

Spectrometry of the Earth core using Hyper-K

: Sensitivity study

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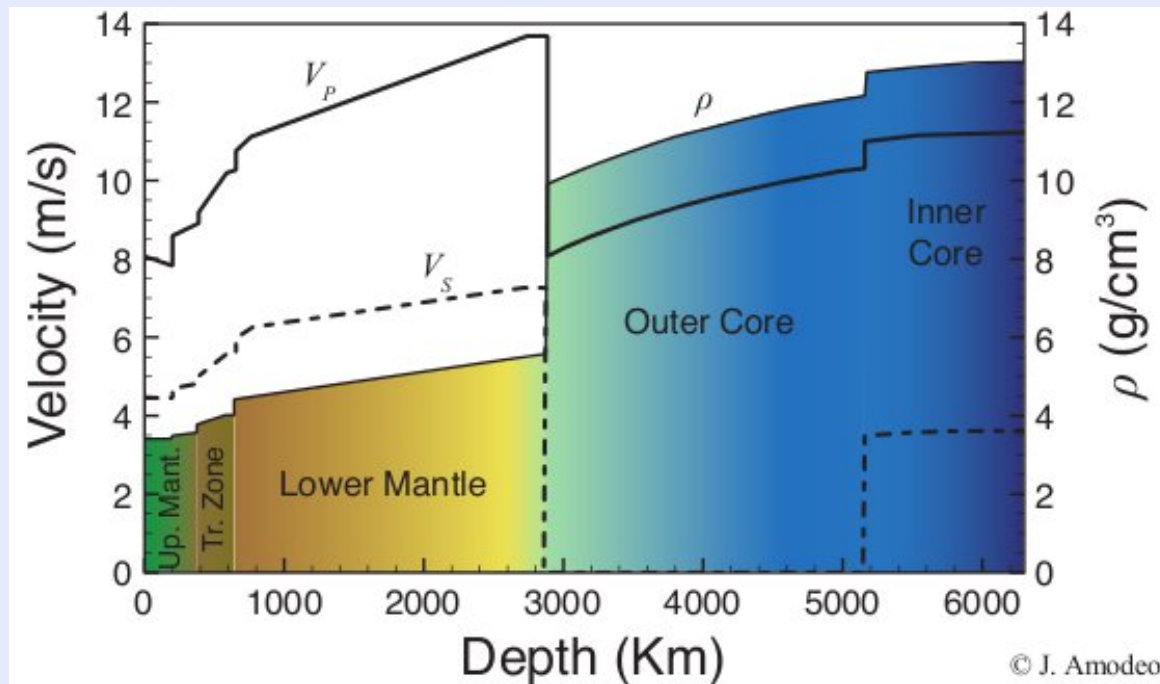
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Current understanding

- ◆ Matter density profile is well known
 - ◆ Seismic measurement and free oscillation

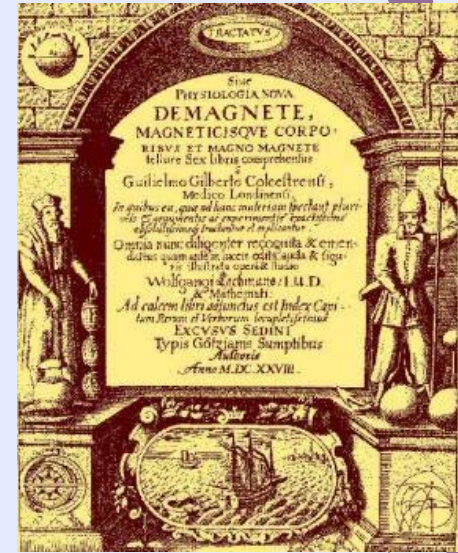
Outer core is assumed to be liquid iron (+Ni)

- ◆ And some other light element
- ◆ **But it's not measured**

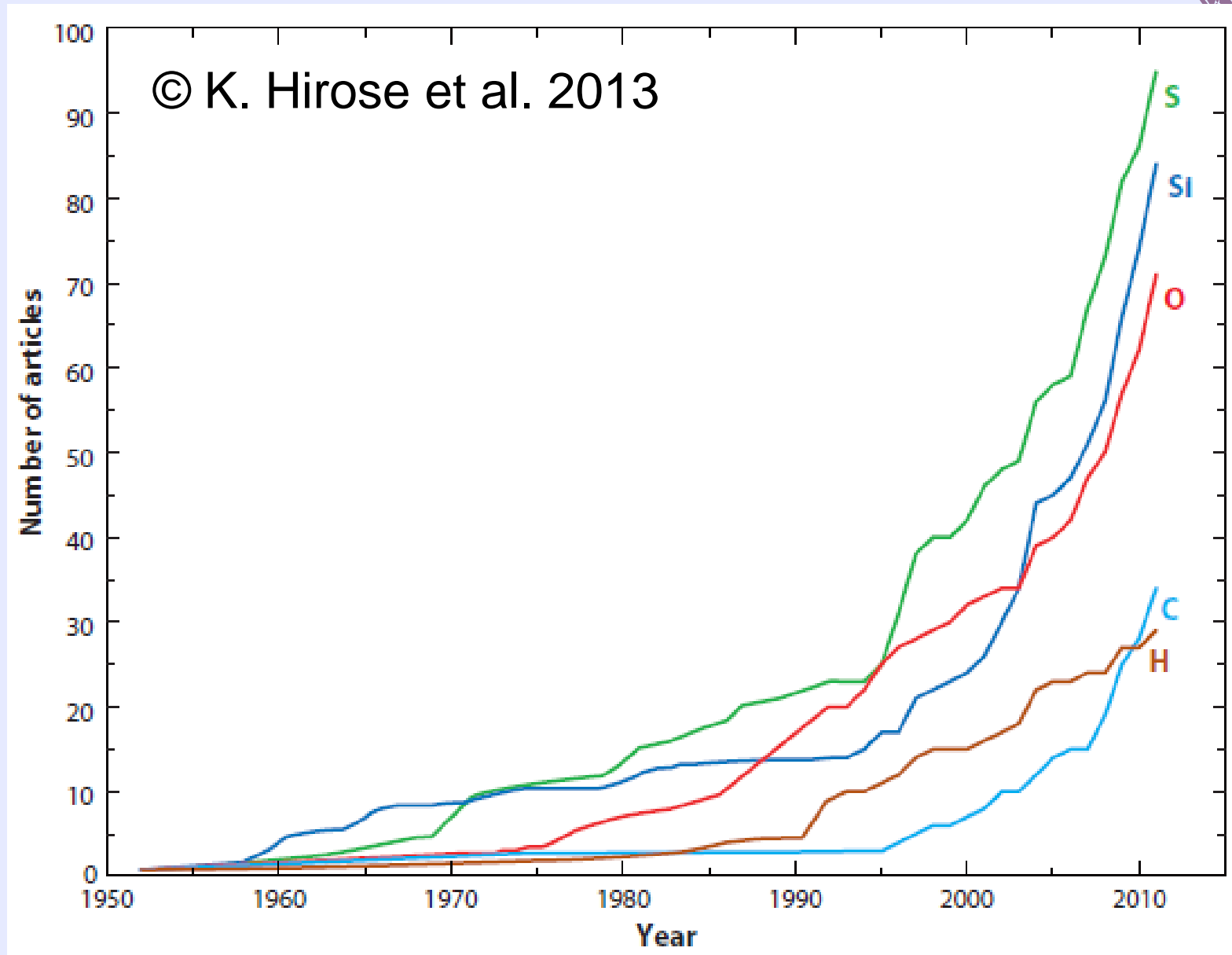


Why is light element important ?

- ◆ History of the Earth's evolution
- ◆ What is **the geomagnetic field** ?
Who order it ?
 - ◆ W. Gilbert 1600
 - ◆ A. Einstein 1905
 - ◆ It requires the metal convection : dynamo theory
- ◆ Outer core composition is essentially important
 - ◆ Especially light component
 - ◆ Pure iron cannot maintain the convection
 - ◆ H? O? C? Si? S?



Cumulative number of articles



Neutrino physics to geophysics

- ◆ Neutrino oscillation probability depend on **electron density, not matter density**
- ◆ By using neutrino oscillation, we can measure the electron density of the medium
 - ◆ If we know the neutrino property very well
- ◆ We have the precise matter density of the earth
 - ◆ From seismic wave tomography and free oscillation
- ◆ Combining matter density and electron density, **we can measure the average chemical composition of the deep earth !**
 - ◆ Ratio of atomic number to mass number (Z/A)

Hypothesis

- ◆ Z/A ratio of materials
 - ◆ Fe :0.466, Light material :~0.5, Hydrogen :1
 - ◆ **More sensitive to Hydrogen**

Matter density model : modified PREM

Initial neutrino flux : Honda flux 2011

Mantle is pyrolite (Z/A=0.496)

Inner Core is Pure Iron (Z/A=0.467)

- ◆ **Z/A of outer core : free parameter**

- ◆ **Normal hierarchy**

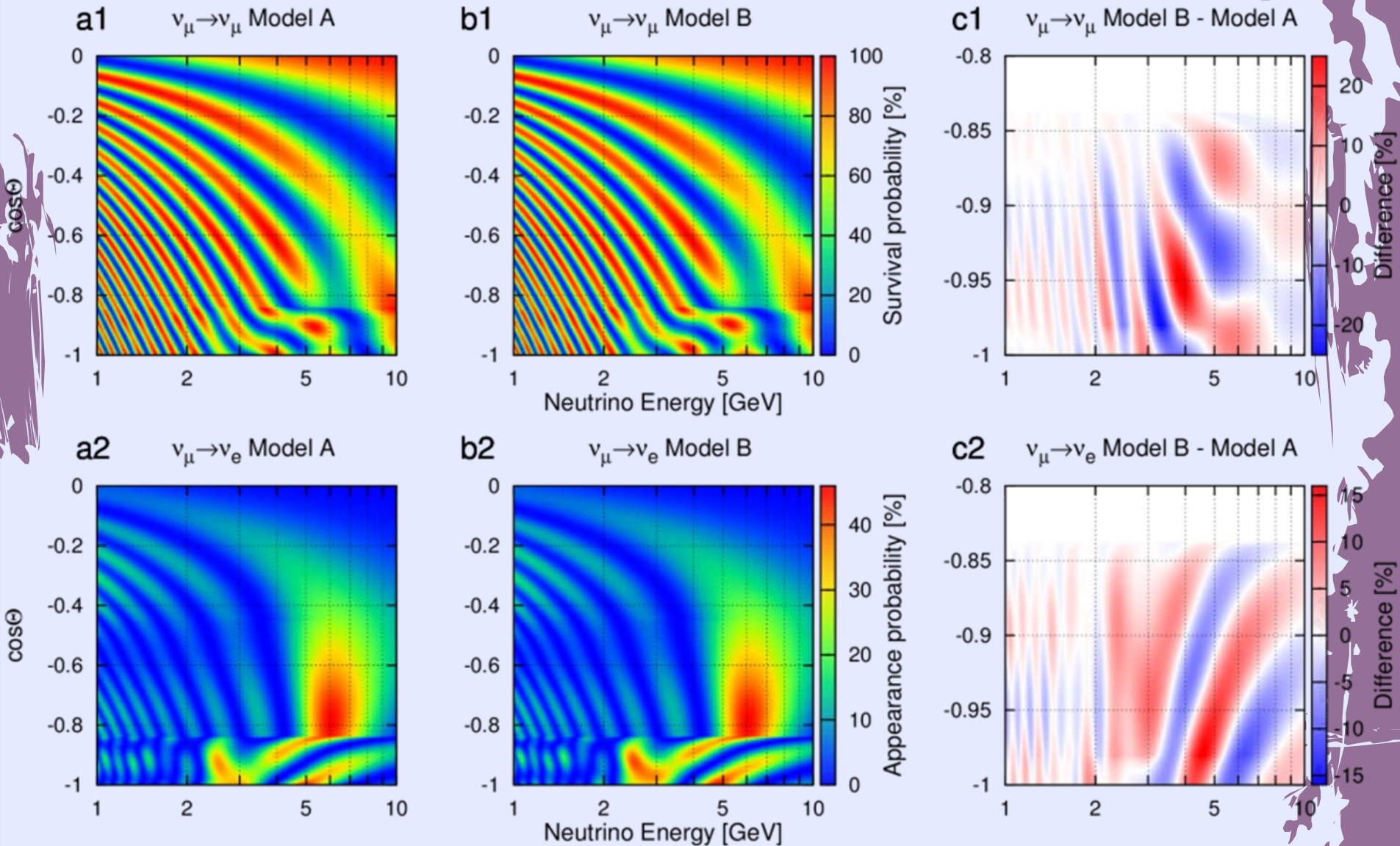
- ◆ Oscillation parameter : Capozzi et al. 2014

$$\sin^2 \theta_{12} = 0.308_{-0.017}^{+0.017} \quad \delta_{CP} = 1.39_{-0.27}^{+0.38} \times \pi$$

$$\sin^2 \theta_{13} = 0.0234_{-0.0019}^{+0.0020} \quad \Delta m_{21}^2 = (7.54_{-0.22}^{+0.26}) \times 10^{-5}$$

$$\sin^2 \theta_{23} = 0.437_{-0.023}^{+0.033} \quad \Delta m_{32}^2 = (2.39_{-0.06}^{+0.06}) \times 10^{-3}$$

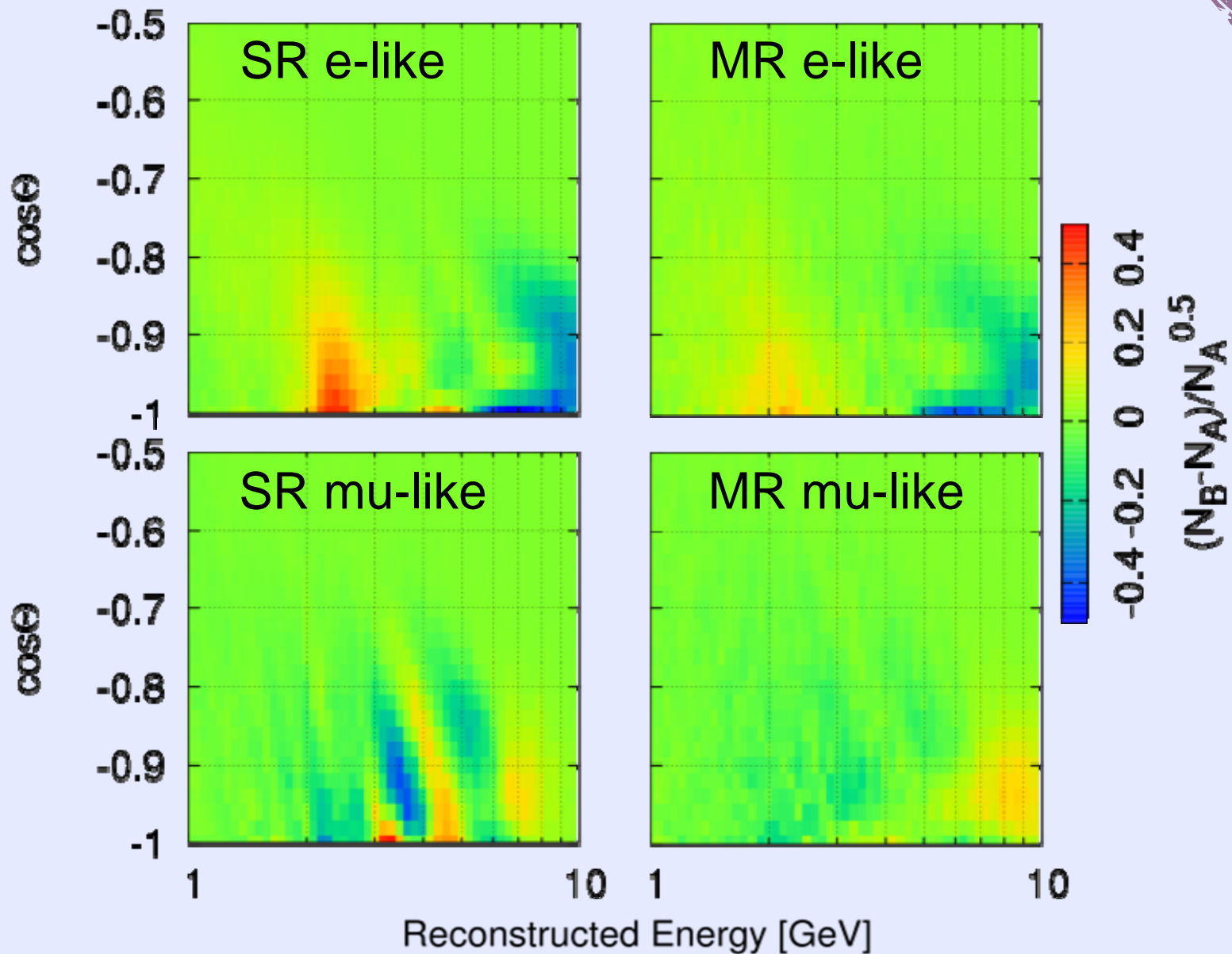
Oscillograms (Fe OC VS rock OC)



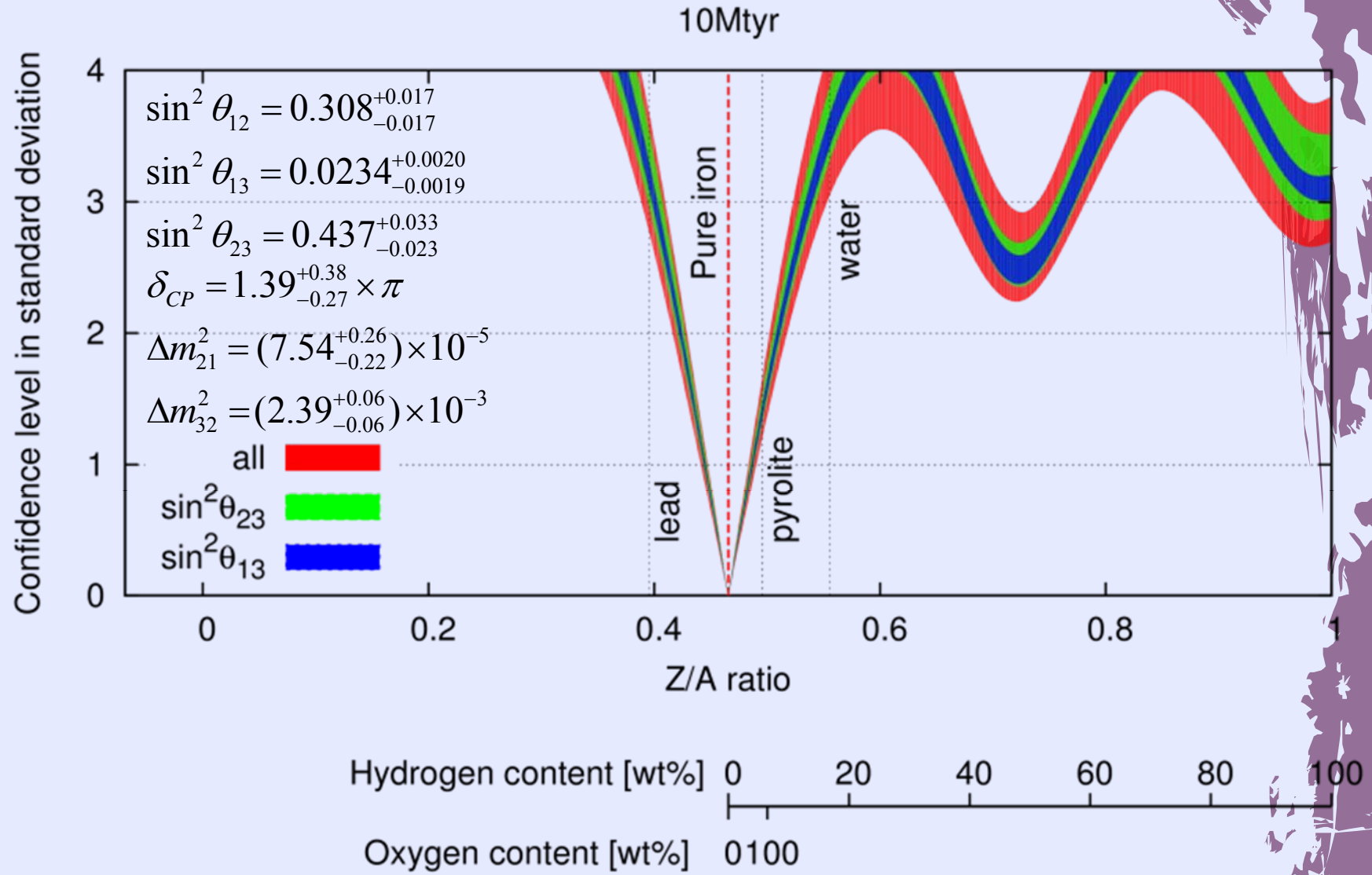
Sensitivity estimation

- ◆ Event template generation
 - ◆ Reconstructed SK-MC event : only 11.25Mt yr
 - ◆ Not enough for likelihood calculation using pseudo experiment
 - ◆ Event enhancement (1Gtyr)
 - ◆ Assume efficiency and resolution are independent to arrival direction
 - ◆ Calculate oscillation weight and add the weight to each bin
 - ◆ Bin size is 0.02 for $\log(P_{\text{rec}})$ and 0.02 for $\cos\theta_{\text{rec}}$
 - ◆ A template contain 4 event type : (SR mu, SR e, MR mu, MR e)
- ◆ P-value(confidence level) calculation
 - ◆ Generate likelihood ratio (LR) distribution
 - ◆ Generate pseudo experiments using 2 different templates A, B
 - ◆ Calculate the median of LR_B : M_B
 - ◆ Integrate $\text{LR}_A > M_B$

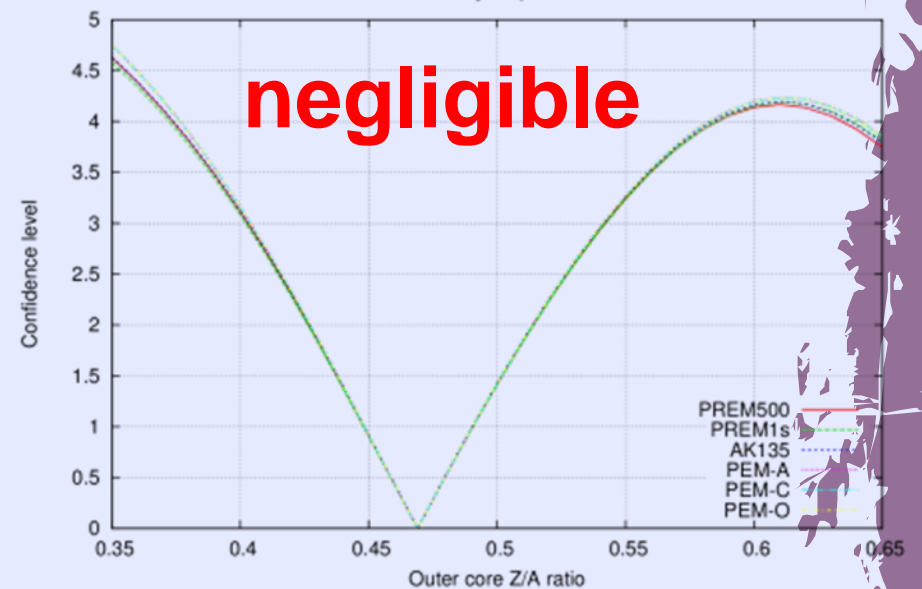
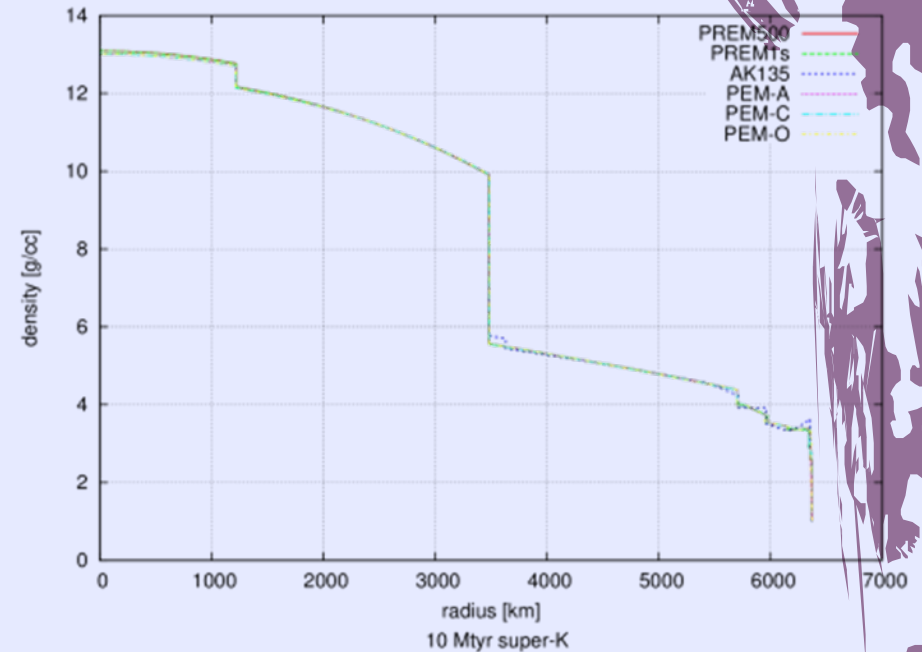
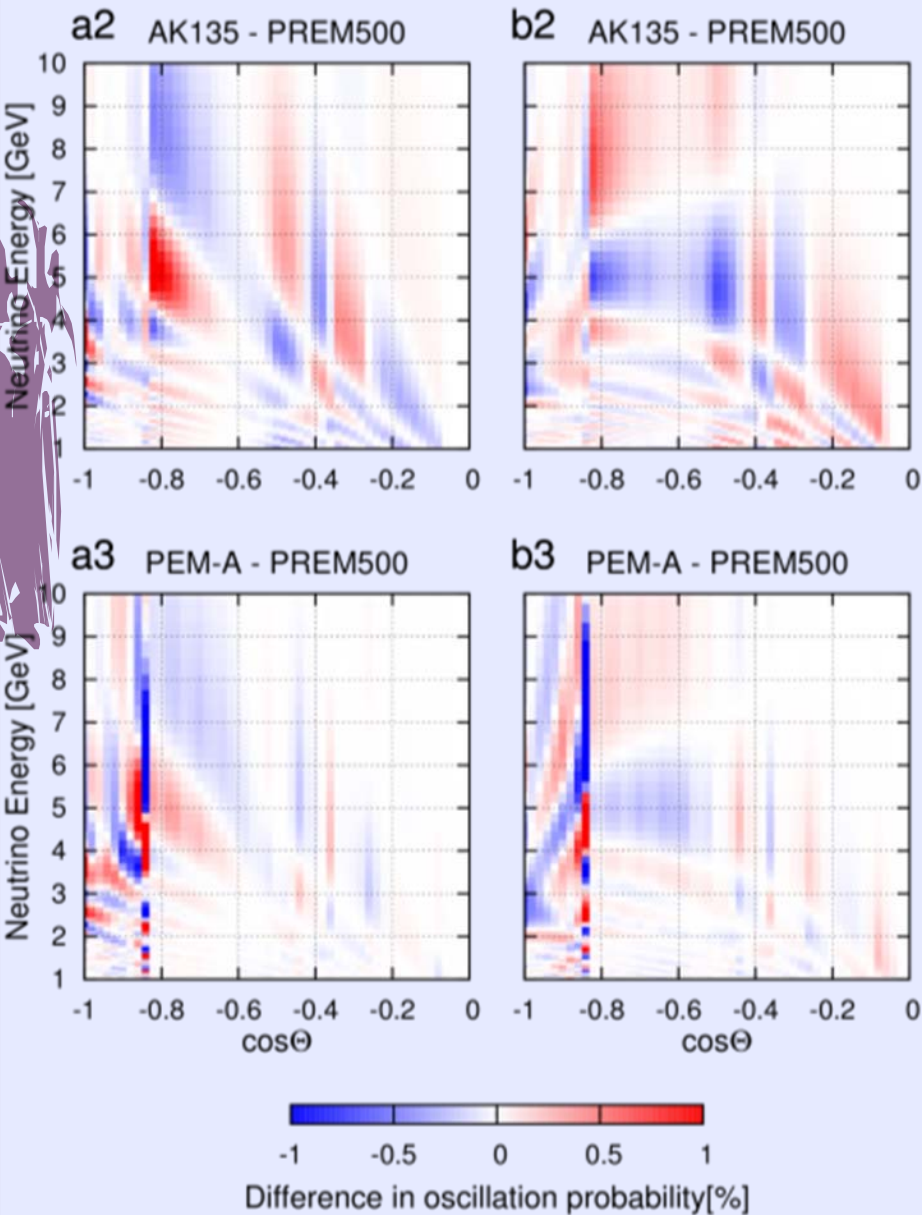
Significance map (Fe OC VS rock OC)



Expected sensitivity (HK 10Mtyrs)



Uncertainty from matter density model



TODO

- ◆ Try fiTQun and add event classification
 - ◆ Currently single-ring e-like (SREL), MREL, SRML, MRML
 - ◆ Following classification can be tested
 - ◆ Single particle : e, μ , π^0
 - ◆ Double particle : $e\pi^0$, $\mu\pi^0$, other
 - ◆ More than 2 particle
 - ◆ But this method increases MC sys. uncertainty
- ◆ Include upward muon event and PC event
- ◆ Systematic uncertainty estimation
 - ◆ Neutrino flux, Cross sections, Detector

Conclusion

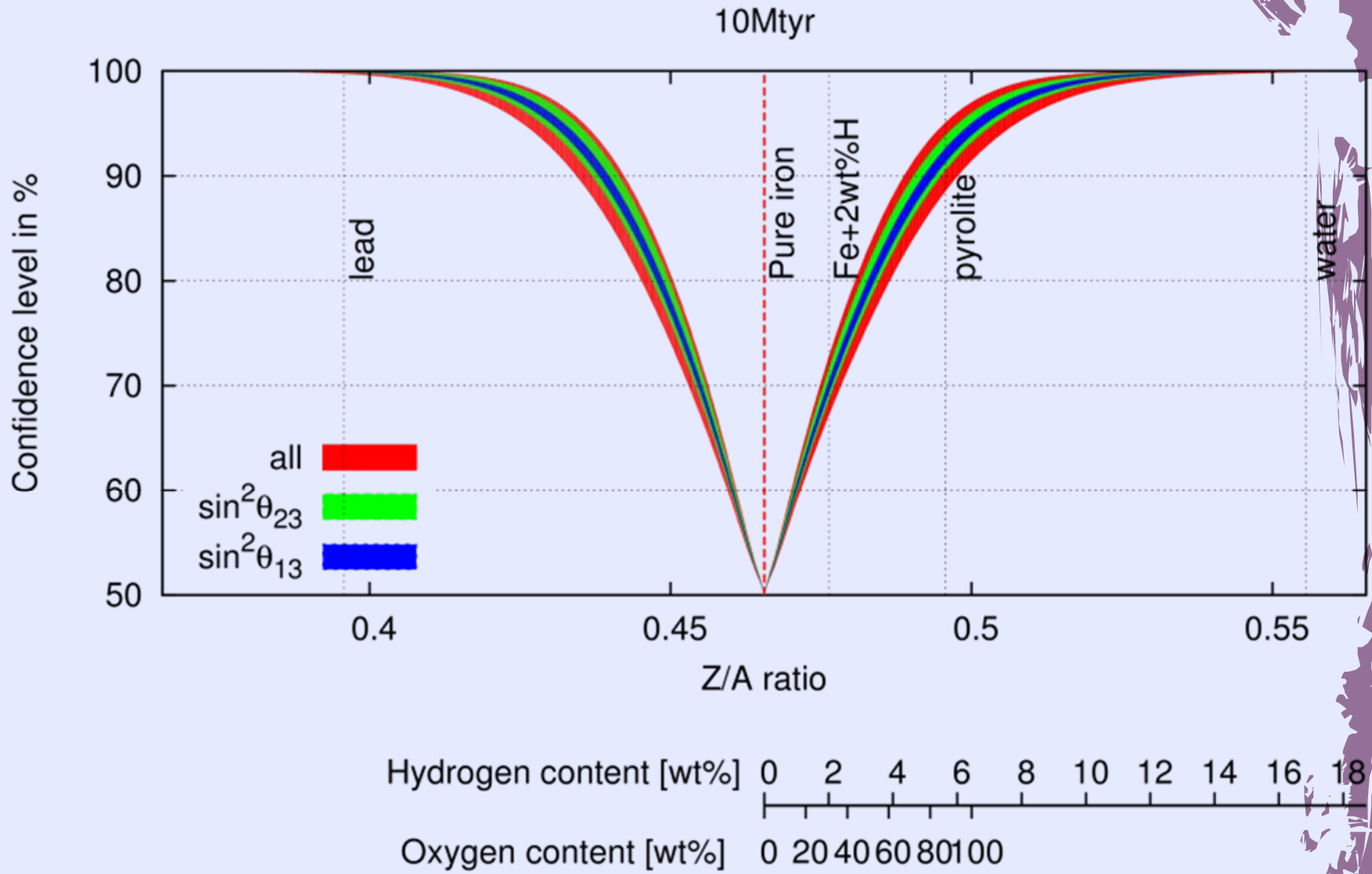
- ◆ Neutrino oscillation is applicable to geophysics
 - ◆ Only way to measure the chemical composition of the deep Earth
 - ◆ Larger detector volume is essentially important for this study
 - ◆ Sensitive energy range : **>2GeV**
 - ◆ **Precise oscillation parameter measurements are useful for geophysics**
- ◆ Uncertainty from matter density models is negligible
- ◆ Hyper-K can start spectrometry of the Earth
 - ◆ Fe, Pb, water can be resolved
 - ◆ Excludable Z/A range : $0.389 >$, $0.553 <$ @ 3σ , $0.434 >$, $0.498 <$ @90%
 - ◆ **Geophysicist's request : $0.466 >$, $0.472 <$ @ 1σ**
 - ◆ **HK provides the first data point(s) of Z/A**
 - ◆ **Larger detector volume is essentially important**

Geophysicist meets physicists...

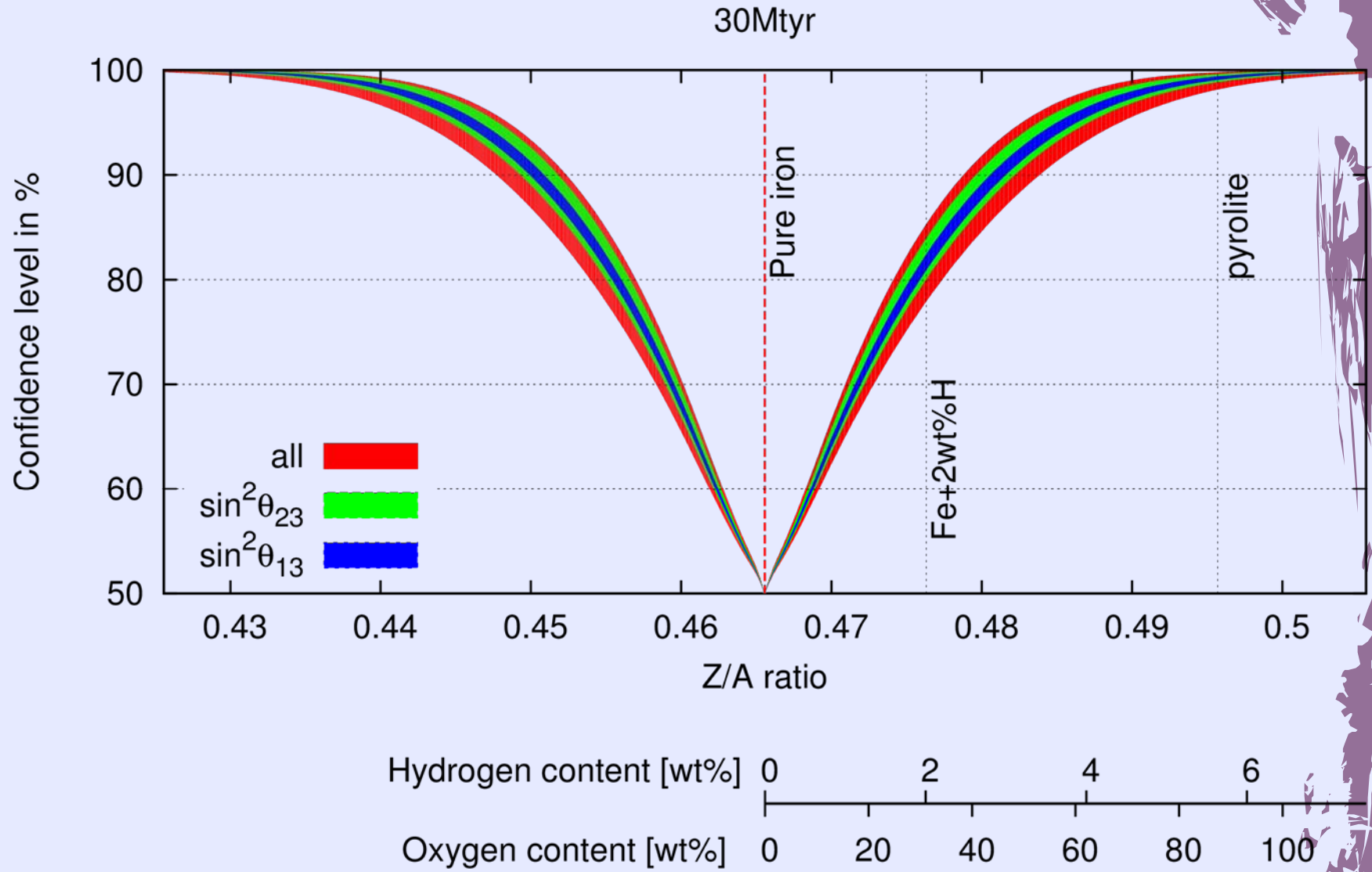


Thank you for your attention.

Expected sensitivity (HK 10Mtyrs)



Expected sensitivity (HK 30Mtyrs)



Event template generation

- ◆ Reconstructed SK-MC event : only 11.25Mt yr
 - ◆ Not enough for likelihood calculation using pseudo experiment

Event enhancement (1Gtyr)

- ◆ Assume det./reco. efficiency is independent to arrival direction
- ◆ Select an event : $(P_v^0, \Theta_v^0, \Phi_v^0, P_{rec}^0, \Theta_{rec}^0, \Phi_{rec}^0)$
- ◆ Search 1Gt/11.25Mt(=888 events) : $(P_v^i, \Theta_v^i, \Phi_v^i, P_{rec}^i, \Theta_{rec}^i, \Phi_{rec}^i)$
 - ◆ Same flavour and similar momentum
 - ◆ Maximum difference of momentum was 3% (1 – 10 GeV range)
- ◆ Scale and scramble each event : $(P_v^0, \Theta_v^0, \Phi_v^0, P_{rec}^i \times P_v^0 / P_v^i, \vartheta_{rec}^i, \phi_{rec}^i)$
- ◆ Calculate oscillation weight : $P(x \rightarrow x) + P(x \rightarrow y) F_y(P_v^0, \Theta_v^0) / F_x(P_v^0, \Theta_v^0)$
- ◆ Add the weight to correspond bin $M(P_{rec}^i \times P_v^0 / P_v^i, \vartheta_{rec}^i)$
 - ◆ Bin size is 0.02 for $\log(P_{rec}^i \times P_v^0 / P_v^i)$ and 0.02 for $\cos\vartheta_{rec}$
 - ◆ A template contain 4 event type : (SR mu, SR e, MR mu, MR e)

Log likelihood ratio (LLR) calculation

- ◆ Generate pseudo experiments using 2 different templates (m_A, m_B)
 - ◆ $N_A(E, \cos\Theta) \leftarrow \text{Poisson}(m_A(E, \cos\Theta))$
 - ◆ $N_B(E, \cos\Theta) \leftarrow \text{Poisson}(m_B(E, \cos\Theta))$
 - ◆ $m(E, \cos\Theta) = (M t \text{ yr}) * M(E, \cos\Theta)$
 - ◆ 4 data set is created for each psuedo experiment
 - ◆ SR mu, SR e, MR mu, MR e
- ◆ $\ln L_{\text{total}} = \ln L_{\text{SR mu}} + \ln L_{\text{SR e}} + \ln L_{\text{MR mu}} + \ln L_{\text{MR e}}$
 $\ln L = \sum (N(P, \cos\Theta) \ln(m(P, \cos\Theta)) - m(P, \cos\Theta) - \ln(N(P, \cos\Theta)!))$
- ◆ Calculate log likelihood ratio
 - ◆ $\text{LLR}_A = \ln L(N_A, m_B) - \ln L(N_A, m_A)$
 - ◆ $\text{LLR}_B = \ln L(N_B, m_B) - \ln L(N_B, m_A)$

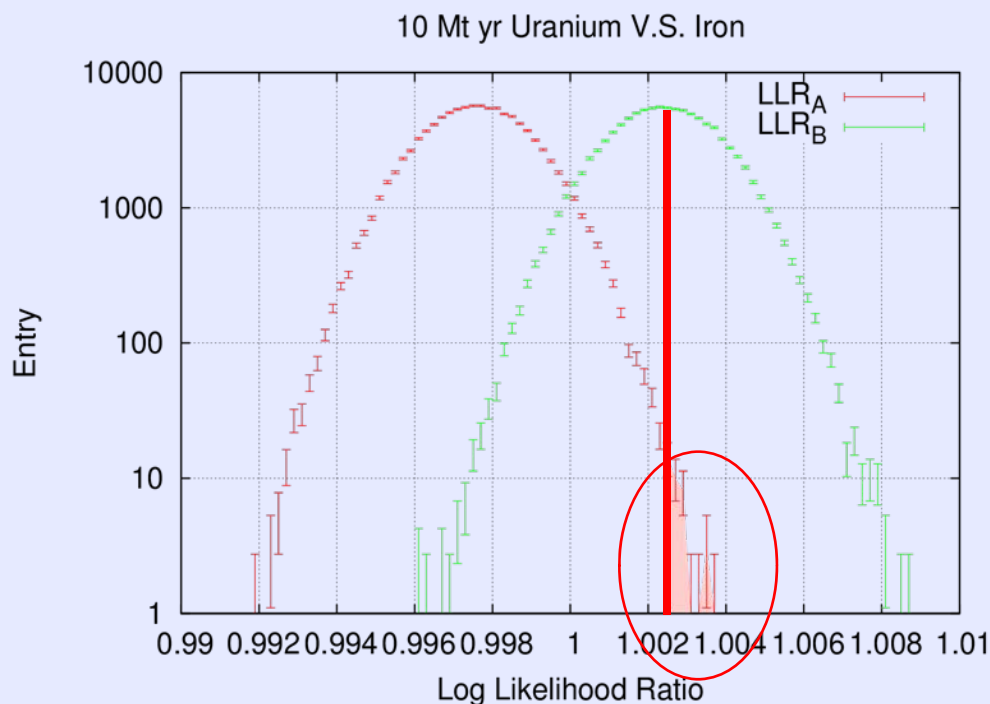
P-value (confidence level)

- ◆ Calculate the median of LR_B : $M(LR_B)$
- ◆ count the number of $LR_A > M(LR_B)$: N_{over}

$P = N_{\text{over}} / N_{\text{tot}}$ determines the probability of 2 different model exclusion

Actual calculation was done by convolution of probabilities

- ◆ Not by MC



The Earth's core models



Model name	Z/A ratio	Si(wt%)	O(wt%)	S(wt%)	C(wt%)	H(wt%)	reference
Single light element model (maximum abundance)							
Fe+18wt%Si	0.4715	18	-	-	-	-	Poirier ²⁸
Fe+11wt%O	0.4693	-	11	-	-	-	Poirier ²⁸
Fe+13wt%S	0.4699	-	-	13	-	-	Li and Fei ⁵
Fe+12wt%C	0.4697	-	-	-	12	-	Li and Fei ⁵
Fe+1wt%H	0.4709	-	-	-	-	1	Li and Fei ⁵
Multiple light elements model							
Allegre2001	0.4699	7	5	1.21	-	-	Allègre et al. ²⁵
McDonough2003	0.4682	6	0	1.9	0.2	0.06	McDonough ²⁶
Huang2011	0.4678	-	0.1	5.7	-	-	Huang et al. ²⁷

