

vPRISM: Disappearance Analysis

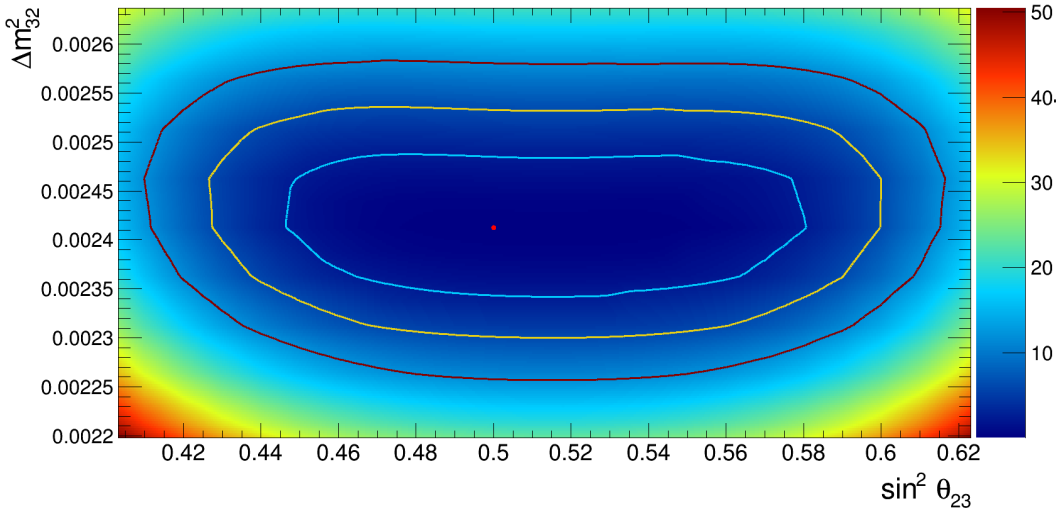
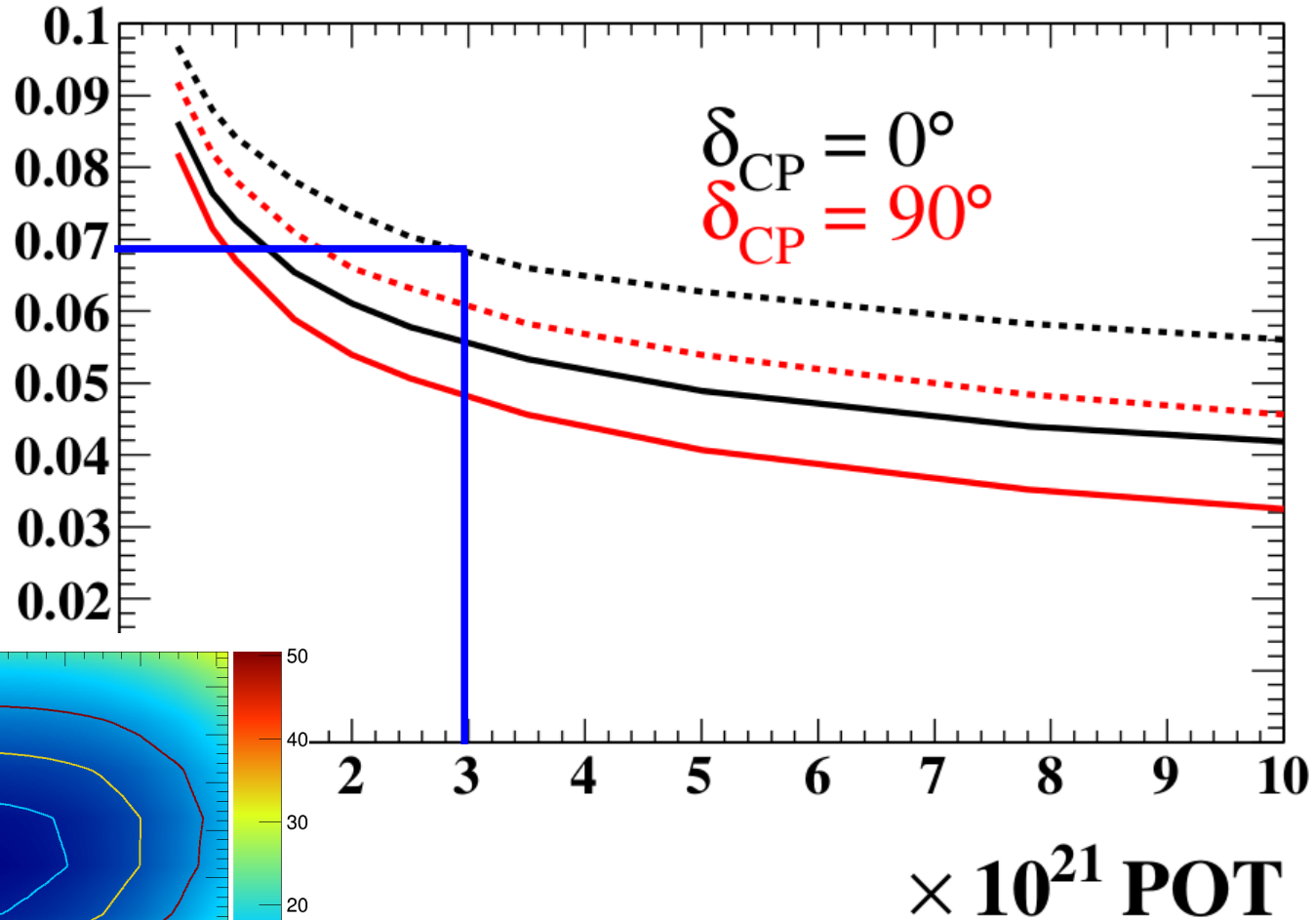
Mark Scott for the vPRISM collaboration
vPRISM workshop
17th Mar 2015 - IPMU

- At 6th HK meeting showed initial sensitivity plots for T2K statistics - 1 sigma width = 0.065, with current systematics

- Compare to T2K projection:

- 1 sigma width = 0.068
- nuPRISM slightly better than T2K

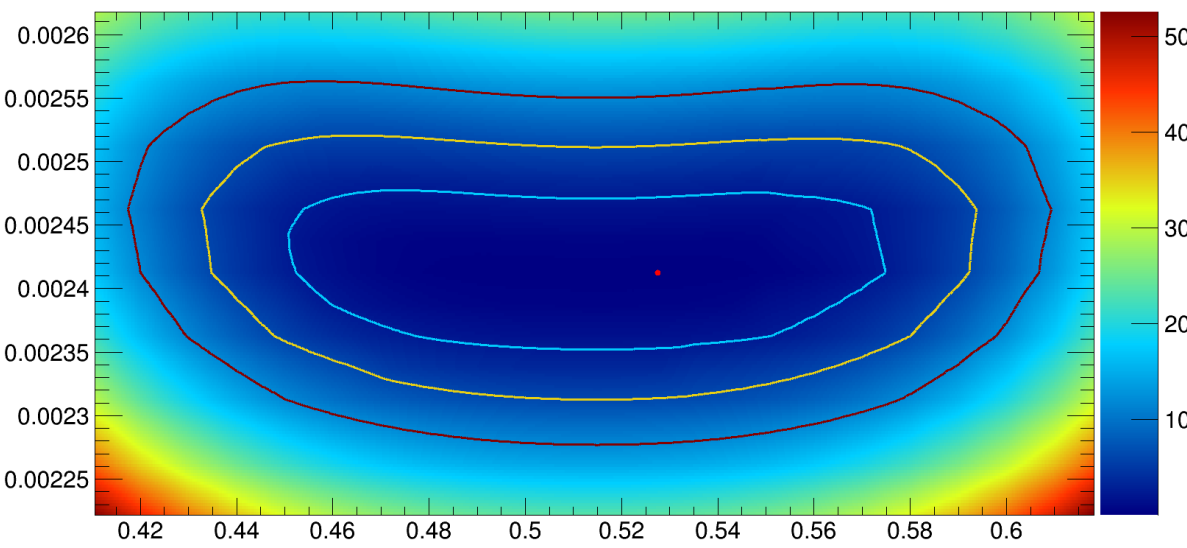
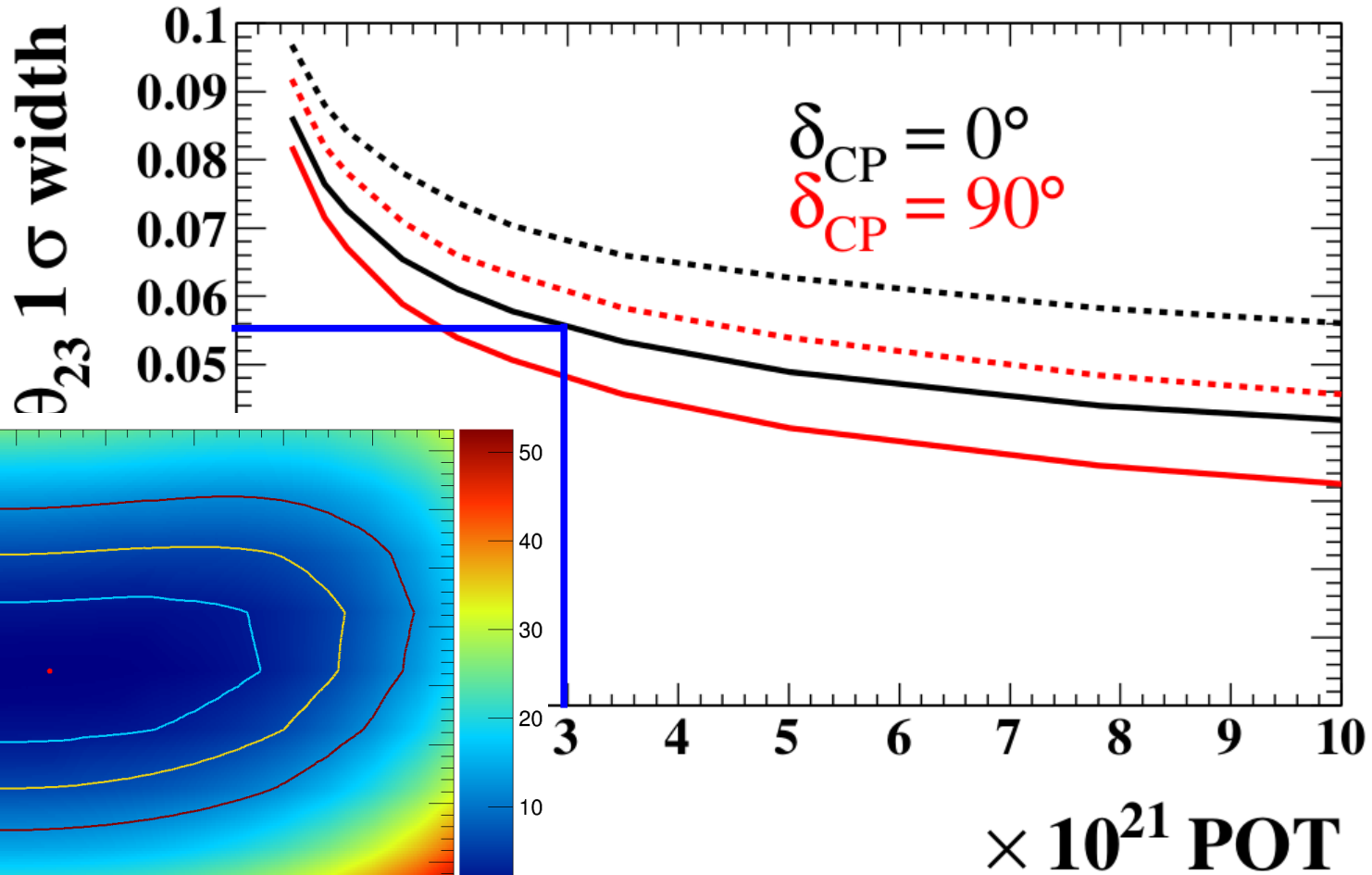
$\sin^2\theta_{23}$ 1 σ width



- nuPRISM naïve 1 sigma width = 0.062

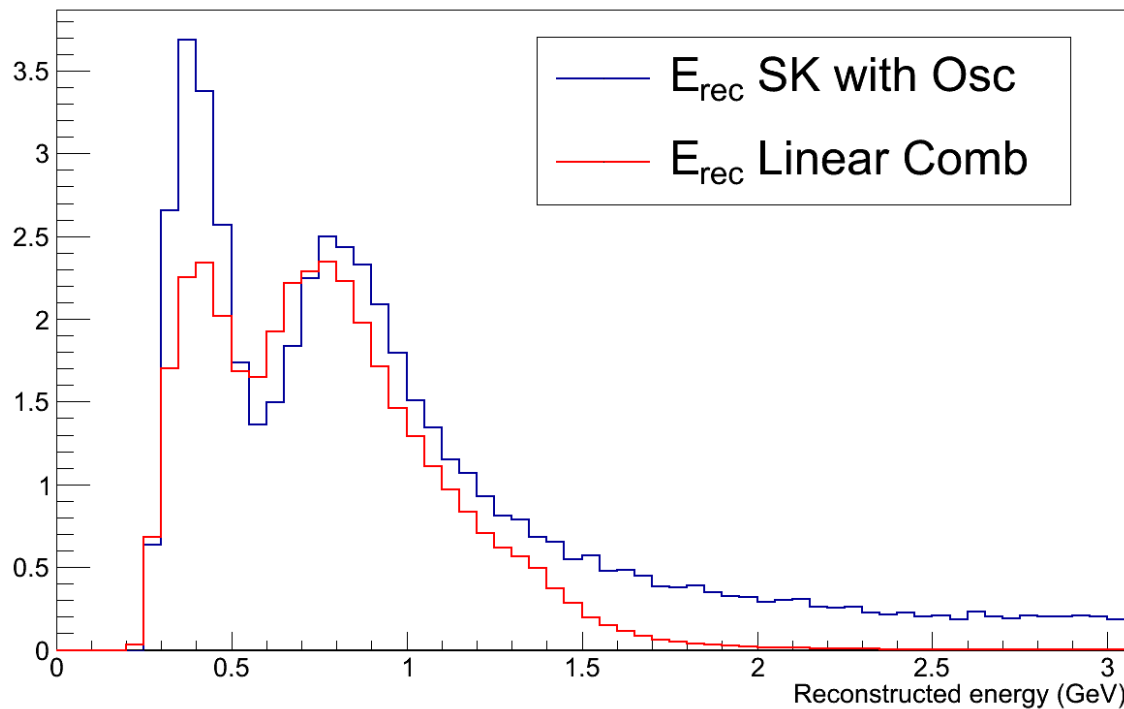
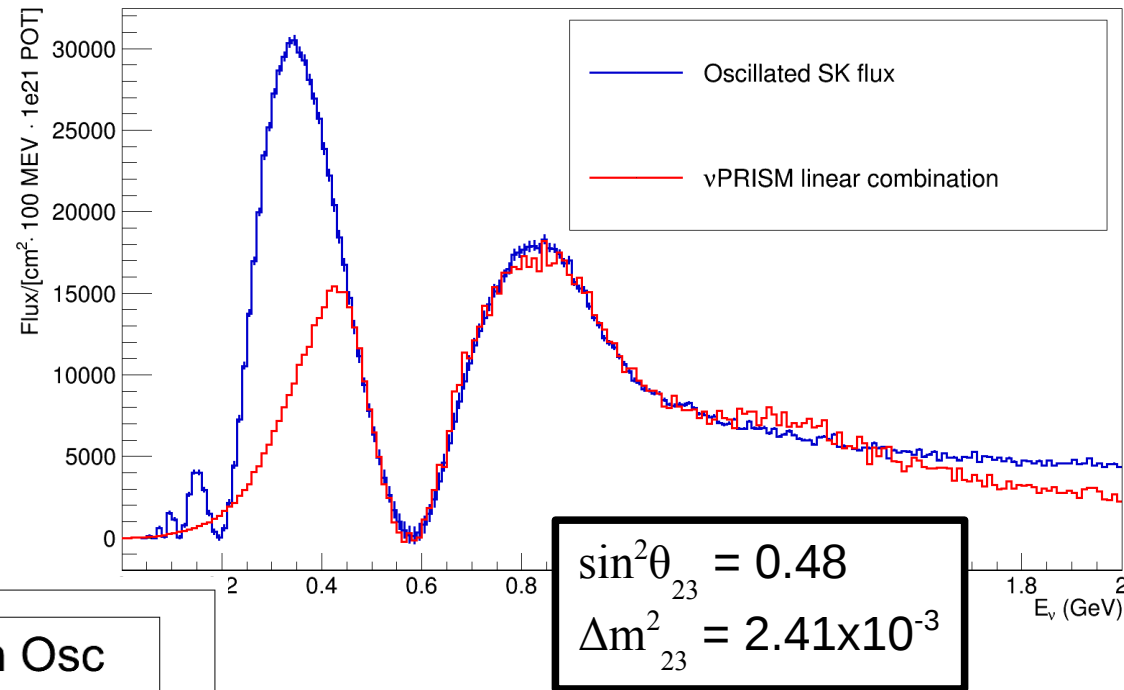
- Compare to T2K projection:

- 1 sigma width = 0.055



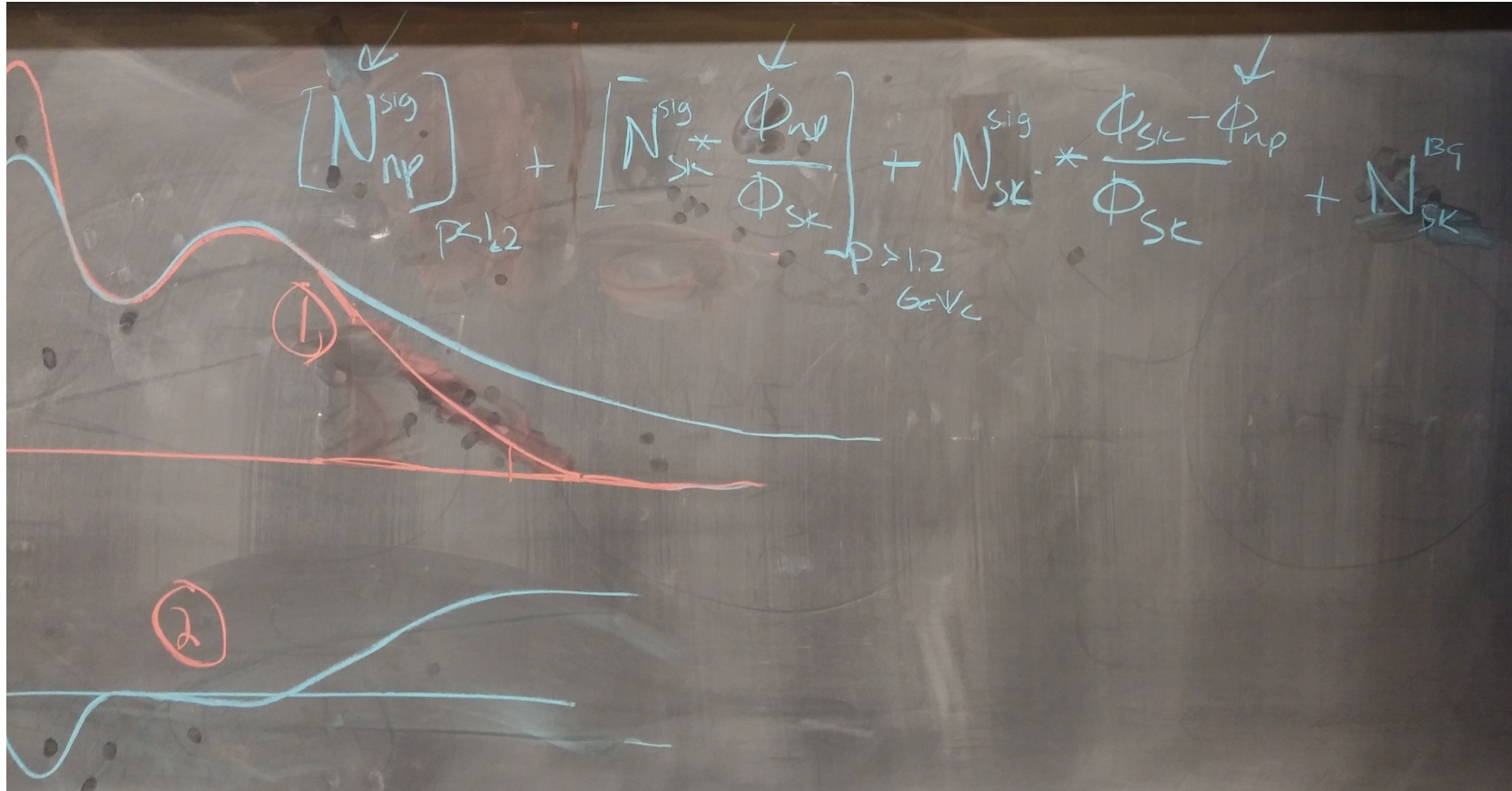
- Need to understand why the nuPRISM statistics only fit performs worse than the T2K statistics only fit

- Current flux fit matches SK flux very well around oscillation region
- Even so, there are differences in the reconstructed energy spectra

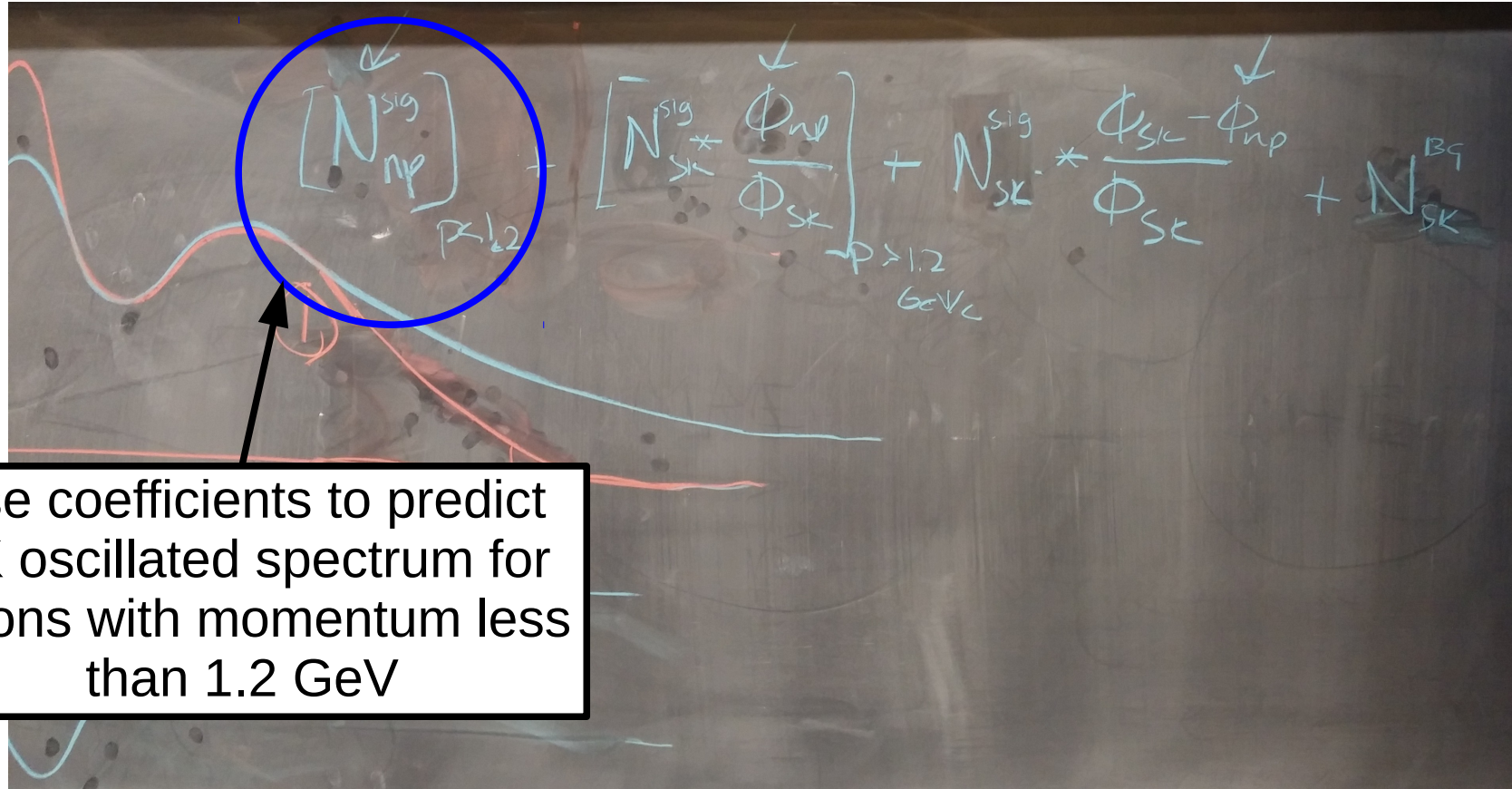


- Correct for these differences
 - Efficiency corrections
 - Add SK background
 - **Additive correction** – just brute force agreement

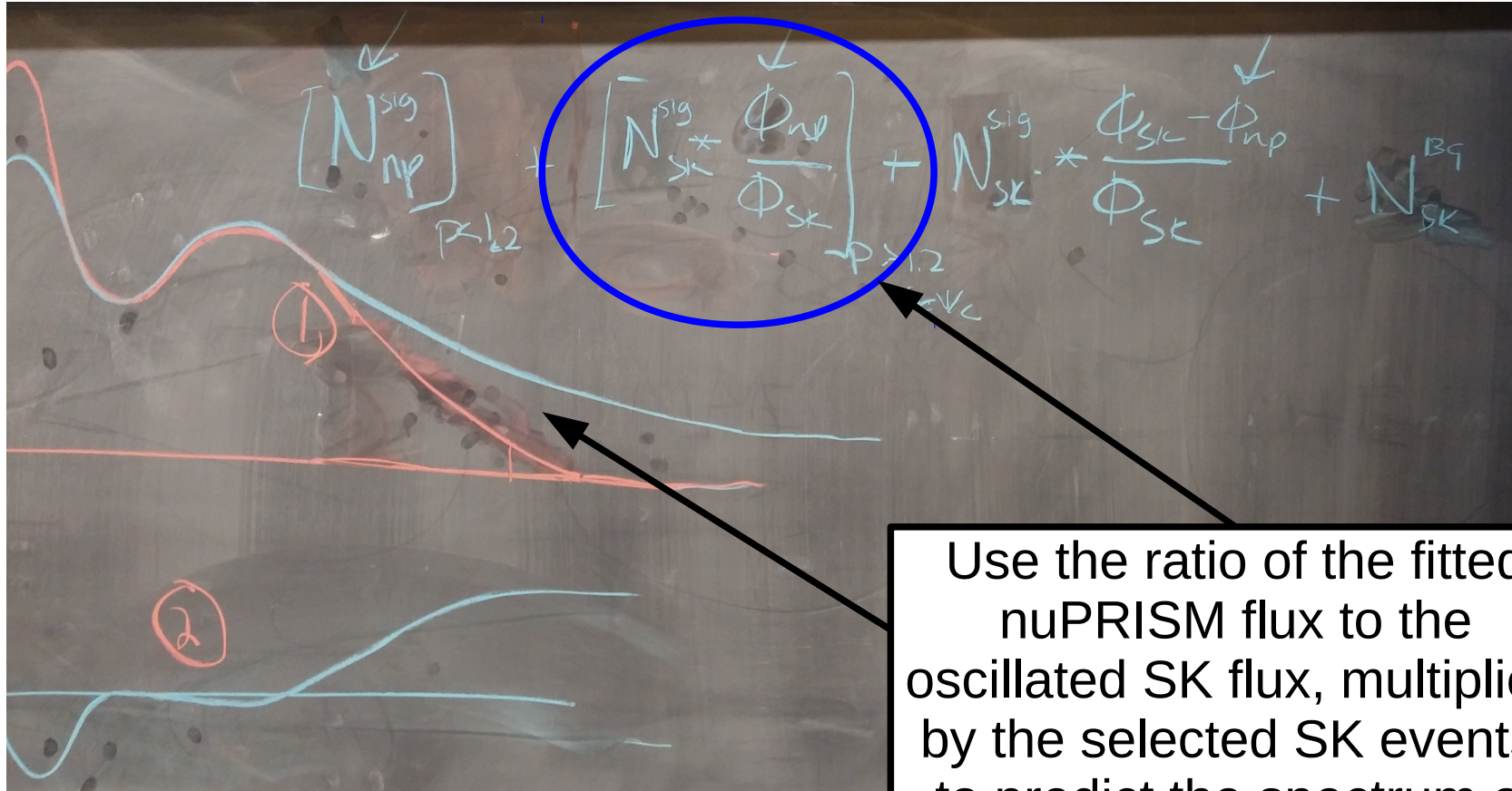
- Following M.Hartz's suggestion, want to try performing analysis in a different way



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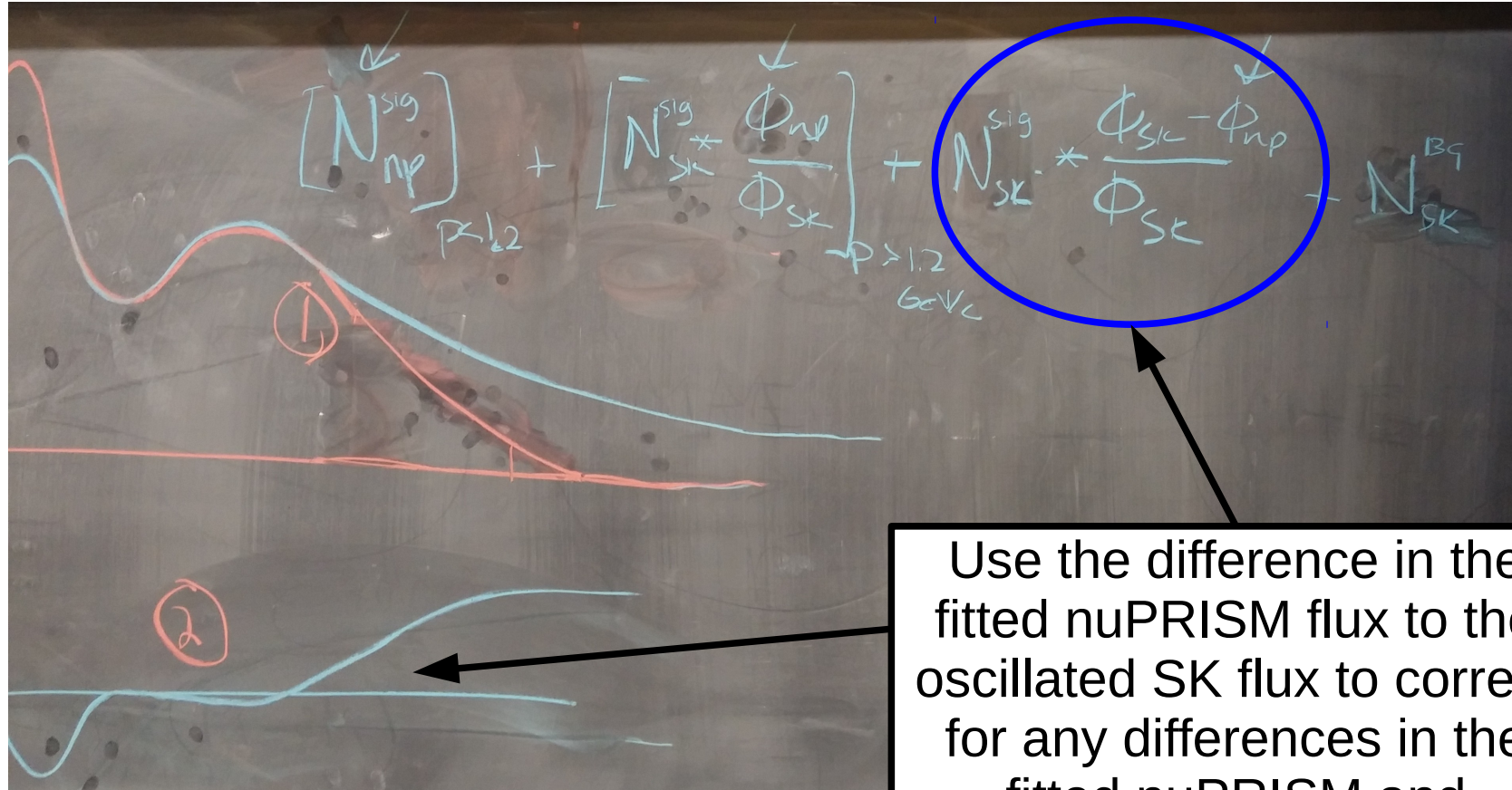


- Following M.Hartz's suggestion, want to try performing analysis in a different way



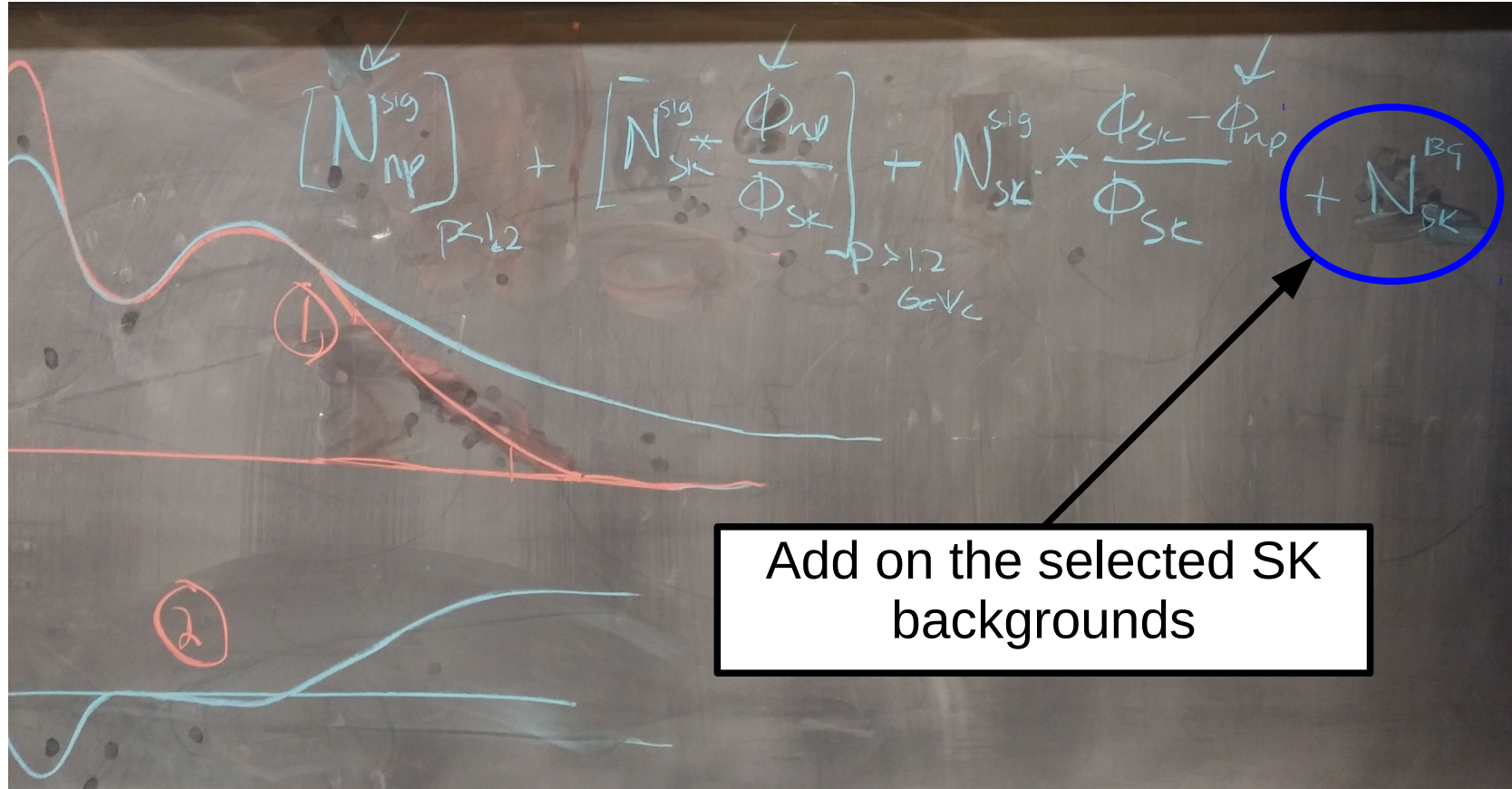
Use the ratio of the fitted nuPRISM flux to the oscillated SK flux, multiplied by the selected SK events, to predict the spectrum of muons above 1.2 GeV

- Following M.Hartz's suggestion, want to try performing analysis in a different way



Use the difference in the fitted nuPRISM flux to the oscillated SK flux to correct for any differences in the fitted nuPRISM and oscillated SK fluxes

- Following M.Hartz's suggestion, want to try performing analysis in a different way



Add on the selected SK backgrounds

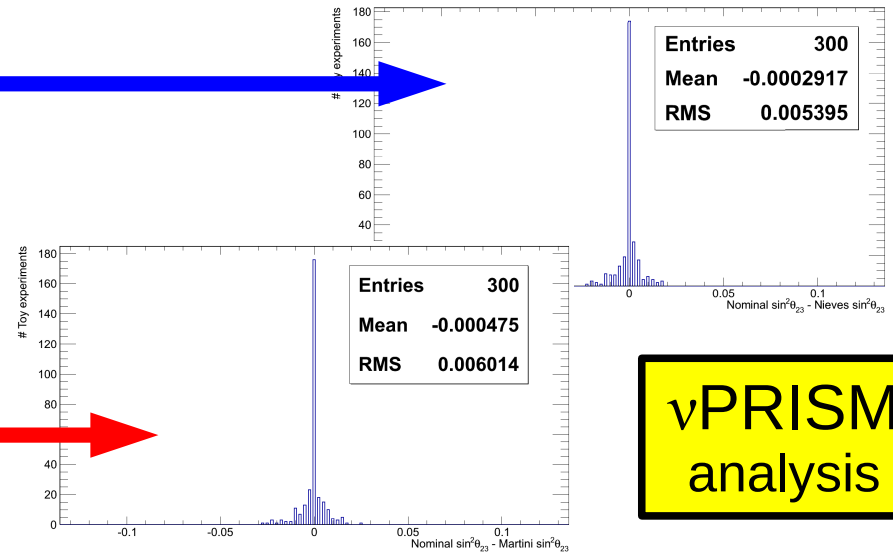
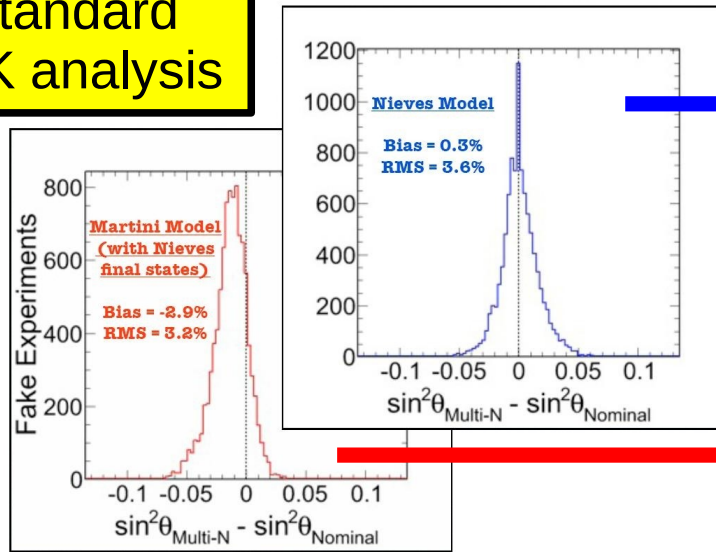
- Only have efficiency corrections for muons below 1.2 GeV, where nuPRISM can contain and reconstruct them
- No 'additive' correction – everything comes from the SK MC or flux

- nuPRISM is insensitive to mis-modelled nuclear effects in MC
- NuPRISM $\sin^2\theta_{23}$ sensitivity is as good as T2K, but should be better
 - Investigate analysis method
 - Examine each correction in detail
- Also want to produce T2K-like precision vs POT curves and sensitivity to non-maximal $\sin^2\theta_{23}$
- Quite a bit of work for 3 months!
- If anyone wants to get involved feel free to email/speak to me

ν_{μ} disappearance analysis status

- At 5th HK meeting - vPRISM disappearance analysis unaffected by unknown nuclear model
- Incomplete MC model did not change the fitted oscillation parameters

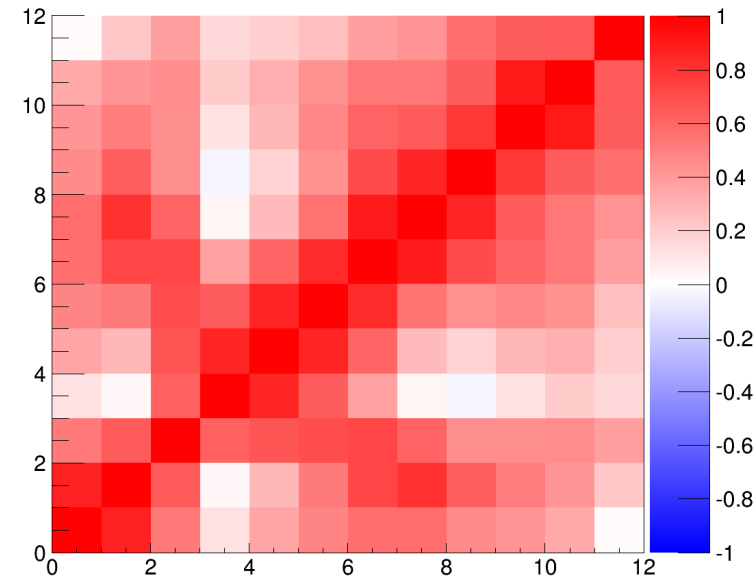
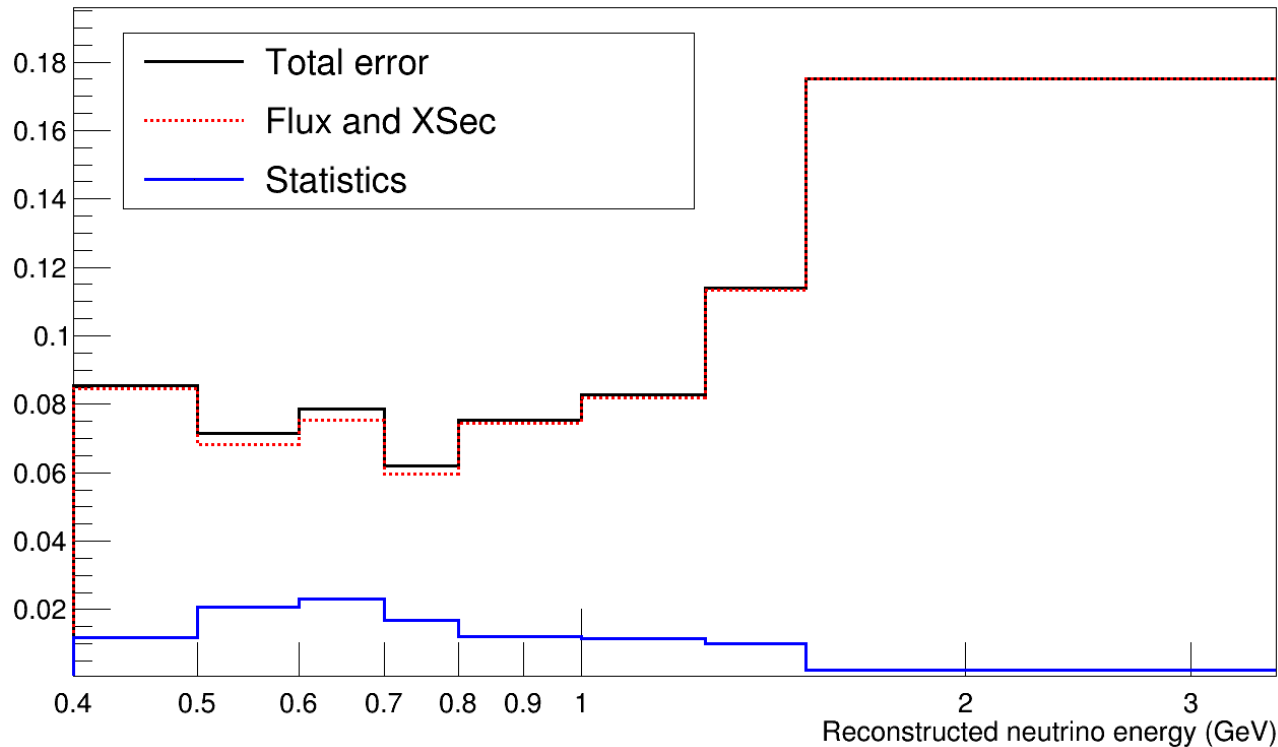
Standard T2K analysis



vPRISM analysis

- Bug found in statistical uncertainty calculation:
 - Bug scaled down number of events at vPRISM by factor of 40

- With new flux fit - total uncertainty in oscillation dip $< 8\%$

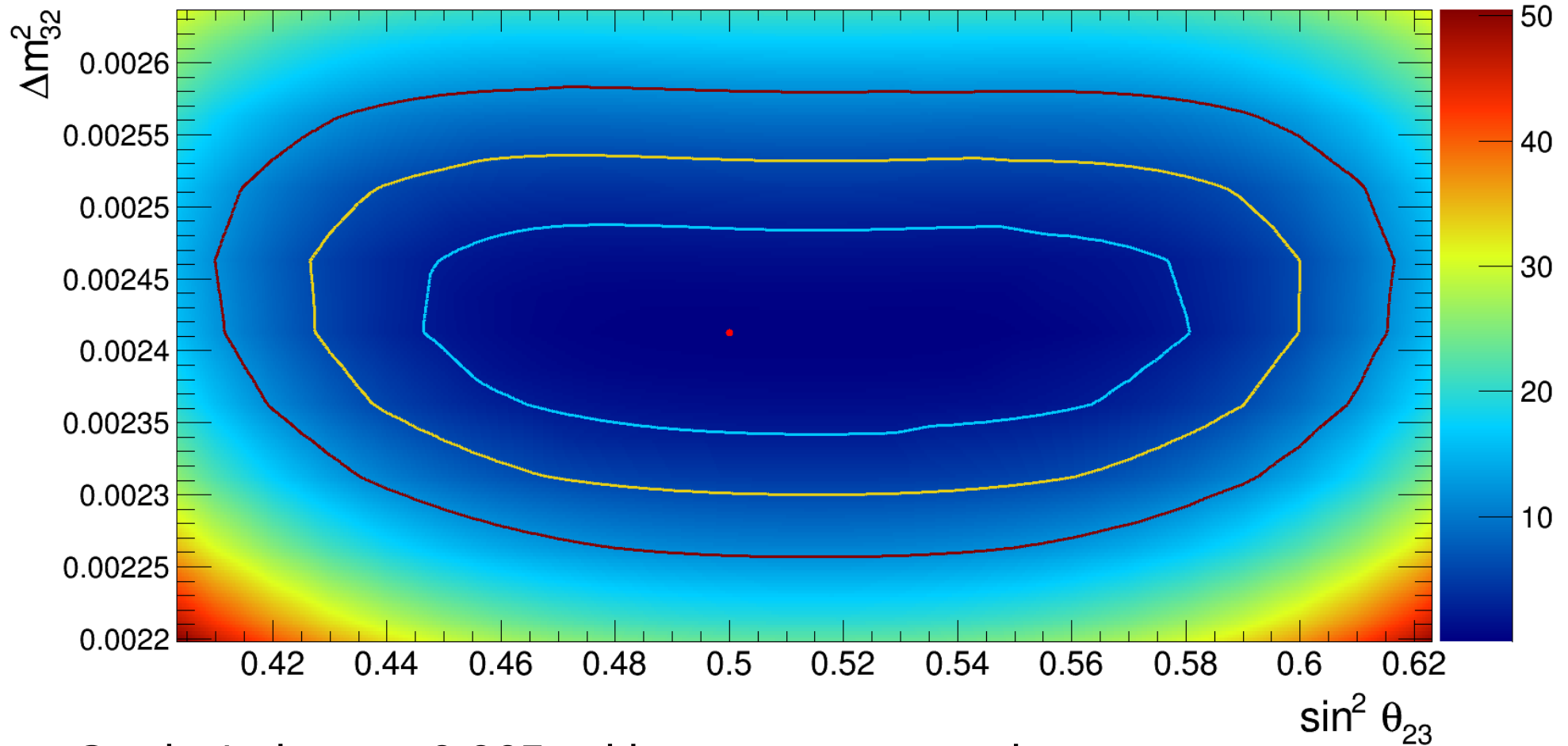


- Correlation matrix very similar shape to previous analysis
- Statistical uncertainty dropped, but flux uncertainty increased – this was expected (see backup slides)
- Now systematics (which are reducible) dominate

- With new flux fit examine sensitivity to disappearance parameters
- Use same methodology as Future Sensitivities Task Force (FSTF)
- For the Asimov data set (nominal MC) choose $dM^2 = 2.41e-3$ and $\sin^2 \theta_{23} = 0.5$ as the 'fake data'
- Perform grid scan over dM^2 and $\sin^2 \theta_{23}$ space, calculating the nuPRISM prediction and covariance matrix at each point
- Calculate the chi-squared value between prediction and the fake data at each point
- Use Delauney interpolation to smooth the delta chi-squared map this creates
- Create 1, 2 and 3 sigma contours

- In all plots there is $2.25e21$ POT of nuPRISM neutrino MC
- There are currently no SK uncertainties

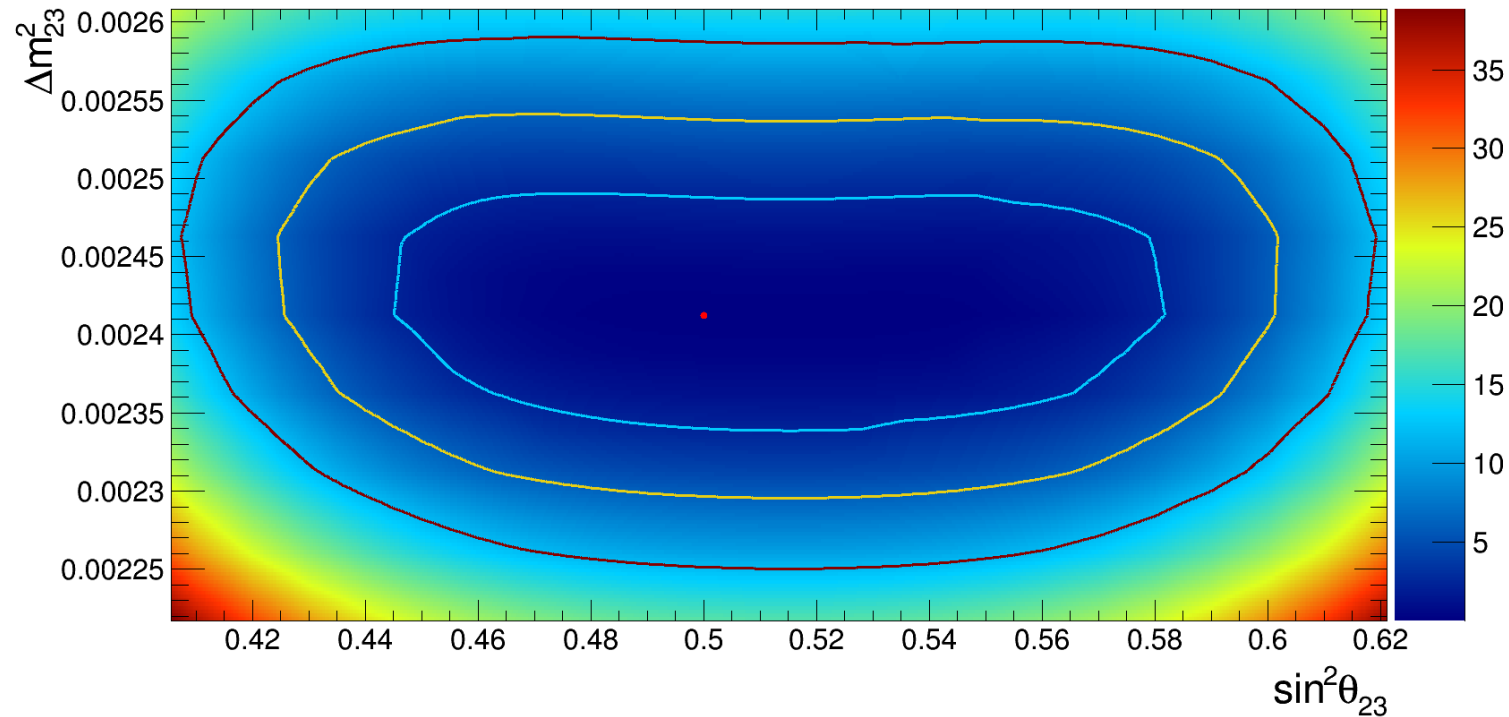
- For $3e21$ POT at SK – roughly the expected T2K POT
- $\sin^2 \theta_{23} = 0.5$, $\Delta m^2 = 2.41e-3$



- Crude 1 sigma = 0.065, with current systematics

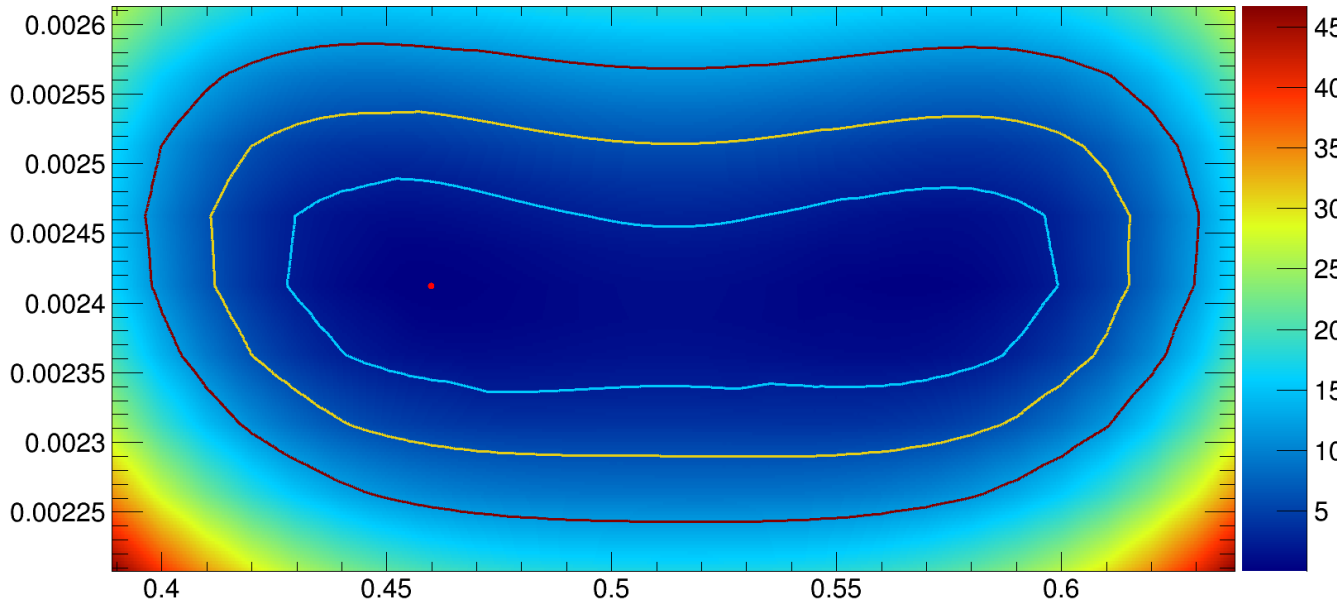
- Add in a 5% uncorrelated error to mock-up SK detector uncertainty
- $\sin^2 \theta_{23} = 0.5$, $\Delta m^2 = 2.41e-3$

- NuPRISM 1
sigma width =
0.0675
- NuPRISM = T2K
- SK detector
errors smaller
than this?
- Correlations?



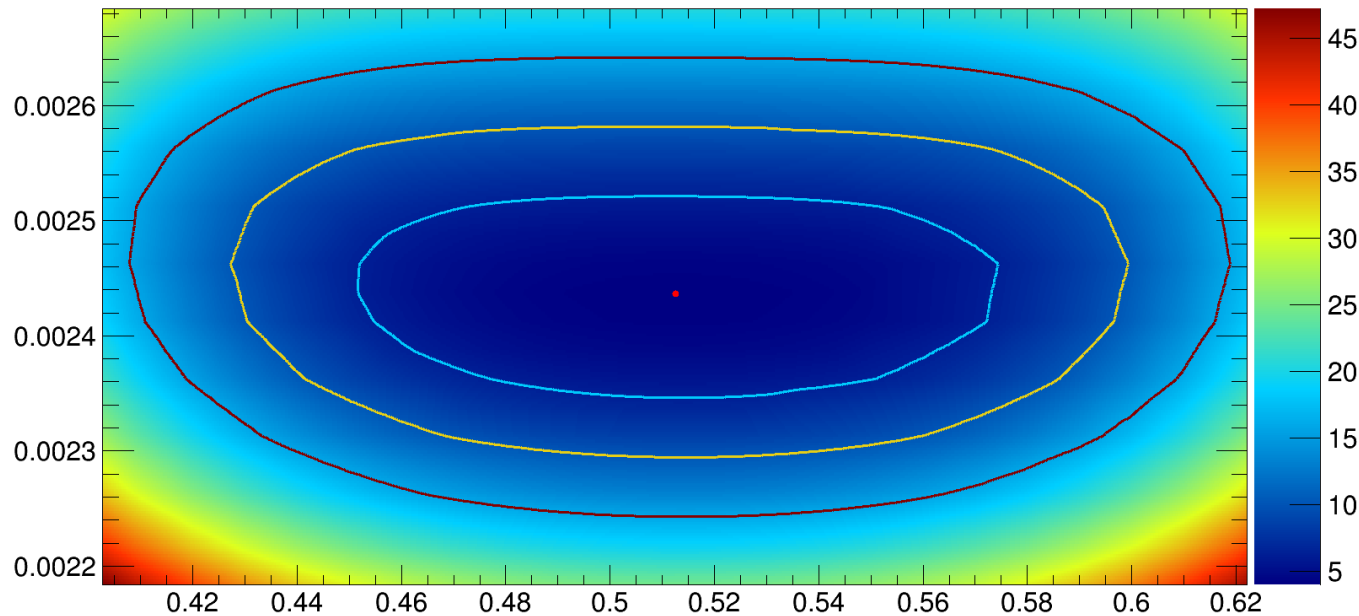
- Conservatively, NuPRISM analysis is equal to the current T2K projected sensitivities

- $\sin^2 \theta_{23} = 0.5$, now check at $\sin^2 \theta_{23} = 0.46$ (top)
- Using Asimov data set before, want to check what thrown data looks like

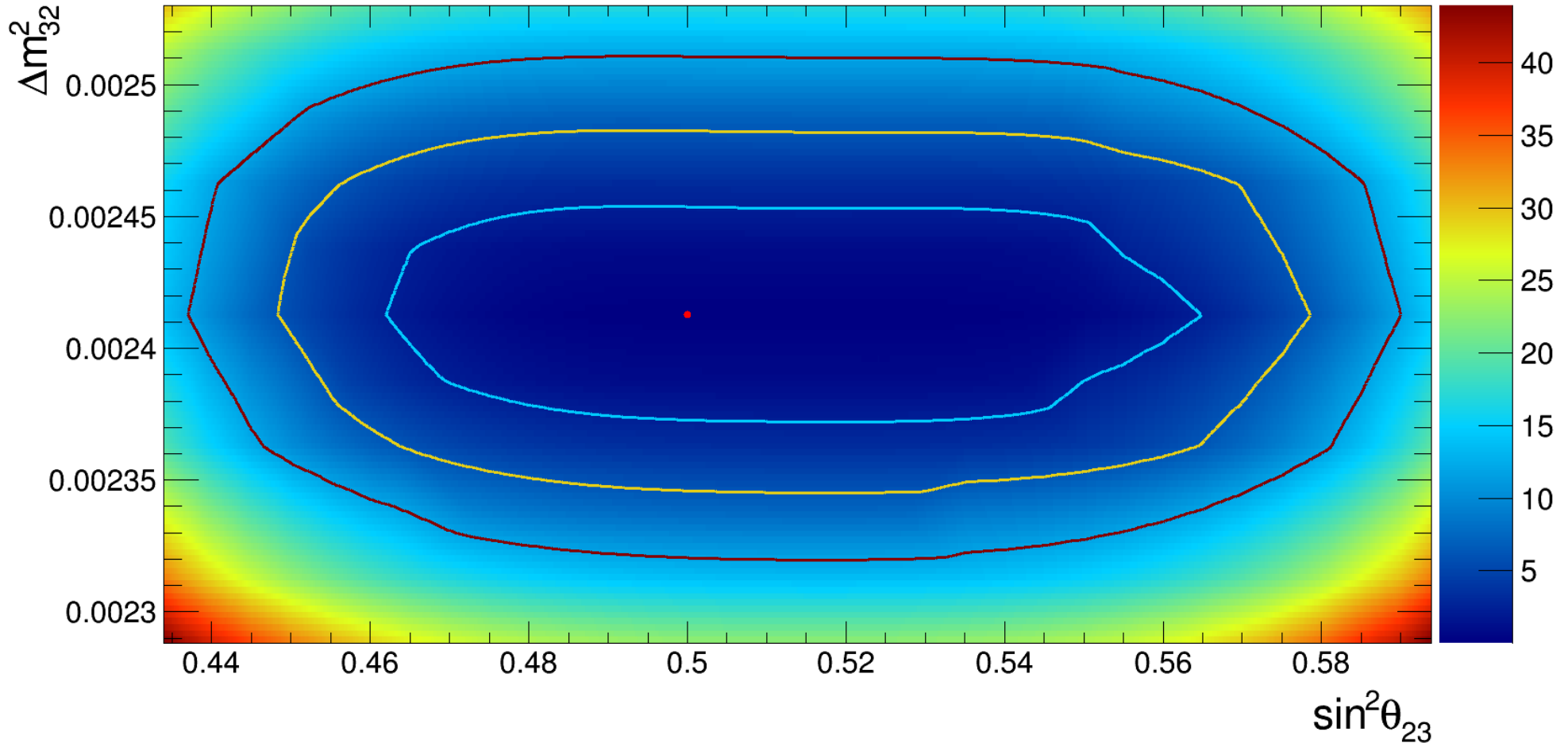


- See expected two-lobed structure when moving away from maximal mixing

- For a single throw, contours are compatible with Asimov data set



- $\sin^2 \theta_{23} = 0.5$, but at $1.56e22$ POT at SK
- Maybe \sim HK statistics?



- 1 sigma width < 0.050 , compared to ~ 0.056 at T2K

ν PRISM sensitivity summary and future work

- NuPRISM, with rough SK errors is as good as the T2K projected sensitivity
- Asimov sensitivities are compatible with thrown sensitivities, and different values of $\sin^2 \theta_{23}$
- Increasing statistics leads to reduction in 1 sigma width, at ~same rate as FSTF showed
- Still to do:
 - Convert 2D contours into 1D widths
 - Work out why stats only width is wider than FSTF stats only width
 - Do full SK systematics + HK statistics sensitivities
 - Octant sensitivity

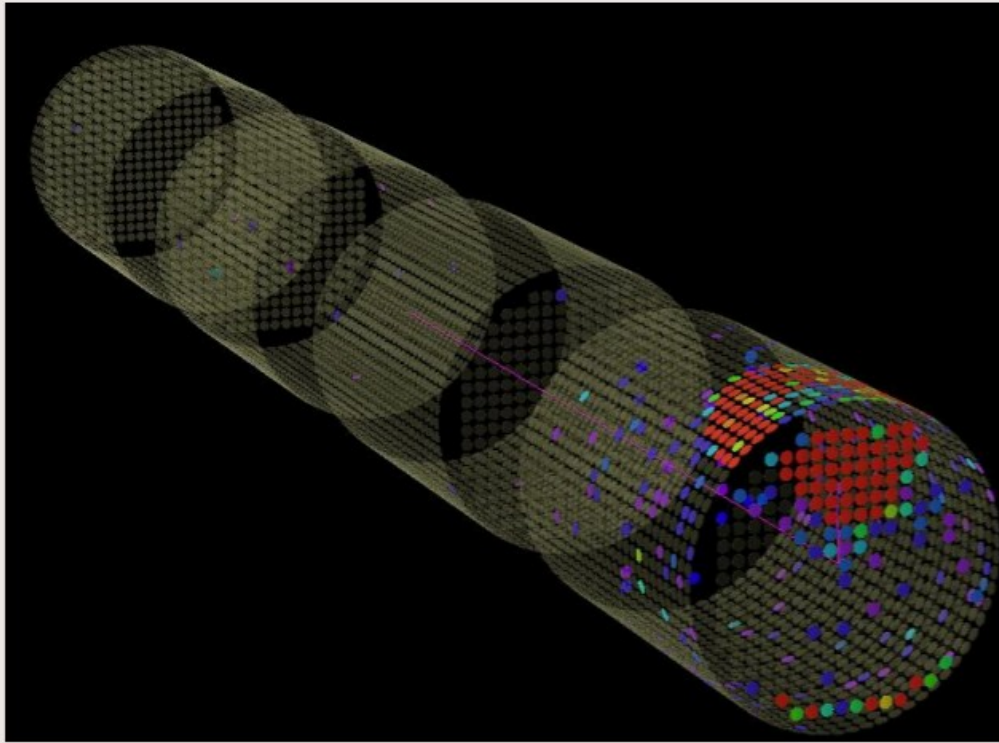
Future ideas

- Identify which flux systematics dominate – can these be reduced?
- Near detector constraint
 - Lets use them
 - Just take BANFF output – easy, but probably wrong
 - Joint fit? - Can't think how to combine with nuPRISM linear combinations
 - Use INGRID + ND280 to constrain flux
 - CC inclusive samples in both
 - INGRID = 0 – 0.8 degrees off axis
 - ND280 = 2 – 3 degrees off axis if we use ECal events
 - Good for T2K, cross section analyses

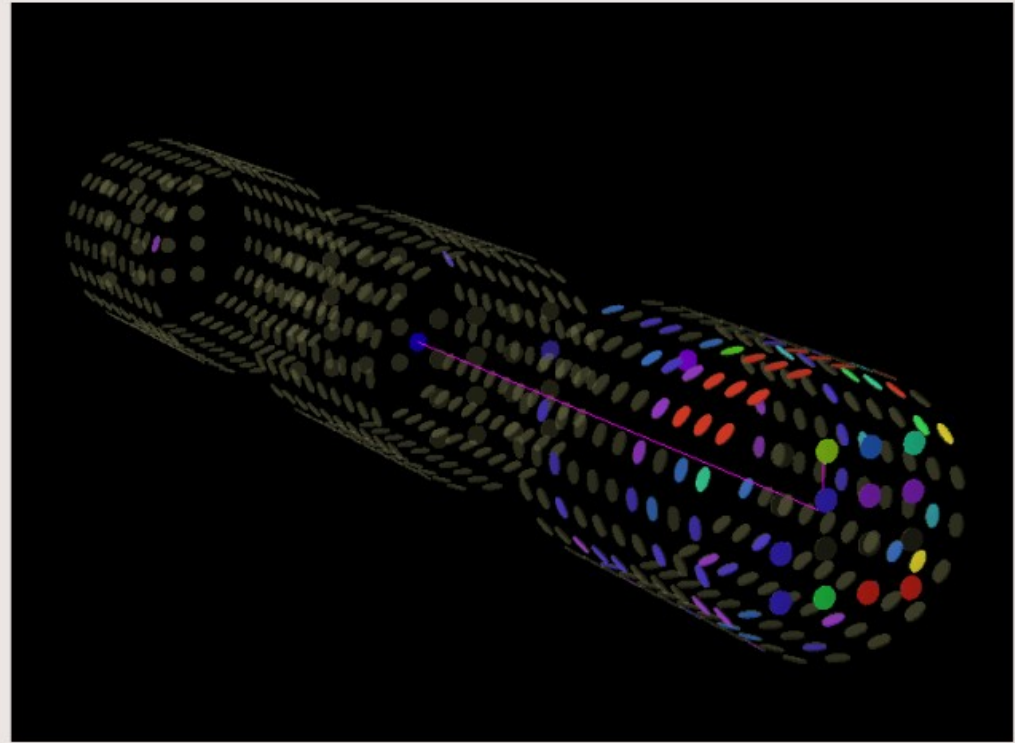
vPRISM simulation

- Working to create full detector simulation using WCSim and fiTQun
- Majority of work performed by two Co-op students at TRIUMF – Alex Lam and Carl Rethmeier
- Current status:
 - Variable vPRISM detector geometry present in older version of WCSim
 - Can vary detector size, PMT coverage, PMT type, number of compartments
 - Can generate fiTQun scattering table with any vPRISM geometry
 - Transformation from vPRISM coordinate system (z-axis along neutrino beam) to WCSim/fiTQun coordinates (z-axis along centre of detector)
 - Code to extract reconstructed and true information from fiTQun
- Have full software chain set up to create files for analysis

- Example of two extreme vPRISM configurations in WCSim



Radius = 0.127m (10inch diameter)
Percent Coverage = 60%
nSections = 4
SectionHeight = 8.0m
SectionGap = 4.0m



Radius = 0.254m (20inch diameter)
Percent Coverage = 20%
nSections = 3
SectionHeight = 10.0m
SectionGap = 7.5m

- fiTQun tuned to SK geometry in WCSim by Mike W.
- The WCSim SK tuned values were expected to be valid for all geometries - Super-K, Hyper-K, vPRISM, TITUS
- However, if PMTs are changed, retuning is necessary
- Need to solve reconstruction for real detector uncertainties!
- Mike studying the QE and angular acceptance of these PMTs to determine how different they are from the Super-K 20" PMTs
- Work ongoing, hopefully will get some manpower focussed on nuPRISM reconstruction:
 - Need to migrate WCSim changes to latest version
 - NuPRISM specific version of fiTQun?
 - Upload everything to GitHub
 - Scripts to run full chain and process

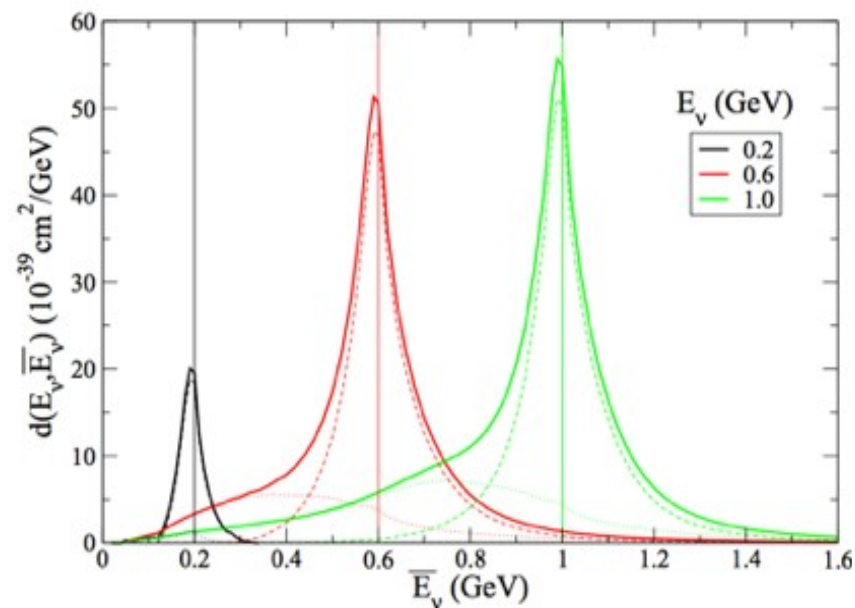
Summary

- Disappearance sensitivity studies are encouraging
 - Soon will have much more stringent analysis
 - Appear to be more capable than ND280 fit
- Full simulation and reconstruction needed to estimate detector systematics and reconstruction ability
 - More work to start soon
 - Want to see how reconstruction is affected, hopefully improved

Backup slides

Motivation for mono-chromatic Beams

- The modelling of multi-nucleon reactions, pion absorption, the nuclear initial state, etc., introduce uncertainties on:
 - The absolute normalization of the cross section for CC events with only visible leptons
 - The relationship between the lepton (or other final state) kinematics and the neutrino energy (important for oscillation measurements)
- Measuring the effect of nuclear effects on the final state kinematics is challenging in a conventional beams due to the width of the neutrino spectrum
- Ideally, a monochromatic neutrino beam would allow one to study how nuclear effects contribute to the final state particle distributions
- We can make “mono-chromatic” neutrino beams in nuPRISM



Martini et. al.
Phys.Rev. D87 (2013) 013009

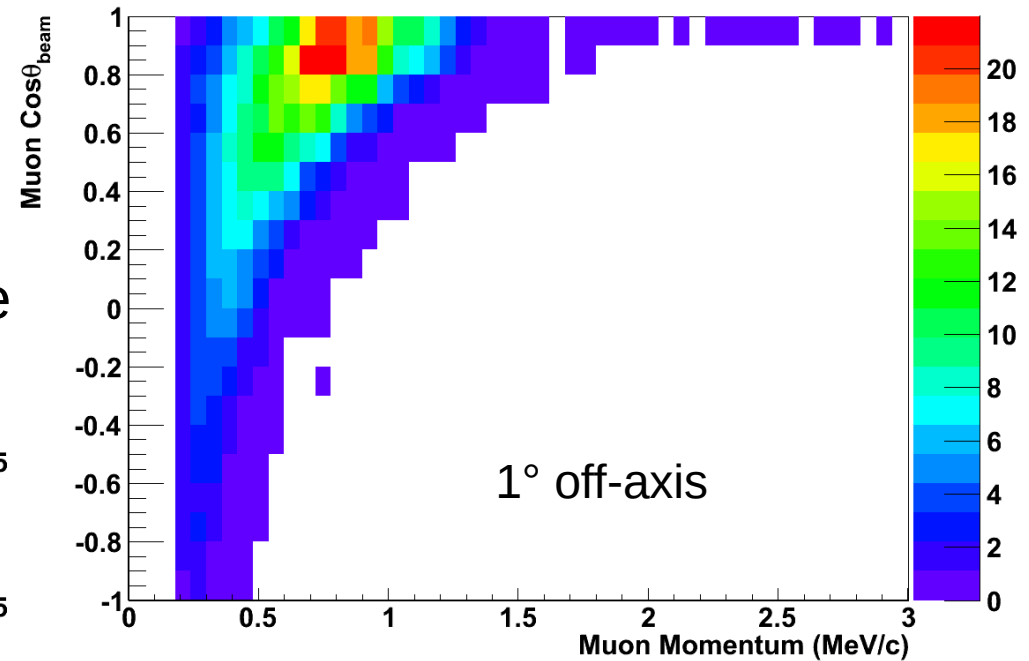
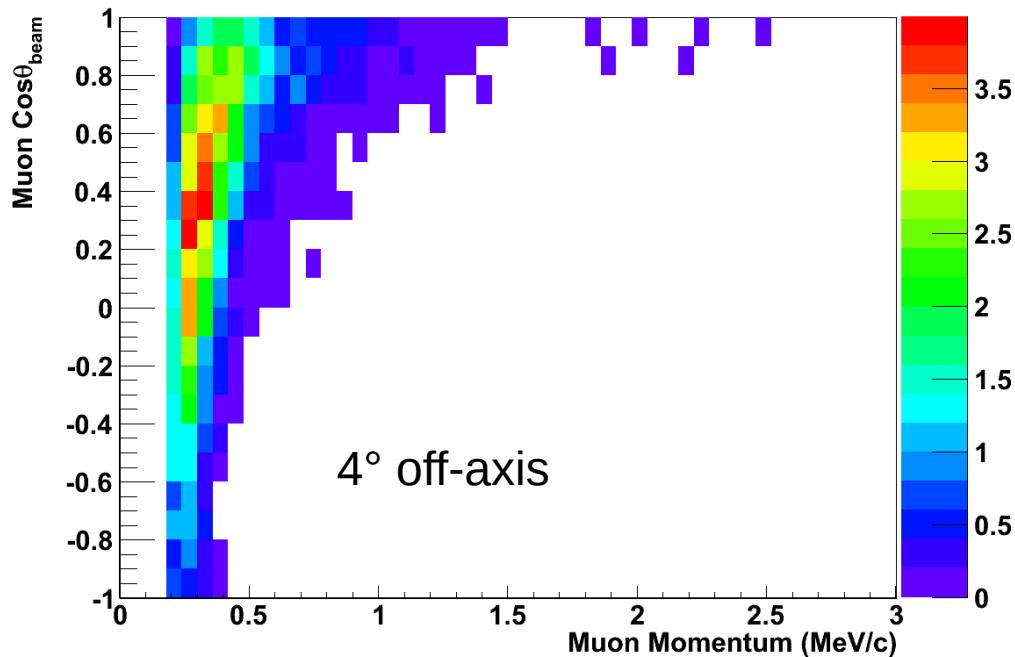
Mono-chromatic Beams with nuPRISM

- Using the linear combination method, we can produce Gaussian beams with widths significantly less than an off-axis spectrum peaked at the same energy

$$G(E_\nu; \mu, \sigma) = \sum_{i=1}^{\text{\# of Off-axis bins}} c_i \varphi_i(E_\nu)$$

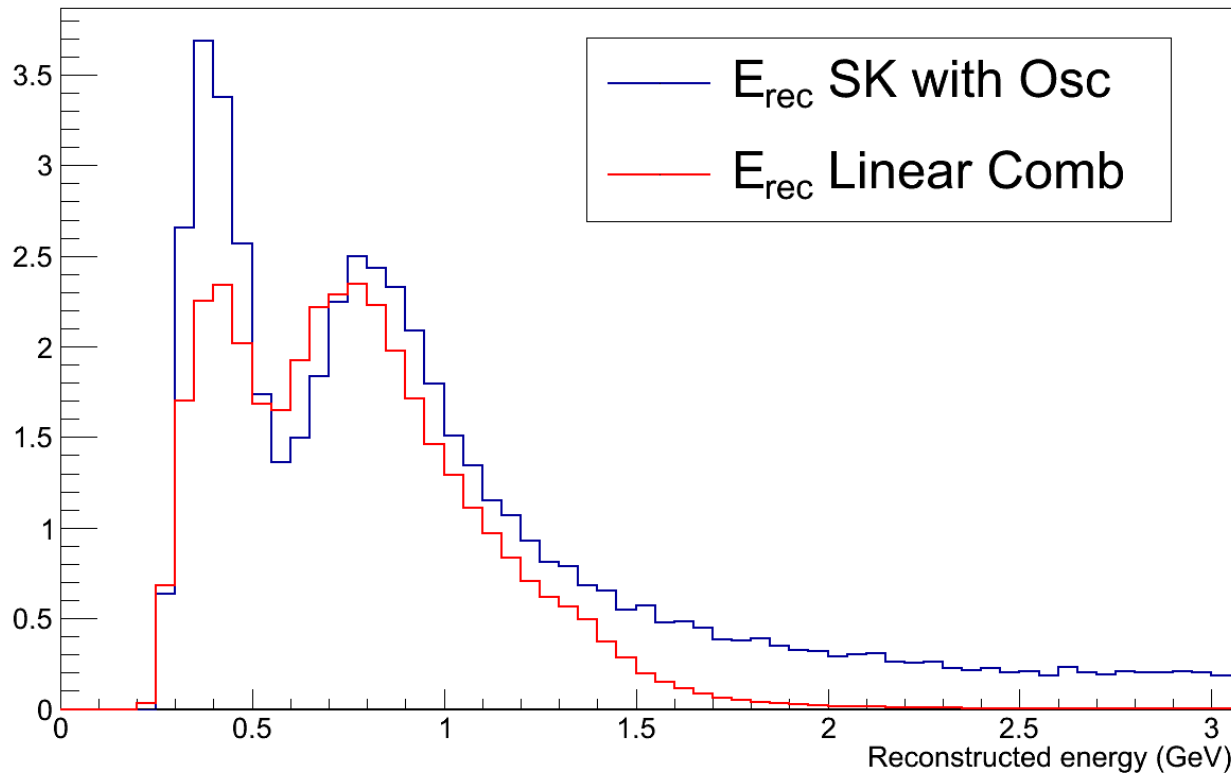
- Here the c_i are chosen to give the desired mean μ and width σ of the Gaussian
- In practice, the range of μ that can be achieved is limited by the range of peak energies in the off-axis fluxes that nuPRISM observed, $\sim 0.4\text{-}1.2$ GeV
- The width of the mono-chromatic beam, σ , is limited by the level of statistical and systematic error that can be propagated in the linear combination

- Same event selection as at SK:
 - Single ring
 - Muon-like
 - Fully contained in fiducial volume



- Record the off-axis angle of the interaction, using the reconstructed vertex position

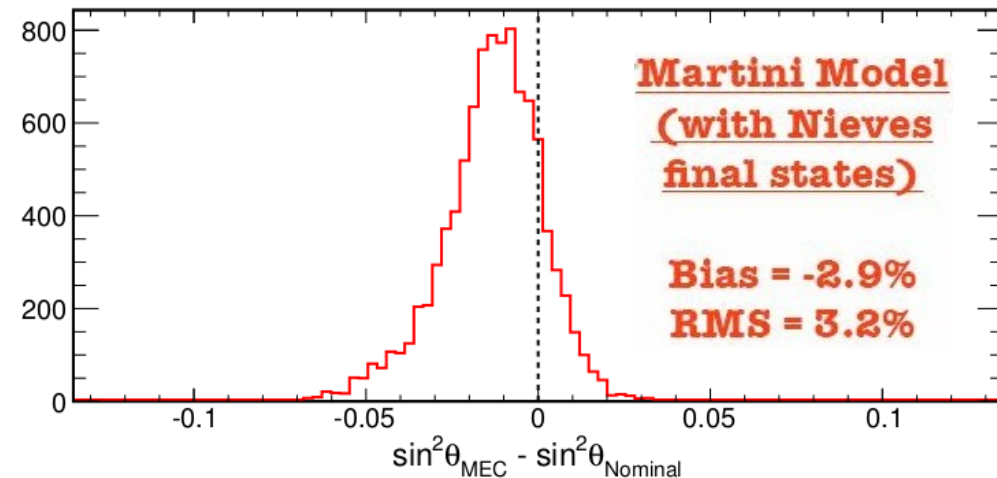
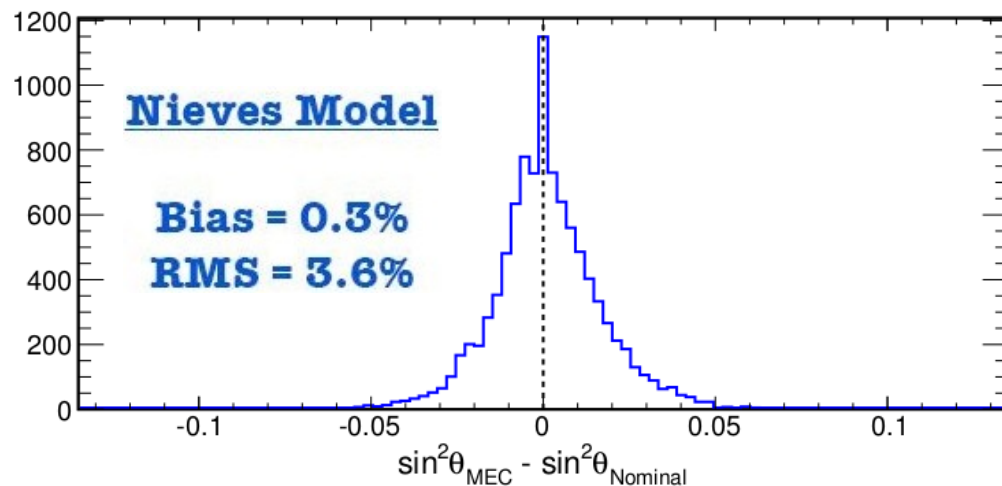
- Apply these weights to the selected events in each off-axis slice of ν PRISM
- Now looking at reconstructed neutrino energy - events smeared into oscillation dip by nuclear effects and energy resolution



- To ν PRISM data:
 - Background subtraction
 - Efficiency correction
 - Addition of selected SK background
- Introduce some model dependence

- Every correction made to the ν PRISM prediction is calculated from our nominal MC – all are constant corrections
- To calculate systematic uncertainties:
 - Apply a variation to the ν PRISM and SK MC
 - Changes number of selected events at both detectors
 - Apply corrections (from the unvaried, nominal MC)
 - Calculate difference between the ν PRISM prediction and the varied SK MC
 - Use this to calculate fractional covariance matrix for ν PRISM prediction
- This analysis takes flux and cross section uncertainties into account
 - Initial detector systematics being studied

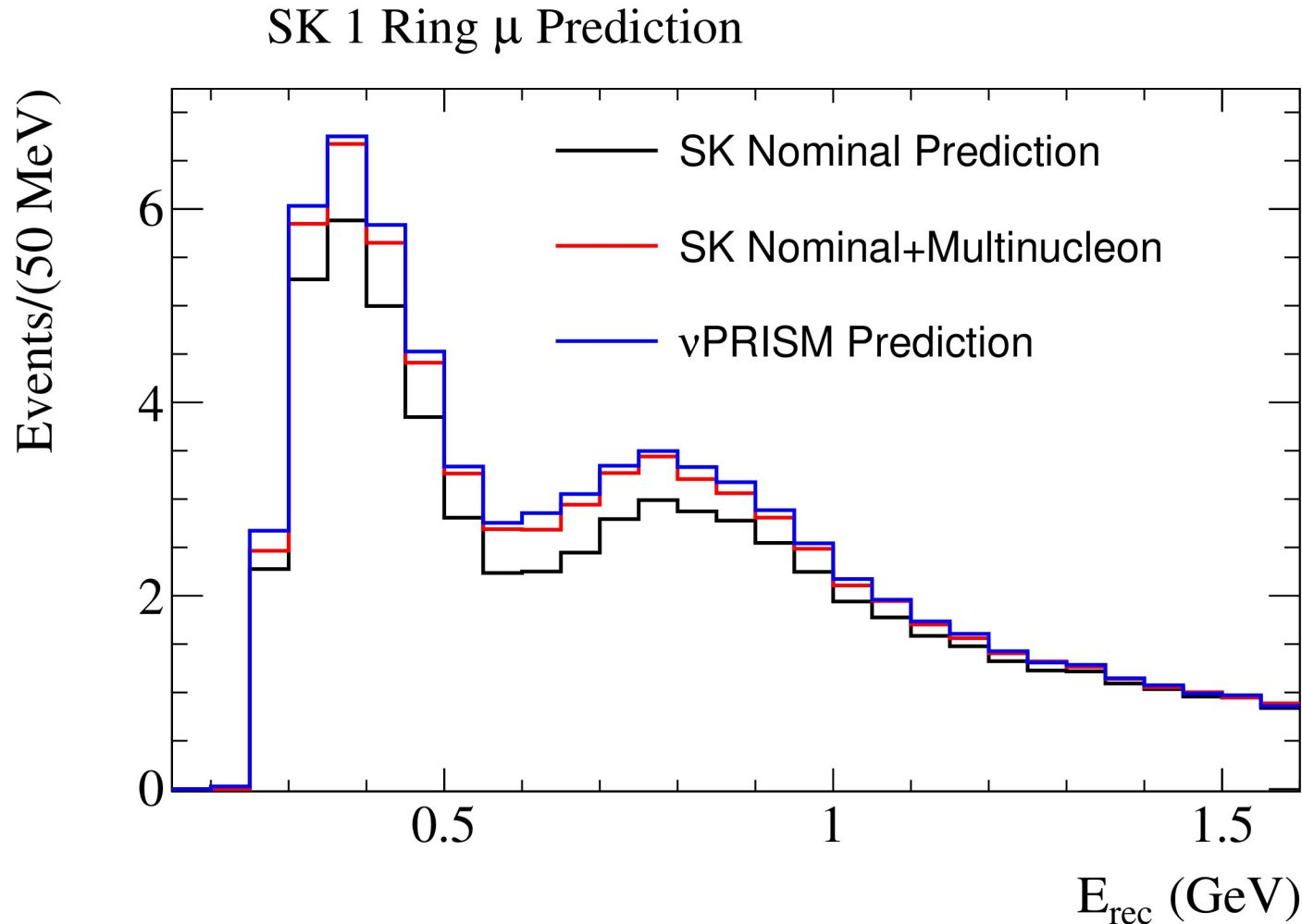
- MC-based analysis using full detector simulation, full systematics etc.
- Three fake datasets
 - Nominal NEUT MC
 - NEUT + meson exchange current (MEC) events from [Nieves' model](#) - [Phys. Rev. C, 83:045501, Apr 2011](#)
 - NEUT + MEC events based on [Martini's model](#) - [Phys. Rev. C, 81:045502, Apr 2010](#)
- Perform disappearance fit to extract θ_{23} in each case and compare



- Both models give $\sim 3.5\%$ RMS in $\sin^2\theta_{23}$, Martini model introduces $\sim 3\%$ bias
- Effects much smaller than current statistical uncertainty, but maybe large for future analyses

Multi-Nucleon example

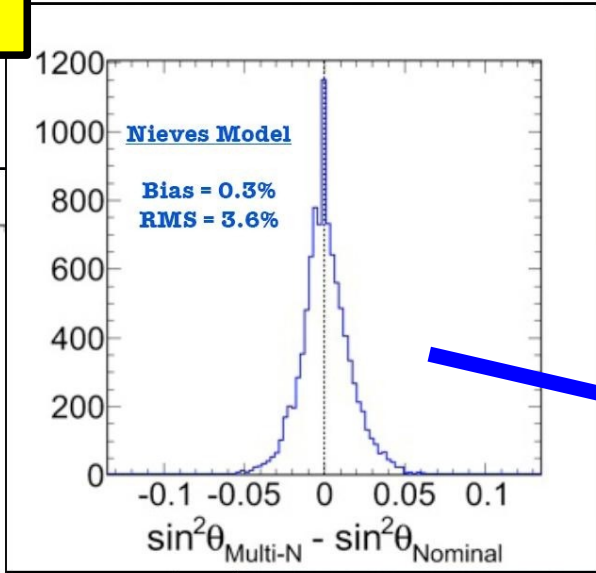
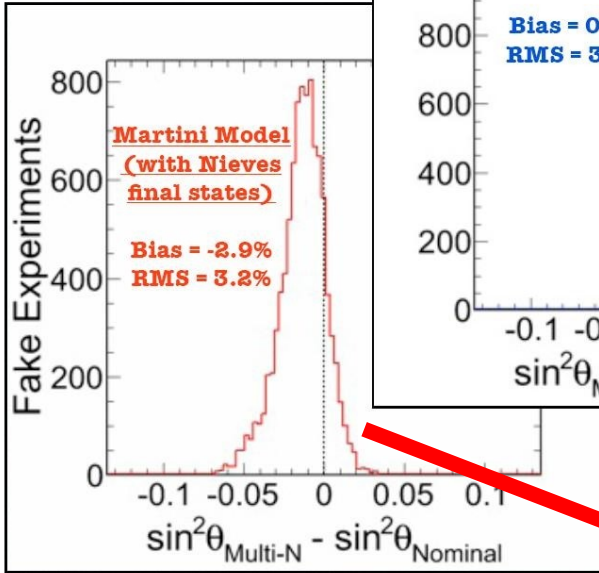
- Add multi-nucleon events to the nominal MC to make fake data



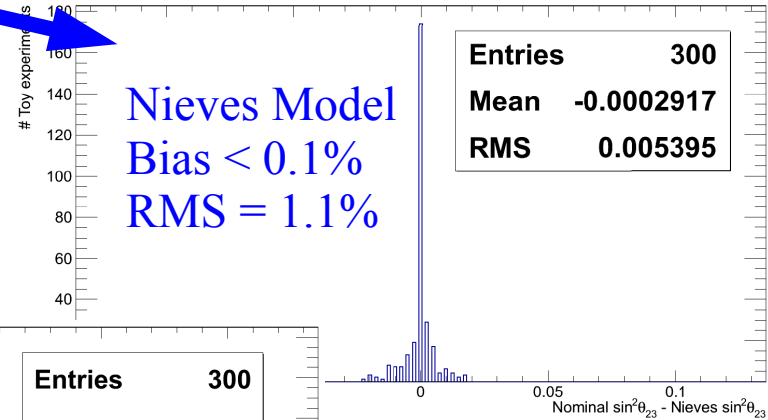
- See vPRISM prediction still reproduces oscillated SK spectrum when multi-nucleon events are present

Effect of multi-nucleon events at vPRISM

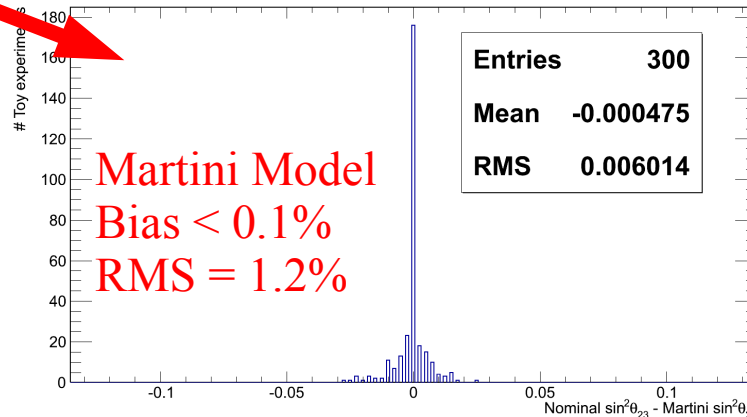
Standard T2K analysis



- Performed same MEC study as T2K
- Bias and RMS greatly reduced

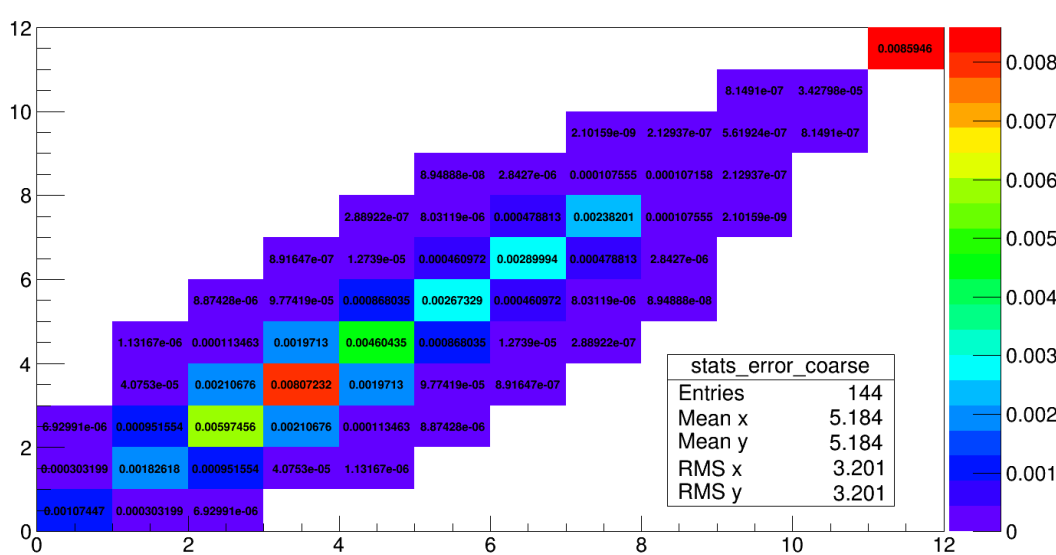


- vPRISM analysis largely independent of cross section model

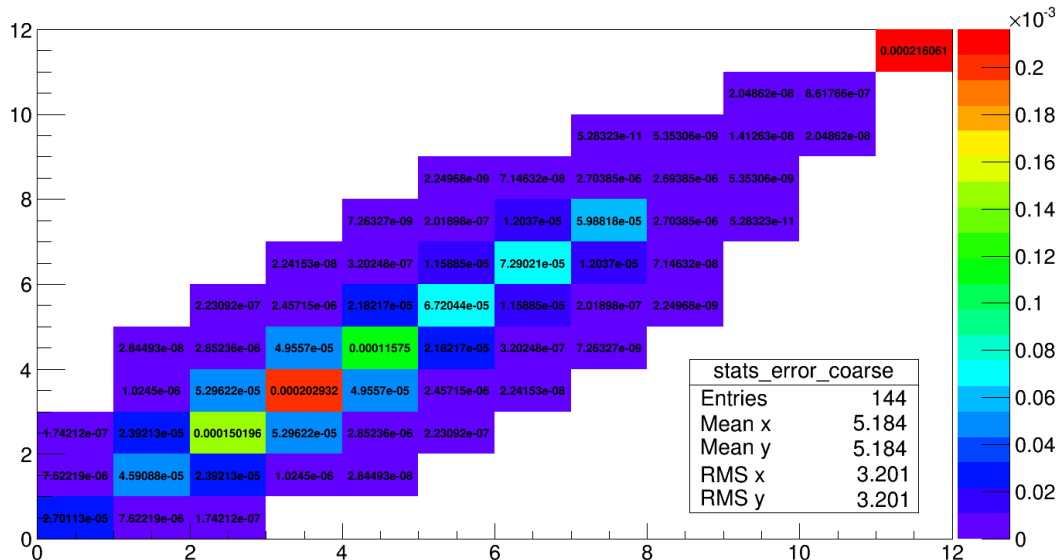


vPRISM analysis

- Ran analysis code with 'flux_mass_correction_factor' set to its current value and set to 1

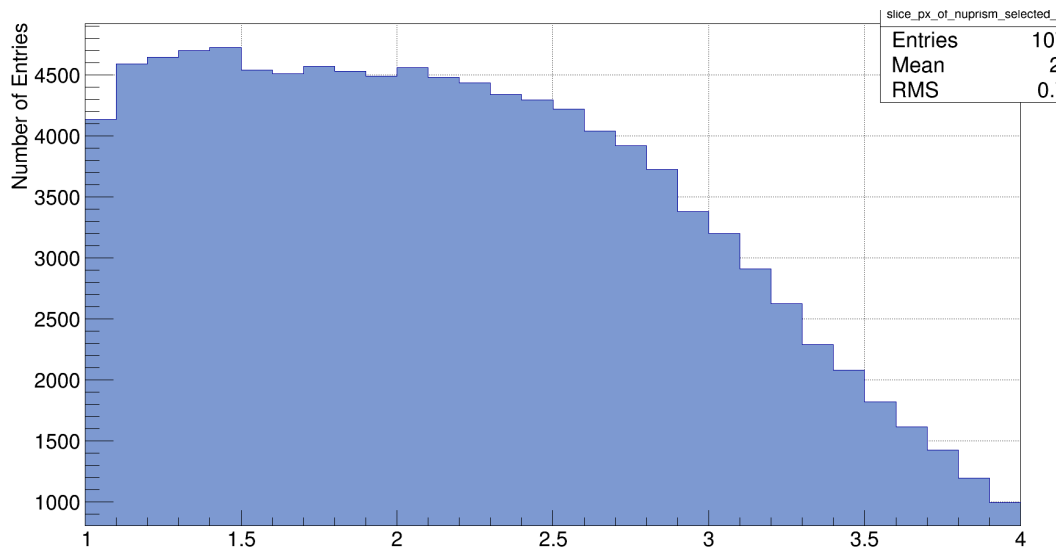
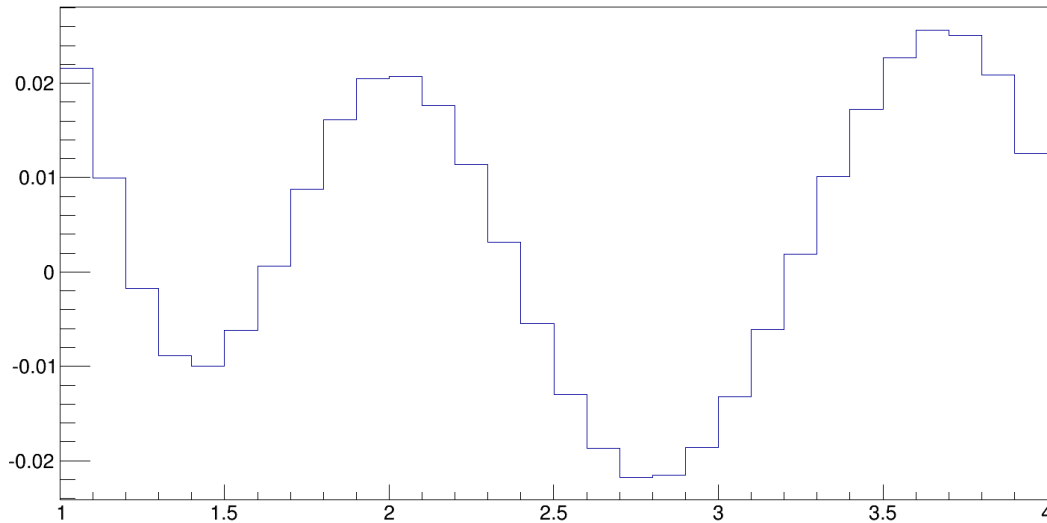


- Statistical uncertainty covariance matrix
- Buggy code on top, fixed code beneath
- In buggy code see statistical uncertainty of 8.8% in largest bin



- Roughly same size as in analysis presented at last T2K meeting
- Difference due to different oscillation parameters
- In fixed code, uncertainty $\sim 1.4\%$
- $1.4 * 6.3 = 8.8$, as expected

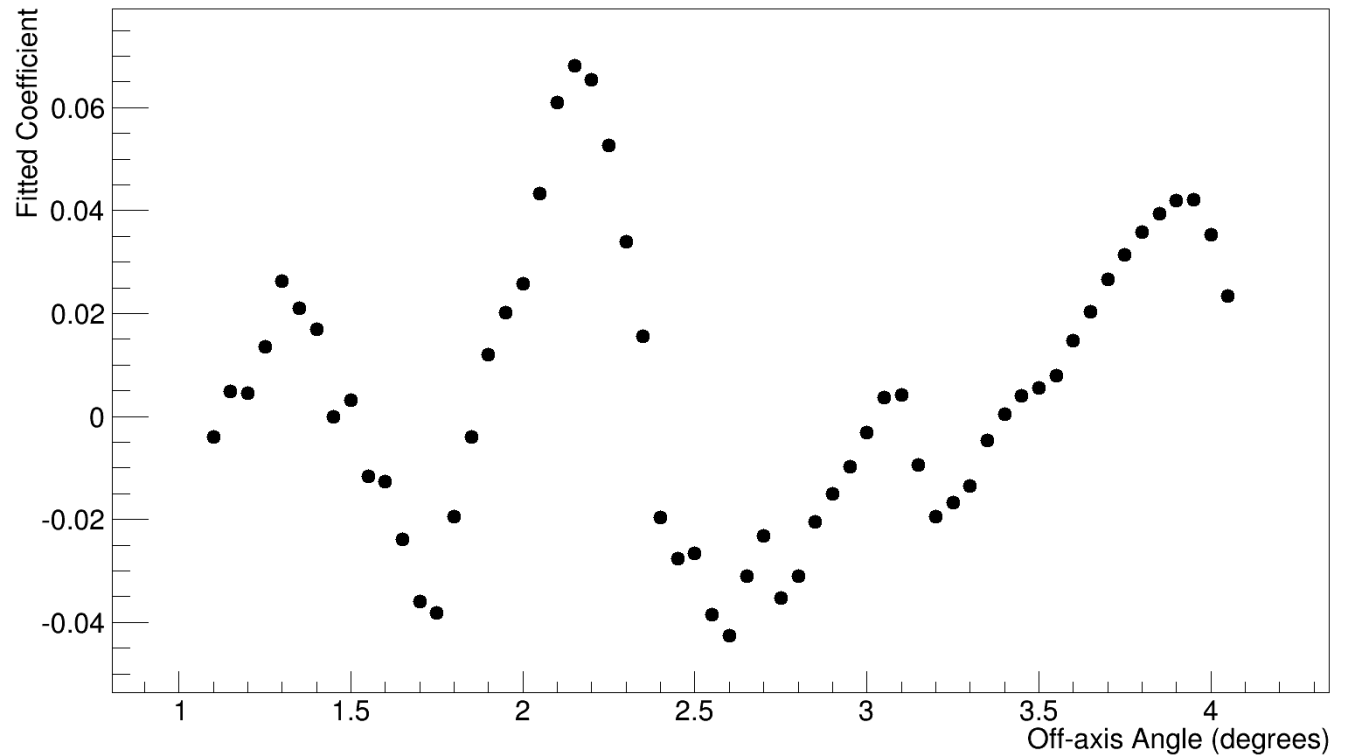
- Calculated naïve statistical uncertainty for 500-600MeV reconstructed energy bin by hand



- Flux fit coefficients versus off-axis angle on top
- Selected nuPrism events versus off-axis angle on bottom
- $N_Events = \text{Sum over oaa bins} (N_{Selected} * \text{Coefficient})$
- $Error^2 = \text{Sum over oaa bins} (N_{Selected} * \text{Coefficient}^2)$
- $N_Events = 220$
- $Error = 4.8$
- $\text{Fractional error} = 2.2\%$

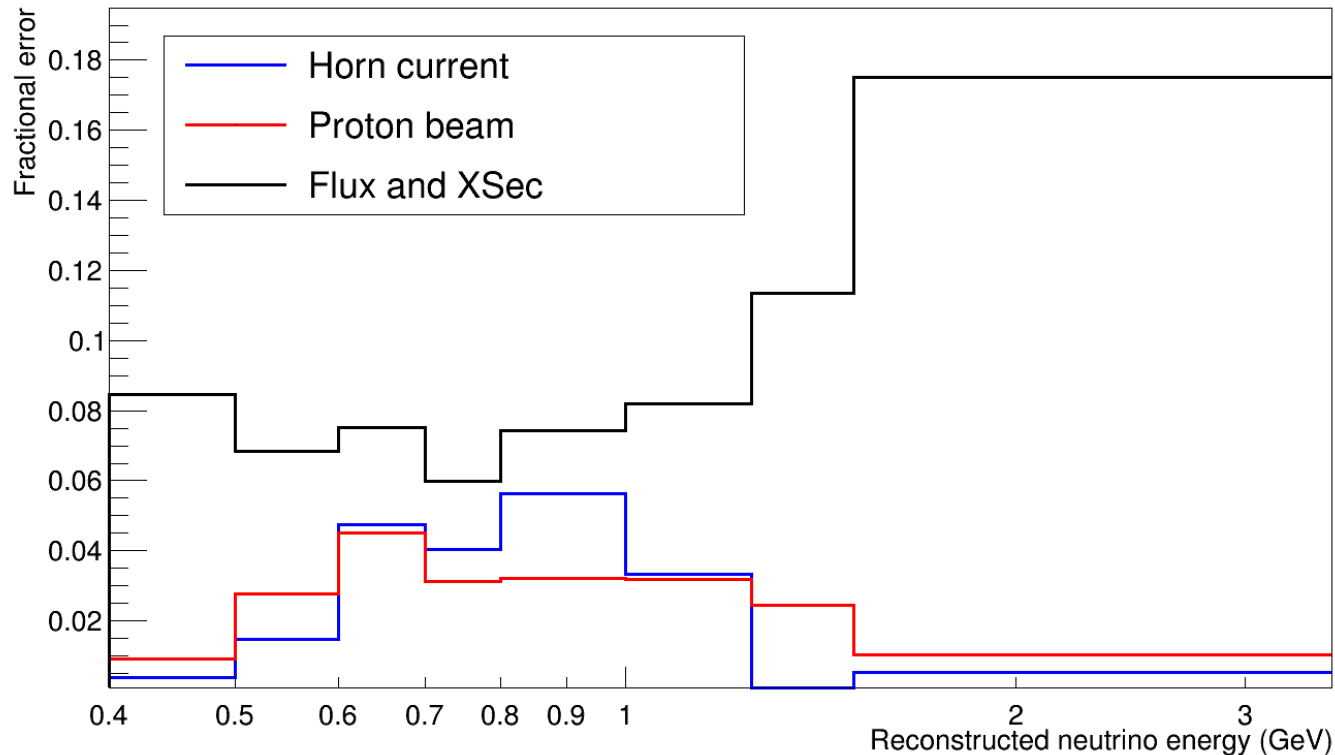
- Calculated naïve statistical uncertainty for 500-600MeV reconstructed energy bin by hand - 2.2% fractional statistical uncertainty
- Also calculated same uncertainty using c++ – almost exact agreement with above results
- Analysis code gave uncertainty of 1.2% for this bin
 - Analysis applies corrections, so do not expect identical numbers
 - Only change to code was to remove scaling number – all previous statistical variation validation (Poisson throws) still valid

- Coefficients not as smooth as before
- Built on previous work to improve fit



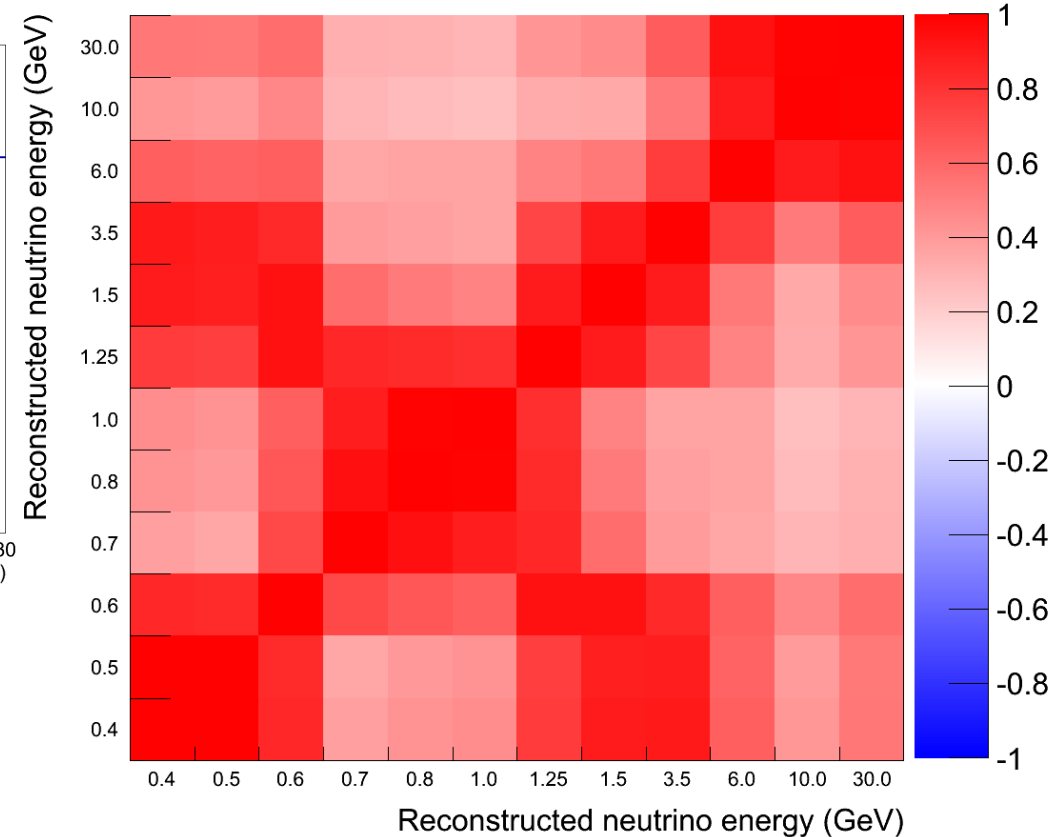
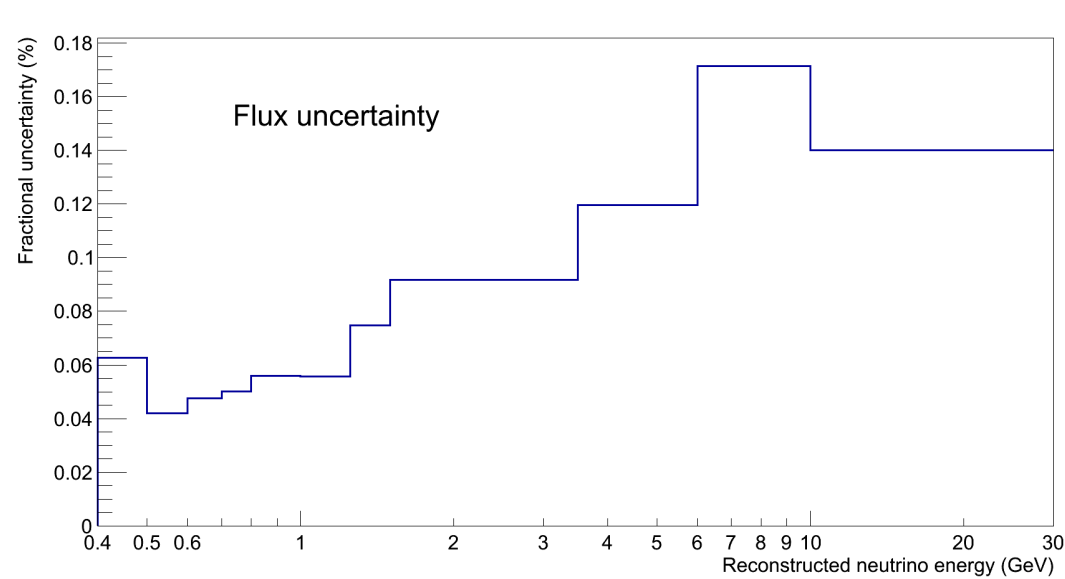
- Fit from 450MeV – 1200MeV
- Use 10MeV bins in chi-squared calculation
- Use 60 off-axis parameters
- Smooth with denominator = 0.01

- Flux uncertainty increased from 5% to 7% at oscillation dip
- Largely driven by proton beam and horn current uncertainties



- Not unexpected:
 - More variation between neighbouring coefficients
 - Systematics that effect off-axis angle flux have bigger effect

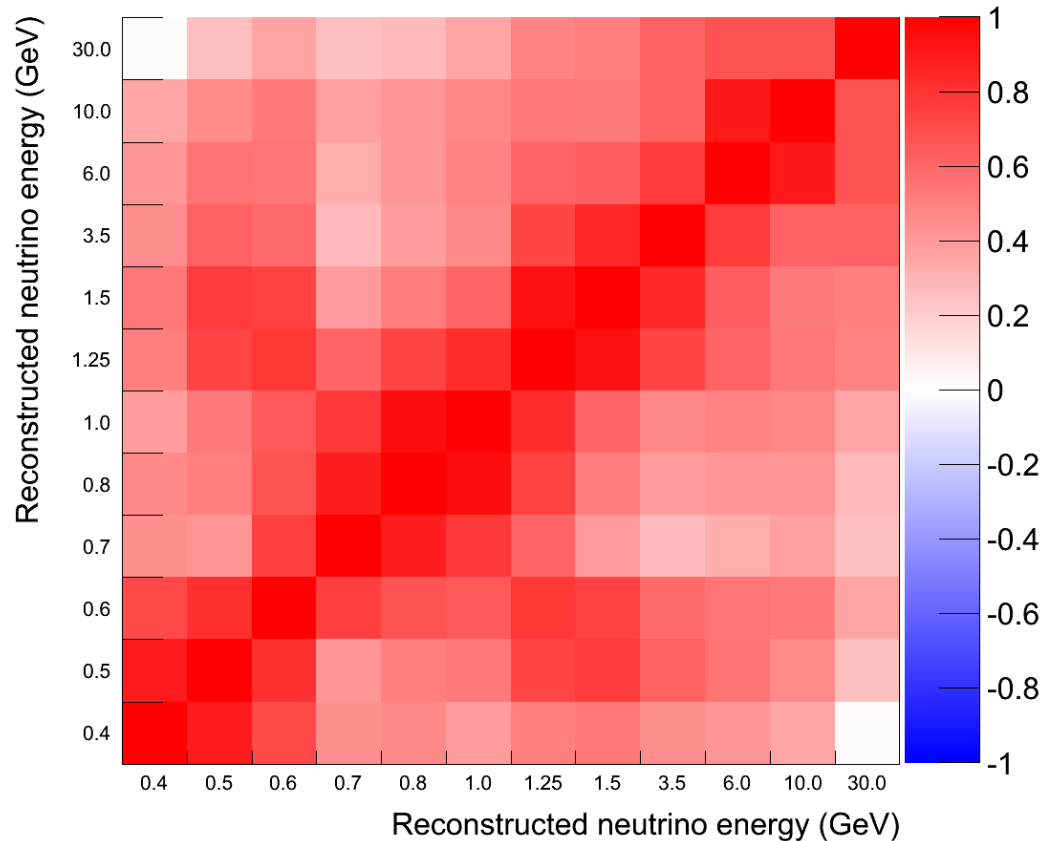
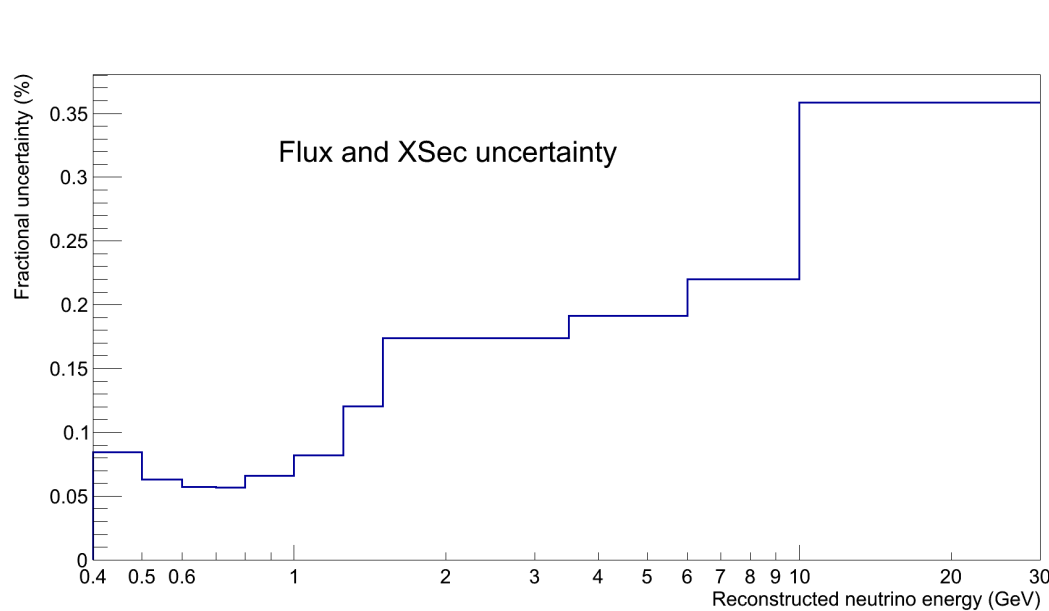
- Flux uncertainties calculated in same ways as for T2K, evaluated at 1km
- Fractional error on left, correlation matrix on right



- Larger errors at high energy – no ν PRISM events
- Error at oscillation dip around 4-5%

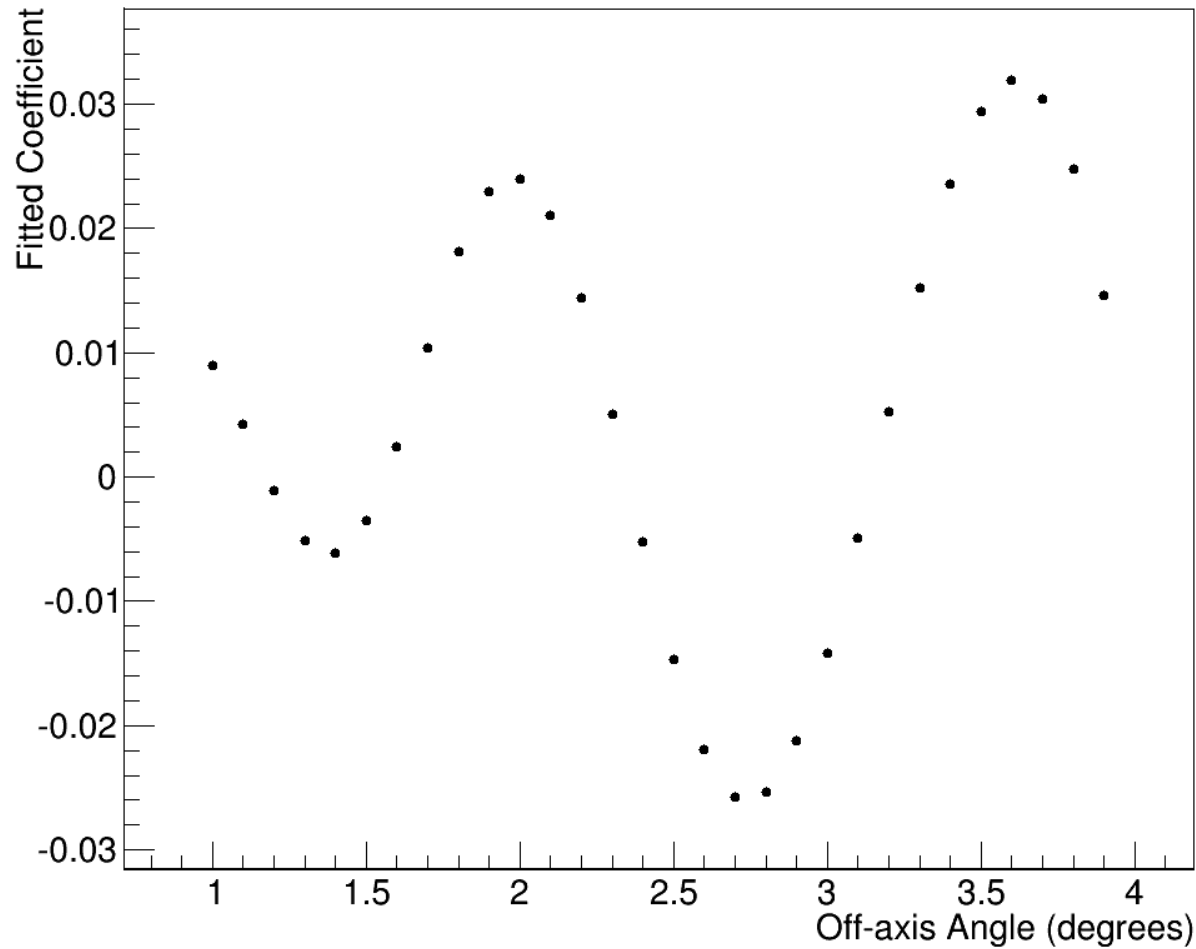
Old flux and Xsec uncertainty

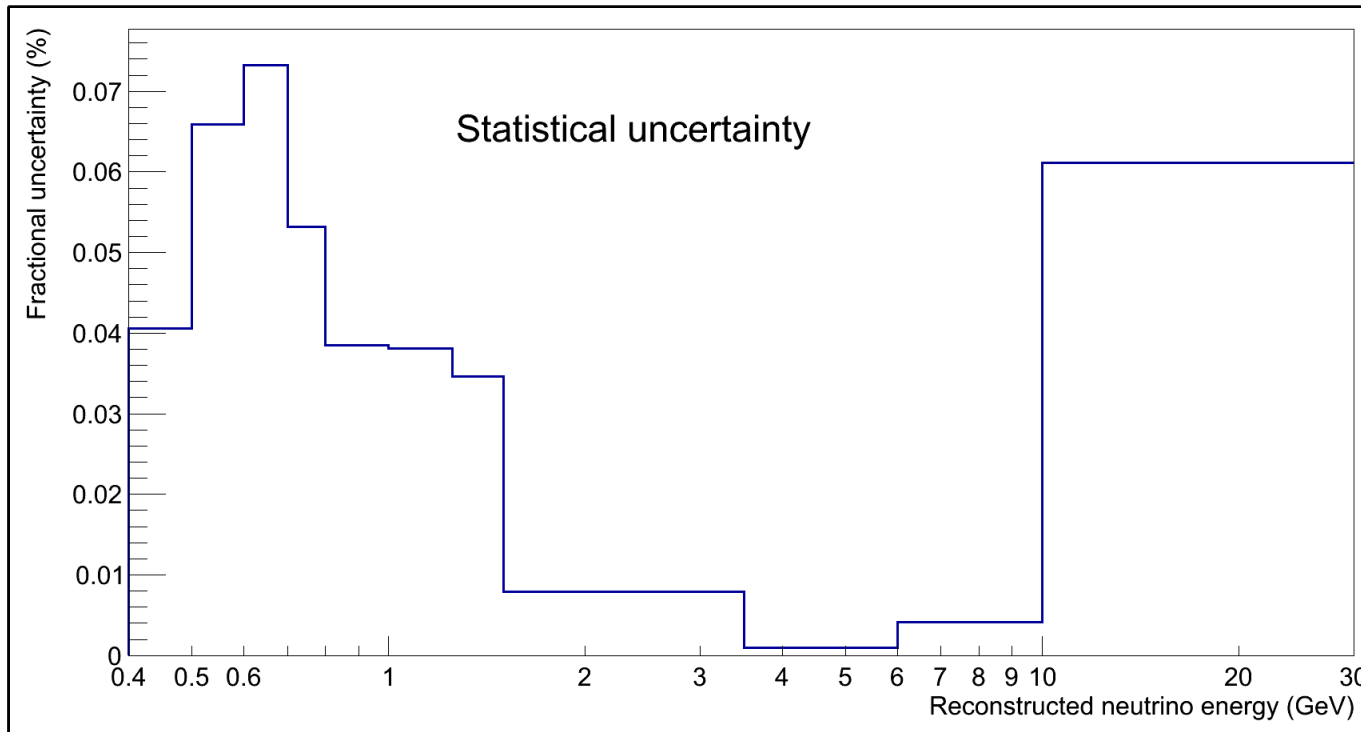
- Xsec uncertainties should largely cancel at vPRISM – amount of cancellation depends on how well flux combination matches SK flux
- Need to throw flux and cross section uncertainties together



- Combined flux and cross section uncertainty around 5% at the oscillation dip

- Smooth linear combination – variations in neighbouring slices cancel out to large extent





- Uncertainty maximal in oscillation dip – subtracting distributions to get zero events
- Statistical uncertainty ~7% in oscillation dip

