High-scale validity of 2HDM scenarios: Study using LHC data.

Nabarun Chakrabarty HRI, Allahabad

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- Dark matter and a radiative neutrino mass in a 2HDM.

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- Pros: Rich phenomenology, Dark Matter (DM) canditate in special cases.
- Cons: Additional doublets → Additional quartic couplings → Fast rise of such couplings → threat to perturbativity.
- Our goal: To see if there is a balance between these extremes, modulo constraints from LHC and DM experiments.

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Model: Type II 2HDM

 Most general renormalizable scalar potential for two doublets Φ₁ and Φ₂, each having hypercharge (+1)

$$\begin{split} V(\Phi_{1},\Phi_{2}) &= m_{11}^{2} \Phi_{1}^{\dagger} \Phi_{1} + m_{22}^{2} \Phi_{2}^{\dagger} \Phi_{2} - m_{12}^{2} \left(\Phi_{1}^{\dagger} \Phi_{2} + \Phi_{2}^{\dagger} \Phi_{1} \right) \\ &+ \frac{\lambda_{1}}{2} \left(\Phi_{1}^{\dagger} \Phi_{1} \right)^{2} + \frac{\lambda_{2}}{2} \left(\Phi_{2}^{\dagger} \Phi_{2} \right)^{2} \\ &+ \lambda_{3} \Phi_{1}^{\dagger} \Phi_{1} \Phi_{2}^{\dagger} \Phi_{2} + \lambda_{4} \Phi_{1}^{\dagger} \Phi_{2} \Phi_{2}^{\dagger} \Phi_{1} + \frac{\lambda_{5}}{2} \left[\left(\Phi_{1}^{\dagger} \Phi_{2} \right)^{2} + \left(\Phi_{2}^{\dagger} \Phi_{1} \right)^{2} \right] \\ &+ \lambda_{6} \Phi_{1}^{\dagger} \Phi_{1} \left(\Phi_{1}^{\dagger} \Phi_{2} + \Phi_{2}^{\dagger} \Phi_{1} \right) + \lambda_{7} \Phi_{2}^{\dagger} \Phi_{2} \left(\Phi_{1}^{\dagger} \Phi_{2} + \Phi_{2}^{\dagger} \Phi_{1} \right) \end{split}$$

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 Enlarged scalar spectrum: Mutually conjugate pair of charged scalars (H[±]), two neutral scalars (H, h) and a neutral pseudoscalar (A).

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- The Analytic forms of the 2HDM beta functions depend on the 2HDM "Type".
- We illustrate our findings in context of a Type II 2HDM.

• We demand $m_h \sim 125$ GeV to conform with the Higgs discovery@LHC and $m_{H^+} \geq 315$ GeV to avoid flavor constraints.

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- Higgs signal strength constraints at 2σ further imposed.

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• Motivates one to look beyond exact \mathbb{Z}_2 symmetry.

Include $m_{12} \neq 0$ in the scalar potential while keeping $\lambda_6, \lambda_7 = 0$. tan $\beta = 2, 10, 20$ and $m_{12} = 200, 1000$ GeV are the benchmarks chosen.

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Figure: The allowed parameter spaces in the soft \mathbb{Z}_2 breaking case for $\Lambda_{UV} = 10^{11}$ (green), 10^{16} (grey) and 10^{19} GeV (red).

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• Fixing m_{12} induces a high degree of correlation amongst the non-standard scalar masses. They loom around the TeV scale for the benchmarks taken.

• Unlike the SM, the electroweak vacuum is rendered stable even for high values of M_t .



Figure: The first two figures describe regions in the $m_{H^{\pm}}-\alpha$ plane allowed by the Higgs data in the soft \mathbb{Z}_2 breaking case. The last one depicts parameter spaces in $m_{H^-}m_A$ plane for two different values of M_t .

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• Higgs data from the LHC imposes $\alpha \sim \beta - \frac{\pi}{2}$, i.e. takes the 2HDM near the *alignment limit*.

Results: \mathbb{Z}_2 symmetry broken by *hard* terms

• λ_6 , $\lambda_7 \neq 0$. $\lambda_1(M_t) = 0.02$ and $\lambda_6(M_t) = \lambda_7(M_t)$ for computational convenience.



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• Larger parameter space opens up ensuring validity up to high scales. Thus, a Type-II 2HDM could be valid upto M_{Pl} only if the \mathbb{Z}_2 symmetry is broken either by *soft* or *hard* terms.

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The RH neutrinos play the role of generating a neutrino mass radiatively.

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The relevant Yukawa and mass terms are

$$-\mathcal{L}_{Y} = (y_{ij}\bar{N}_{i}\tilde{\Phi}_{2}^{\dagger}\ell_{j} + h.c) + \frac{M_{i}}{2}(\bar{N}_{i}^{c}N_{i} + h.c), (i, j = 1, 2, 3)$$
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In the inert limit,

$$\Phi_1 = \begin{pmatrix} G^+ \\ \frac{1}{\sqrt{2}}(v+h+iG) \end{pmatrix} \text{ and, } \Phi_2 = \begin{pmatrix} H^+ \\ \frac{1}{\sqrt{2}}(H+iA) \end{pmatrix}$$
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• $M_H > 500$ GeV: Co-annihilations among H, A and H^{\pm} (Arhrib 2013).

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$$\mathcal{M}_{ij}^{\nu} = \sum_{k=1}^{3} \frac{y_{ik} y_{jk} M_k}{16\pi^2} \left[\frac{M_H^2}{M_H^2 - M_k^2} \ln \frac{M_H^2}{M_k^2} - \frac{M_A^2}{M_A^2 - M_k^2} \ln \frac{M_A^2}{M_k^2} \right]$$
(4)

Benchmarks of *M* consistent with leptogenesis, (a) M = 110 TeV and (b) $M = 10^9$ TeV (Pilaftsis,1997; Hambye,2009). N_i -*H* co-annihilation ruled out.

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Benchmarks of *M* consistent with leptogenesis, (a) M = 110 TeV and (b) $M = 10^9$ TeV (Pilaftsis,1997; Hambye,2009). N_i -*H* co-annihilation ruled out.

• Effect of N_i in the beta functions only considered above the threshold M. $M = 10 \text{ TeV} \longrightarrow y_{\nu} = O(10^{-5}) \longrightarrow \text{ No effect on RG evolution.}$ $M = 10^9 \text{ TeV} \longrightarrow y_{\nu} = O(1) \longrightarrow \text{ Appreciable effect on RG evolution.}$

$$16\pi^2 \frac{d\lambda_2}{dt}\bigg|_{IDM+RH} = 16\pi^2 \frac{d\lambda_2}{dt}\bigg|_{IDM} + 4\lambda_2 y_{\nu}^2 - 4y_{\nu}^4 \tag{5}$$

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- Latest PLANCK data gives

$$\Omega_{\rm DM} h^2 = 0.1199 \pm 0.0027 \tag{6}$$

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• Higgs signal strengths from ATLAS and CMS satisfied at 2σ .

Results

- \blacksquare Vary 50 $GeV < M_H < 90~GeV$, 80 $GeV < M_{H^\pm}, M_A.$ Fix λ_2 and \emph{M} at two chosen benchmarks.
- Stability upto the Planck scale achieved, a radiative neutrino mass, relic density and direct detection cross section in the correct ballpark.



Results: 50 ${\rm GeV} < {\rm M_H} < 90~{\rm GeV}$



Figure: Regions compatible with the theoretical constraints for M = 110 TeV (left panel) and 10^9 TeV (right panel) with three different choices of Λ_{UV} and two values of λ_2 . The regions denoted by A (red), B (cyan) and C (green) obey these constraints up to $\Lambda_{UV} =$ 10^6 , 10^{16} and 10^{19} GeV respectively. The grey region denoted by D keeps the Higgs to diphoton signal strength within 2σ limits of the current data.

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2HDM high-scale

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Results: 50 ${\rm GeV} < {\rm M_H} < 90~{\rm GeV}$



Perturbative unitarity \rightarrow *Tight upper bound on scalar masses* + *upper bound on* λ_2

• Vacuum stability
$$\rightarrow$$
 Affected by λ_2

Results: $M_H > 500 \text{ GeV}$



• $M = 10^9$ TeV $\longrightarrow y_{\nu} = O(1) \rightarrow \lambda_2(\mu)$ starts approaching instability beyond the scale M.

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- Sizable parameter space over which electroweak vacuum remains stable upto M_{Pl} . perturbativity and unitarity constraints put stringent limits on the DM-SM Higgs coupling and the Z_2 -odd scalar masses.
- Moreover, the role played by the heavy neutrinos in the RG evolution demonstrated. They induce corrections to the purely IDM beta functions, an effect which in turn reflected in the parameter spaces so obtained.

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- In the second part, we have examined the high-scale validity of a scenario that (a) offers a scalar dark matter, (b) radiatively generates mass for the light neutrinos.
- Sizable parameter space over which electroweak vacuum remains stable upto M_{Pl} . perturbativity and unitarity constraints put stringent limits on the DM-SM Higgs coupling and the Z_2 -odd scalar masses.
- Moreover, the role played by the heavy neutrinos in the RG evolution demonstrated. They induce corrections to the purely IDM beta functions, an effect which in turn reflected in the parameter spaces so obtained.

Thank you for your attention

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16.07.2015 20 / 30

Model features: Scalar sector

• m_{11} and m_{22} eliminated using the tadpole conditions.

The masses of the physical scalars,

$$\begin{split} m_A^2 &= \frac{m_{12}^2}{s_\beta c_\beta} - \frac{1}{2} v^2 \left(2\lambda_5 + \frac{\lambda_6}{t_\beta} + \lambda_7 t_\beta \right), \\ m_{H^\pm}^2 &= m_A^2 + \frac{1}{2} v^2 \left(\lambda_5 - \lambda_4 \right), \\ m_h^2 &= \frac{1}{2} \left[(A + B) - \sqrt{(A - B)^2 + 4C^2} \right], \\ m_H^2 &= \frac{1}{2} \left[(A + B) + \sqrt{(A - B)^2 + 4C^2} \right], \\ \tan 2\alpha &= \frac{2C}{A - B}, \end{split}$$

where we have defined,

$$\begin{aligned} A &= m_A^2 s_\beta^2 + v^2 (\lambda_1 c_\beta^2 + \lambda_5 s_\beta^2 + 2\lambda_6 s_\beta c_\beta), \\ B &= m_A^2 c_\beta^2 + v^2 (\lambda_2 s_\beta^2 + \lambda_5 c_\beta^2 + 2\lambda_7 s_\beta c_\beta), \\ C &= -m_A^2 s_\beta c_\beta + v^2 \left[(\lambda_3 + \lambda_4) s_\beta c_\beta + \lambda_6 c_\beta^2 + \lambda_7 s_\beta^2 \right]. \end{aligned}$$

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Constraints: Theoretical

 Vacuum stability: Stability of the electroweak vacuum is ensured up to some specified energy scale if the following conditions are satisfied for all scales Q up to that scale,

$$\begin{array}{ll} \operatorname{vsc1} & : & \lambda_1(Q) > 0 \\ \operatorname{vsc2} & : & \lambda_2(Q) > 0 \\ \operatorname{vsc3} & : & \lambda_3(Q) + \sqrt{\lambda_1(Q)\lambda_2(Q)} > 0 \\ \operatorname{vsc4} & : & \lambda_3(Q) + \lambda_4(Q) - |\lambda_5(Q)| + \sqrt{\lambda_1(Q)\lambda_2(Q)} > 0 \end{array}$$

Perturbativity:

$$\lambda_i(Q) < 4\pi$$

The corresponding constraints for the Yukawa and gauge interactions are,

$$y_i(Q), \ g_i(Q) < \sqrt{4\pi}$$

Results: $M_H > 500 \text{ GeV}$ regime

BP	M _H	$M_{H^{\pm}}$	M _A	λ_L	λ_2
BP1	850.0 GeV	854.0 GeV	858.0 GeV	0.02	0.1
BP2	710.0 GeV	712.0 GeV	711.0 GeV	0.11	0.1

Table: Benchmark values (BP) of parameters affecting the RG evolution of the quartic couplings. For each BP, two values of M, namely, 110 TeV and 10⁹ TeV, have been used.



Figure: RG running of different scalar quartic couplings corresponding to BP1. The solid, dashed, dashed dotted and dotted lines denote the evolution curves of the stability conditions vsc1, vsc2, vsc3 and vsc4 respectively.

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2HDM high-scale

- The theoretically allowed parameter space is fully consistent with the Higgs to diphoton LHC data in this case.
- The parameter space valid until the Planck scale and corresponding to M = 110 TeV shrinks significantly for $M = 10^9 \text{ TeV}$.
- Interestingly, y_{ν} becomes large $\mathcal{O}(10^{-1})$ for $M = 10^9$ TeV. Such a large yukawa coupling contributes to the beta function of λ_2 through the terms $+\lambda_2 y_{\nu}^2$ and $-y_{\nu}^4$ that either makes λ_2 perturbative in some cases or λ_2 negative and the vacuum unstable in other.
- \blacksquare The above argument best demonstrated through a display of RG evolution of $\lambda_2.$

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Results: $M_H > 500 \text{ GeV}$ regime



Figure: RG running of different scalar quartic couplings corresponding to BP2. The solid, dashed, dashed dotted and dotted lines denote the evolution curves of the stability conditions vsc1, vsc2, vsc3 and vsc4 respectively.

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Constraints: Theoretical

• Unitarity: A further set of conditions come on demanding unitarity of the scattering matrix comprising all $2 \rightarrow 2$ channels involving, by the optical theorem. Each distinct eigenvalue (shown below) of the aforementioned amplitude matrix be bounded above at 8π .

$$\begin{aligned} a_{\pm} &= \frac{3}{2}(\lambda_1 + \lambda_2) \pm \sqrt{\frac{9}{2}(\lambda_1 - \lambda_2)^2 + (2\lambda_3 + \lambda_4)^2} \\ b_{\pm} &= \frac{1}{2}(\lambda_1 + \lambda_2) \pm \sqrt{\frac{1}{4}(\lambda_1 - \lambda_2)^2 + \lambda_5^2}, \\ c_{\pm} &= d_{\pm} = \frac{1}{2}(\lambda_1 + \lambda_2) \pm \sqrt{\frac{1}{4}(\lambda_1 - \lambda_2)^2 + \lambda_5^2}, \\ e_1 &= (\lambda_3 + 2\lambda_4 - 3\lambda_5), \\ e_2 &= (\lambda_3 - \lambda_5), \\ f_1 &= f_2 = (\lambda_3 + \lambda_4), \\ f_+ &= (\lambda_3 + 2\lambda_4 + 3\lambda_5), \\ f_- &= (\lambda_3 + \lambda_5). \end{aligned}$$

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16.07.2015 27 / 30

The RG equations for the gauge couplings, for this model, are given by [?],

$$\begin{split} 16\pi^2 \frac{dg_s}{dt} &= -7g_s^3, \\ 16\pi^2 \frac{dg}{dt} &= -3g^3, \\ 16\pi^2 \frac{dg'}{dt} &= 7{g'}^3. \end{split}$$

Here g', g and g_s denote the U(1), SU(2)_L and SU(3)_c gauge couplings respectively.

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One-loop beta functions

$$\begin{split} 16\pi^2 \frac{d\lambda_1}{dt} &= 12\lambda_1^2 + 4\lambda_3^2 + 4\lambda_3\lambda_4 + 2\lambda_4^2 + 2\lambda_5^2 + \frac{3}{4}(3g^4 + g'^4 + 2g^2g'^2) \\ &-\lambda_1(9g^2 + 3g'^2 - 12y_t^2 - 12y_b^2 - 4y_\tau^2) - 12y_t^4 \,, \\ 16\pi^2 \frac{d\lambda_2}{dt} &= 12\lambda_2^2 + 4\lambda_3^2 + 4\lambda_3\lambda_4 + 2\lambda_4^2 + 2\lambda_5^2 \\ &+ \frac{3}{4}(3g^4 + g'^4 + 2g^2g'^2) - 3\lambda_2(3g^2 + g'^2 - \frac{4}{3}y_{\nu}^2) - 4y_{\nu}^4 \,, \\ 16\pi^2 \frac{d\lambda_3}{dt} &= (\lambda_1 + \lambda_2)(6\lambda_3 + 2\lambda_4) + 4\lambda_3^2 + 2\lambda_4^2 + 2\lambda_5^2 + \frac{3}{4}(3g^4 + g'^4 - 2g^2g'^2) \\ &-\lambda_3(9g^2 + 3g'^2 - 6y_t^2 - 6y_b^2 - 2y_\tau^2 - 2y_{\nu}^2) \,, \\ 16\pi^2 \frac{d\lambda_4}{dt} &= 2(\lambda_1 + \lambda_2)\lambda_4 + 8\lambda_3\lambda_4 + 4\lambda_4^2 + 8\lambda_5^2 + 3g^2g'^2 \\ &-\lambda_4(9g^2 + 3g'^2 - 6y_t^2 - 6y_b^2 - 2y_\tau^2 - 2y_{\nu}^2) \,, \\ 16\pi^2 \frac{d\lambda_5}{dt} &= (2\lambda_1 + 2\lambda_2 + 8\lambda_3 + 12\lambda_4)\lambda_5 - \lambda_5(9g^2 + 3g'^2 - 6y_b^2 - 2y_\tau^2 - 6y_t^2 - 6y_t^2 - 6y_t^2 - 6y_{\nu}^2) \,. \end{split}$$

$$\begin{split} &16\pi^2 \frac{dy_b}{dt} &= y_b \left(-8g_s^2 - \frac{9}{4}g^2 - \frac{5}{12}g'^2 + \frac{9}{2}y_b^2 + y_\tau^2 + \frac{3}{2}y_t^2 \right) \,, \\ &16\pi^2 \frac{dy_t}{dt} &= y_t \left(-8g_s^2 - \frac{9}{4}g^2 - \frac{17}{12}g'^2 + \frac{9}{2}y_t^2 + y_\tau^2 + \frac{3}{2}y_b^2 \right) \,, \\ &16\pi^2 \frac{dy_\tau}{dt} &= y_\tau \left(-\frac{9}{4}g^2 - \frac{15}{4}g'^2 + 3y_b^2 + 3y_t^2 + \frac{1}{2}y_\nu^2 + \frac{5}{2}y_\tau^2 \right) \,. \\ &16\pi^2 \frac{dy_\nu}{dt} &= y_\tau \left(-\frac{9}{4}g^2 - \frac{3}{4}g'^2 - \frac{3}{4}y_\tau^2 + \frac{5}{2}y_\nu^2 \right) \,. \end{split}$$

16.07.2015 30 / 30

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