Overview (Welcome) KEK & IPMU

Mihoko M. Nojiri

Welcome to IPMU

- IPMU(Institute of Physics and Mathematics of the Universe started Oct 1st 2007 (initially 5+5years)
- University support (permanent positions through TODIAS)
- March 2015 Kavli IPMU has been nominated for a 5-year extension, as a "highly exceptional case whose achievements are far beyond the very high WPI standard".
- Good start up to the transition to the long term future
- First IPMU workshop is on LHC Dec 17 2007, even before "Opening Symposium
- Since then, we have been interested in the physics relevant to the collider physics

Thanks to Durham and KIAS



HOME JAPANESE

Focus week : Facing LHC data

Dates: Dec 17 to 21, 2007

Contact:

Mihoko M. Nojiri (nojiri _at_ kek.jp)

(_a_ should be replaced by @)

The meeting aims to discuss the issues related to the discovery of the new physics signature at LHC, ideas to measure the parameters, identify

experimental and theoretical reality that should be overcome by the start of the experiments. Following researchers are agreed to come.

Teruki Kamon (Texas A&M) Tomasso Lari (Milan) Patrick Meade (Harvard) Tilman Plhen (Edinburgh) Giacomo Polesello (Pavia) Maxim Perelstein (Cornell) Steffen Schumann (Edinburgh) Jay Wacker (SLAC) C.-P. Yuan (Michigan State)

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LHC Run Lachievement Higgs discovery





consistency among the channels



Higgs coupling in future Peskin 1312.4974

μ values	300 fb^{-1}		3000 fb^{-1}	
	CMS	here	CMS	here
$\gamma\gamma$	[6,12]	[6.2 , 11.3]	[4,8]	[3.7 , 8.0]
WW	[6, 11]	[7.6, 12.7]	[4, 7]	[5.2, 11.9]
ZZ	[7, 11]	$[\ 6.2\ ,\ 12.7\]$	[4, 7]	$[\ 3.0 \ , \ 7.0 \]$
$b\overline{b}$	[11,14]	[13.6, 16.7]	[5, 7]	[4.7 , 8.6]
$\tau^+\tau^-$	[8,14]	$[\ 6.2 \ , \ 12.0 \]$	[5, 8]	[2.8, 7.2]
invis.	[11, 17]	[11.2, 16.6]	[4,11]	[4.1, 10.9]
κ values	300 fb^{-1}		3000 fb^{-1}	
	CMS	here	CMS	here
γ	CMS [5,7]	here [5.7 , 9.0]	CMS [2,5]	here [2.9 , 6.5]
$\frac{\gamma}{W}$	$ \begin{array}{c} \text{CMS} \\ [5,7] \\ [4,6] \end{array} $	$\begin{array}{r} \text{here} \\ \hline [5.7 , 9.0] \\ \hline [4.2 , 5.4] \end{array}$	CMS [2,5] [2,5]	here [2.9 , 6.5] [1.6 , 3.3]
$\frac{\gamma}{W}_{Z}$	$\begin{array}{c} \text{CMS} \\ [5,7] \\ [4,6] \\ [4,6] \\ [4,6] \end{array}$	$\begin{tabular}{ c c c c c c c c c c c c c c c c c c c$	$\begin{array}{c} \text{CMS} \\ [2,5] \\ [2,5] \\ [2,4] \end{array}$	here [2.9 , 6.5] [1.6 , 3.3] [2.8 , 6.3]
$\begin{array}{c} \gamma \\ W \\ Z \\ g \end{array}$	$\begin{array}{c} \text{CMS} \\ \hline [5,7] \\ [4,6] \\ [4,6] \\ [6,8] \end{array}$	here [5.7 , 9.0] [4.2 , 5.4] [5.7 , 8.5] [4.9 , 6.9]	$\begin{array}{c} \text{CMS} \\ [2,5] \\ [2,5] \\ [2,4] \\ [3,5] \end{array}$	$\begin{tabular}{ c c c c c c c c c c c c c c c c c c c$
$\gamma \\ W \\ Z \\ g \\ b$	$\begin{array}{c} \text{CMS} \\ \hline [5,7] \\ [4,6] \\ [4,6] \\ [6,8] \\ [10,13] \end{array}$	here [5.7 , 9.0] [4.2 , 5.4] [5.7 , 8.5] [4.9 , 6.9] [11.4 , 14.9]	$\begin{array}{c} \text{CMS} \\ \hline [2,5] \\ [2,5] \\ [2,4] \\ [3,5] \\ [4,7] \end{array}$	$\begin{tabular}{ c c c c c c c c c c c c c c c c c c c$
$egin{array}{c} \gamma & & \ W & \ Z & \ g & \ b & \ t & \ \end{array}$	$\begin{array}{c} \text{CMS} \\ \hline [5,7] \\ [4,6] \\ [4,6] \\ [6,8] \\ [10,13] \\ [14,15] \end{array}$	here [5.7 , 9.0] [4.2 , 5.4] [5.7 , 8.5] [4.9 , 6.9] [11.4 , 14.9] [17.3 , 20.5]	$\begin{array}{c} \text{CMS} \\ \hline [2,5] \\ [2,5] \\ [2,4] \\ [3,5] \\ [4,7] \\ [6,8] \end{array}$	$\begin{array}{r} \text{here} \\ \hline [2.9 , 6.5] \\ [1.6 , 3.3] \\ [2.8 , 6.3] \\ [2.3 , 4.8] \\ [4.2 , 8.5] \\ [5.7 , 12.9] \end{array}$
$\begin{array}{c} \gamma \\ W \\ Z \\ g \\ b \\ t \\ \tau \end{array}$	$\begin{array}{c} \text{CMS} \\ \hline [5,7] \\ [4,6] \\ [4,6] \\ [6,8] \\ [10,13] \\ [14,15] \\ [6,8] \end{array}$	$\begin{array}{r} \text{here} \\ \hline [5.7 , 9.0] \\ [4.2 , 5.4] \\ [5.7 , 8.5] \\ [4.9 , 6.9] \\ [11.4 , 14.9] \\ [17.3 , 20.5] \\ [5.8 , 9.5] \end{array}$	$\begin{array}{c} \text{CMS} \\ \hline [2,5] \\ [2,5] \\ [2,4] \\ [3,5] \\ [4,7] \\ [6,8] \\ [2,5] \end{array}$	$\begin{tabular}{ c c c c c c c c c c c c c c c c c c c$

4% at 3000fb⁻¹ O(0.1)% at future e+e-

Table 2: Comparison of the results of fits with the inputs in Table 1 to the fit results given in [7]. All numbers are given as 1 σ uncertainties, in %. In expressions in brackets, the first entry is for Scenario 2, the second is for Scenario 1.



Some (opposite) tension between the two channels but very good agreement in the central values

$$egin{aligned} m_{H}^{\gamma\gamma} &= 125.07 \pm 0.29 \,\,\, {
m GeV} \ &= 125.07 \pm 0.25 \,({
m stat.}) \pm 0.14 \,({
m syst.}) \,\,\, {
m GeV} \ m_{H}^{4\ell} &= 125.15 \pm 0.40 \,\,\, {
m GeV} \ &= 125.15 \pm 0.37 \,({
m stat.}) \pm 0.15 \,({
m syst.}) \,\,\, {
m GeV} \end{aligned}$$

Sep 1, 2015

Marco Pieri UC San Diego

35

ATLAS $\gamma\gamma 2\sigma$ away from the best . ATLAS 4I best fit is 2σ away from ATLAS 2γ

Is Effective theory always efficient way?

- effective theory approach is often useful to parametrize signal distribution, but the care must be taken so that we do not misuse it.
- \Rightarrow ex: Higgs width in pp \rightarrow ZZ using off shell amplitude $\sigma_{\rm gg \to H \to ZZ^*}^{\rm on-shell} \sim \frac{g_{\rm ggH}^2 g_{\rm HZZ}^2}{m_{\rm H} \Gamma_{\rm H}} \text{ and } \sigma_{\rm gg \to H^* \to ZZ}^{\rm off-shell} \sim \frac{g_{\rm ggH}^2 g_{\rm HZZ}^2}{(2m_7)^2}.$ C.Englert et al 1410.5440 $p(g)p(g) \rightarrow e^+e^-\mu^+\mu^-$ 10 $\mathrm{d}\sigma/\mathrm{d}m(4\ell)~[\mathrm{ab}/20~\mathrm{GeV}]$ **FH**<20MeV $\sqrt{s} = 8 \text{ TeV}$ 10 but the result $tt \to ZZ$ full 10^{-1} $\sigma \, [pb]$ $\rightarrow ZZ$ Gauge comes from $t\bar{t} \rightarrow ZZ$ Yukawa 10^{-2} CMS search region CMS exclusion 0.1 10^{-3} non-Unitarity of $\Gamma_h = 10 \times \Gamma_h^{\text{tot}}$, couplings rescaled $t\bar{t} \rightarrow ZZ$ 10^{-4} the amplitude 0.01 0.6 0.2 0.3 0.1 0.4 0.5 $m(4\ell)$ [TeV] 2 3 7 \sqrt{s} [TeV]

More information in higgs distributions?

Higgs boson pT, ttH... typically require high pT





Typical cross section vs cross section limit



Stop searches ~ theoretical understanding of SM



Theory needs to prepare high scale SUSY calculation

- ◆ Higgs mass (now much less than 1% error) → SUSY scale, stop mass and mixing, tanβ
 - Higgs mass, Higgs branching ratios (resum large logs and nondecoupling corrections
 - rare decays (public tools mostly uses parameters at mZ...)

ATL-PHYS-PUB-2013-007 (W, Z channel)



Anomaly driven life



CMS 2.6 σ



ATLAS 3σ in ee none in $\mu\mu$



Stay calm and look

Global search results

Compare expected P-value distributions with observed
 channel gets diluted by look elsewhere effect



ATLAS-CONF-2014-006 from Ben Kilminster's talk

Use SUSY data, HT, M_event, Etmiss compare with pseudo experiment and get p-values compare with 697 class of events with SM bkg>0.1evnets

vista Point

Anomalies are not significant "looking elsewhere" effect ATLAS-CONF-2014-006



di-b 20 0^E 10 15 20 25 30 35 40 5

0.004

0.002



QCD backgrounds estimate

- number of charged tracks ← large uncertainty in MC prediction for gluon jet (n<30)</p>





1507.01681 Abe Kitahara Nojiri

Still something need to be done to understand the cut flow

dark matter at LHC



Dark Matter at LHC: Mono-Jet



26.8%

Dark Matter

8TeV→Toward original design

Accident on Sep 19 2008 → half
 recovered in 2010 → trying to achieve
 the original design



- ◆ 4TeV, 50 ns 1380 bunch → 25ns to avoid pile up(experimental request) 2500~2800 bunchs/ring 1.15 10^11 p per bunch
- It is not trivial: Electronic cloud: more radiation-> more electrons-> more electron acceleration-> more secondary electrons -> heat to cryogenic system-> (><)
 - beam scrubbing (long beam operation to reduce electron clouds) Many ideas are tried

Some issues of machine and experiments

Unidentified Falling Objects (UFO)

microparticle falling to the beam, ionized -> beam dump and quench

- ✤ … ULO, Earth faults …
- Experimental side: CMS magnet system (I think having two sound experiments are important)







Run 2



- EYETS Extended Year End Technical Stop 19 weeks CMS pixel upgrade
- Start LS2 at the end of 2018

Very intersting results already and will be reviewed tomorrow

	Peak lumi E34 cm ⁻² s ⁻¹	Days proton physics	Approx. int lumi [fb ⁻¹]
2015	~0.5	65	3
2016	1.2	160	30
2017	1.5	160	36
2018	1.5	160	36

2015 Lum is not as large as we hoped

And beyond



is mandatory:

FG EPS 15

Overview

New stage (factor 2 energy and factor 5 luminosity) is ahead.

 Meantime, Theory have gone from fundamental theory to effective theory.

 3000fb-1 ahead? and 100TeV or ILC?
 Particle physics need to answer big question to justify its costs.