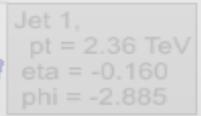


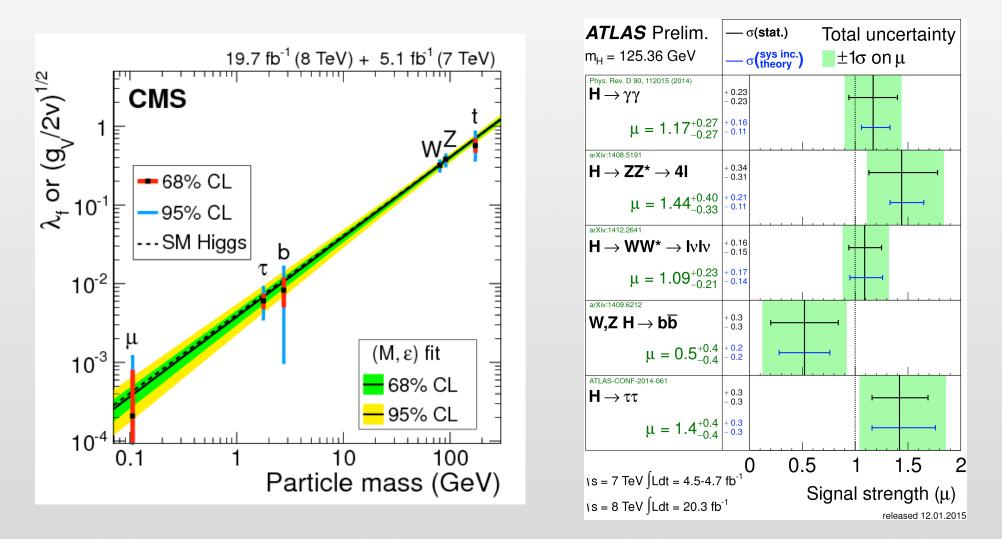
Non-relativistic Heavy Particles at LHC Run 2

Yang Bai

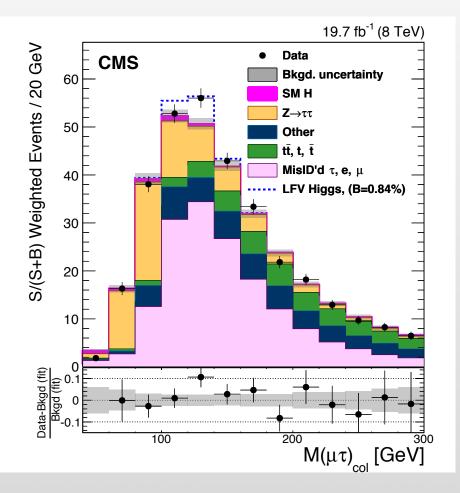
University of Wisconsin-Madison Kavli-IPMU-Durham-KIAS workshop September 7, 2015

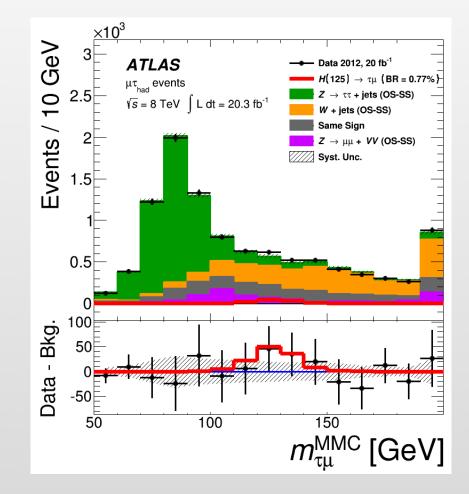


Discovered the Higgs boson with SM properties



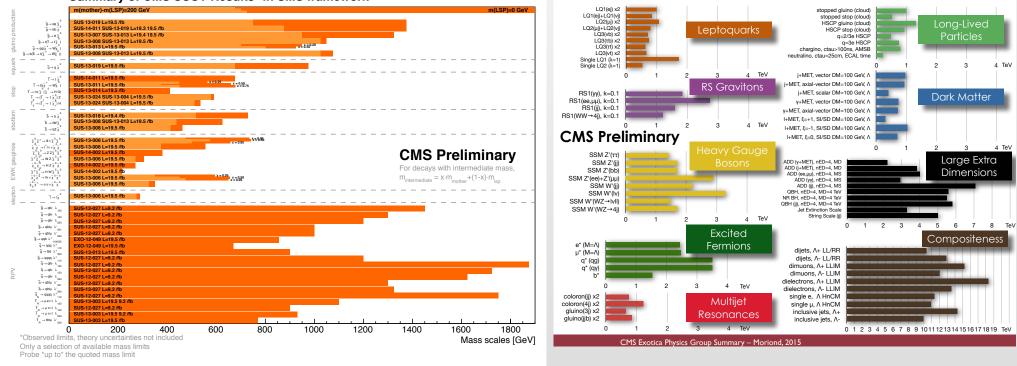
• Flavor-changing decay?





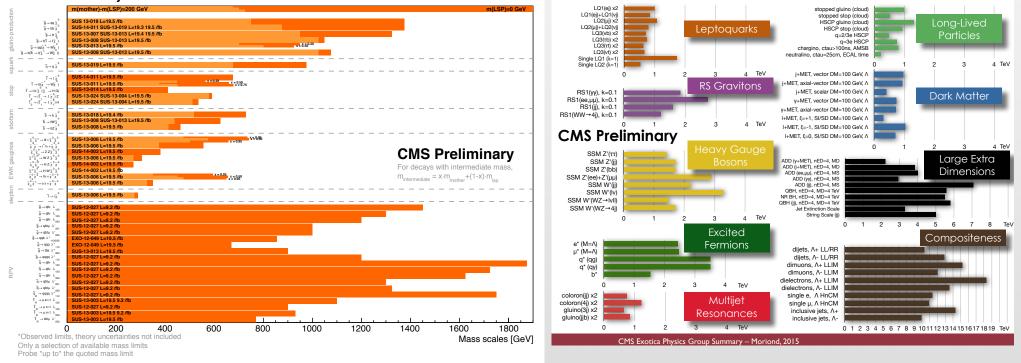
No strong sign of BSM physics

Summary of CMS SUSY Results* in SMS framework



No strong sign of BSM physics

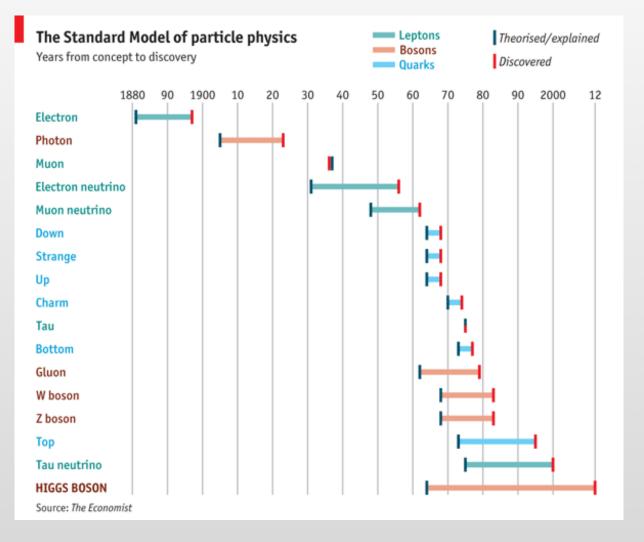
Summary of CMS SUSY Results* in SMS framework



• Is this normal?

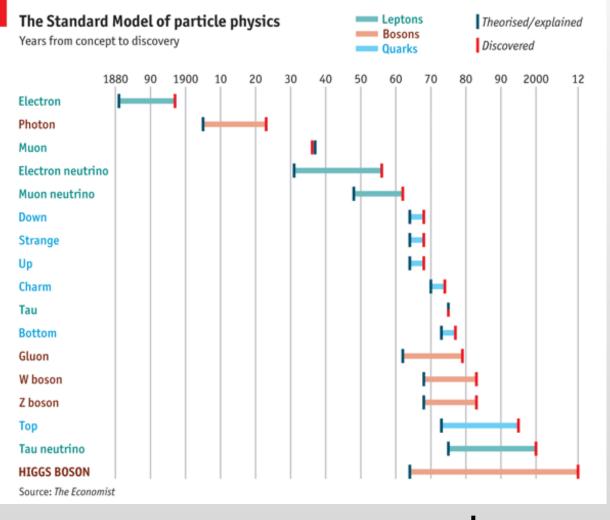
Next Particle from Colliders

• Longer time interval between theory and experiment



Next Particle from Colliders

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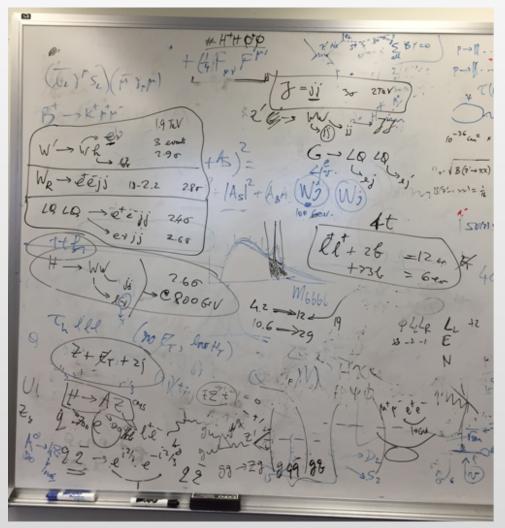


MSSM

> 2029 ?

.

• A few tentative hints



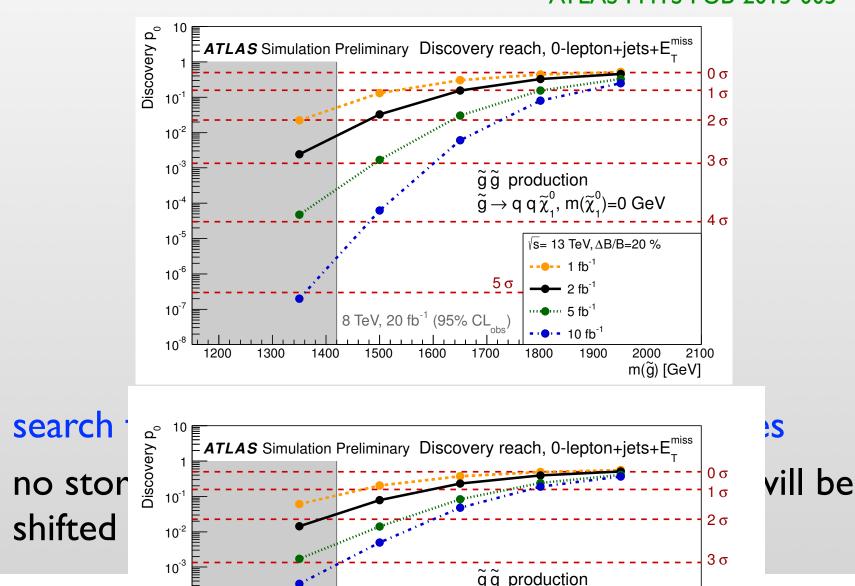
- no cross confirmation from CMS and ATLAS
- + lepton flavor dependent
- understand more of the boosted objects

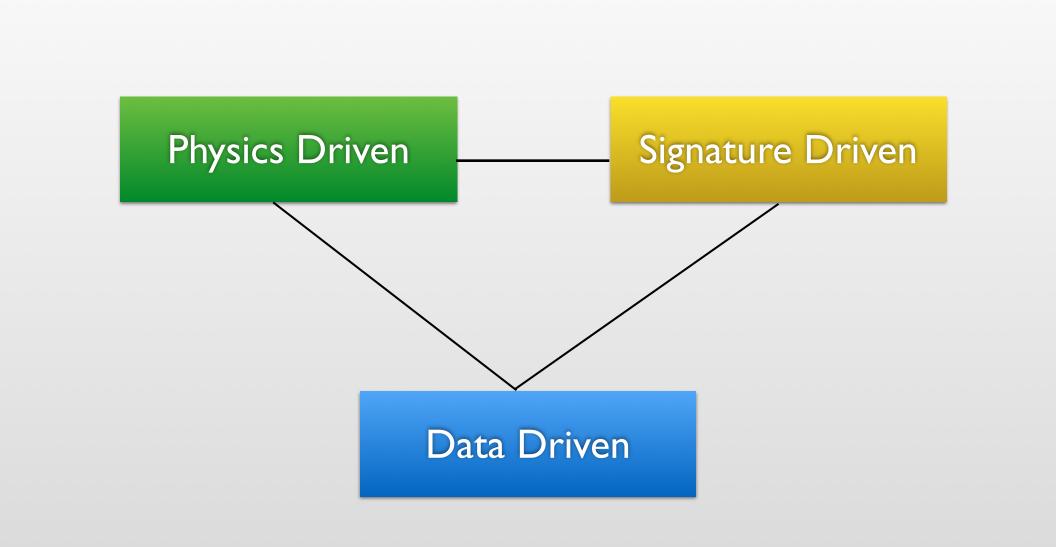
see more talks about them here

from discussion with B. Dobrescu@Fermilab in May 2015

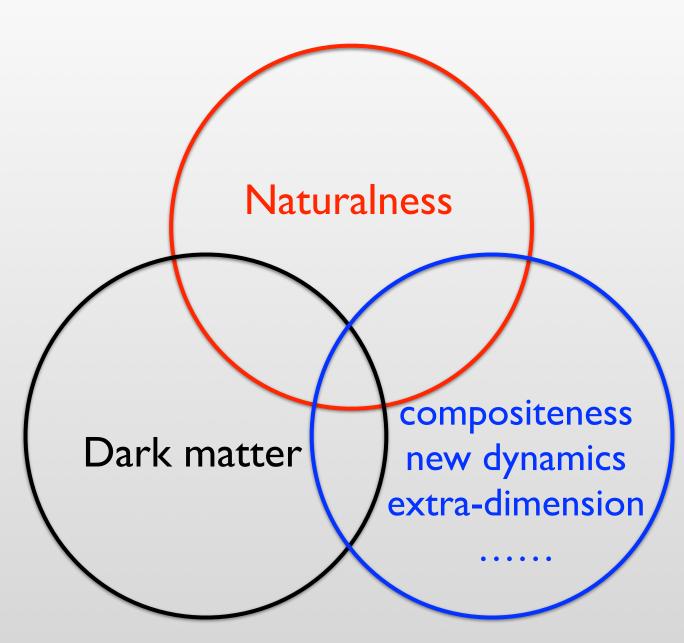
What to Anticipate for LHC Run II?

 Increase the reach for well-motivated new particle masses by a factor of few
 ATLAS-PHYS-PUB-2015-005

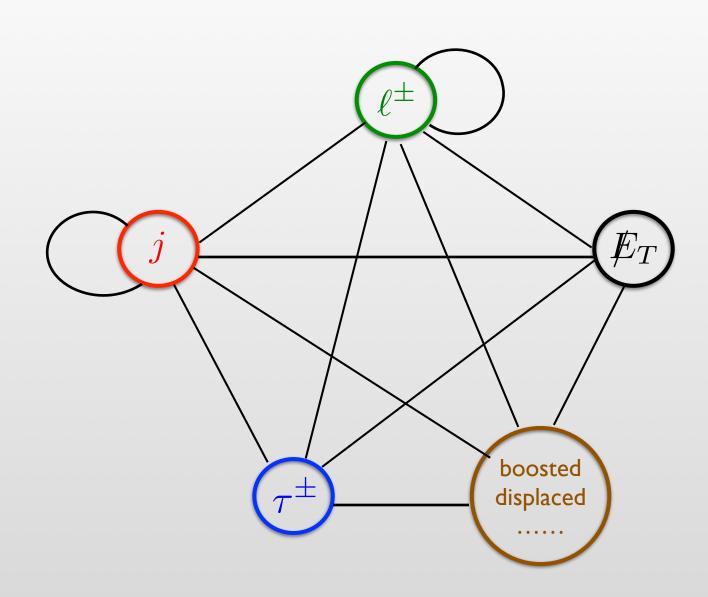




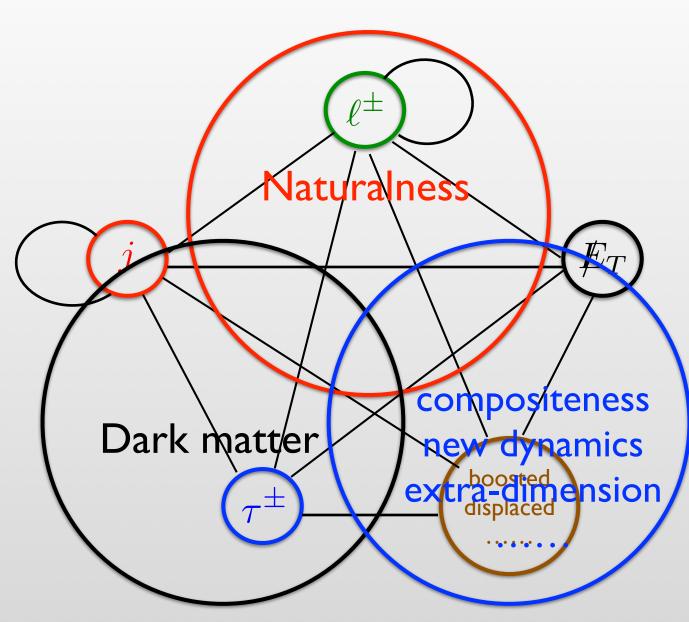
Physics Driven



Signature Driven



Signature Driven



П

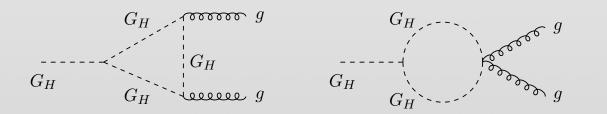


- The new physics scale is generically at ~TeV
- LHC Run 2 will be the right one to discover them
- Hope Nature will not have a meta tuning such that the LHC is just below the threshold



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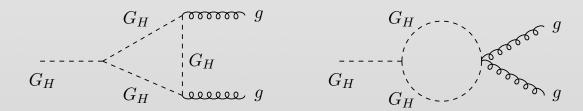
$$\frac{g_s^2}{2} f^{abc} f^{ade} G^b_\mu G^{\mu d} G^c_H G^e_H + g_s f^{abc} G^a_\mu G^b_H \partial^\mu G^c_H + \mu_G d_{abc} G^a_H G^b_H G^c_H$$





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$$\frac{g_s^2}{2} f^{abc} f^{ade} G^b_\mu G^{\mu d} G^c_H G^e_H + g_s f^{abc} G^a_\mu G^b_H \partial^\mu G^c_H + \mu_G d_{abc} G^a_H G^b_H G^c_H$$



$$\Gamma_0(G_H \to gg) = \frac{15 \,\alpha_s^2 \,\mu_G^2}{128 \,\pi^3 \,M_{G_H}} \,\left(\frac{\pi^2}{9} - 1\right)^2$$



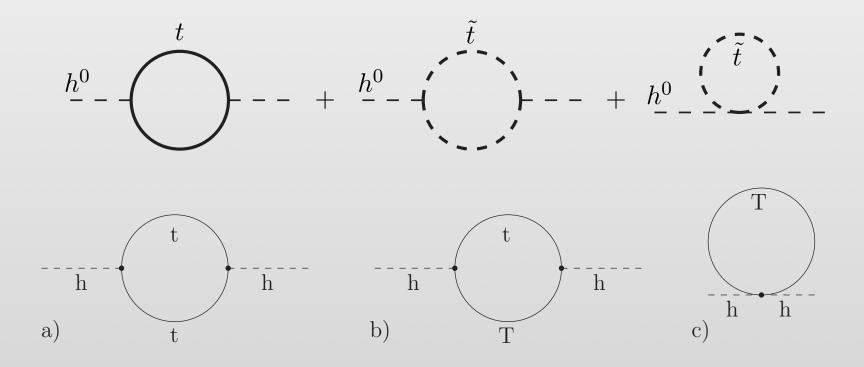
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$$\frac{g_{s}^{r}}{2}f^{abc}f^{ade}G_{\mu}^{b}G^{\mu d}G_{H}^{c}G_{H}^{e} + g_{s}f^{abc}G_{\mu}^{a}G_{H}^{b}\partial^{\mu}G_{H}^{c} + \mu_{G}d_{abc}G_{H}^{a}G_{H}^{b}G_{H}^{c}$$

$$G_{H} = G_{H} = G$$

Naturalness: top partner

- The leading model is supersymmetry. The Higgs field could also be a composite particle.
- To cancel the top-quark induced divergence, we can have bosonic or fermionic top partners



13

Top Partners at Run I

$\tilde{t}_1 \rightarrow t + \tilde{\chi}_1^0$ the signal is ttbar+MET

.....

Early work:

Meade and Reece, hep-ph/0601124 Kong and Park, hep-ph/0703057 Han, Mahbubani, Walker, Wang, 0803.3820

Endpoints:

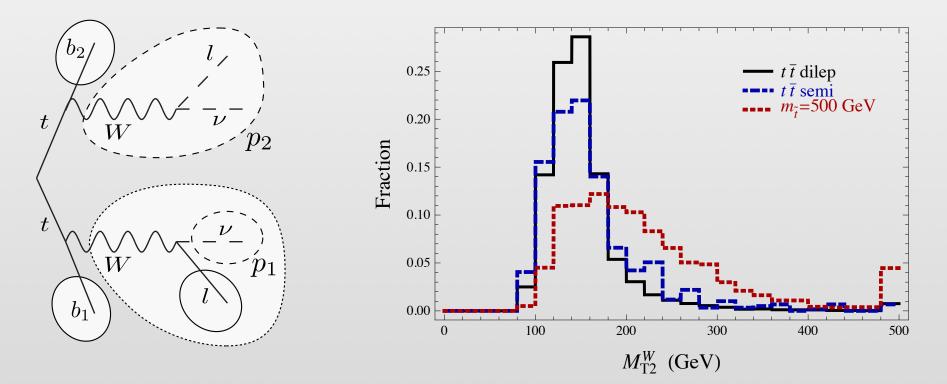
Topness:

Spin-corre	lations:
Top-tagging	g:

YB, Cheng, Gallichio, Gu, 1203.4813 Cao, Han, Wu, Yang, Zhang, 1206.3865 Killic and Tweedie, 1211.6106 Han, Katz, Krohn, Reece, 1205.5808 Plehn, Spannowsky, Takeuchi, 1205.2696 Kaplan, Rehermann, Stolarski, 1205.5816 Dutta, Kamon, Kolev, Sinha, Wang, 1207.1893 Shapes of missing Et: Alves, Buckley, Fox, Lykken, Yu, 1205.5805 Graesser and Shelton, 1212.4495

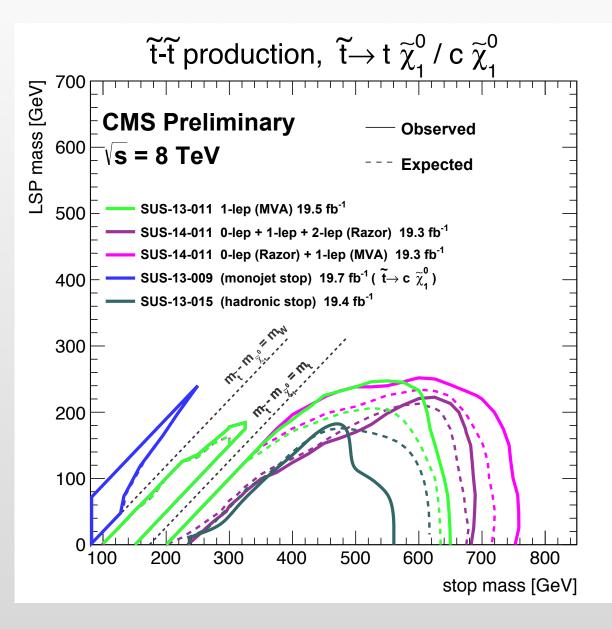
Kinematical Variables for ttbar+MET

 Noticing the main background is dileptonic ttbar, one can construct the asymmetric MT2 variable. The background is bounded by the top quark mass.



YB, Cheng, Gallichio, Gu, 1203.4813

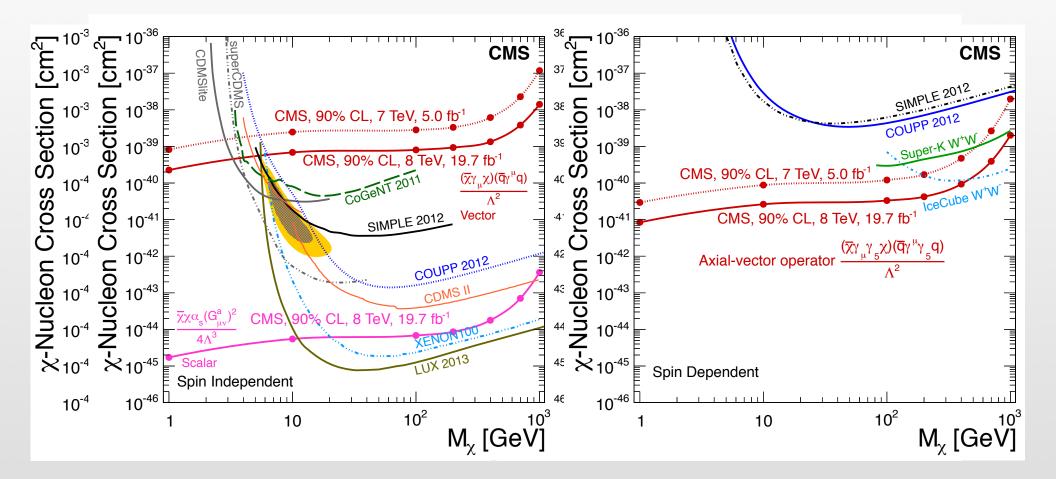
Top Partners at Run I





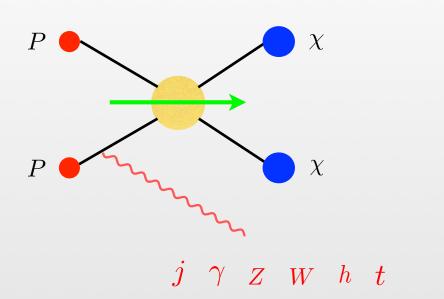
- WIMP dark matter has a scale ~ TeV and may show up at LHC
- Only subset of asymmetric dark matter may has a TeV scale
- the generic EFT-based mono-X signatures have been searched for at the LHC Run I
- LHC Run 2 will concentrate more on simplified dark matter models
- there could be more exotic dark matter signatures

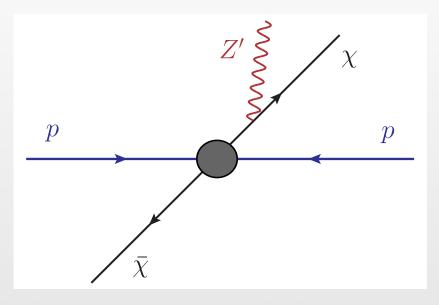
Dark Matter@LHC Run I



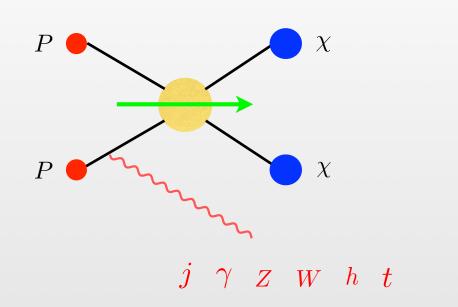
based on dark matter effective operators

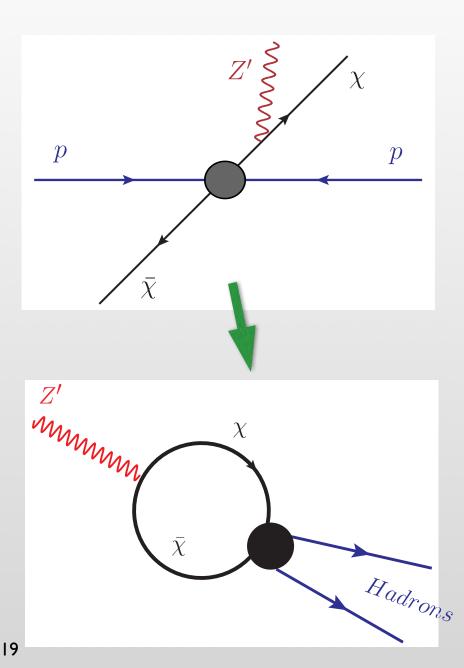
Dark Matter FSR Signature





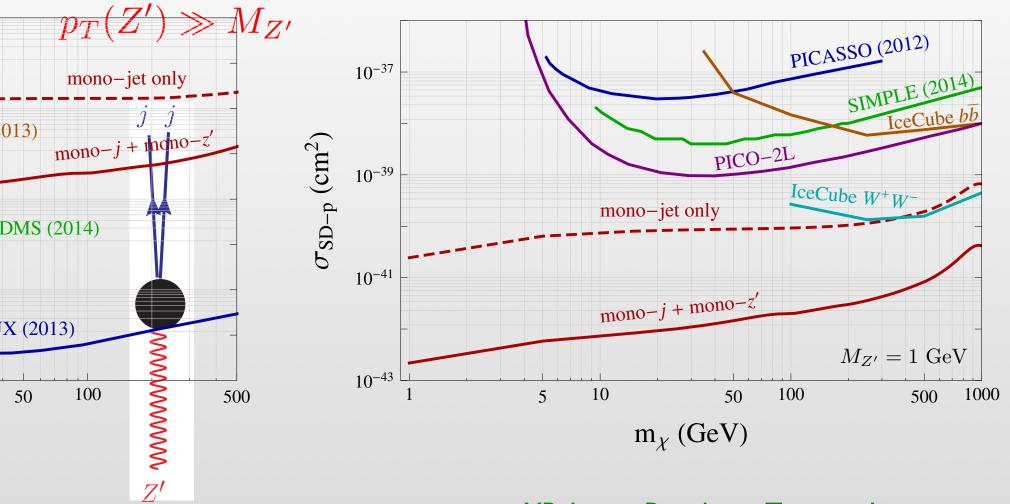
Dark Matter FSR Signature





Given the fact that dark matter can be produced at the LHC and couple to Z'; Z' must decay into hadrons

Mono-Z' **Jet**



YB, James Bourbeau, Tongyan Lin 1504.01395

Simplified Dark Matter Models

- Better description, but more models. One may have several hundred models to work with.
- One should use LHC signatures to organize different models. Finding signatures beyond the MSSM could be useful.

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Chromo-Rayleigh Interactions

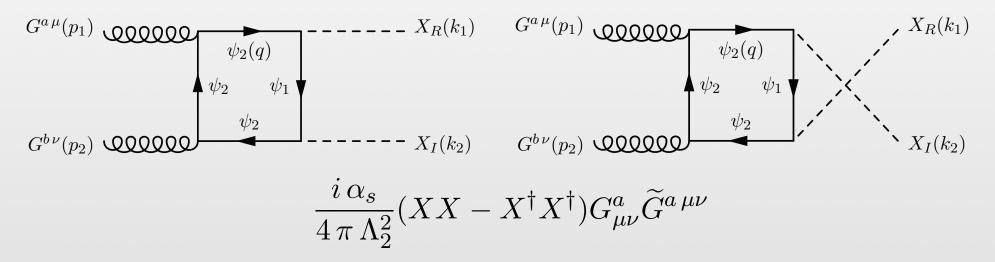
$$\frac{\alpha_s}{4\pi\Lambda_1^2} X^{\dagger} X G^a_{\mu\nu} G^{a\,\mu\nu}$$

$$\frac{i\,\alpha_s}{4\,\pi\,\Lambda_2^2}(XX - X^{\dagger}X^{\dagger})G^a_{\mu\nu}\widetilde{G}^{a\,\mu\nu}$$

"UV"-Completed Simplified Models

 $\mathcal{L} \supset -y_1 \left(X + X^{\dagger} \right) \left(\overline{\psi}_1 \psi_2 + \overline{\psi}_2 \psi_1 \right) - \left(X - X^{\dagger} \right) \left(y_2 \,\overline{\psi}_1 \gamma_5 \psi_2 + y_2 \,\overline{\psi}_2 \gamma_5 \psi_1 \right)$

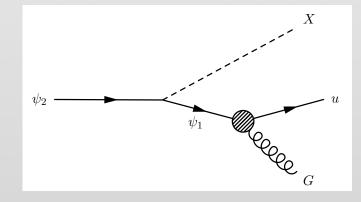
• two representative one-loop diagrams (totally 6)



• LHC signature: a pair of dijetresonance plus MET

$$pp \to \psi_2 \overline{\psi}_2 \to (u \, g \, X)(\bar{u} \, g \, X^\dagger) \to 4j + E_T^{\text{miss}}$$

YB, Osborne, 1506.07110



Asymmetric Dark Baryon Model

 $\frac{\Omega_{\rm DM}}{\Omega_{\rm Baryon}} = \frac{m_{\rm DM} n_{\rm DM}}{m_{\rm p} n_p} \approx 5 \sim 6$

Two conditions: (1): $n_{
m DM} \sim n_p$ (2): $m_{
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• (2) If dark matter is a "dark baryon" from a new QCD-like strong dynamics in the dark matter sector

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 ?

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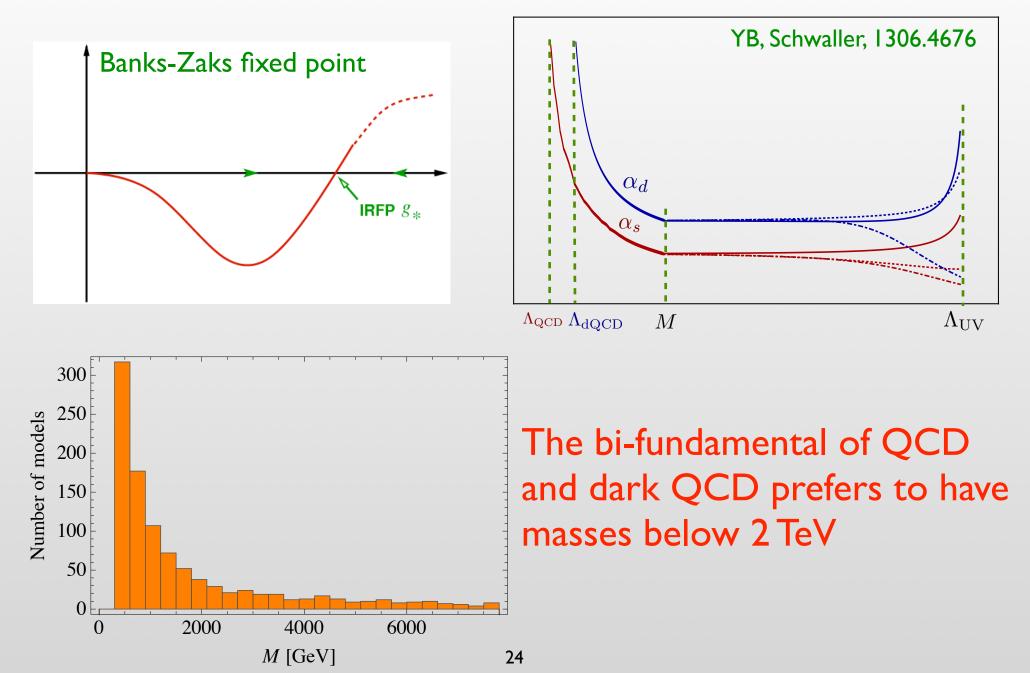
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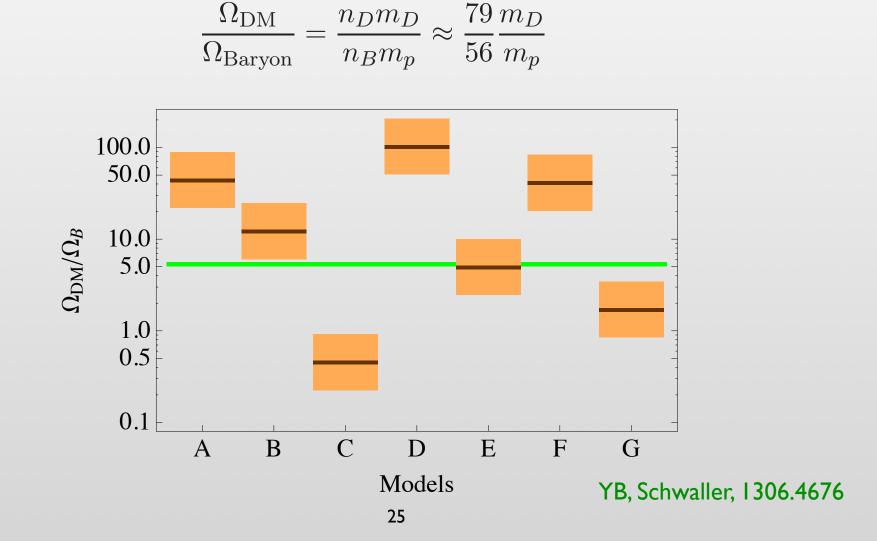
 Need to have QCD and dQCD gauge couplings related to each other

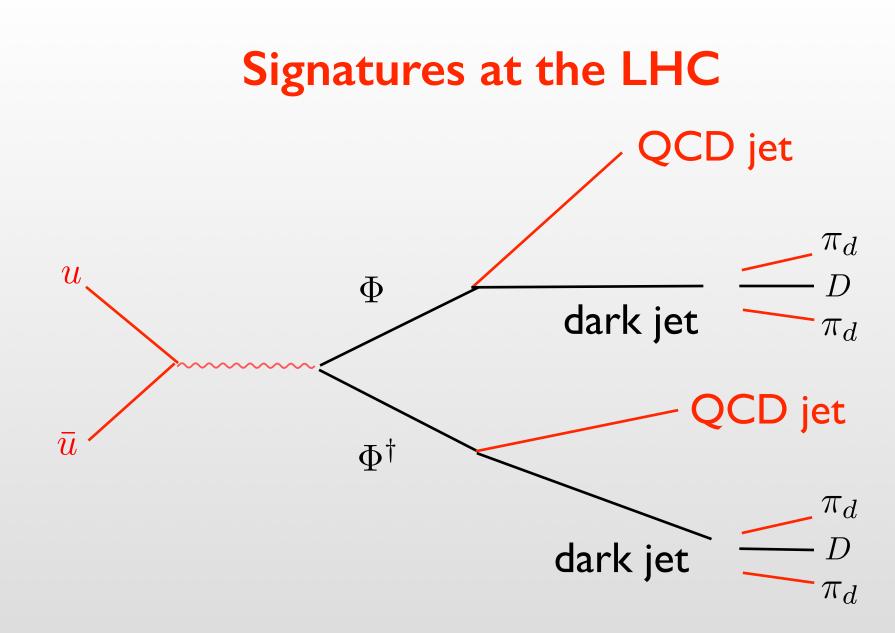
IRFP to Relate QCD and dQCD Scales



Ratios of Energy Densities

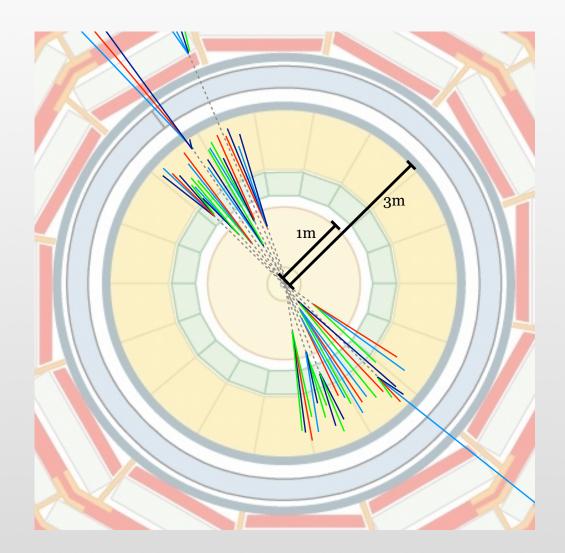
 the bi-fundamental particles can also be used to share the baryon and the dark baryon asymmetries





 dark pions can decay back to SM particles and are generically long-lived

Emerging Dark Jets



Schwaller, Stolarski, Weiler, 1502.05409

compositeness new dynamics extra-dimension

- no obvious scales unless related to naturalness or dark matter
- looking for surprises from LHC; chance is not high
- fully utilize the LHC capability



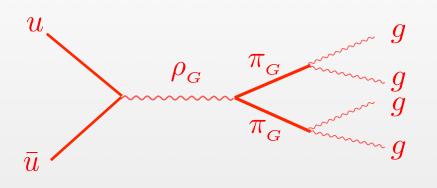
- no obvious scales unless related to naturalness or dark matter
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	$SU(N)_G$	$SU(3)_c$	$SU(2)_W$	$U(1)_Y$
$\overline{\psi_{L,R}}$	N	3	1	0

for another QCD-like dynamics, we anticipate both rho mesons and pions

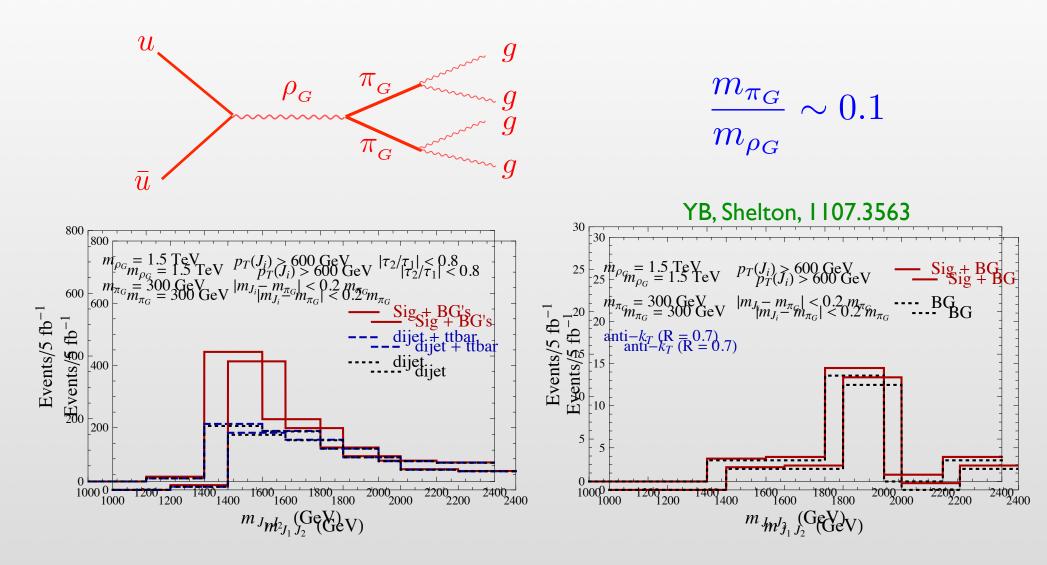
Kilic, Okui, Sundrum, 0906.0577 YB, Shelton, 1107.3563

Di-fat-jet Resonance

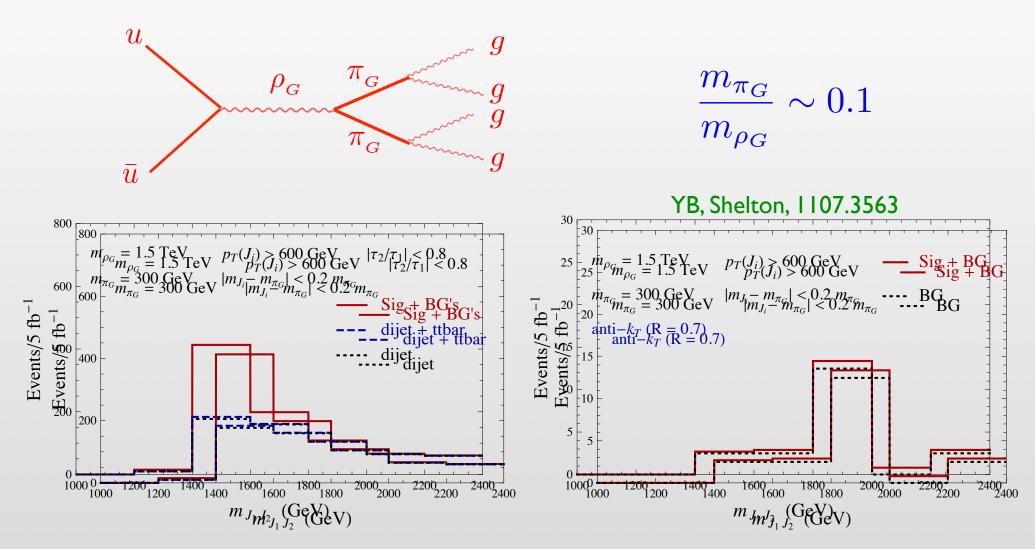


 $\frac{m_{\pi_G}}{m_{\rho_G}} \sim 0.1$

Di-fat-jet Resonance

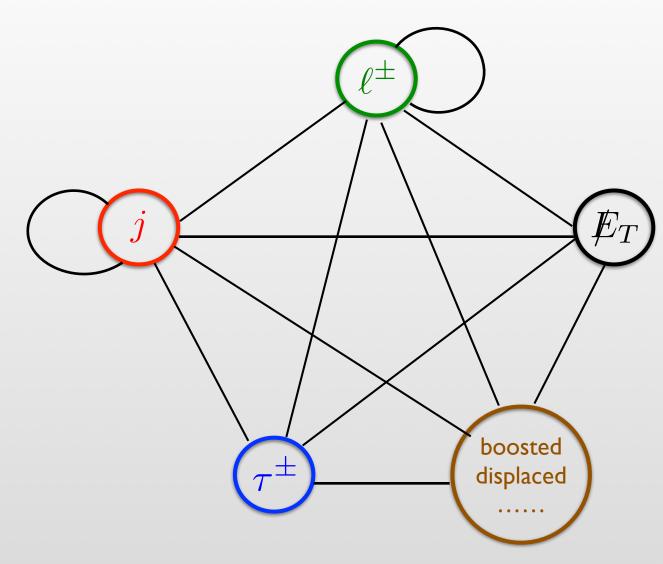


Di-fat-jet Resonance

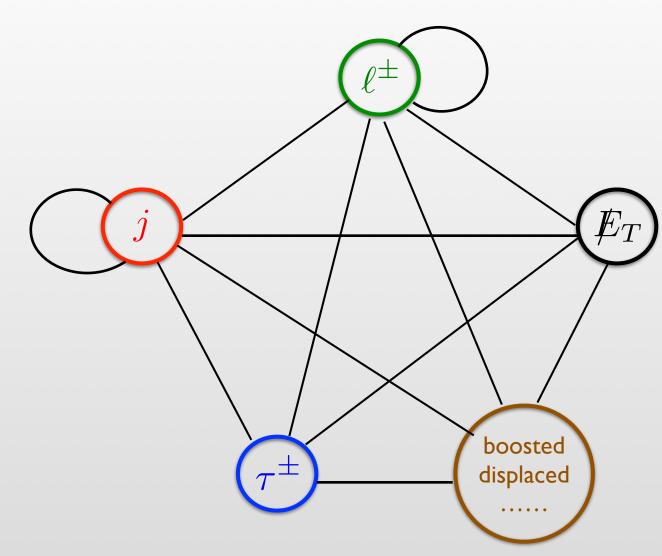


can provide an explanation why the recent di-boson ~ 2 TeV resonance does not show up in leptonic channels

Signature Driven



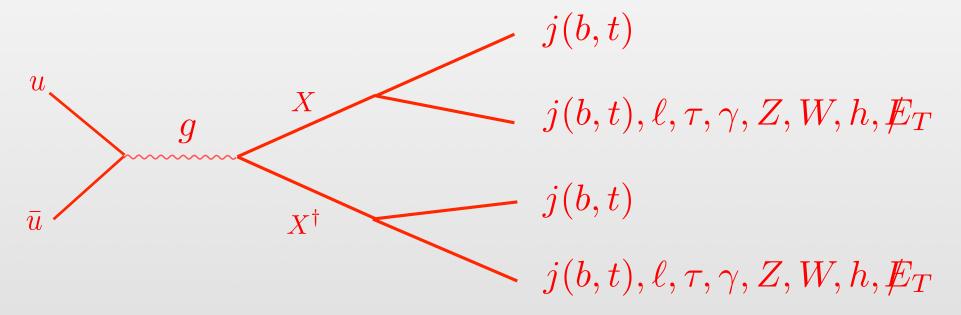
Signature Driven



While there are too many combinations, let us focus QCD charged new particles

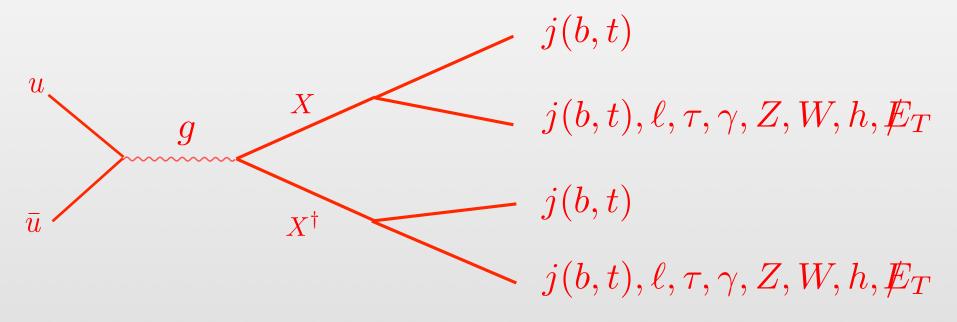
QCD Pair-produced Particles

Focusing on two-body decays, at least two QCD charged particles in the final state



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Focusing on two-body decays, at least two QCD charged particles in the final state

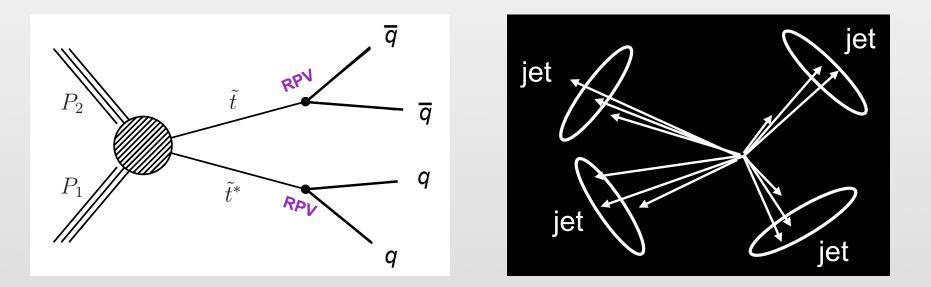


The most challenging one is the four-jet final state. The smallest production cross section is a scalar color-triplet X.

Right-handed Stop Decay

from the following R-parity violating operator

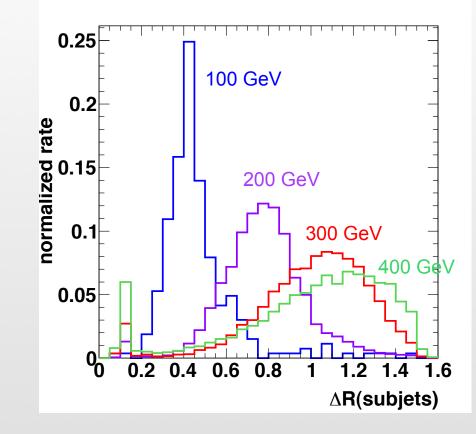
 $\lambda_{3ij}\,\tilde{t}_R\,d_R^i\,d_R^j \qquad (i\neq j)$



Serve as a benchmark for purely jetty pair-production searches (minimal color and spin)

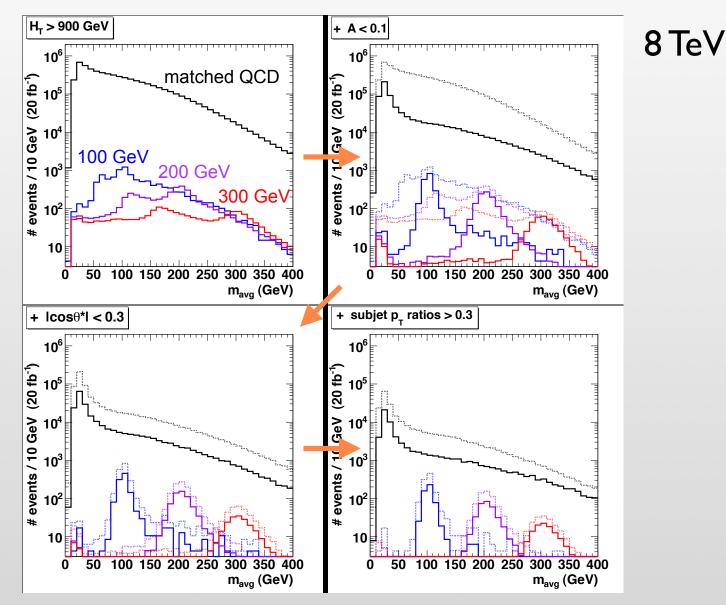
AR DISTRIBUTIONS

Delta R Distribution

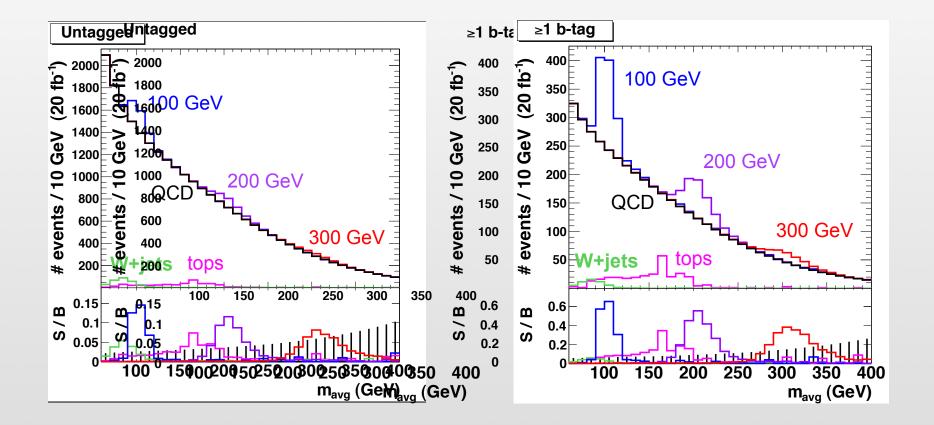


- Jet-Ht trigger: offline Ht > 900 GeV
- Capture stop decays in R = 1.5 C/A jets

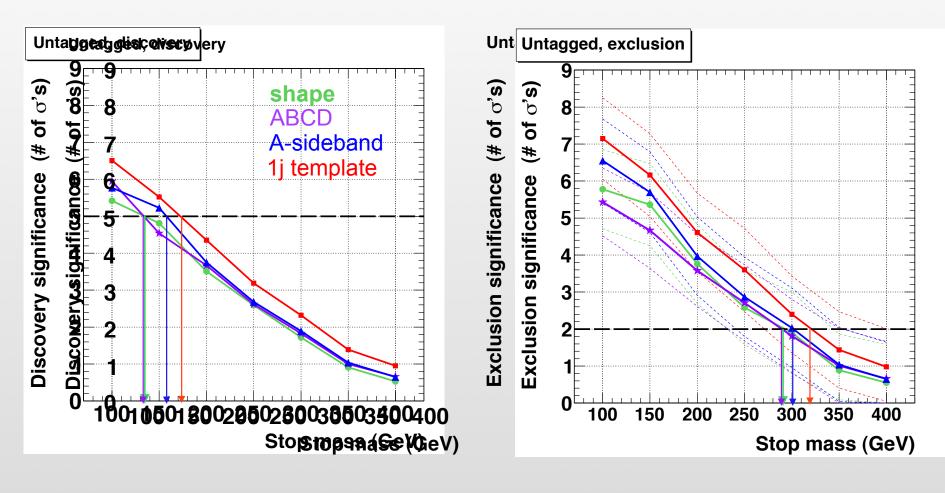
Jet-substructure Cut Flow



Avangegeiddesepenetra



220225888511/v/1885 Unitagged Sensitivities, untagged



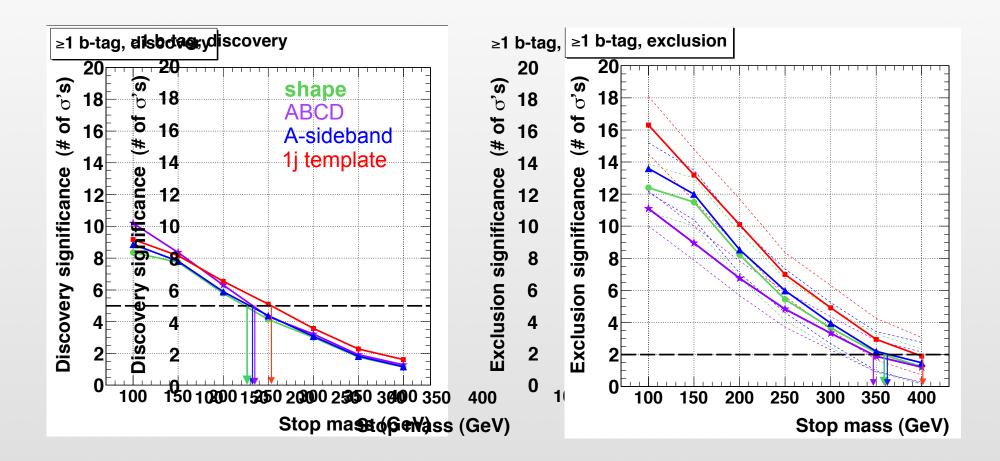
exclusion ~300 GeV

YB, Katz, Tweedie, 1309.6631

discovery ~150 GeV

COTE TOELLENDAN MED'ED' HELDRENDEN

Sensitivities, b-tagged

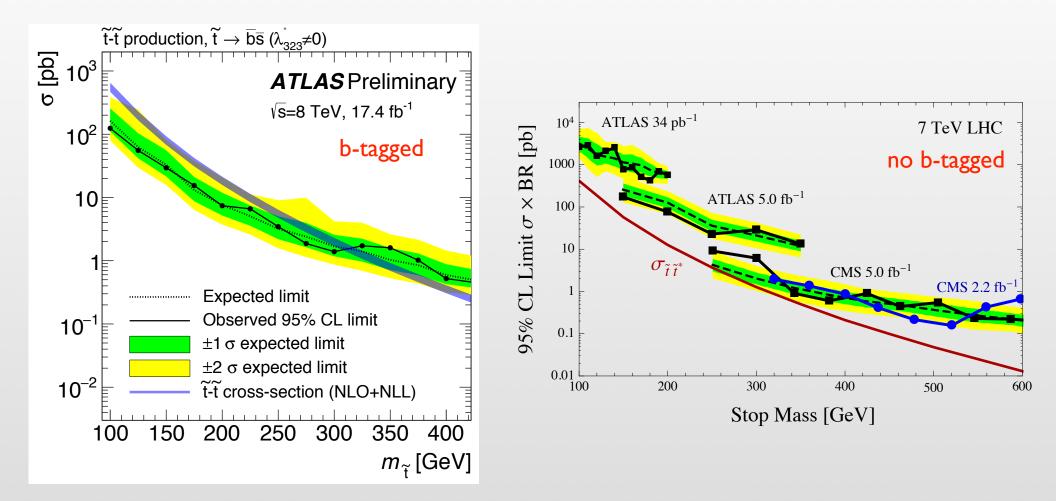


exclusion ~350-400 GeV

discovery ~250 GeV

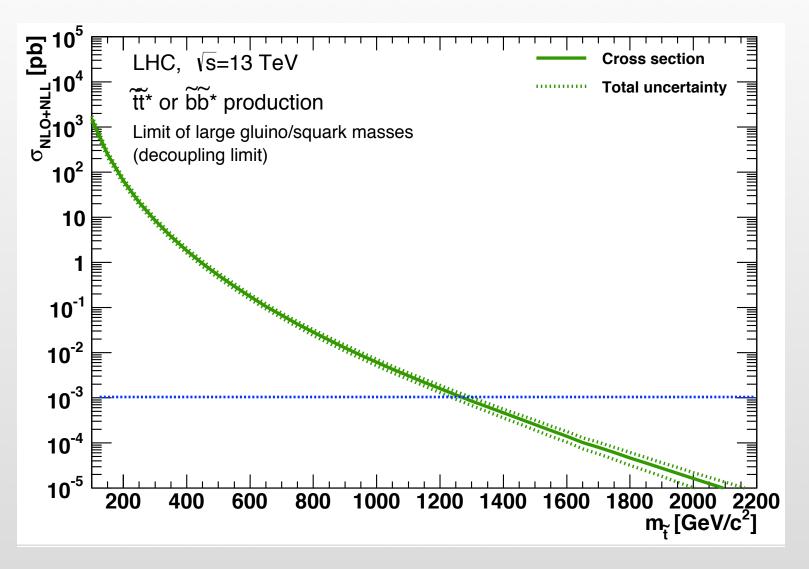
YB, Katz, Tweedie, 1309.6631

Results from LHC Run I



the non-b-tagging case should be searched for at the LHC Run II

ttbar+MET at LHC Run II



the potential limits on the t' mass are higher

Non-Relativistic Limits of Heavy Particles

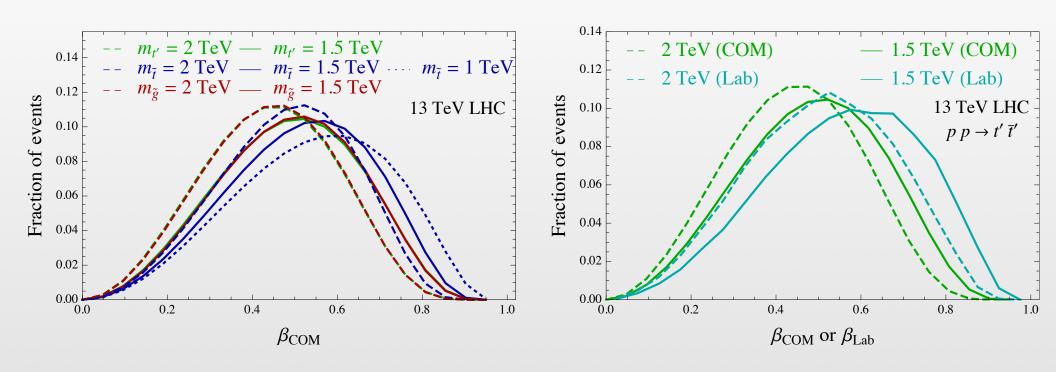
$$\hat{\sigma}(q\bar{q} \to t'\bar{t}') = \frac{4\pi\alpha_s^2\beta}{27\,\hat{s}}\left(3-\beta^2\right) \approx \frac{4\pi\alpha_s^2\beta}{9\,\hat{s}},$$
$$\hat{\sigma}(gg \to t'\bar{t}') = \frac{\pi\alpha_s^2}{48\,\hat{s}}\left[(33-18\beta^2+\beta^4)\ln\left(\frac{1+\beta}{1-\beta}\right)-\beta(59-31\beta^2)\right] \approx \frac{7\pi\alpha_s^2\beta}{48\,\hat{s}}$$

$$\begin{aligned} \hat{\sigma}(q\bar{q} \to \tilde{t}\tilde{t}^*) &= \frac{2\pi\alpha_s^2\beta^3}{27\,\hat{s}}, \\ \hat{\sigma}(gg \to \tilde{t}\tilde{t}^*) &= \frac{\pi\alpha_s^2}{96\,\hat{s}} \left[\beta(41-31\beta^2) - (17-18\beta^2+\beta^4)\ln\left(\frac{1+\beta}{1-\beta}\right)\right] \approx \frac{7\pi\alpha_s^2\beta}{96\,\hat{s}} \end{aligned}$$

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one can understand those β -dependence from s-wave or p-wave production cross sections

Velocity Distributions

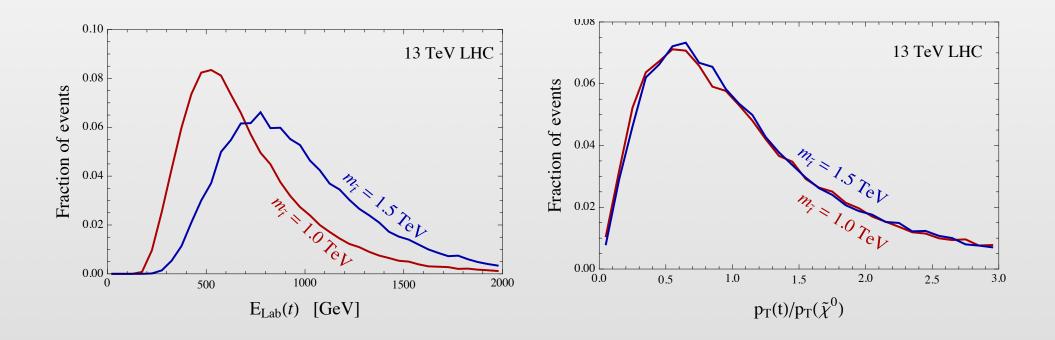


There is no sharp peak for the velocity distributions, but it may help us to understand kinematics

Now, the question is how to use this understanding to improve our searches for heavy particles

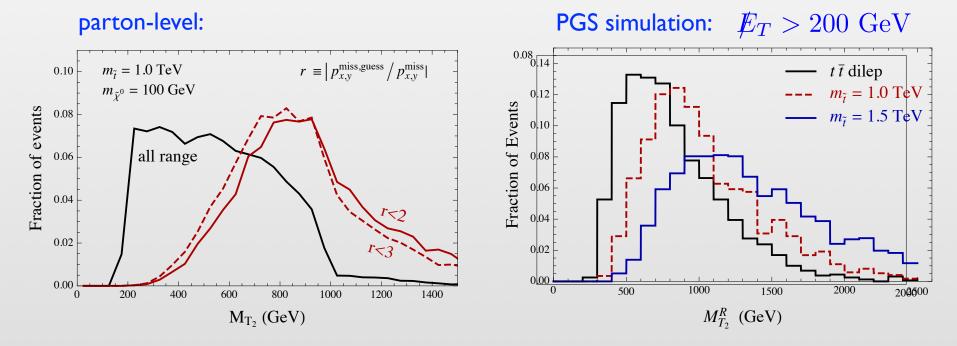
Simple Kinematics

Consider one-step decay: $\tilde{t} \rightarrow t + \tilde{\chi}^0$



so, the missing neutralino particles have similar pt's compared to the top/anti-top pt's

$M_{T_2}^R$ **Distribution**



YB, Osborne, Stefanek, work in progress

the ttbar+jets background has a tail in $M_{T_2}^R$ without a peak structure; the S/B could be improved.

Conclusions

