

ν PRISM:

An Experimental Method to
Remove Neutrino Interaction
Uncertainties from Oscillation
Experiments

Mike Wilking
Stony Brook University
6th Hyper-K Meeting
January 30th, 2015

precision
reaction
independent
spectrum
measurement
 ν PRISM:

An Experimental Method to Remove Neutrino Interaction Uncertainties from Oscillation Experiments

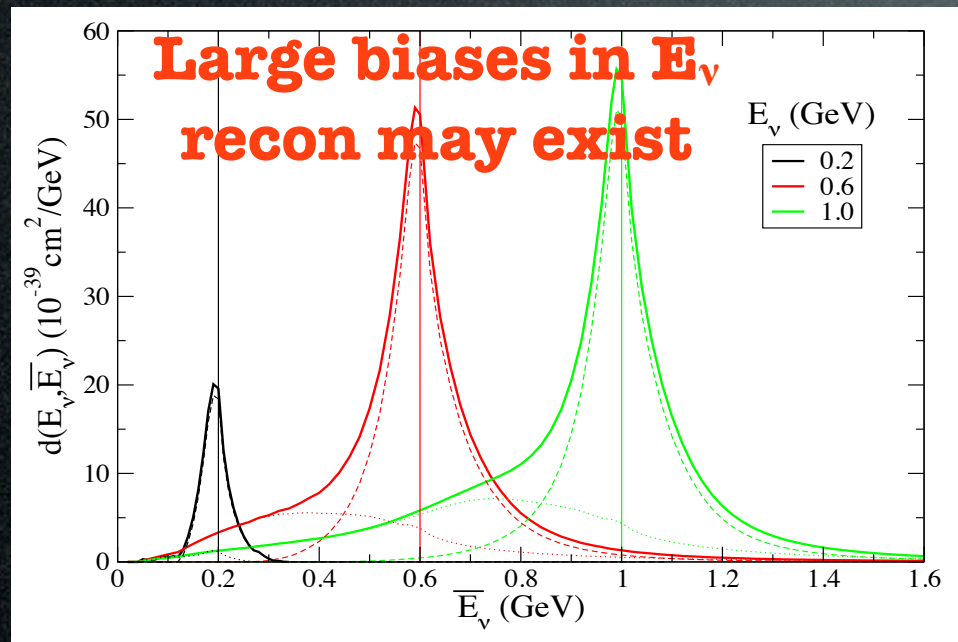
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Overview

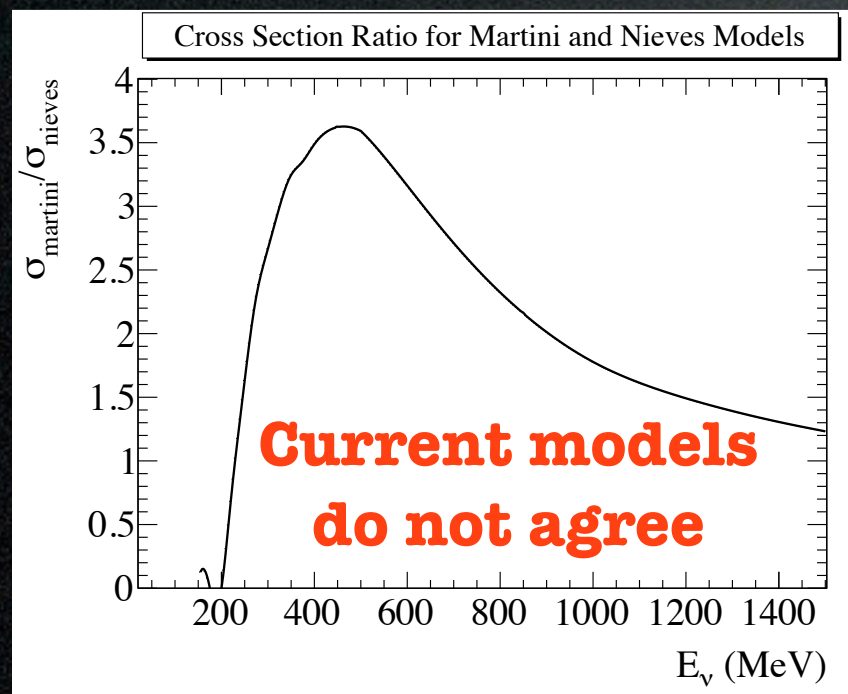
- Brief Reminder of NuPRISM Concept
 - E_ν Measurement Problem
 - Constraining E_ν with Linear Combinations
- Plans for other measurements
 - CPV & ν_e , sterile- ν , cross sections
- NuPRISM-Lite: Current Status
- Next Steps

Why Hyper-K Needs NuPRISM:

The E_ν Measurement Problem



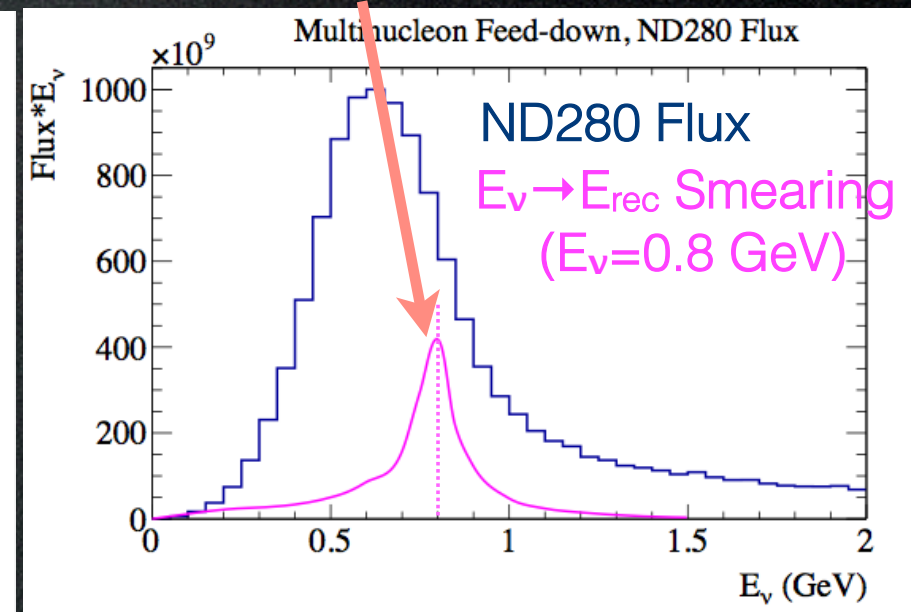
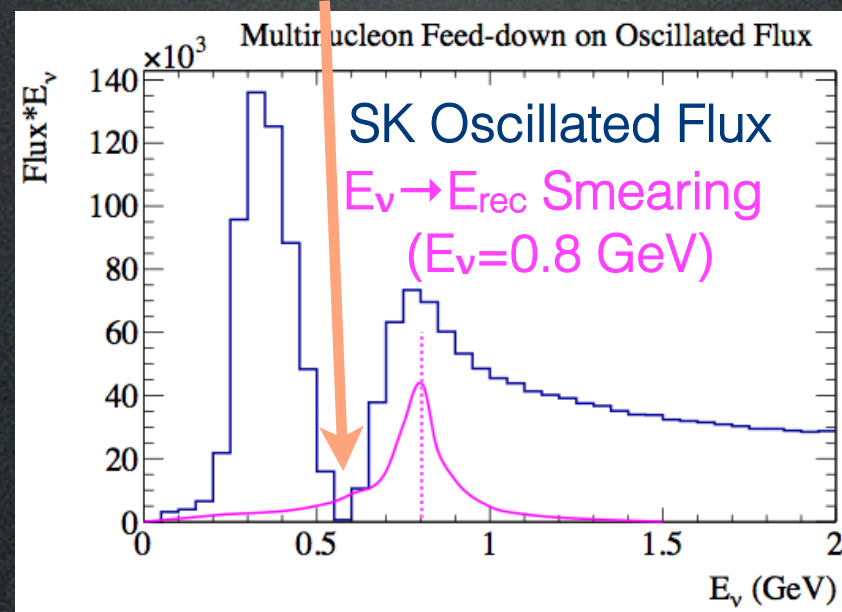
- It is now believed that large E biases can exist due to nuclear and non-nuclear effects (e.g. multinucleon interactions)
- Models are very difficult to produce and show large disagreements
- Without a data-driven constraint, this will likely be a dominant uncertainty for T2HK
- Typical near detectors likely cannot provide a sufficient constraint



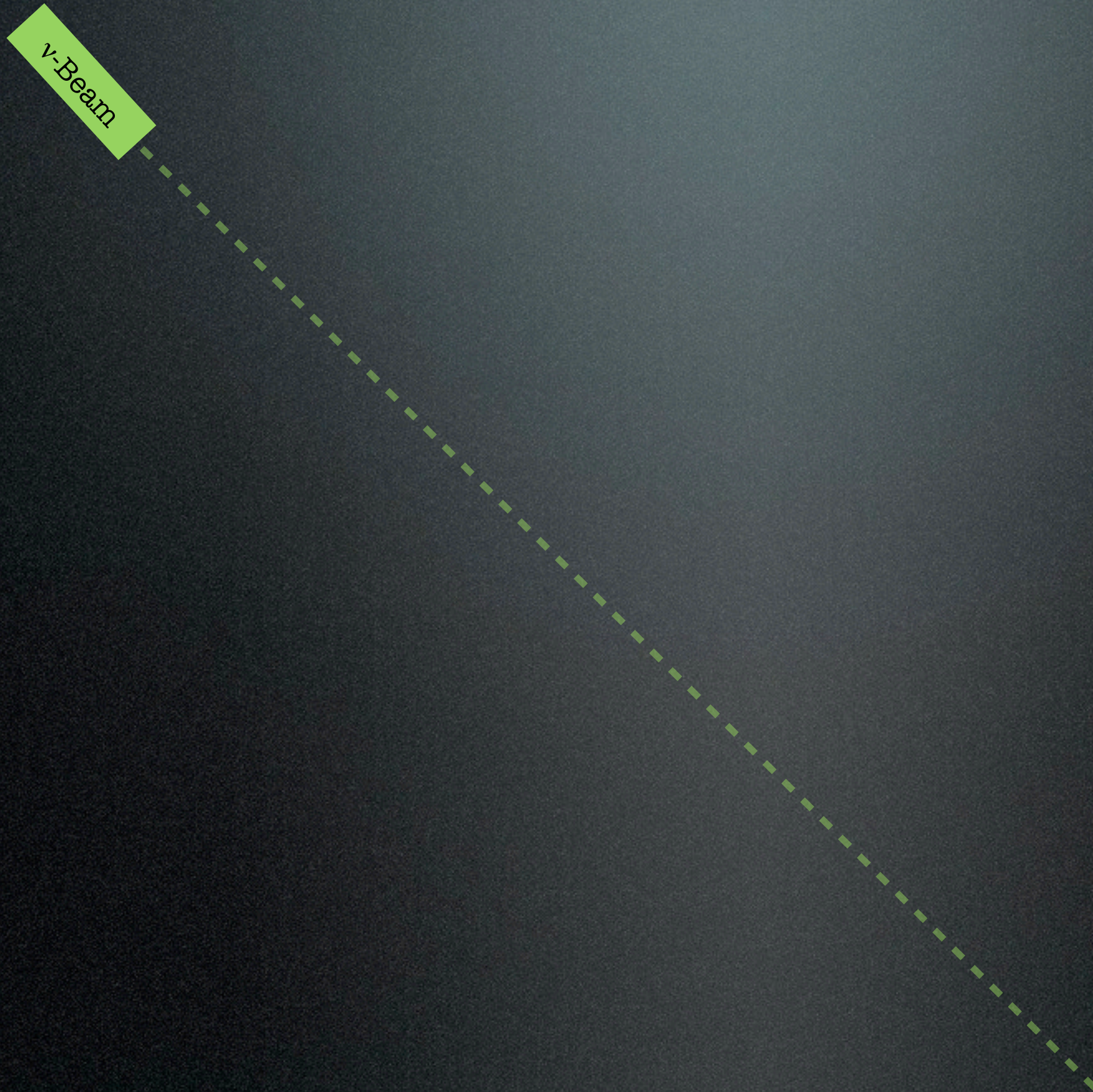
- J. Nieves, I. Ruiz Simo, and M. J. Vicente Vacas, PRC 83:045501 (2011)
- M. Martini, M. Ericson, G. Chanfray, and J. Marteau, PRC 80:065501 (2009)

Mixing Angle Bias!

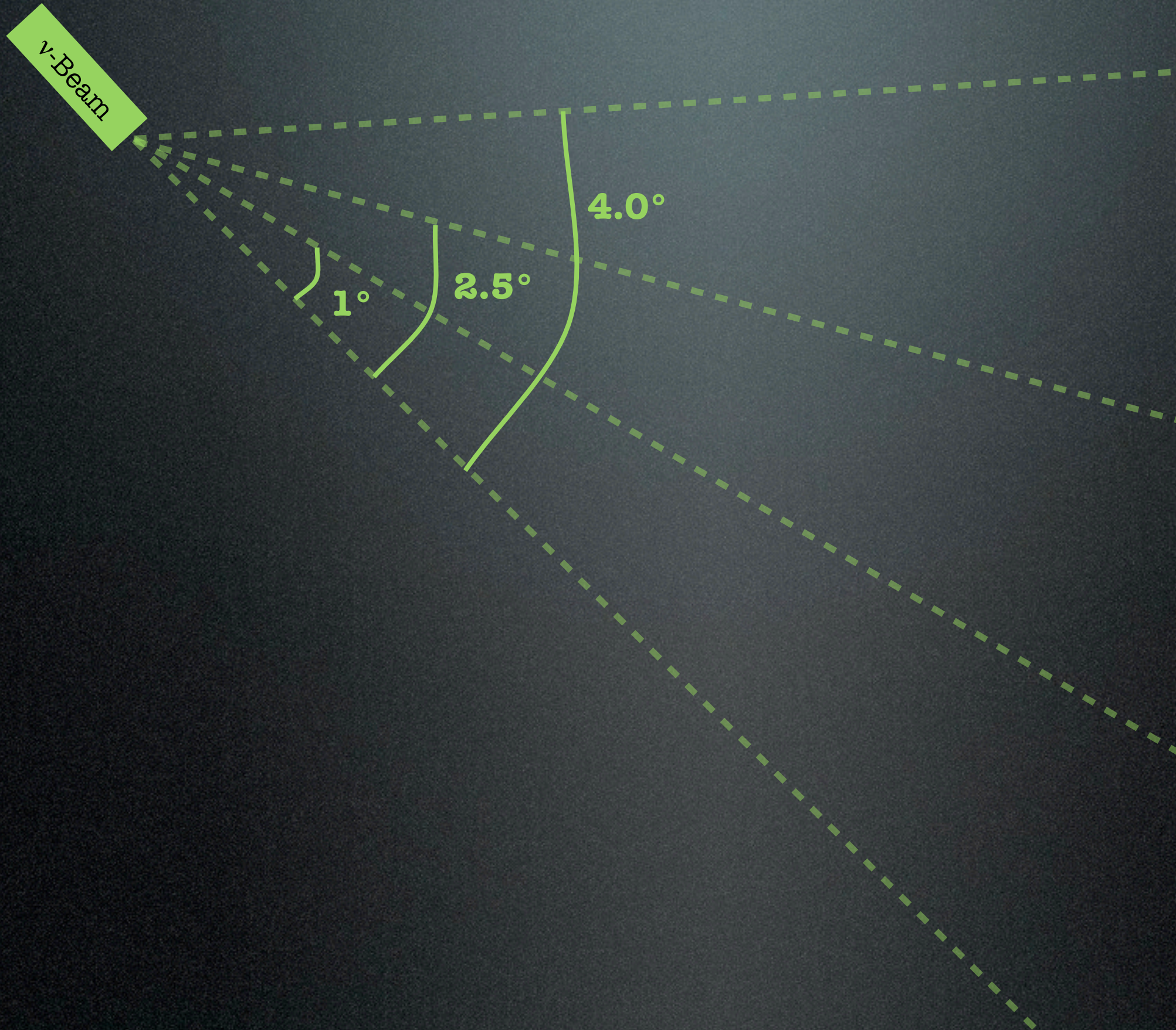
Typical ND lacks sensitivity



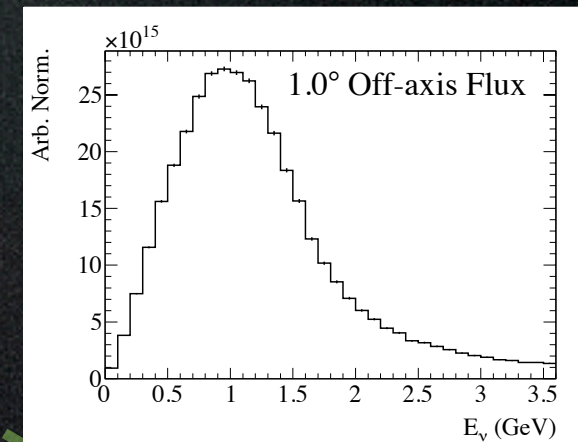
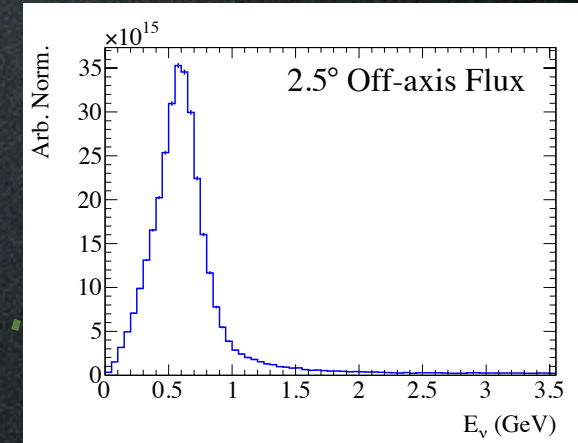
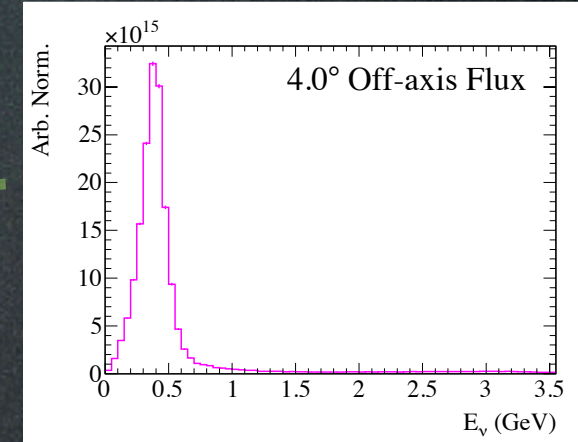
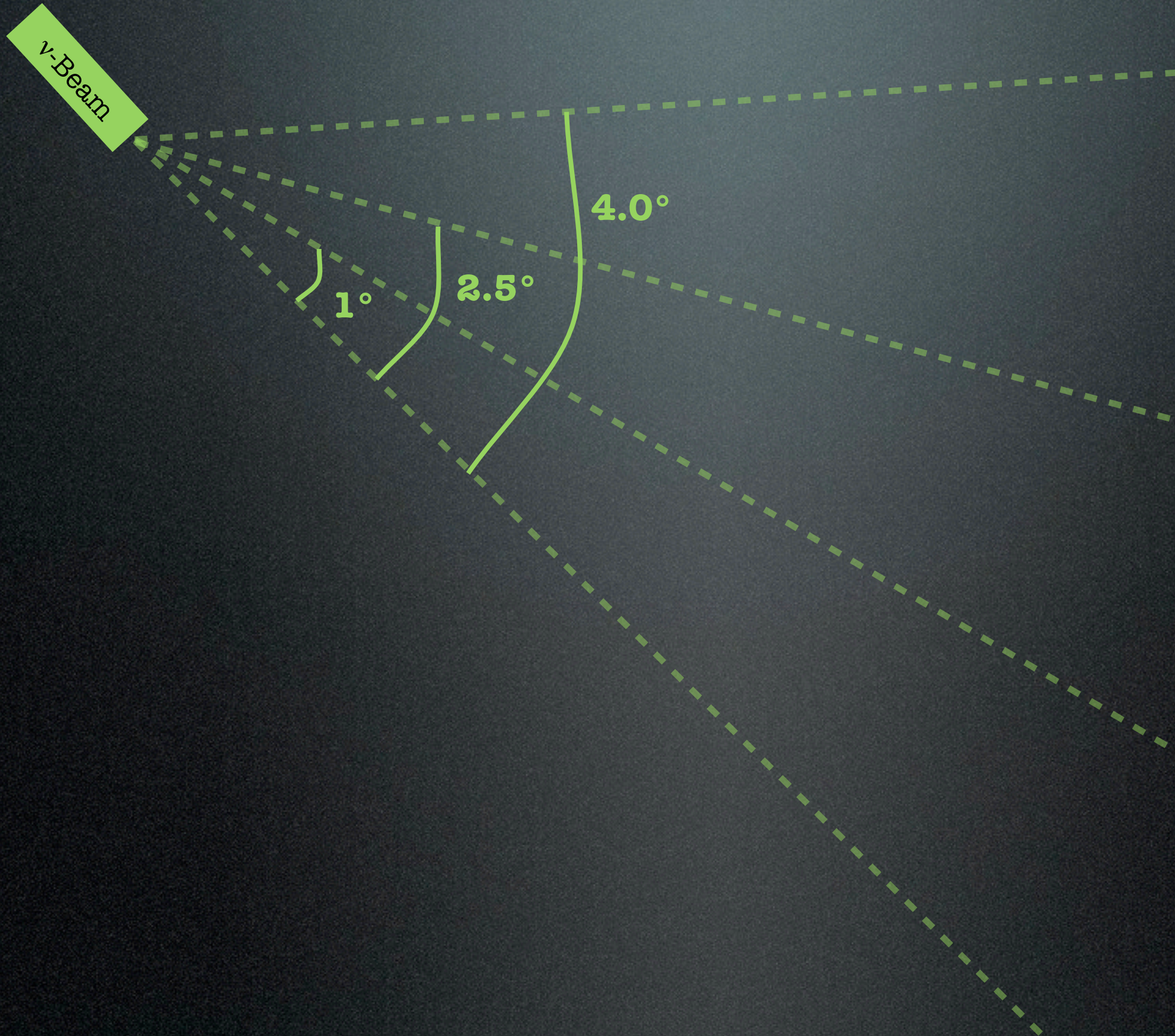
Reminder: NuPRISM Detector Concept



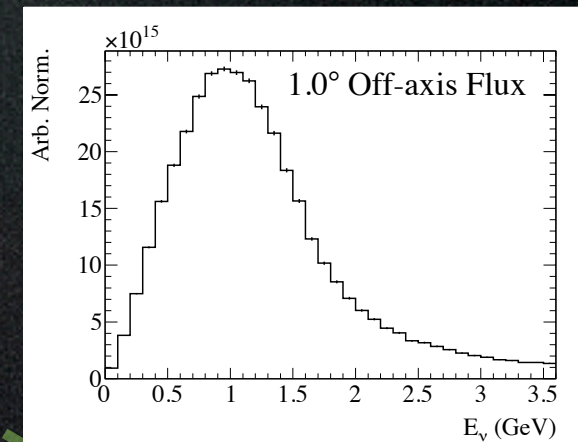
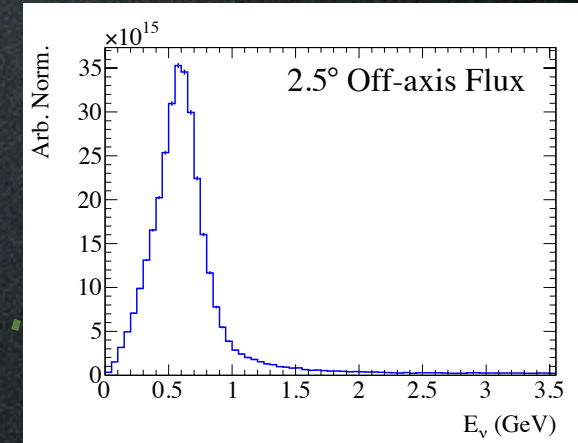
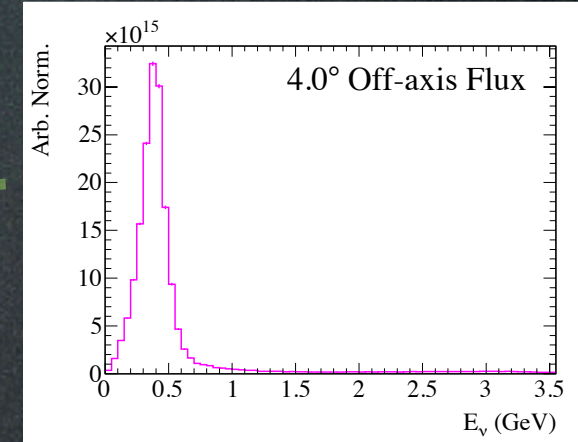
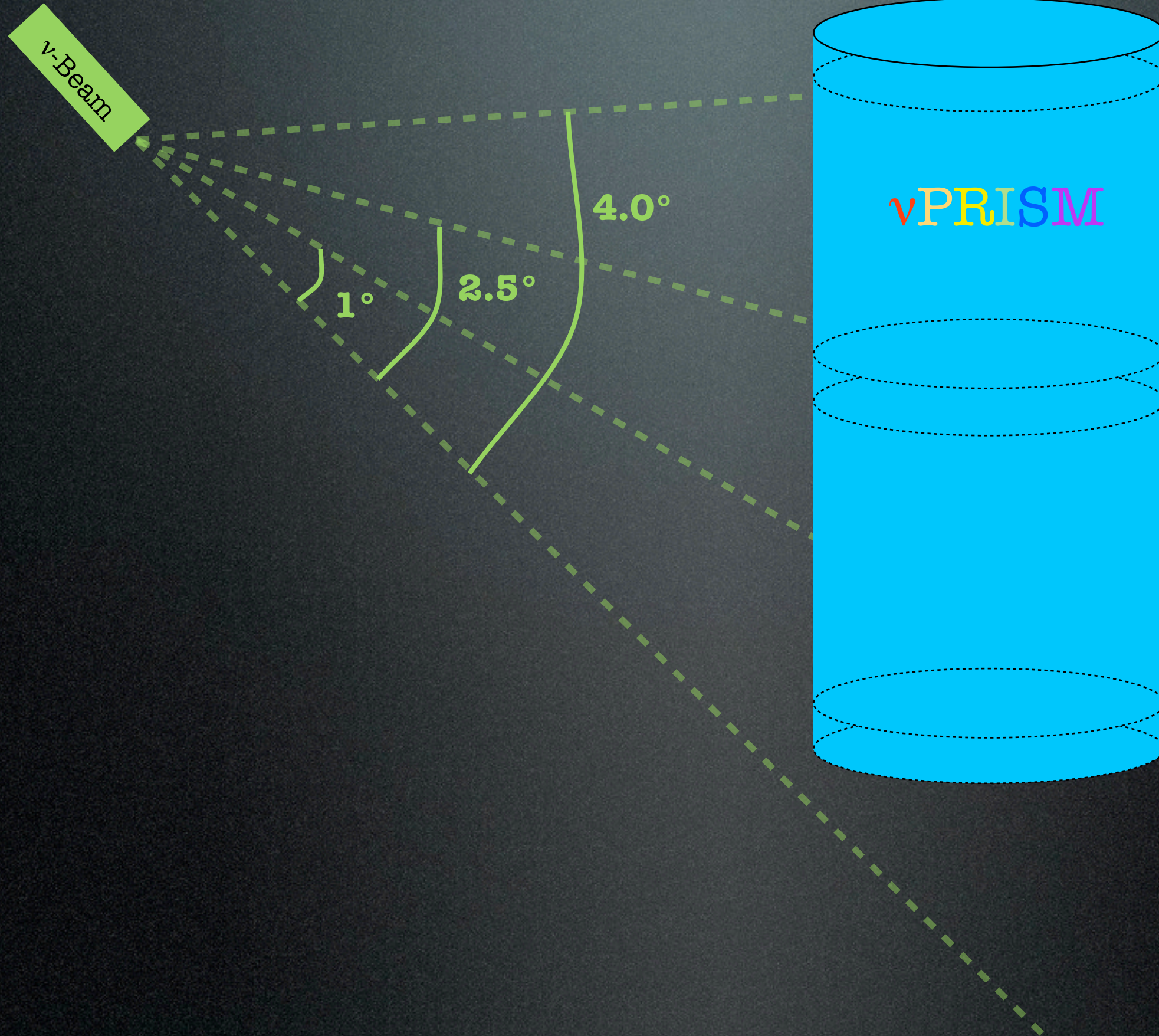
Reminder: NuPRISM Detector Concept



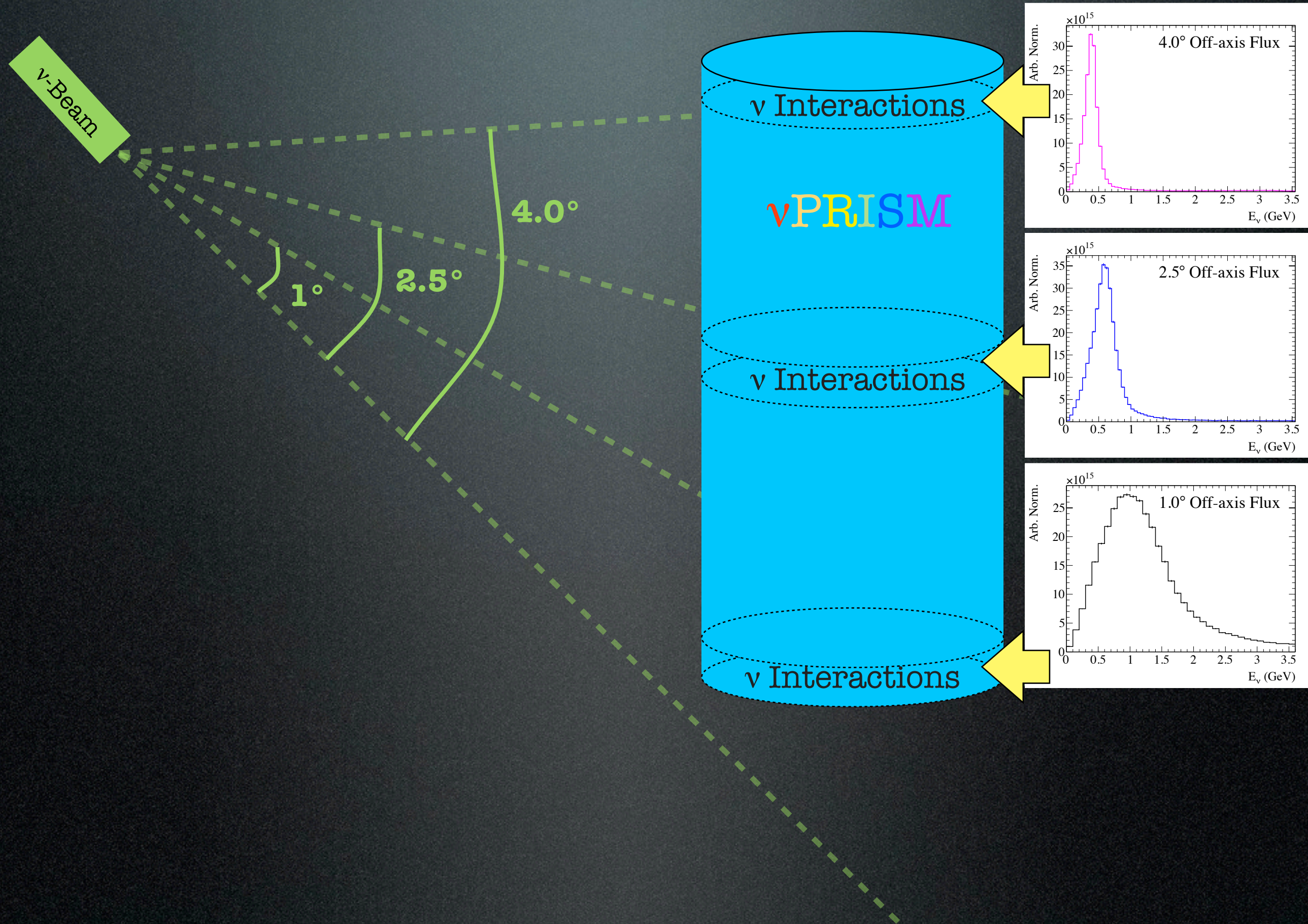
Reminder: NuPRISM Detector Concept



Reminder: NuPRISM Detector Concept

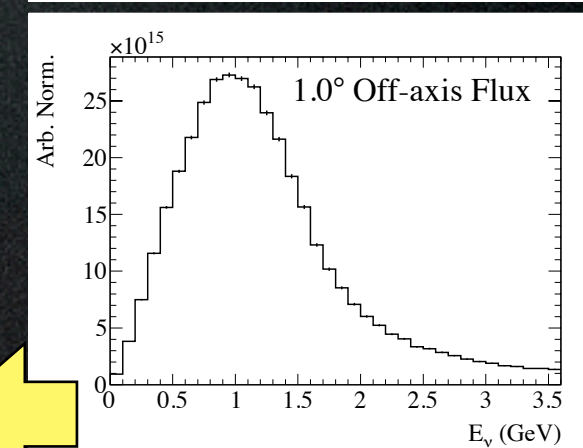
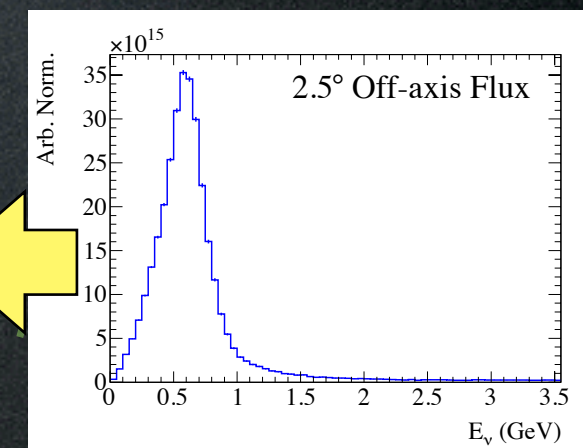
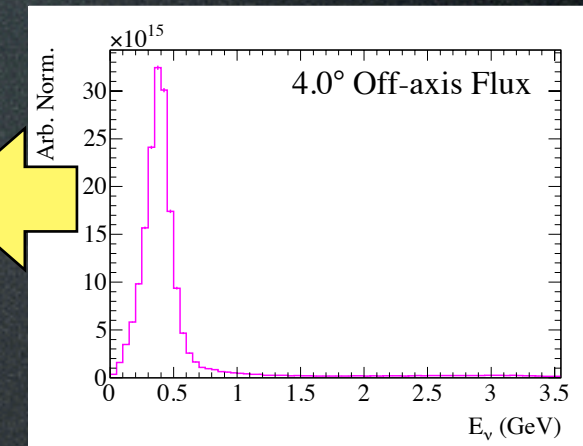
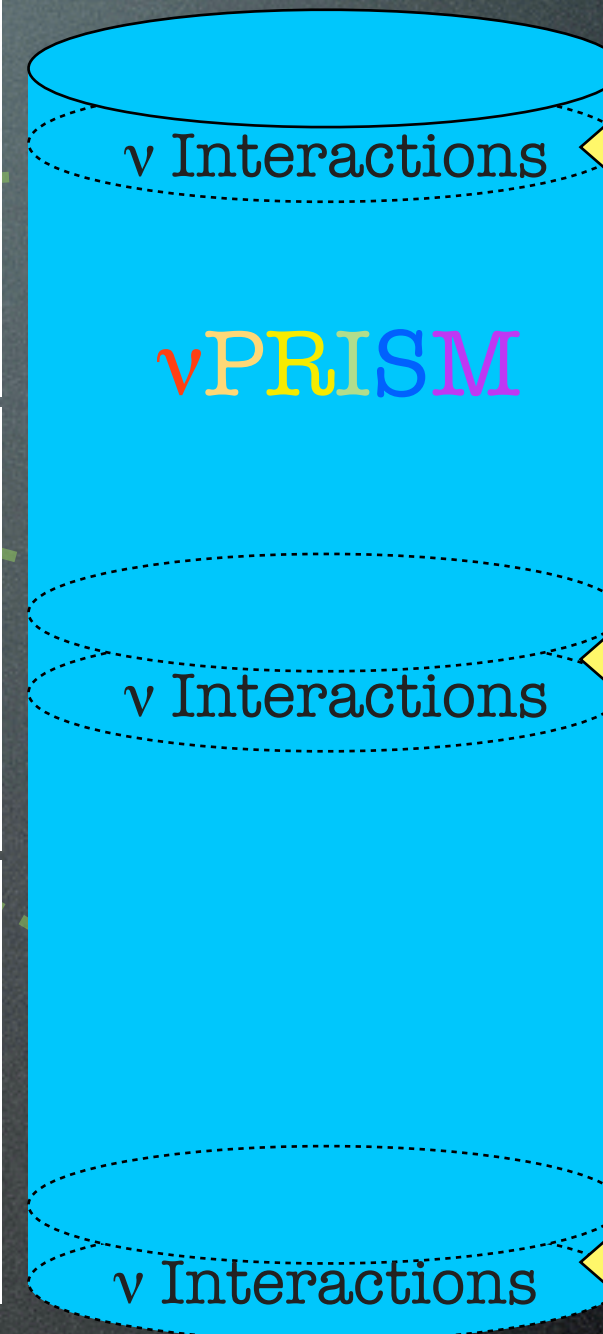
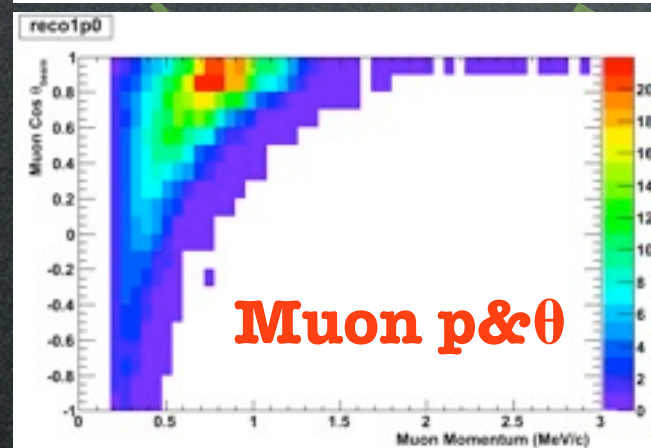
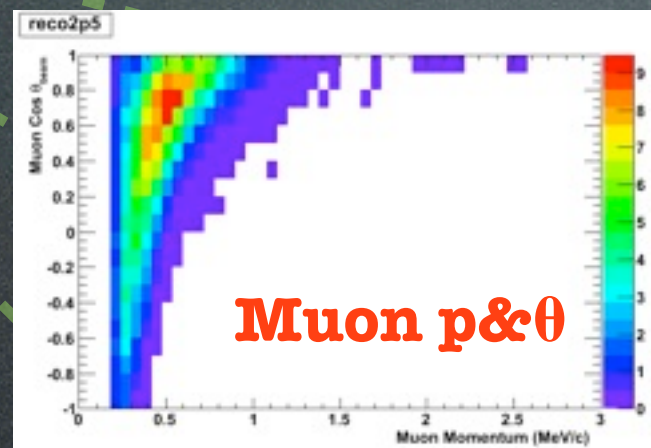
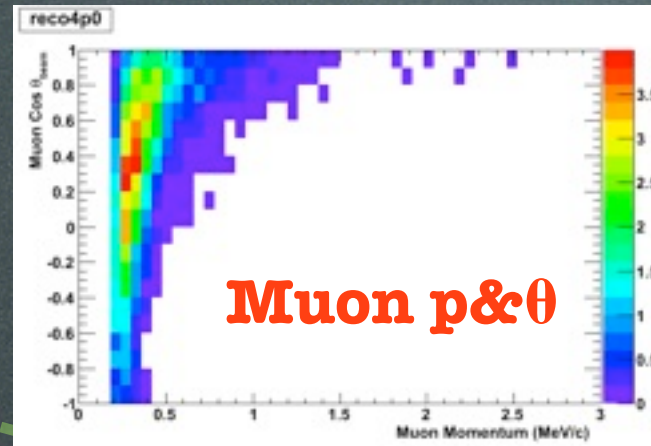


Reminder: NuPRISM Detector Concept

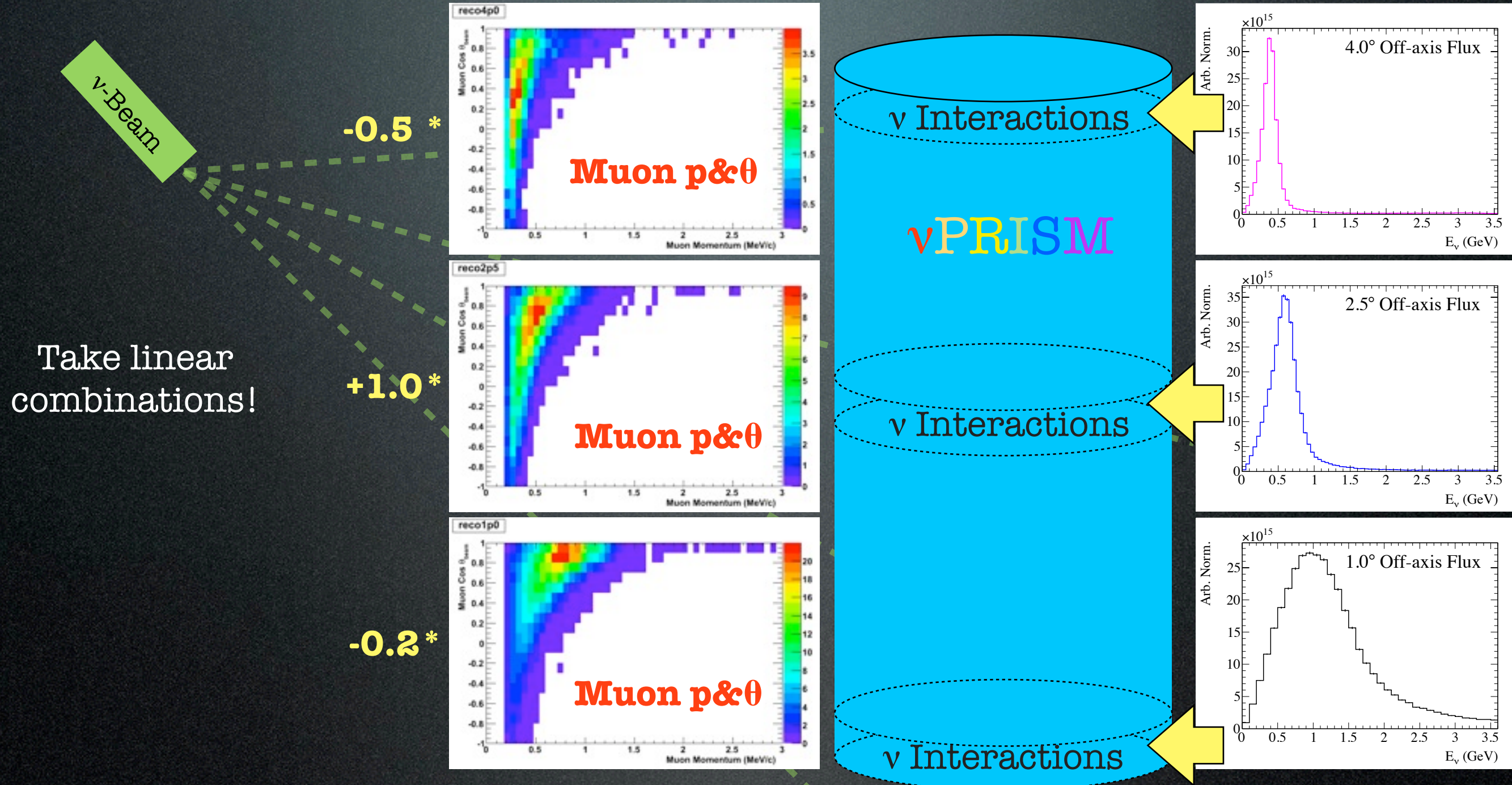


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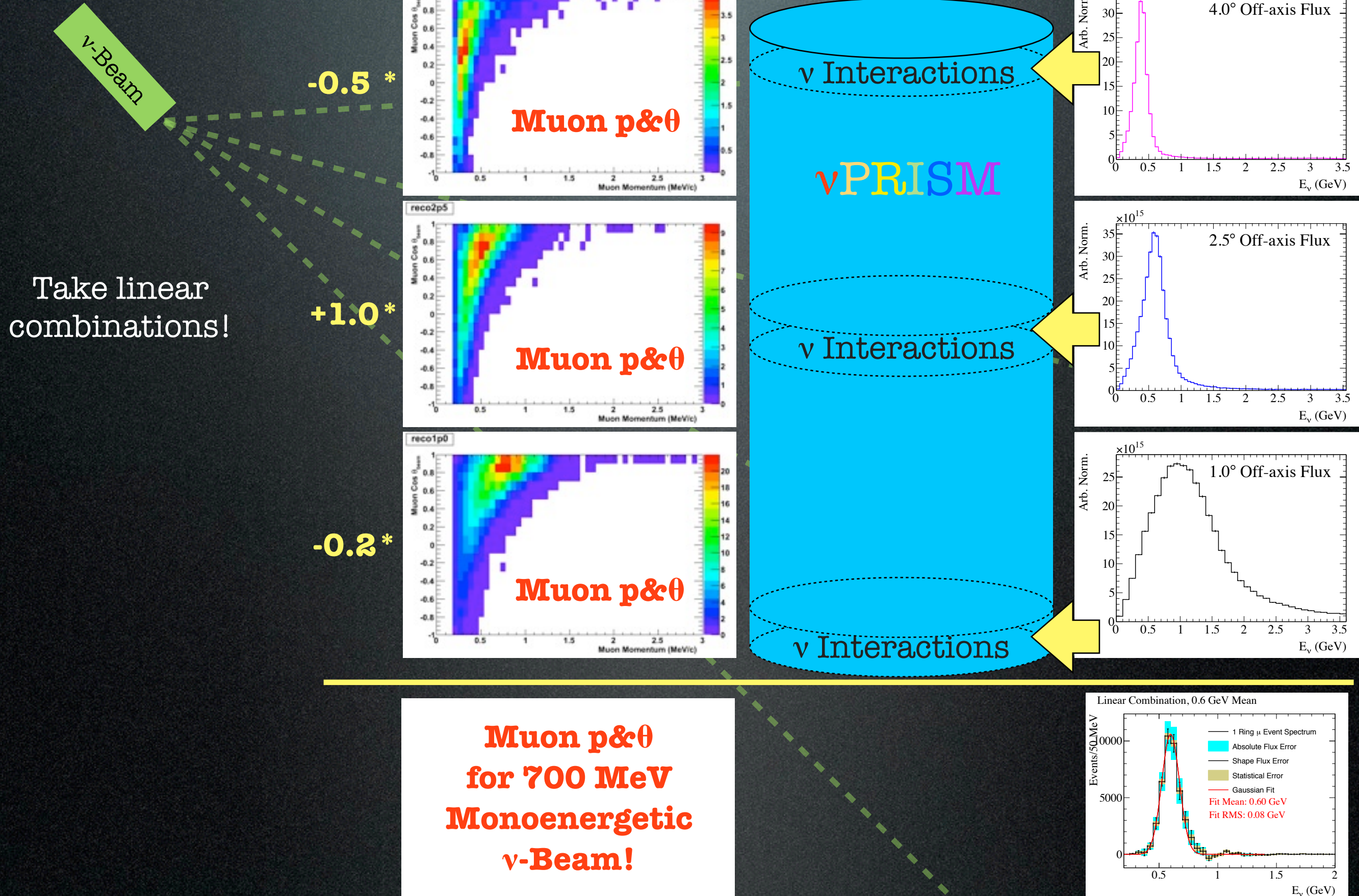
ν -Beam



Reminder: NuPRISM Detector Concept



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Reminder: NuPRISM Detector Concept

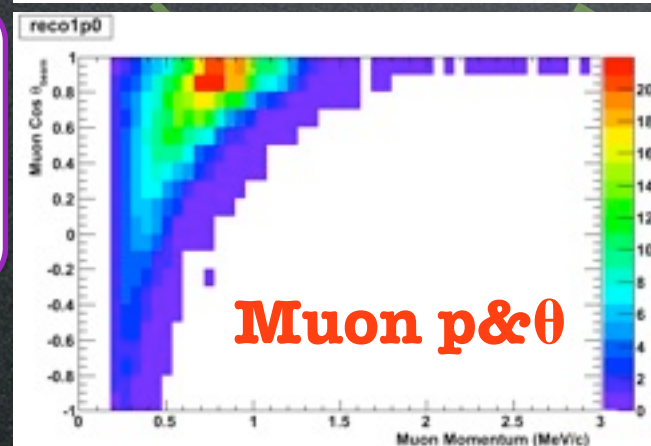
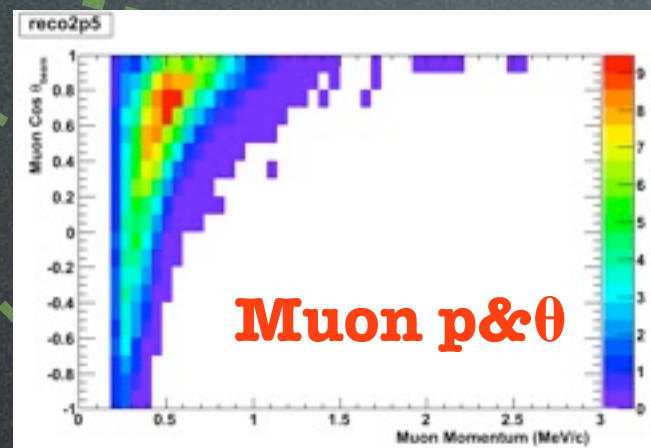
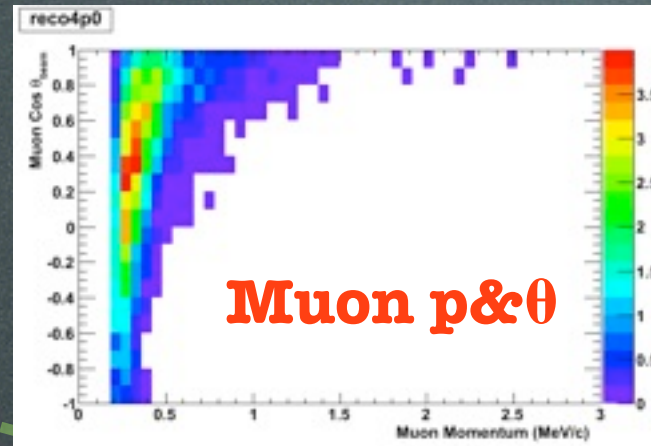
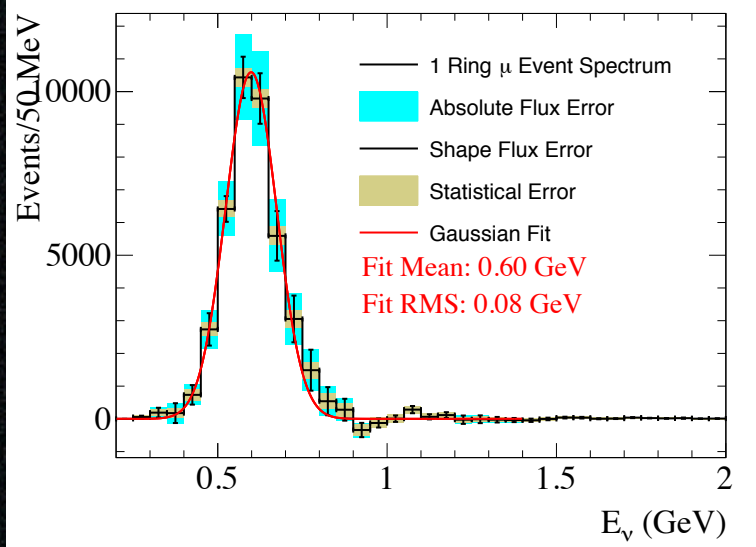
ν -Beam

-0.5 *

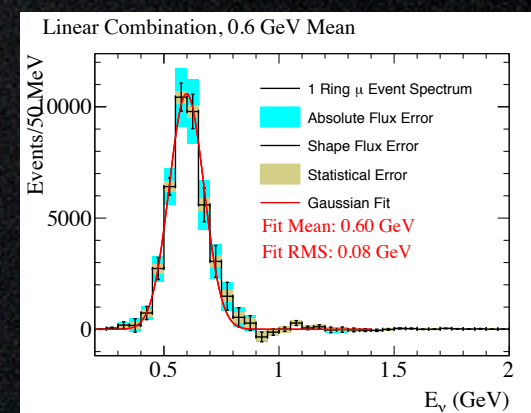
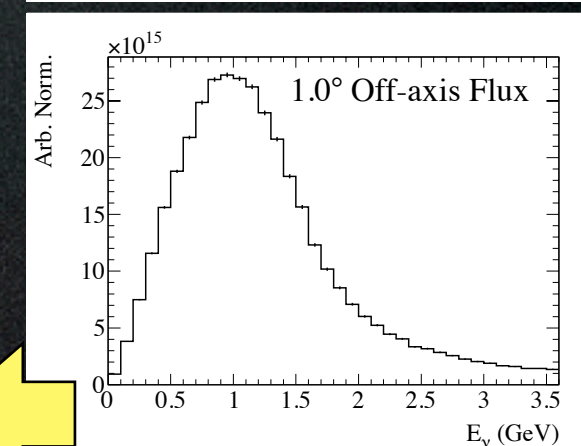
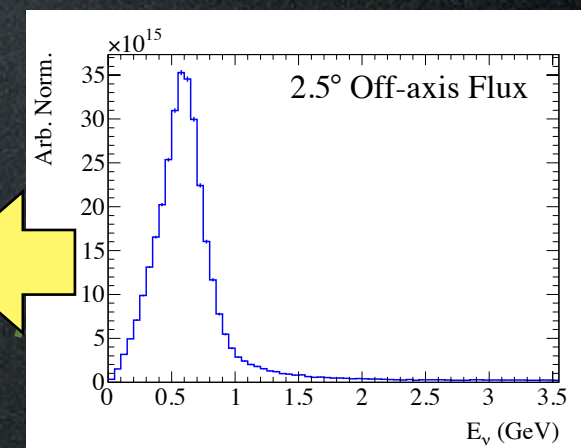
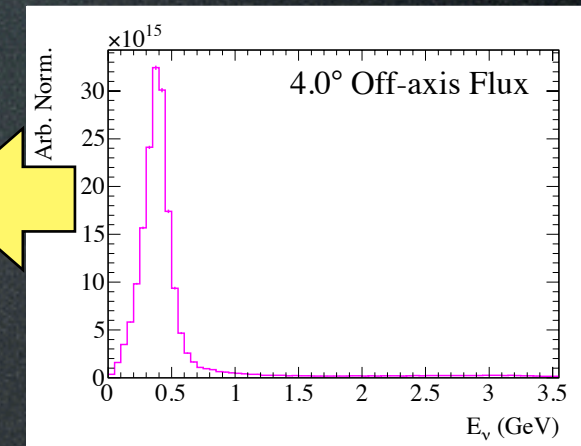
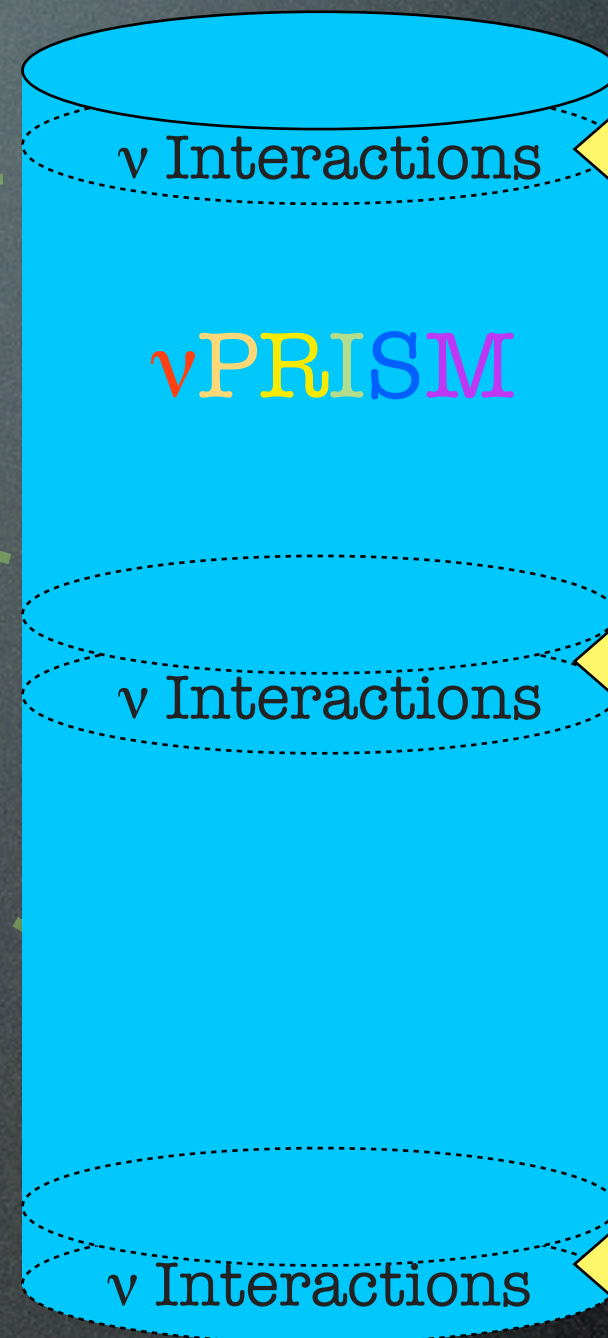
+1.0 *

Take linear combinations!

600 MeV Monoenergetic Beam
using 60 slices
in off-axis angle

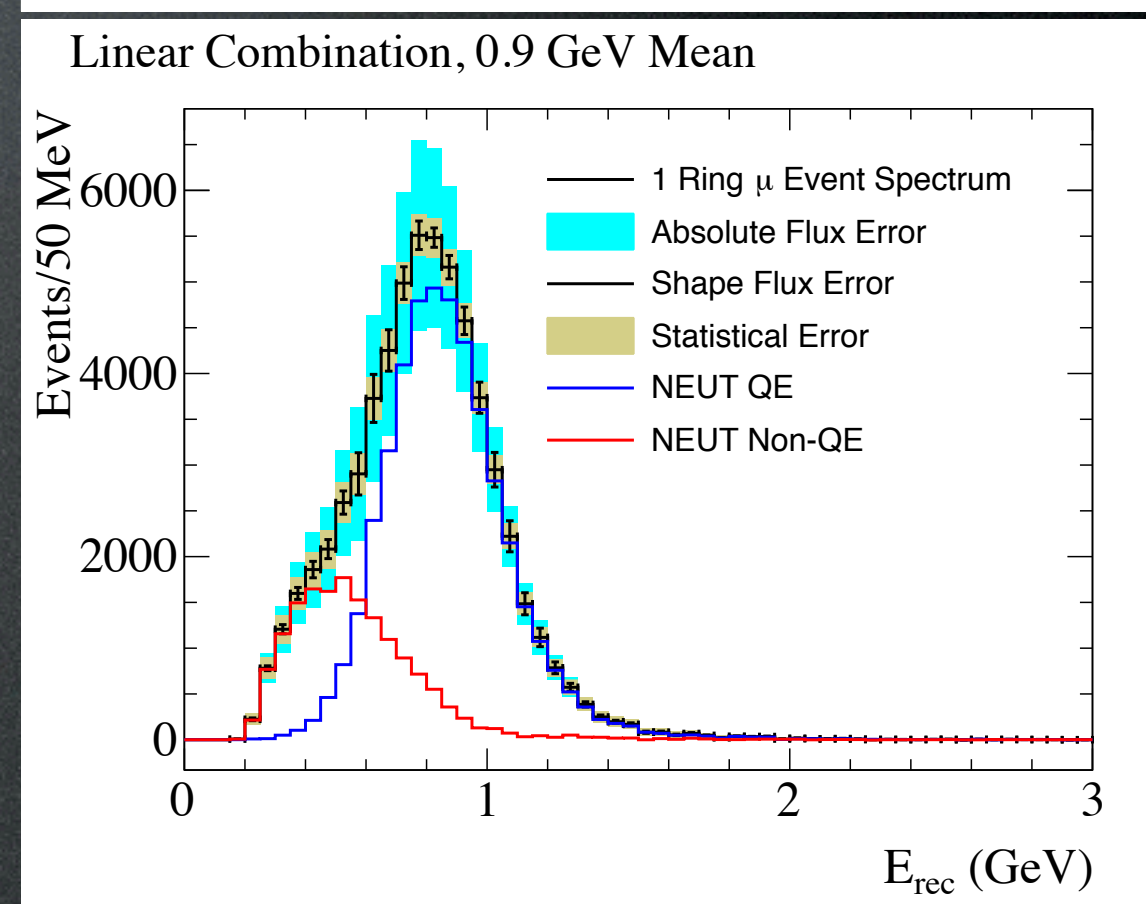
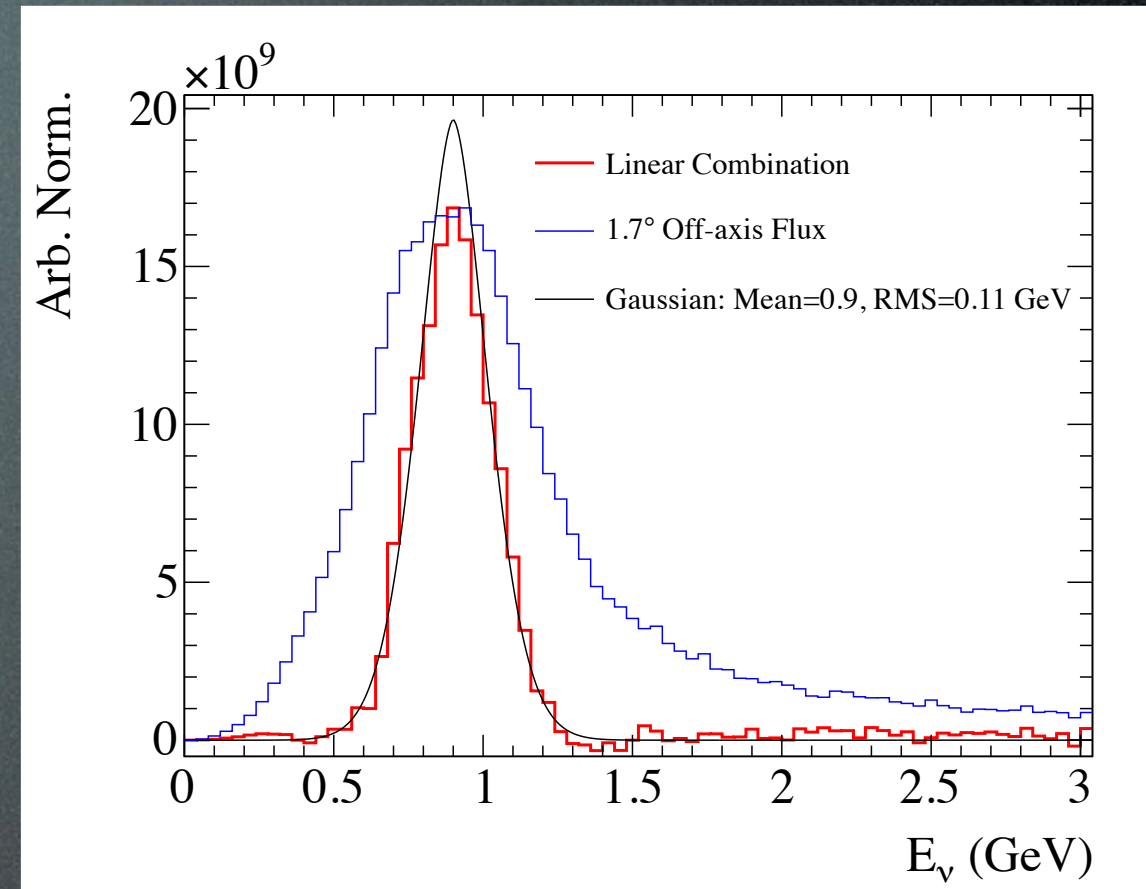


Muon $p \& \theta$
for 700 MeV
Monoenergetic
 ν -Beam!



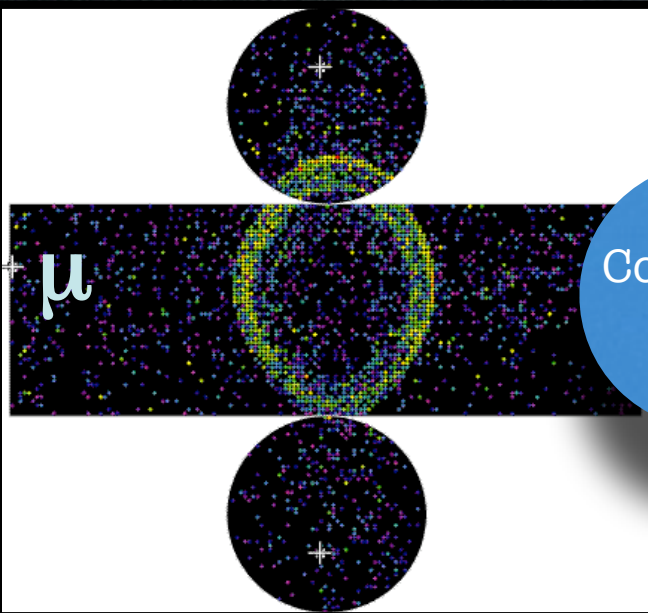
Benefits of a Monoenergetic Beam

- First ever measurements of NC events with E_ν
 - Much better constraints on NC oscillation backgrounds
- First ever “correct” measurements of CC events with E_ν
 - No longer rely on final state particles to determine E_ν
- It is now possible to separate the various components of single- μ events!
- This is also very interesting to the nuclear physics community



How We Typically Perform Oscillation Analyses

Observed far detector signal:
1-ring muon events



Composed Of:

Nuclear model

CCQE: $\mu^- + p$
(p unobserved)

CC π^+ : $\mu^- + N + \pi^+$
(p, π^+ unobserved)

CCDIS: $\mu^- + X$
(X unobserved)

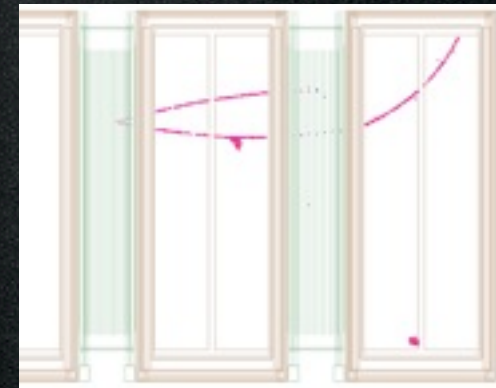
NC π^+ : $\pi^+ + n$
**(π^+ misidentified,
n unobserved)**

...

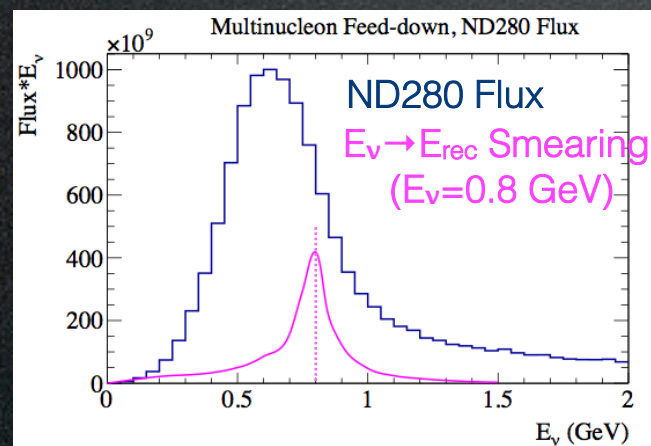
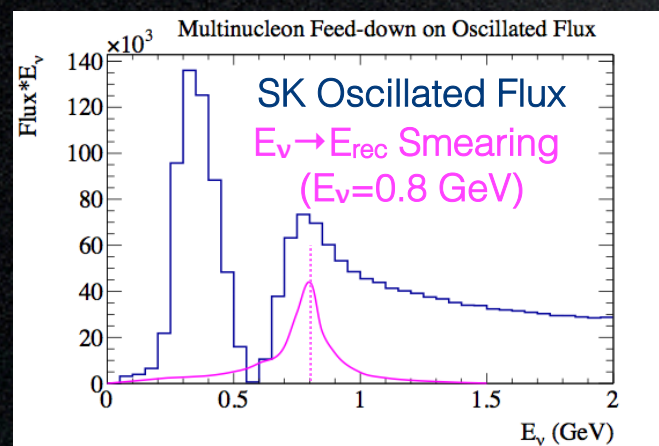
**Predicted by
poorly understood
models**

Parameter	E_ν Range	Nominal	Error	Class
M_A^{QE}	all	1.21 GeV/ c^2	0.45	shape
M_A^{RES}	all	1.41 GeV/ c^2	0.11	shape
p_F^{12C}	all	217 MeV/ c	30	shape
E_B^{12C}	all	25 MeV	9	shape
SF 12C	all	0 (off)	1 (on)	shape
CC Other shape ND280	all	0.0	0.40	shape
Pion-less Δ Decay	all	0.0	0.2	shape
CCQE E1	$0 < E_\nu < 1.5$	1.0	0.11	norm
CCQE E2	$1.5 < E_\nu < 3.5$	1.0	0.30	norm
CCQE E3	$E_\nu > 3.5$	1.0	0.30	norm
CC1 π E1	$0 < E_\nu < 2.5$	1.15	0.43	norm
CC1 π E2	$E_\nu > 2.5$	1.0	0.40	norm
CC Coh	all	1.0	1.0	norm
NC1 π^0	all	0.96	0.43	norm
NC 1 π^\pm	all	1.0	0.3	norm
NC Coh	all	1.0	0.3	norm
NC other	all	1.0	0.30	norm
ν_μ/ν_e	all	1.0	0.03	norm
$\nu/\bar{\nu}$	all	1.0	0.40	norm

**Simultaneously
constrain flux
and cross section
parameters with
a near detector**



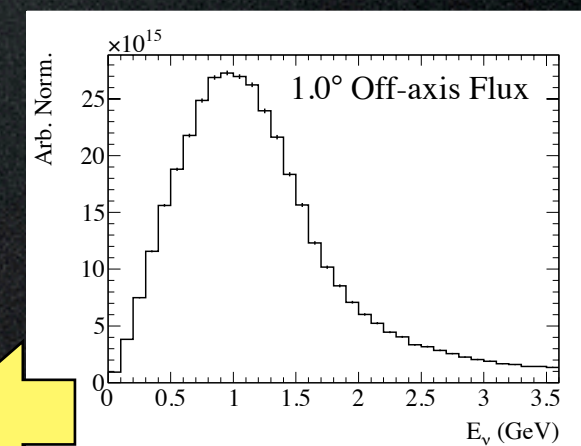
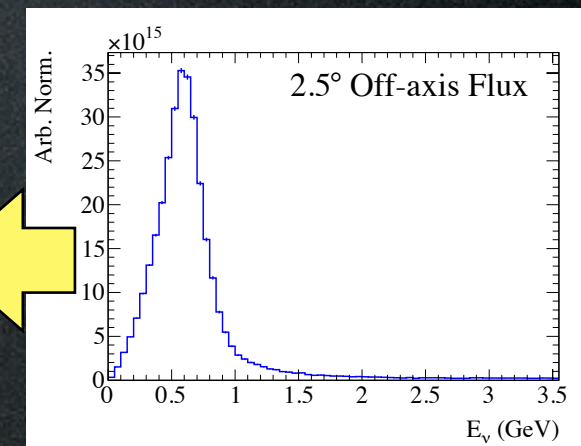
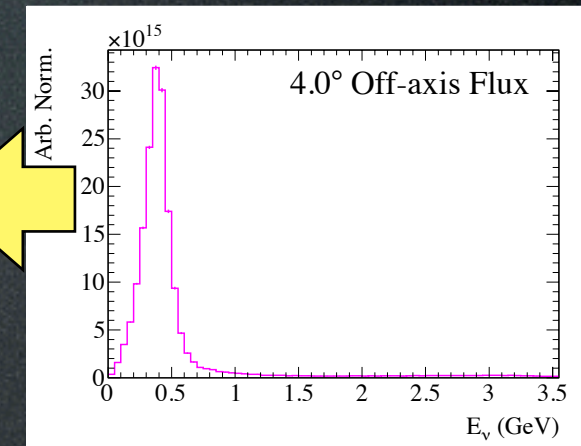
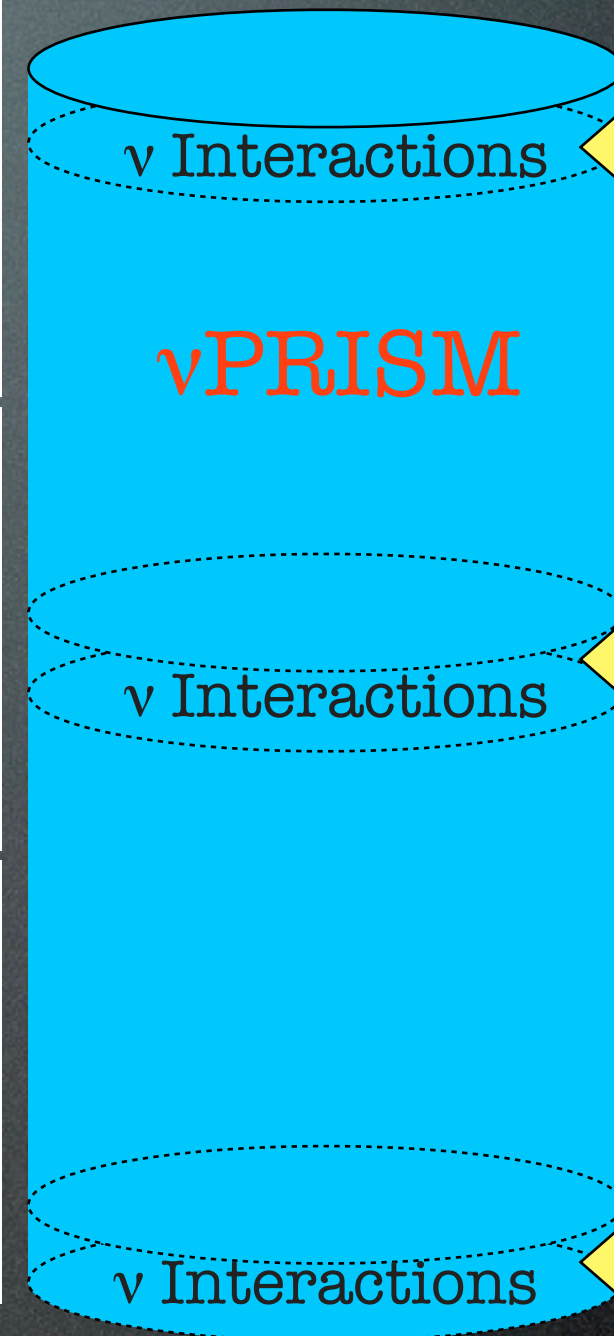
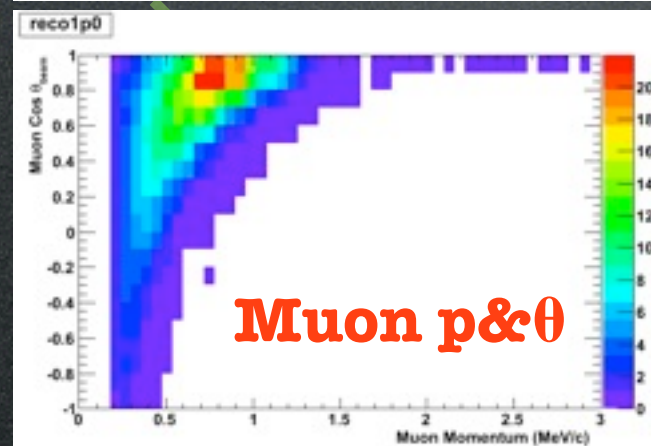
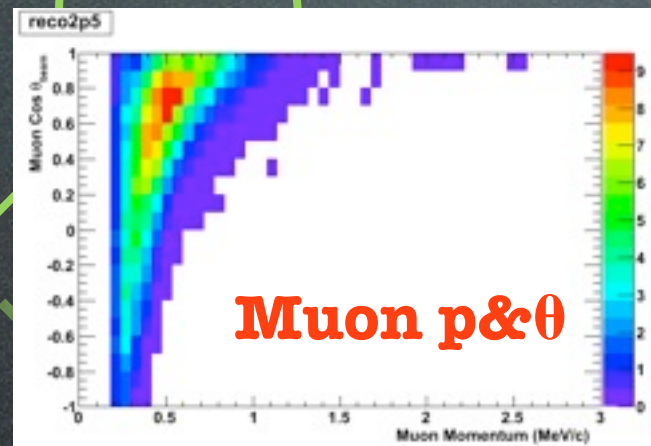
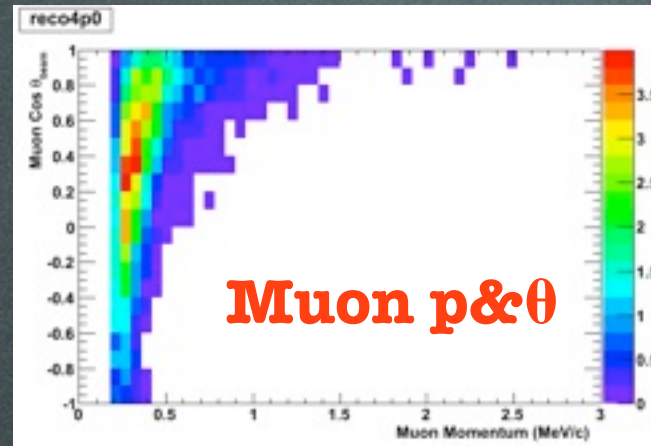
**But the near
and far fluxes
are different!**



**Goal of NuPRISM is to replace
this procedure with a data
measurement (to first order)**

NuPRISM in Oscillation Analyses

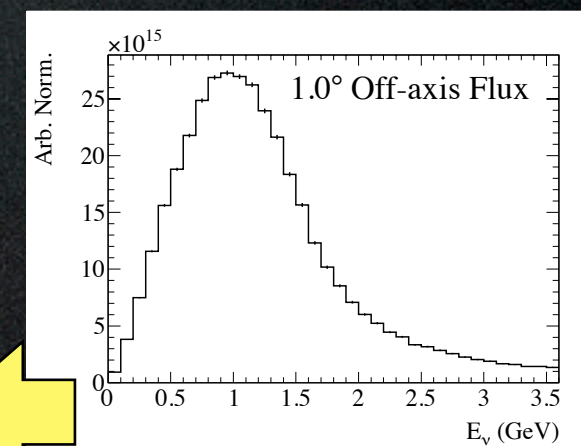
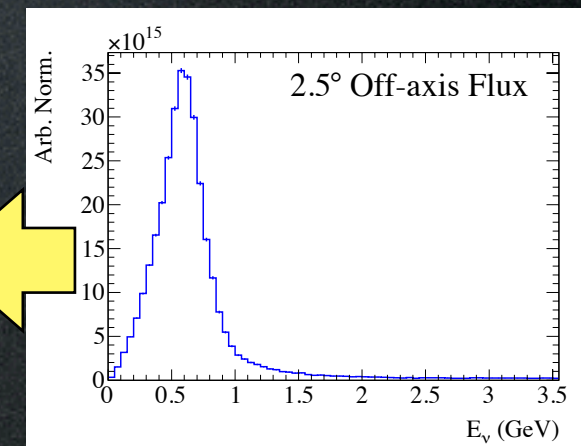
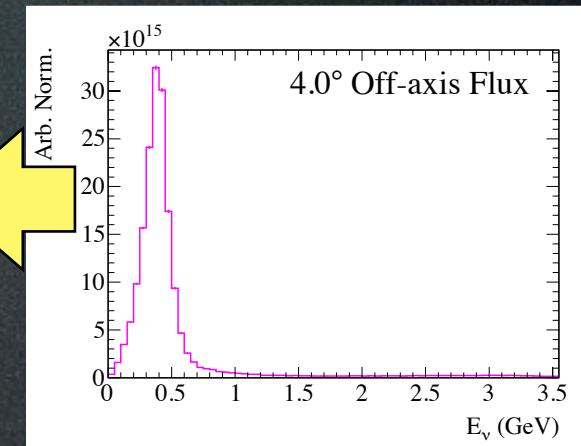
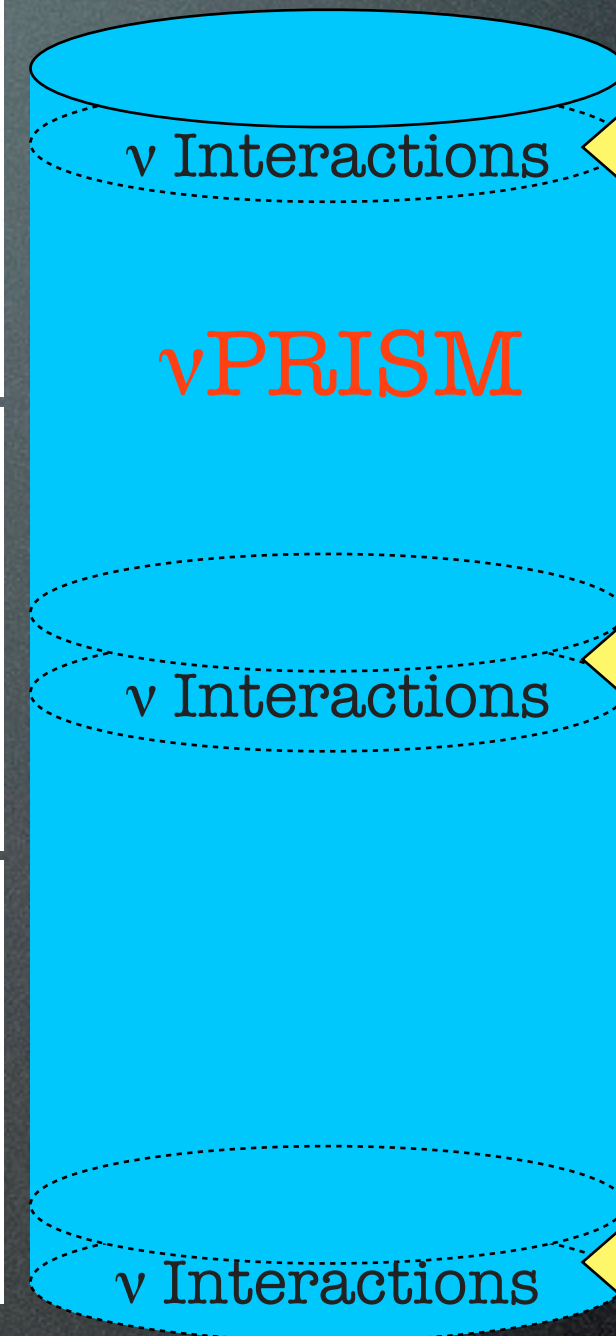
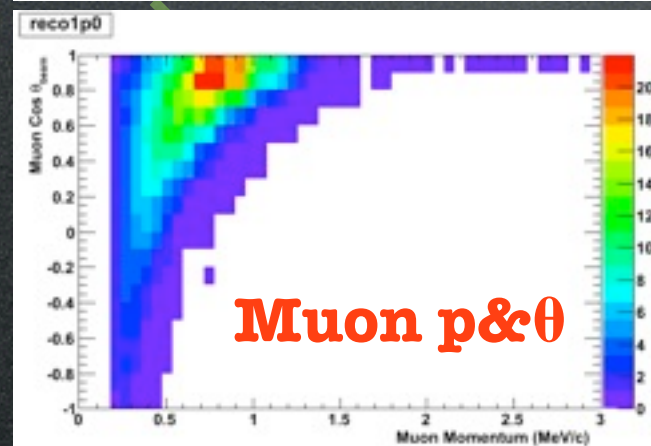
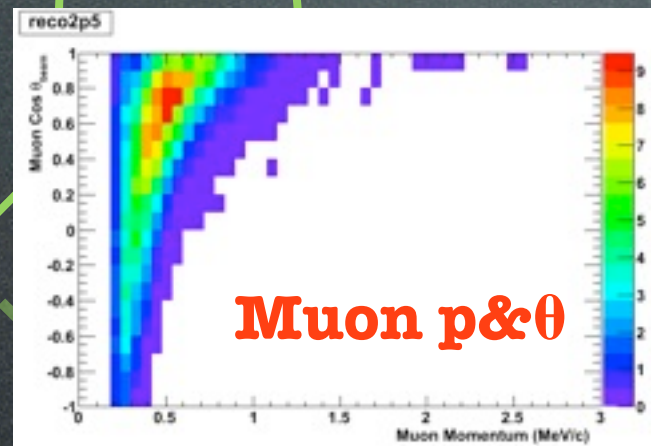
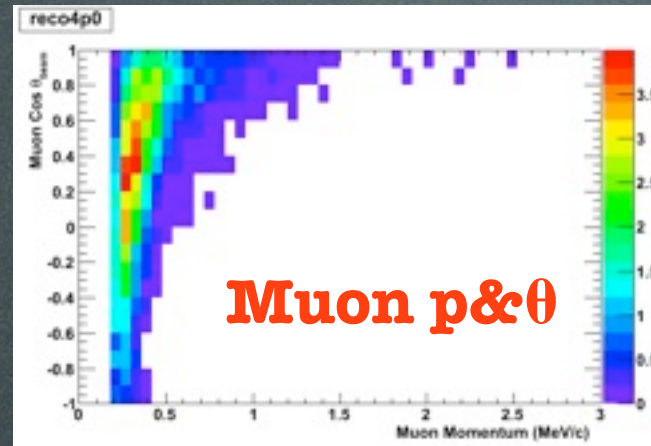
ν -Beam



NuPRISM in Oscillation Analyses

ν -Beam

Take different
linear
combinations!

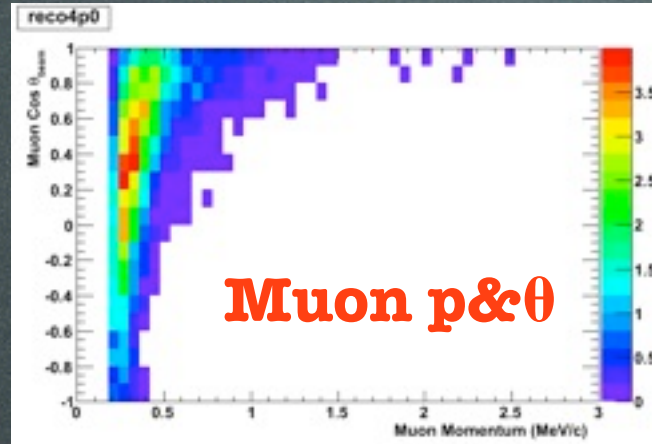


NuPRISM in Oscillation Analyses

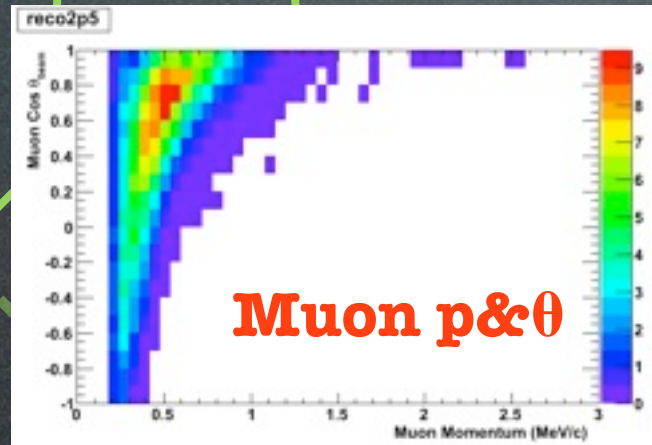
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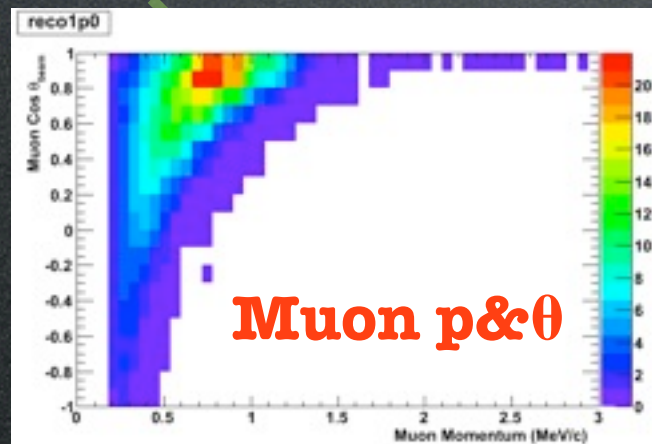
+1.0*



-0.8*



+0.2*

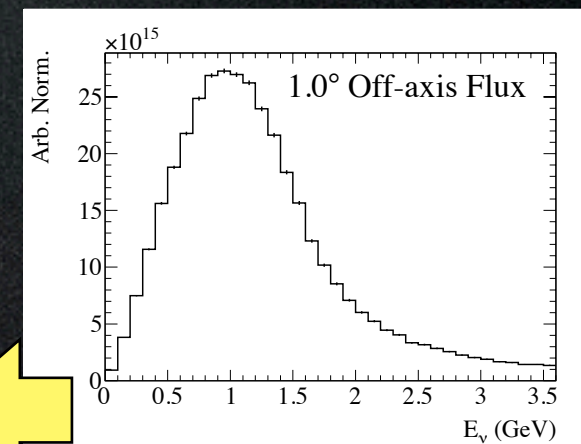
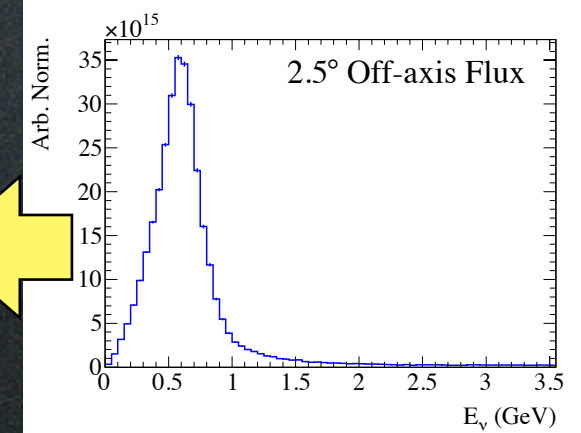
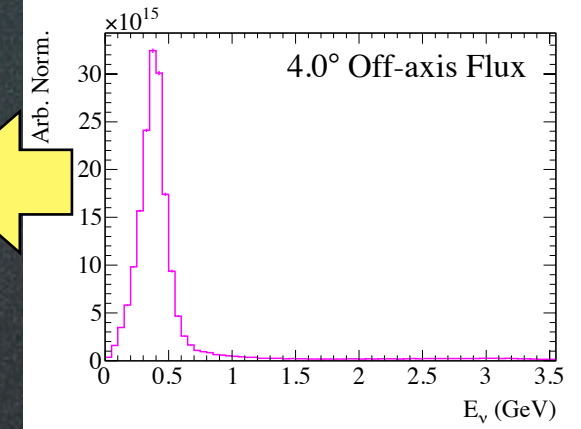


ν Interactions

ν PRISM

ν Interactions

ν Interactions



NuPRISM in Oscillation Analyses

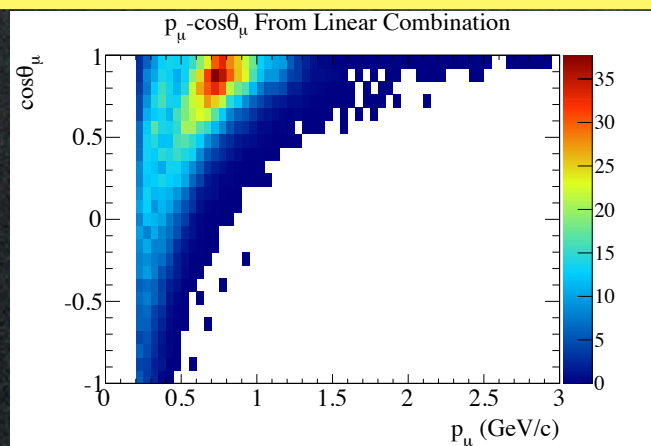
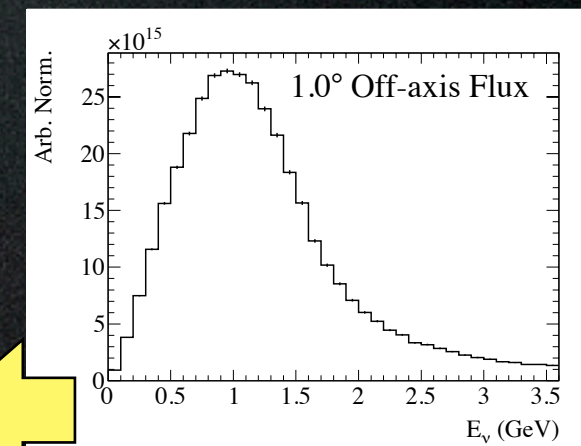
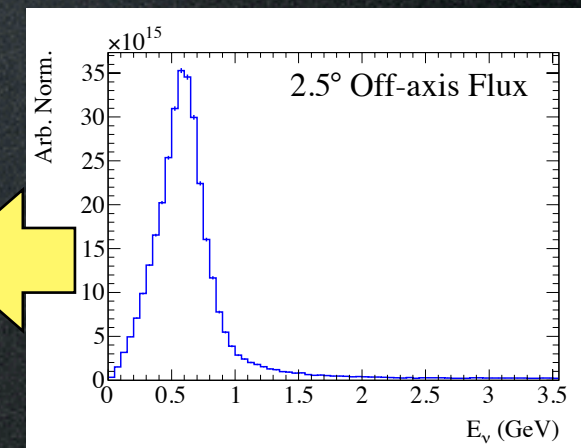
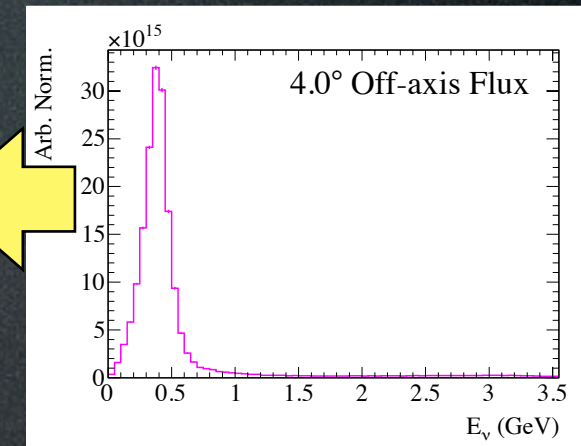
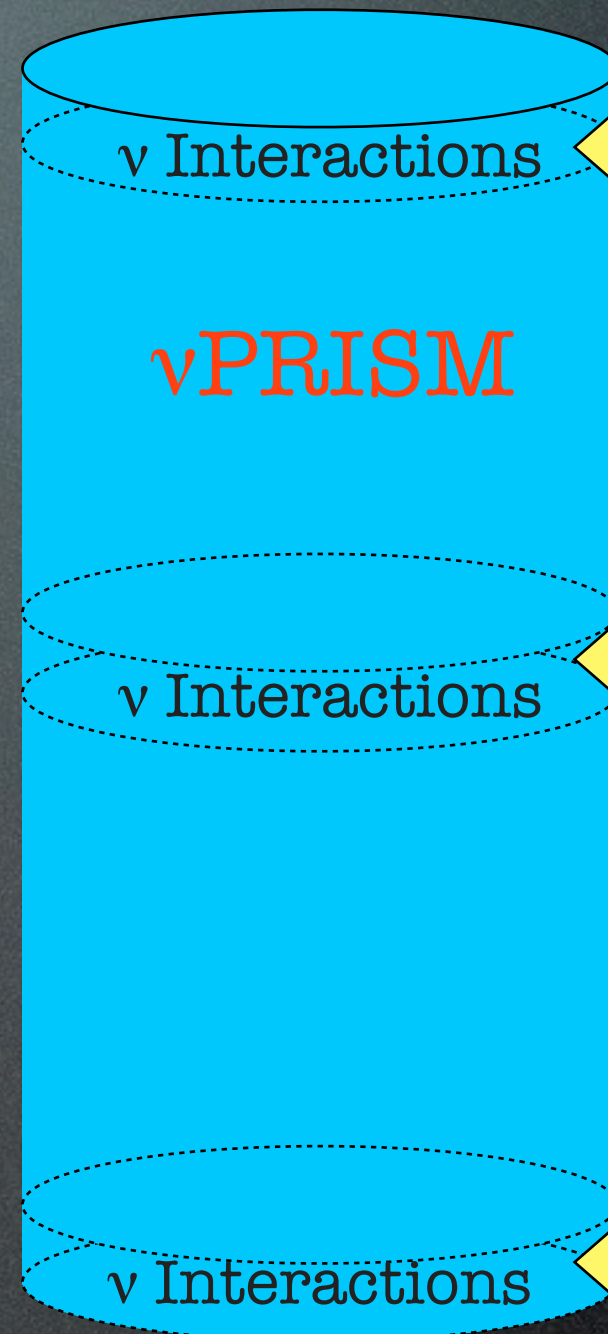
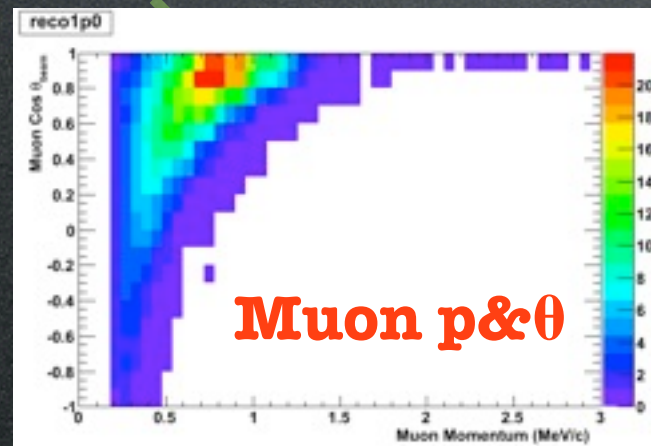
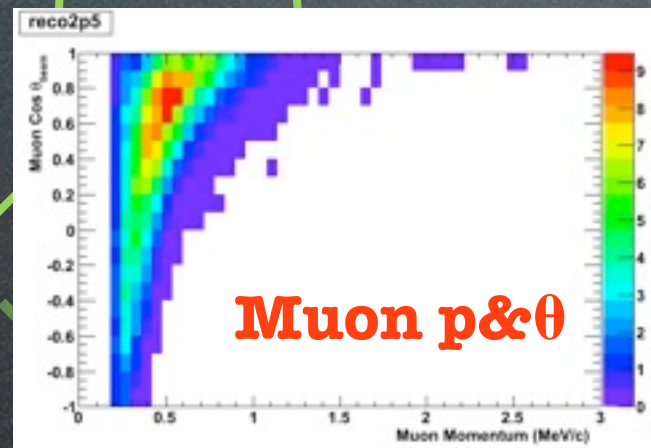
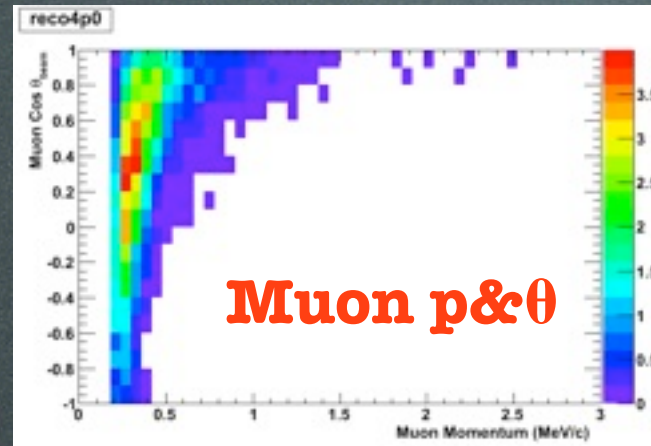
ν -Beam

Take different
linear
combinations!

+1.0*

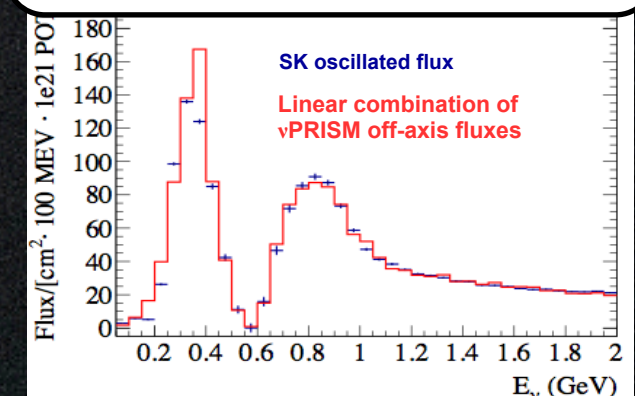
-0.8*

+0.2*



Measured!
Muon $p \& \theta$
for Oscillated
SK Flux!

Match Super-K Oscillated Flux



NuPRISM in Oscillation Analyses

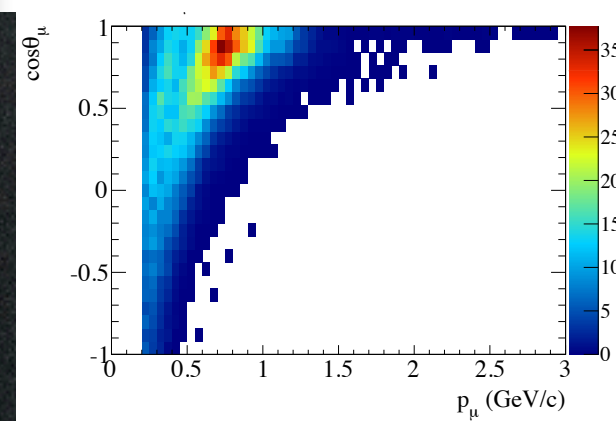
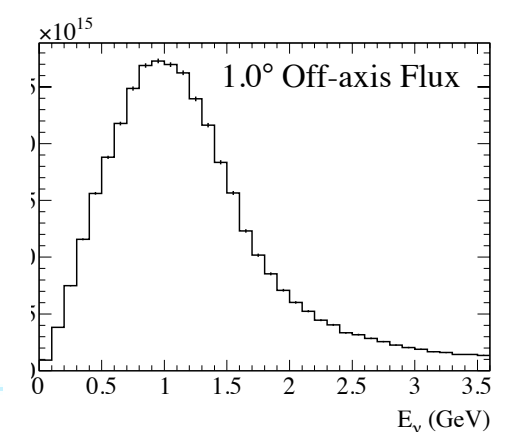
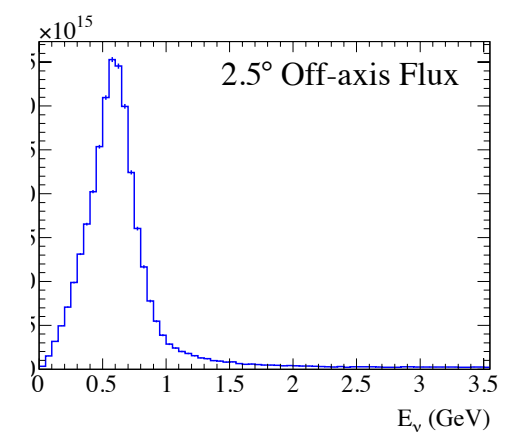
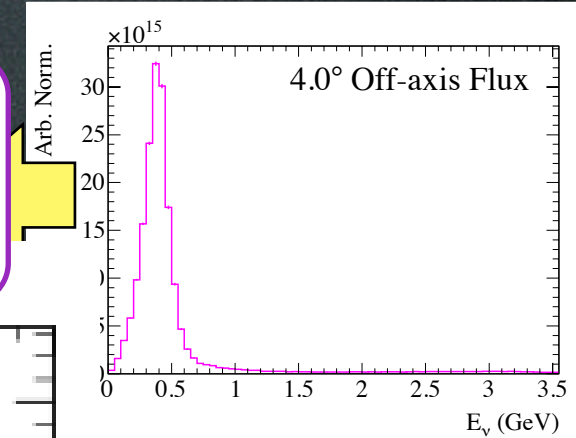
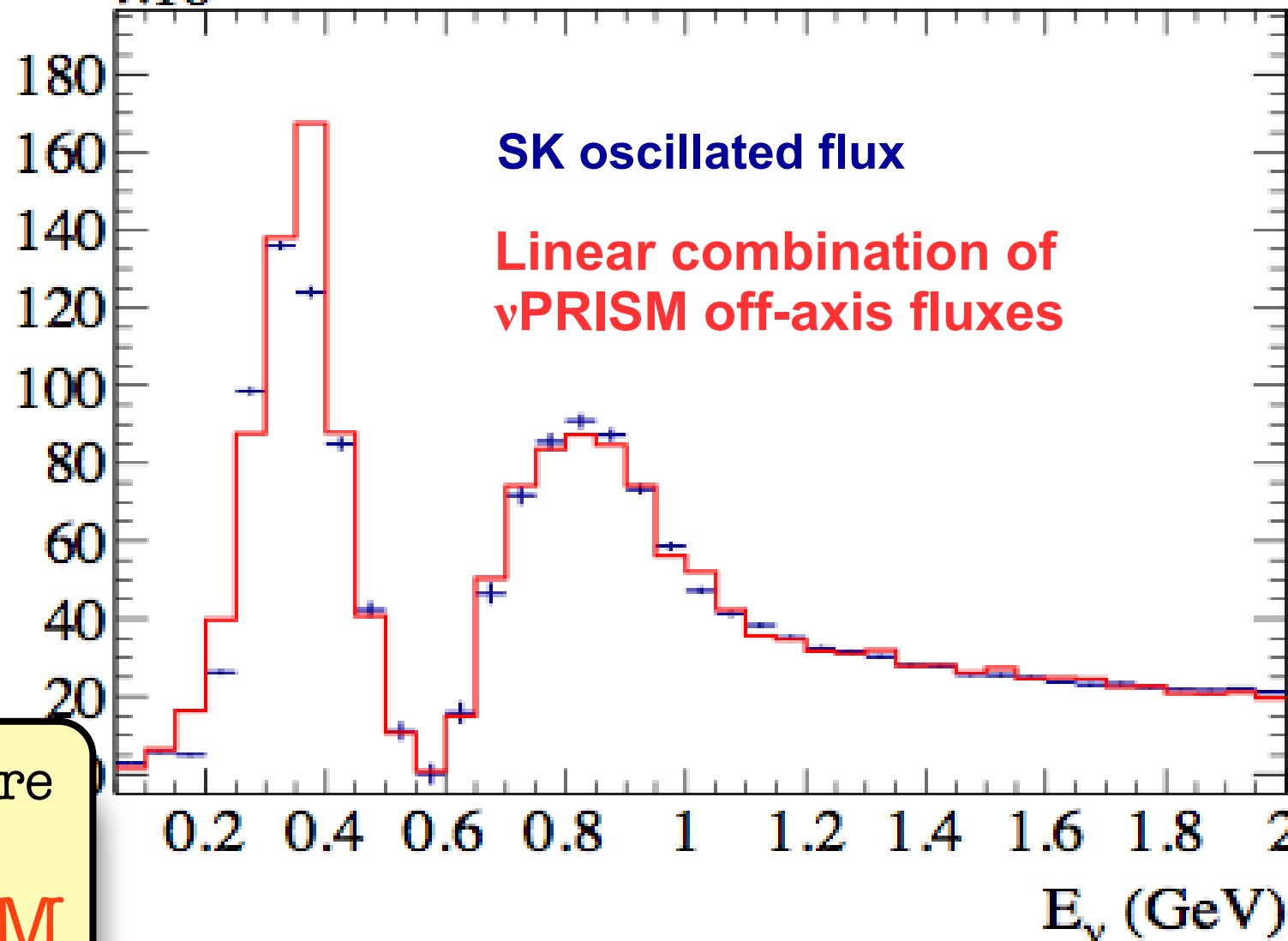
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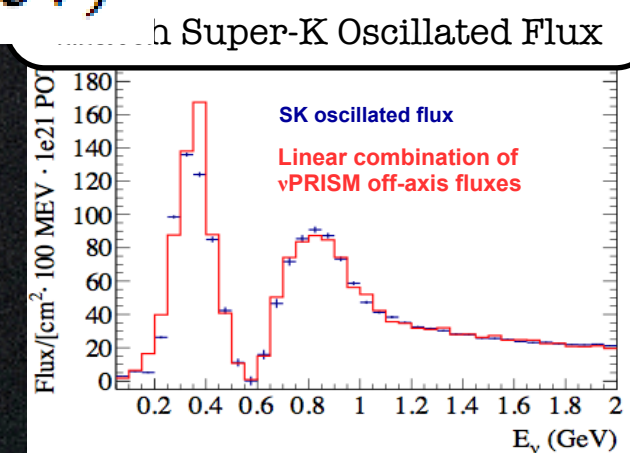
This is the procedure
used for the
T2K/nuPRISM
 ν_μ disappearance
analysis
(Details in the Next
Talk by Mark Scott)

**Reproduce Super-K Oscillation
Pattern at a Near Detector!**

$\times 10^{15}$
Flux/[cm² · 100 MEV · 1e21 POT]



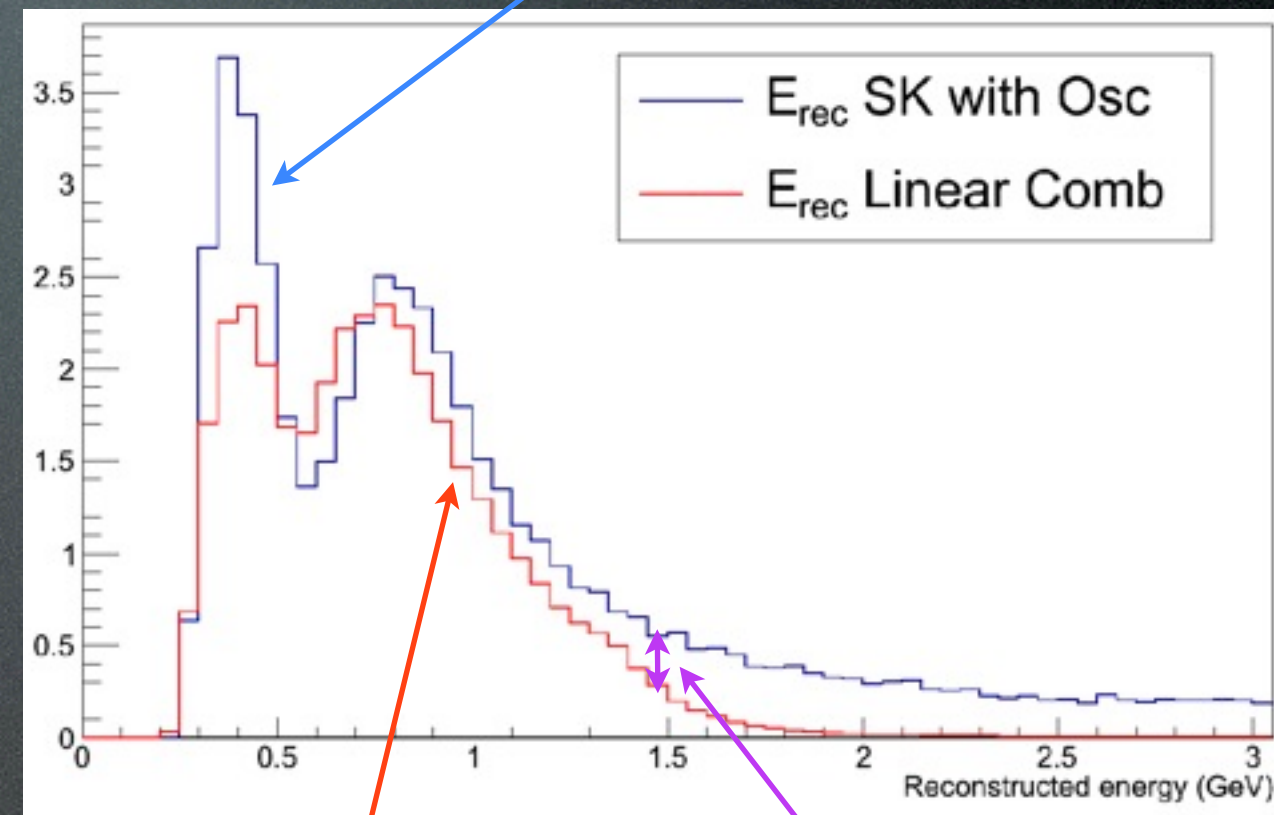
Measured!
Muon p& θ
for Oscillated
SK Flux!



E_{rec} Distribution

- **For now, collapse 2D muon p, θ distribution into 1D E_{rec} plot**
- **Notice the NuPRISM and SK distributions disagree**
 - If they didn't, we would have no cross section systematic errors (modulo variations in the flux)
 - Differences are from detector acceptance & resolution, and imperfect flux fit
- **Super-K prediction is largely based on the directly-measured NuPRISM muon kinematics!**
 - Now, only a small amount of model extrapolation is needed
 - T2K measurements are now largely independent of cross section modeling!

Previously, the entire predicted E_{rec} distribution at Super-K was based on model extrapolation

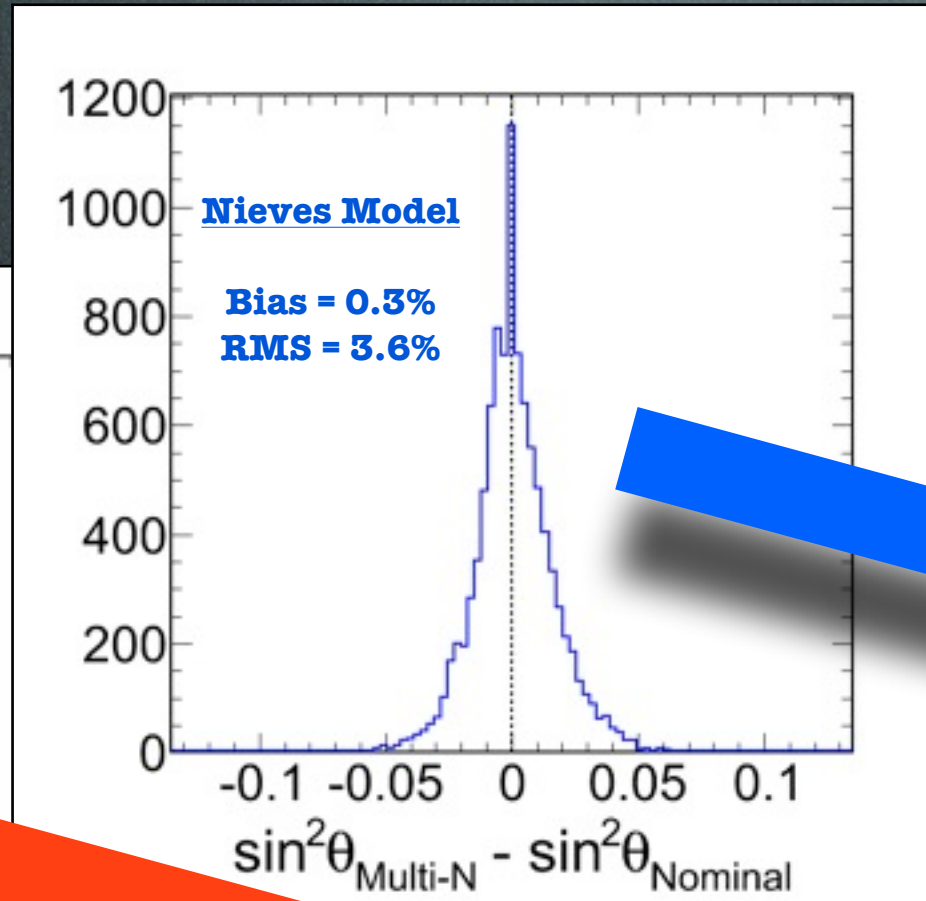
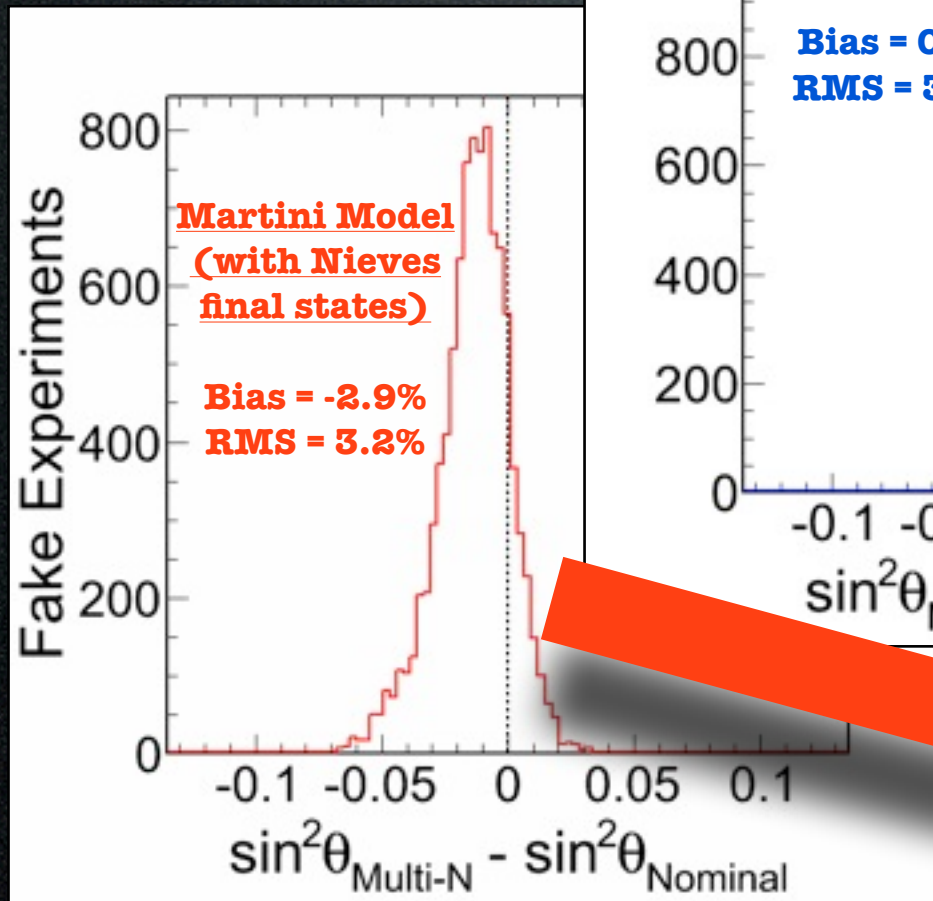


Now, NuPRISM directly measures most of this distribution

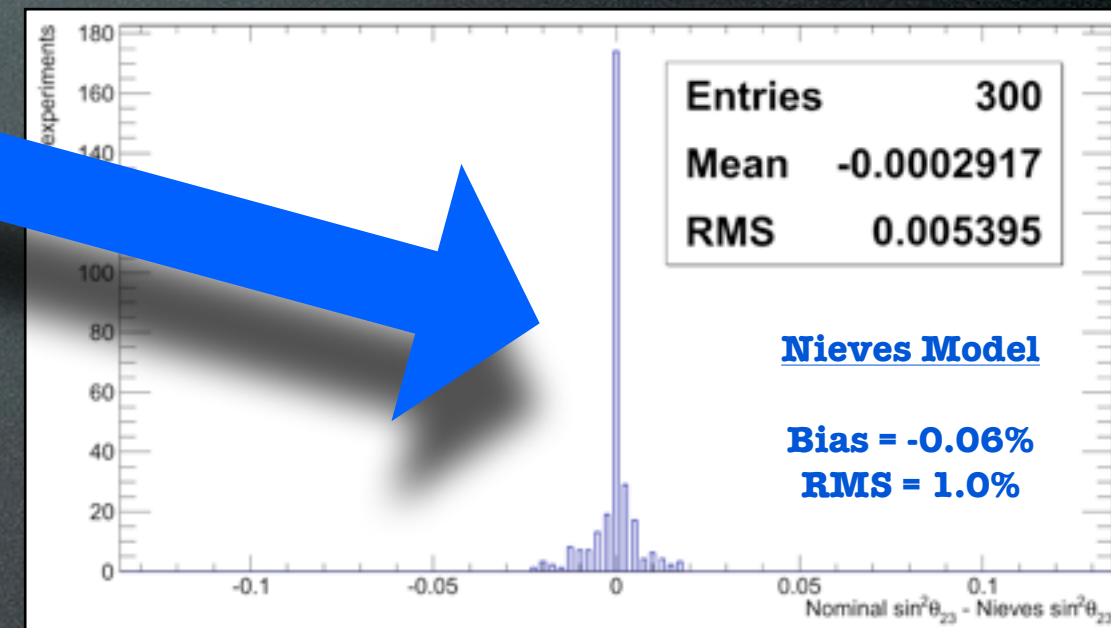
The remaining model-dependent correction factor (i.e. systematic uncertainty) is relatively small

ν PRISM ν_μ Disappearance Constraint

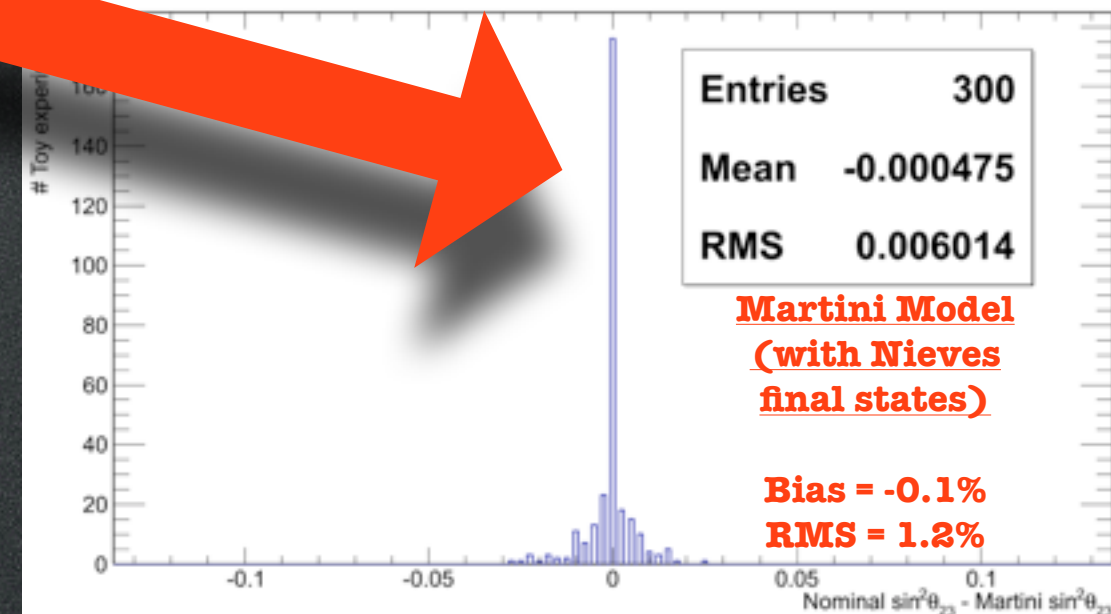
Standard T2K Analysis



ν PRISM Analysis



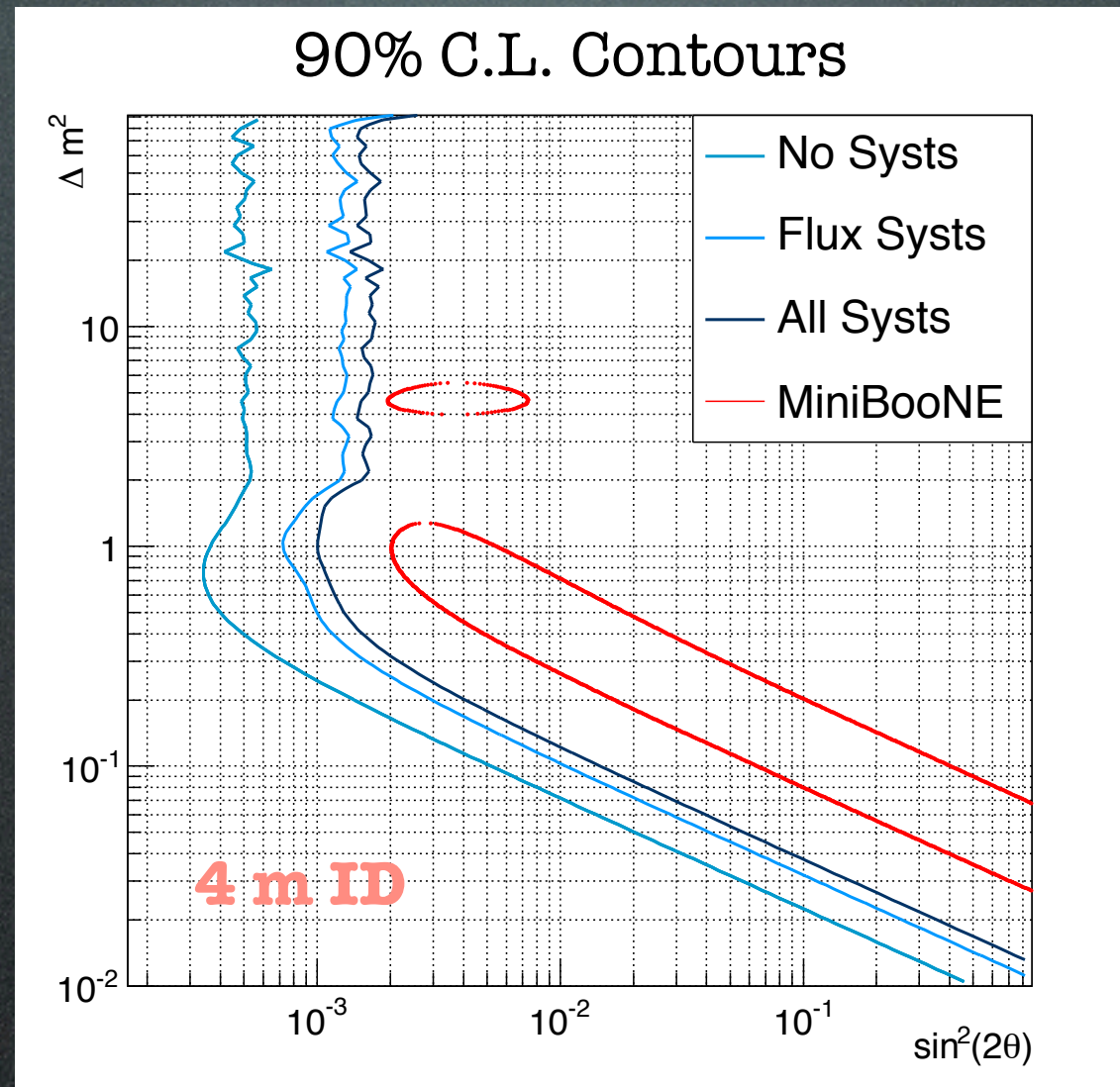
- Fake data studies show the bias in θ_{13} is reduced from **4.3%/3.6%** to **1.2%/1.0%**
- More importantly, this is now based on a **data constraint**, rather than a model-based guess
- Expect the NuPRISM constraints to get significantly better as additional constraints are implemented (very conservative errors)



Sterile Neutrino Analysis

- To compute first sensitivities, make several conservative assumptions
- **No constraint from the existing near detector (ND280)**
 - Eventually, a powerful 2-detector constraint will be incorporated
- **No constraints on background processes**
 - nuPRISM should provide control samples for all of the major backgrounds to impose strong data-driven constraints
- **No combined $\nu_\mu + \nu_e$ fit**
 - MiniBooNE results would not have been possible without normalizing the ν_e signal to the observed ν_μ spectrum
- **Assume Super-K detector efficiencies and resolutions**
 - nuPRISM has smaller phototubes, and should perform better closer to the wall (which is important, since the diameter is much smaller)
 - Significant increase in ν_e statistics is expected
- With such conservative assumptions, is a measurement still possible?

(Very) Conservative Sterile- ν Sensitivities

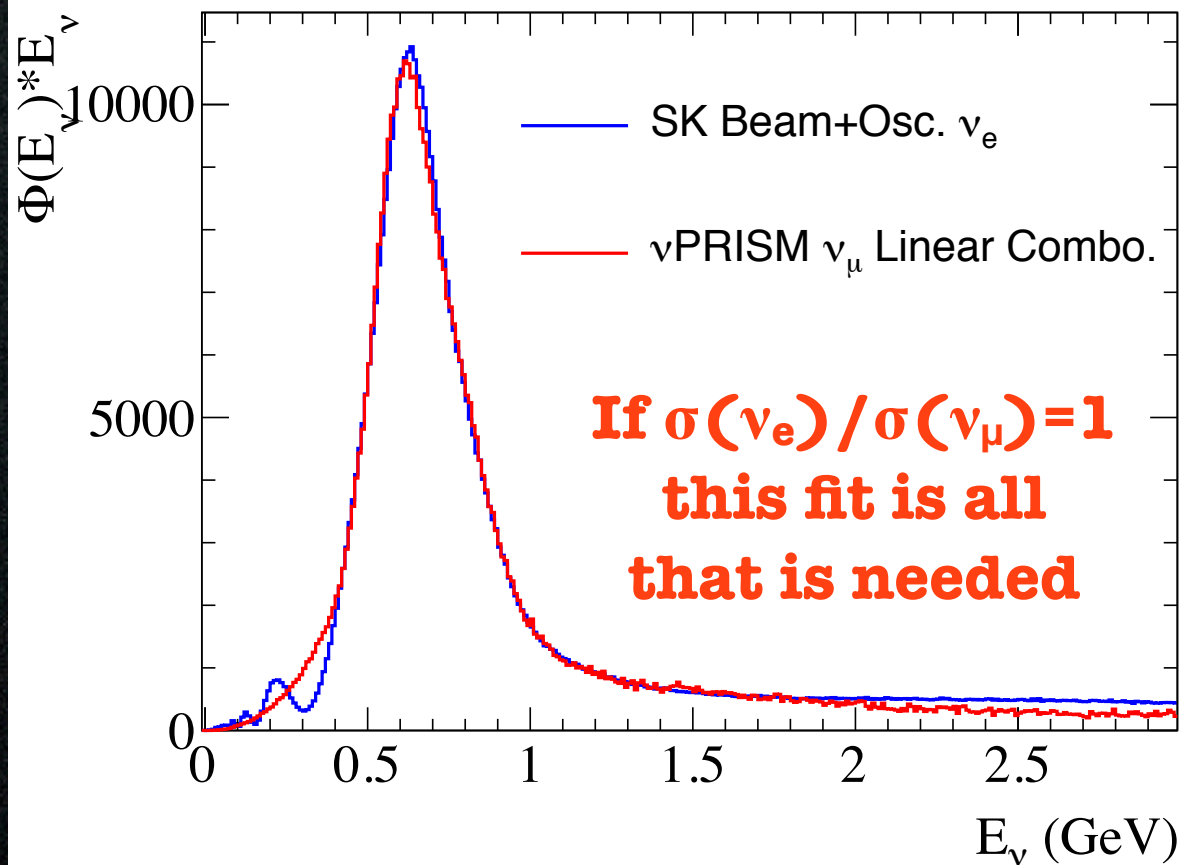


- Can already exclude currently allowed MiniBooNE regions at 90% C.L.
- Much better limits expected as the analysis improves

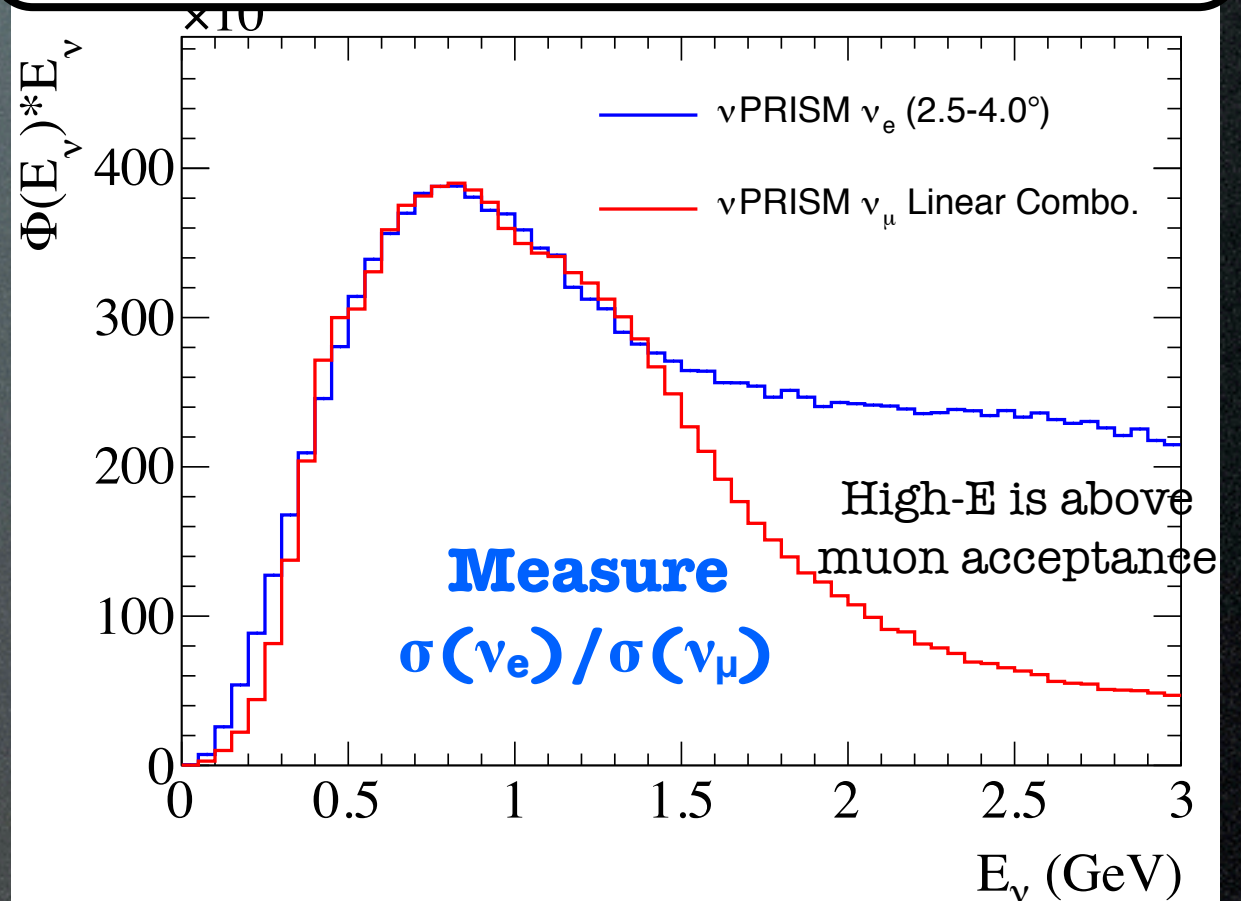
nuPRISM CPV (ν_e Appearance)

2 step approach:

Step 1: Measure **Super-K** ν_e response
with nuPRISM ν_μ



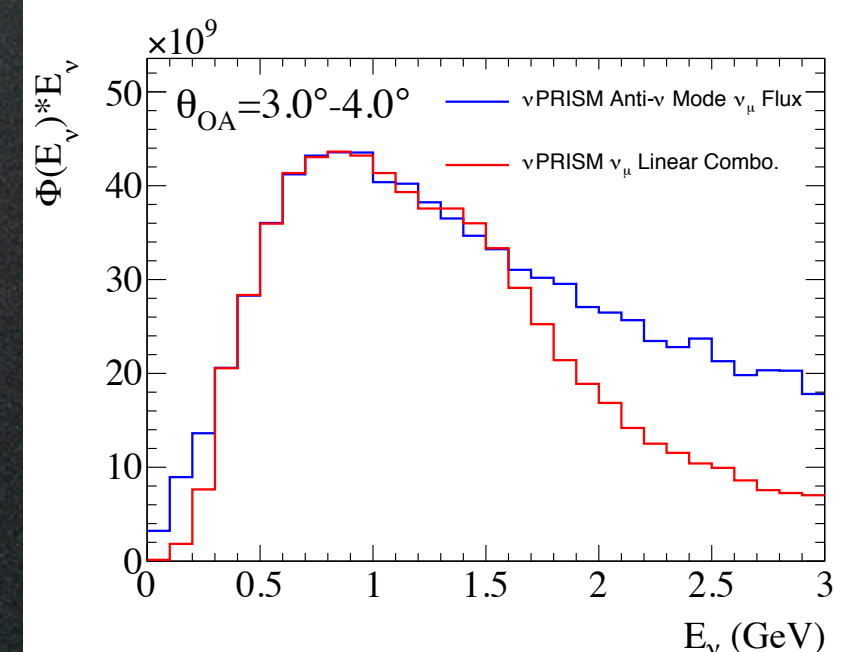
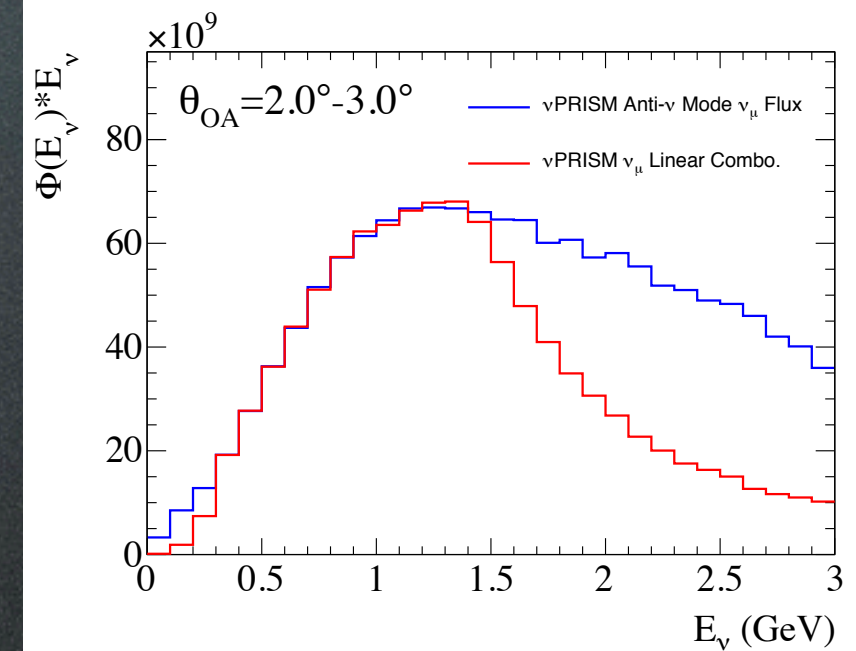
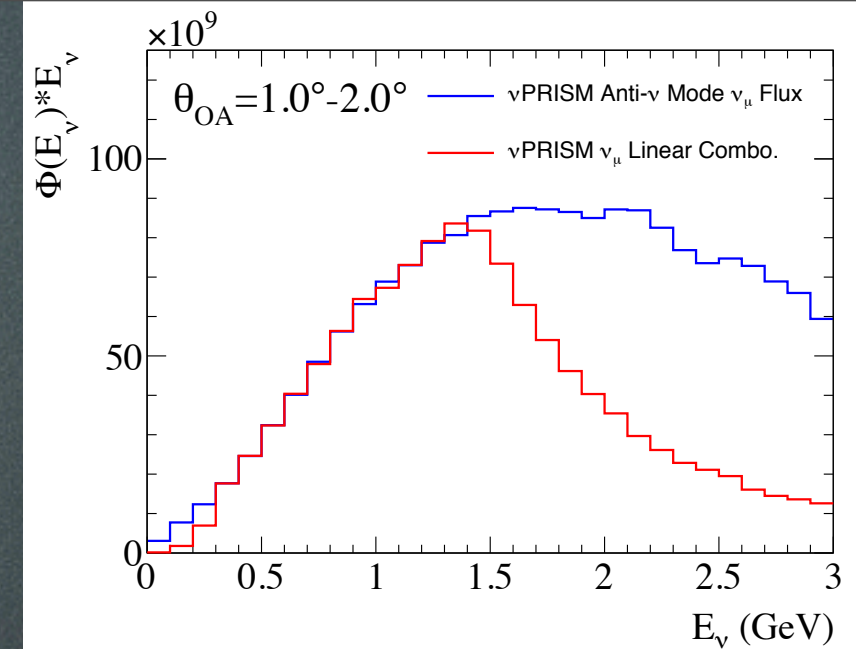
Step 2: Measure **nuPRISM** ν_e response
with nuPRISM ν_μ



- Step 1 is the ν_e version of the ν_μ disappearance analysis
- Step 2 uses only nuPRISM to measure $\sigma(\nu_e)/\sigma(\nu_\mu)$
 - High energy disagreement is above muon acceptance
 - These plots show flux * E_ν , so difference is 1-ring μ events is smaller

Anti-neutrinos

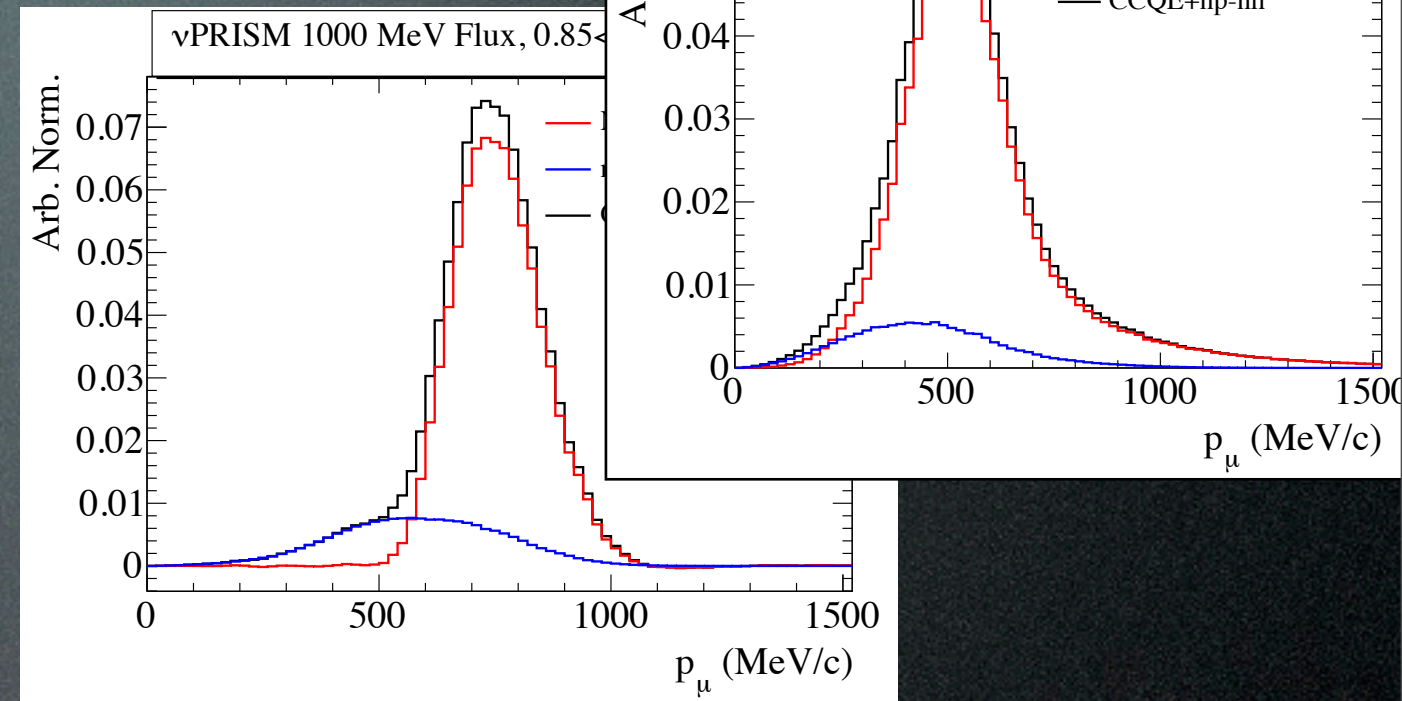
- T2K can switch between ν -mode and anti- ν -mode running by switching the beam focusing
- Anti- ν -mode analysis is the same as for neutrinos
 - Except with a much larger neutrino contamination
- Can use **ν -mode ν_μ data** to construct the **ν_μ background** in the **anti- ν -mode anti- ν_μ data**
 - Statistical separation of neutrinos from anti-neutrinos, rather than event-by-event sign selection
- After subtracting neutrino background, standard NuPRISM oscillation analyses can be applied to anti-neutrinos



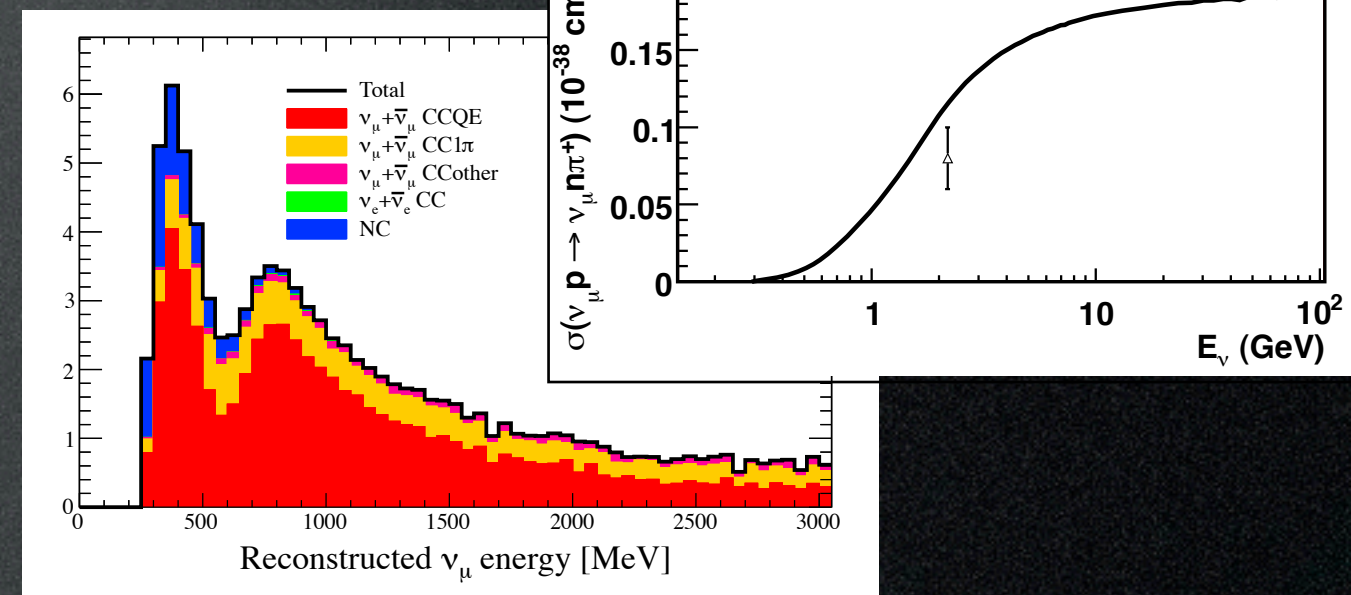
ν Cross Section Measurements

- Mono-energetic neutrino beams are ideal for measuring neutrino cross sections
 - Can provide a strong constraint on new models
- T2K ν_μ disappearance is subject to large $\text{NC}\pi^+$ uncertainties
 - 1 existing measurement
 - NuPRISM can place a strong constraint on this process vs E_ν

Example MEC event separation

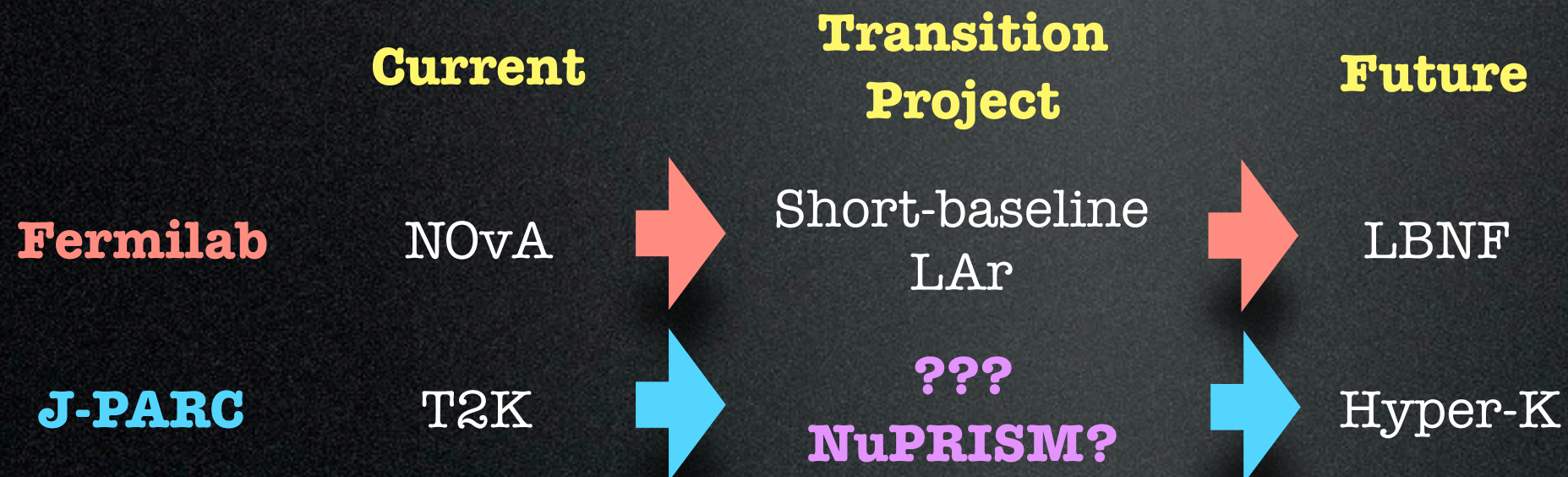
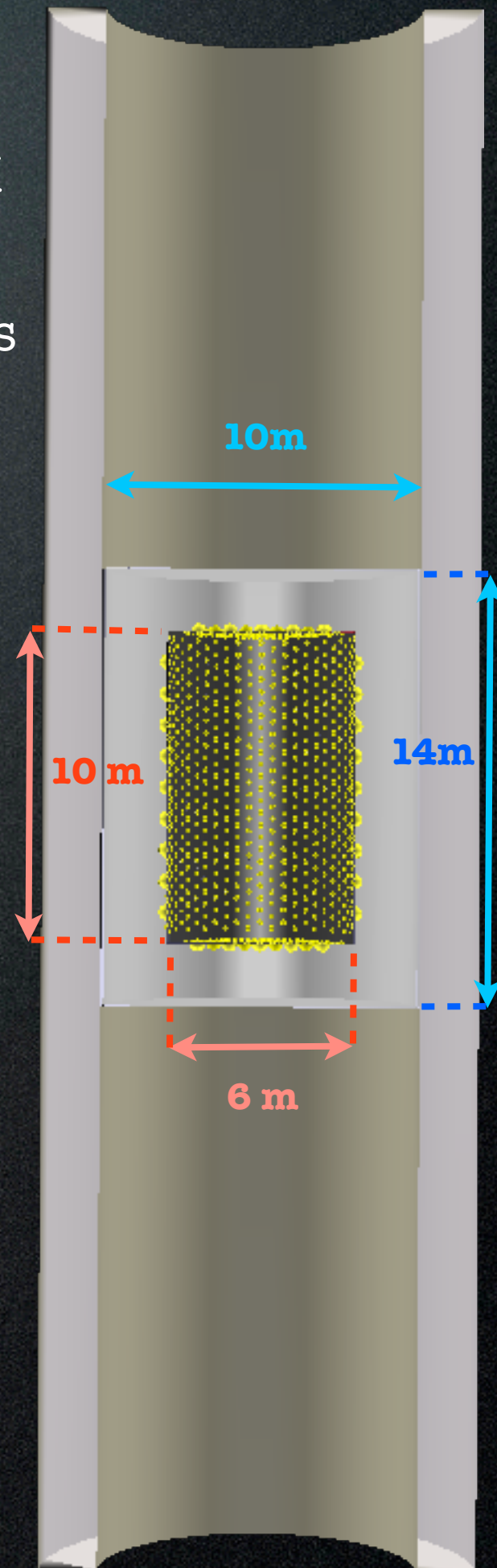


$\text{NC}\pi^+$ at T2K



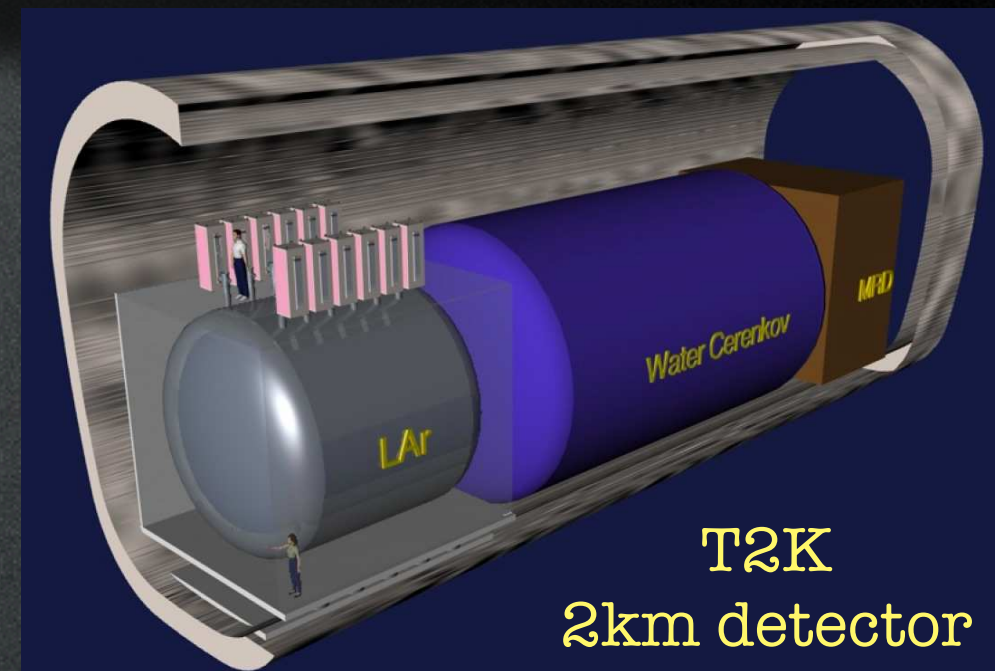
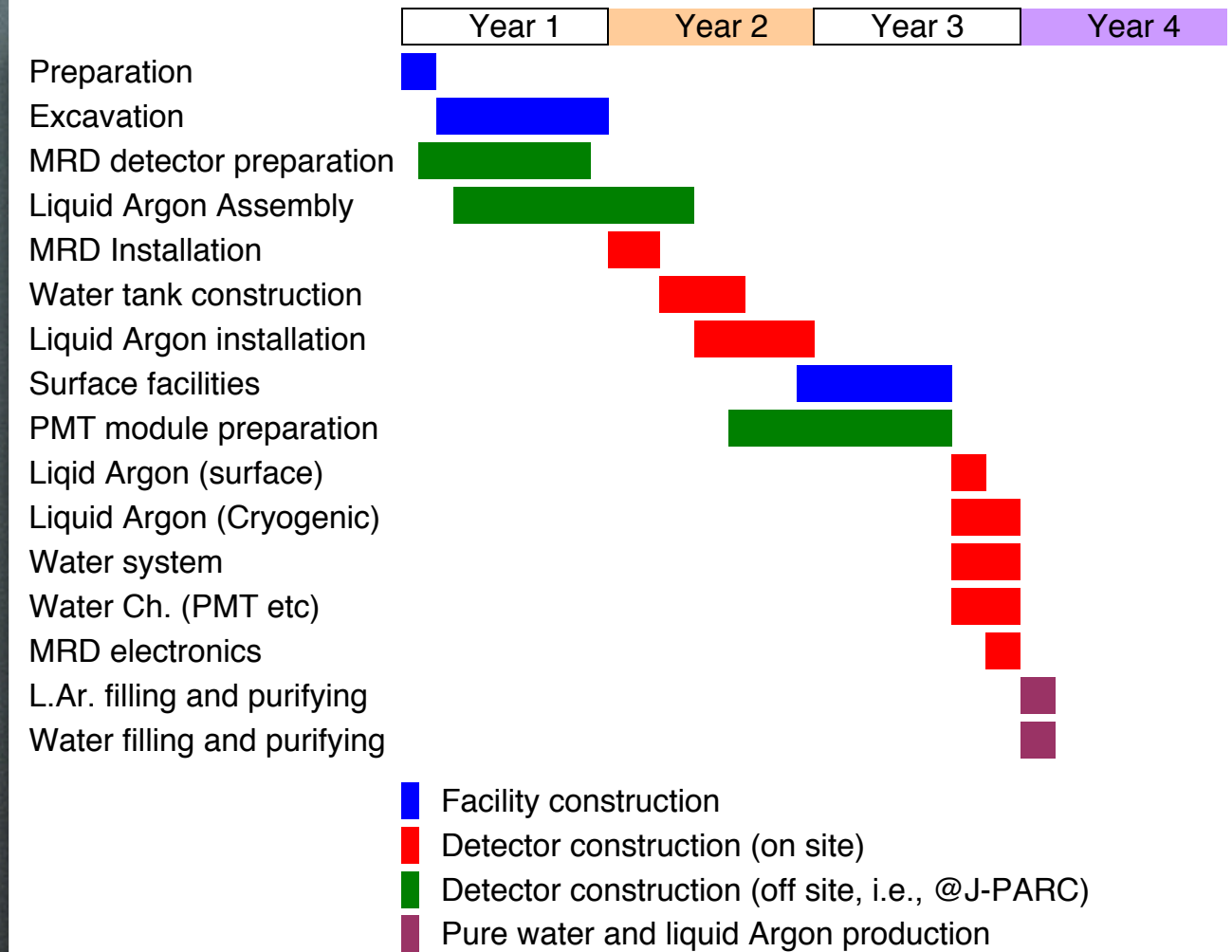
NuPRISM-Lite

- Goal is to construct the first NuPRISM detector during the T2K era
 - Moveable detector that samples full off axis range in 5 steps
 - After J-PARC beam upgrade (2018?) T2K will double its POT
- Provides an ideal environment for Hyper-K detector R&D
 - Detector can be lifted out of the water for maintenance or replacements
- Provides a mechanism to grow the Japanese neutrino physics community toward Hyper-K
 - Large, engaged, international user base will be needed



Timescales

- The T2K 2 km detector provides a
- NuPRISM construction time is faster
 - Same pit depth as the 2km detector, but no excavation of a large cavern at the bottom of the pit
 - Smaller instrumented volume
 - No LAr or MRD detector
- < 3 year timescale from approval to data taking
- Goal is to start data taking in time for the J-PARC 700kW beam (2018?)
 - Ideally, ground breaking would start in 2016



Current Status

- A Letter of Interest (LoI) was submitted to the J-PARC PAC in November 2014
 - arXiv:1412.3086
- Total cost is \$15-\$20M
 - Cheaper than Fermilab short-baseline program
 - Will need to spend this eventually to build a Hyper-K ND anyway
- Several sources of money already exist to build a WC detector for Hyper-K R&D
 - If timescales are compatible; this money can be used for NuPRISM
 - Even if initial testing is done elsewhere (e.g. EGADs), can transfer to NuPRISM later

Letter of Intent to Construct a nuPRISM Detector in the J-PARC Neutrino Beamline

S. Bhadra,²⁴ A. Blondel,³ S. Bordini,⁵ A. Bravar,³ C. Bronner,⁹ J. Caravaca Rodríguez,⁵ M. Dziewiecki,²³ T. Feusels,¹ G.A. Fiorentini Aguirre,²⁴ M. Friend,^{4,*} L. Haegel,³ M. Hartz,^{8,22} R. Henderson,²² T. Ishida,^{4,*} M. Ishitsuka,²⁰ C.K. Jung,^{11,†} A.C. Kaboth,⁶ H. Kakuno,²⁵ H. Kamano,¹³ A. Konaka,²² Y. Kudenko,^{7,‡} M. Kuze,²⁰ T. Lindner,²² K. Mahn,¹⁰ J.F. Martin,²¹ J. Marzec,²³ K.S. McFarland,¹⁵ S. Nakayama,^{18,†} T. Nakaya,^{9,8} S. Nakamura,¹² Y. Nishimura,¹⁹ A. Rychter,²³ F. Sánchez,⁵ T. Sato,¹² M. Scott,²² T. Sekiguchi,^{4,*} M. Shiozawa,^{18,8} T. Sumiyoshi,²⁵ R. Tacik,^{14,22} H.K. Tanaka,^{18,†} H.A. Tanaka,^{1,§} S. Tobayama,¹ M. Vagins,^{8,2} J. Vo,⁵ D. Wark,¹⁶ M.O. Wascko,⁶ M.J. Wilking,¹¹ S. Yen,²² M. Yokoyama,^{17,†} and M. Ziemicki²³

(The nuPRISM Collaboration)

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⁸Kavli Institute for the Physics and Mathematics of the Universe (WPI), Todai

Institutes for Advanced Study, University of Tokyo, Kashiwa, Chiba, Japan

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²²TRIUMF, Vancouver, British Columbia, Canada

²³Warsaw University of Technology, Institute of Radioelectronics, Warsaw, Poland

²⁴York University, Department of Physics and Astronomy, Toronto, Ontario, Canada

²⁵Tokyo Metropolitan University, Department of Physics, Tokyo, Japan

Next Steps

- Full proposal will be submitted to the J-PARC PAC, July 15-17
- Significant progress has been made in detector simulation (next talk)
- However, reconstruction is not yet available
 - Simple tuning of PMT QE and water attenuation was not sufficient for 8" PMTs
 - Full tuning of PMT pulse shape, angular acceptance, and time PDFs are now underway
 - Same procedure as for Hyper-K (see fitQun talk)
- Aiming for significant progress in physics analyses for the full proposal
 - Full ν_μ disappearance analysis with estimate of all systematic errors
 - Complete ν_e appearance analysis with CP violation constraint
 - Including anti-neutrinos if wrong-sign background constraint can be finished
- Planning a weeklong workshop in mid-March
 - Intensive week of analysis work (very few talks)

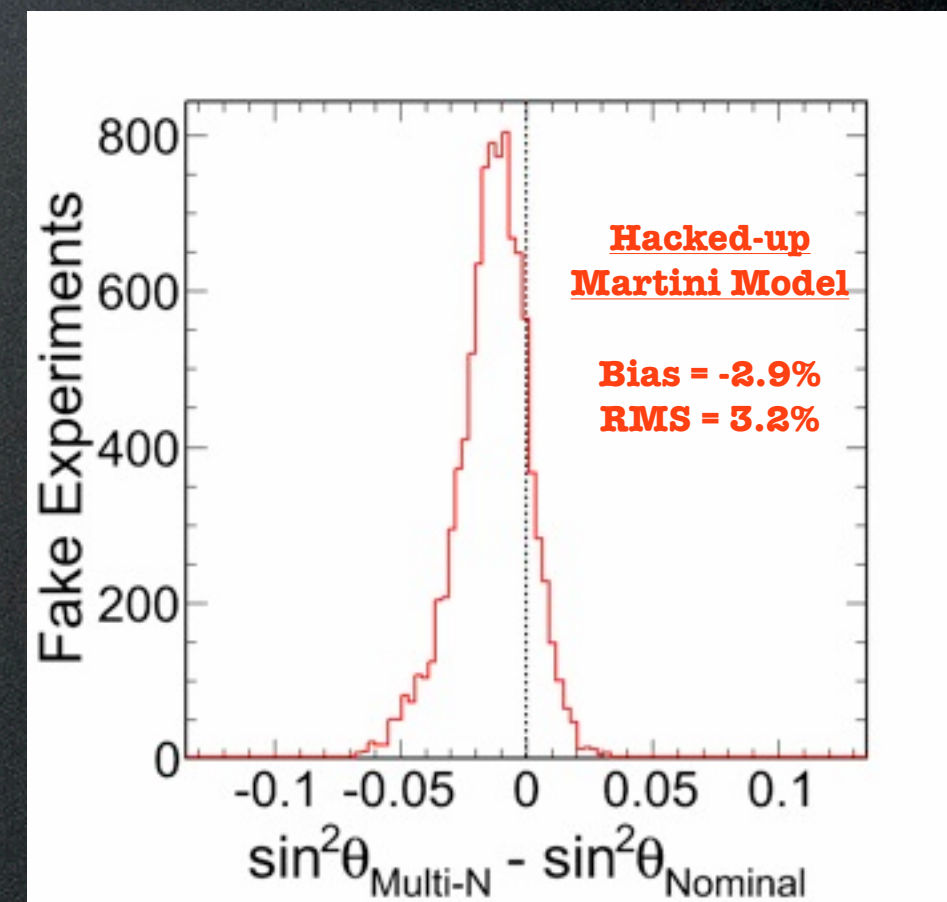
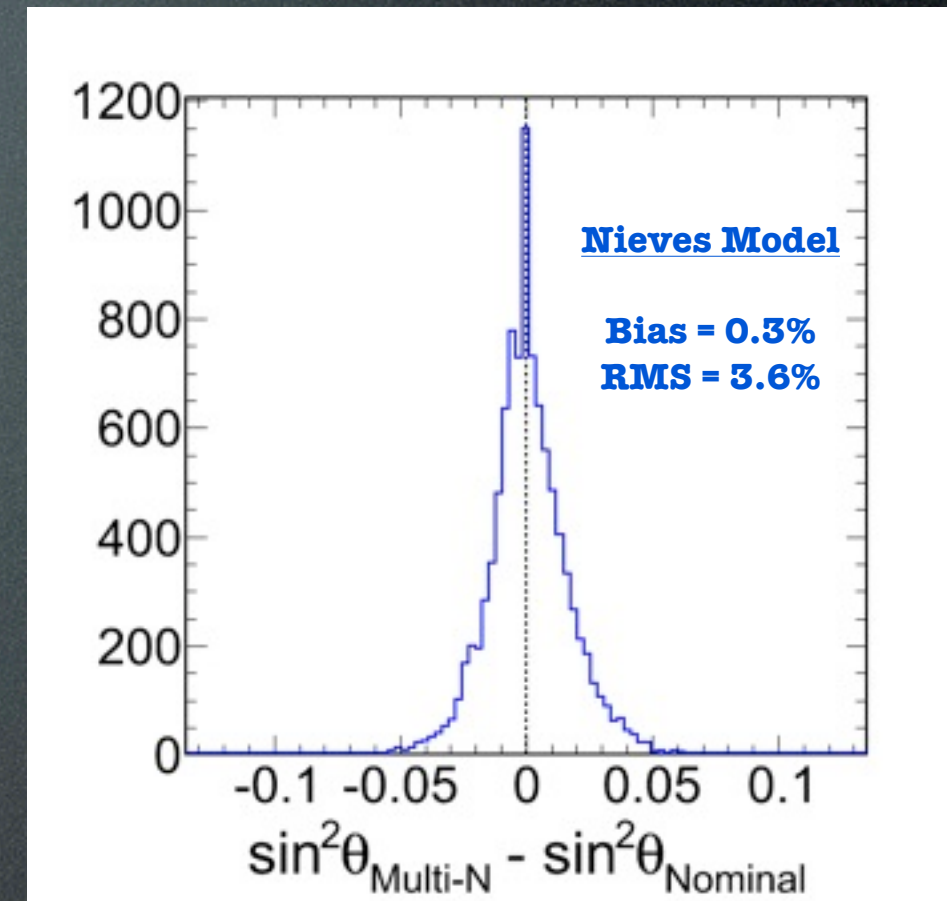
Summary

- To reach ultimate Hyper-K precision, it will be necessary to constrain E_ν reconstruction
 - NuPRISM provides the only data-driven mechanism for achieving this
- NuPRISM can also measure many other important physical processes
 - Sterile neutrinos and a variety of unique cross section measurements
- It is important to build the first version now! (NuPRISM-Lite)
 - A lot of interesting physics in the next 5 years!
 - Ideal tool for Hyper-K R&D
 - Intermediate project to expand Hyper-K involvement
 - We need a detailed understanding of NuPRISM to ensure it will achieve Hyper-K goals (calibration requirements, etc.)
- A full proposal will be submitted to J-PARC in June
- Additional collaborators are welcome!
 - Consider attending the NuPRISM meeting on Sunday and/or the week-long workshop in March

Supplement

Effect on T2K ν_μ Disappearance

- Create “fake data” samples with flux and cross section variations
 - With and without multi-nucleon events
- For each fake data set, full T2K near/far oscillation fit is performed
 - For each variation, plot difference with and without multi-nucleon events
- For Nieves model, “average bias” (RMS) = **3.6%**
- For Martini model, mean bias = -2.9%, RMS = 3.2%
 - Full systematic = $\sqrt{(2.9\%^2 + 3.2\%^2)} = \mathbf{4.3\%}$
 - **This would be one of the largest systematic uncertainties**
- But this is just a comparison of 2 models
 - How much larger could the actual systematic uncertainty be?
- **We need a data-driven constraint!**

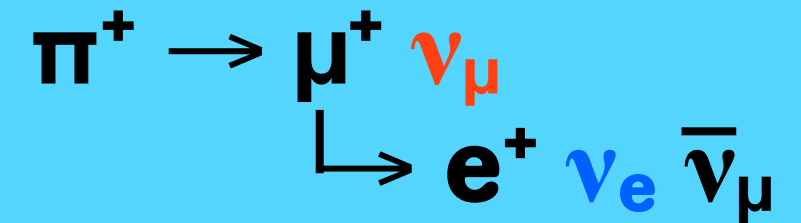


Interpreting Linear Combinations

- After ν PRISM linear combination:
 - CC- ν_μ spectrum should reproduce oscillated far detector spectrum:
Good!
 - NC- ν_μ backgrounds will also appear “oscillated”:
Bad!
 - NC events are unaffected by oscillations at Super-K
- **NC events must be subtracted** at both Super-K and nuPRISM
 - Introduces cross section model dependence
- However, **NC backgrounds can be very well measured** using mono-energetic beams
 - Significantly reduces cross section model dependence
- In current analysis (see later slides), NC constraint has not yet been applied
 - **Conservative errors**

ν Energy Spectrum

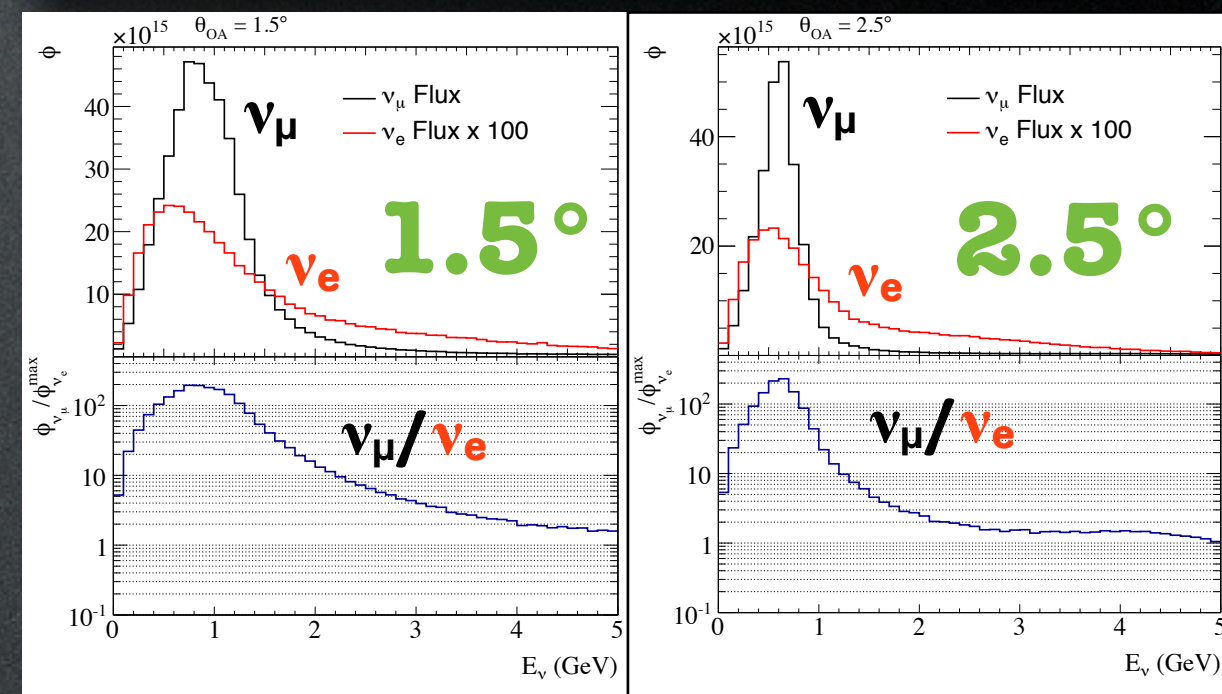
Flux < 1 GeV is dominated by π^+ decay



ν_μ produced in 2-body decay

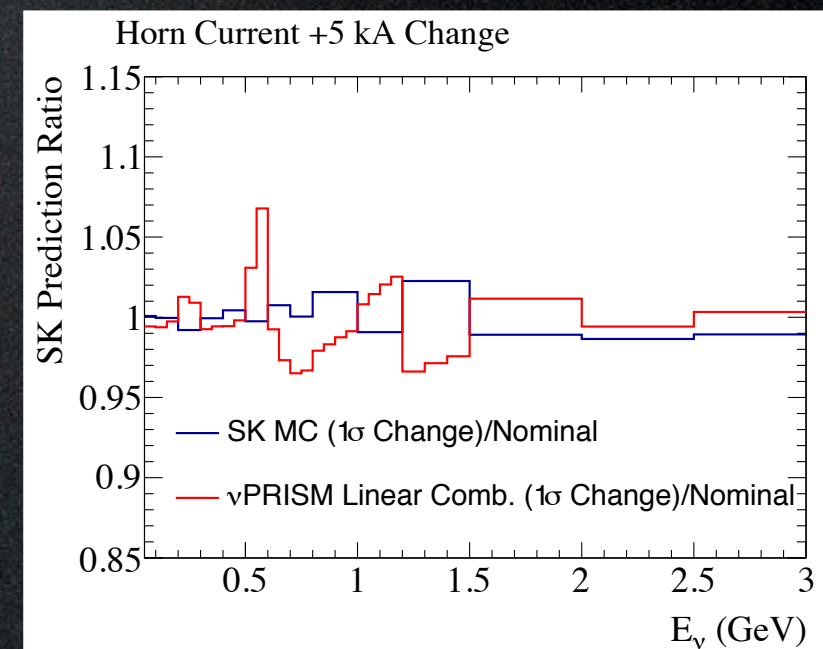
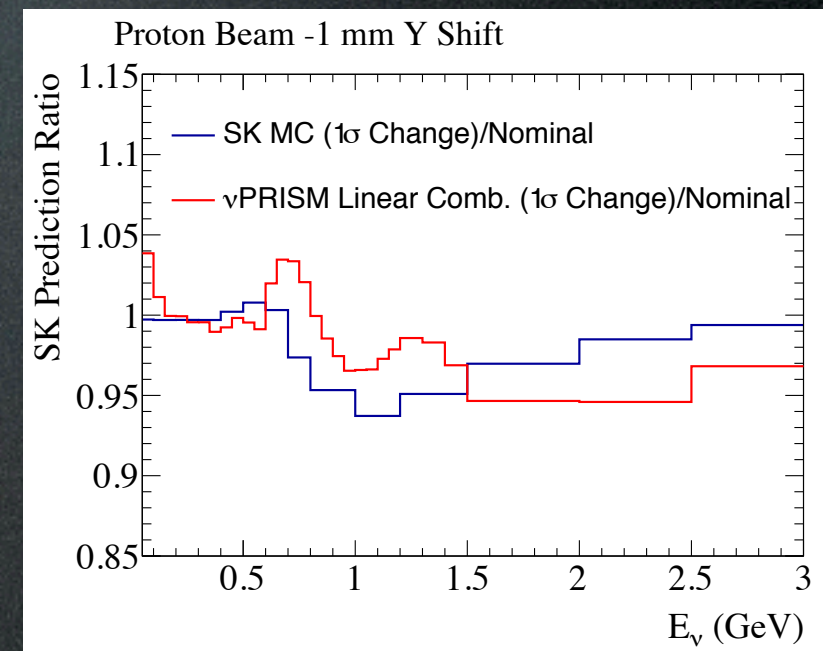
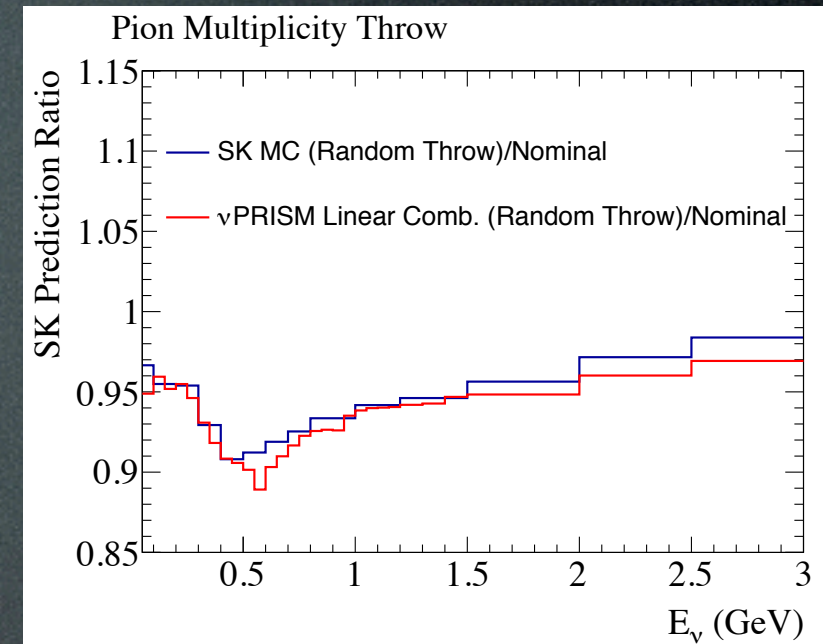
ν_e produced in 3-body decay

☞ ν_μ experience more off-axis affect



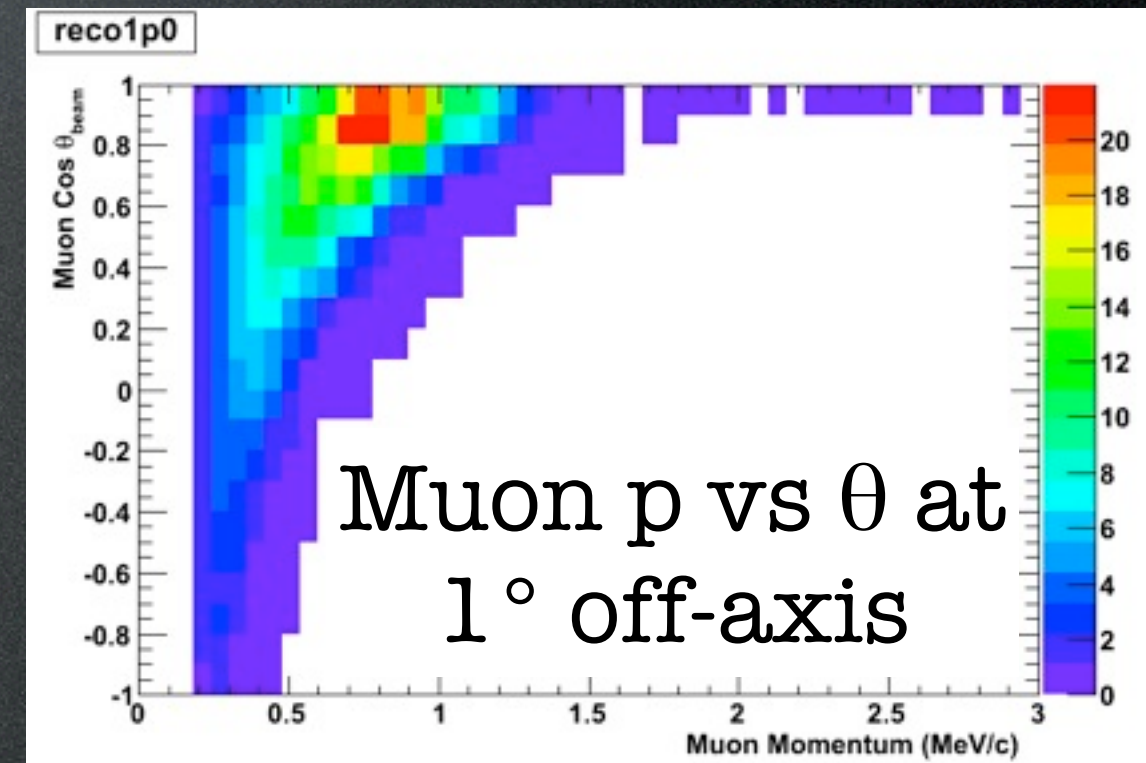
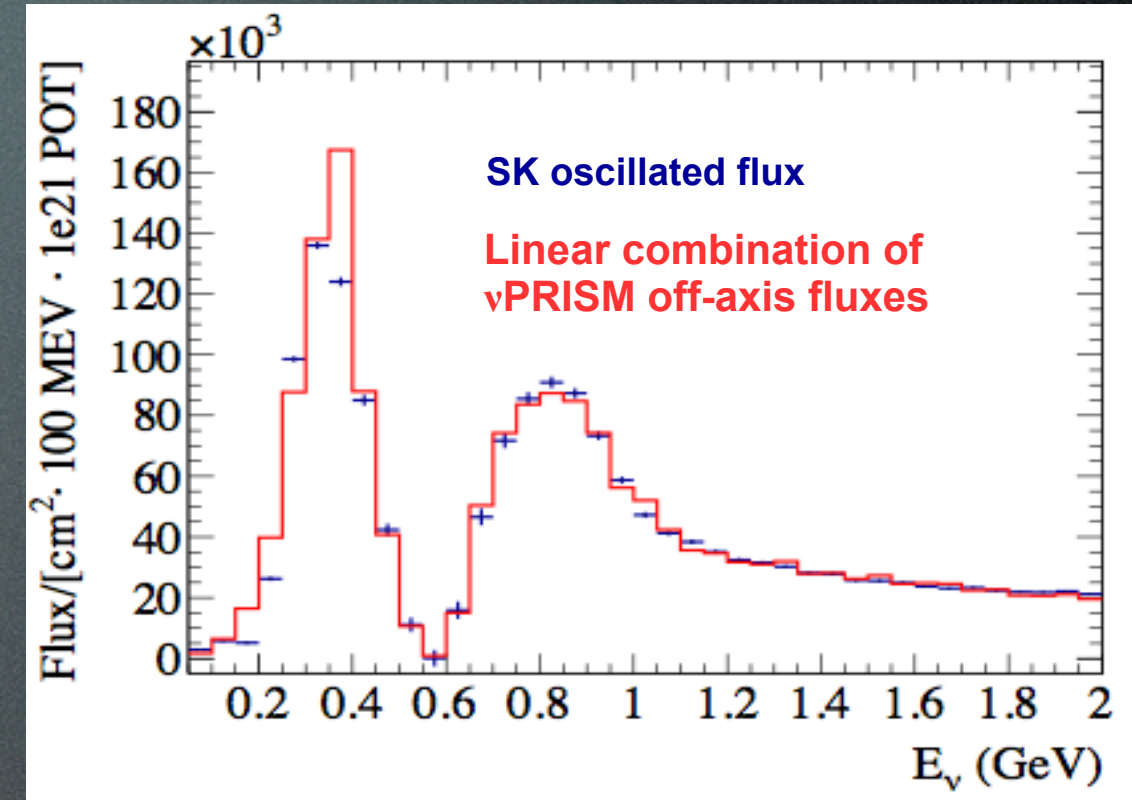
More on Beam Errors

- Haven't we just replaced **unknown cross section errors** with **unknown flux errors**?
 - Yes! But only relative flux errors are important!
 - Cancellation exist between nuPRISM and far detector variations
- **Normalization uncertainties will cancel** in the ν PRISM analysis
 - Cancellations persist, even for the ν PRISM linear combination
 - Shape errors are most important
- For scale, **10% variation** near the dip means **~1% variation** in $\sin^2 2\theta_{23}$
 - Although this region is dominated by feed down
- Full flux variations are reasonable
 - No constraint used (yet) from existing near detector!

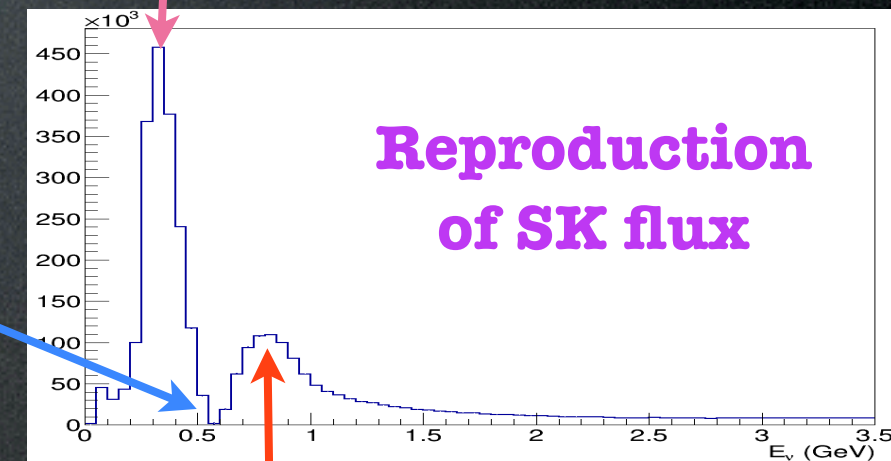
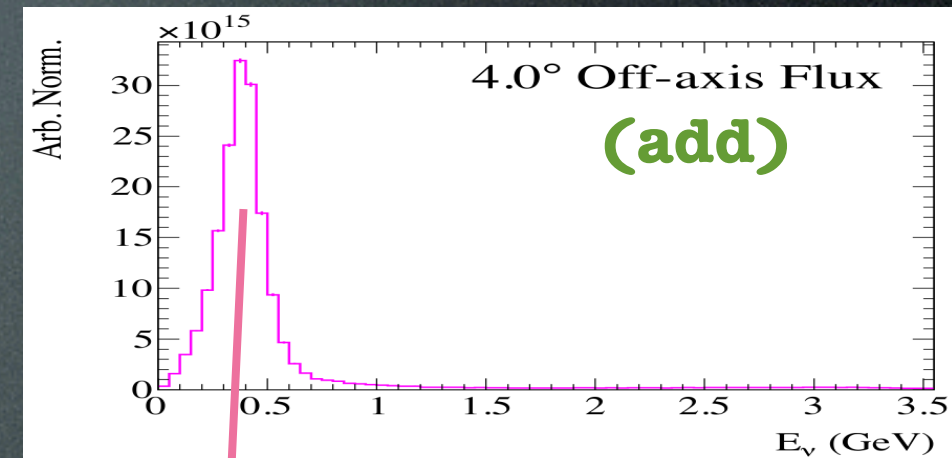
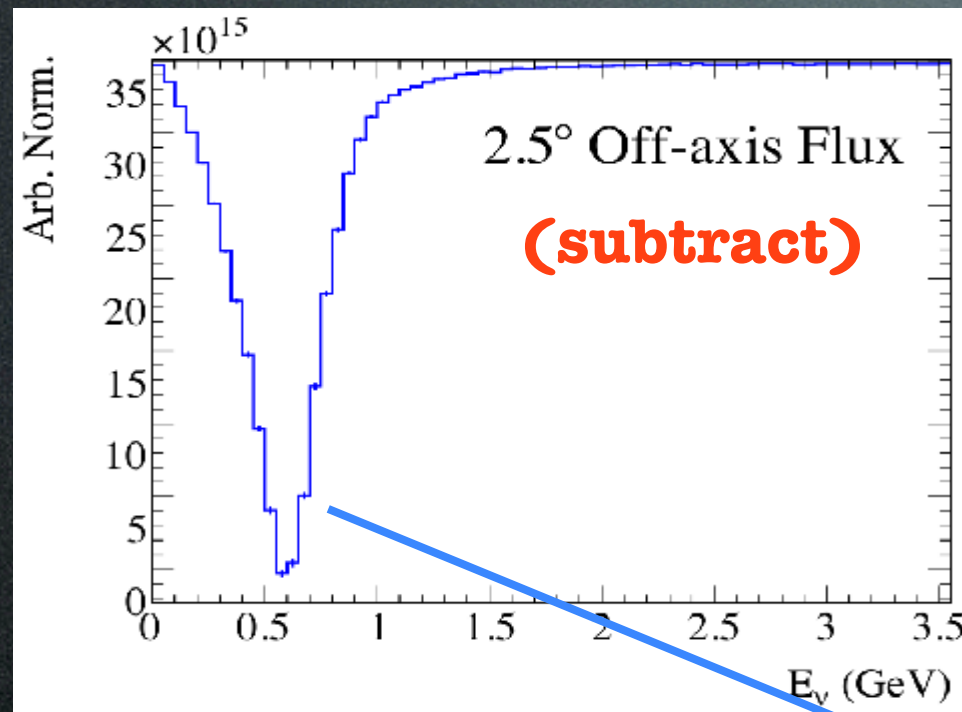


nuPRISM Technique

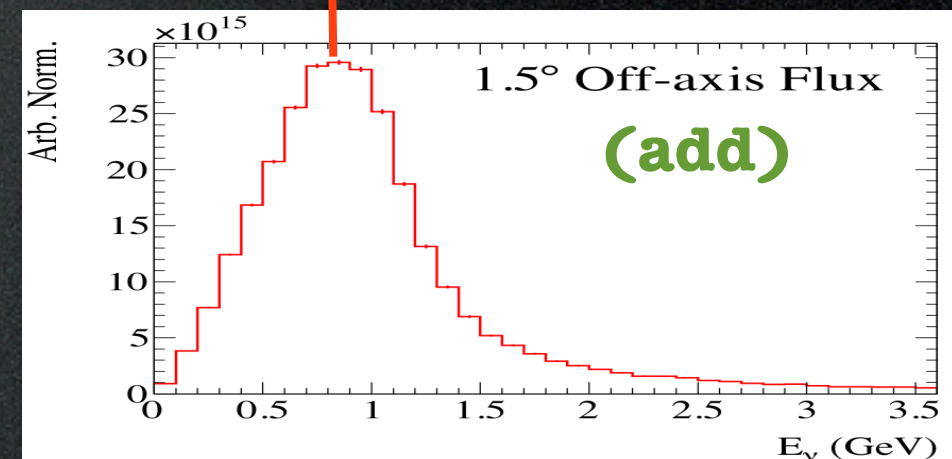
- Flux is now the same at the near and far detector
 - **Can just measure observed muon p vs θ for any oscillated flux**
- Same signal selection as used at Super-K
 - Single, muon-like ring
- **Signal events** are defined as **all true single-ring, muon-like events**
 - A muon above Cherenkov threshold
 - All other particles below Cherenkov threshold
 - Signal includes CCQE, multi-nucleon, $\text{CC}\pi^+$, etc.
- **No need to make individual measurements of each process and extrapolate to T2K flux**



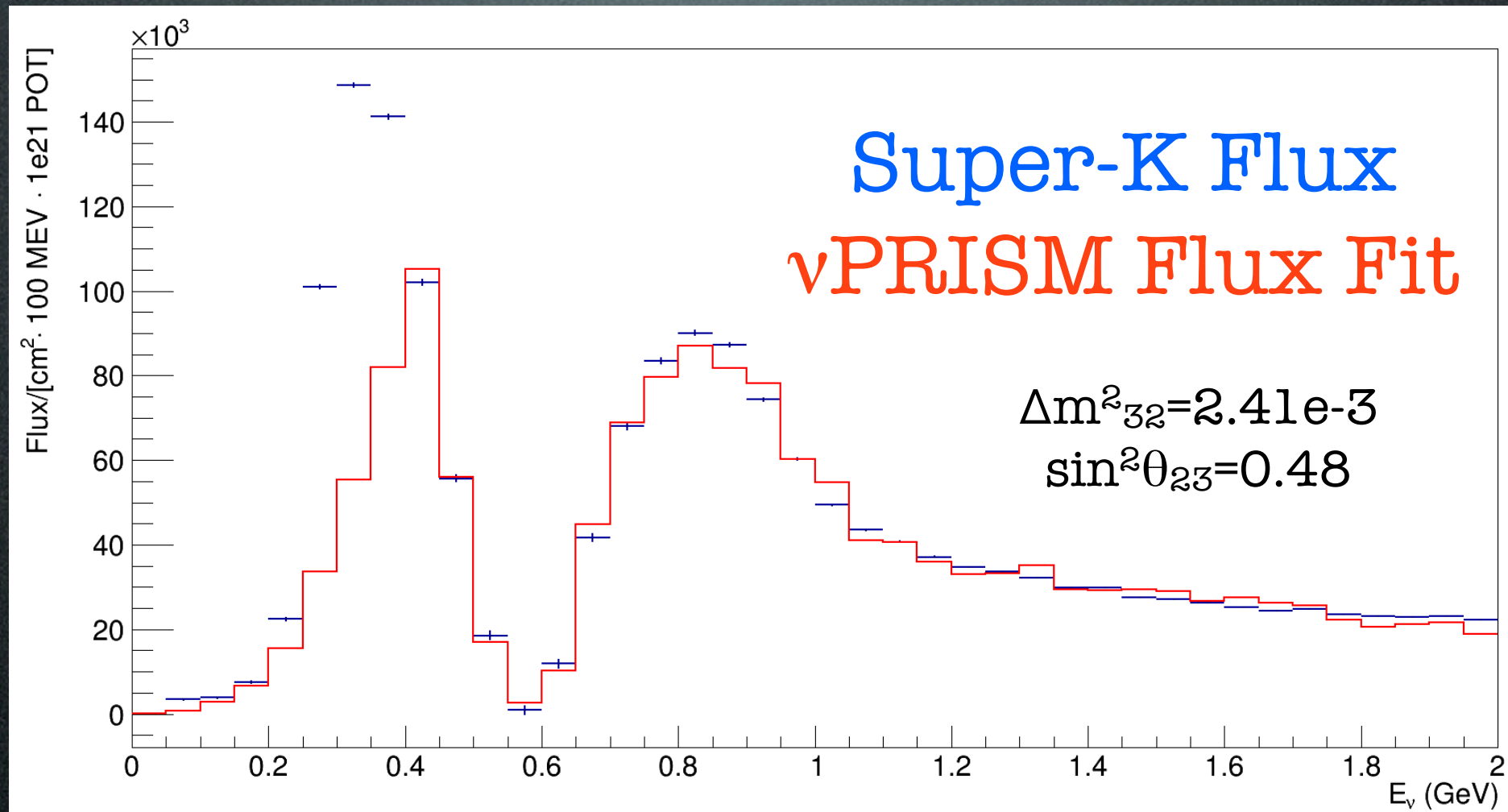
Reminder: Analysis Concept



- Different slices of nuPRISM are combined to reproduce an oscillated SK flux
 - **Flux only!** No cross sections or detector response at this point
- For simplicity, only 3 slices are shown here
 - The default analysis **uses 60 slices**



Flux Fit



- Fit for coefficients of 60 off-axis ν PRISM slices to match a chosen Super-K oscillated spectrum
 - Fit between 400 MeV and 2 GeV
 - **Repeat this fit for every set of oscillation parameters**
- Notice disagreement at low energy
 - The most off-axis flux (4°) peaks at 380 MeV, so difficult to fit lower energies
 - Could extend detector further off-axis, but the low energy region is not very important to extract oscillation physics (e.g. nuclear feed-down not an issue)

nuPRISM Prediction for Super-K

- **Efficiency correction** is still needed for both ν PRISM and Super-K
- ν PRISM and Super-K have **different detector geometries**
 - Particles penetrate ID wall (and get vetoed) more often in nuPRISM
 - Particle ID degrades near the tank wall
- The efficiency correction is performed in **muon momentum and angle** to be as **model independent** as possible
 - This should be nearly a pure geometry correction
- For now, fit in Super-K E_{rec} distribution (in future, just use muon p, θ)

$$E_{\text{rec},j}^{SK}(\Delta m_{32}^2, \theta_{23}) = \sum_{p,\theta} \left[\sum_i^{OAangles} c_i(\Delta m_{32}^2, \theta_{23}) (N_{p\theta i}^{obs} - B_{p\theta i}) \frac{\epsilon_{p\theta}^{SK}}{\epsilon_{p\theta}^{\nu\text{PRISM}}} \right] * M_{p\theta j}$$

predicted
Super-K E_{rec}
distribution

weight for
off-axis slice, i

events in
muon p, θ bin
in slice, i

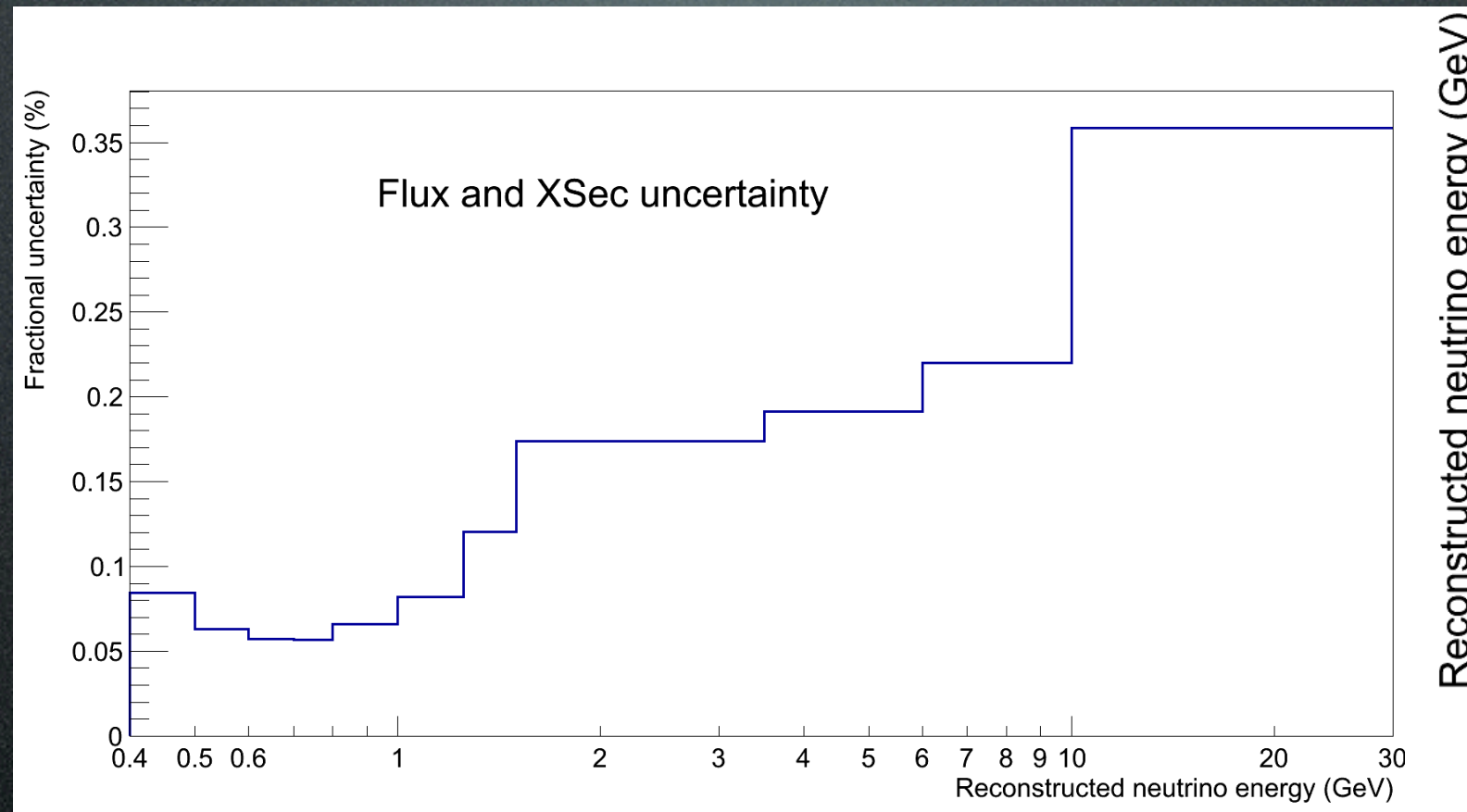
background
subtraction

efficiency
ratio

translation
matrix
 $p, \theta \rightarrow E_{\text{rec}}$

Systematic Covariance Matrices

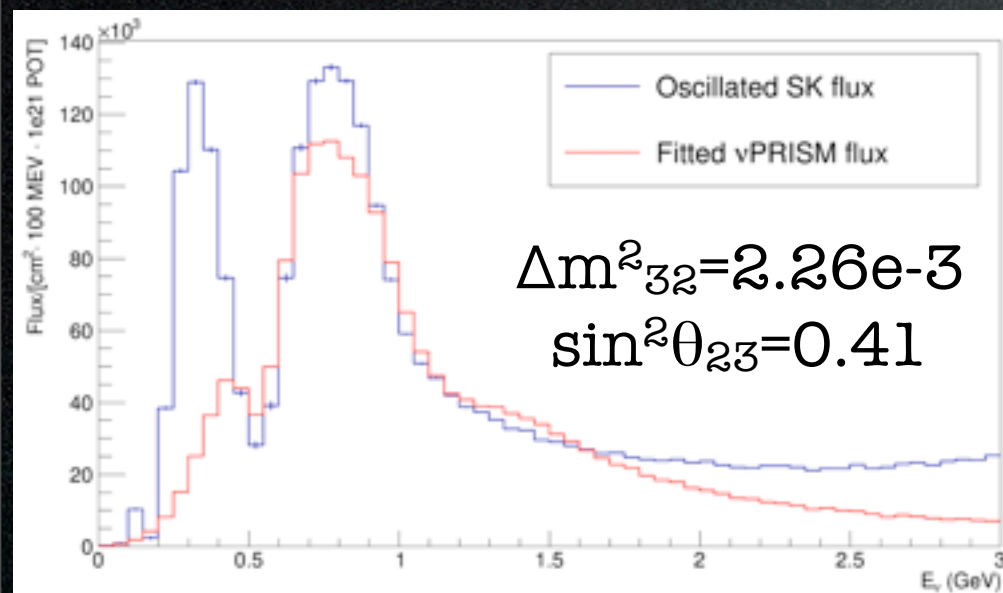
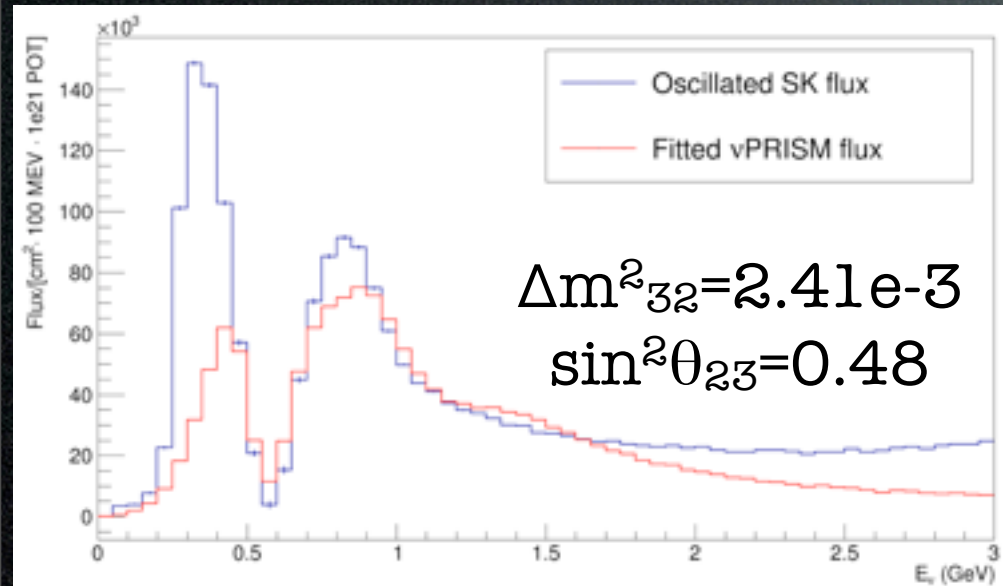
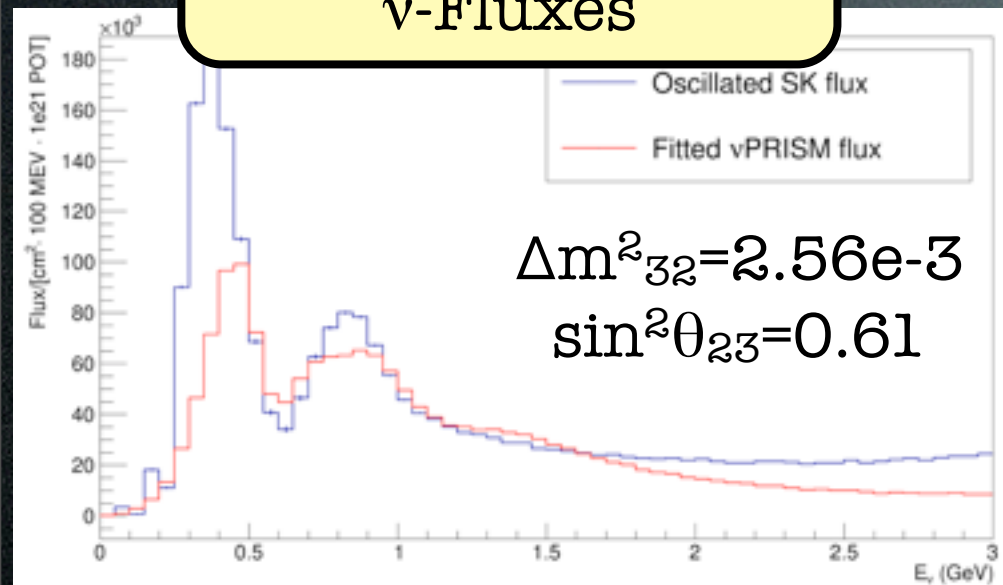
Analysis is performed in unequal-sized E_{rec} bins



- Fractional uncertainties are shown (normalized to bin content)
- At high energies, ν PRISM provides no constraint
 - Detector acceptance: all muons exit the inner detector
 - Subject to full flux & cross section uncertainties
- Bin 3 (600-700 MeV) has a 6% uncertainty

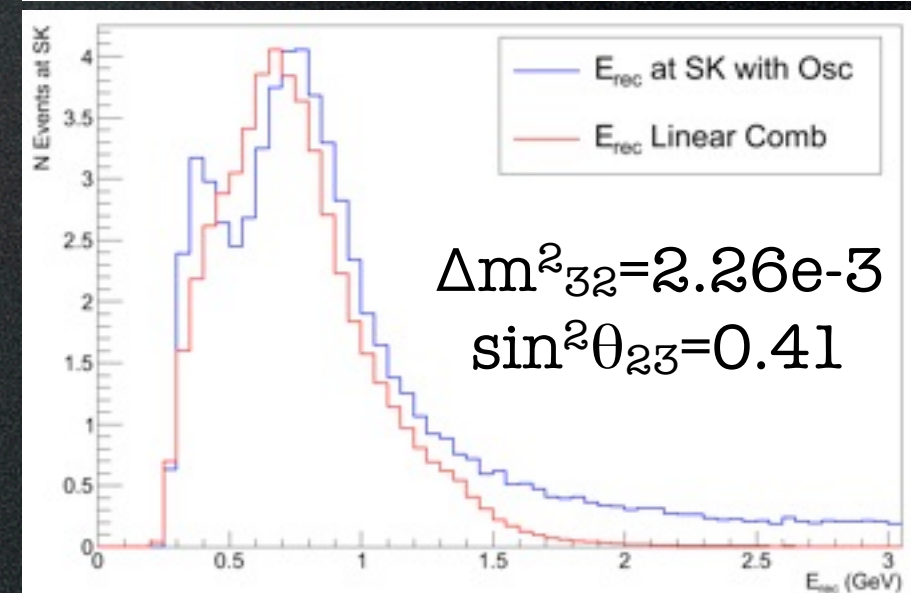
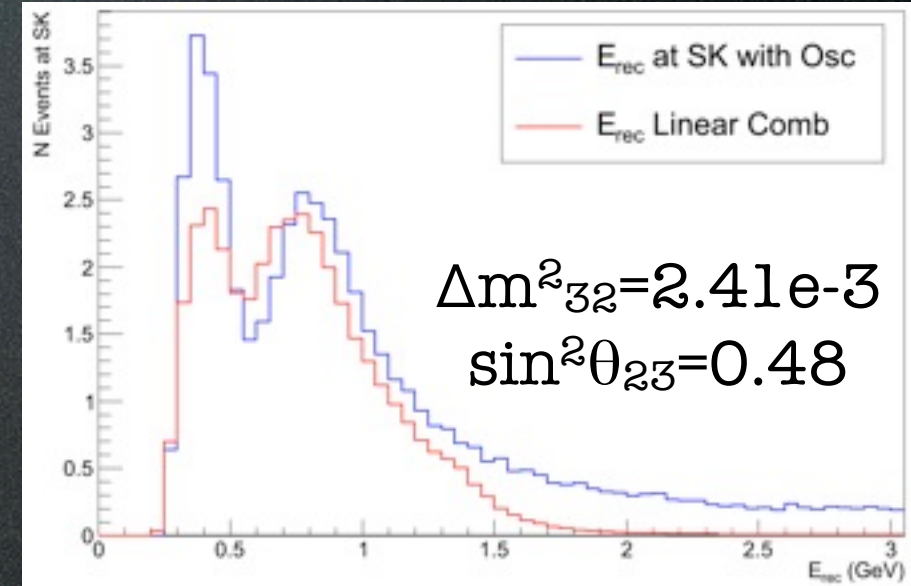
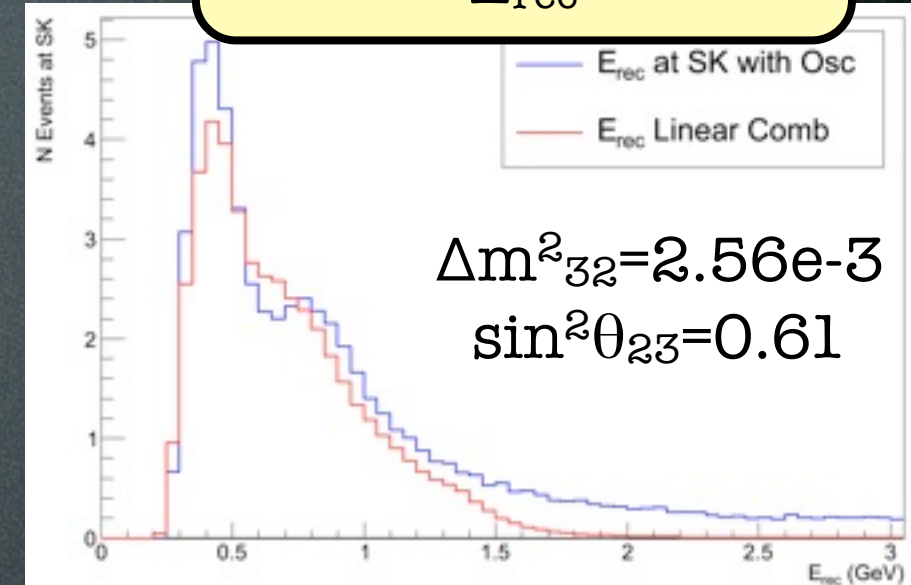
Smoothed ν -Flux Fits

ν -Fluxes

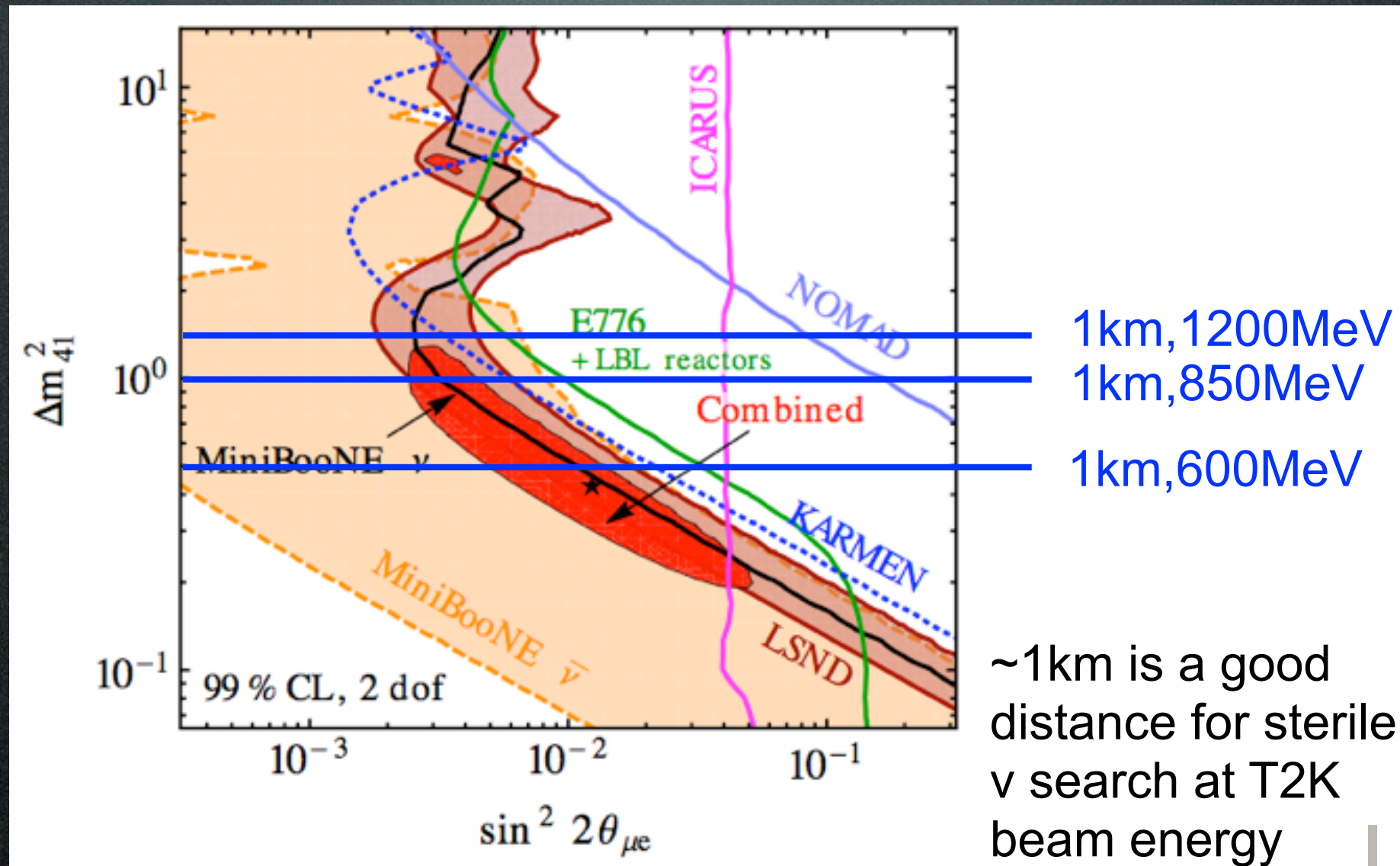


- Smoothed flux fits do not match as well
 - Easy to improve, if necessary
- However, very small increase to systematic uncertainties
 - Flux systematic variations are large
- Fits can be improved
 - Smoothness can be relaxed near fast-changing features
 - Off-axis angle bins need not be equal size

E_{rec}

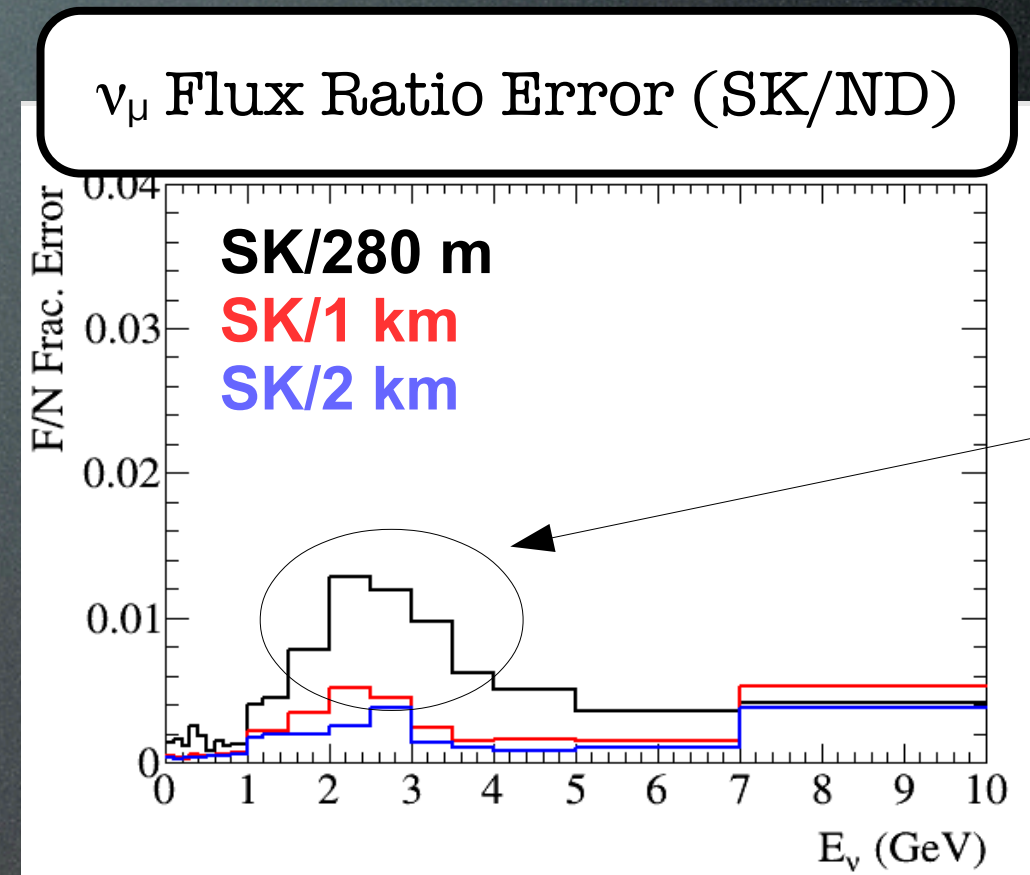
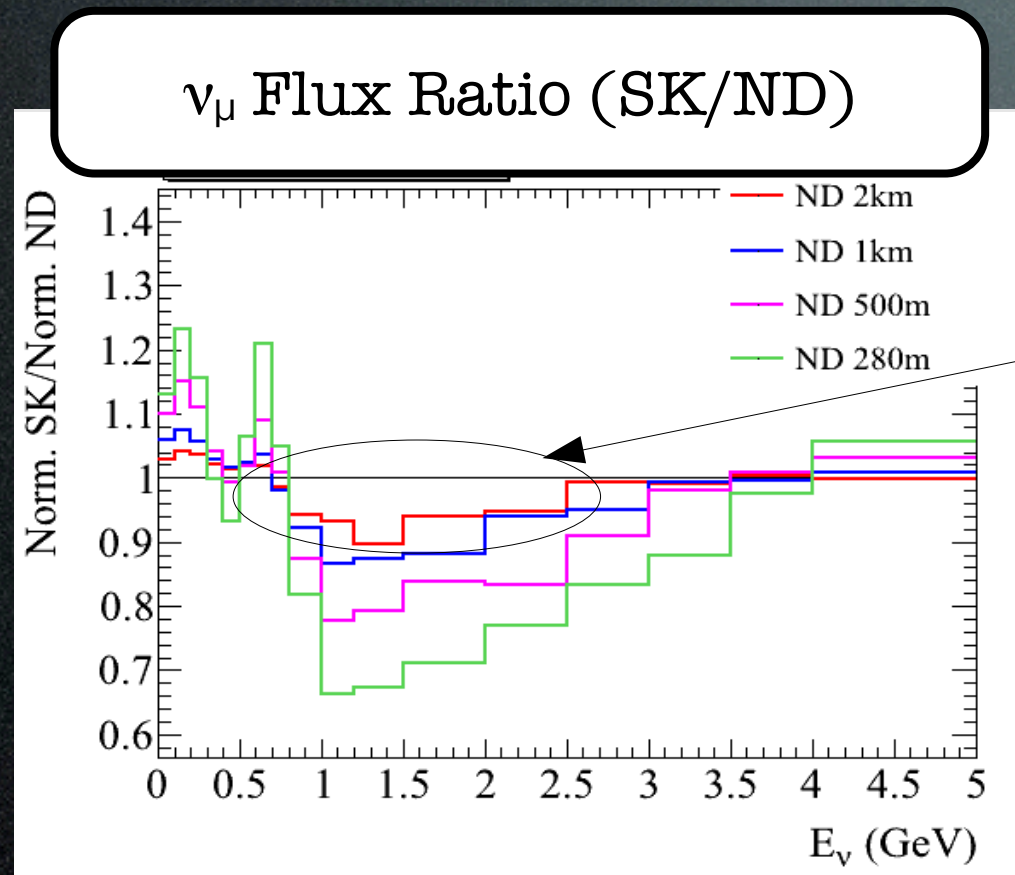


Sterile Neutrinos



- The 1 km baseline is ideal for sterile neutrinos
 - Many repeated measurements for varying energy spectra
 - Continuously sample a variety of L/E values

Detector Location: Energy Spectrum Ratio

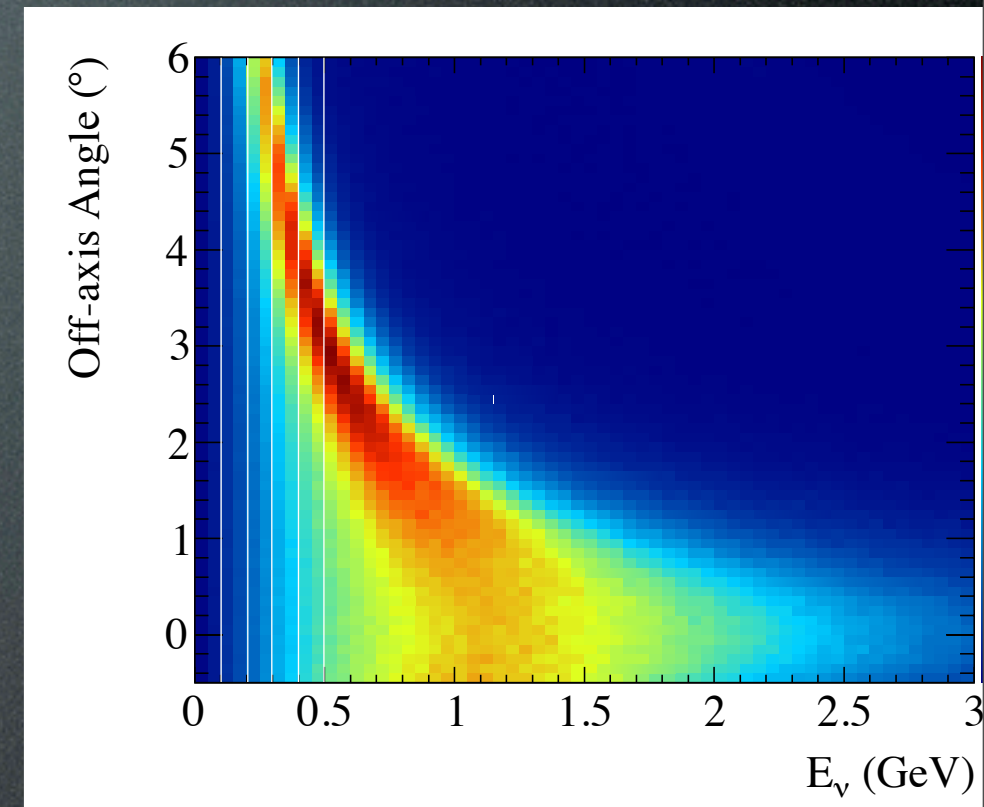


- At 280 m, the flux shape has 20-30% differences below 1 GeV
 - Uncertainty in the ratio is noticeably larger, but mostly above 1 GeV
- The difference between 1km and 2km is small in both shape and shape uncertainty

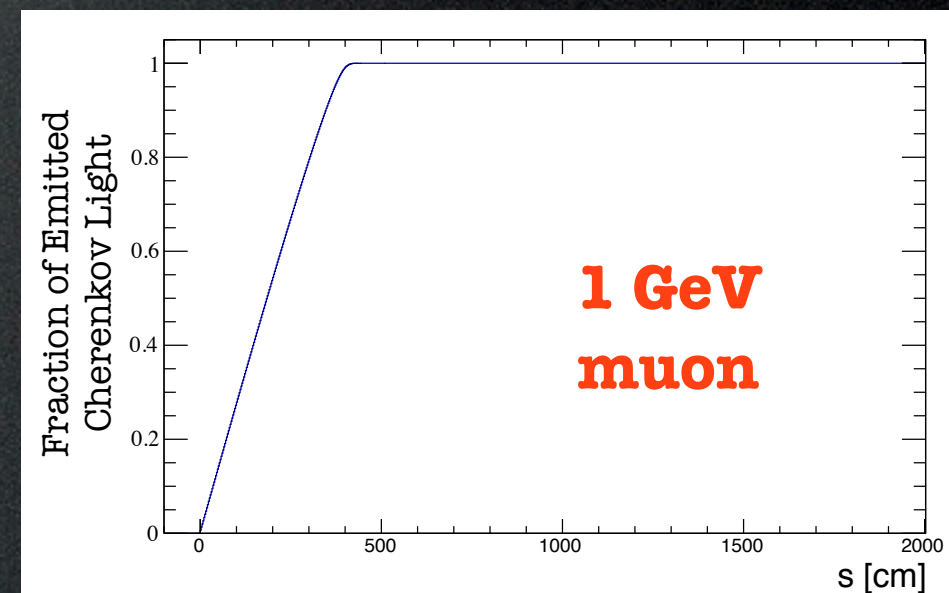
Other Design Considerations

- **Civil construction is expensive!**
 - Smaller hole = More affordable
- **Off-axis angle range (i.e. E_ν range)**
 - On-axis flux peaks at 1.2 GeV
 - 4° (6°) off-axis peaks at ~ 380 (~ 260) MeV
 - Beam points 3.63° below horizon, so get $\sim 4^\circ$ for free
- **Distance to target**
 - At 1 (1.2) km, need 54 (65) m deep pit to span 1° - 4°
 - Event pileup must be manageable (see later slides)
- **Tank diameter**
 - Determines maximum muon contained
 - 4 m (+ FV cut) for 1 GeV/c muon
 - PID degrades near the wall
 - Important for selecting e-like events
 - Larger = more stats, but also more pileup
 - Larger = more PMTs = more expensive
 - How much outer detector is necessary?

Off-axis Fluxes

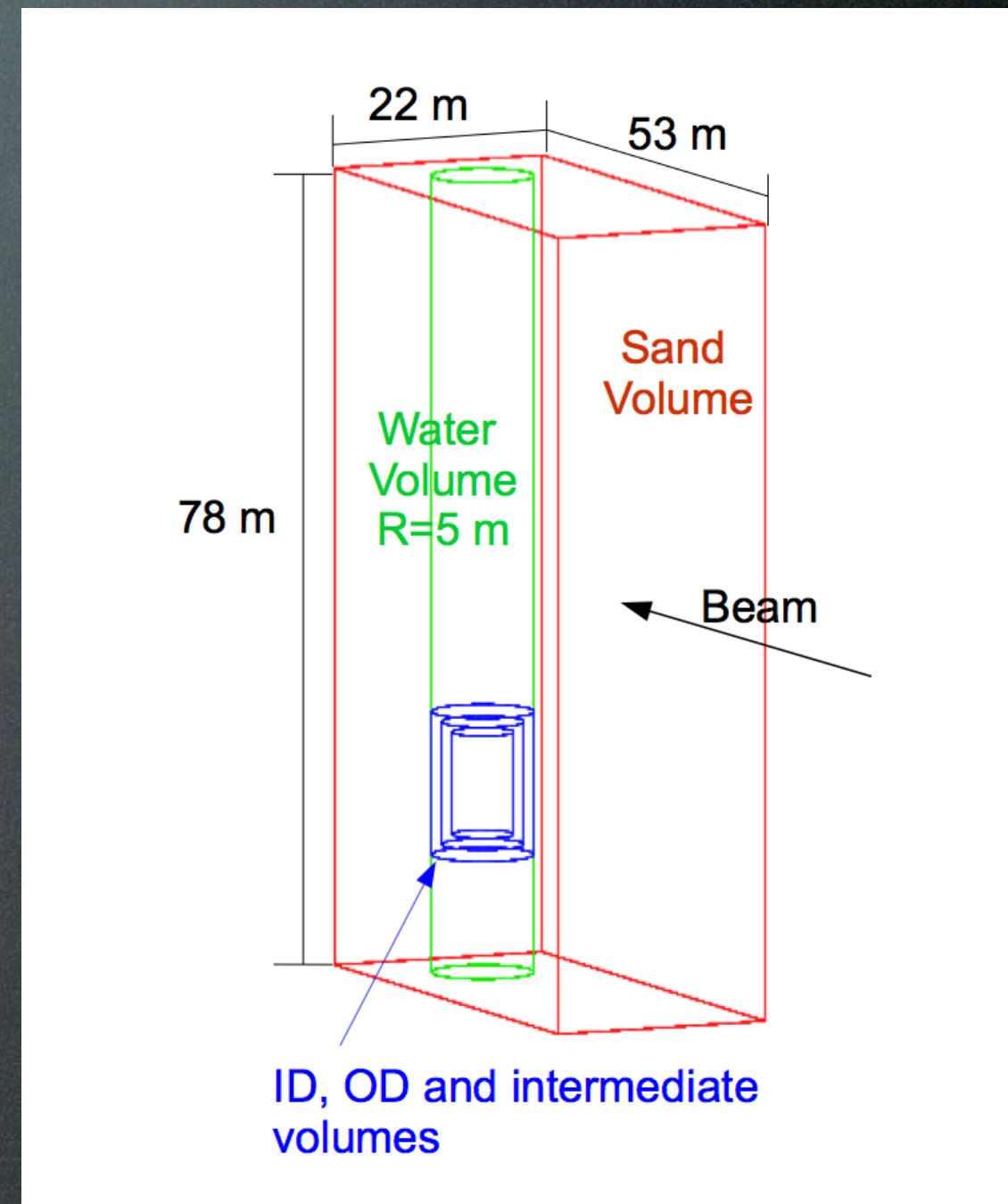


Muon Range



Event Pileup

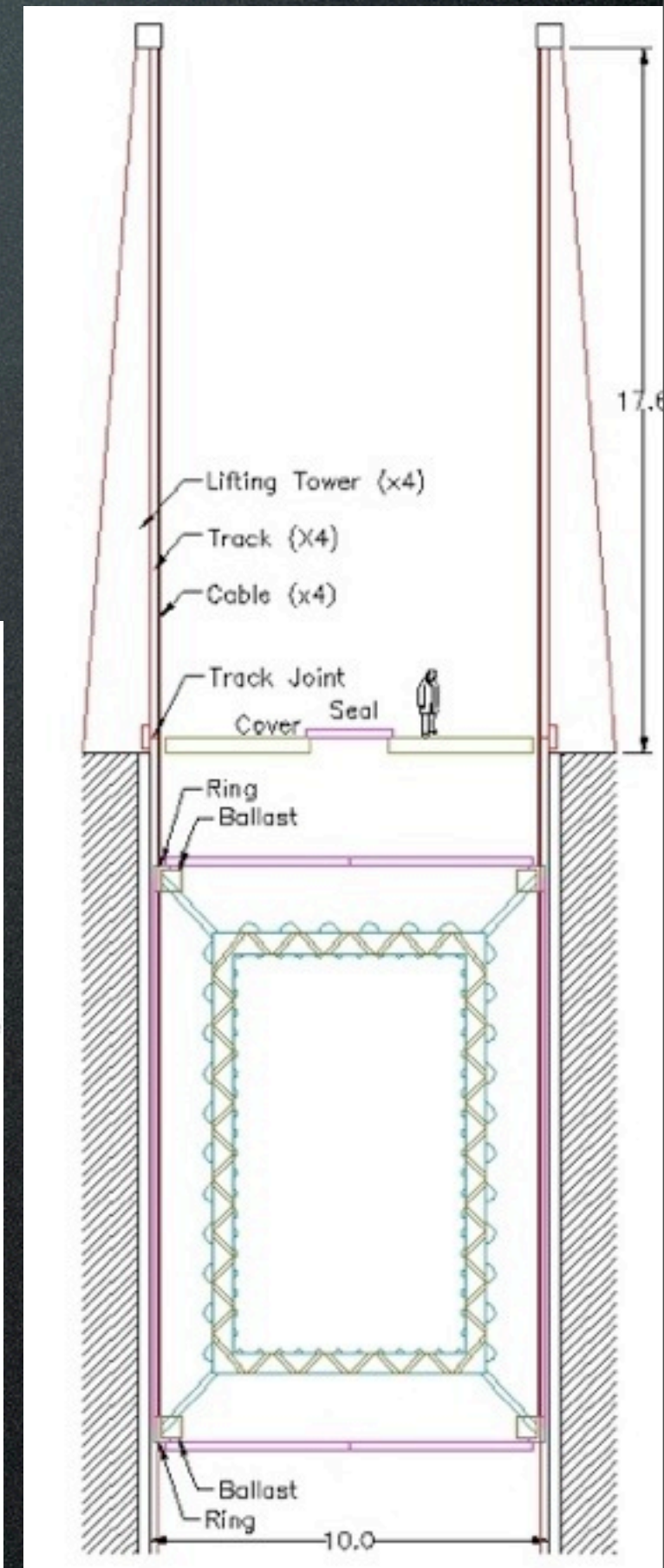
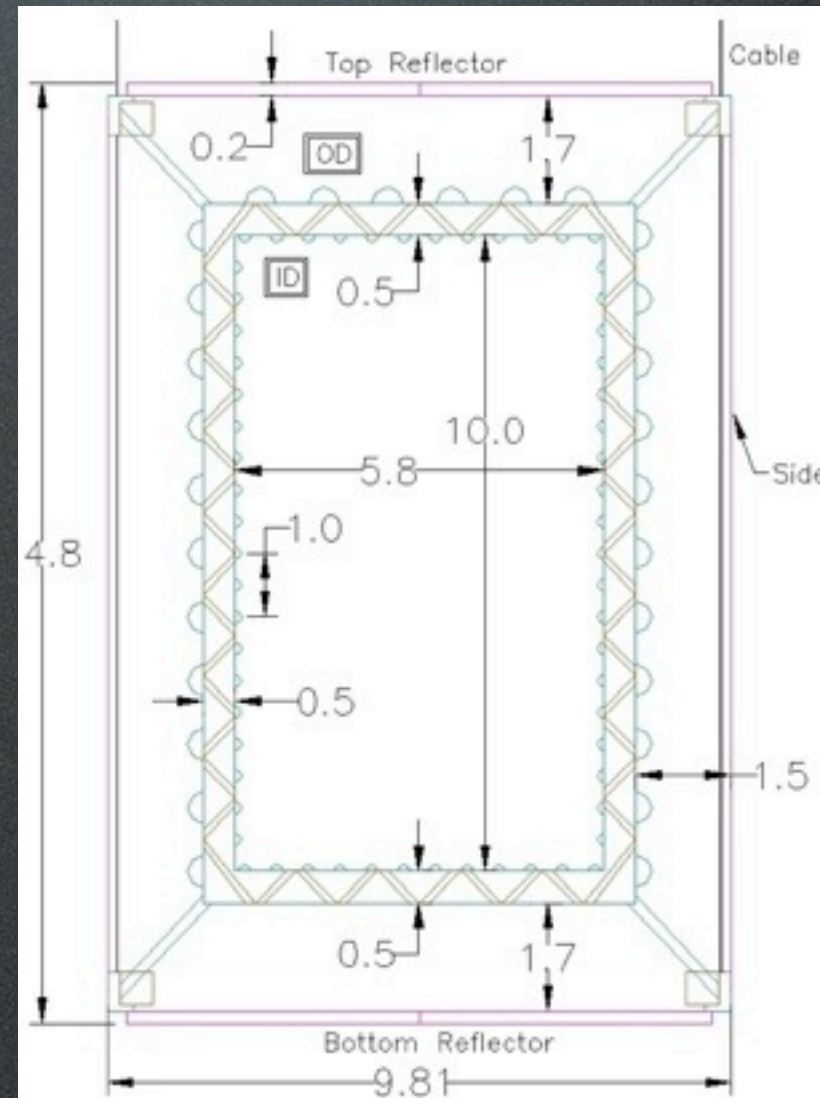
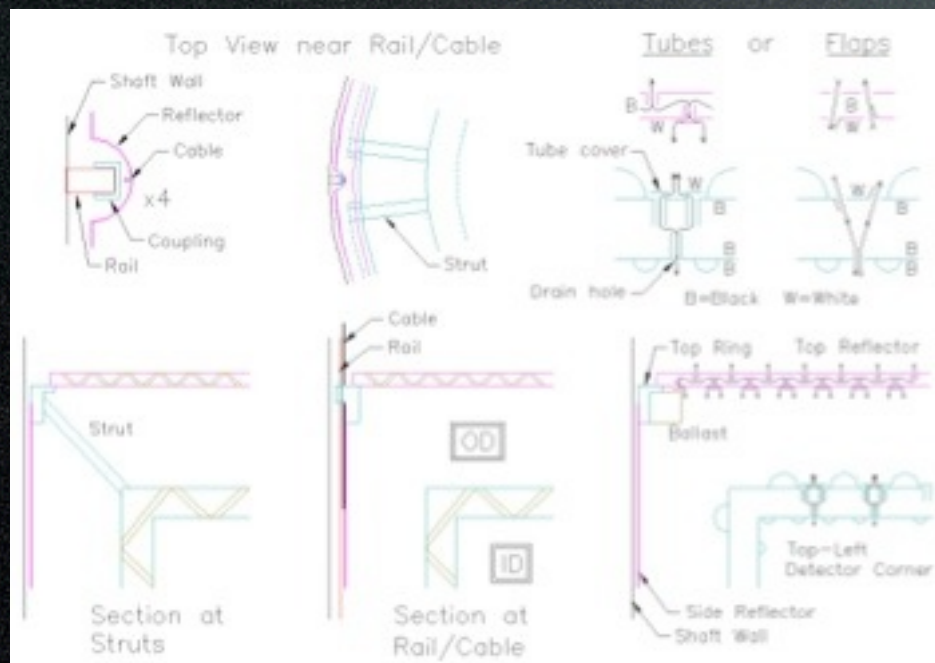
- Full GEANT4 simulation of water and surrounding sand
 - Using T2K flux and neut cross section model
- 8 beam bunches per spill, separated by 670 ns with a width of 27 ns (FWHM)
- **41% chance of in-bunch OD activity during an ID-contained event**
 - Want to avoid vetoing only on OD light (i.e. using scintillator panels)
- **17% of bunches have ID activity from more than 1 interaction**
 - 10% of these have no OD activity
 - Need careful reconstruction studies
 - (but multi-ring reconstruction at Super-K works very well)



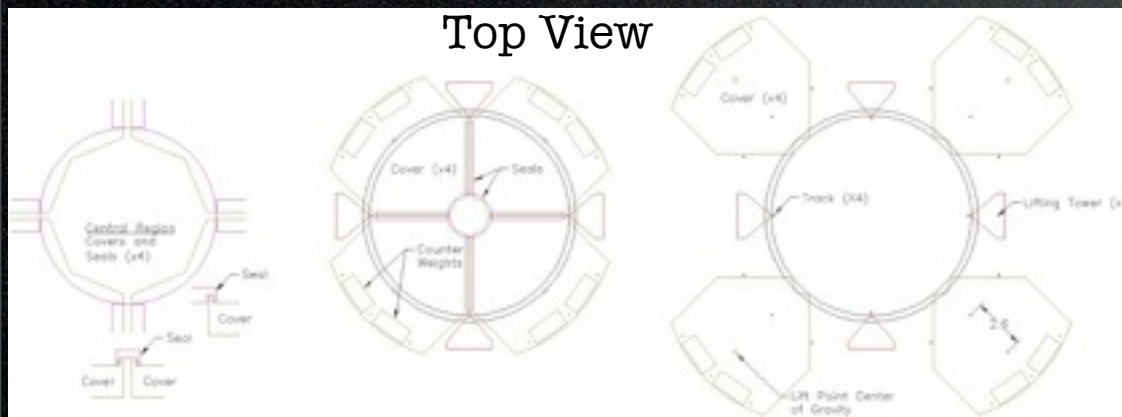
Pileup Rates at 1 km Look Acceptable!

Detector Frame

- Initial proposal for ID/OD frame and lifting mechanism has been produced
- Careful consideration given to water flow rate while in motion
- 4 towers allow the entire detector to be lifted out of the water tank for maintenance



Top View



PMTs

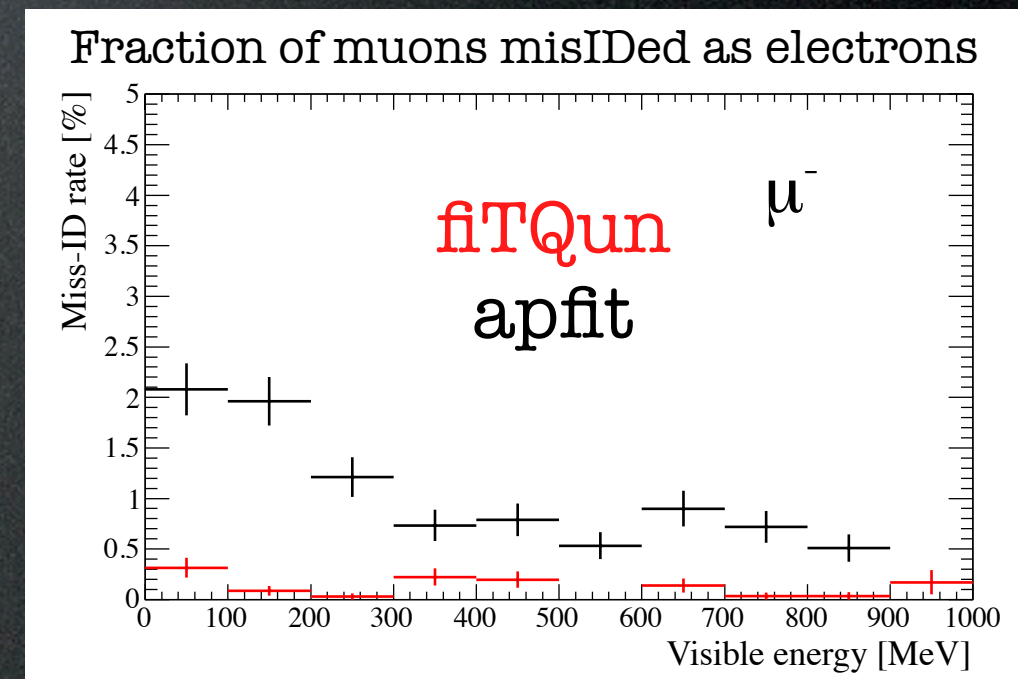
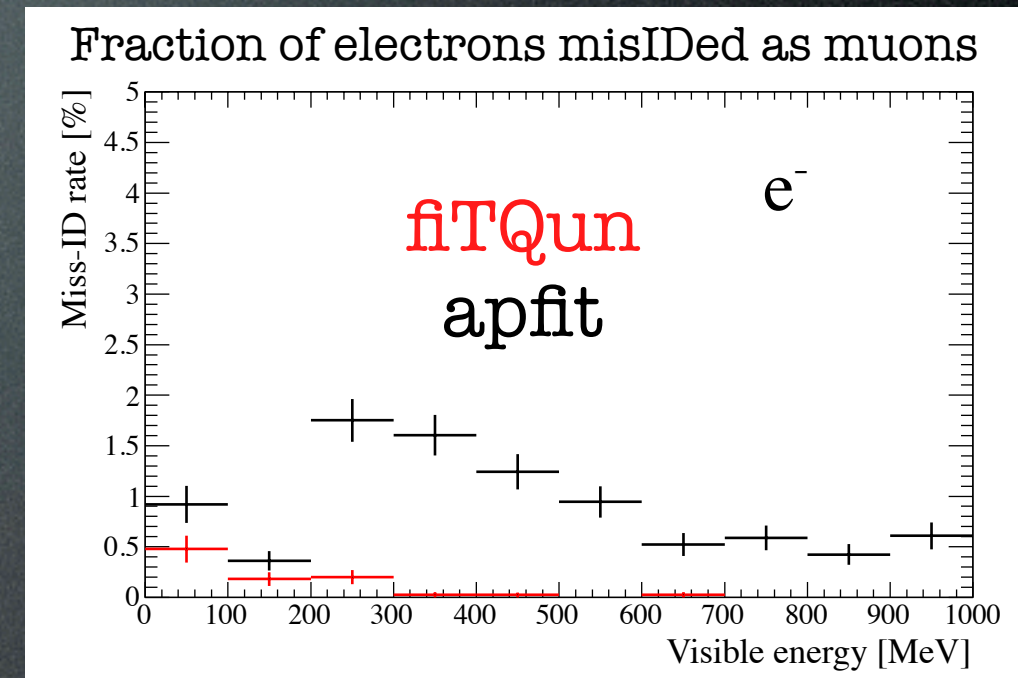
- For the ID, both 8" and 5" PMTs are being considered
 - Perhaps with high-quantum-efficiency (HQE) coating
 - Also considering Hyper-K-style hybrid photodetectors (HPD)
- Initial Hamamatsu estimate for basic 8" R5912 PMT is much more expensive than assumed for 2km detector
 - US \$4.3M for 3,000 PMTs**
- UK/Texas company ETEL/ADIT has also been consulted
 - Basic 8" PMT is \$1775
 - No HQE or HPD option available

Hamamatsu Estimates

Name	Type	QE%	Quantity	Price/PMT	Total Cost	Delivery Year
5" PMT	R6594-WPassy	25	8000	103,500	828M	
5" PMT HQE		35	5714	123,700	707M	
8" PMT	R5912-WPassy	25	3215	143,000	460M	
8" PMT HQE		35	2296	170,500	391M	
8" HPD HQE	R12112-WPmodule	35	2296	264,000	606M	2014
		35	2296	236,500	543M	2015
		35	2296	209,000	480M	2016
20" PMT HQE	R12860-WPassy	30	508	604,500	307M	2014
		30	508	572,000	291M	2015
		30	508	539,500	274M	2016
20" HPD HQE	R12850-WPmodule	30	508	715,000	363M	2014
		30	508	617,500	314M	2015
		30	508	520,000	264M	2016
20" HPD HQE	R12850-WPmodule	30	140	770,000	108M	2014
		30	140	665,000	93M	2015
		30	140	560,000	78M	2016
20" PMT	R12860-WPassy	30	140	651,000	91M	2014
		30	140	616,000	86M	2015
		30	140	581,000	81M	2016

Physics Capabilities

- Direct measurement of the relationship between lepton kinematics and neutrino energy
 - No longer rely solely on models
- 4π detector (like Super-K)
- Target material is water (like Super-K)
 - Can directly measure NC backgrounds
- Very good e/μ separation
- Can make a precise measurement of beam ν_e
 - π^0 background is well separated
 - Can also constrain ν_e cross sections



T2K Uncertainties

ND280 Analysis	ND280 Data	SK Selection	$\sin^2 2\theta_{13}=0.1$	$\sin^2 2\theta_{13}=0.0$	
No Constraint	--	Old	22.6%	18.3%	
No Constraint	--	New	26.9%	22.2%	
2012 method*	Runs 1-2	Old	5.7%	8.7%	Factor 2.4 more ND280 POT
2012 method**	Runs 1-3	Old	5.0%	8.5%	
2012 method	Runs 1-3	New	4.9%	6.5%	Improved SK π^0 rejection
2012 method***	Runs 1-3	New	4.7%	6.1%	New ND280 reconstruction, selection, binning
2013 method	Runs 1-3	New	3.5%	5.2%	
2013 method	Runs 1-4	New	3.0%	4.9%	Factor 2.2 more ND280 POT

*Results presented at Neutrino 2012 conference
 **Published results, arXiv:1304.0841v2
 ***Update to NEUT tuning with MiniBooNE data

These are very nice constraints!
 (if the current parametrization is to be believed)