

# Frame invariant analysis of first order hydrodynamics as IR effective theory

Atsuhisa Ota

Chongqing U.

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- ① Motivation: the long-standing stability puzzle in first order hydrodynamics
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## Motivation: the stability puzzle

*First, the textbook problem: first-order theory looks unstable.*

# What is “first-order” hydrodynamics?

## Zeroth order: ideal fluid

$$T_{(0)}^{\mu\nu} = \rho u^\mu u^\nu + P \gamma^{\mu\nu},$$
$$J_{(0)}^\mu = n u^\mu.$$

- No derivatives of  $\beta, u^\mu, \mu$
- Reversible: no entropy production
- The relativistic Euler equations

## First order: dissipative fluid

$$T^{\mu\nu} = T_{(0)}^{\mu\nu} + T_{(1)}^{\mu\nu},$$
$$J^\mu = J_{(0)}^\mu + J_{(1)}^\mu.$$

- Exactly *one* gradient of  $\beta, u^\mu, \mu$  per term
- Irreversible: viscosity, heat/charge diffusion
- A relativistic Navier–Stokes

“First order” means first order in  $\nabla_\mu$ .

[Landau, Lifshitz (1987)]

Everyone wants to evolve a dissipative relativistic fluid as a closed system of PDEs built from conservation laws.

$$\nabla_\mu T^{\mu\nu} = 0, \quad \nabla_\mu J^\mu = 0.$$

The problem: naive first-order hydrodynamics is unstable; simulations fail.

# The classical diagnosis for first-order theory

**Step 1**– Plane-wave expansion of perturbations around equilibrium:

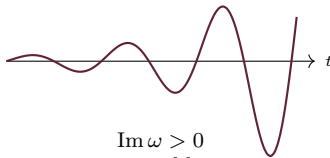
$$\delta\Phi \propto e^{-i\omega t + i\mathbf{k}\cdot\mathbf{x}}$$

**Step 2**– Solve the dispersion relation and find

$$\omega = \omega(k).$$



$\text{Im } \omega < 0$   
stable



$\text{Im } \omega > 0$   
unstable

[Hiscock, Lindblom (1985)]

# Two kinds of “frame”: (1) hydrodynamic frame

**Hydrodynamic frame** – the freedom in *redefining* the hydrodynamic variables:

$$\beta \rightarrow \beta + \delta\beta, \quad u^\mu \rightarrow u^\mu + \delta u^\mu, \quad \mu \rightarrow \mu + \delta\mu.$$

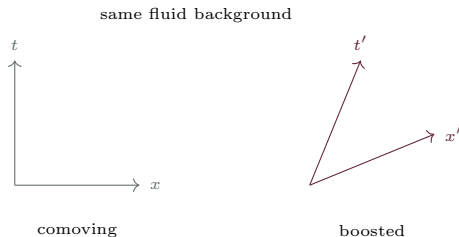
- Some first order coefficients can be absorbed in  $\delta\beta$ ,  $\delta u^\mu$  and  $\delta\mu$ .
- Landau frame: velocity defined by energy flow
- Eckart frame: velocity defined by particle flow

[Landau, Lifshitz (1987); Eckart (1940)]

## Two kinds of “frame”: (2) inertial frame

**Inertial frame** – which *coordinate* describes the same fluid background:

- Comoving frame: moves together with the fluid
- Non-comoving frame: boosted relative to the fluid





# The classical mode analysis: verdict

Only a special combination is stable:

*Landau frame + comoving inertial frame*

Shifting either frame then appears to produce unstable modes. [Hiscock, Lindblom (1985)]

The Landau frame was regarded as “a peculiar singular limit”.

*But a physical IR pole should not depend on how we choose to label the fluid nor coordinates.*



## Reframing hydrodynamics as an IR effective theory

*The fix: stop treating hydrodynamics as exact — treat it as an IR EFT.*

# Hydrodynamics is not a fundamental theory

Hydrodynamics is an **IR effective theory** describing the low-frequency, long-wavelength regime – it is not fundamental.

**Analogy: chiral perturbation theory.** Truncating the derivative expansion at finite order and then solving *exactly* can generate spurious UV poles. But the expansion is only trusted for

$$p \ll \Lambda_{\text{EFT}},$$

and UV poles produced this way are not treated as physical degrees of freedom.

[Weinberg (1979)]

# Is first-order hydrodynamics any different?

- A truncated first-order equation should *not* be treated as an exact PDE valid at all frequencies and wavenumbers
- Pathological poles appearing in the UV should not disqualify the theory as a valid IR effective theory
- Nonlinear response and loop corrections only require IR poles that are controlled order-by-order in the derivative expansion

**Question:** Can first-order hydrodynamics be treated consistently as an IR effective theory?

[AO (2026)]



## Effective theory for hydrodynamic modes

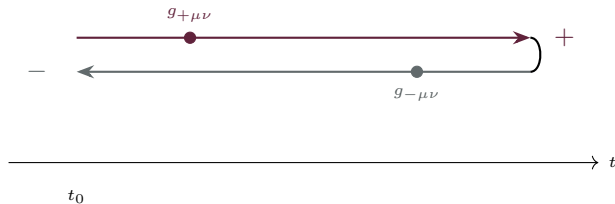
*Build the EFT action explicitly, with hydrodynamic fluctuations as NG modes.*

# Schwinger–Keldysh: doubling the metric

Hydrodynamics is realized by integrating fast modes ( $\chi$ ) on the closed-time-path:

$$e^{iI[g_{+\mu\nu}, g_{-\mu\nu}]} = \oint_{\mathcal{C}} \mathcal{D}\chi_+ \mathcal{D}\chi_- e^{iS[\chi_+, \chi_-; g_{+\mu\nu}, g_{-\mu\nu}]}$$

The closed time path doubles every external source:



We will develop  $I[g_{+\mu\nu}, g_{-\mu\nu}]$  as the effective action for the hydrodynamic modes.

The generating functional is invariant under  $\text{Diff}_+ \times \text{Diff}_-$ .

$$\begin{aligned} e^{iI[g_{+\mu\nu}, g_{-\mu\nu}]} &= \oint_{\mathcal{Q}} \mathcal{D}\chi_+ \mathcal{D}\chi_- e^{iS[\chi_+, \chi_-; g_{+\mu\nu}, g_{-\mu\nu}]} \\ &= \oint_{\mathcal{Q}} \mathcal{D}\chi'_+ \mathcal{D}\chi'_- e^{iS[\chi'_+, \chi'_-; g'_{+\mu\nu}, g'_{-\mu\nu}]} \\ &= e^{iI[g'_{+\mu\nu}, g'_{-\mu\nu}]} \end{aligned}$$

But, a realized fluid select a preferred rest frame:  $\text{Diff}_+ \times \text{Diff}_-$  is *fixed*.

**Hydro EFT:** Non linear realization of  $\text{Diff}_+ \times \text{Diff}_-$ .



# Symmetry structure

The Keldysh basis:

$$g_{\mu\nu} \equiv \frac{g_{+\mu\nu} + g_{-\mu\nu}}{2}, \quad g_{A\mu\nu} \equiv g_{+\mu\nu} - g_{-\mu\nu},$$

The symmetry structure:

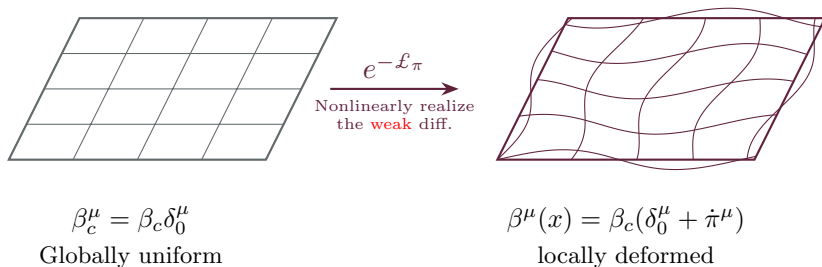
| Symmetry               | Generators                           | Acts on                                  |
|------------------------|--------------------------------------|------------------------------------------|
| <b>Strong</b>          | $\text{Diff}_+ \times \text{Diff}_-$ | $g_{+\mu\nu}, g_{-\mu\nu}$ independently |
| <b>Weak (physical)</b> | diagonal, $\xi_+ = \xi_- \equiv \xi$ | $g_{\mu\nu}$ and $g_{A\mu\nu}$           |
| <b>Noise</b>           | remaining part                       | $g_{A\mu\nu}$ only                       |

[Crossley, Glorioso, Liu (2017); Glorioso, Liu (2018)]

# Weak diffeomorphisms

A fluid emerges with a rest frame.

→ a local hydrodynamic configuration is generated by a diff. Stückelberg field  $\pi^\mu$



$\pi^\mu$  contains the NG fields associated with spontaneously broken boosts.

[Dubovsky, Hui, Nicolis, Son, (2011), Crossley, Glorioso, Liu (2017)]

Physical and noise coordinates:

$$x^\mu \equiv \frac{x_+^\mu - x_-^\mu}{2}, \quad x_A^\mu \equiv x_+^\mu - x_-^\mu$$

Under the noise diffeomorphism

$$\begin{aligned} T^{\mu\nu} &\rightarrow T^{\mu\nu}, \\ g_{A\mu\nu} &\rightarrow g_{A\mu\nu} + \nabla_\mu \xi_{A\nu} + \nabla_\nu \xi_{A\mu} \end{aligned}$$

Introduce a Stückelberg field  $\xi_{A\nu} \rightarrow -\pi_{A\nu}$  to nonlinearly realize the noise diff:

$$G_{A\mu\nu} = g_{A\mu\nu} - \nabla_\mu \pi_{A\nu} - \nabla_\nu \pi_{A\mu} \quad \Rightarrow \quad G_{A\mu\nu} \rightarrow G_{A\mu\nu}.$$

[Crossley, Glorioso, Liu (2017)]

The stress tensor couples to the noise metric in the SK path integral:

$$e^{iI[g,g_A]} \propto \exp\left[\frac{i}{2} \int d^4x \sqrt{-g} T^{\mu\nu} g_{A\mu\nu}\right].$$

Replacing  $g_A \rightarrow G_A$ , we have

$$I = \frac{1}{2} \int d^4x \sqrt{-g} \pi_{A\nu} \nabla_\mu T^{\mu\nu} + \dots$$

An “equation of motion” in  $\pi_{A\mu}$  is the conservation law:

$$\frac{\delta I[g, G_A]}{\delta \pi_{A\nu}} = 0 \quad \implies \quad \nabla_\mu T^{\mu\nu} = 0.$$

[Crossley, Glorioso, Liu (2017)]

# Schwinger–Keldysh effective action

The effective action for the hydrodynamic modes:

$$I = \int d^4x \sqrt{-\det g} \sum_{n=1} \mathcal{L}_n,$$

with ( $C_{A\mu}$  for the U(1) source)

$$\mathcal{L}_1 = \frac{1}{2} T^{\mu\nu} G_{A\mu\nu} + J^\mu C_{A\mu},$$

$$\mathcal{L}_2 = \frac{i}{4} W^{\mu\nu\rho\sigma} G_{A\mu\nu} G_{A\rho\sigma} + i X^{\mu\nu\rho} G_{A\mu\nu} C_{A\rho} + i Y^{\mu\rho} C_{A\mu} C_{A\rho}.$$

Constraints inherited from the UV generating functional:

- Conjugation condition
- Unitarity
- Local KMS symmetry

# KMS and unitarity fix the structure

| Condition               | Implies                           |
|-------------------------|-----------------------------------|
| <b>Unitarity + KMS</b>  | local second law                  |
| <b>KMS</b> at 0th order | thermodynamic relations           |
| <b>KMS</b> at 1st order | fluctuation–dissipation relations |

3 thermodynamic + 18 dissipative + 10 noise coefficients

$\Downarrow$  KMS + unitarity

**12** independent parameters

[Crossley, Glorioso, Liu (2017)]

# The general first-order constitutive relations

Most general U(1) first order hydrodynamics:

$$T^{\mu\nu} = (\rho + \rho^{(1)})u^\mu u^\nu + (P + P^{(1)})\gamma^{\mu\nu} + q^\sigma u^{(\mu}\gamma^{\nu)}_\sigma - \Sigma^{\mu\nu},$$
$$J^\mu = (n + n^{(1)})u^\mu + \gamma^{\mu\nu}j_\nu.$$

The first-order pieces  $\rho^{(1)}, P^{(1)}, n^{(1)}, q_\mu, j_\mu, \Sigma^{\mu\nu}$  split into a scalar sector and a vector/tensor sector – shown next.

[Kovtun (2019)]

# Structure of the derivative expansion: scalar sector

Possible building blocks:

$$u^\rho \nabla_\rho \beta, \quad u^\rho \nabla_\rho (\beta \mu), \quad \nabla_\rho u^\rho.$$

The first order scalars:

$$\begin{aligned}\rho^{(1)} &= \beta^{-1} \epsilon u^\rho \nabla_\rho \beta - \lambda_1 \nabla_\rho u^\rho - \nu_4 u^\rho \nabla_\rho (\beta \mu), \\ P^{(1)} &= \beta^{-1} \lambda_2 u^\rho \nabla_\rho \beta - \zeta \nabla_\rho u^\rho - \nu_5 u^\rho \nabla_\rho (\beta \mu), \\ n^{(1)} &= \nu_1 u^\rho \nabla_\rho \beta - \beta \nu_2 \nabla_\rho u^\rho - \beta \nu_3 u^\rho \nabla_\rho (\beta \mu).\end{aligned}$$

$\epsilon, \lambda_1, \lambda_2, \zeta, \nu_1, \nu_2, \nu_3, \nu_4, \nu_5$ : scalar transport coefficients, functions of  $\beta, \mu$  [Kovtun (2019)]



# Structure of the derivative expansion: vector & tensor sector

The heat flux  $q_\mu$ , charge flux  $j_\mu$ , and shear stress  $\Sigma^{\mu\nu}$  follow the same logic:

$$\begin{aligned}q_\mu &= \beta^{-1} \kappa_1 \nabla_\mu \beta - \kappa_2 u^\rho \nabla_\rho u_\mu - \tau_3 \nabla_\mu (\beta \mu) - \beta \tau_4 u^\rho F_{\rho\mu}, \\j_\mu &= \tau_1 \nabla_\mu \beta - \beta \tau_2 u^\rho \nabla_\rho u_\mu - \beta \chi_1 \nabla_\mu (\beta \mu) - \beta^2 \chi_2 u^\rho F_{\rho\mu}, \\\Sigma^{\mu\nu} &= \eta \gamma^{\mu\alpha} \gamma^{\nu\beta} \left( \nabla_\alpha u_\beta + \nabla_\beta u_\alpha - \frac{2}{3} \gamma_{\alpha\beta} \nabla_\gamma u^\gamma \right).\end{aligned}$$

- $\kappa_1, \kappa_2, \tau_1, \dots, \tau_4, \chi_1, \chi_2$ : vector transport coefficients.
- $\eta$ : shear viscosity.

[Kovtun (2019)]

# Hydrodynamic-frame invariants: energy sector

Physical predictions must be built from **frame-invariant combinations**, e.g.

$$\tilde{\lambda} = \lambda_2 - \left( \frac{\partial P}{\partial \rho} \right)_n \epsilon - \left( \frac{\partial P}{\partial n} \right)_\rho \nu_1,$$

$$\tilde{\zeta} = \zeta - \left( \frac{\partial P}{\partial \rho} \right)_n \lambda_1 - \left( \frac{\partial P}{\partial n} \right)_\rho \nu_2,$$

$$\tilde{\nu} = \nu_5 - \left( \frac{\partial P}{\partial \rho} \right)_n \nu_4 - \left( \frac{\partial P}{\partial n} \right)_\rho \nu_3,$$

$$\tilde{\tau}_1 = \tau_1 - \frac{n}{\beta(\rho + P)} \kappa_1,$$

$$\tilde{\tau}_2 = \tau_2 - \frac{n}{\beta(\rho + P)} \kappa_2,$$

$$\tilde{\chi}_1 = \chi_1 - \frac{n}{\beta(\rho + P)} \tau_3.$$

KMS and unitarity of the UV theory imply

$$\tilde{\tau}_1 = \tilde{\tau}_2, \quad \eta \geq 0, \quad \tilde{\chi}_1 - \frac{n}{\beta(\rho + P)} \tilde{\tau}_2 \geq 0, \quad \tilde{\zeta} + \left( \frac{\partial P}{\partial \rho} \right)_n \tilde{\lambda} + \left( \frac{\partial P}{\partial n} \right)_\rho \tilde{\nu} \geq 0.$$

These are exactly the ingredients that will guarantee stability later.

[Crossley, Glorioso, Liu (2017); AO (2026)]



## Mode analysis: a frame-invariant dispersion relation

*Solve the mode equations and show the dispersion relation is frame invariant.*

# From the effective action to a dispersion relation

Two steps turn the effective action into the mode equation.

**Step 1 – constraints.** Solve

$$\frac{\delta I}{\delta \pi_{A0}} = \frac{\delta I}{\delta \theta_A} = \frac{\delta I}{\delta \pi_0} = \frac{\delta I}{\delta \theta} = 0$$

to eliminate the non-dynamical fields  $\pi_0$ ,  $\pi_{A0}$ ,  $\theta$ ,  $\theta_A$  in favor of  $\pi_i$  and  $\pi_{Ai}$ .

**Step 2 – equation of motion.** Substituting back and varying with respect to the remaining field,

$$\frac{\delta I}{\delta \pi_{Ai}} = 0,$$

gives the dispersion relations.

[AO (2026)]

# Eliminating redundant modes

A formally sixth-order dispersion relation:

$$\sum_{n=0}^6 d_n \omega^n = 0.$$

- Solving this exactly reproduces the known gapped, unstable mode with  $\text{Im } \omega \propto \kappa_2^{-1}$
- But this over-interprets a truncated first-order theory

We must instead solve *perturbatively* in the derivative-counting parameter  $\varepsilon$ :

$$\omega = \omega_0 + \varepsilon \omega_1 + O(\varepsilon^2).$$

# Leading order: the sound speed

The leading-order solution is fixed purely by zeroth-order thermodynamics:

$$\omega_0 = \pm c_{s/n} k, \quad c_{s/n}^2 \equiv \left( \frac{\partial P}{\partial \rho} \right)_{s/n}.$$

- The adiabatic sound speed is completely insensitive to any first-order (dissipative) coefficient
- Zeroth-order stability requires  $c_{s/n}^2 \geq 0$

[AO (2026)]

# First-order correction: a messy start...

A direct evaluation from the general constitutive relations at first order looks hopeless – a long expression in many frame-dependent coefficients:

$$\omega_1 = -ik^2 F(\epsilon, \lambda_{1,2}, \zeta, \nu_{1..5}, \kappa_{1,2}, \tau_{1..4}, \chi_{1,2}, \eta).$$

[AO (2026)]



...but it collapses to frame invariants

Remarkably, everything reorganizes into

$$\omega_1 = -i \Gamma_L k^2,$$

with

$$\Gamma_L = \frac{1}{2(\rho + P)} \left[ \frac{4}{3} \underbrace{\eta}_{\geq 0} + \underbrace{\zeta + \left( \frac{\partial P}{\partial \rho} \right)_n \tilde{\lambda} + \left( \frac{\partial P}{\partial n} \right)_\rho \tilde{\nu}}_{\geq 0} \right. \\ \left. + f^2 \left( \left( \frac{\partial P}{\partial \rho} \right)_{s/n} \underbrace{\left( \tilde{\chi}_1 - \frac{n}{\beta(\rho + P)} \tilde{\tau}_2 \right)}_{\geq 0} + \beta \left( \frac{\partial P}{\partial n} \right)_\rho \underbrace{(\tilde{\tau}_1 - \tilde{\tau}_2)}_{=0} \right) \right].$$

This is **hydrodynamic-frame invariant**.

[AO (2026)]

# The main result: stability from positivity

$$\rho + P > 0 \implies \Gamma_L \geq 0 \implies \text{Im } \omega \leq 0$$

- This holds for **any** choice of hydrodynamic frame – Landau, Eckart, or otherwise
- The transverse mode gives the same conclusion, with

$$\Gamma_T = \frac{\eta}{\rho + P} \geq 0$$

[AO (2026)]



## Lorentz boosted frames

*Check the same conclusion holds in any inertial frame, not just comoving.*

# Non comoving frame

Lorentz boost between comoving ( $\tilde{k}_\mu$ ) and non comoving ( $k_\mu$ ) frames:

$$k_\mu \rightarrow \tilde{k}_\mu = \Lambda_\mu{}^\nu k_\nu.$$

For a boost in  $t$ - $x$  plain ( $\gamma \equiv 1/\sqrt{1-v^2}$ ),

$$\tilde{\omega} = \gamma(\omega - vk_x),$$

$$\tilde{k}_x = \gamma(k_x - v\omega),$$

$$\tilde{k}_y = k_y,$$

$$\tilde{k}_z = k_z.$$

In comoving frame, we found

$$\tilde{\omega} = \pm c_{s/n} \tilde{k} - i\Gamma_L \tilde{k}^2$$

Then what about  $\omega = \omega_0 + \varepsilon\omega_1 + \mathcal{O}(\varepsilon^2)$ ?

# First order dispersion

The leading order dispersion depends on the relative direction:

$$\omega_0 = \frac{v \left(1 - c_{s/n}^2\right) k_x \pm c_{s/n} \sqrt{1 - v^2} \sqrt{(1 - v^2) k_x^2 + \left(1 - c_{s/n}^2 v^2\right) k_\perp^2}}{1 - c_{s/n}^2 v^2}.$$

If  $\vec{k} \propto \vec{x}$  and then  $k_\perp = 0$ ,

$$\omega_0 = \frac{v \pm c_{s/n}}{1 \pm c_{s/n} v} k_x.$$

The first order dispersion relation:

$$\omega_1 = - \frac{i \Gamma_L \tilde{k}^2}{\gamma \left( 1 \pm c_{s/n} v \cos \tilde{\theta} \right)},$$

with

$$\cos \tilde{\theta} \equiv \frac{\gamma (k_x - v \omega_0)}{|\tilde{k}|},$$
$$\tilde{k}^2 \Big|_{\omega_0} \equiv \gamma^2 (k_x - v \omega_0)^2 + k_{\perp}^2.$$

This is stable.



## Conclusions and outlook

*Put it together: first-order hydrodynamics is a healthy IR effective theory.*

# First-order hydrodynamics is a healthy IR effective theory

- The physical IR pole does not depend on the hydrodynamic frame
- Under KMS + unitarity,  $\rho + P > 0 \Rightarrow \text{Im } \omega \leq 0$
- Lorentz boosts maintain the stability of the IR spectrum.

As an **IR effective theory**, no instability issue in first-order hydrodynamics.

[AO (2026)]



# Outlook: curved and time-dependent backgrounds

- Coupling hydrodynamic modes to gravitational perturbations
- Extension to FLRW, Einstein static universe, black hole backgrounds
- NG modes and stability under different spacetime symmetry breaking
- Effects of background curvature and time dependence on the IR pole

# Outlook: nonlinear fluctuations and loop corrections

- Loop corrections sourced by the hydrodynamic modes themselves
- IR renormalization of transport coefficients; long-time tails
- A hydrodynamic-frame-invariant nonlinear effective action

**Next goal:** a frame-independent, nonlinear, quantum effective theory of relativistic fluids.

Thank you for your attention.

## References

- A. Ota, *Frame-independent analysis of relativistic hydrodynamics via effective theory*, Phys. Rev. D 114, L011902 (2026).
- A. Ota, *Mode analysis of Nambu-Goldstone modes in  $U(1)$  charged first-order relativistic viscous hydrodynamics*, Phys. Rev. D 114, 016001 (2026).

Contact: aota@cqu.edu.cn