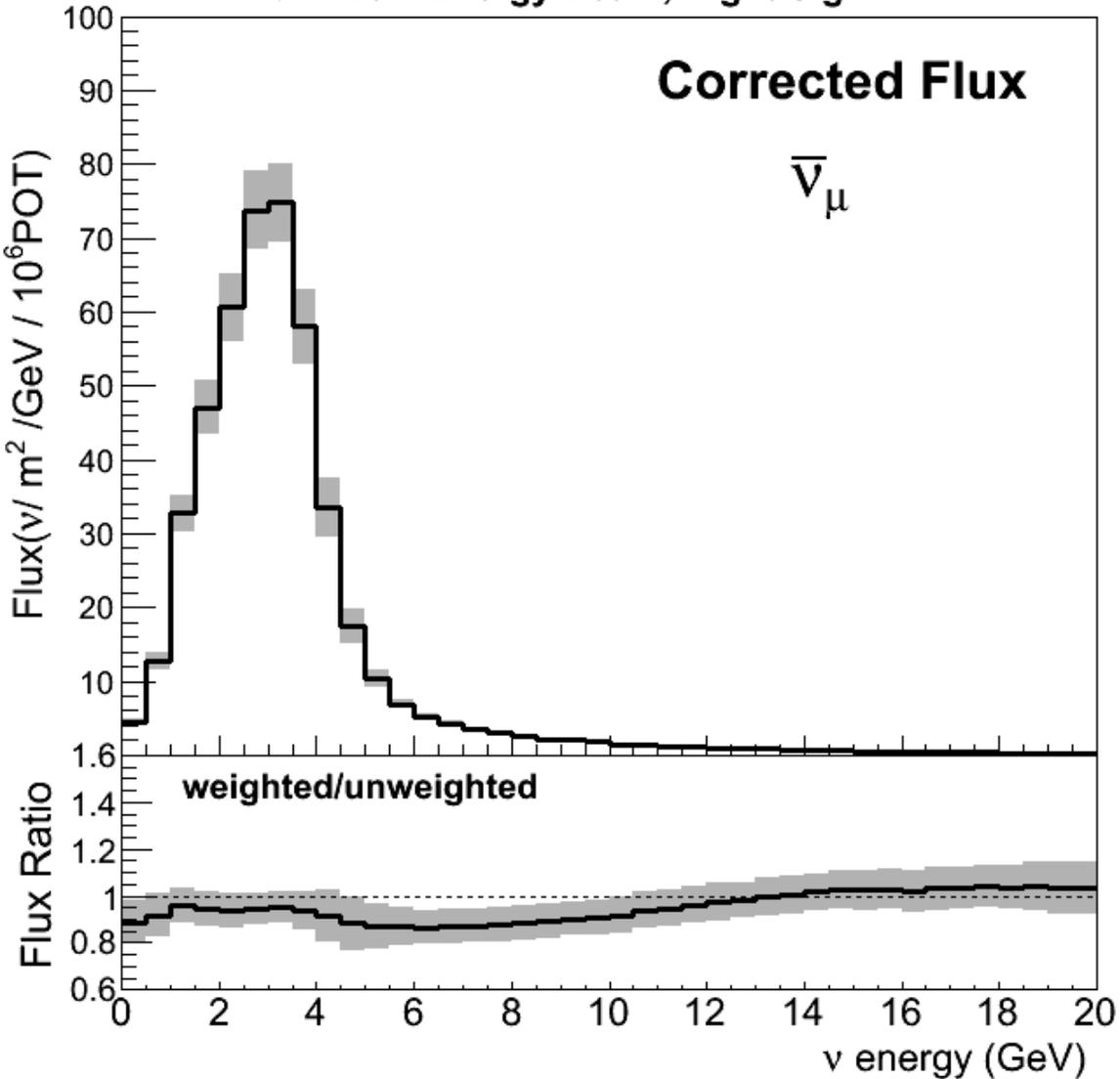


NuMI Low Energy Beam, HP Uncertainties, ν_μ

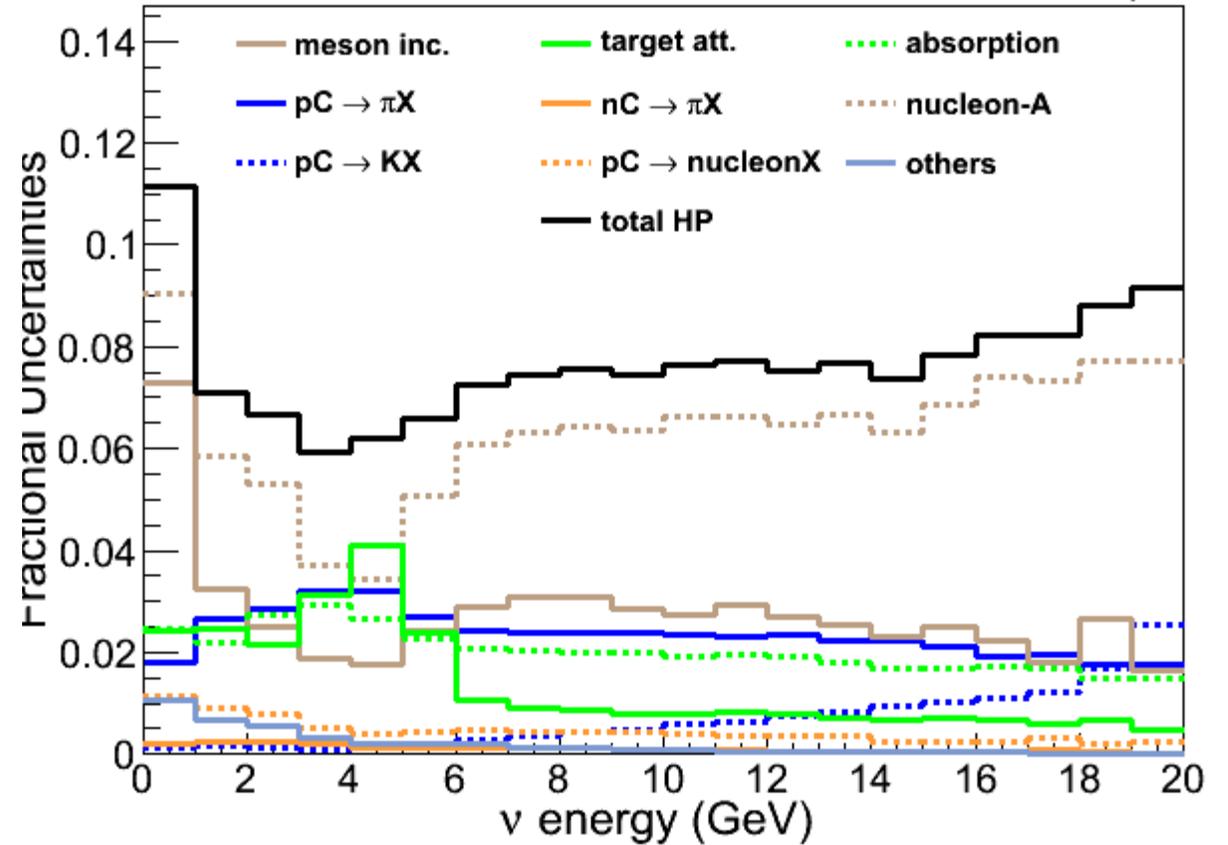


Muon Anti-Neutrino

NuMI Low Energy Beam, Right Sign

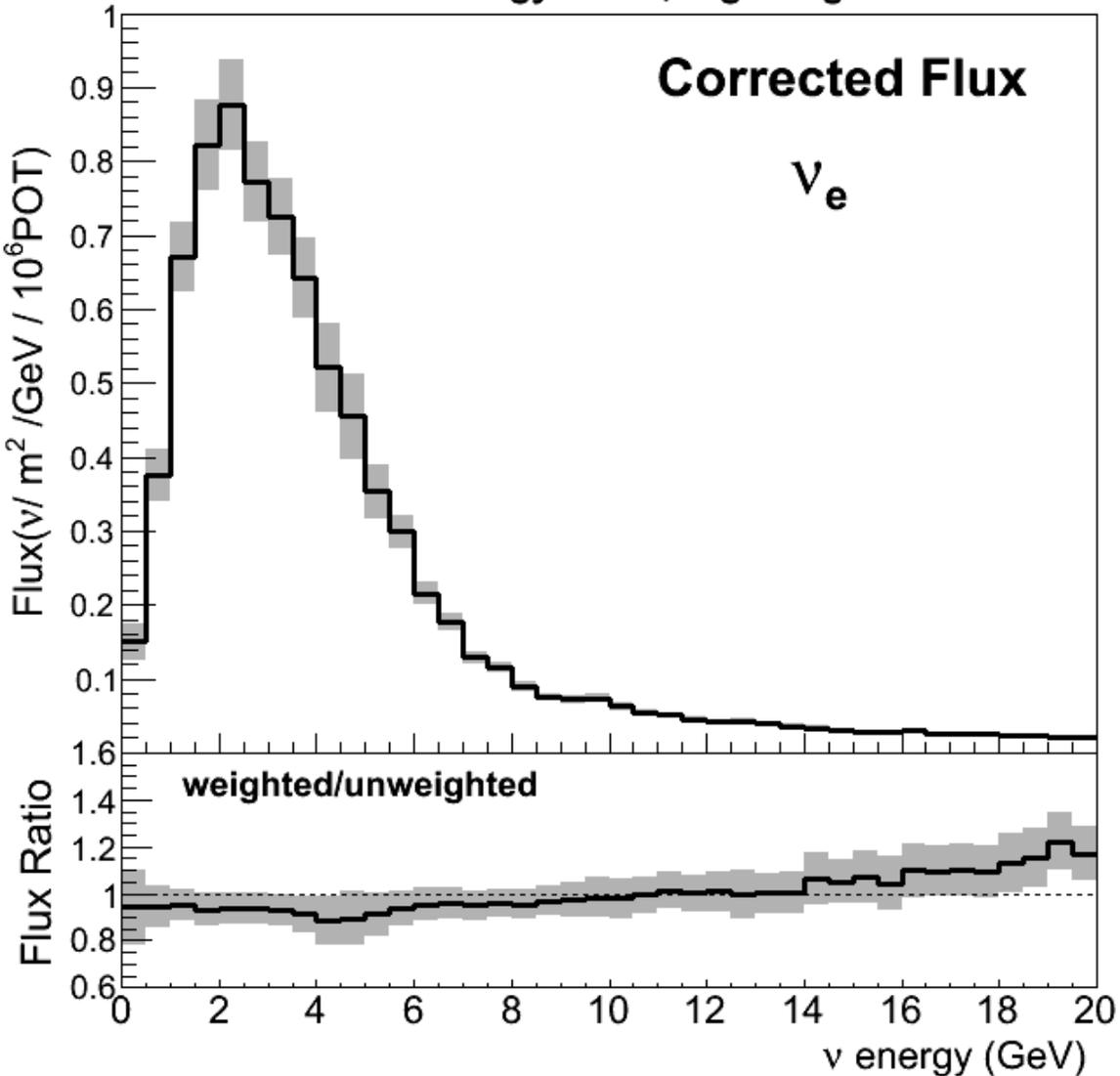


NuMI Low Energy Beam, HP Uncertainties, $\bar{\nu}_\mu$

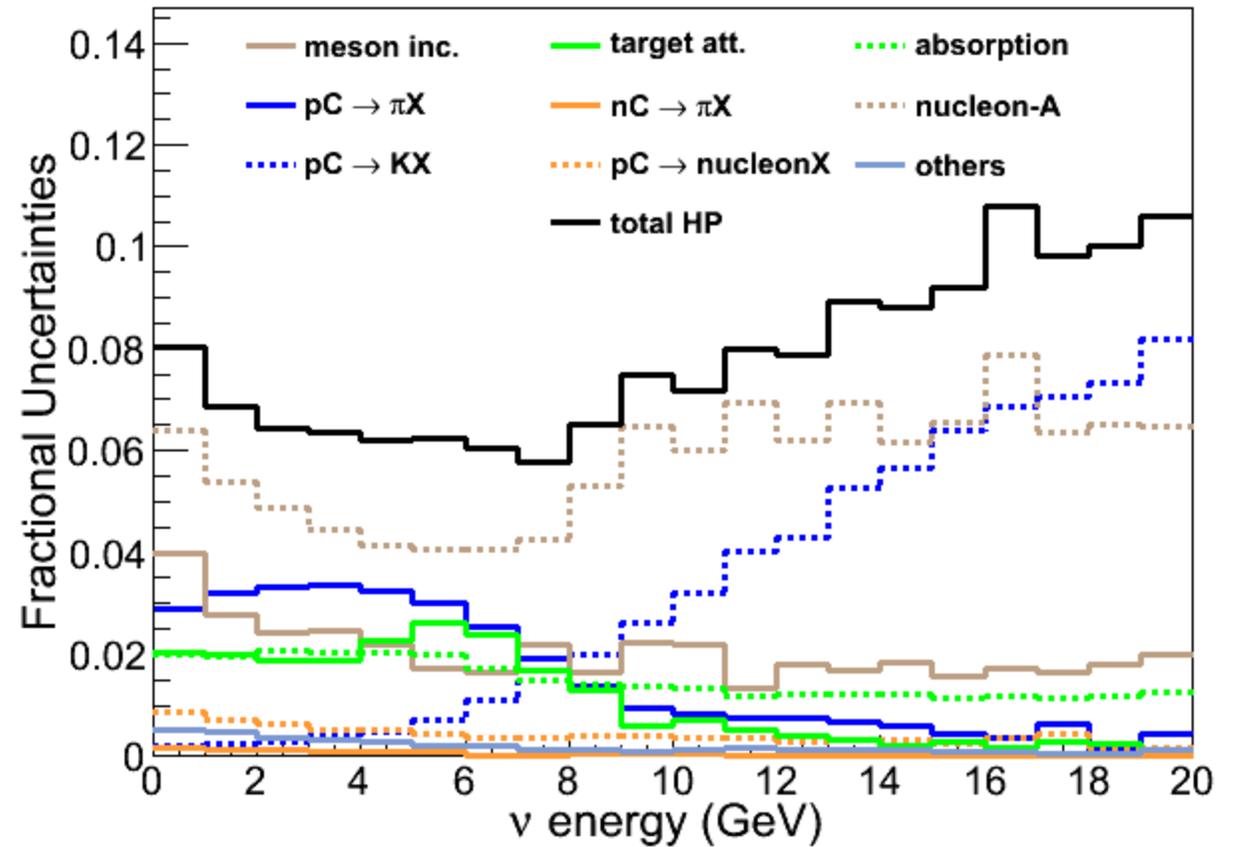


Electron Neutrino

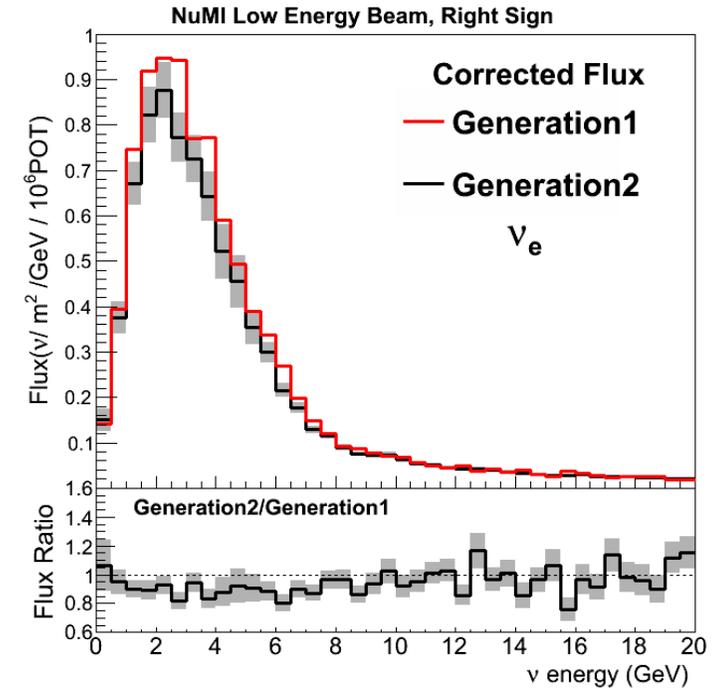
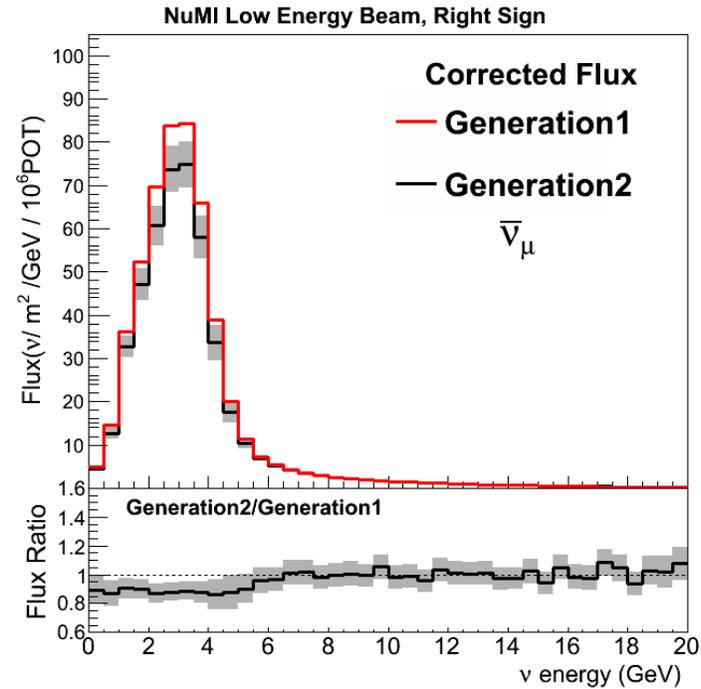
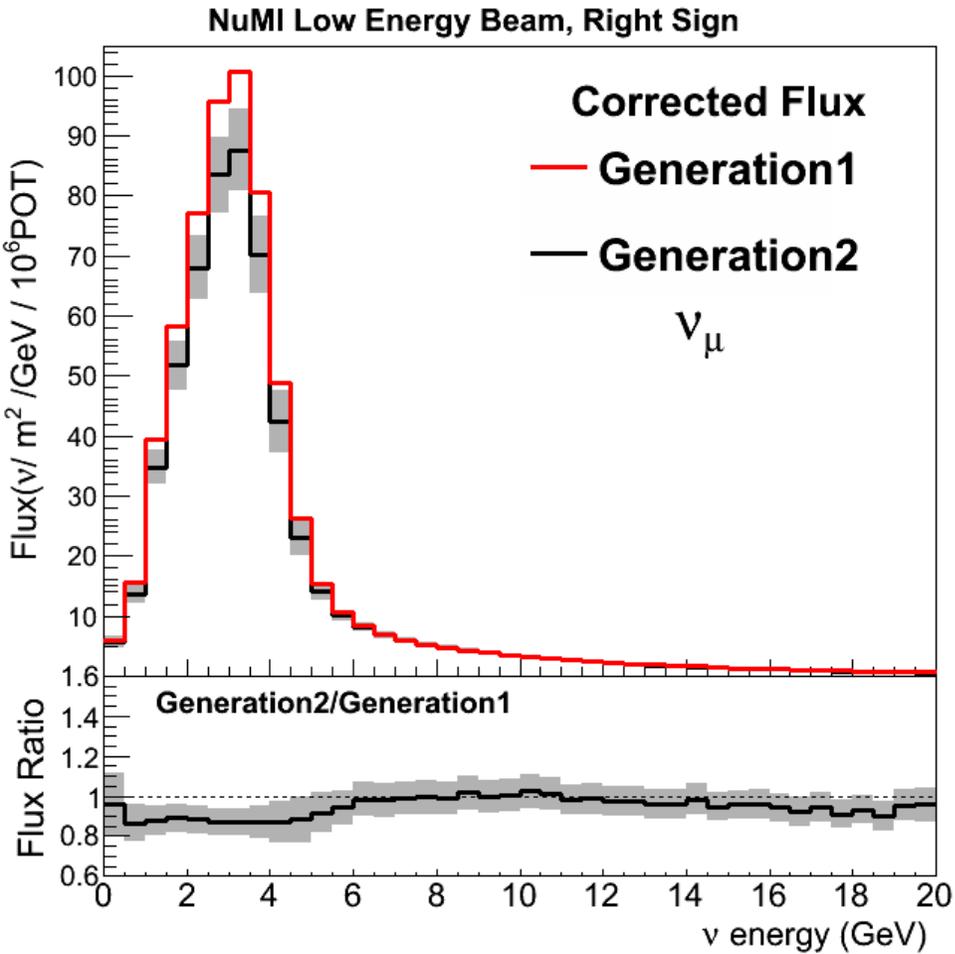
NuMI Low Energy Beam, Right Sign



NuMI Low Energy Beam, HP Uncertainties, ν_e

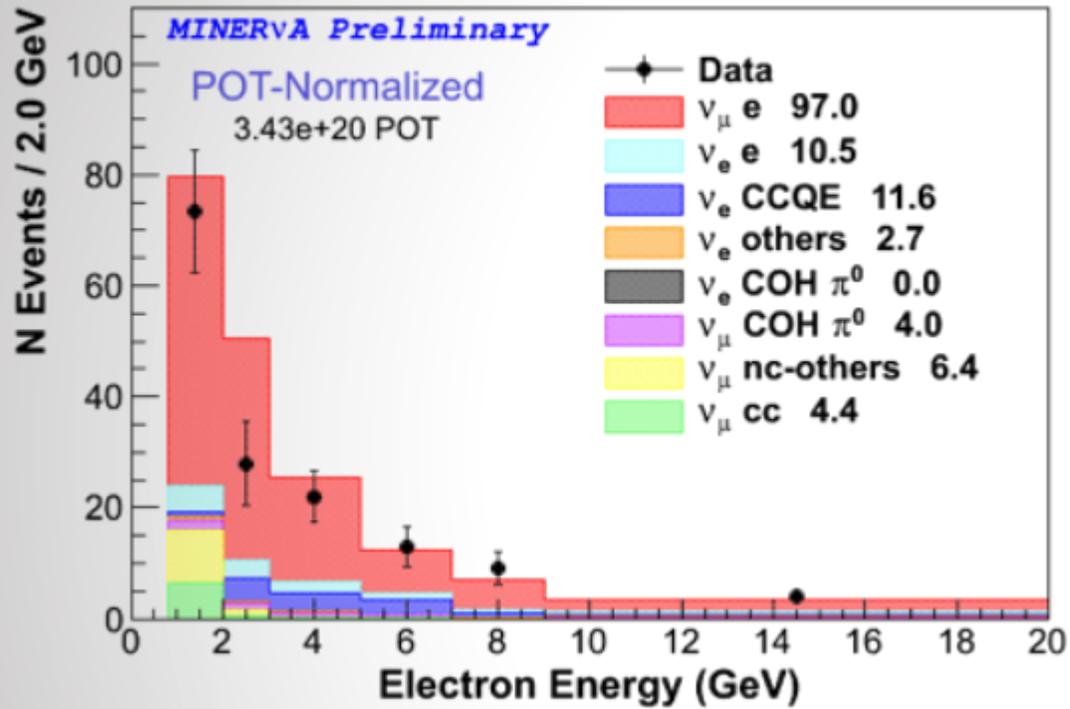


Flux comparisons



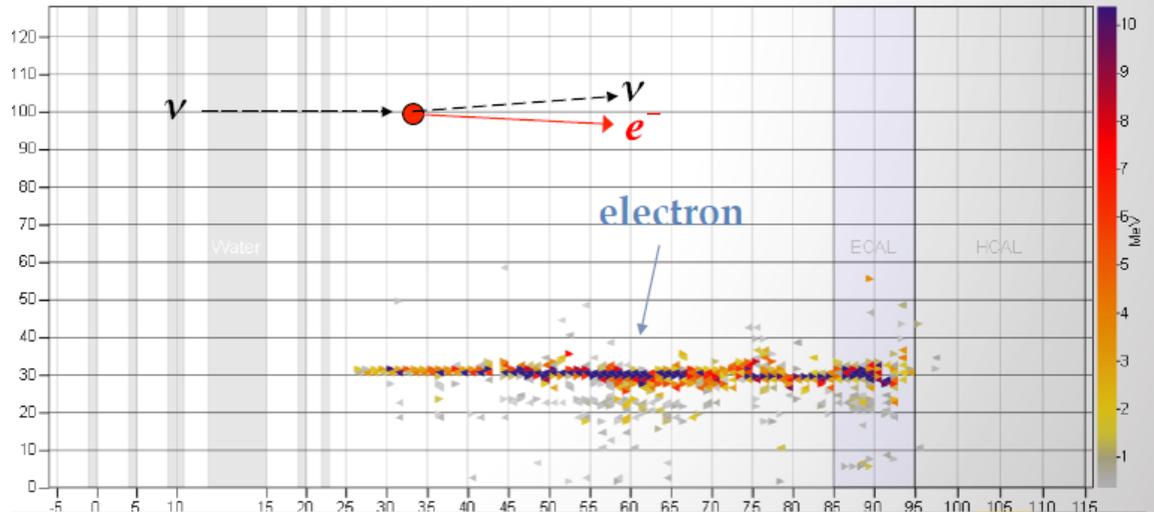
ν -e scattering constraint

First analysis which uses this constraint:
 ν_e CCQE, see J. Wolcott talk on Tue

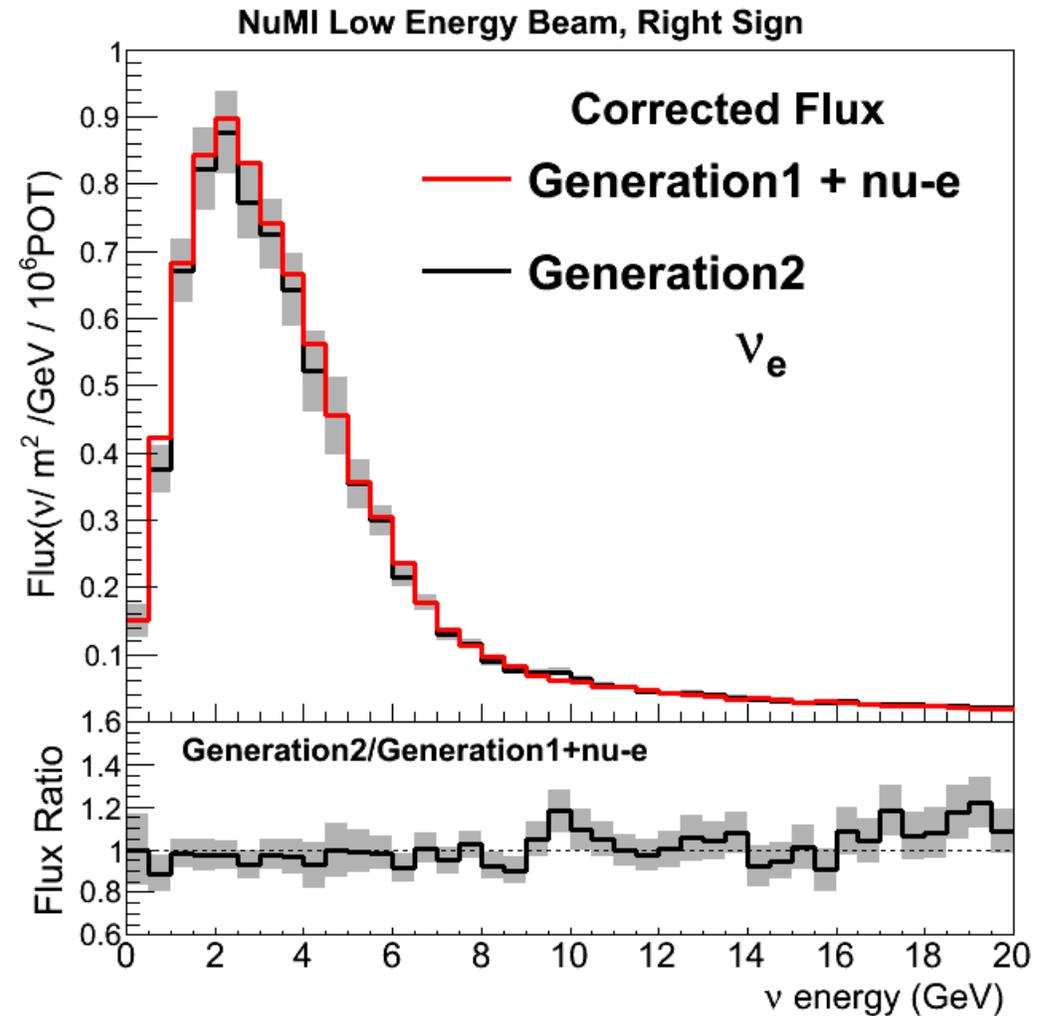
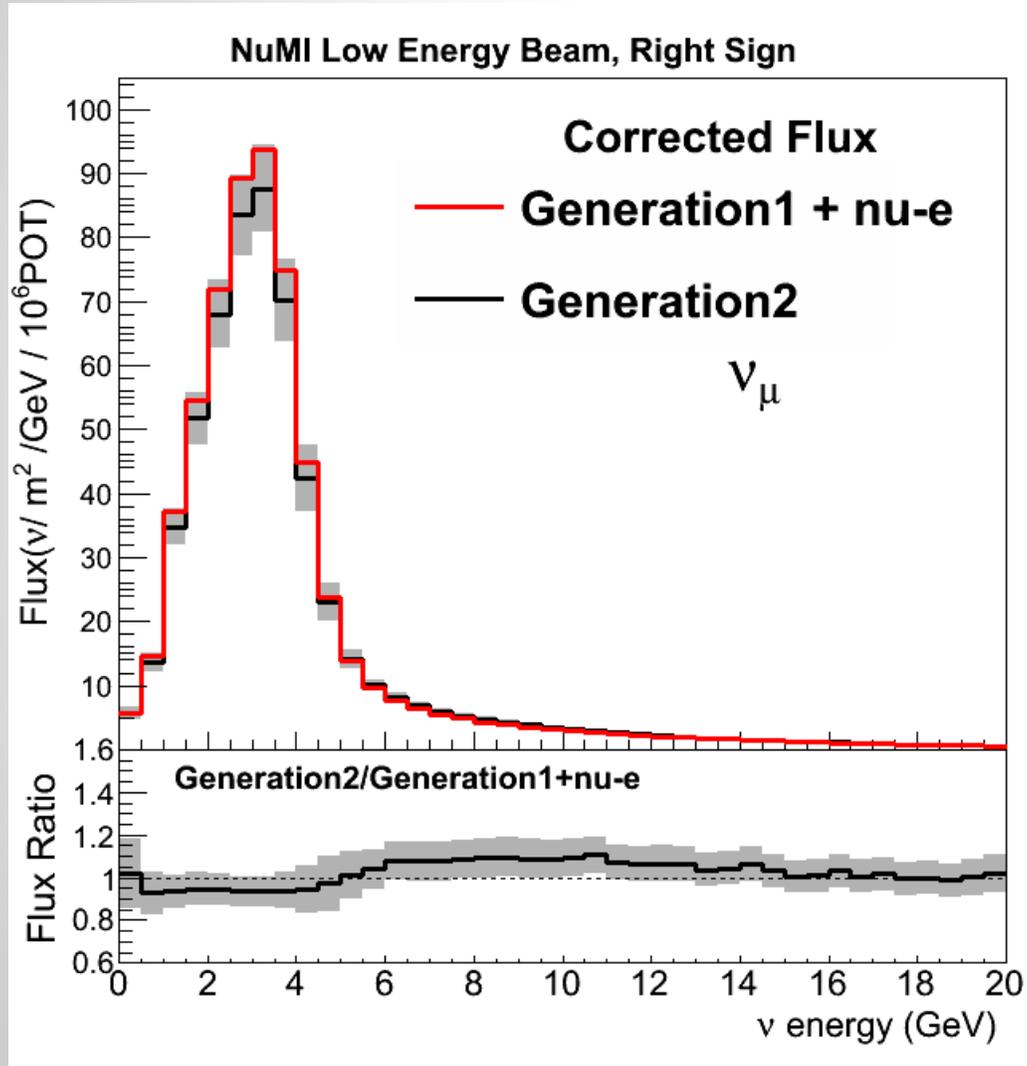


- Electroweak theory predicts precisely cross section for neutrino-electron scattering
- Experimental signature is a very forward single electron in the finale state

Predicted number of signal events is higher than indicated by data (weighting up universes that agree better with data)



ν -e scattering constraint



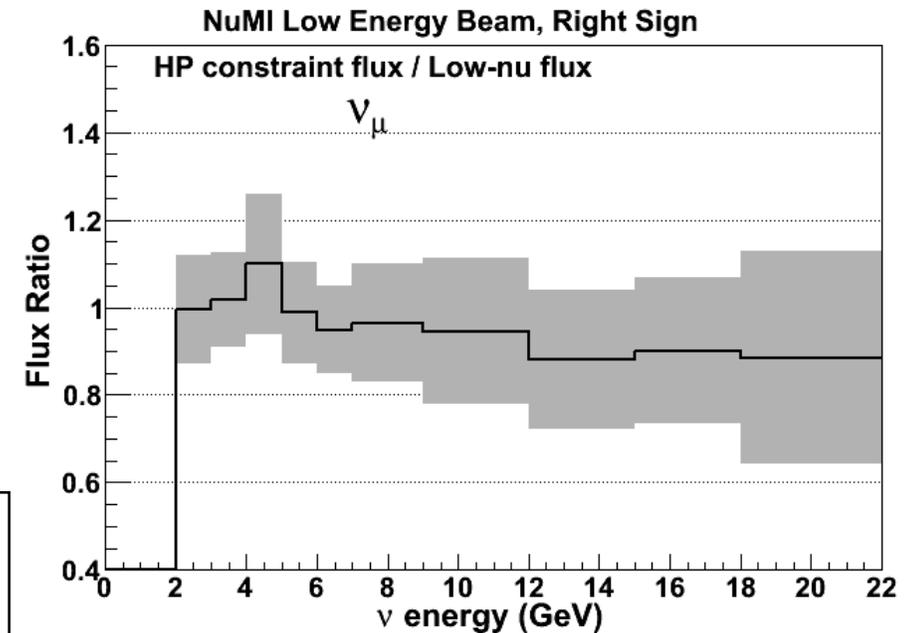
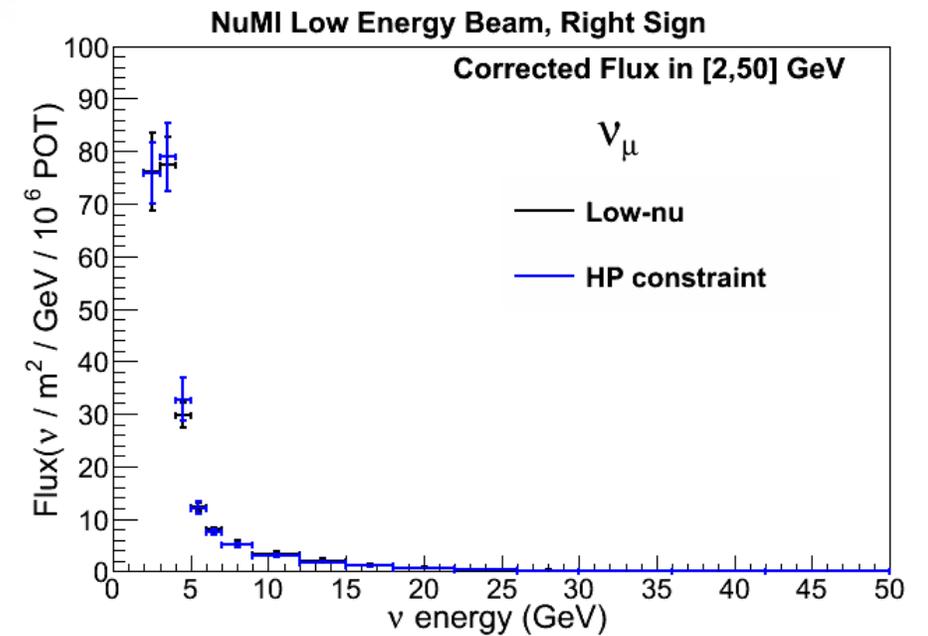
Low- ν constraint

- Differential cross section can be approximated by

$$\frac{d\sigma}{d\nu} = A \left(1 + \frac{B \nu}{A E} - \frac{C \nu^2}{A 2E^2} \right)$$

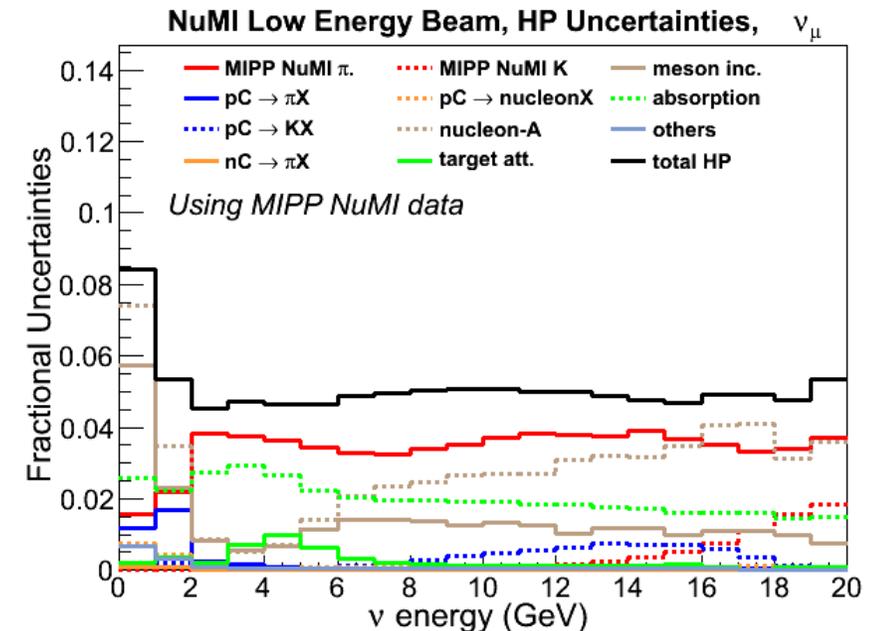
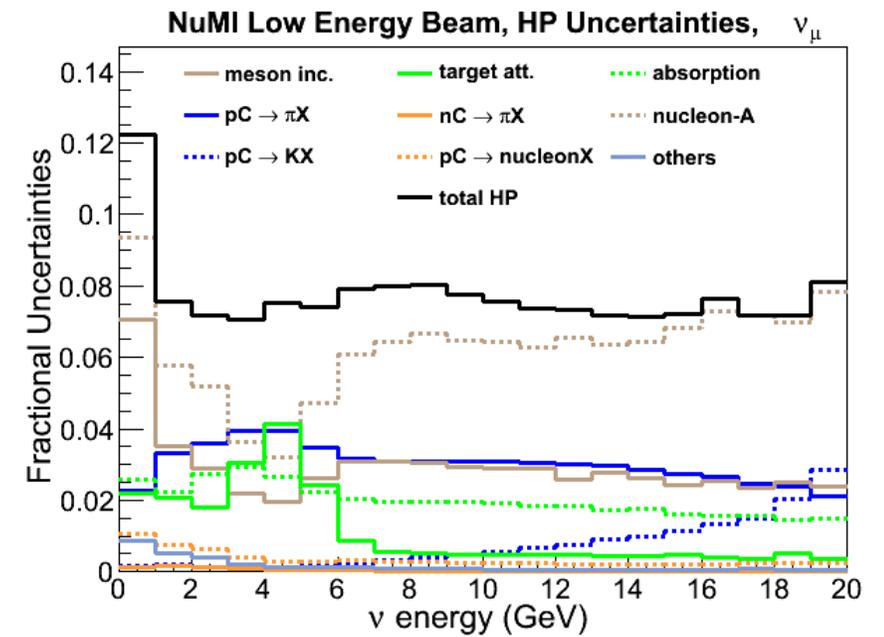
- It is constant for: $\frac{\nu}{E} \rightarrow 0$
- It can be used to constraint flux prediction
- At high energies normalization is fixed using high precision cross section measurements (like NOMAD)

More details on Jan 8
Wine&Cheese by J. Nelson
and J. DeVan PhD thesis



Summary

- MINERvA's flux prediction is significantly improved
- Geometry of NuMI beamline is updated
- Geant4-based simulations are corrected using external hadron production data
- ν -e scattering analysis is done
- Work on low- ν constraint is done as well as inclusion of MIPP thick target data
- Do not miss Wine&Cheese by L. Aliaga on Dec 17 and by J. Nelson on Jan 8



Backup slides

Flux playlists

Beam mode	Playlist	New Z position (cm)
LE FHC	1	9.50
nominal:	7	9.18
10 cm	9	10.40
	13	9.17
LE RHC	downstream	9.50
	5	8.85
	10	9.18
LE 0HC	6	9.18
pME FHC	2	99.57
nom:100cm	11	99.17
pME RHC	3	99.57
	12	99.17
pHE FHC	4	249.57
nom:250cm	8	250.09

Interaction map legend

pC \rightarrow π X: NA49 data scaled using FLUKA (checked by comparing with NA61)

pC \rightarrow KX: NA49 data and K/ π ratios from MIPP

pC \rightarrow NX: NA49 data for nucleon production

nC \rightarrow pX: assuming isoscalar symmetry

nucleon-A: for targets other than C

A-scaling is applied for kinematics covered by data

otherwise 40% assumed

meson-inc: meson as the projectile, educated guess (uncorrelated 40% uncertainty)

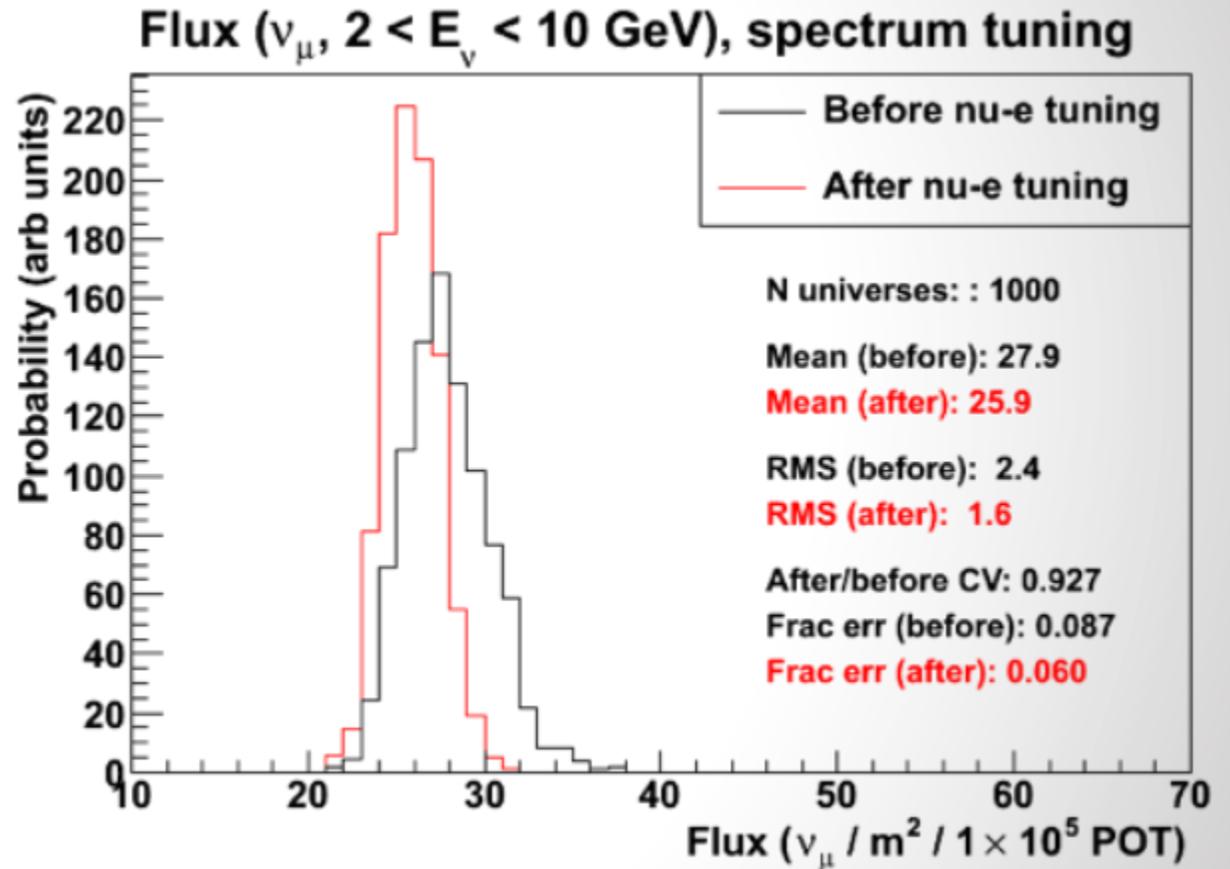
others: anything else, really rare, assuming 40% uncertainty

ν -e scattering constraint

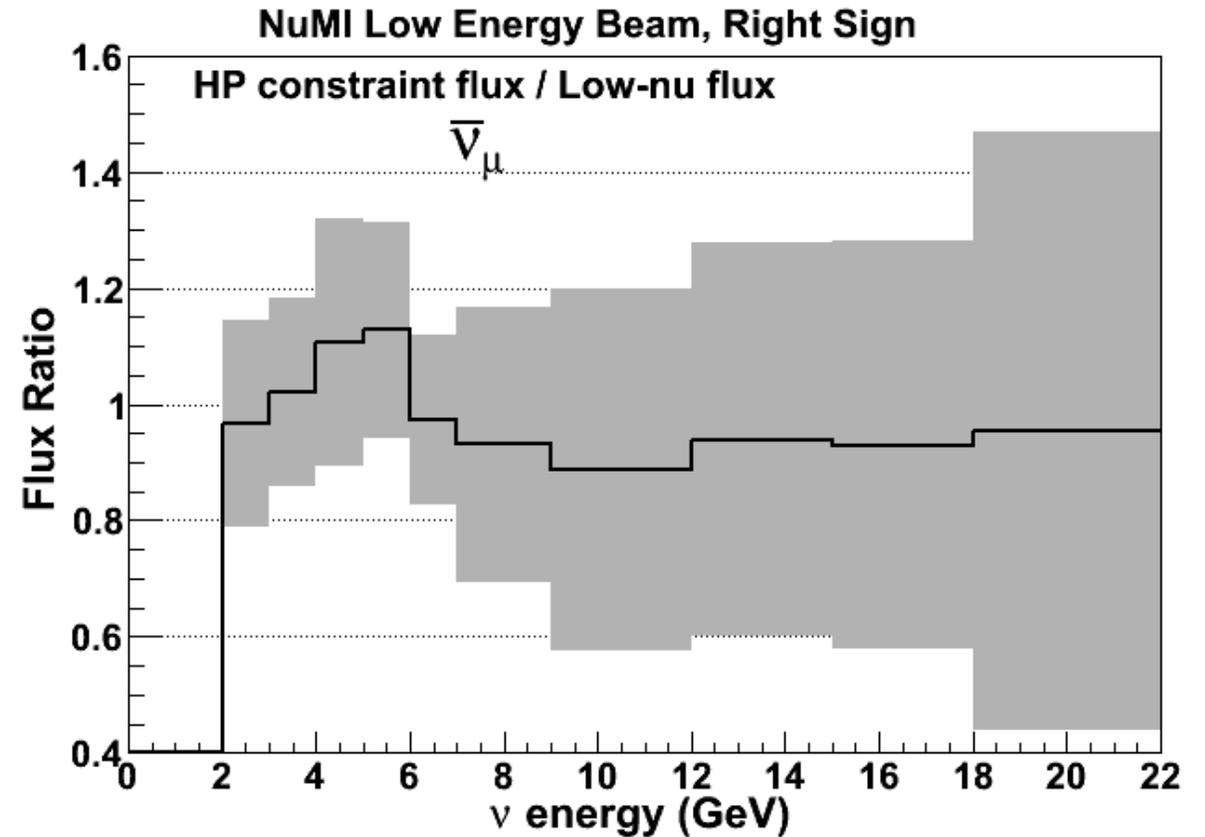
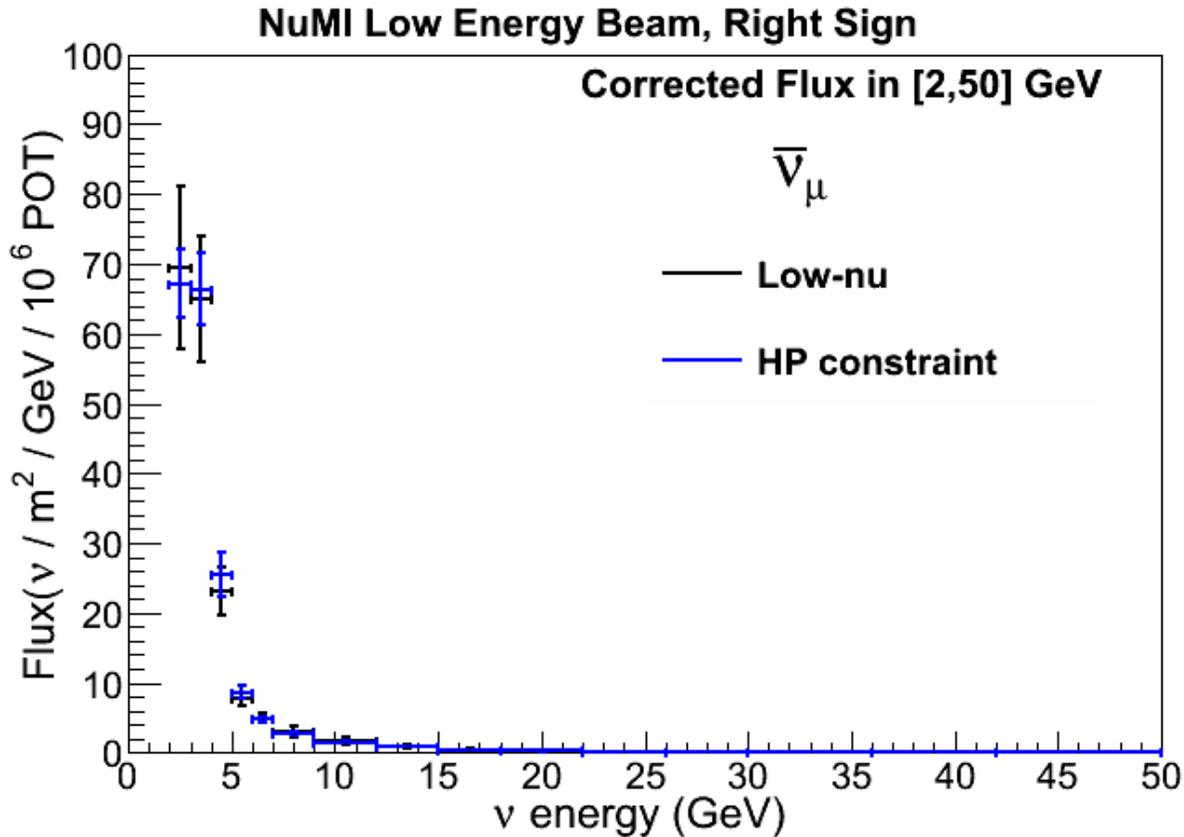
Probability of a universe:

$$f(x, \mu, \sigma) = \frac{1}{\sigma\sqrt{2\pi}} e^{-\frac{(x-\mu)^2}{2\sigma^2}}$$

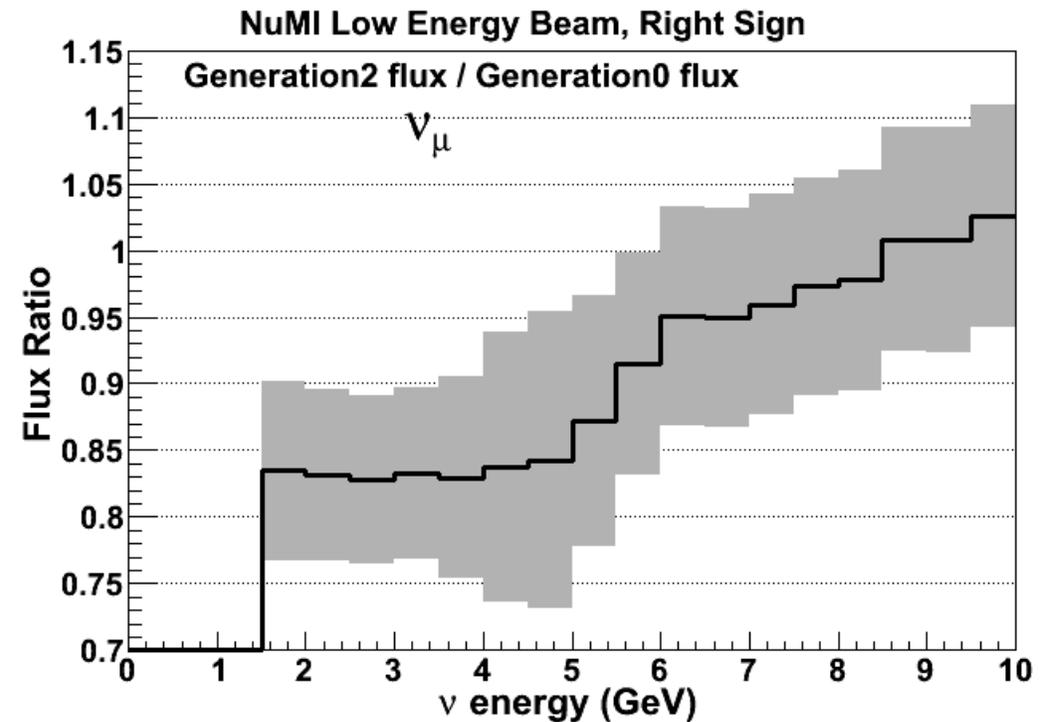
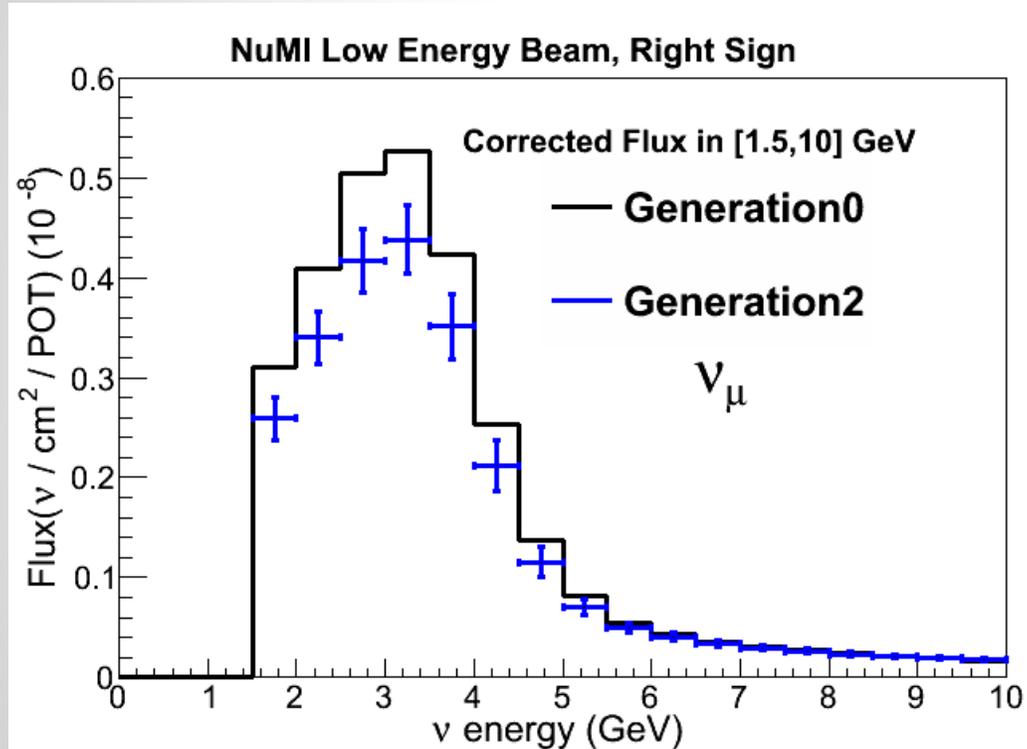
- μ : data events
- σ : error on data (fraction uncertainty = 13%)
- x : MC events in i -th multi-universe



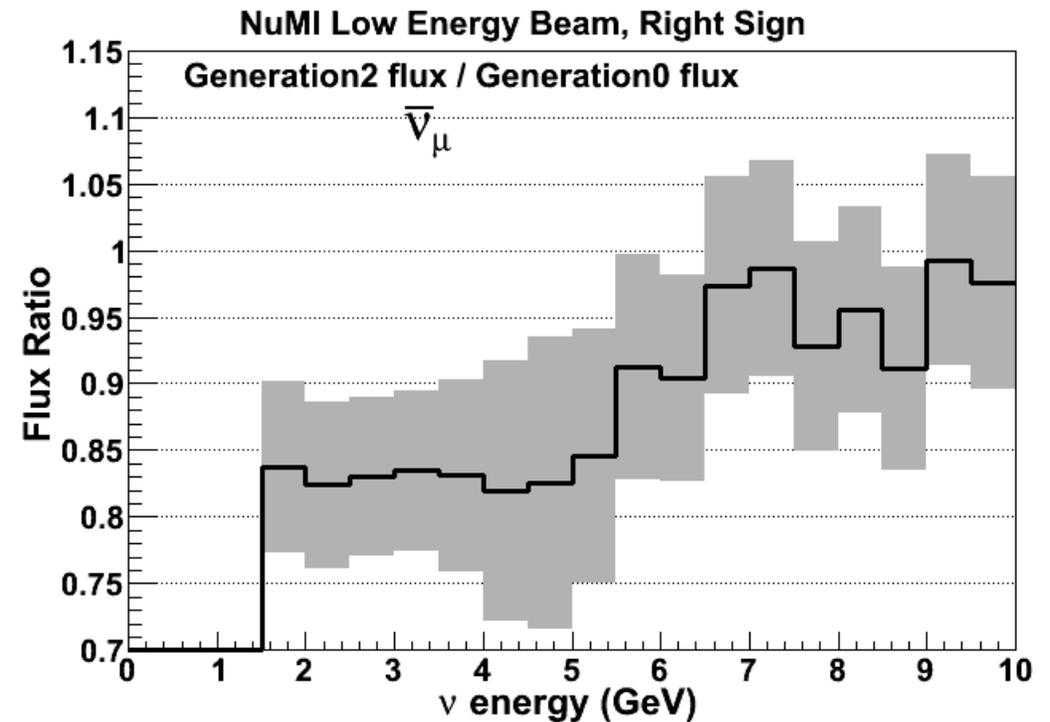
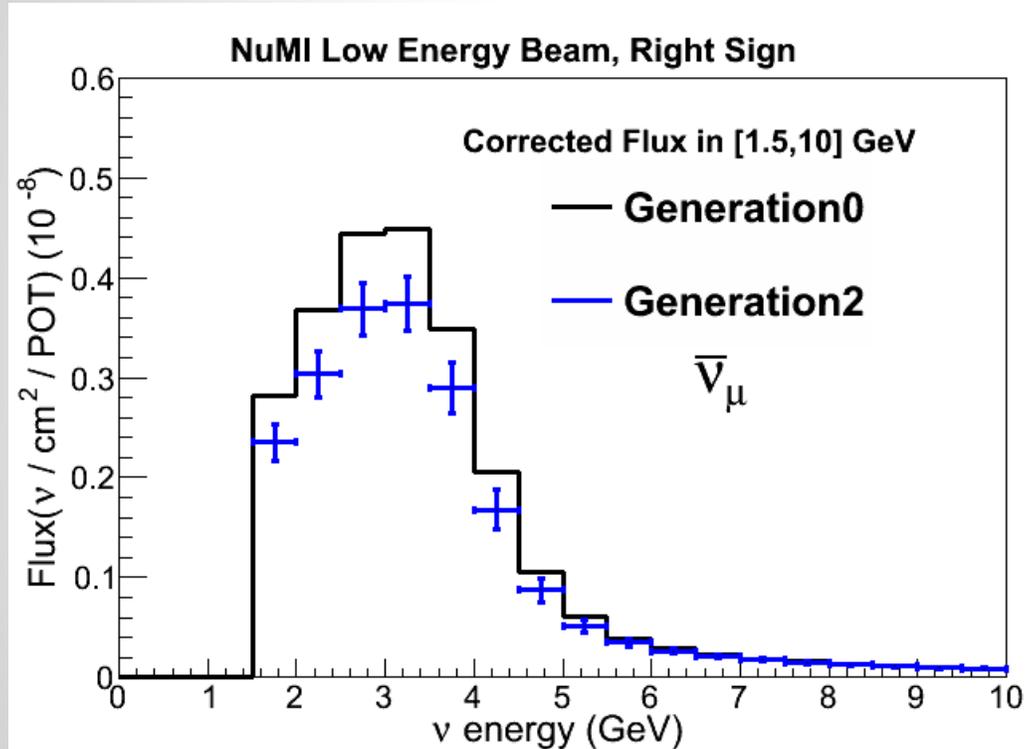
Low- ν constraint



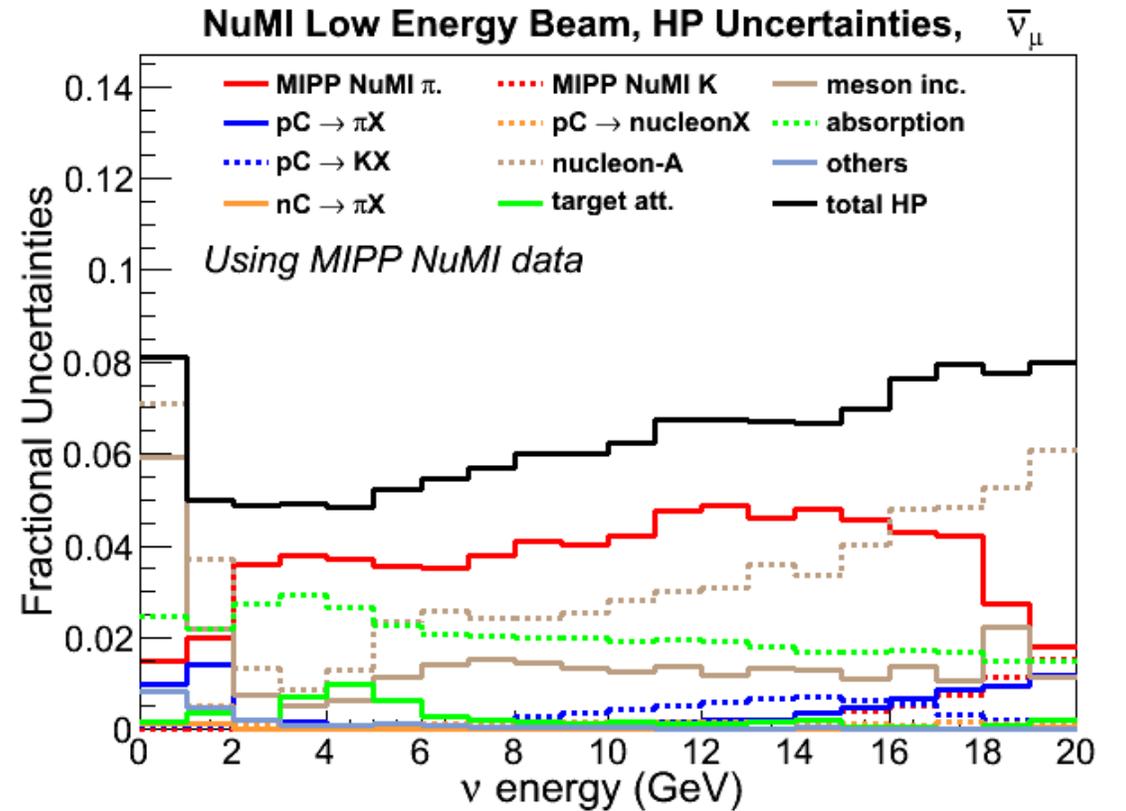
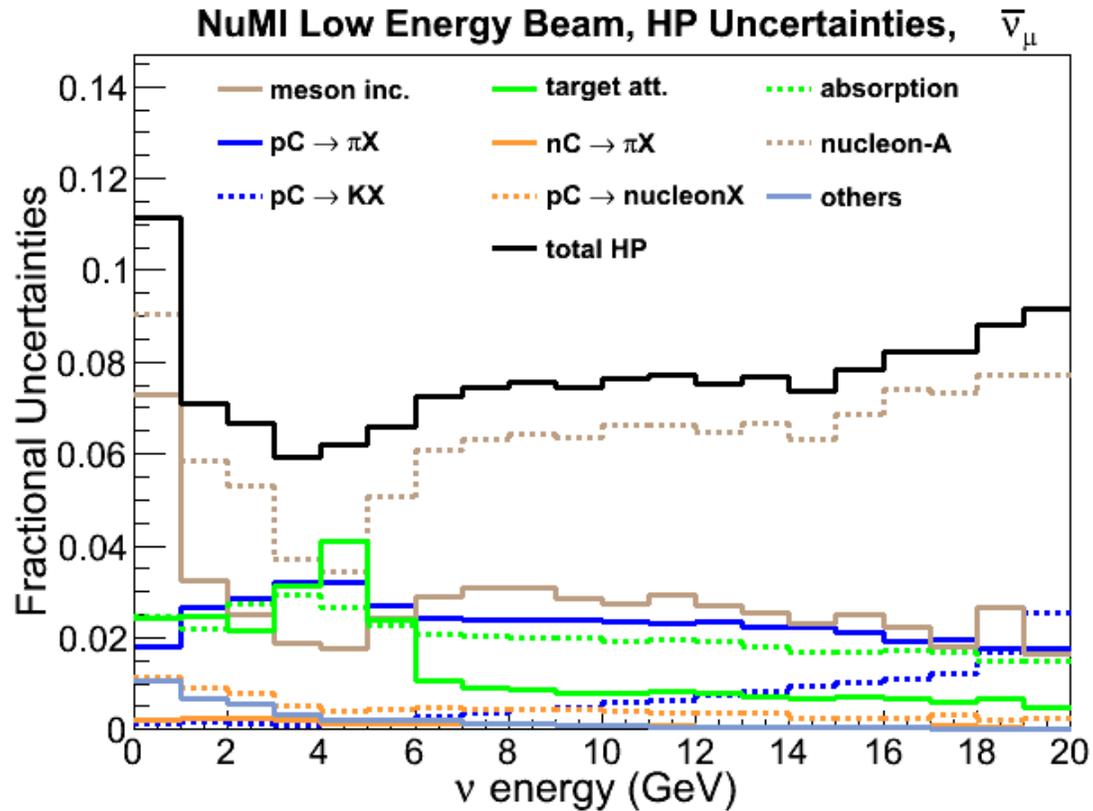
Gen 0 vs Gen 2 (muon neutrino)



Gen 0 vs Gen 2 (muon antineutrino)



MIPP thin data



MIPP thin data

