

European neutrino project(s): status

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Open Meeting for the Hyper-Kamiokande Project
21-23 August 2012 Kavli Institute for the Physics and Mathematics
of the Universe (Kavli IPMU), The University of Tokyo

Foreword

- Given limited time available, this talk will **focus on the most relevant topic for this meeting**, namely the future long baseline option in Europe, presenting in particular the LAGUNA-LBNO project status and plans.
- Importance of the LAGUNA/LAGUNA-LBNO systematic approach in the decision process:
 - ➔ Seven sites visited and studied in much detail
 - ➔ Three detector technologies (WCD, LAr and LSc)
- For a recent discussion and general recommendations, see:
 - ➔ **S. Bertolucci et al., “European Strategy for Accelerator-Based Neutrino Physics”, arXiv:1208.0512.**
 - ➔ Submitted as input to the European Strategy Preparatory Group.
 - ➔ To be presented and discussed at the Krakow Symposium, September 10-12th 2012 (<http://indico.cern.ch/conferenceCFA.py?confId=175067>).
- Message: Neutrino physics is an important field, tightly related to physics beyond the SM. The possibility to host a new long baseline experiment beyond the present CERN-Gran Sasso (CNGS) is directly or indirectly mentioned by several input documents. See in particular the national or institutional roadmaps to the CERN Strategy.

CERN European Strategy for Accelerator-Based Neutrino Physics

arXiv:1208.0512

1. **Neutrinos must be part of the CERN Roadmap.**
2. **Large discovery potential:** The determination of the neutrino mass hierarchy and the determination of the CP phase are the next steps in long baseline neutrino experiments. These fundamental measurements require and justify dedicated long baseline accelerator-based experiments.
3. **LAGUNA-LBNO and CERN→Pyhäsalmi:** The next step should be an experiment which can start now and be constructed in a reasonable time (less than about 10 years), maintains the community healthy, with a real chance of discovery and long term upgrade possibilities. The existence of a possible long baseline in Europe from CERN to Pyhäsalmi (2300 km) is unique in this regard.
4. **Incremental approach:** The LBNO project, considering an initial 20 kton fine grain LAr tracking-calorimeter (GLACIER) and a magnetized muon detector (MIND) is the first priority of the LAGUNA-LBNO consortium and is endorsed by the Neutrino Factory community. An **Expression of Interest**, signed by enlarged consortium, has been submitted to the CERN SPSC and is presently being reviewed.
5. **Preparing for longer term, precision experiments:** The European Strategy for Particle Physics must provide for European participation in the programme required for a Neutrino Factory proposal (in particular NuSTORM) to be prepared in time for the next update of the European Strategy (2018 ?).

LBNO

**Expression of Interest for a
very long baseline neutrino
oscillation experiment**

CERN-SPSC-2012-021 ; SPSC-EOI-007

~230 authors, 51 institutions

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The LBNO long baseline goals

- LBNO is a next generation long baseline experiment which aims at a significantly better sensitivity than what is achievable with the combined T2K, NOvA and reactors experiments.
- **LBNO will explicitly observe MH induced matter effects and CP-violation, which is different from simply extracting the hierarchy or δ_{CP} value from global fits of all available data:**
 - ★ Large detectors and intense beam for a significant increase in statistics
 - ★ Measure all active-active transitions (e / mu / tau CC) and active-sterile (NC) at long baseline
 - ★ A precise investigation of the oscillation probabilities as a function of energy (L/E) and a direct comparison of neutrino and antineutrino behaviors to verify the expectations from 3-generations neutrino mixing.
 - ★ A very long baseline to have an excellent separation of the asymmetry due to the matter effects (i.e. the mass hierarchy measurement) and the CP asymmetry due to the δ_{CP} complex phase, and thus to break the parameter degeneracies, and to “see” the 1st and 2nd maxima !
 - ★ To directly observe the different MH induced matter- and CP-phase induced effects in oscillation probabilities for neutrinos and antineutrinos !
- Extend nucleon decay searches, a unique probe for BSM up to the Grand Unification Scale
- Perform very compelling and complementary atmospheric and astrophysical neutrino detection programs, which become accessible when the detector is deep underground.



A new massive deep underground neutrino observatory for long baseline neutrino studies, capable of proton decay searches, atmospheric and astrophysical neutrino detection

The LBNO requirements

Beam

Fully exploit long baseline neutrino oscillation pattern



perform L/E analysis over large energy range
(1st and 2nd maxima)



Wide Band Beam (WBB)

$$E_{\nu}^{2nd\ max} \gtrsim 0.5\ \text{GeV} \implies L \gtrsim 1000\ \text{km}$$

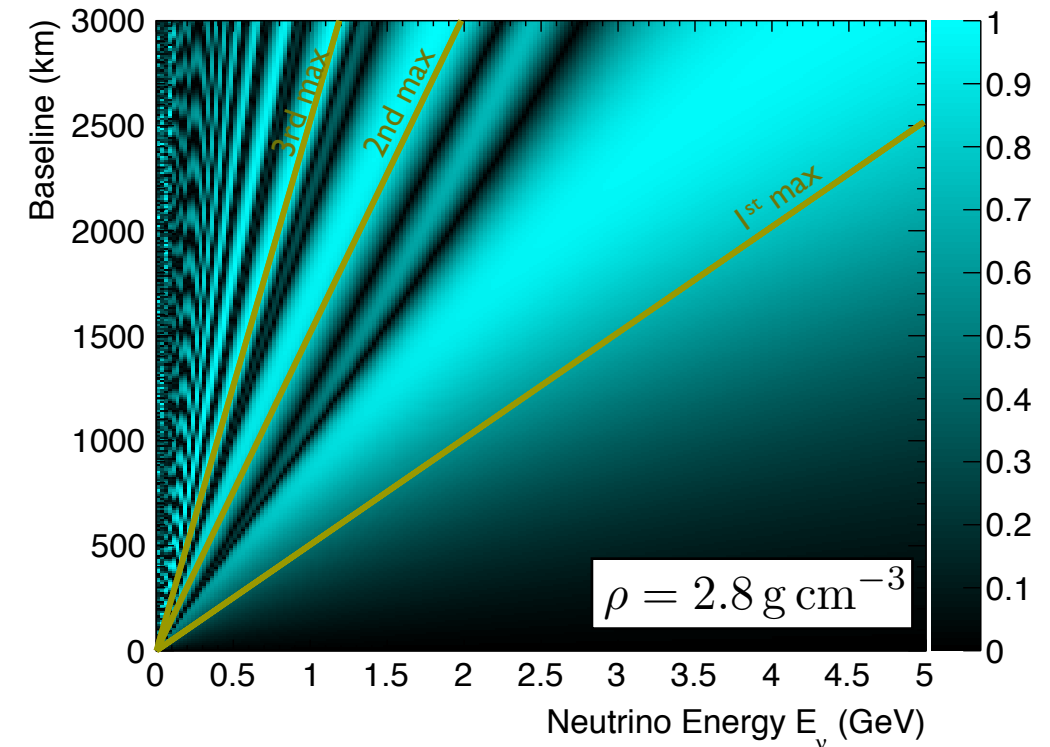
Detector

Better signal efficiency and background rejection
with a comparable mass

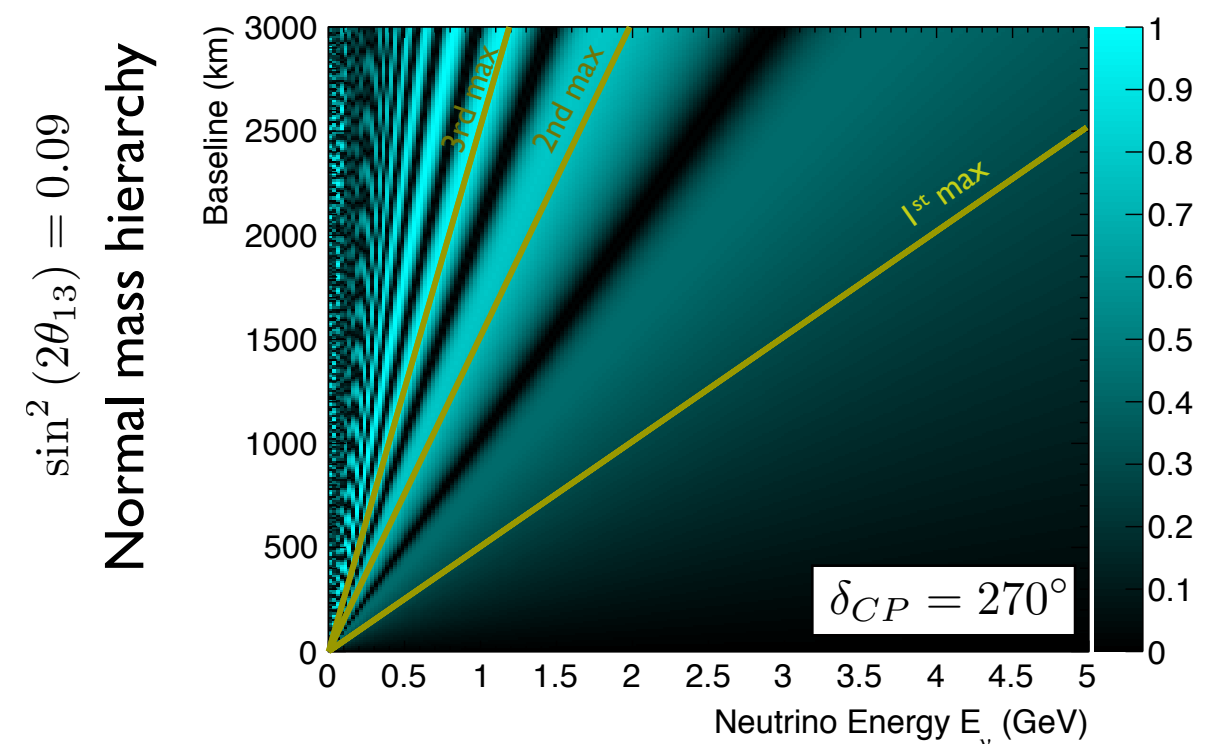


20 kton fine sampling tracking device
and magnetized muon detector

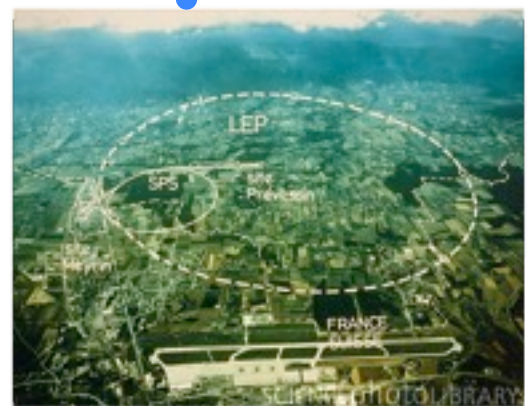
$$\mathcal{A}_{CP}(\rho) \equiv \text{abs} \left(\frac{P^{mat}(\nu) - P^{mat}(\bar{\nu})}{P^{mat}(\nu) + P^{mat}(\bar{\nu})} \right)$$



$$\mathcal{A}_{CP}^{vac}(\delta_{CP}) \equiv \text{abs} \left(\frac{P^{vac}(\nu) - P^{vac}(\bar{\nu})}{P^{vac}(\nu) + P^{vac}(\bar{\nu})} \right)$$



Pyhäsalmi far site location



2100 km from RAL, 1500 km from DESY, and 1160 km from Protvino.

- ▶ CUPP : Centre for Underground Physics in Pyhäsalmi (www.cupp.fi)
- ▶ Location: $63^{\circ} 39' 31''\text{N} - 26^{\circ} 02' 48''\text{E}$
- ▶ Distances (by roads)
 - ▶ Oulu – 165 km
 - ▶ Jyväskylä – 180 km
 - ▶ Helsinki – 450 km
- ▶ Distance to CERN 2300 km
- ▶ Good traffic connections
 - ▶ the main highway: Helsinki – Jyväskylä – Oulu – ...
 - ▶ the second busiest airport in Oulu
 - ▶ rail yard at the mine
- ▶ Inhabitants: ~6000

Being extensively investigated in LAGUNA DS since 2008

Extended site investigation foreseen for 2013-2014



This pump alone takes all the water from 645 m to the surface



250 m long tunnel and a cavern at 1400m excavated for LAGUNA R&D



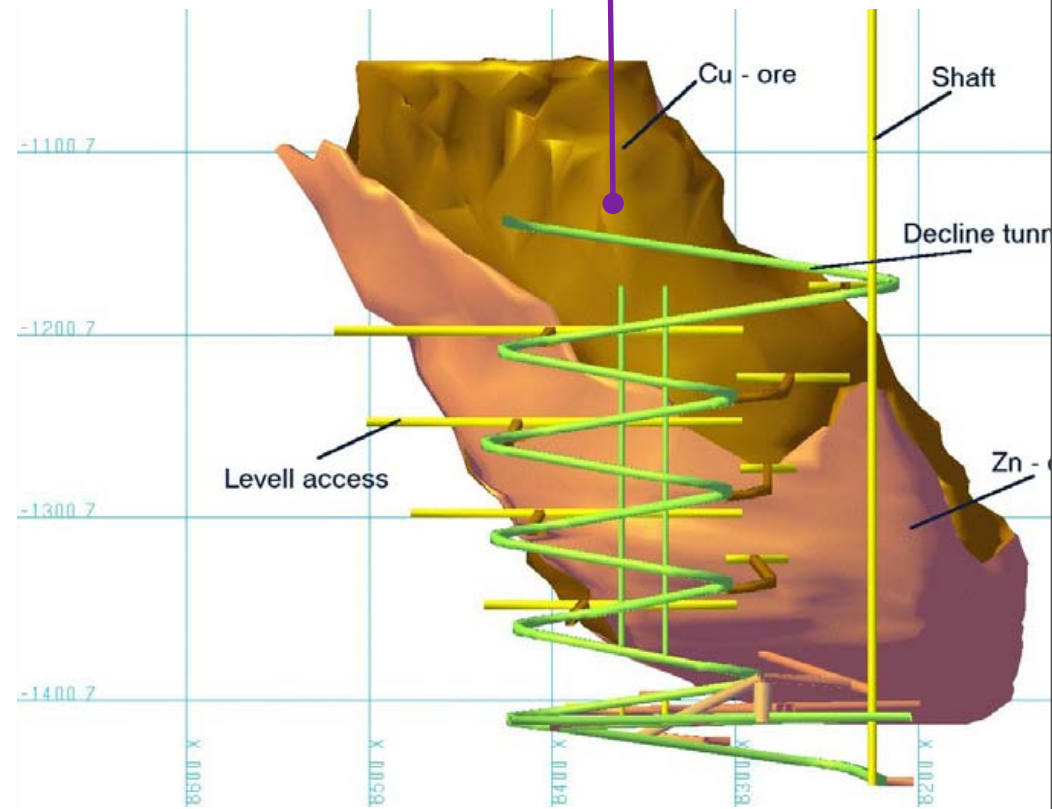
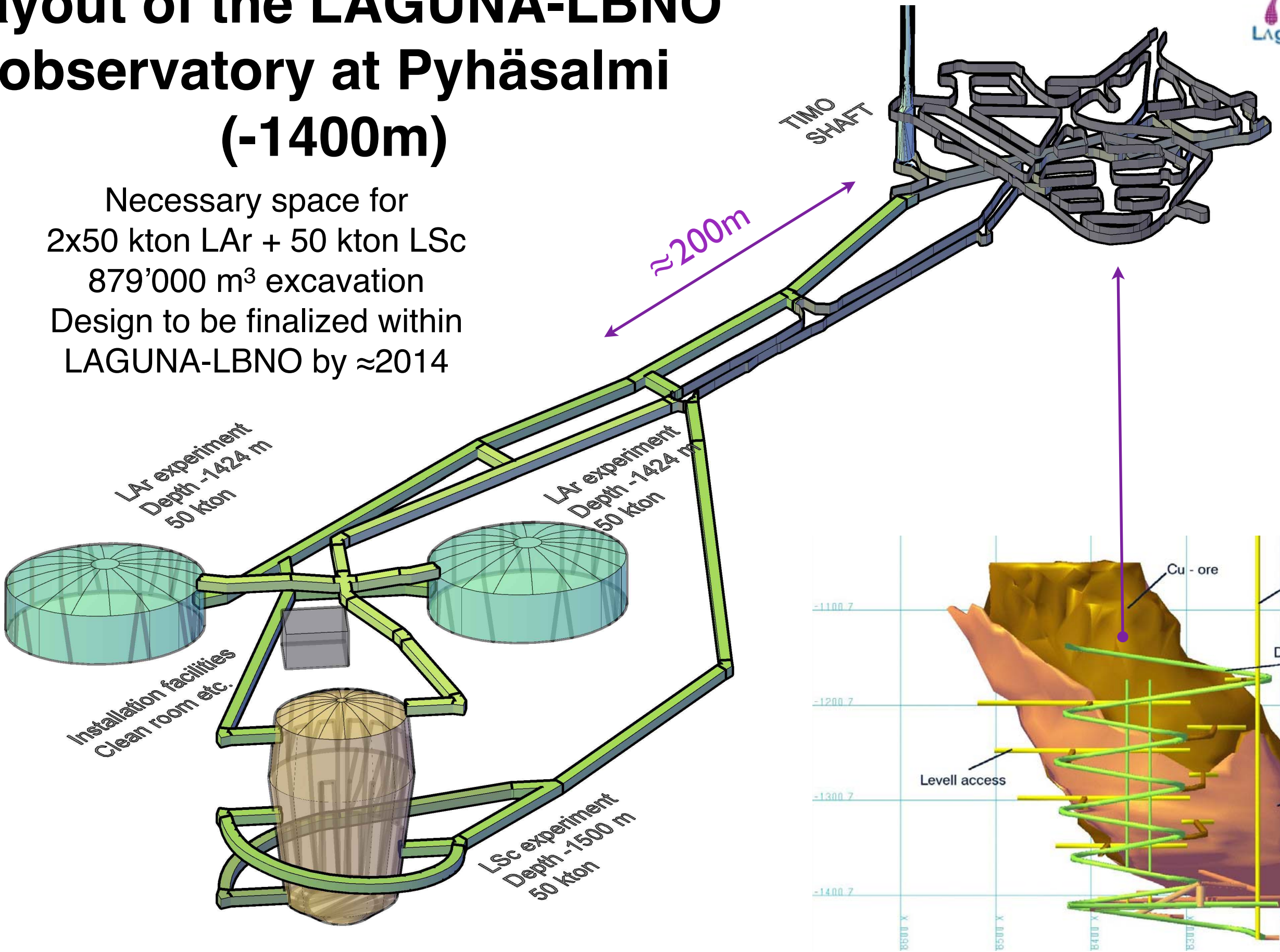
Cafeteria, meeting room and sauna at 1400 m below ground



Mobile phones work and internet available also at 1400 m

Layout of the LAGUNA-LBNO observatory at Pyhäsalmi (-1400m)

Necessary space for
 2x50 kton LAr + 50 kton LSc
 879'000 m³ excavation
 Design to be finalized within
 LAGUNA-LBNO by ≈ 2014



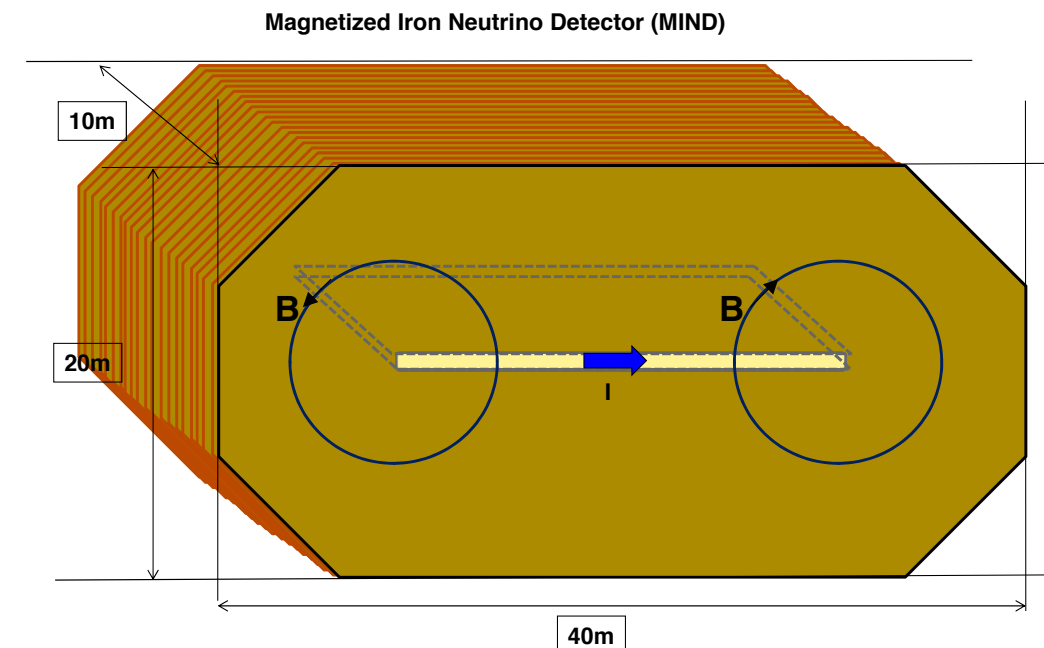
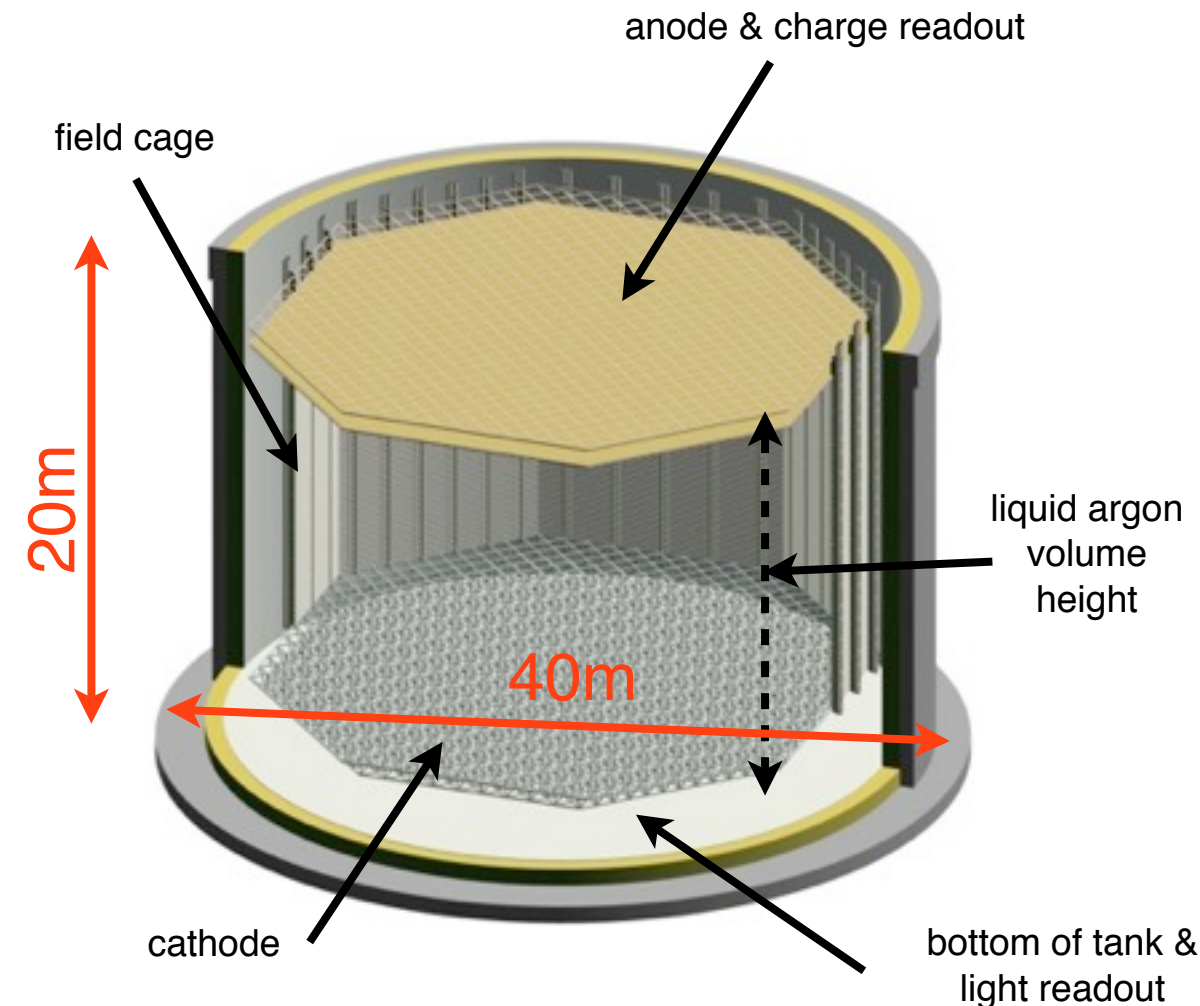
Far underground detectors

- **20 kton double phase LAr LEM TPC (GLACIER): best detector for electron appearance measurements with excellent energy resolution and small systematic errors**

- ▶ Very fine grain tracking-calorimeter
- ▶ Exclusive final states, low energy threshold on all particles
- ▶ Excellent ν energy resolution and reconstruction ability from sub GeV to a few GeV, from single prong to high multiplicity
- ▶ Suitable for spectrum measurement with needed wide energy coverage
- ▶ Excellent π^0 /electron discrimination
- ▶ Best detector for baselines $> 300\text{km}$

- **35 kton magnetized Muon Detector (MIND): conventional and well-proven detector for muon CC, and NC**

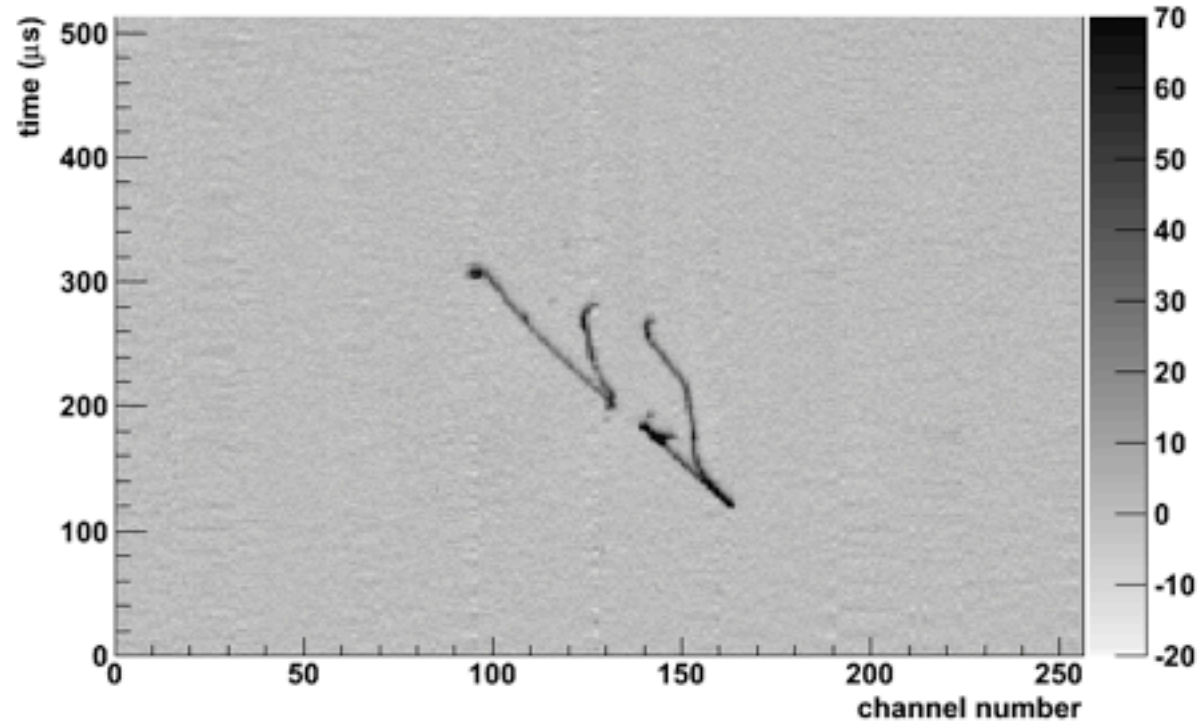
- ▶ muon momentum & charge determination, inclusive total neutrino energy
- ▶ $\text{rs}\mu/\text{ws}\mu$ with Neutrino Factory
- ▶ 3cm Fe plates, 1cm scintillator bars, $B=1.5\text{-}2.5\text{ T}$



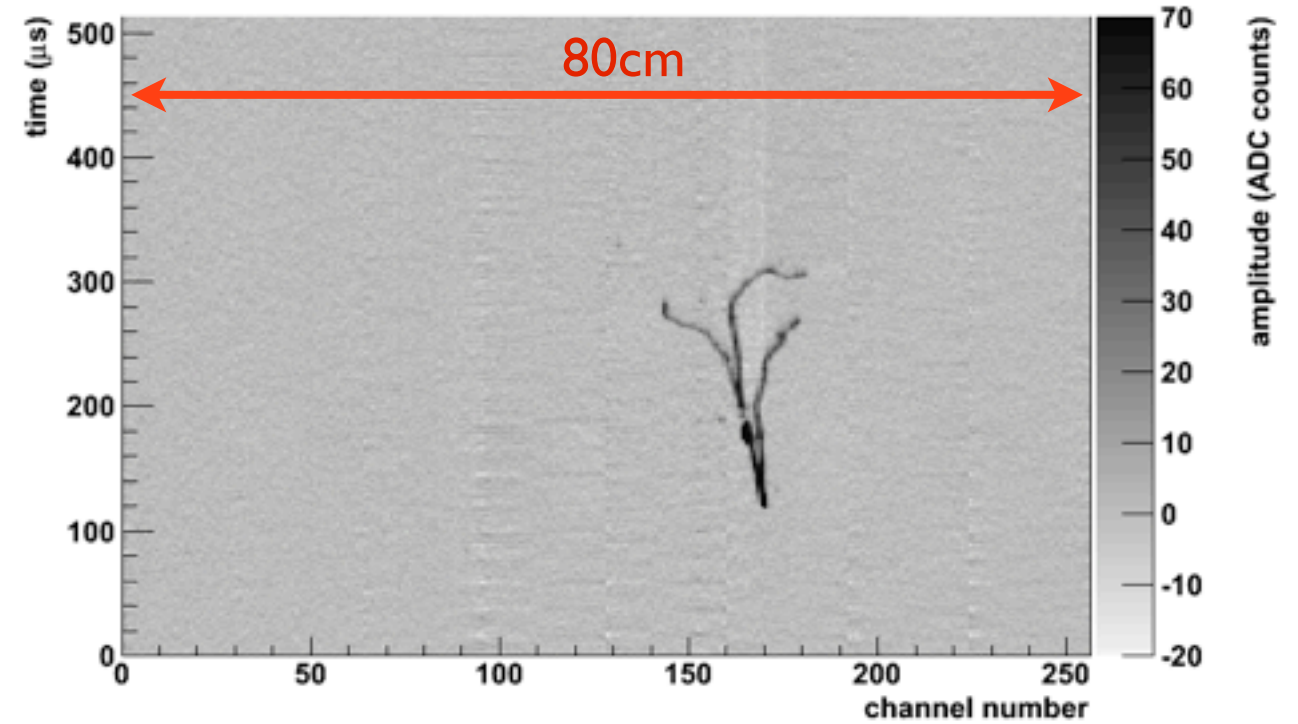
Real cosmic rays in LAr LEM-TPC

Cosmic track in double phase 80x40cm² LAr-LEM TPC with adjustable gain : $S/N > 100$ for m.i.p !!

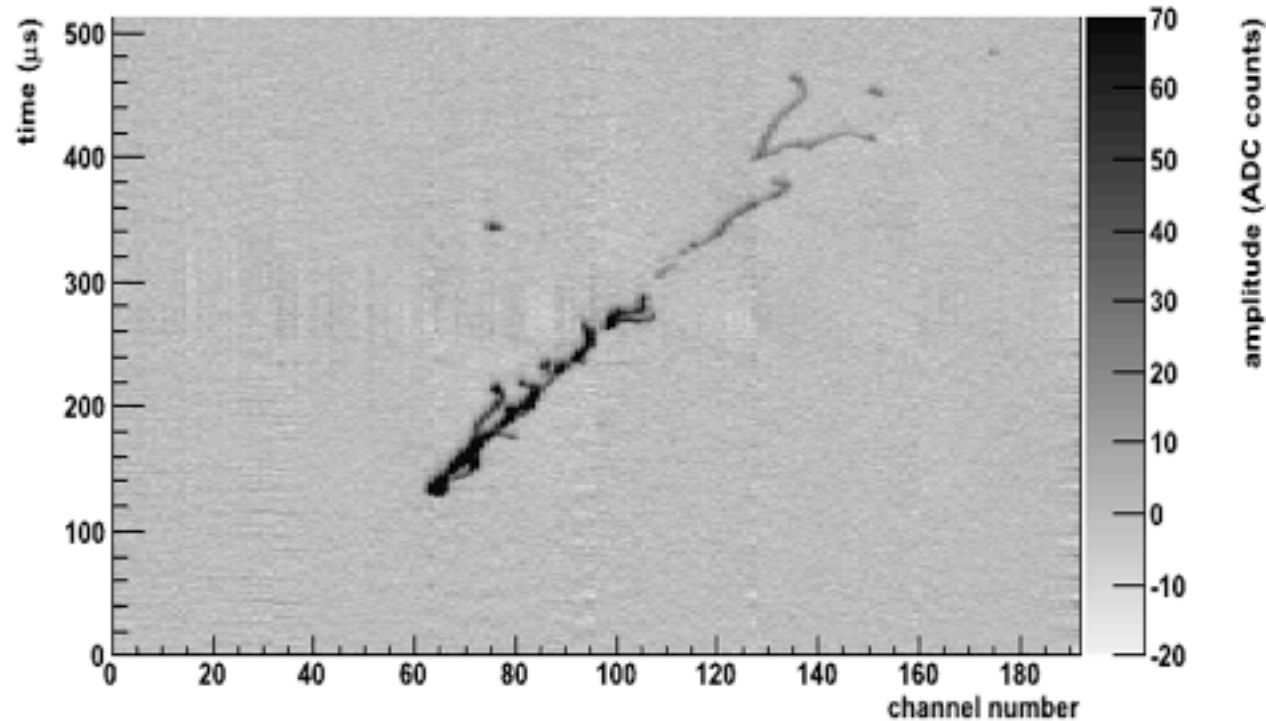
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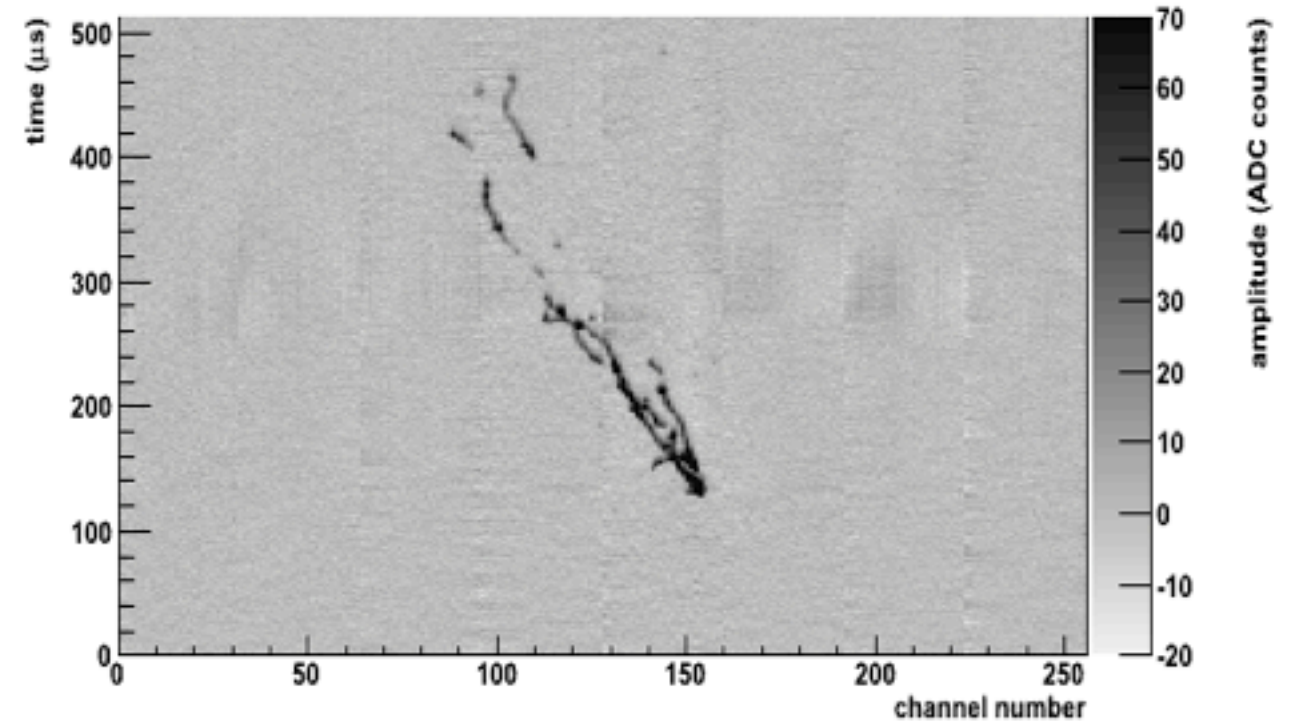
View 1: Event display (run 14456, event 8044)



View 0: Event display (run 14450, event 1511)



View 1: Event display (run 14450, event 1511)

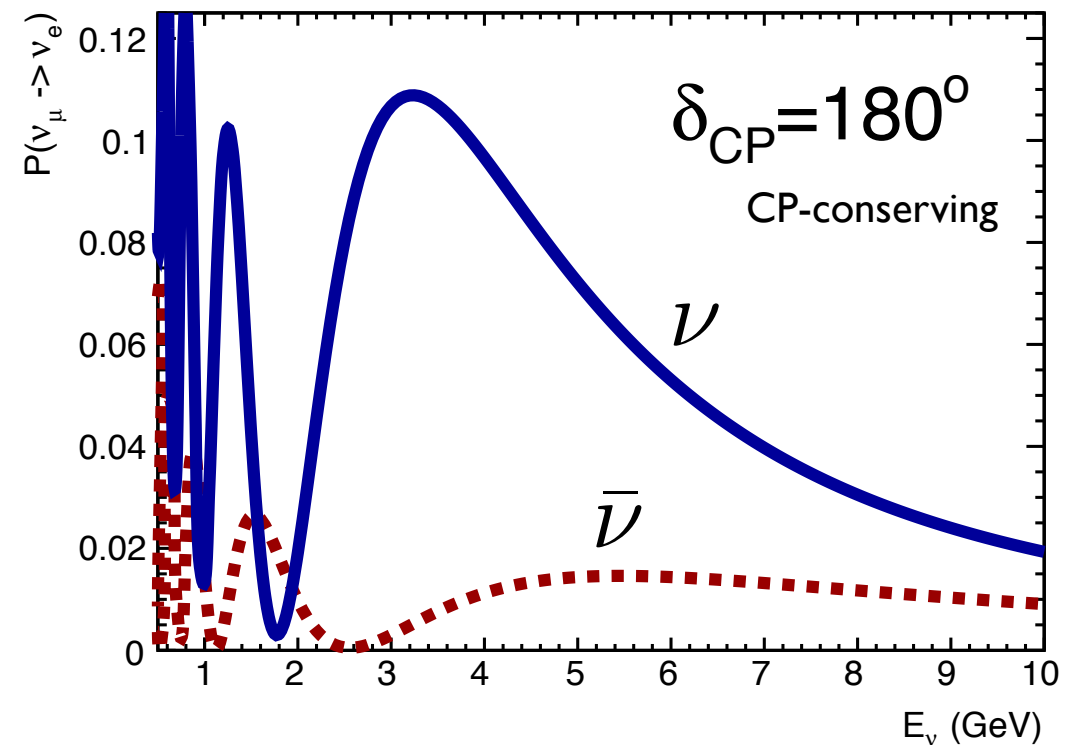
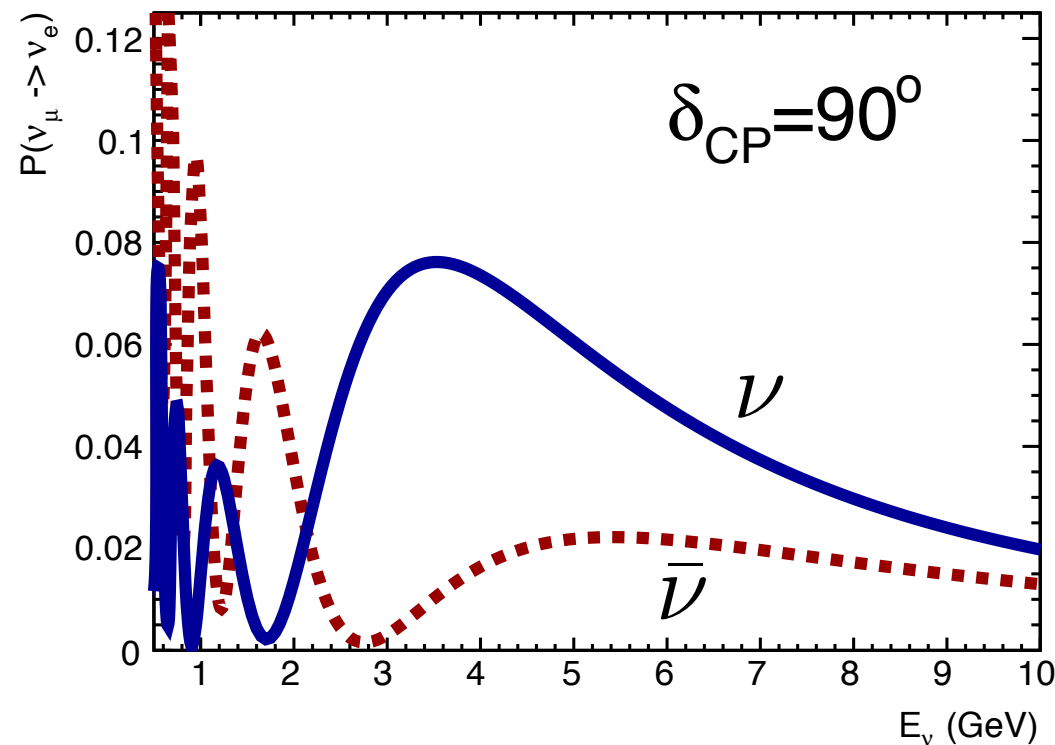
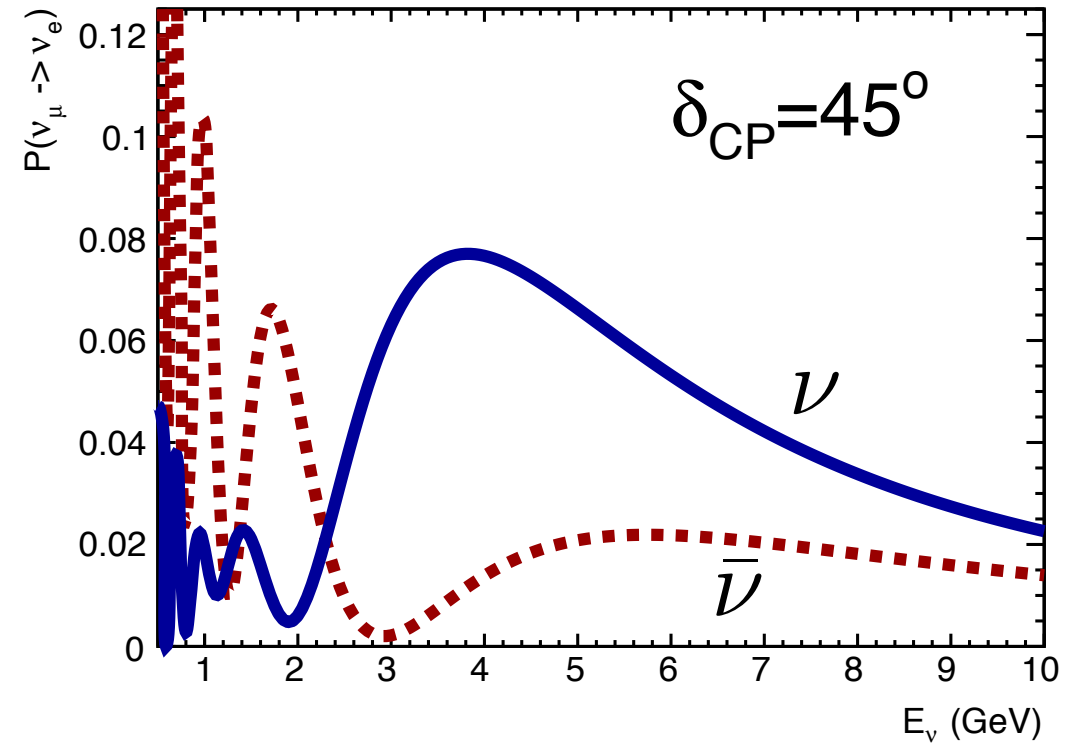
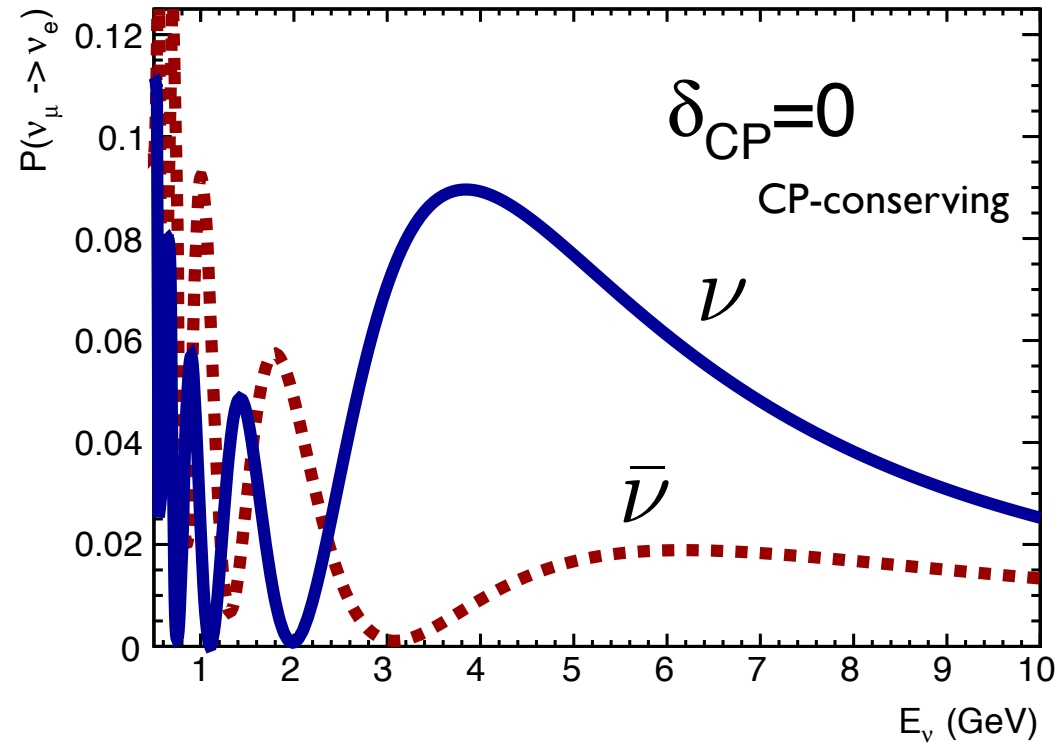


CERN-Pythäsalmi: spectral information $\nu_\mu \rightarrow \nu_e$

★ Normal mass hierarchy

$L=2300$ km

$$\sin^2(2\theta_{13}) = 0.09$$

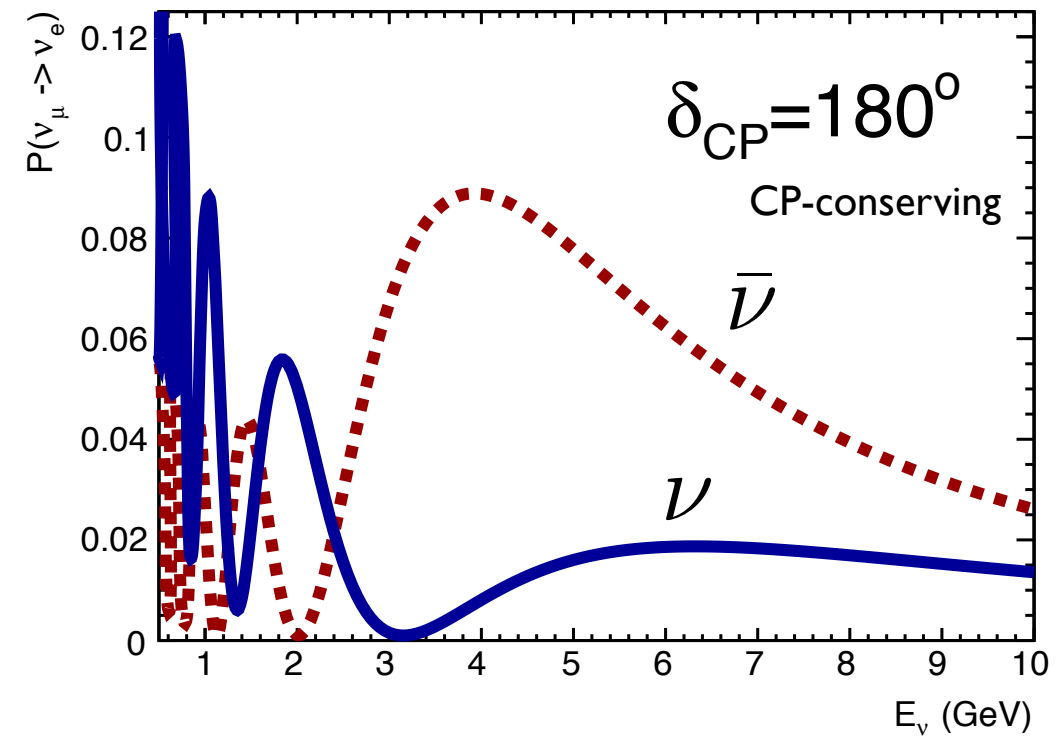
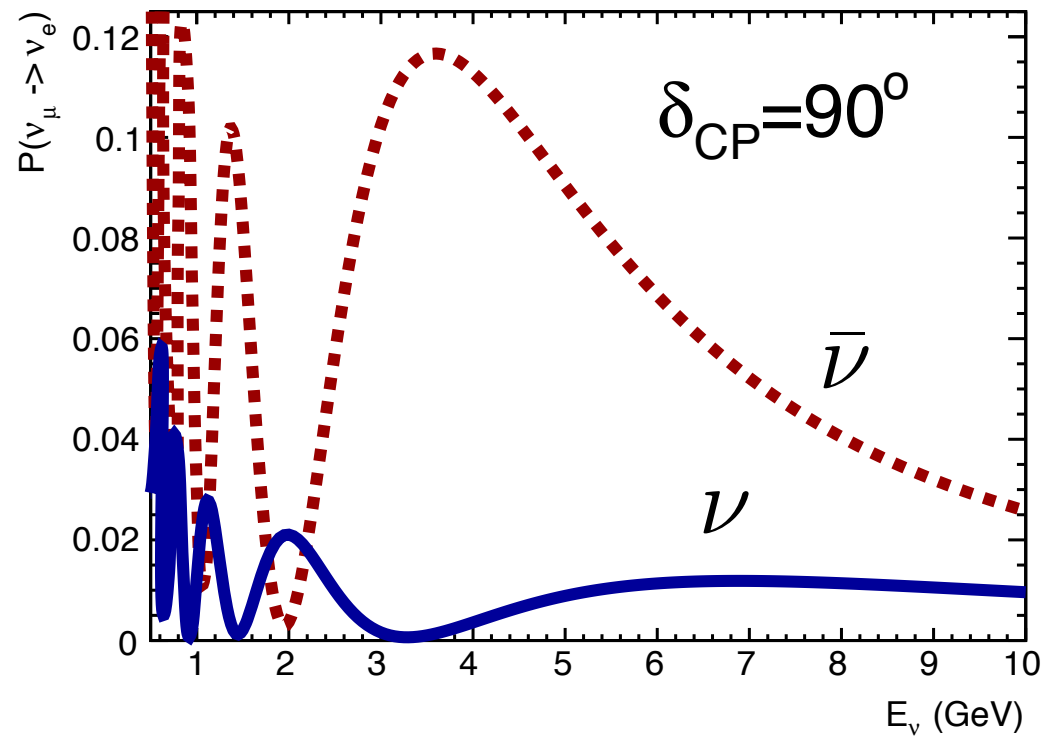
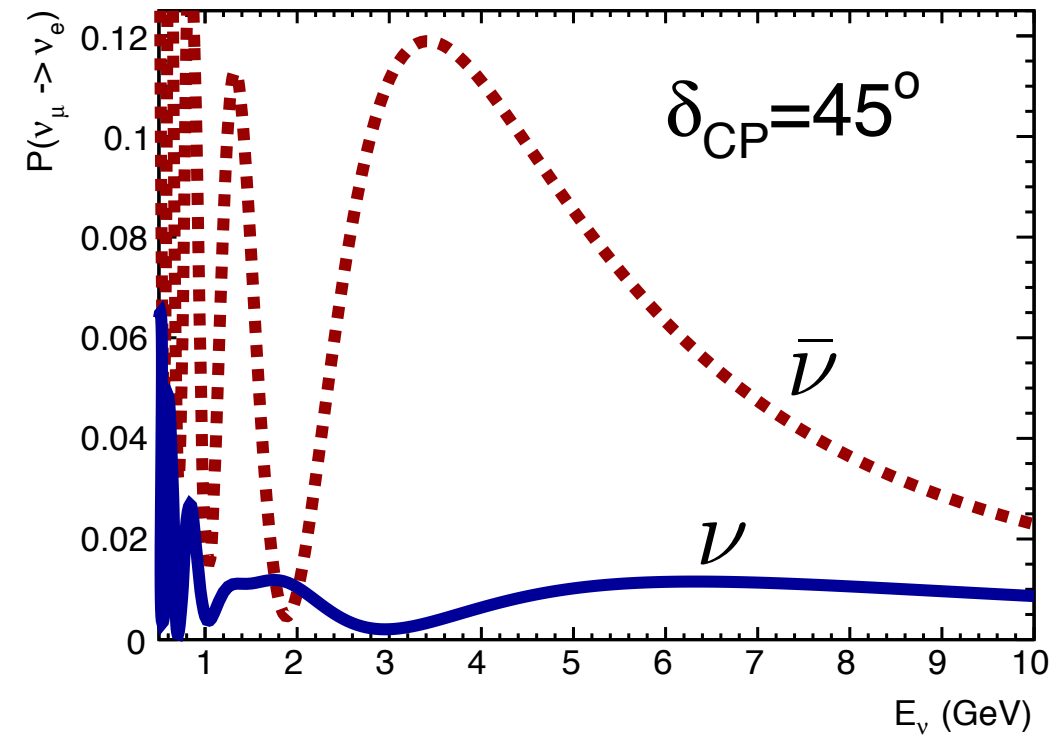
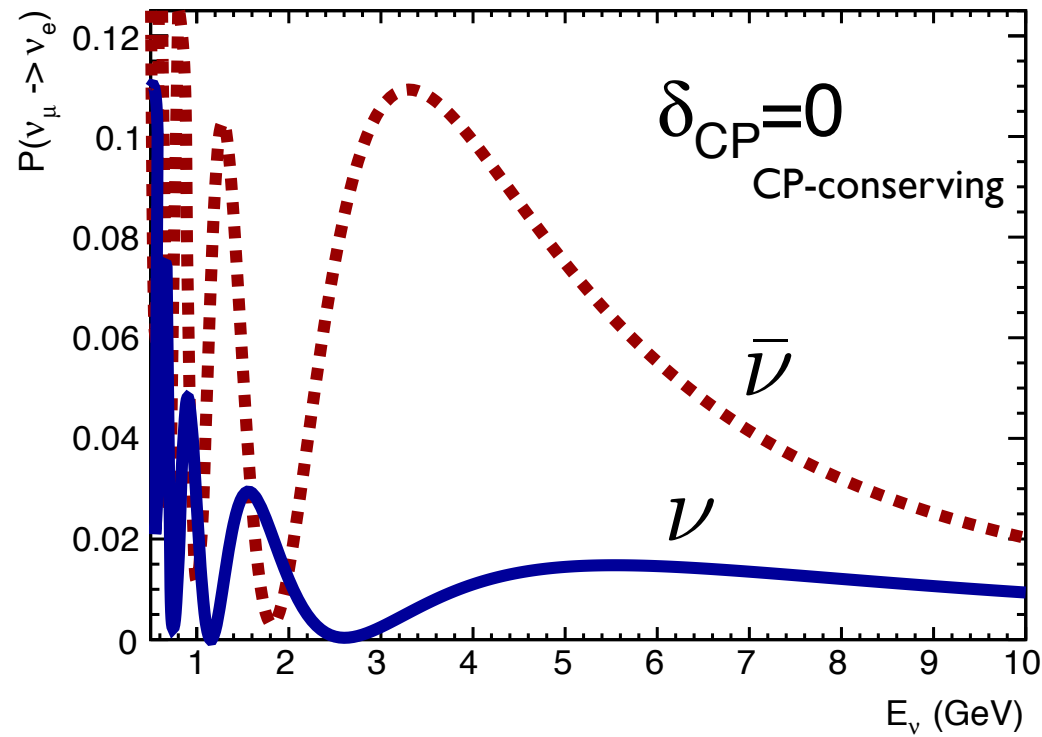


CERN-Pythäsalmi: spectral information $\nu_\mu \rightarrow \nu_e$

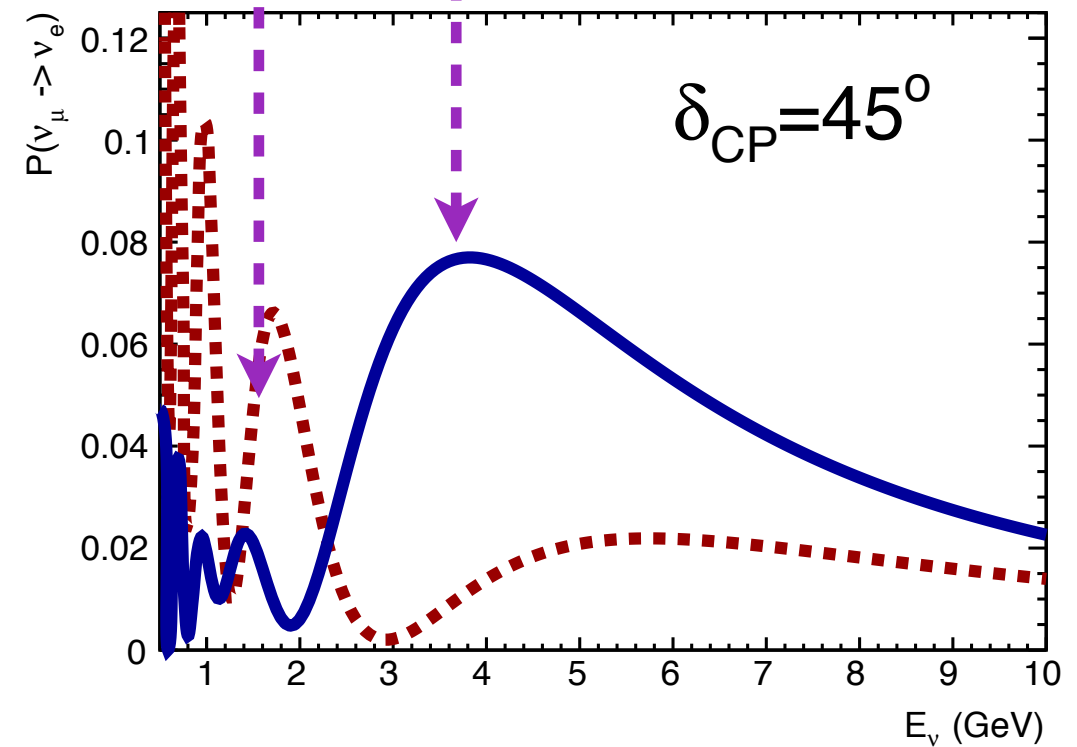
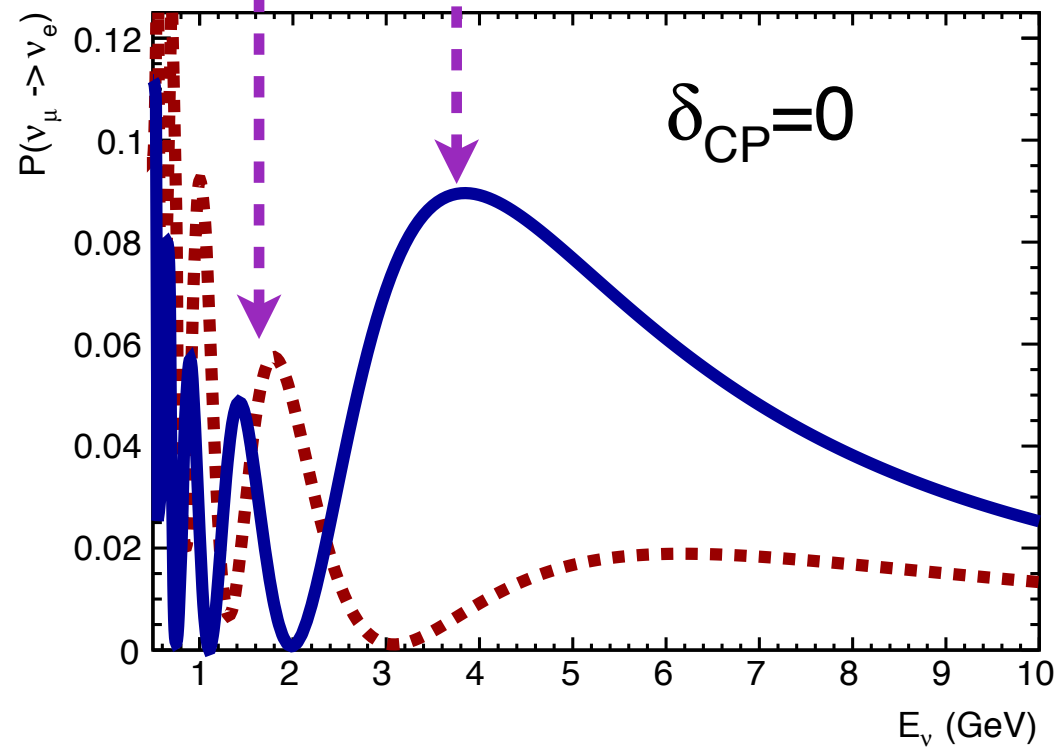
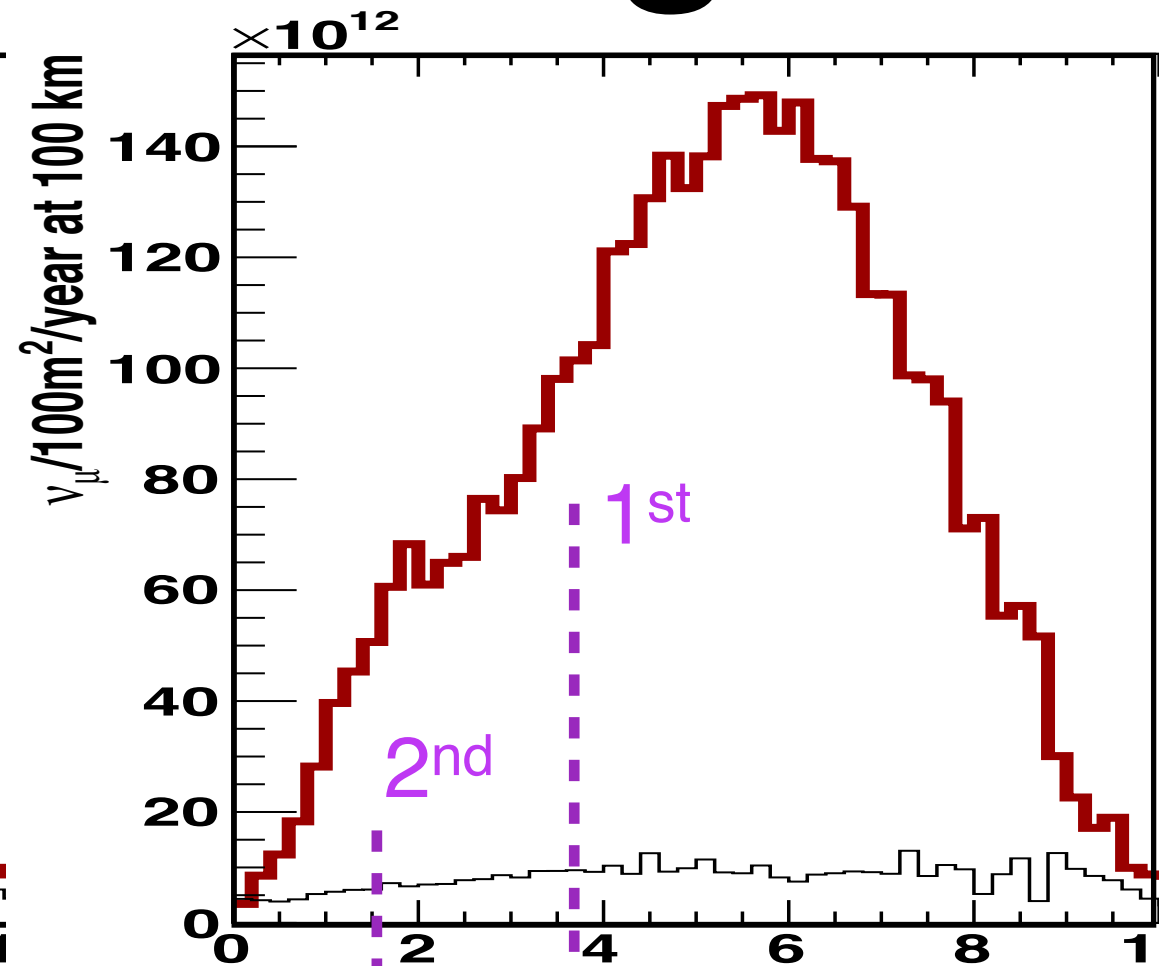
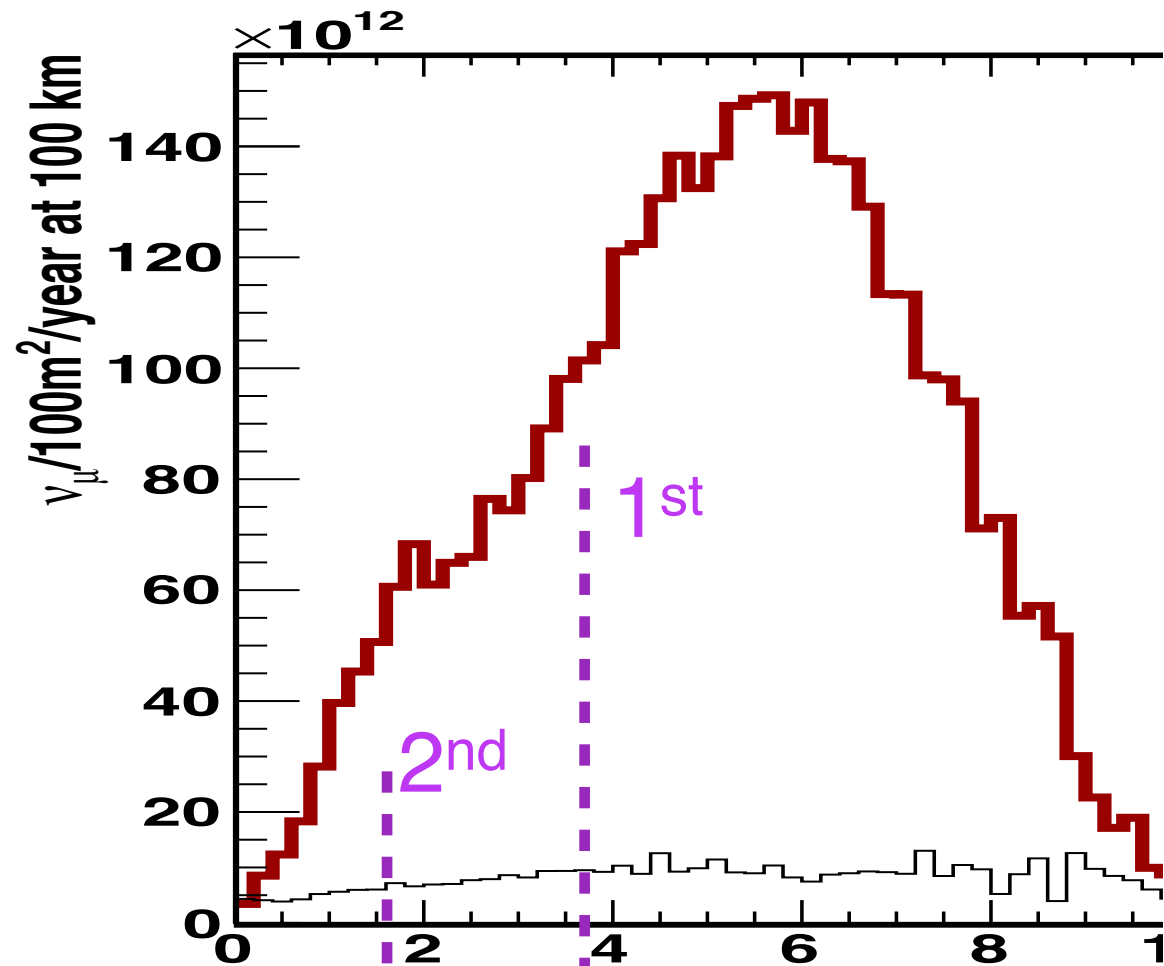
★ Inverted mass hierarchy

L=2300 km

$$\sin^2(2\theta_{13}) = 0.09$$



Flux matching



LAGUNA-LBNO neutrino beamline

- **CN2PY horn focused neutrino beam towards Pyhäsalmi**

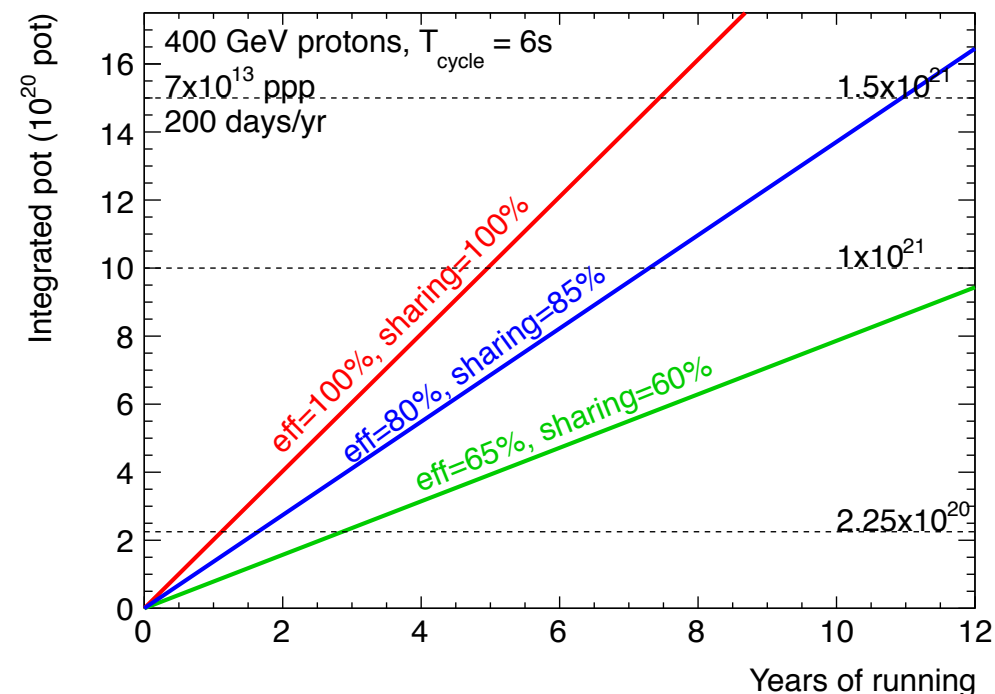
- ▶ Starting point is SPS and CNGS operation (achieved 420kW)
- ▶ Design optimized target and horn focusing systems.
- ▶ Afford relatively short decay tunnel $\approx 300\text{m}$, but 10deg dip angle
- ▶ Near detector station to achieve target systematic errors
- ▶ Consider dedicated set of hadron-production measurements

- **Benefit from improved performance of SPS+injectors for LHC-HL; consider further options to upgrade power of SPS:**

- ▶ SPS intensity is upgraded to 7×10^{13} ppp @ 400 GeV (6 s cycle).
- ▶ Yearly integrated pot = **$(0.8-1.3) \times 10^{20}$ pot / yr**
- ▶ Total integrated (12 years) = **$(1-1.5) \times 10^{21}$ pot**
- ▶ Range corresponds to sharing 60–85%
- ▶ Studies ongoing within CERN acc. team

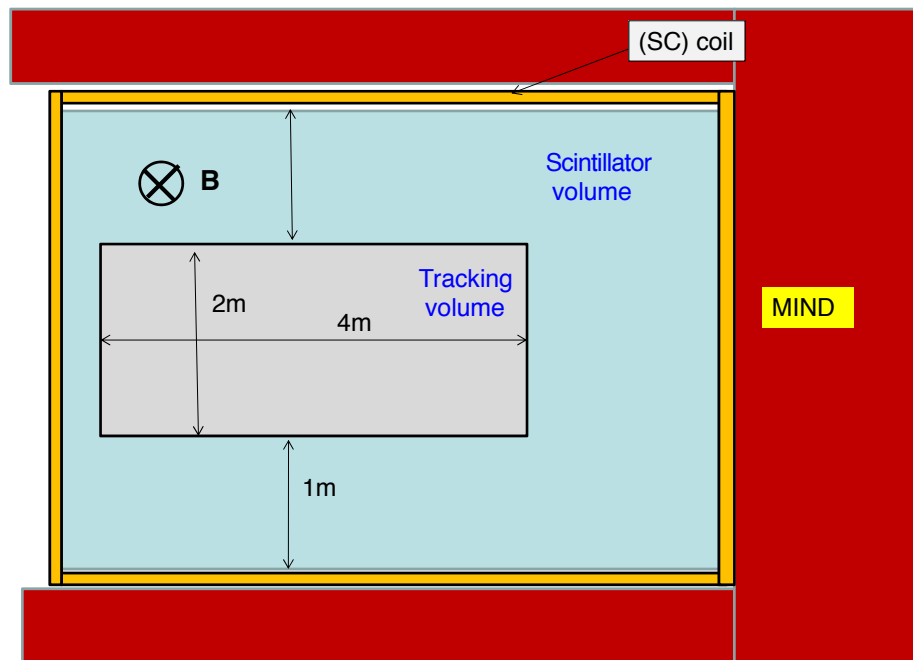
- **Upgrade path (three options):**

- **SPS upgrades (800 GeV) \rightarrow 2 MW**
- **New HP-PS accelerator (50 GeV) \rightarrow 2 MW**
- **NF storage ring**

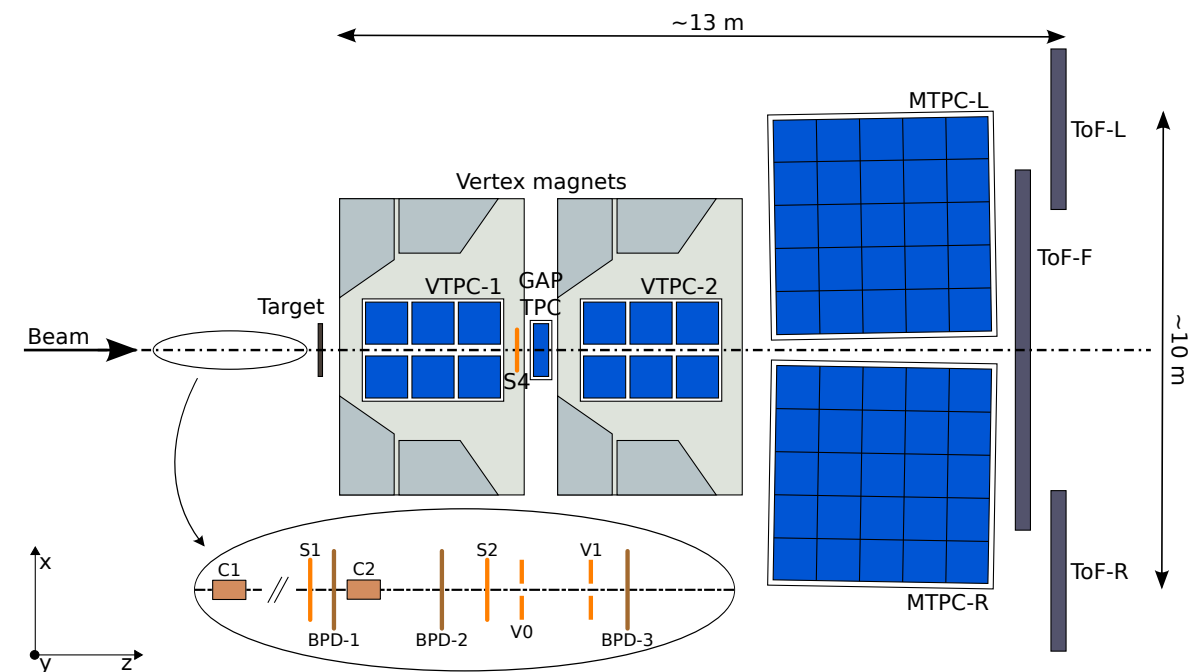


Near detector and hadro-production

- Aim: systematic errors for signal and backgrounds in the far detectors below $\pm 5\%$, possibly at the level of $\pm 2\%$ \Rightarrow control of fluxes, cross-sections, efficiencies,...



- Concept: 10 bar gas argon-mixture TPC surrounded by scintillator bar tracker embedded in an instrumented magnet with field 0.5T
- 270 kg argon mass, of which ≈ 100 kg fiducial
- 0.2 event/spill @ 700 kW
- $O(100'000)$ events/year

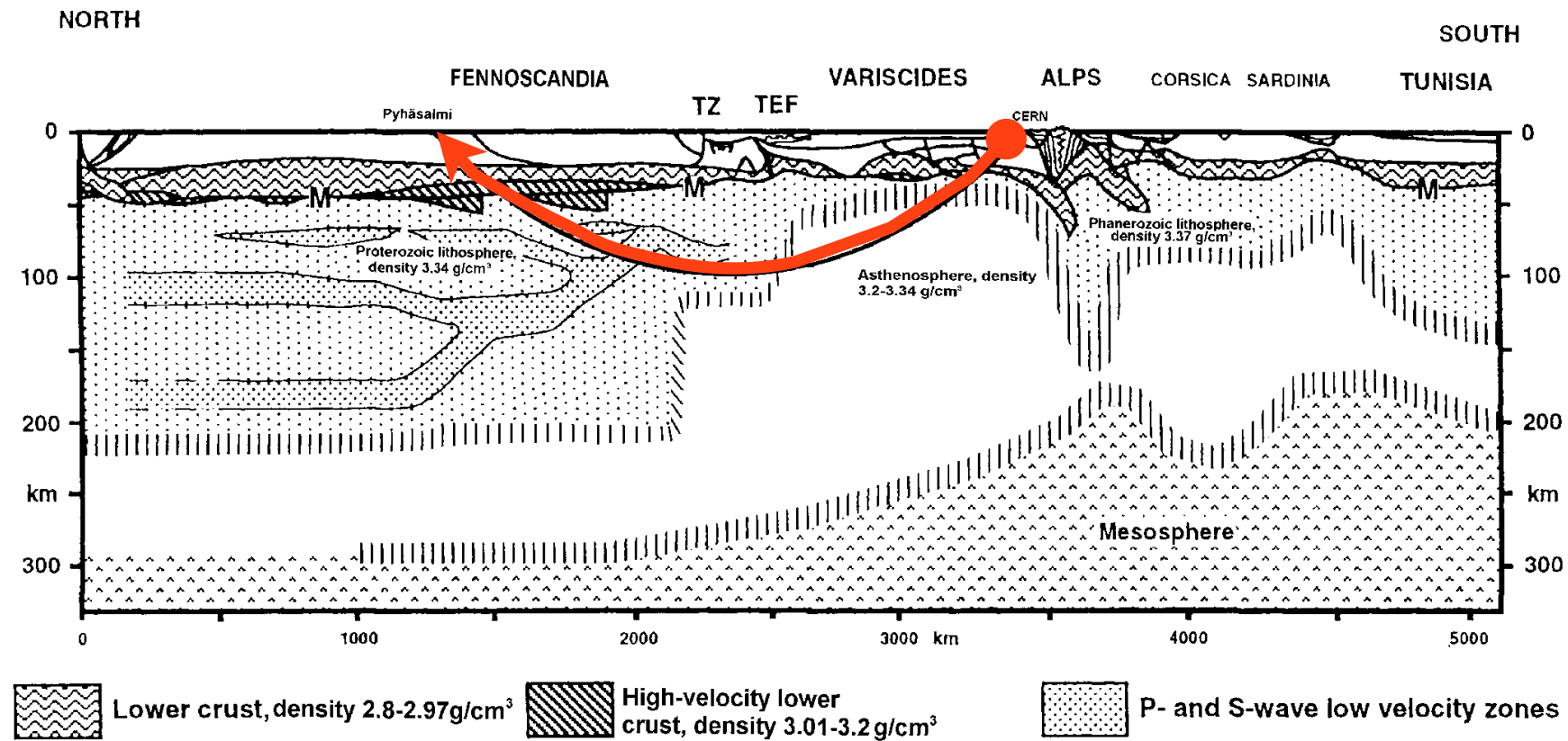
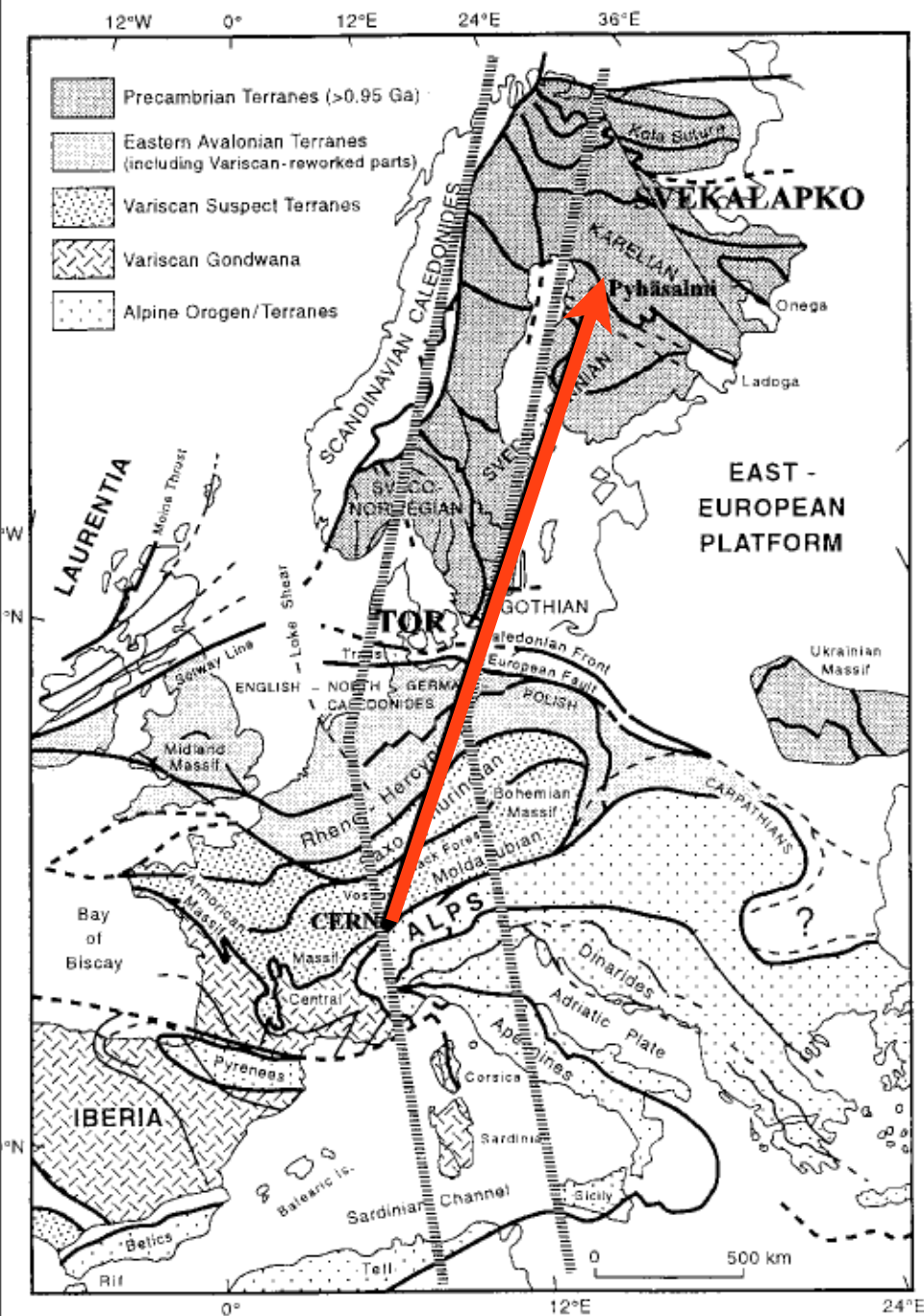


- It is widely recognized that hadro-production measurements with thin or replica target are really crucial for precision neutrino experiments (eg. K2K, T2K, MINOS).
- CERN NA61 acceptance study for 400 GeV incident protons

- Precision neutrino cross-section measurements: e.g. MINERVA, T2K-ND280, also nuSTORM (FNAL LoI)

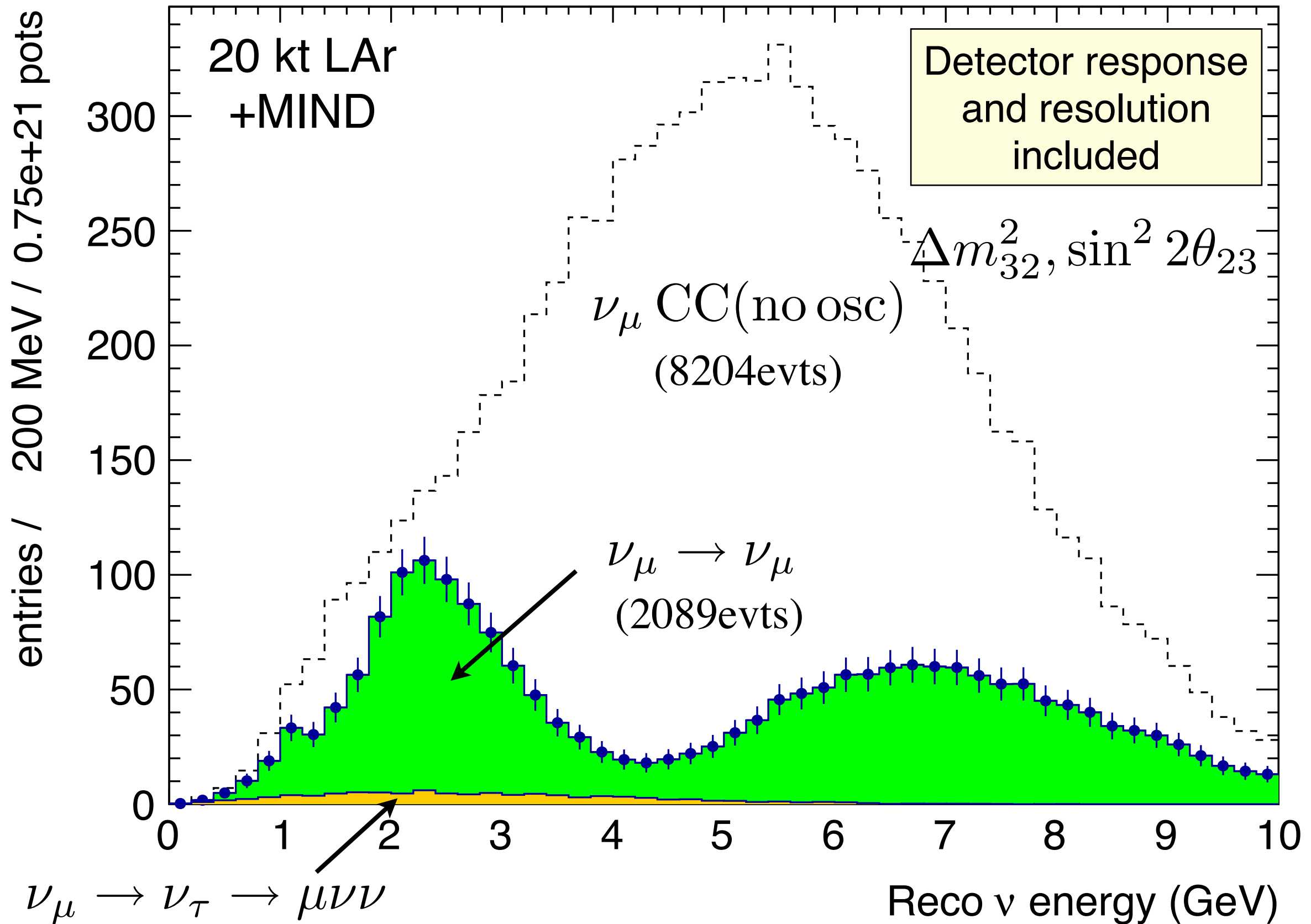
Neutrinos from CERN to Pyhäsalmi

[arXiv:hep-ph/0305042v1](https://arxiv.org/abs/hep-ph/0305042v1)

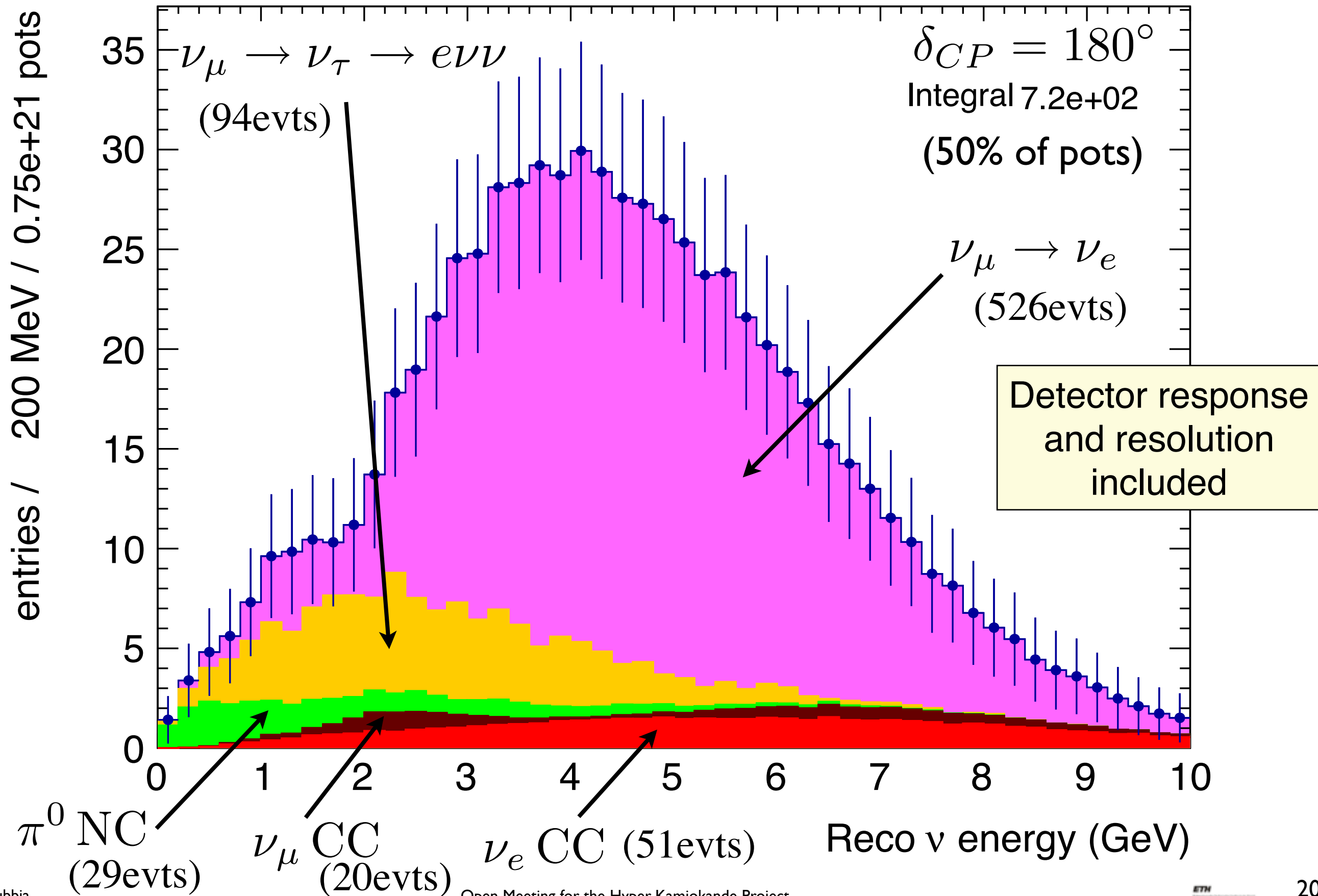


- Distance CERN-Pyhäsalmi = 2288 km
- Deepest point = 103.8 km
- Abundant geophysical data about crust and upper mantle available
- Densities = 2.4 ÷ 3.4 g/cm³
- Remaining uncertainty has small effect on neutrino oscillations (assumed equivalent to ±4% global change in matter density)

μ -like CC sample (+)



e-like CC sample (+)

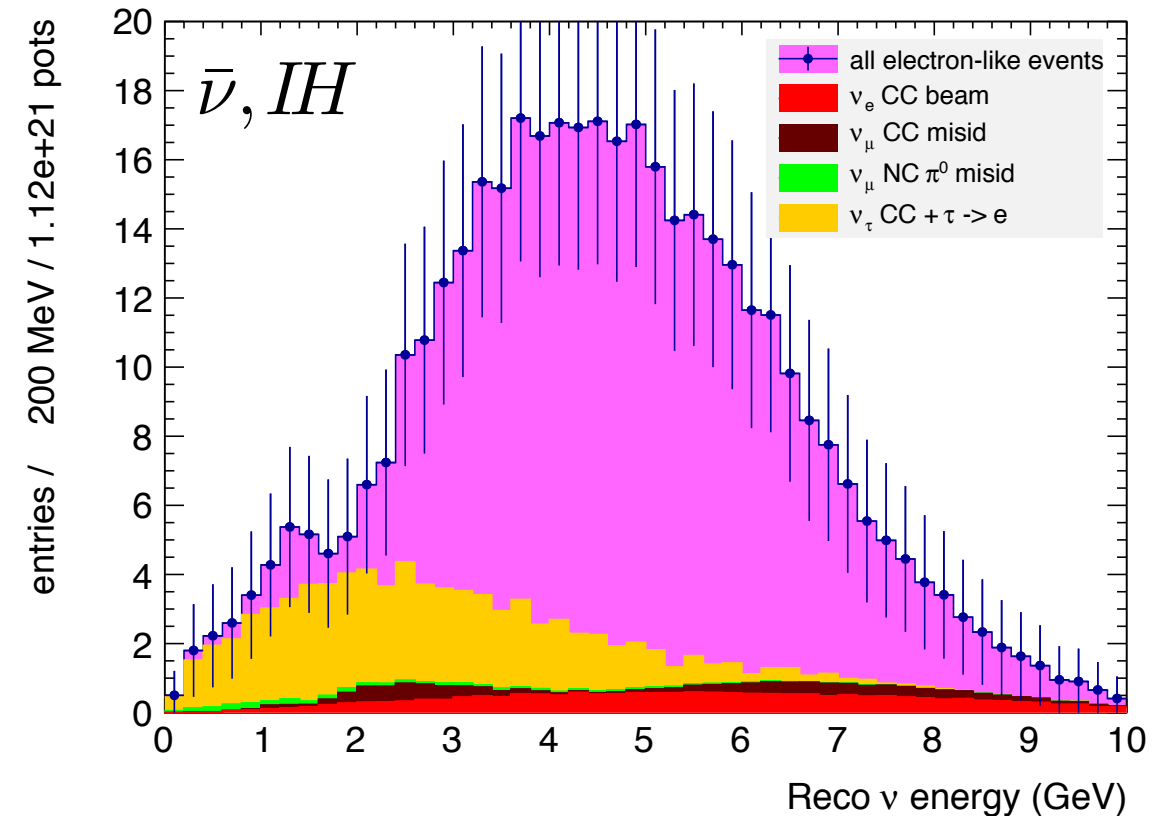
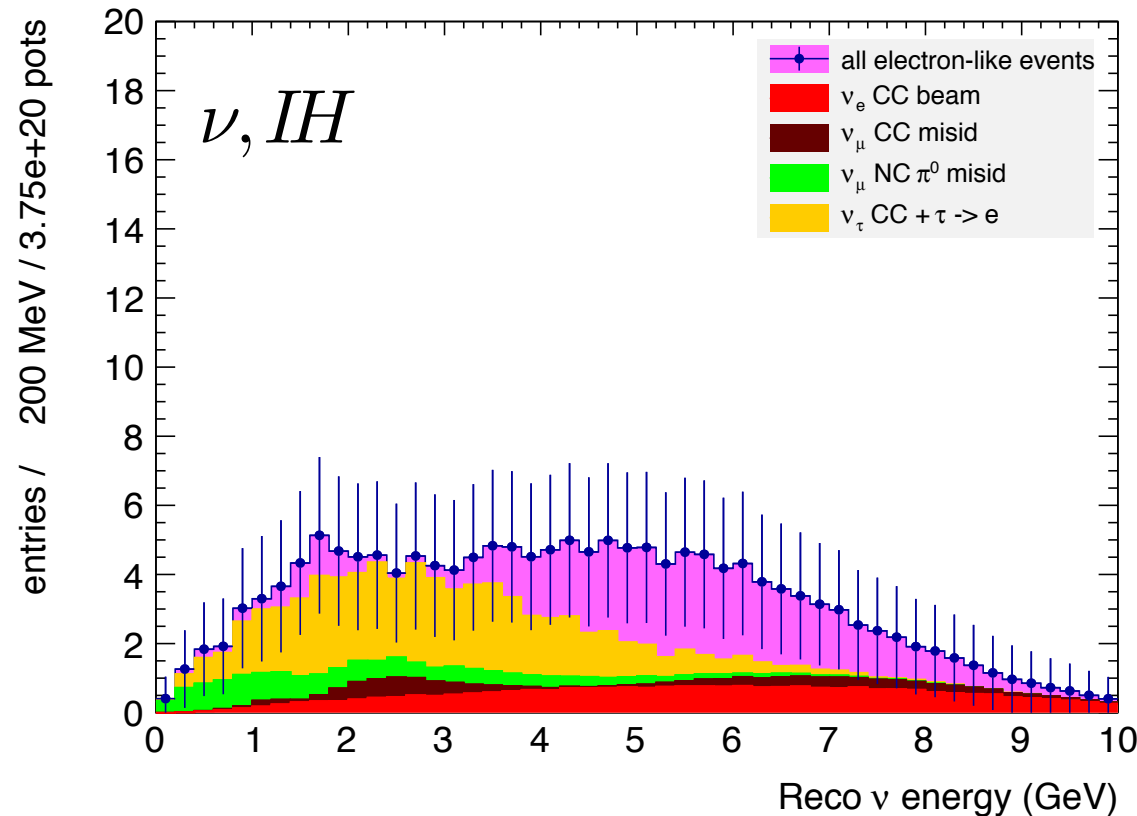
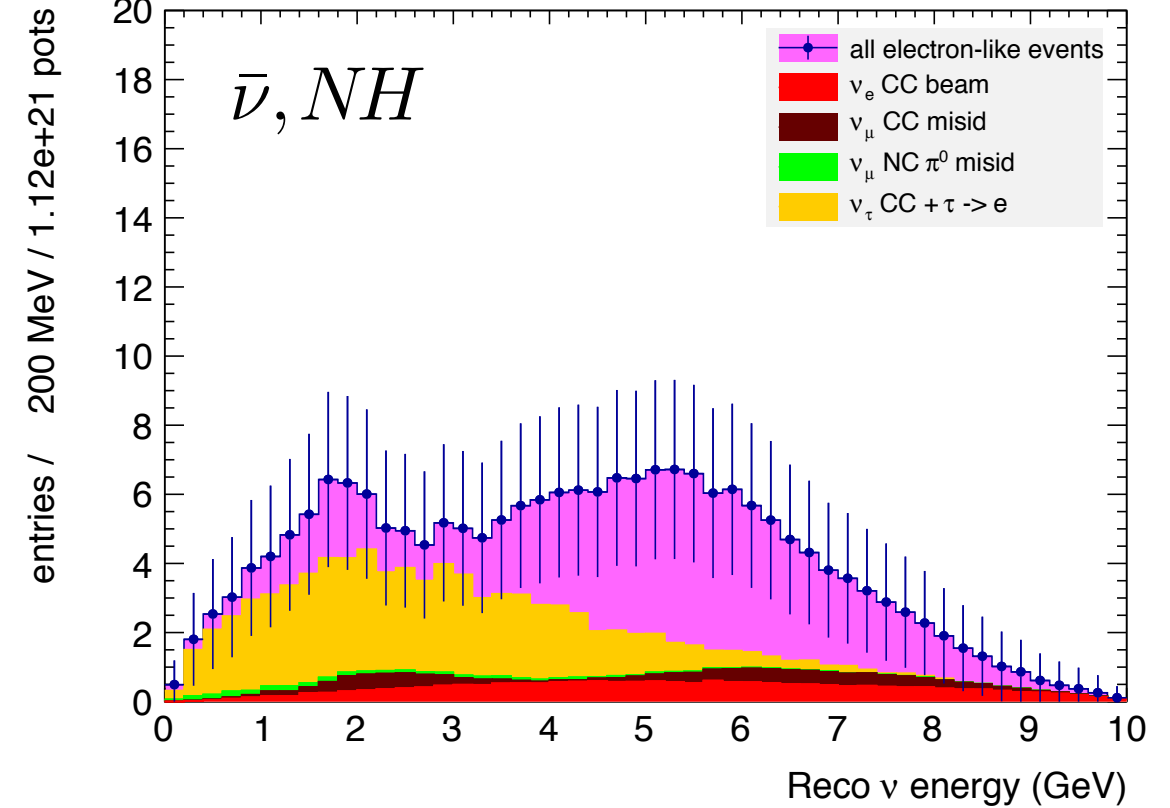
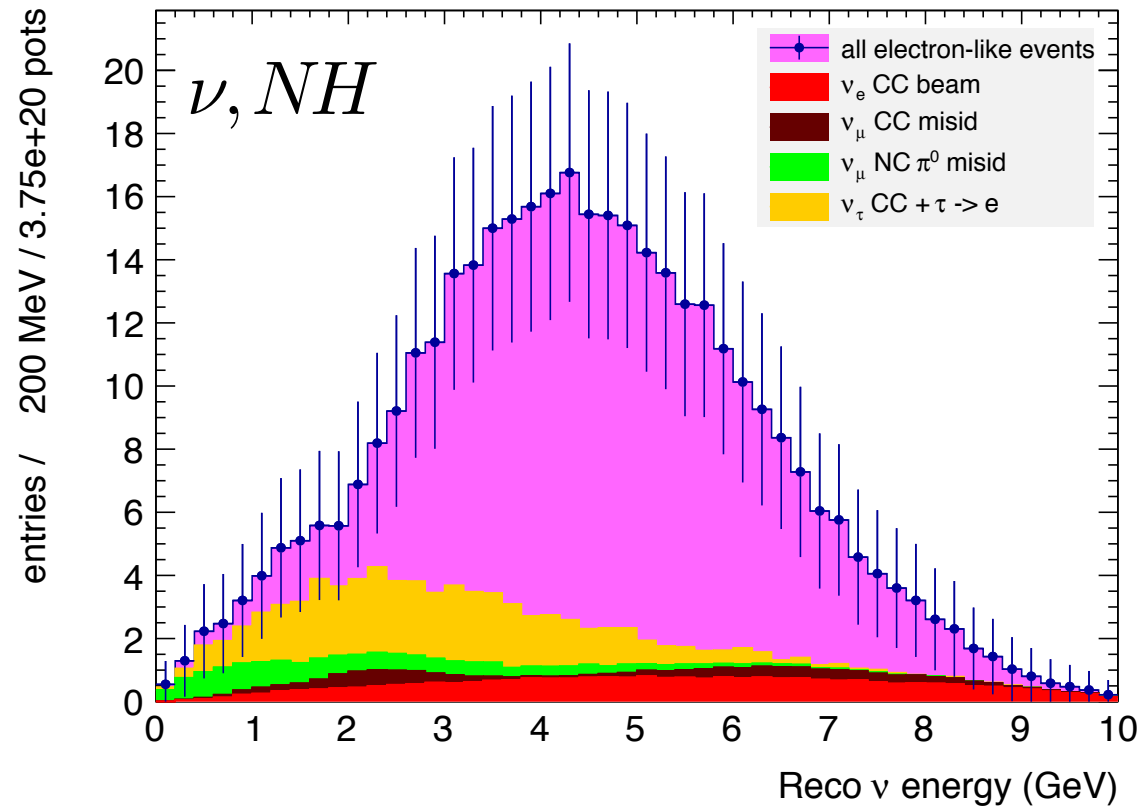


Neutrino/antineutrinos and MH

Detector response and resolution included

Running mode:

$\nu/\text{anti-}\nu$:25%/75%



LBNO sensitivity for MH&CPV

- We estimate the significance C.L. with a chi2sq method, with which we can
 - 1) exclude the opposite mass hierarchy and
 - 2) exclude $\delta_{CP} = 0$ or π (CPV)
- We minimize chi2sq w.r.t to the known 3-flavor oscillations and the nuisance parameters using Gaussian constraints

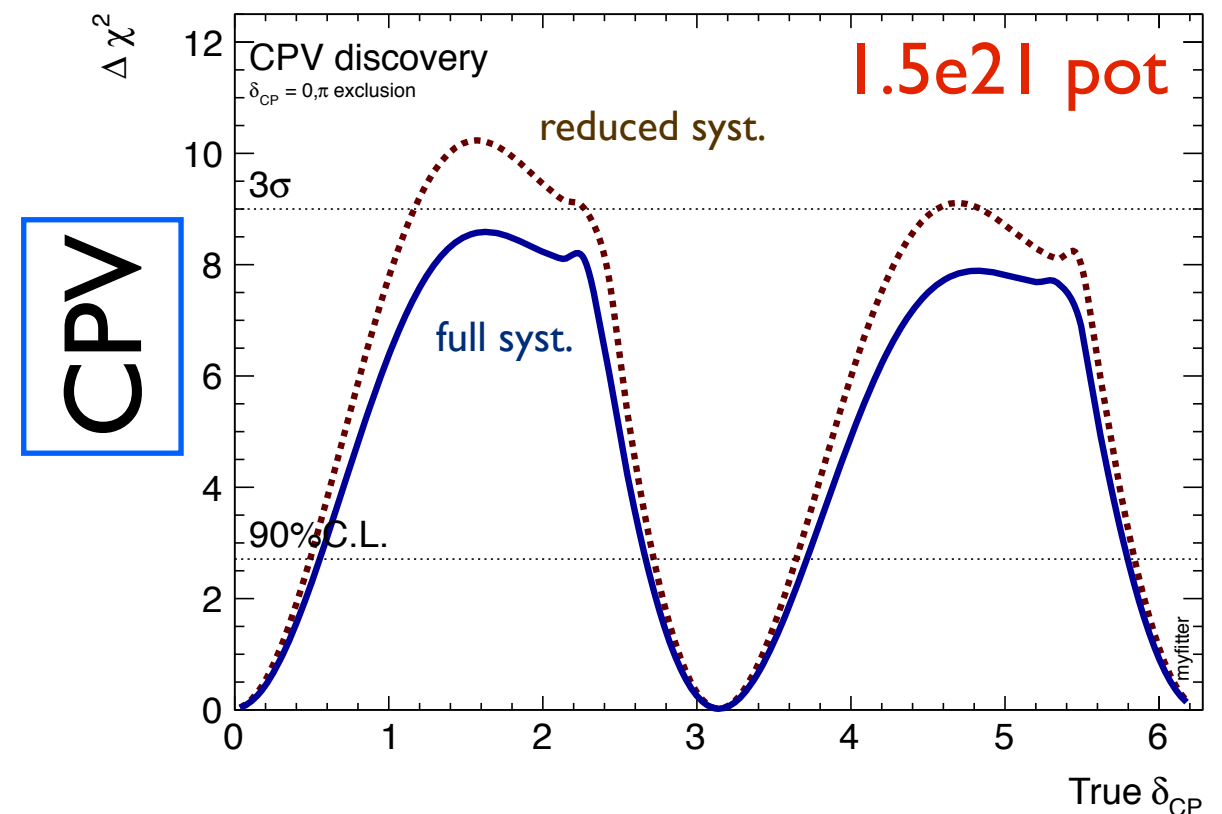
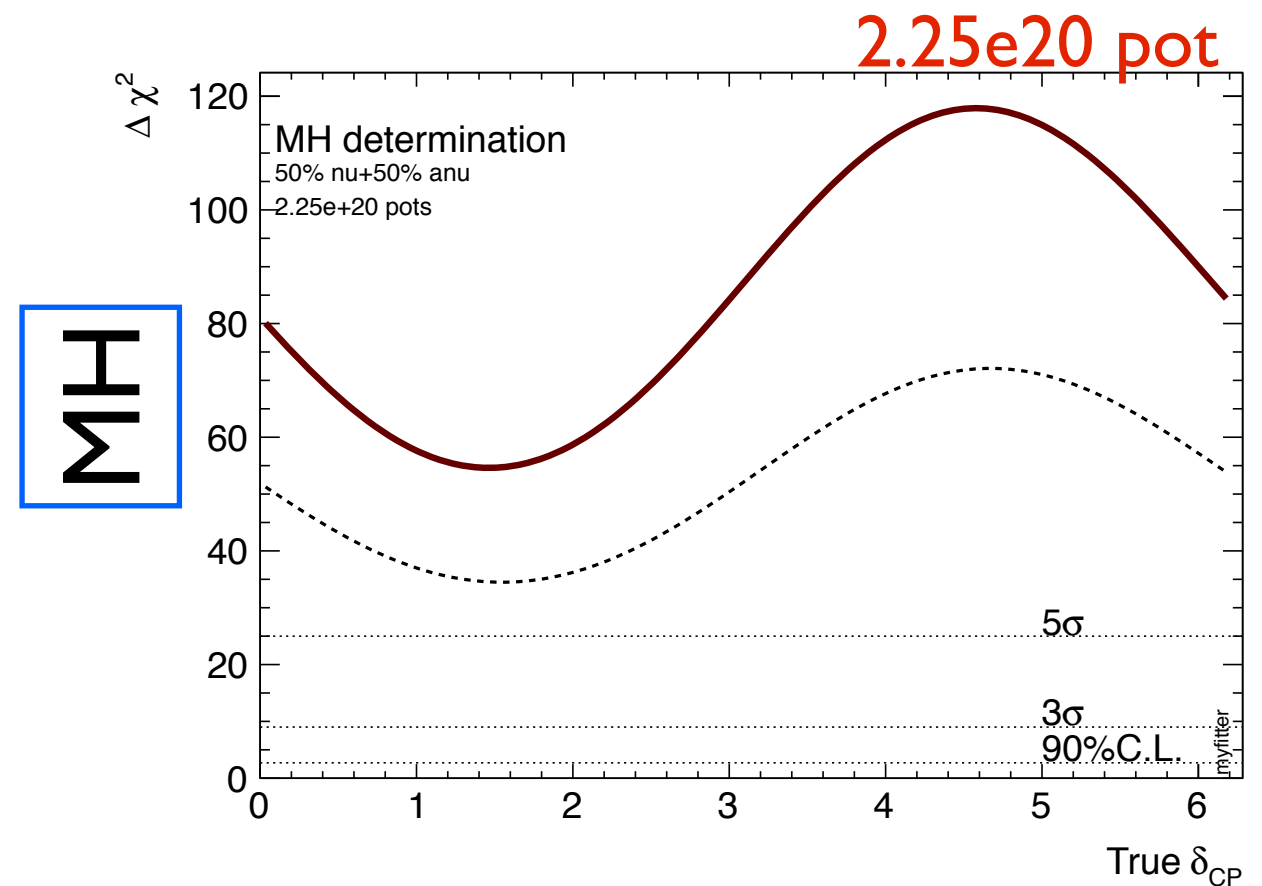
Name	Value	Error (1σ)
L	2300 km	exact
Δm_{21}^2	$7.6 \times 10^{-5} \text{ eV}^2$	exact
$ \Delta m_{32}^2 \times 10^{-3} \text{ eV}^2$	2.40	± 0.09
$\sin^2 \theta_{12}$	0.31	exact
$\sin^2 2\theta_{13}$	0.10	± 0.02
$\sin^2 \theta_{23}$	0.50	± 0.06
Average density of traversed matter (ρ)	3.2 g/cm^3	$\pm 4\%$

Name	MH determination Error (1σ)	CP determination Error (1σ)
Bin-to-bin correlated:		
Signal normalization (f_{sig})	$\pm 5\%$	$\pm 5\%$
Beam electron contamination normalization ($f_{\nu_e CC}$)	$\pm 5\%$	$\pm 5\%$
Tau normalization ($f_{\nu_\tau CC}$)	$\pm 50\%$	$\pm 20\%$
ν NC and ν_μ CC background ($f_{\nu_{NC}}$)	$\pm 10\%$	$\pm 10\%$
Relative norm. of “+” and “-” horn polarity ($f_{+/-}$)	$\pm 5\%$	$\pm 5\%$
Bin-to-bin uncorrelated	$\pm 5\%$	$\pm 5\%$

Control of
systematic
errors will be
fundamental

MH & CPV sensitivities

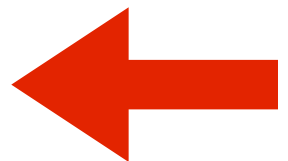
- Estimation using all systematic errors mentioned previously.
- Nominal beam power scenarios (700kW).
- For $\sin^2 2\theta_{13}=0.1$, approximately (at 90%C.L.):
 - MH: 100% coverage at $>5\sigma$ in a few years of running
 - CPV: $\approx 60\%$ coverage and evidence for maximal CP ($\pi/2, 3\pi/2$) at $\sim 3\sigma$ in 10 years
- CPV coverage already sensitive to systematic errors.
- With more details studies and a better definition of the near detector, hadron production measurements, and other auxiliary measurements, they might be reduced.
- In case of negative result, the CPV sensitivity can be improved with longer running periods and/or an increase in beam power and far detector mass. For instance, CPV becomes accessible at $> 3\sigma$'s C.L. for 75% of the δ_{CP} parameter space with a three-fold increase in exposure, provided that systematic errors can be controlled well below the 5% level.



Milestones - Timescale



LAGUNA Design Study funded for site studies:	2008-2011
Categorize the sites and down-select:	Sept. 2010
Start of LAGUNA-LBNO	2011
Submission of LBNO EoI to CERN	2012
Pyhäsalmi extended site investigation	2013
End of LAGUNA-LBNO DS: technical designs, layouts, liquids handling&storage, safety, ...	2014
Critical decision	2015 ?
Excavation-construction (incremental):	2016-2021 ?
Phase 1 LBL physics start:	2023 ?
Phase 2 incremental step implementation:	>2025 ?



Conclusions

- **LBNO, to be located underground at Pyhäsalmi 2300km away from CERN, has truly unique scientific opportunities. EoI submitted to CERN SPSC.**
 - ➔ all transitions ($e/\mu/\tau$) measurable in neutrino/antineutrino in a single experiment
 - ➔ a fully conclusive mass hierarchy determination, in a cleaner and more significant way than any other methods/proposals
 - ➔ a very good chance to find CPV with the spectral information providing unambiguous oscillation parameters sensitivity. With 10 years at 700kW SPS and 20 kton LAr +MIND (=initial phase), the reach is $\approx 60\%$ CPV coverage at 90% C.L. This step will inform future investigations (e.g. is CPV maximal ? systematic errors ?).
 - ➔ $> \times 10$ better sensitivity in several nucleon decay channels.
$$Br(p \rightarrow \bar{\nu} K) > 2 \times 10^{34} y (90\% C.L.) \quad Br(n \rightarrow e^- K^+) > 2 \times 10^{34} y (90\% C.L.)$$
 - ➔ detection of several astrophysical sources (SN,...) and fresh new look at atmospheric neutrinos with high granularity and resolution (atm tau app., atm MH, ...).
- **5600 atmospheric evts/yr > 10000 's evts SN@10kpc relic SN, WIMP?, ...**
- **LBNO defines a clear upgrade path** (long term vision / incremental approach) to fully explore CPV. E.g., a three-fold exposure yields 75% CPV coverage at 3σ C.L. LBNO is a possible first step towards the Neutrino Factory.
- **We have called on CERN to engage in a collaborative effort with the LBNO Collaboration to prepare a full engineering design of the CN2PY beam and to promptly support the necessary detector prototyping and test beams needed to develop a Proposal by the end of 2014.**

Acknowledgements

- FP7 Research Infrastructure “Design Studies” LAGUNA (Grant Agreement No. 212343 FP7-INFRA-2007-1) and LAGUNA-LBNO (Grant Agreement No. 284518 FP7-INFRA-2011-1)
- We are grateful to the CERN Management for supporting the LAGUNA-LBNO design study.
- We thank the CERN staff participating in LAGUNA-LBNO, in particular M.Benedikt, M.Calviani, I.Efthymiopoulos, A.Ferrari, R.Garoby, F.Gerigk, B.Goddard, A.Kosmicki, J.Osborne, Y.Papaphilippou, R.Principe, L.Rossi, E.Shaposhnikova and R.Steerenberg.
- The contributions of Anselmo Cervera are also recognized.

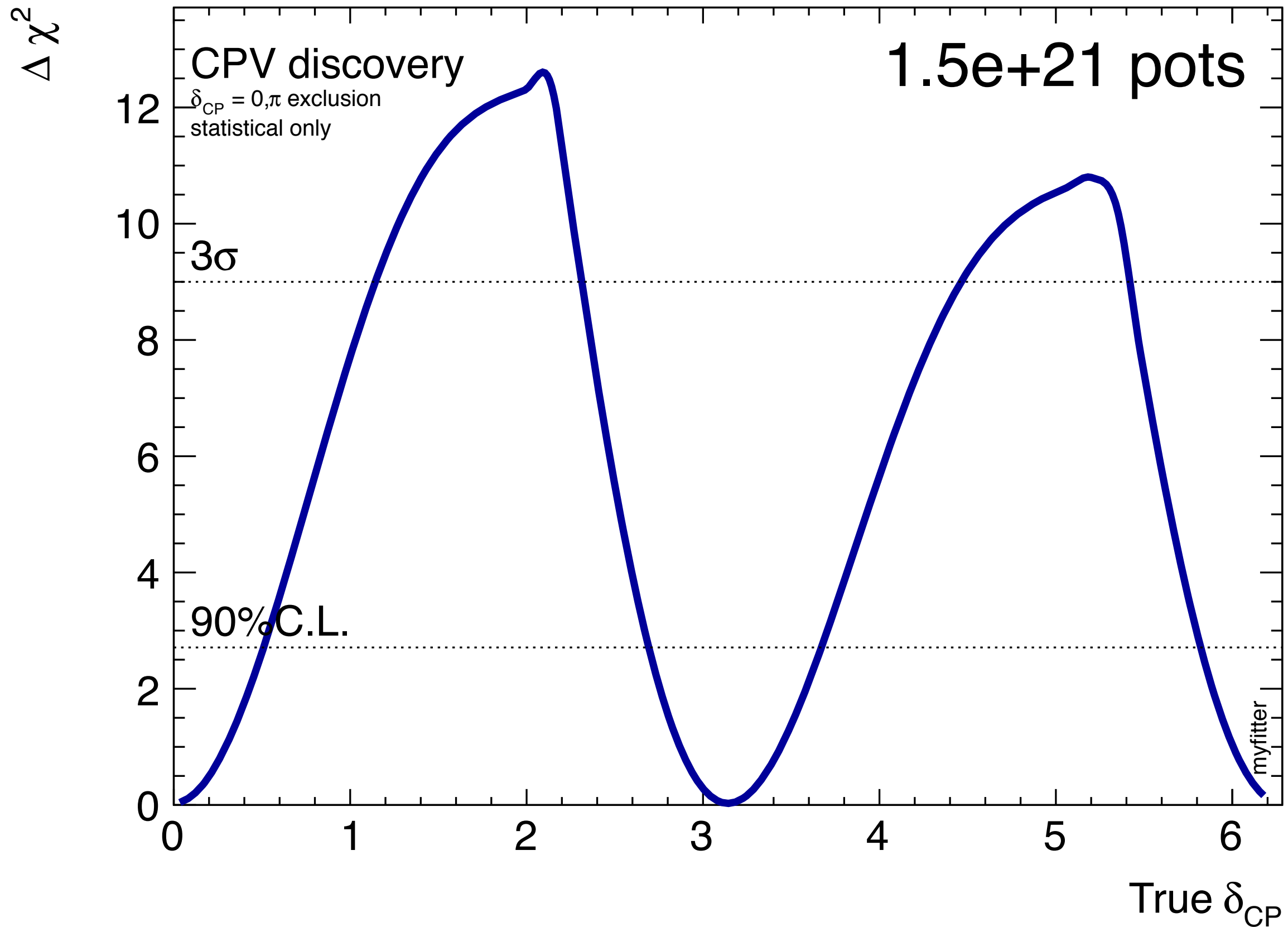


Courtesy PvZ

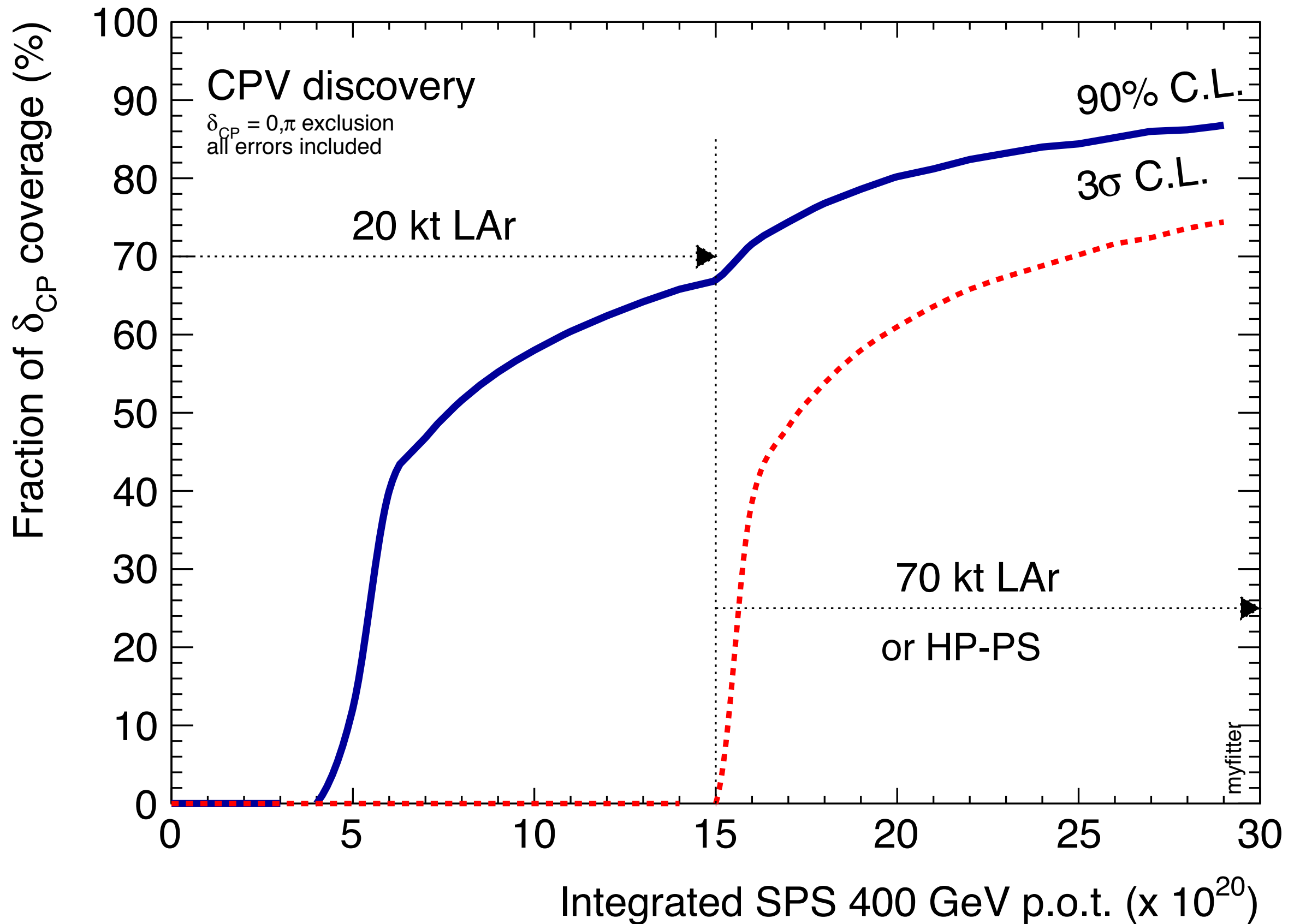
Backup slides



CPV discovery - statistical only



Incremental approach with conventional beams



CP and Matter Asymmetries

★ CP-asymmetry in vacuum:

$$\mathcal{A}_{CP}^{vac}(\delta_{CP}) \equiv abs \left(\frac{P^{vac}(\nu) - P^{vac}(\bar{\nu})}{P^{vac}(\nu) + P^{vac}(\bar{\nu})} \right)$$

★ Asymmetry due to matter effects:

$$\mathcal{A}_{CP}(\rho) \equiv abs \left(\frac{P^{mat}(\nu) - P^{mat}(\bar{\nu})}{P^{mat}(\nu) + P^{mat}(\bar{\nu})} \right)$$

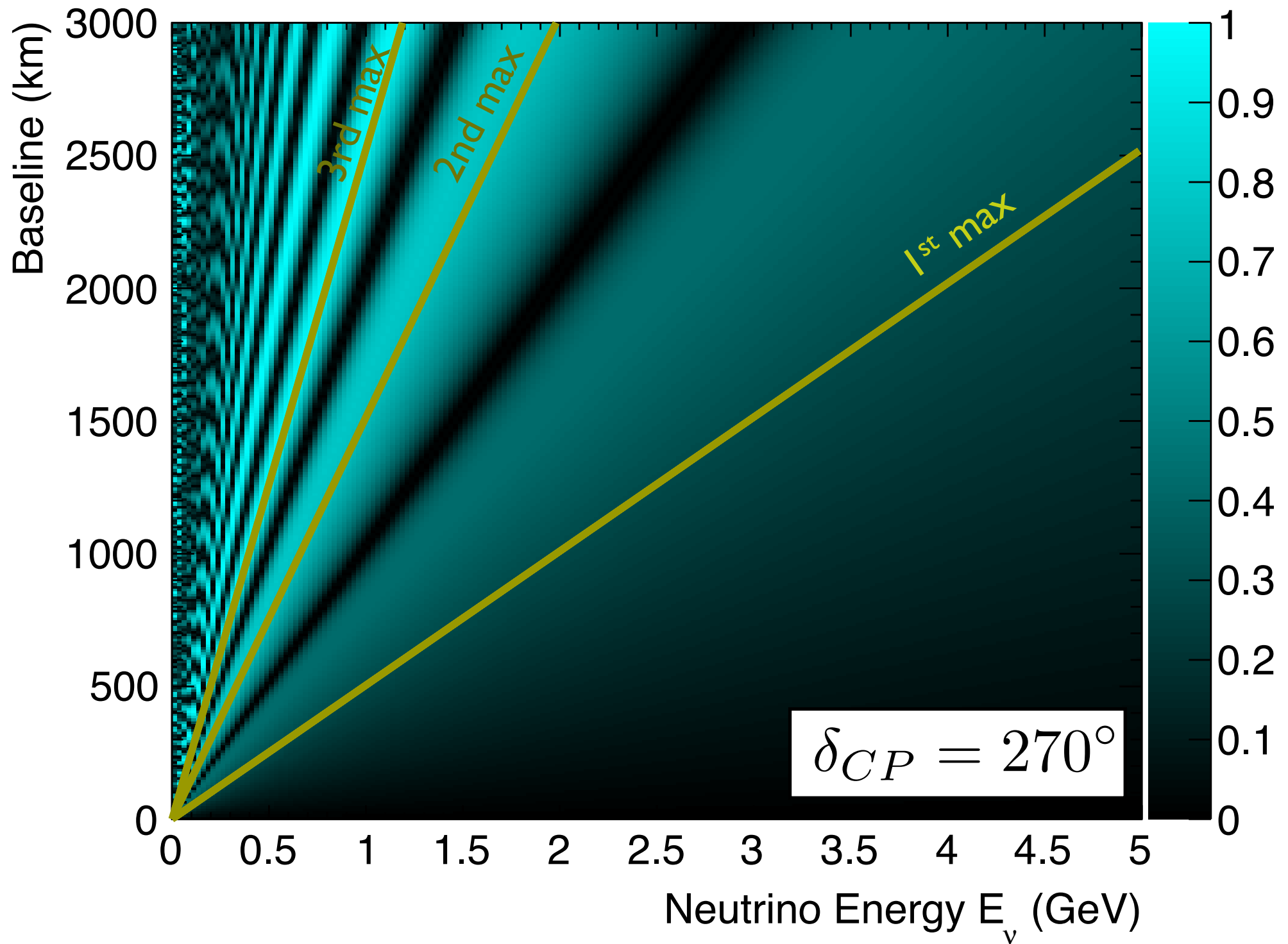
- CP asymmetries are largest at the 2nd, 3rd, ... maxima.
- Matter asymmetry dominates around the 1st maximum.
- Long(er) baselines, wide-band beams to cover several maxima are needed to resolve degeneracies.
- Experimentally: $E_{\nu}^{2nd\ max} \gtrsim 0.5\ \text{GeV} \implies L \gtrsim 1000\ \text{km}$
(fluxes, cross-sections, ...)

CP-violation determination (CPV)

$$\mathcal{A}_{CP}^{vac}(\delta_{CP}) \equiv \text{abs} \left(\frac{P^{vac}(\nu) - P^{vac}(\bar{\nu})}{P^{vac}(\nu) + P^{vac}(\bar{\nu})} \right)$$

$\sin^2(2\theta_{13}) = 0.09$

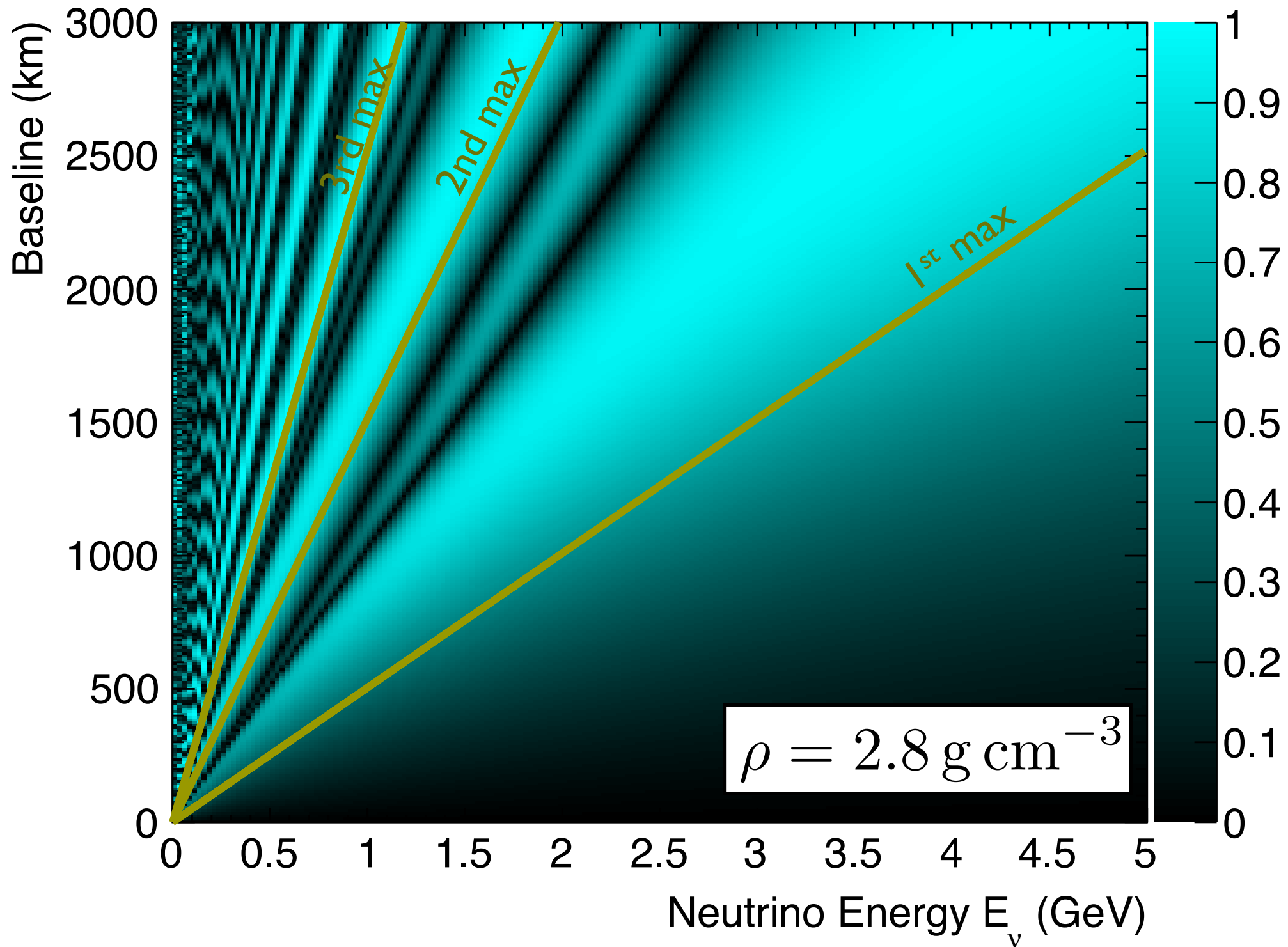
Normal mass hierarchy



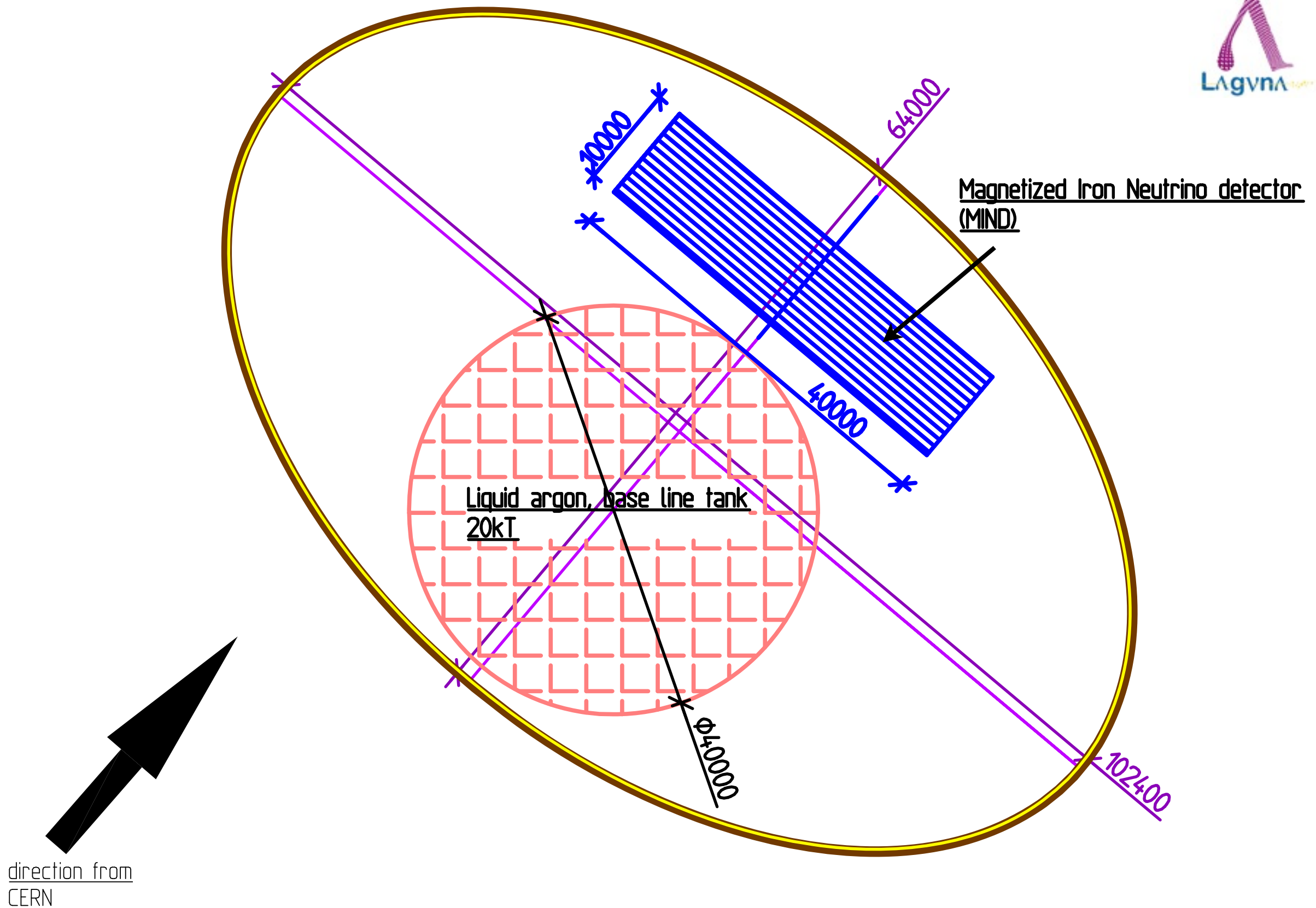
Mass hierarchy determination (MH)

$$\mathcal{A}_{CP}(\rho) \equiv \text{abs} \left(\frac{P^{mat}(\nu) - P^{mat}(\bar{\nu})}{P^{mat}(\nu) + P^{mat}(\bar{\nu})} \right)$$

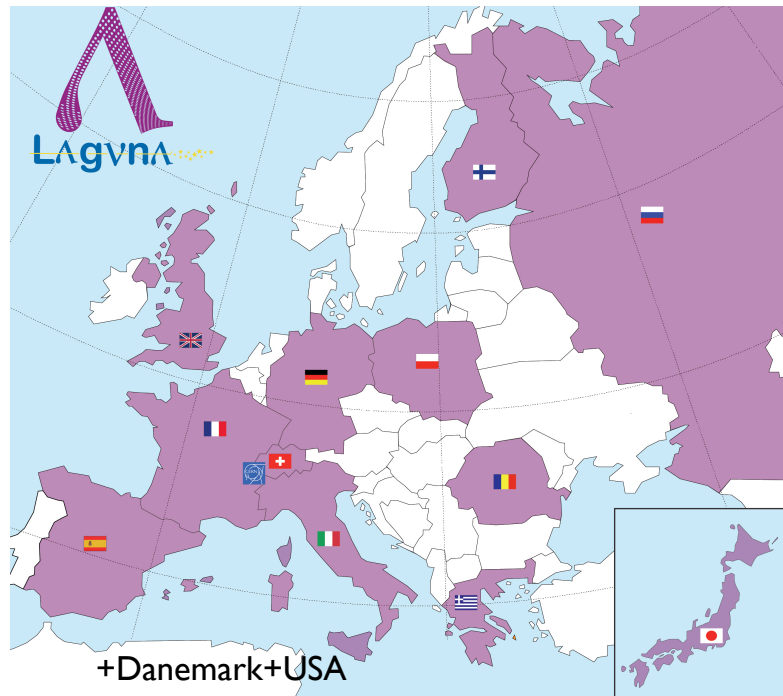
$\sin^2(2\theta_{13}) = 0.09$
Normal mass hierarchy



Top view of far detector cavern



LAGUNA-LBNO consortium



**14 countries, 47 institutions,
~300 members**

France

CEA
CNRS-IN2P3
Sofregaz*

Spain

LSC
UA Madrid
CSIC/IFIC
ACCIONA*

Romania

IFIN-HH
University Bucharest

Denmark

Aarhus

Italy

AGT*

Russia

INR
PNPI

Japan

KEK

USA

Virginia Tech

Germany

TU Munich
University Hamburg
Max-Planck-Gesellschaft
Aachen
University Tübingen

United Kingdom

Imperial College London
Durham
Oxford
QMUL
Liverpool
Sheffield
Sussex
RAL
Warwick
Technodyne Ltd*
Alan Auld Ltd*
Ryhal Engineering*

Poland

IFJ PAN
IPJ
University Silesia
Wroclaw UT
KGHM CUPRUM*

Greece

Demokritos

Switzerland

University Bern
University Geneva
ETH Zürich (coordinator)
Lombardi Engineering*

Finland

University Jyväskylä
University Helsinki
University Oulu
Rockplan Oy Ltd*

CERN

(* = industrial partners)



Outcome of LAGUNA DS

- All seven sites visited and studied in much detail
- Required caverns are technologically feasible
 - The required caverns can be constructed at a reasonable cost and in relatively short time
 - Safety requirements can be fulfilled at all stages of the project (excavation, instrumentations, operation)
 - There were no show-stoppers encountered
- Cavern construction is not the dominant cost
(10 – 20 %) → full excavation? yes but...
- Physics factors should determine site selection

Main events over the past 6 months

- Visit & evaluation of the Pyhäsalmi mine, January 2012
- Regular Technical board meetings: every two weeks
- Dedicated detector meetings in February 2012
 - WCD on the 15th in Paris
 - LSc on the 20th in Hamburg
 - LAr on the 22nd at CERN
- General meeting in Paris, 12 – 14 March 2012
- General meeting in Oxford, 2 – 4 July 2012

Next General Meeting → **CERN, 1-3 October 2012**

What is new in LAGUNA-LBNO vs. DS?

- Input from LAGUNA DS
- Focus on LBNO
- Magnetization (bimagic baseline)
- Clear milestones in the timetable
 - 2013 – European Strategy Update for Particle Physics
 - 2014 – end of LAGUNA-LBNO
 - 2018 – expected end of excavations in the Pyhäsalmi mine; depletion of the main ore will free the infrastructure to become fully dedicated to the construction and instrumentation of LAGUNA
- Prioritization
- Incremental approach

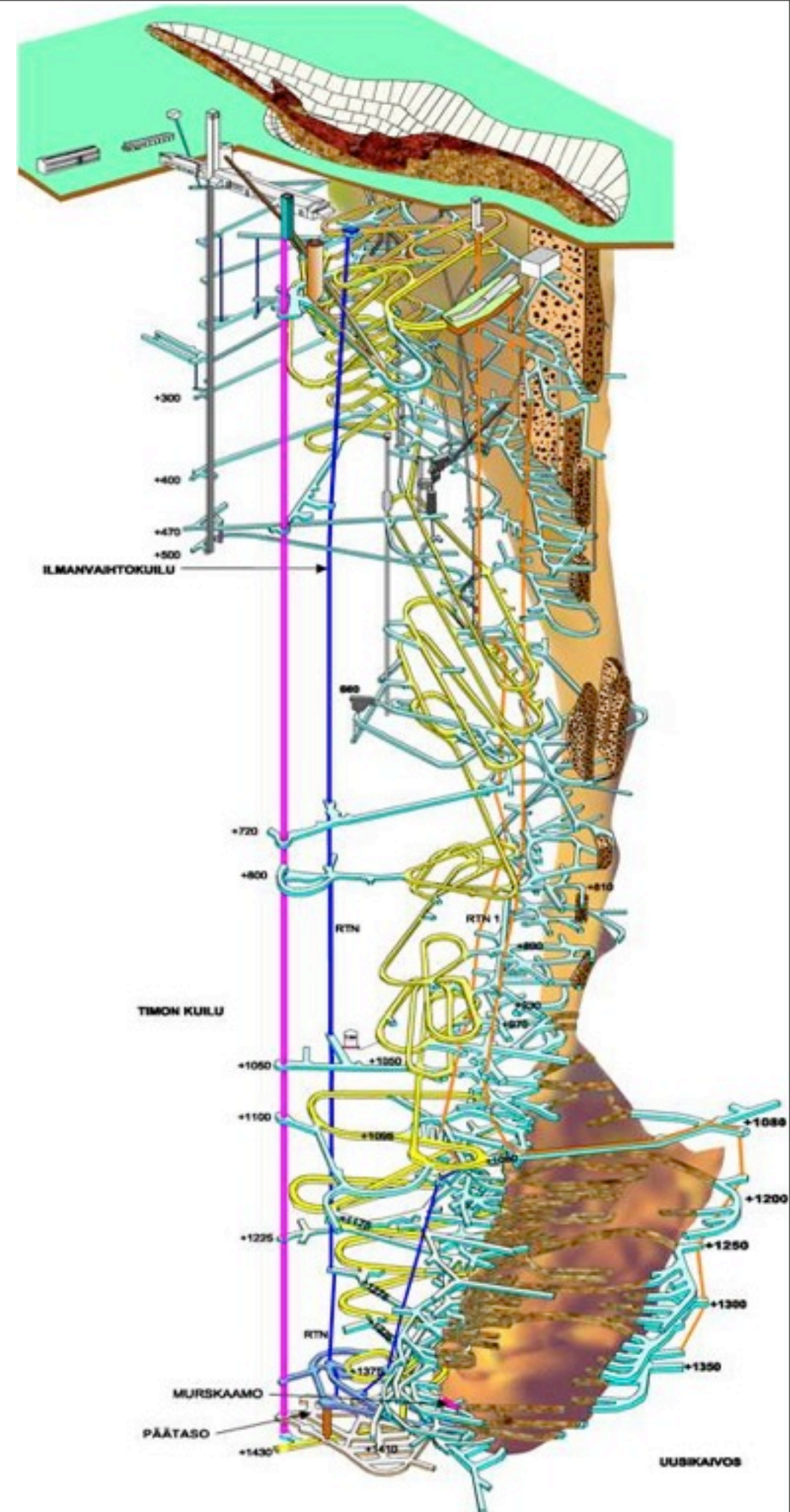
Timescale is tight !

The main developments

- Consensus among the technical partners about the suitability of the Pyhäsalmi site in Finland
- Both detectors would be located at the deepest level (1400 m)
 - Presence of excellent infrastructure at that depth
- Modified layout of the caverns was proposed
- Expression of Interest to SPSC
 - Submitted in June 2012; presented at 106th SPSC
 - Addendum by September-October 2012 to describe R&D program ? to be also submitted as input to European Strategy ?
 - Presented at the ASPERA SAC, Berlin, June 2012
- Input to the update of European Strategy for Particle Physics
 - To be ready by the end of July 2012
 - Discussed at IB yesterday – draft to be finalized (Th. Patzak)
- Funding in Finland for extensive Site Investigation

Pyhäsalmi mine

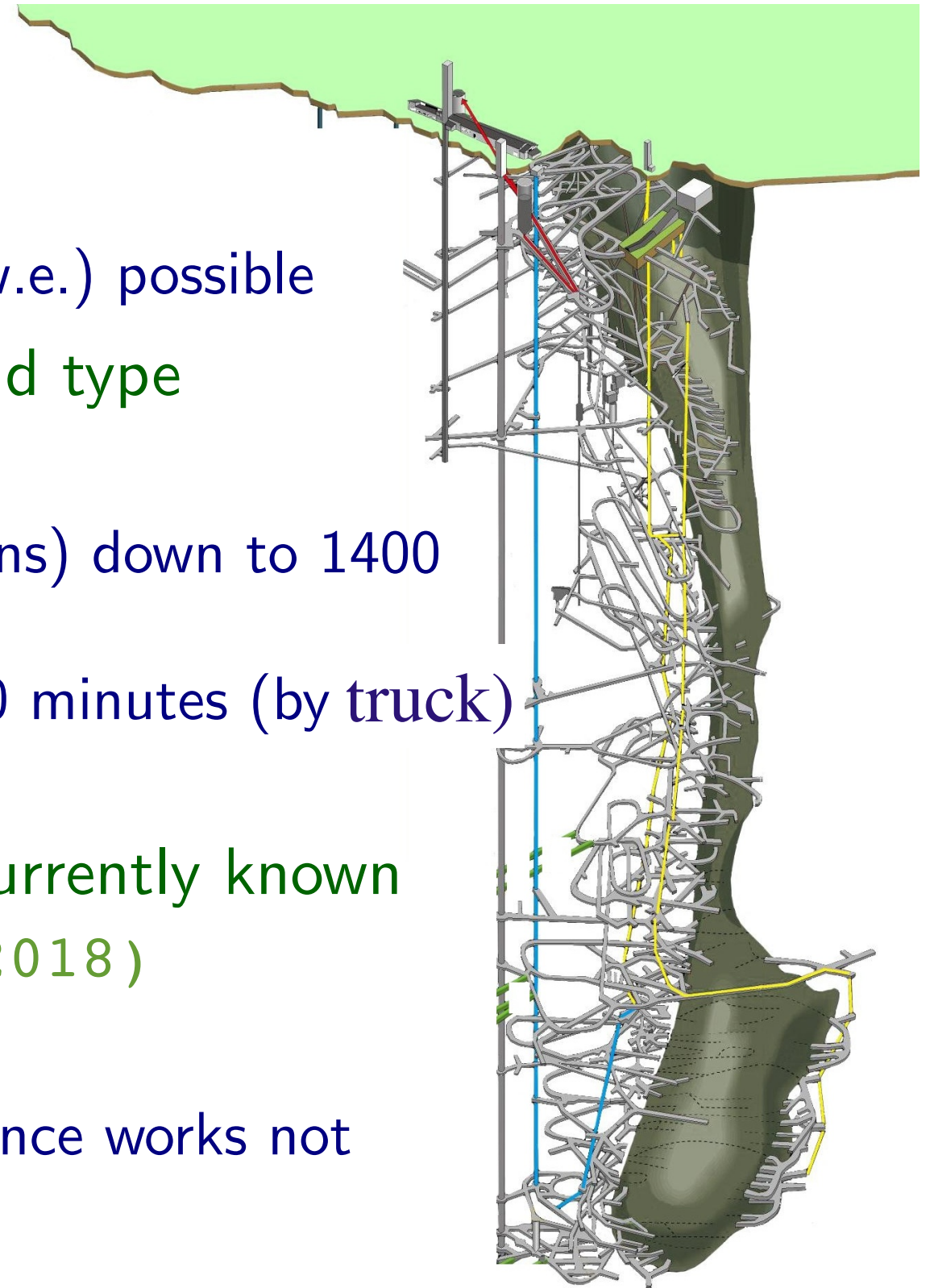
Basic facts and the
essential infrastructure



Present state of mine

Present: The Pyhäsalmi mine (Inmet Mining Ltd., Canada)

- ▶ Produces Cu, Zn, and FeS_2
- ▶ The deepest mine in Europe
 - ▶ Depths down to 1400 m (4000 m.w.e.) possible
- ▶ The most efficient mine of its size and type
- ▶ Very modern infrastructure
 - ▶ lift (of 21.5 tons of ore or 20 persons) down to 1400 metres takes ~ 3 minutes
 - ▶ via 11-km long decline it takes ~ 40 minutes (by truck)
 - ▶ good communication systems
- ▶ Operation time still 7–8 years with currently known ore reserves (presumably until 2018)
- ▶ Compact mine, small 'foot print'
 - ▶ water pumping and other maintenance works not major issues



Some unique features of Pyhäsalmi

- ▶ **Many optimal conditions satisfied simultaneously:**
 - ▶ Infrastructure in **perfect state** because of current exploitation of the mine
 - ▶ **Unique assets available** (shafts, decline, services, sufficient ventilation, water pumping station, pipes for liquids, underground repair shop...)
 - ▶ **Very little environmental water**
 - ▶ Could be **dedicated to science activities** after the mine exploitation ends (around 2018)
- ▶ **One of the deepest location in Europe (4000 m.w.e.)**
- ▶ **The distance from CERN (2300 km) offers unique long baseline opportunities.**
- ▶ **The site has the lowest reactor neutrino background in Europe, important for the observation of very low energy MeV neutrinos.**
- ▶ **Extensive site investigation with rock drilling and detailed analysis planned during the period 2012-2014 (Finnish contribution).**

Pyhäsalmi Mine main parameters and assets	
Owner	Inmet Mining Corporation, Canada
Mine composition	copper and zinc concentrate mine
Level of main ore body	between the levels 1050 and 1400 m
Underground access	guaranteed
Main infrastructure	A decline tunnel, a hoist (Timo shaft) and 3 ventilation shafts
Decline tunnel dimension	inclination, steepness: 1:7; length of tunnel: 11 km, size of tunnel: width 5 m and smallest height 3,3 m; radius of curves: min. 70,5 m.
Decline use	suitable for traffic (even heavy such as dumpers or trucks)
Present condition	good conditions, in use during the mine operations.
Maintenance	Tunnel scaled twice a year (operation takes 1 week/time with one worker)
Timo shaft Specification of hoist	size: circular with diameter 5 m, pulley diameter 4,5m manufacturer ABB, Installation by Pyhsalmi Mine passenger capacity: 20 persons maximum speed: 12 m/s (man hoisting), max. 15,5 m/s (ore hoisting) with acceleration 0,3 g Hoisting capacity: 275 ton/hour, cycle 280s Motor 2,5 MW
Ventilation shafts	1 fresh, 2 exhaust air raises
Capacity	137 m ³ /s (specs)
Temperature	23°C at –1400 m, gradient about 1.2°C/100m
Present infrastructure at the -1430m service level:	maintenance halls for equipment and material parking lots for personal vehicles parking lots for equipment and trucks electricity repair workshop equipment washing lanes (small and big) safety area / oxygen supply area telecommunications and data communication room kitchen and lunch room sauna and showers



TIMO shaft

Old shaft no
longer in use

VT (ramp) entrance

CERN, 14-16 May 2012

W.H.Trzaska

43

Tower of the (new) TIMO shaft



CERN, 14-16 May 2012

W.H.Trzaska

44

VT (ramp or decline) exit near the bottom level



VT (ramp or decline) exit near the bottom level

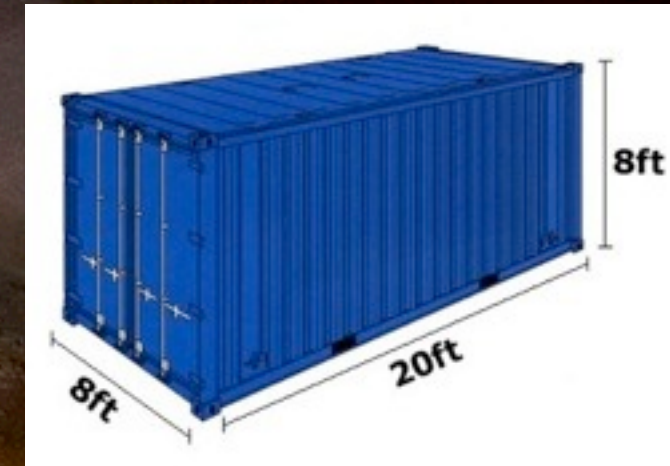


CERN, 14-16 May 2012

W.H.Trzaska

45

VT (ramp or decline) exit near the bottom level



Delivery trucks drive down to the 1400 m level on a daily basis



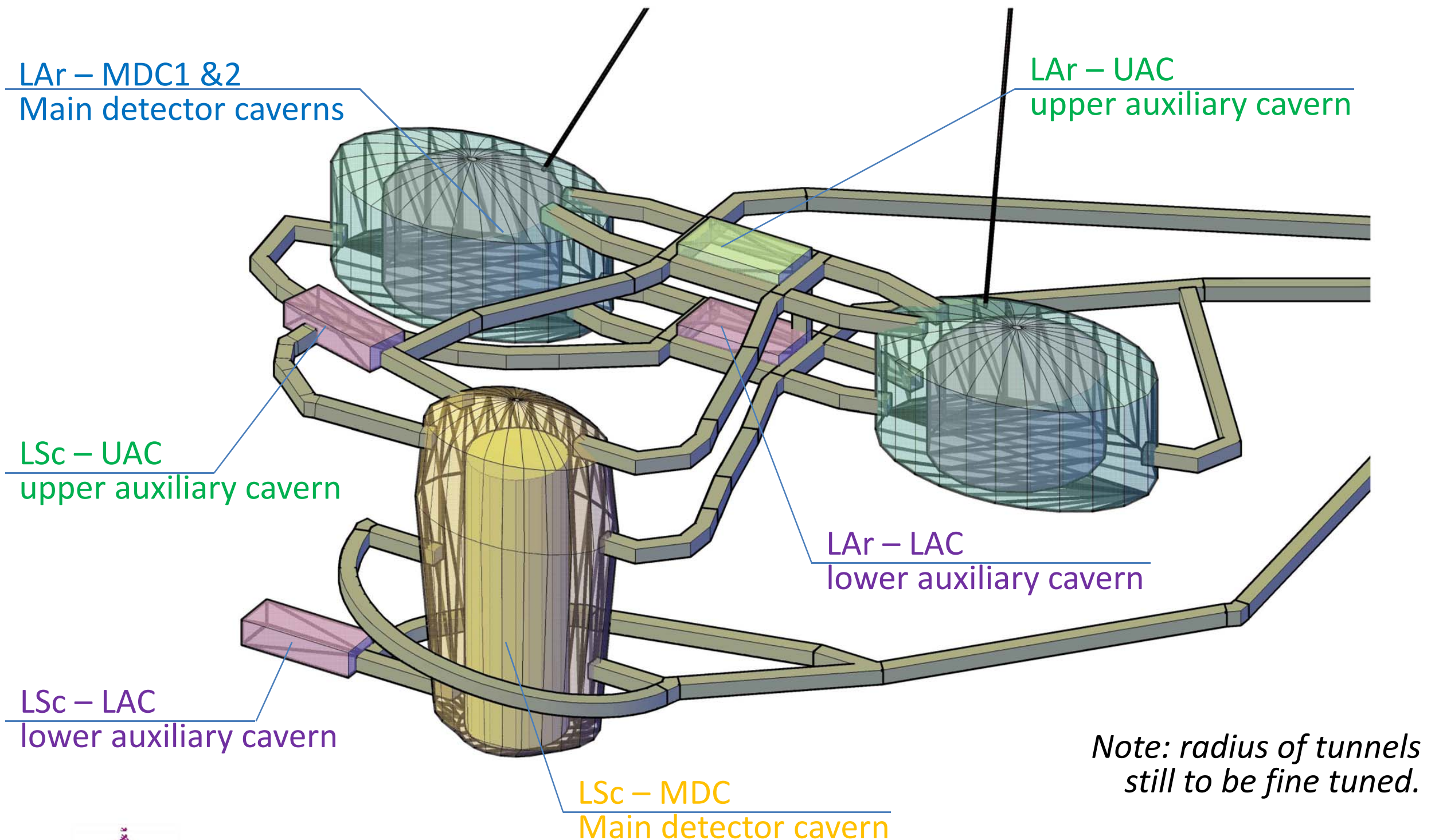
CERN, 14-16 May 2012

W.H.Trzaska

46

Main detector caverns

LAGUNA-LBNO: LAr + LSc LAYOUT @ PYHÄSALMI

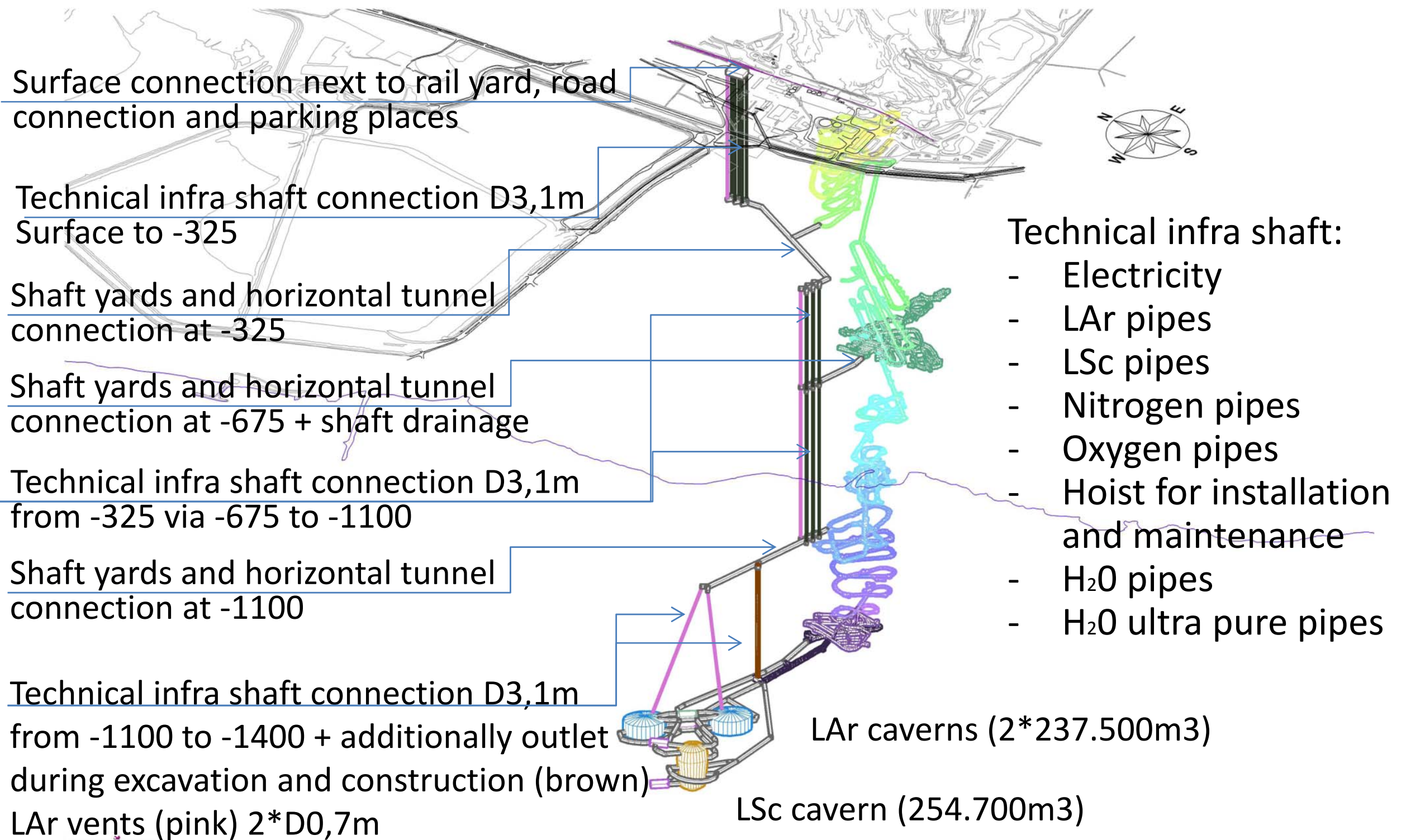


AXONOMETRIC VIEW, LAr AND LSc SOUTH - WEST
2.7.2012

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Mine infrastructure extension

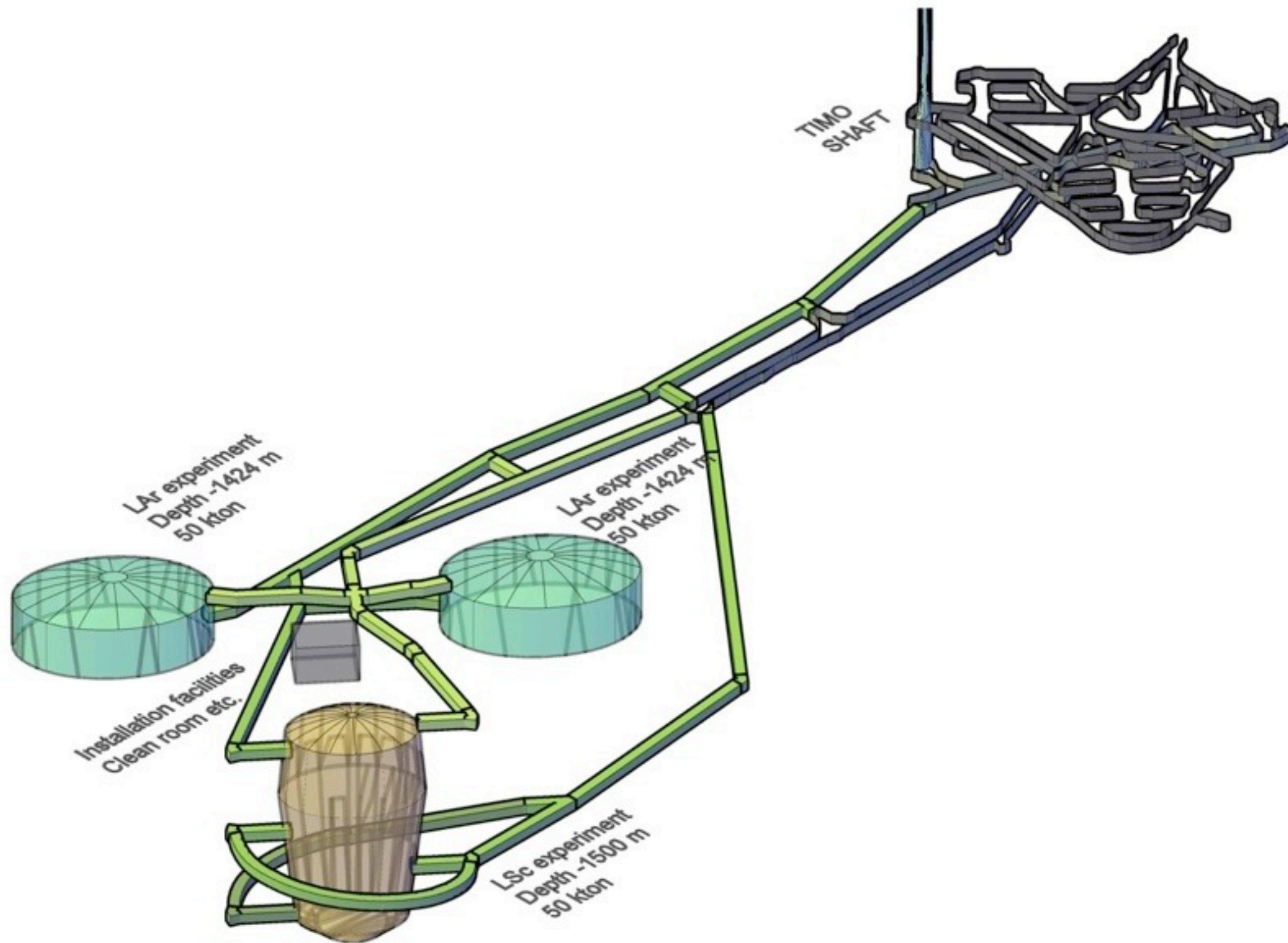


Total LAr+LSc underground infrastructure (879.000m³)

AXONOMETRIC VIEW, SOUTH - WEST

2.7.2012

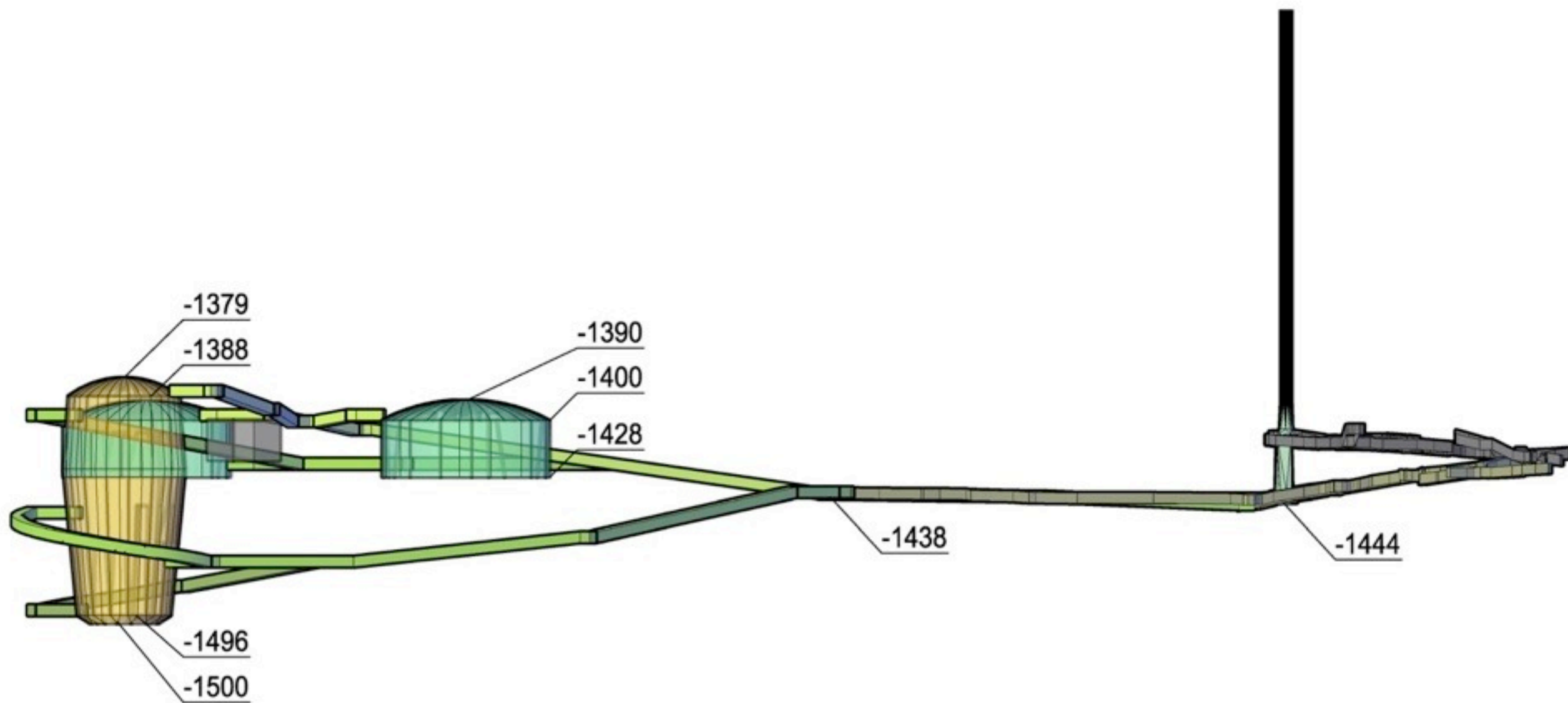
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LAGUNA LBNO

AXONOMETRIC VIEW
PHASE X
12.3.2012

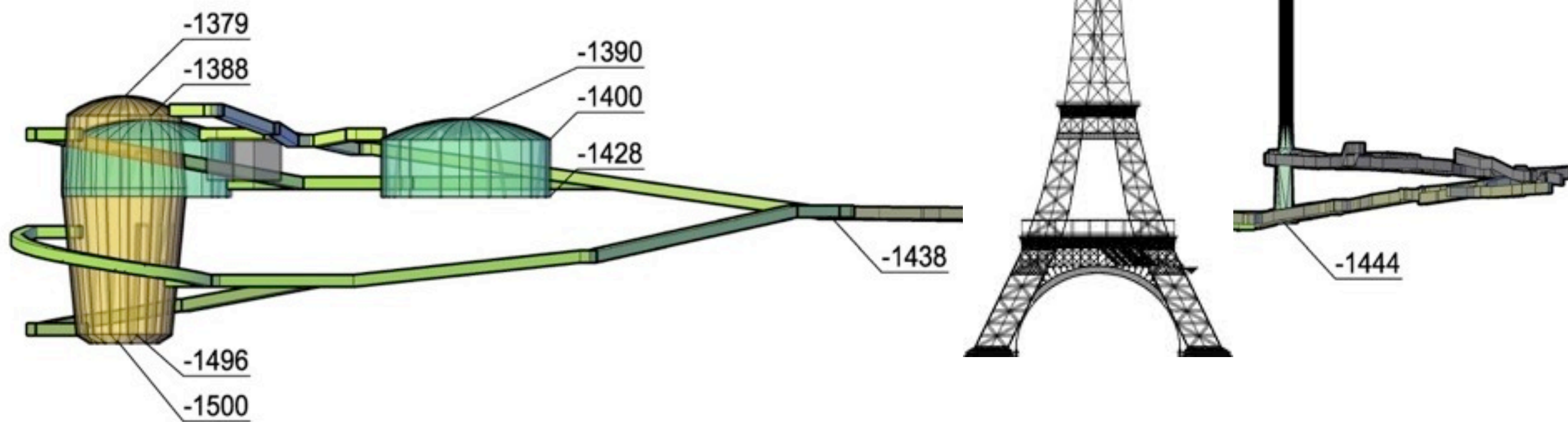
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LAGUNA LBNO

VIEW SOUTH
PHASE X
12.3.2012

 ROCKPLAN



LAGUNA LBNO

VIEW SOUTH
PHASE X
12.3.2012

 ROCKPLAN

Neutrinos at the frontier

- A wealth of new results over the last years are clarifying the landscape in Particle Physics at the various frontiers and confirm the “invincible” Standard Model (SM). The discovery of a Higgs boson at ATLAS/CMS will crown the successful SM and will call for a verification of the Higgs boson couplings to the gauge bosons and to the fermions.
- In this rapidly emerging picture, neutrino masses and oscillations are today the only experimentally established evidence of physics **Beyond the Standard Model (BSM)**.
- Very likely new BSM physics at a yet-unknown high-energy scale is a key ingredient to resolve these questions that the SM cannot answer:
 - What is the origin of the gauge structure of strong and electroweak interactions ?
 - Does a bigger gauge symmetry exist in Nature?
 - Is there a unique theory of family and flavor?
- **Being the only elementary fermions whose basic properties are still largely unknown, neutrinos must naturally be one of the main priorities in the quest to complete our knowledge of the SM.**
 - Their understanding has progressed considerably, but deeper studies are still needed to answer these profound questions.
 - The mixings among leptons have different values and are larger than those among quarks. And the smallness of the neutrino rest masses compared to those of other elementary fermions points to the preferred scenario of Majorana neutrinos and the see-saw mechanism.
 - The above observations are yet to be clarified within a unique and appropriate theoretical framework, and addressing such questions has therefore significant potential to offer new insights into the BSM physics at the very high-energy scale.
 - Is, as Bruno Pontecorvo said, the neutrino “the prototype of all other fermions” ?

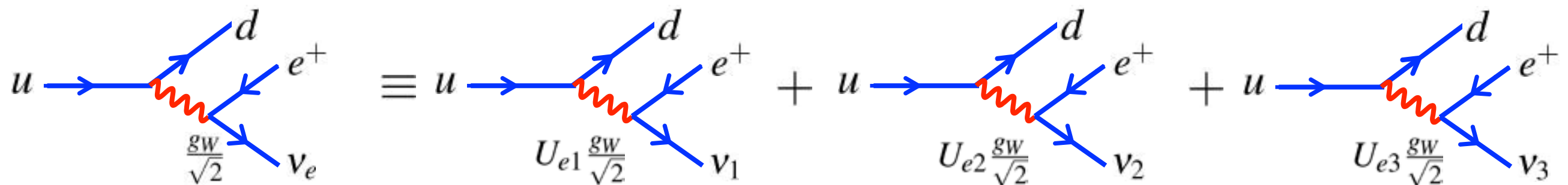
CP violation in leptonic sector

- CP-violation is an essential aspect of our understanding of the Universe and is related to the question of the matter dominance.
- A natural question is whether the SM can provide the necessary CP-violation to explain the baryon asymmetry ($\approx 10^{-9}$).
- Today we are certain that there are two places where CP-violation can enter: the **CKM matrix** and the newly found **PMNS matrix** !
- To date CP violation has **only** been observed in the quark sector

Mixing between weak and mass eigenstates:

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

Pontecorvo-Maki-Nakagawa-Sakata



T	$\nu_e \rightarrow \nu_\mu$	$\xrightarrow{\hat{T}}$	$\nu_\mu \rightarrow \nu_e$
CP	$\nu_e \rightarrow \nu_\mu$	$\xrightarrow{\hat{C}\hat{P}}$	$\bar{\nu}_e \rightarrow \bar{\nu}_\mu$
CPT	$\nu_e \rightarrow \nu_\mu$	$\xrightarrow{\hat{C}\hat{P}\hat{T}}$	$\bar{\nu}_\mu \rightarrow \bar{\nu}_e$

Since CKM is known to be a source of CP-violation in Nature, it is natural to expect a similar situation in the lepton sector.

The June 2011 revolution

- The T2K result which indicated electron-neutrino appearance triggered a revolution. The effect was confirmed by MINOS soon after.
- The observation of near/far ratios smaller than unity at long baseline reactor experiments were also interpreted as evidence for the disappearance of electron neutrinos and confirmed the non-zero value of θ_{13} with high statistical significance, as initially reported by Double-CHOOZ and culminating in the later announcement of a 5.2σ result by Daya Bay and 4.9σ by RENO
- With the present level of knowledge, neutrino oscillations are entering the precision era:
 - Δm_{21}^2 (3%), Δm_{31}^2 (4%), $\sin^2 \theta_{12}$ (5%), $\sin^2 \theta_{13}$ (15%), and $\sin^2 2\theta_{23}$ (15%)
 - MH unknown, $0 < \delta_{CP} < 2\pi$ full range at 2σ C.L.
- These exciting results close more than a decade of exploration of oscillations, and clearly define the way forward:
 - All three mixing angles of the Pontecorvo-Maki-Nakagawa-Sakata (PMNS) leptonic mixing matrix are non-vanishing and large. This has ascertained the 3×3 unitary character of the PMNS matrix, opening the possibility to observe its non-trivial complex nature.
 - This raises the intriguing possibility that neutrinos (or their heavy neutrino partners) might have played an important role in the early age of the Universe contributing to the creation of the baryon asymmetry which is responsible today for the matter dominance.
- ➔ These arguments strongly advocate a further exploration of neutrinos, and indeed in a more urgent and prominent way, but yet also more accessible given the large mixing angles.
- ➔ To observe evidence of CP violation in the leptonic sector has become one among the most important topics in Particle Physics today.

A decade after CHOOZ: the θ_{13} revolution

- T2K (Jun 2011):
 $\sin^2 2\theta_{13} = 0.03 - 0.34$ (90% CL).

T2K Collaboration, Phys.Rev.Lett. 107 (2011) 041801

- MINOS (July 2011):
 $\sin^2 2\theta_{13} \neq 0$ at 89% CL.

MINOS Collaboration, Phys.Rev.Lett. 107 (2011) 181802

- Double CHOOZ (Dec 2011):
 $\sin^2 2\theta_{13} = 0.017 - 0.16$ (90% CL).

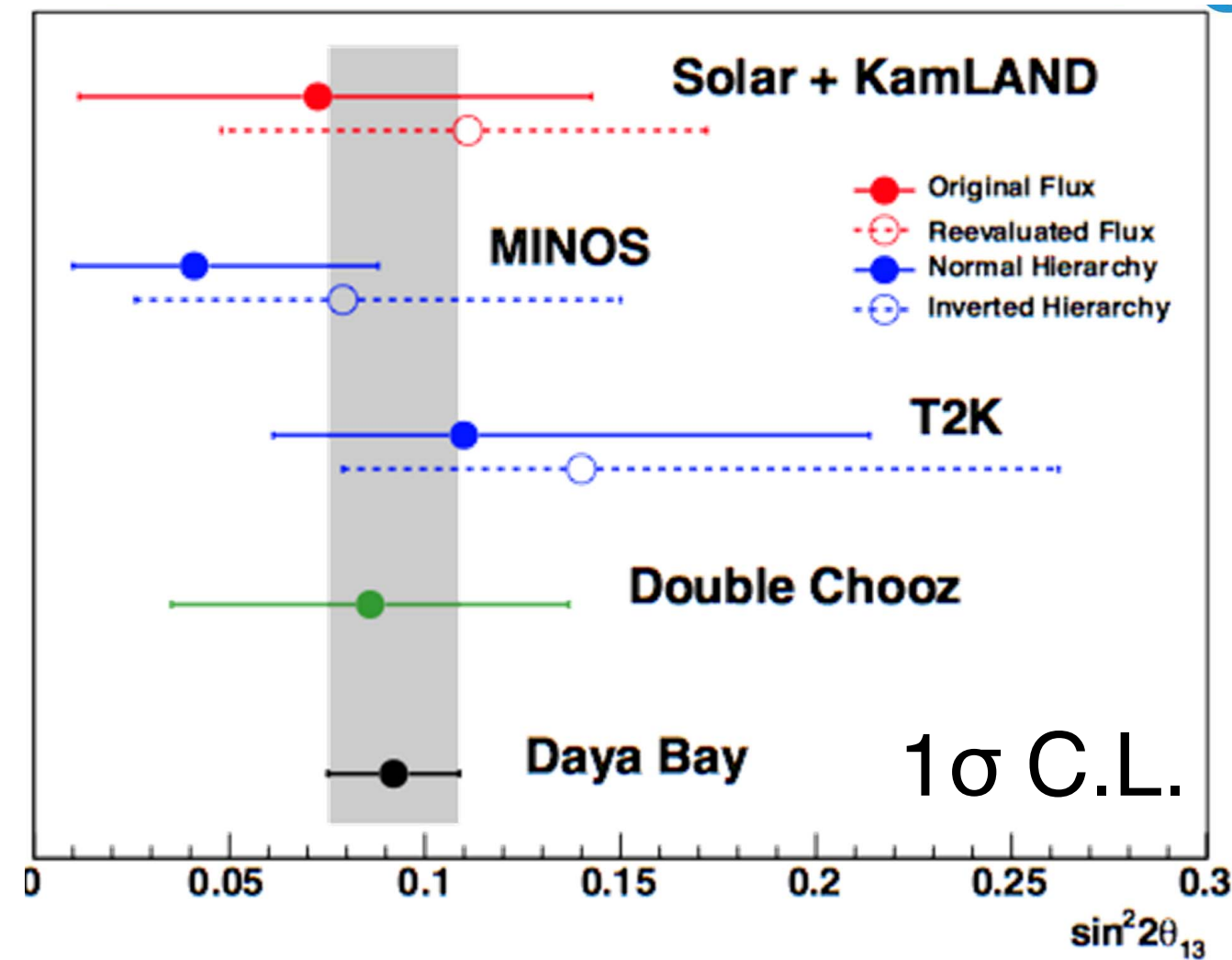
Double CHOOZ Collaboration, arXiv: 1112.6353 [hep-ex].

- **Daya Bay** (Mar 2012):
 $\sin^2 2\theta_{13} \neq 0$ at 5.2σ (!),
best-fit = 0.092.

Daya Bay Collaboration, arXiv: 1203.1669 [hep-ex].

- **Reno** (April 2012):
exclude no oscillations at 4.9σ

Reno Collab, arXiv:1204.0626v2



$$\sin^2 (2\theta_{13}) = 0.092 \pm 0.016(stat) \pm 0.005(syst)$$

$$\sin^2 2\theta_{13} = 0.113 \pm 0.013(stat) \pm 0.019(syst)$$

The knowns and unknowns...

$$\sin^2 \theta_{12} = 0.312^{+0.017}_{-0.015}$$

$$\Delta m_{12}^2 = (7.59^{+0.20}_{-0.18}) \times 10^{-5} \text{ eV}^2$$

$$\sin^2 \theta_{23} = \begin{array}{c} 0.51 \pm 0.06 \\ 0.52 \pm 0.06 \end{array}$$

$$\Delta m_{31}^2 = \begin{array}{c} 2.45 \pm 0.09 \\ -(2.34^{+0.10}_{-0.09}) \end{array} \times 10^{-3} \text{ eV}^2$$

Solar

parameters
SK, SNO, Kamland

Atmospheric

parameters
SK, K2K, MINOS, T2K

→ *Mass ordering is hierarchical or inverted ?*

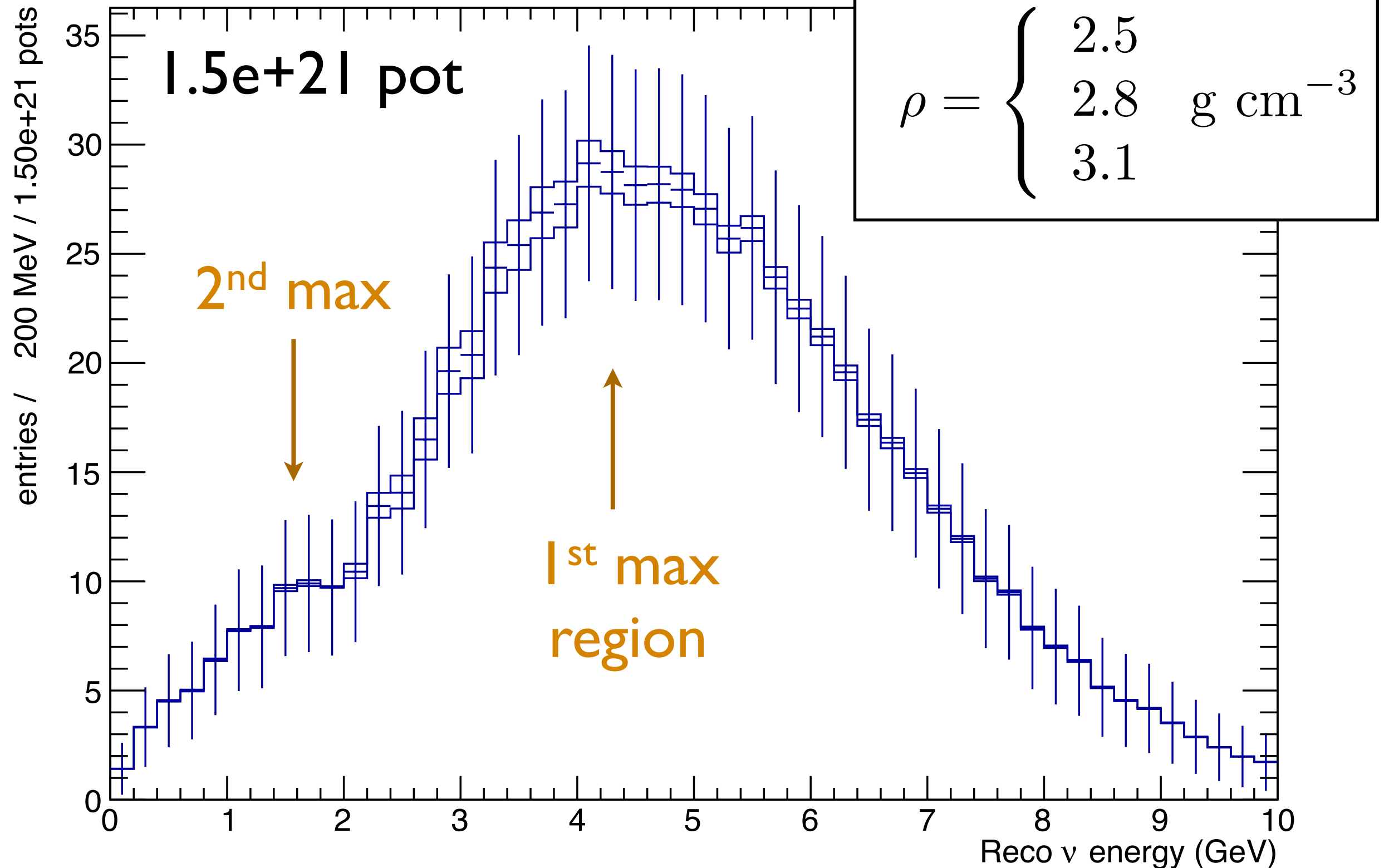
$$\sin^2 (2\theta_{13}) \simeq 0.09 \pm 0.02$$

$$\delta_{CP} = [0, 2\pi]$$

→ *Complex phase is unknown. Because of similarities with CKM matrix, it is natural to expect a CP violation in the lepton sector. But CKM & PMNS angles are very different, what is the size of the CP effect in leptons??*

Matter density effects

e-like sample



CP-phase sensitivity

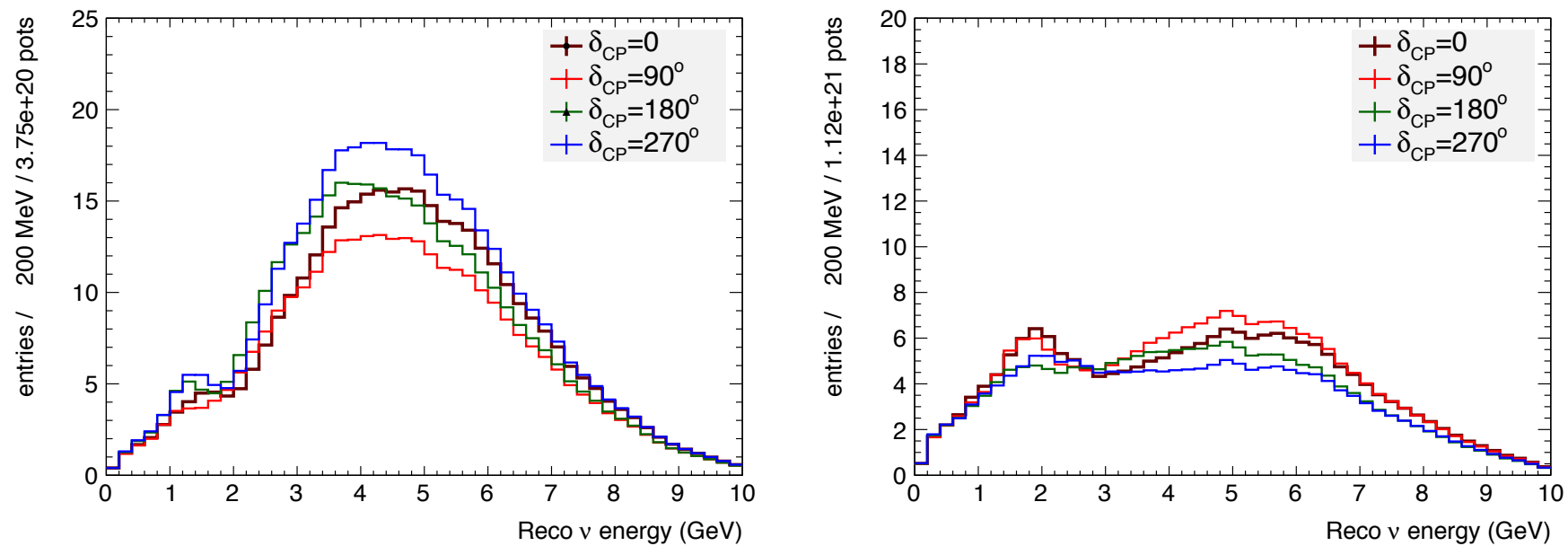


FIG. 73: Reconstructed event energy for (left) neutrino horn polarity running and (right) antineutrino horn polarity running, for different values of true δ_{CP} and for normal mass hierarchy (NH). A 25%-75% share between neutrino and antineutrino running mode and a total of 1.5×10^{21} pot have been chosen.

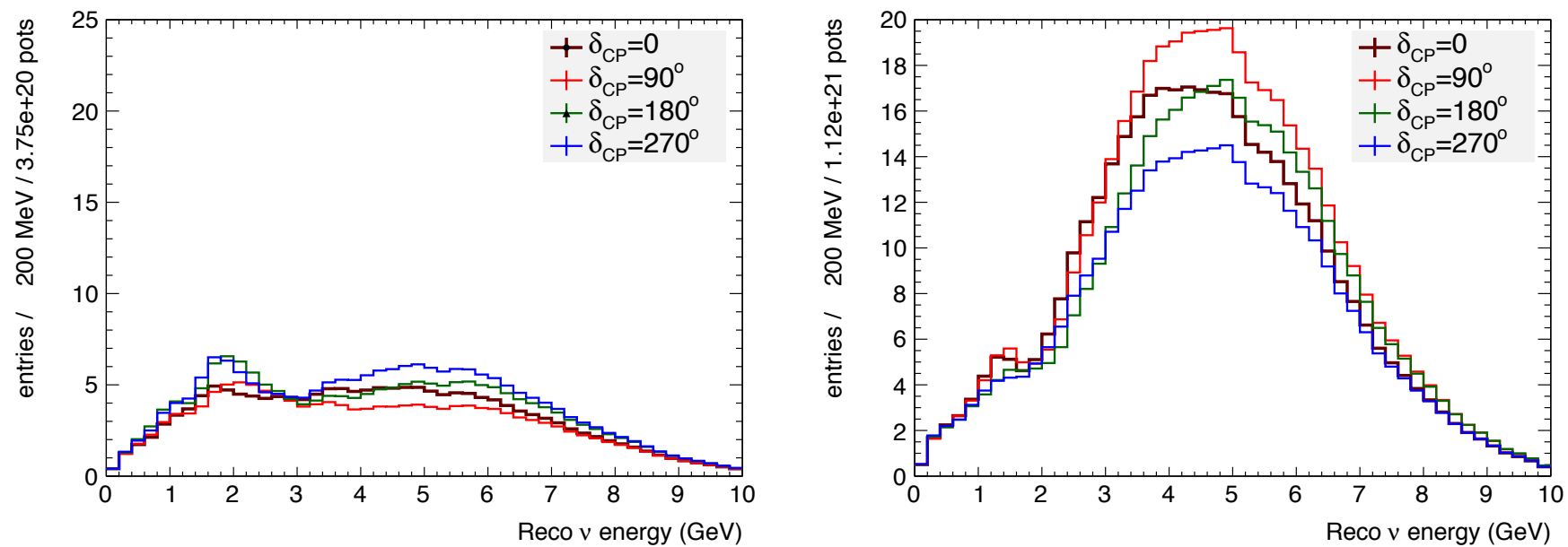
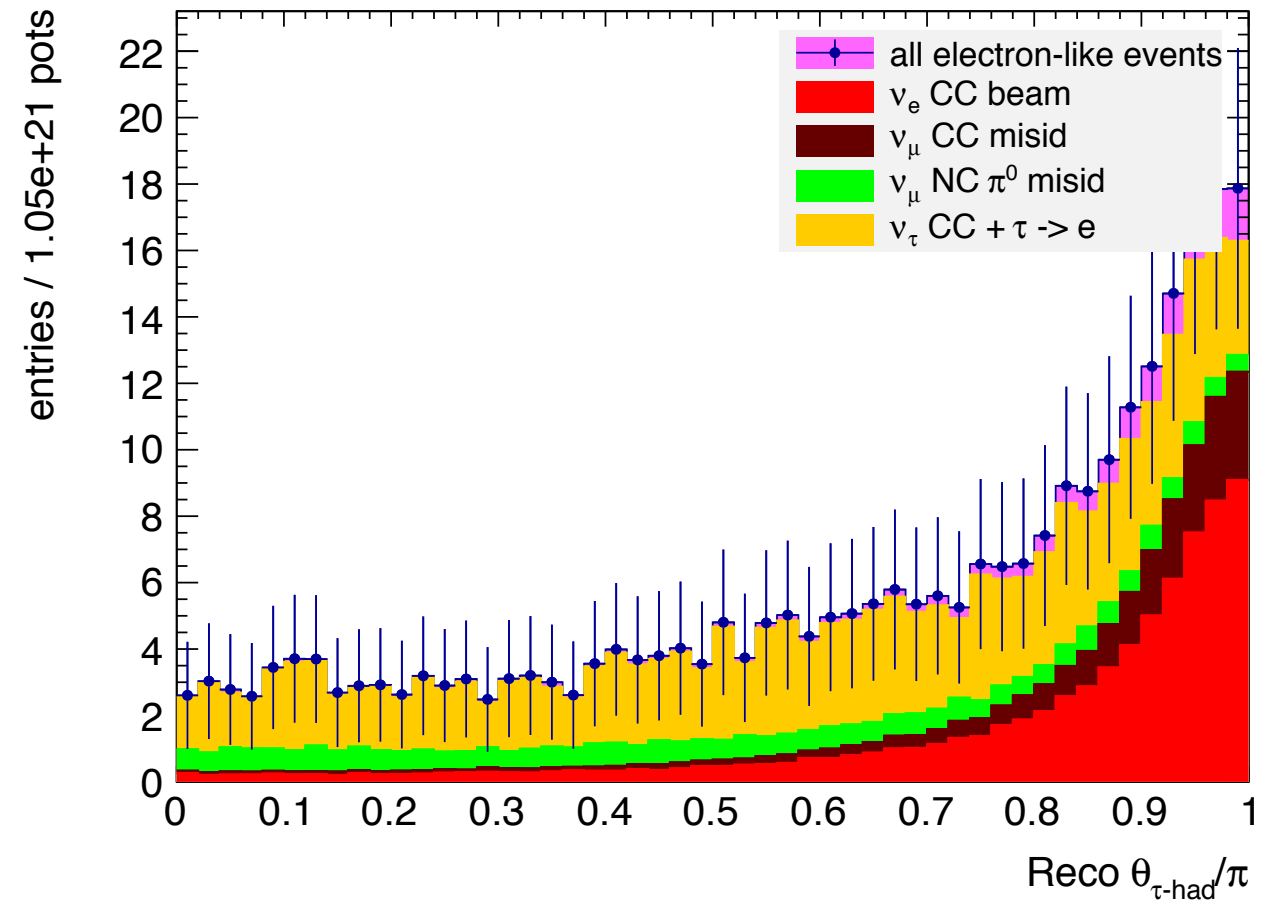
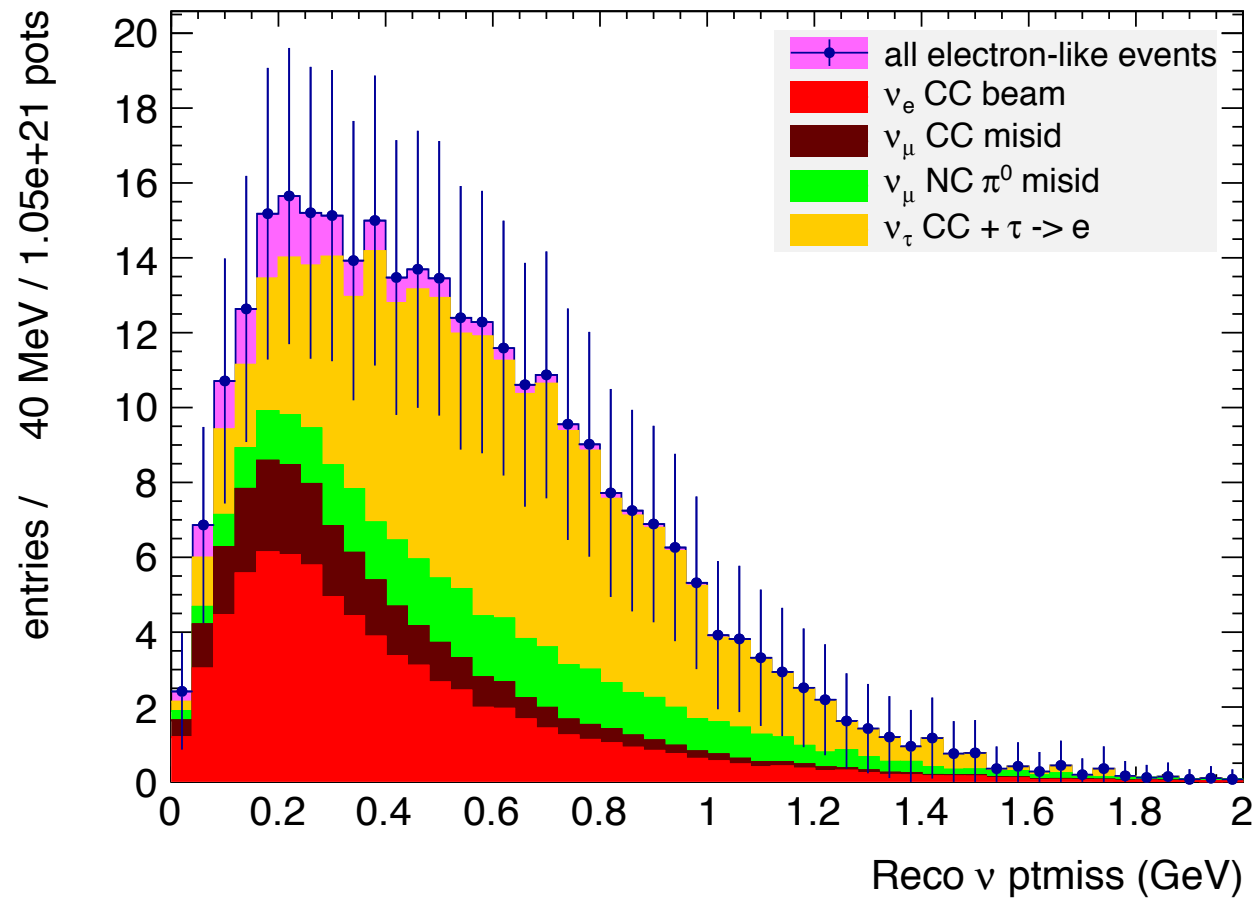
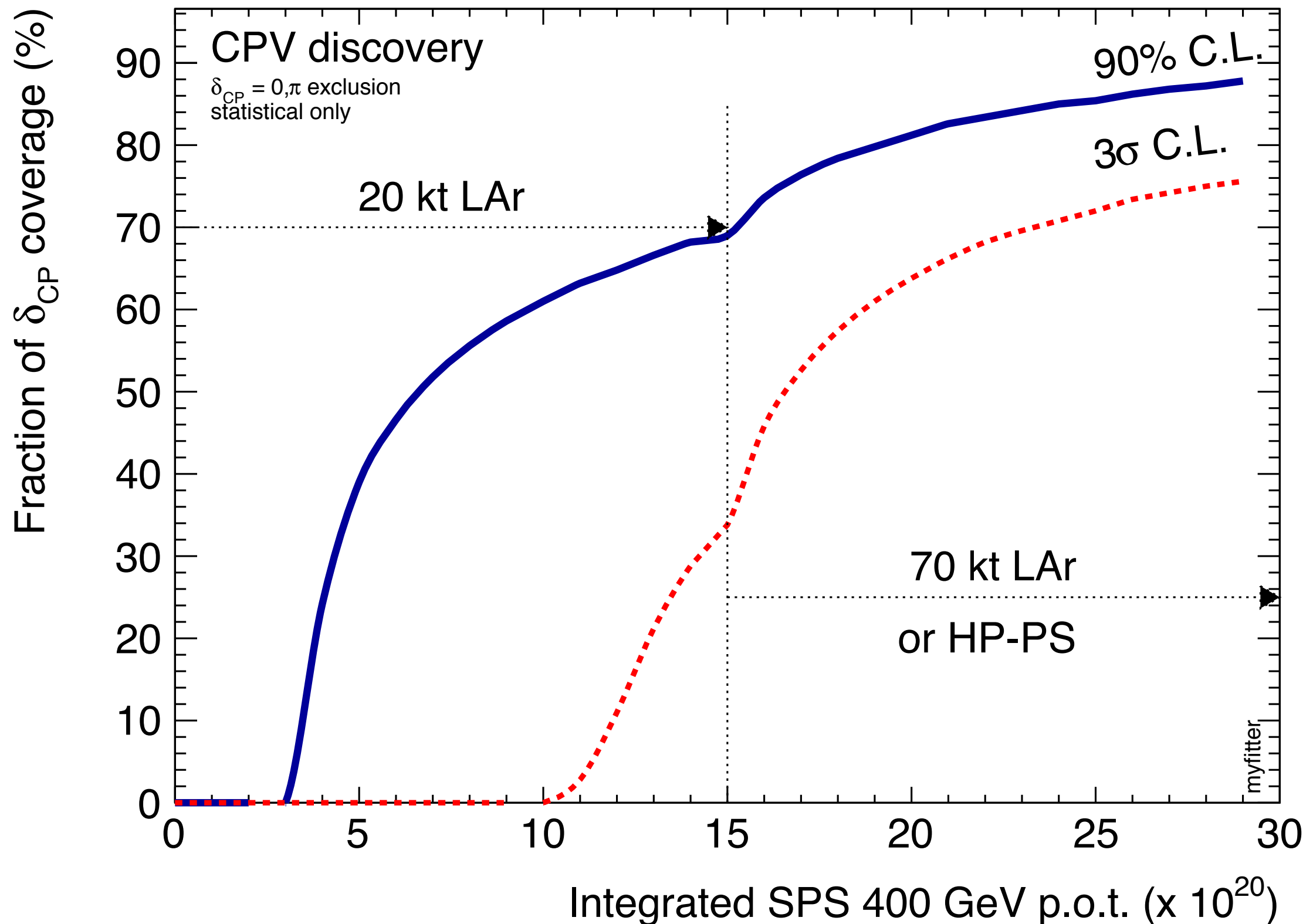


FIG. 74: Same as Figure 73 but for inverted mass hierarchy (IH).

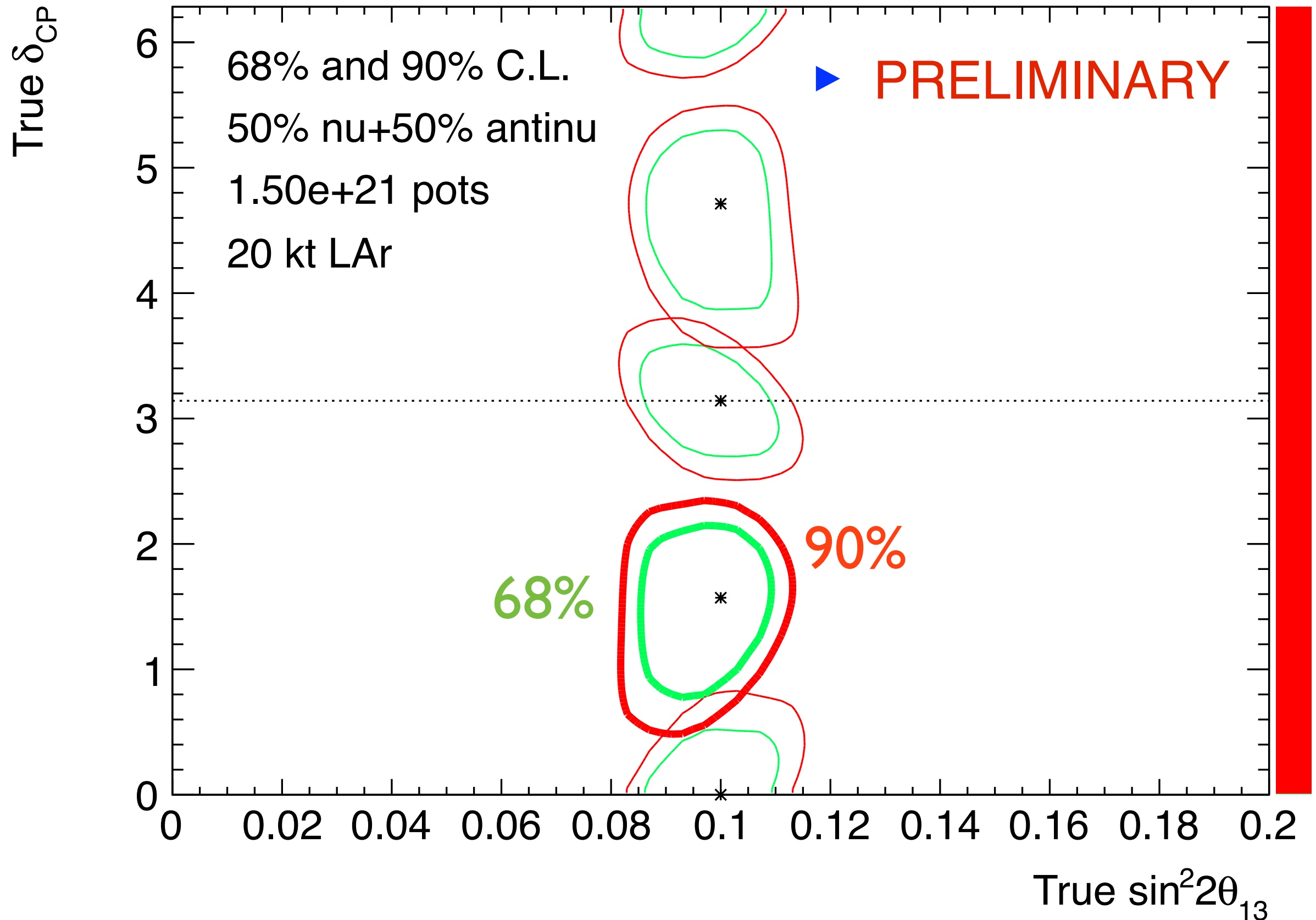
Tau like sample



CPV discovery - statistical only

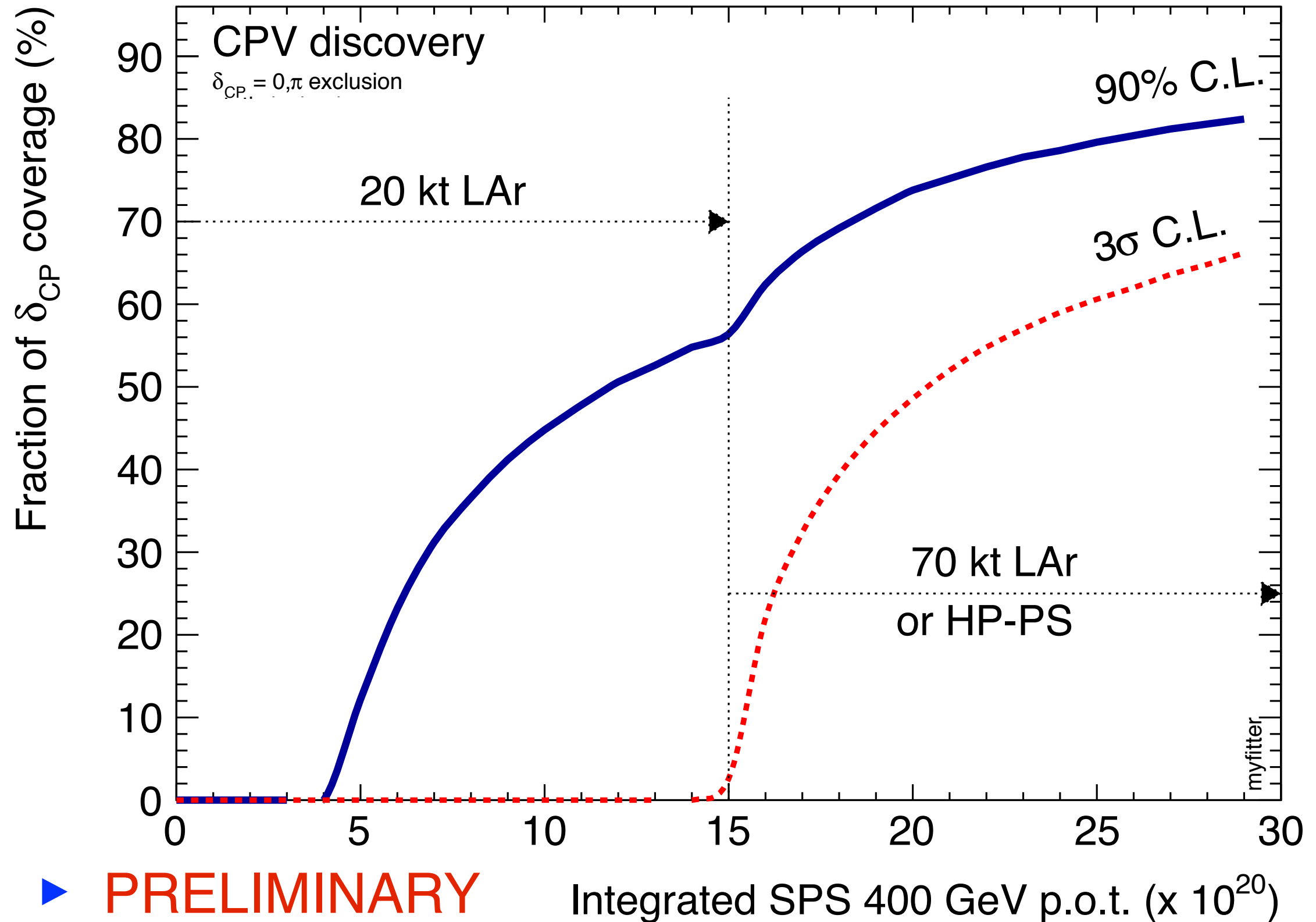


CP-phase determination

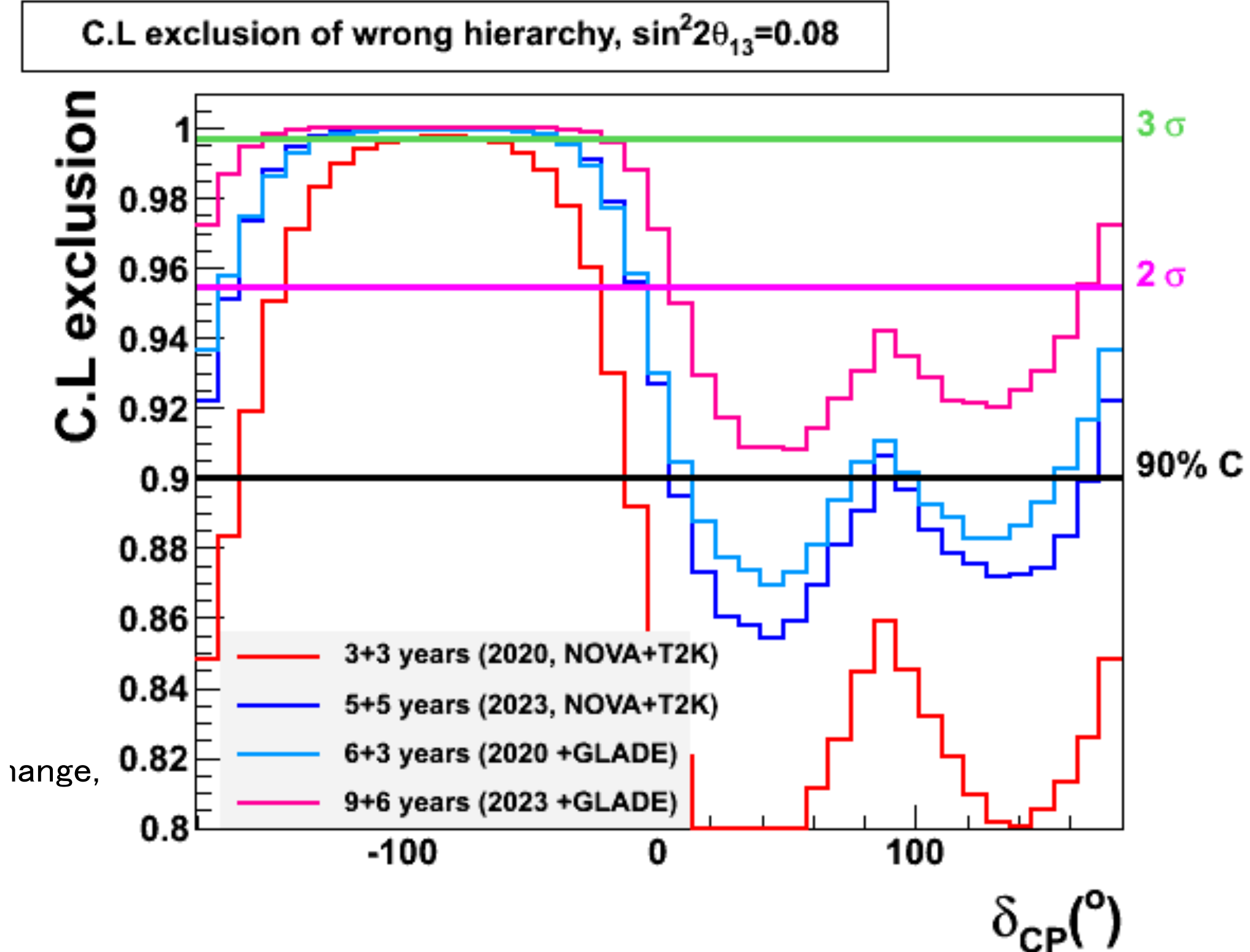


Effect of matter uncertainty

★ INFLATED ERROR ON MATTER DENSITY $\pm 10\%$

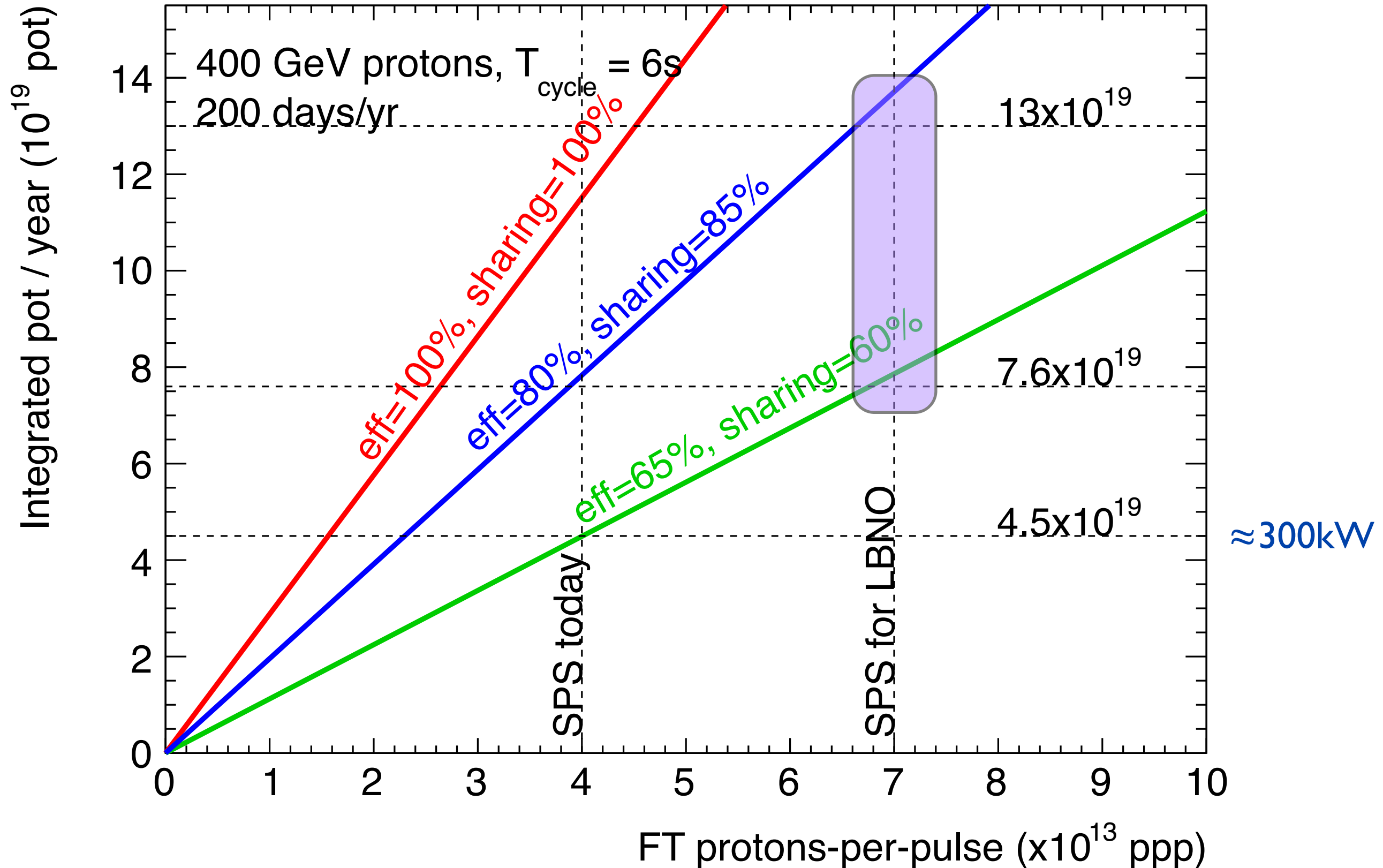


Situation in 2023 ?

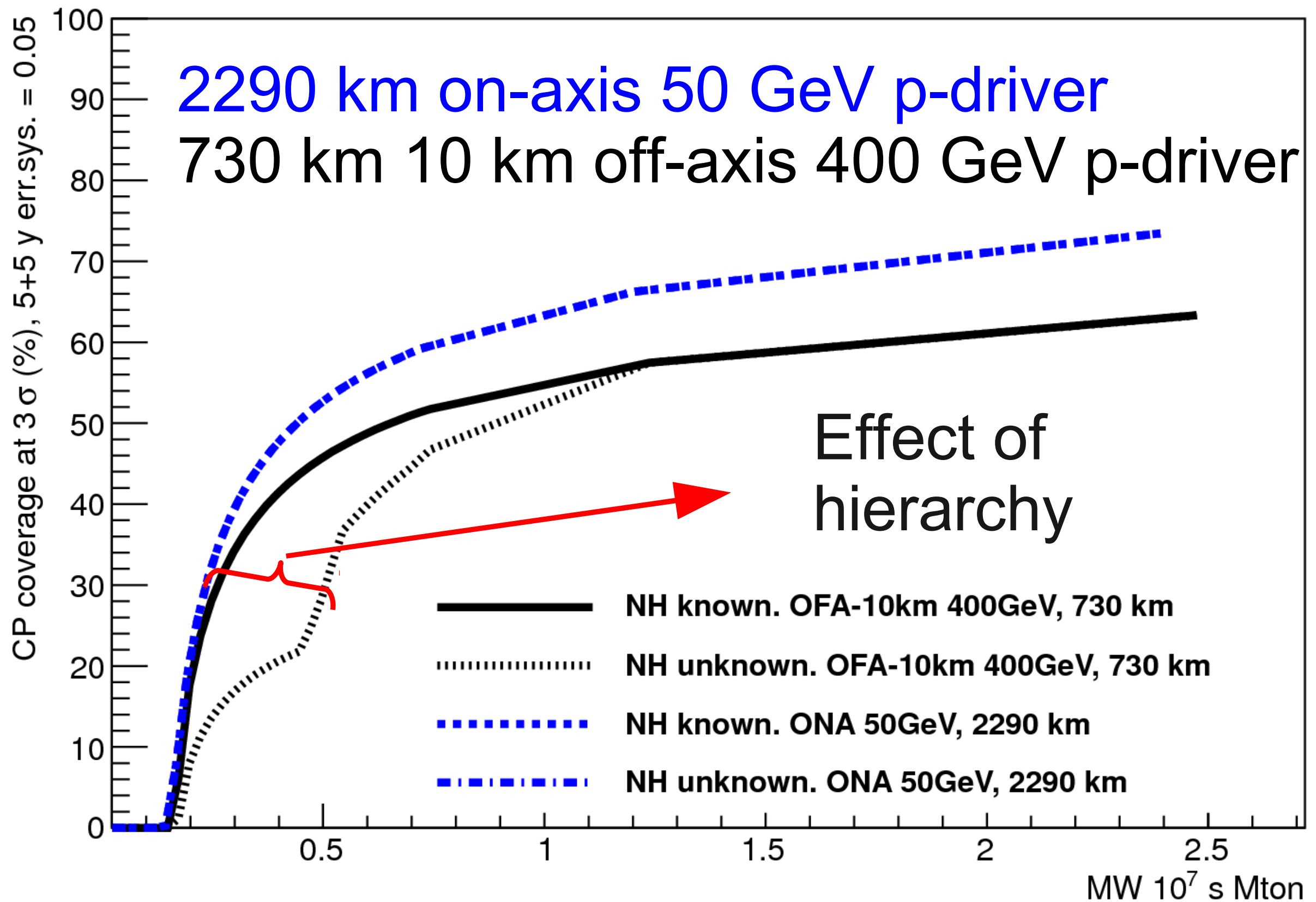


Most likely we will reach a $\approx 2\sigma$ MH determination

SPS 400 GeV p.o.t / year



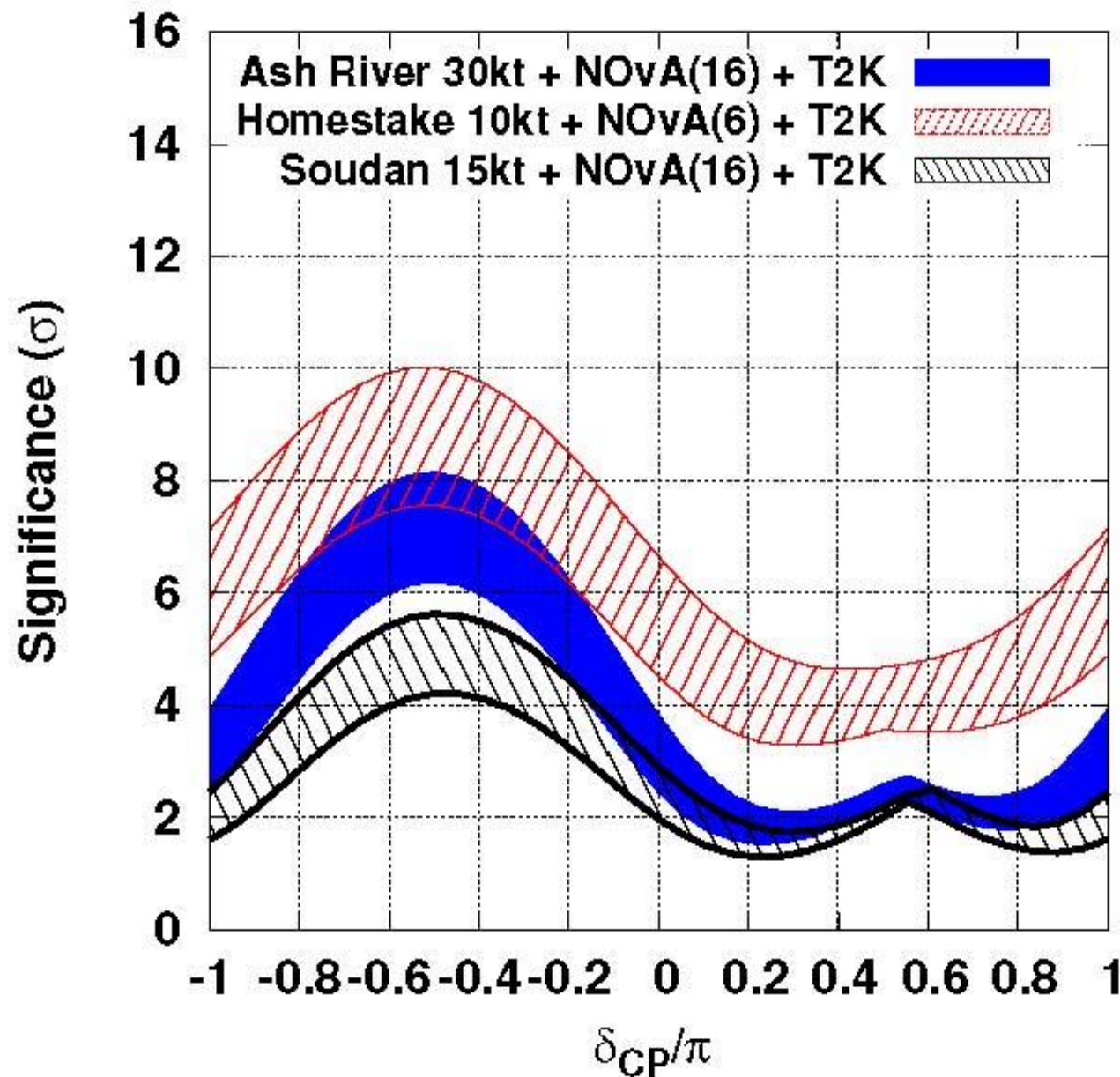
CP coverage at 3σ (%), 5+5 y err.sys. = 0.05



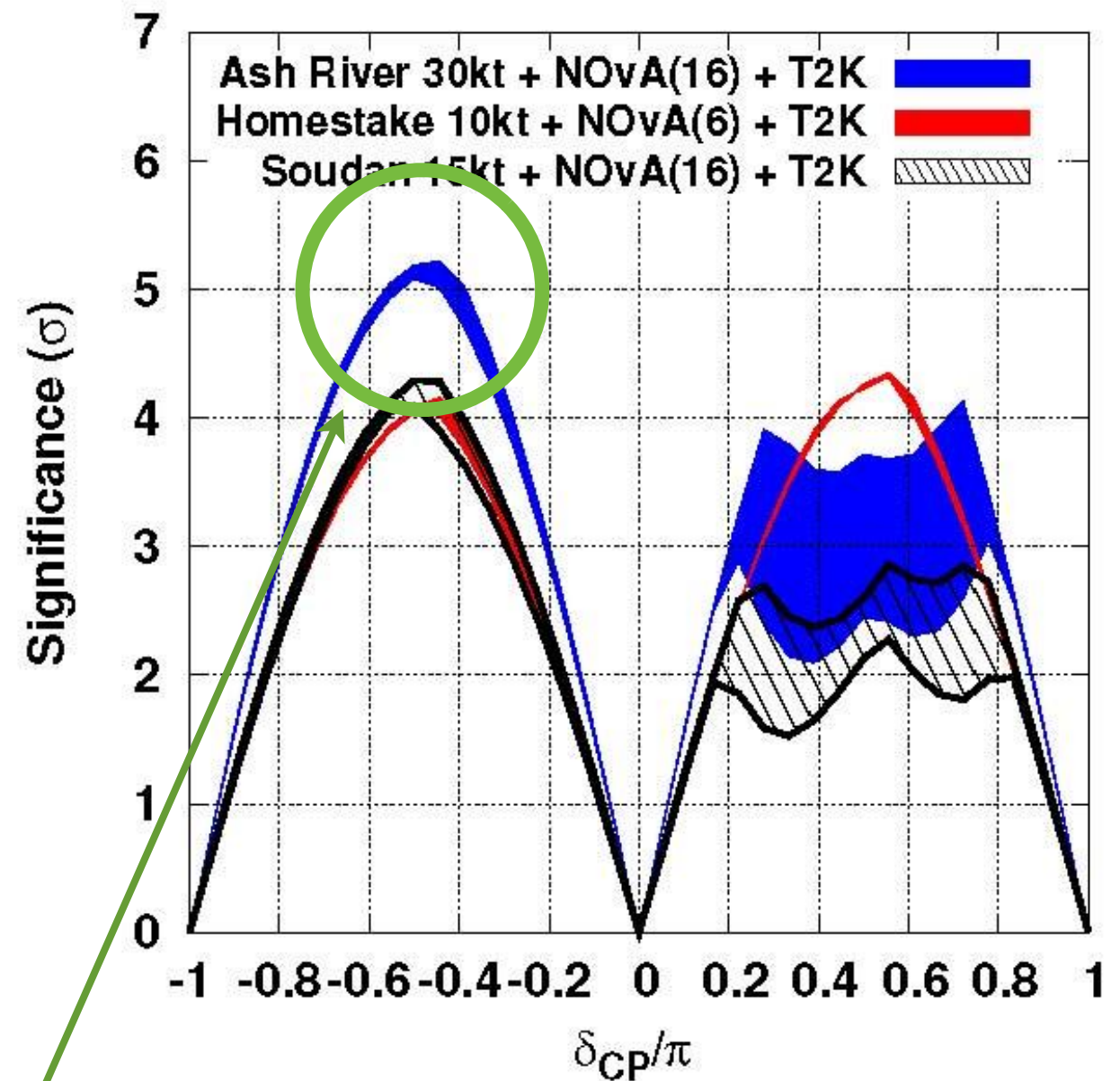
A. Longhin et al., NUTurn12

LBNE: 10 years @ 700kW

Mass Hierarchy Significance vs δ_{CP}
Normal Hierarchy, $\sin^2(2\theta_{13})=0.07$ to 0.12



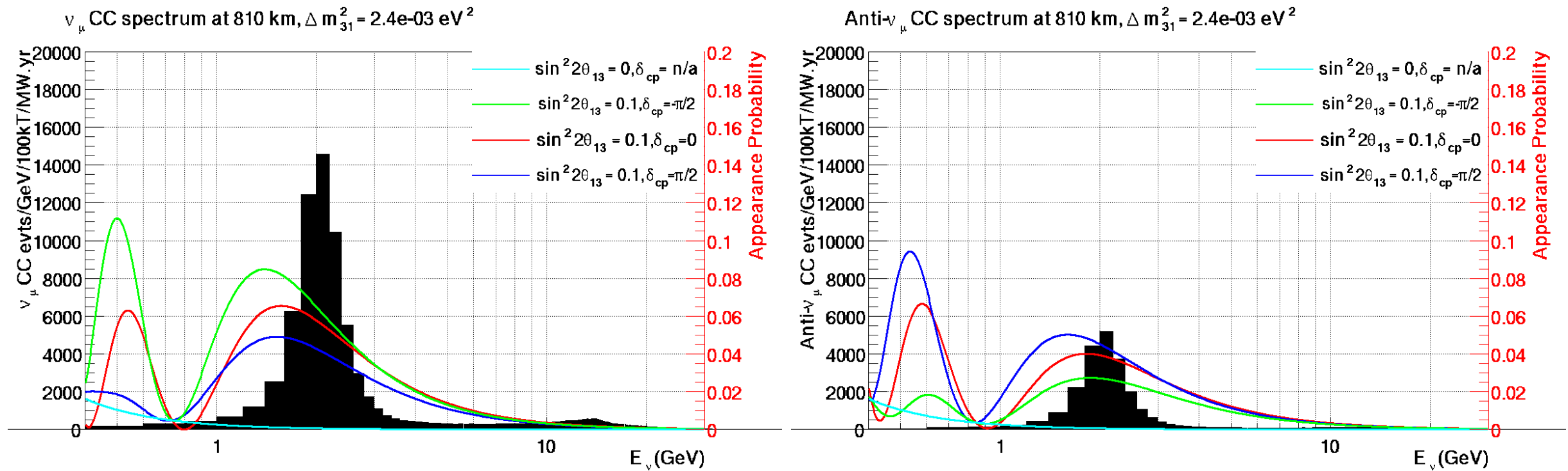
CPV Significance vs δ_{CP}
NH(IH considered), $\sin^2(2\theta_{13})=0.07$ to 0.12



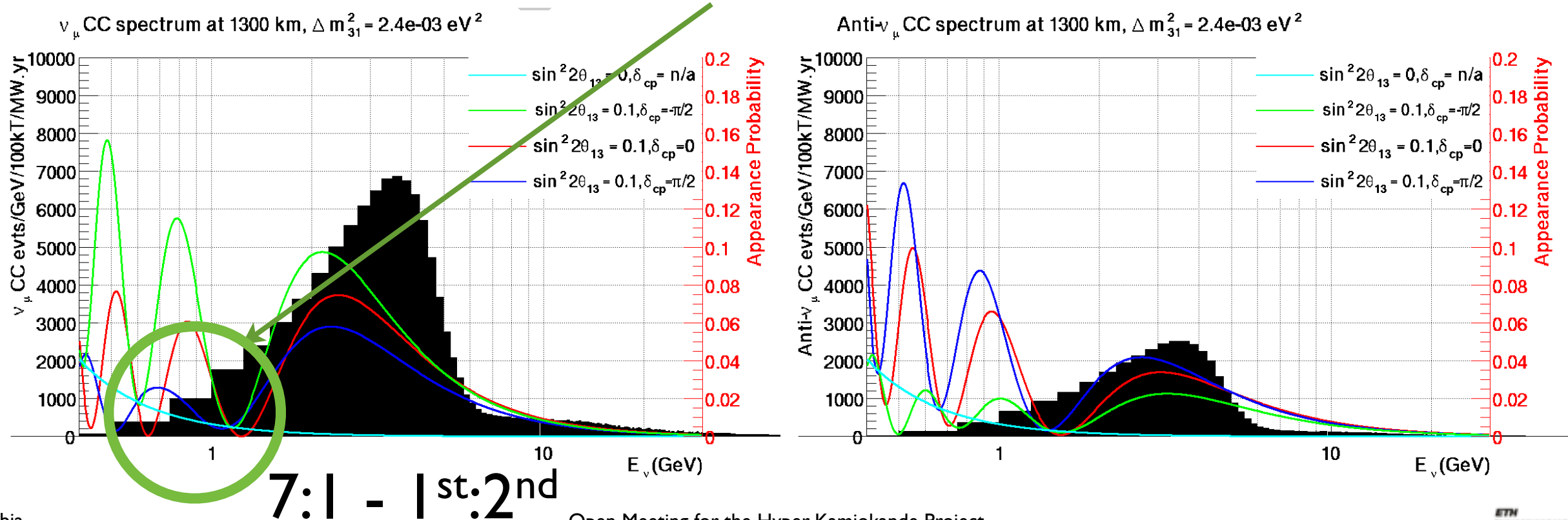
Be aware: Ash River has the best CPV sensitivity when MH is determined ! the displayed sensitivities come mostly from parameter fitting around 1st maximum

LBNE: Ash River & Homestake

810 km off-axis is totally discounted (no E-dependence !)



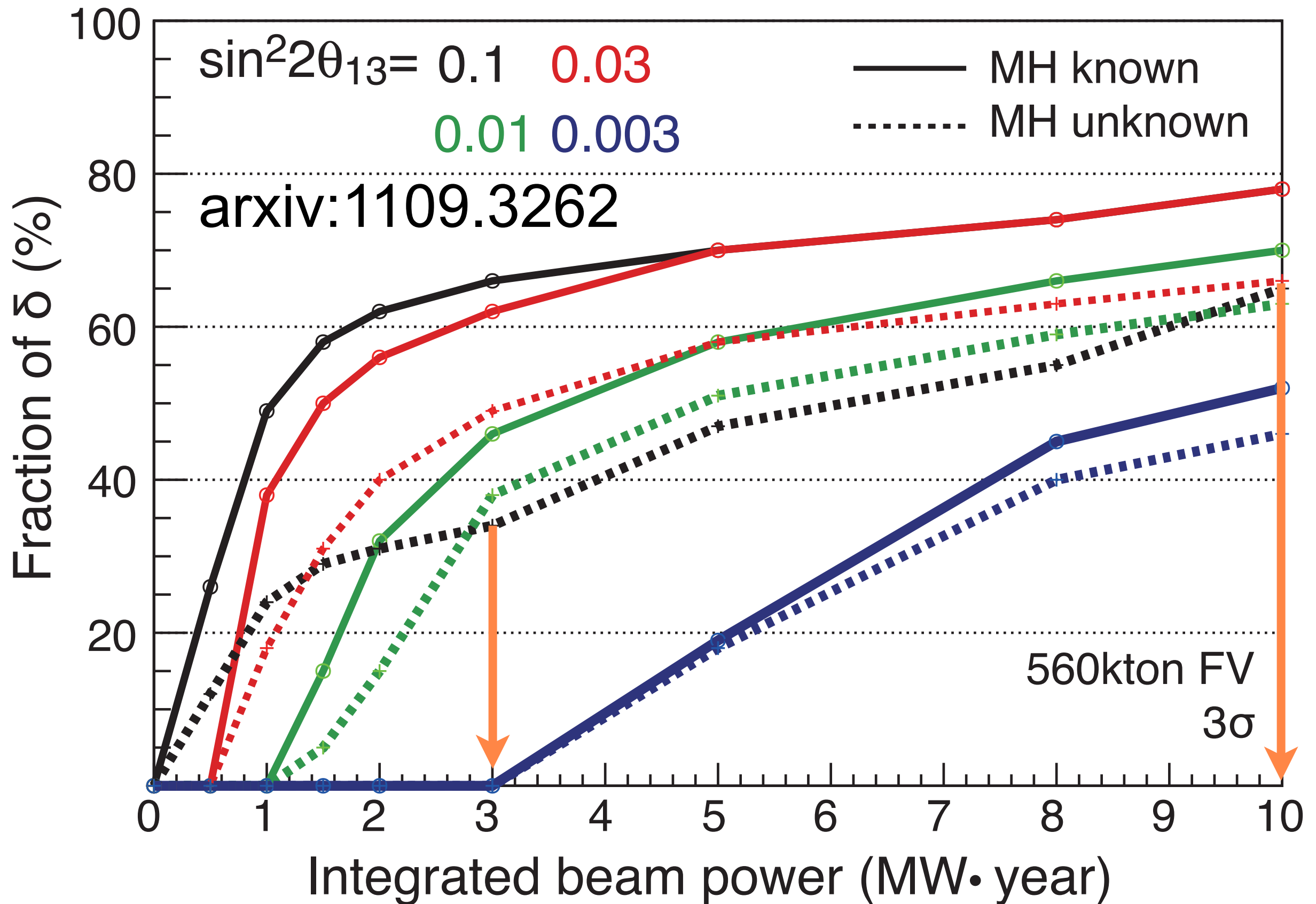
At 1300 km the second maximum is hard to observe !



Why the neutrino mass hierarchy ?

- **CP-violation:** necessary input to solve CPV problem. For example, for the HyperK LOI arxiv:1109.3262 (which considers a 540kton FV and hence has the highest statistical power):
 - ➔ 3 MW×years (note: >10 years at present JPARC MR power)
MH known: 65% coverage → MH unknown: 35% coverage
 - ➔ 10 MW×years needed to reach 65% coverage if MH unknown!
rather unlikely within present JPARC projections.
- **$0\nu\beta\beta$ searches:** necessary input to interpret both negative and positive isotope lifetime results, in terms of neutrinos (as opposed to some other source of lepton number violation).
- **BSM/GUT theories:** important ingredient for model building. An inverted hierarchy would have interesting implications.
- **We need a definitive & conclusive determination of the MH !**

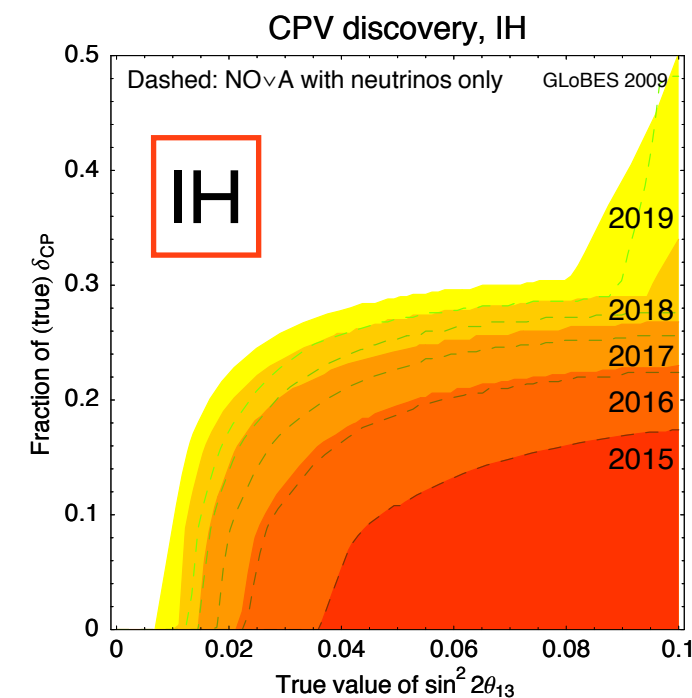
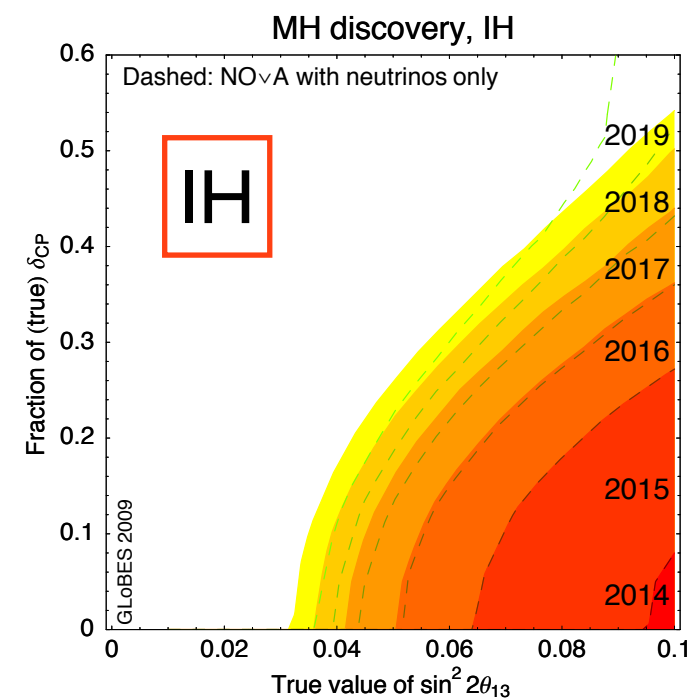
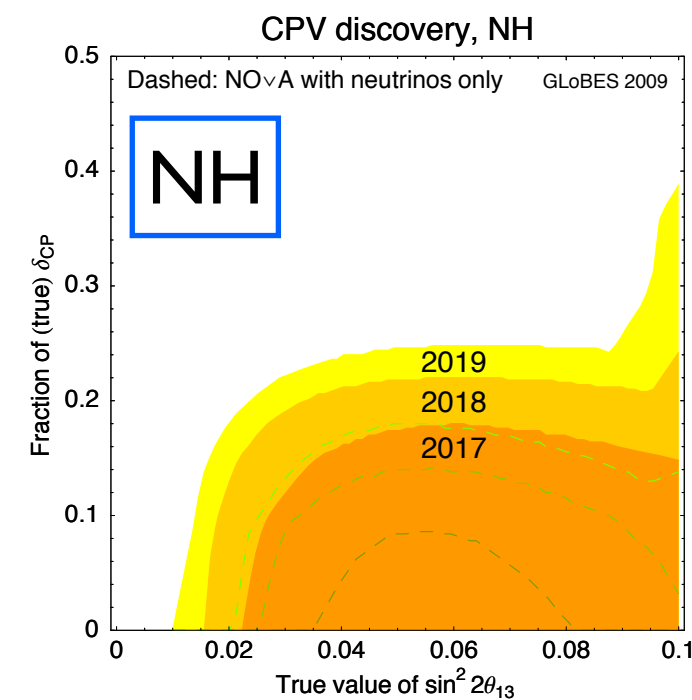
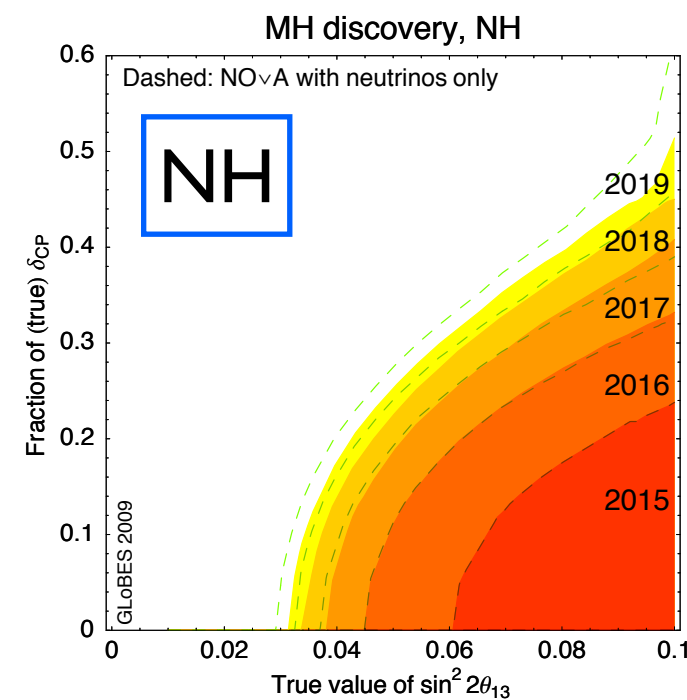
HyperKamiokande CPV



T2K and NOvA: in the future

- Preliminary and not official estimates of the combined T2K, NOvA and reactors sensitivity
- Nominal beam power scenarios (750kW). Need to check beam power assumptions.
- For $\sin^2 2\theta_{13}=0.1$, approximately (at 90%C.L.):
 - MH: $\approx 50\%$ coverage
 - CPV: $\approx 30\text{-}40\%$ coverage (robustness vs MH ?)
- Are these curves too optimistic ?
- Atmosphericics to the rescue ?
- Official predictions to be produced by experiments with revised projections.
- **CPV and MH are “extracted” from a global fit. Not a direct proof of CP nor direct measurement of MH induced matter effect !!**

MH 90%C.L. sensitivity **CPV**



Huber et al., JHEP 0911:044,2009

not official

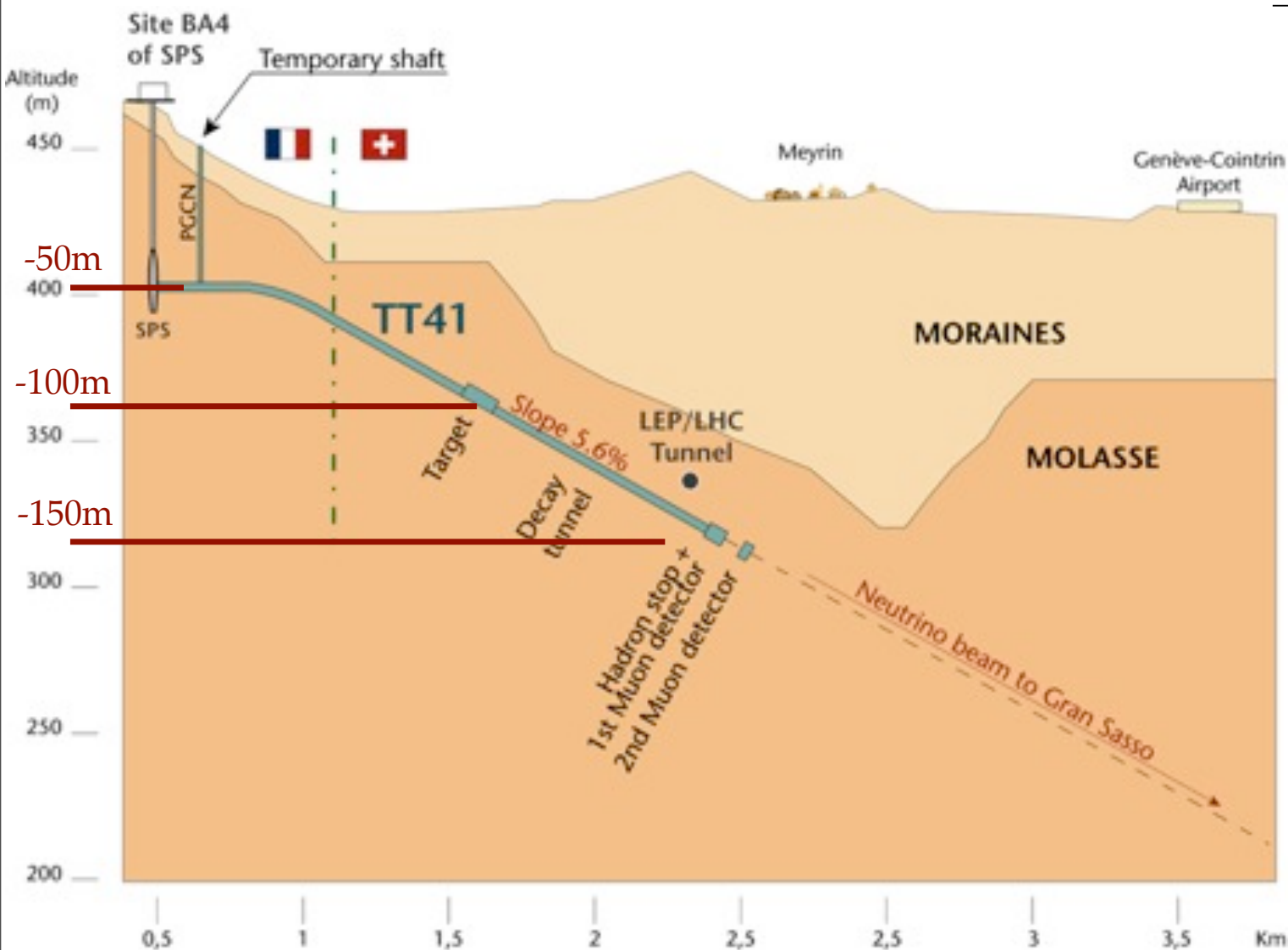
LBNO EoI: the physics reach

- Initial setup 20 kton LAr LEM TPC + MIND + CERN SPS 700kW
- **Ultimate** long baseline oscillations measurements:
 - LBNO can measure all transitions (e/ μ /tau) and determine precisely oscillation parameters. It can achieve a **5 σ C.L. determination of the neutrino mass hierarchy** in a few years. In a 10 years run, it explores a **significant part of the CPV parameter space, namely 60% CPV coverage at 90%C.L.**
 - Both the local situation and the distance make it such that it can evolve into larger detector(s) and a more powerful beams (e.g HP-PS and/or NF) and thus, **offers a long term vision**. For example, with a three-fold increase in exposure, **it reaches 75% CPV coverage at 3 σ C.L..** Competitive with T2HK (even more with JPARC MR at 700kW...) and LBNE.
- **Significantly** extended sensitivity to nucleon decay in several channels. E.g. some channels with sensitivity similar to HK:
$$Br(p \rightarrow \bar{\nu} K) > 2 \times 10^{34} y(90\%C.L.) \quad Br(n \rightarrow e^- K^+) > 2 \times 10^{34} y(90\%C.L.)$$
- **Interesting** astrophysics: LBNO acts as an nu-observatory in the 10 MeV-100 GeV range. **5600 atmospheric events/yr** **relic SN, WIMP annihilation, ...**
>10000's events @ SN explosion@10kpc

Depth considerations for CN2PY

- The depth for the installations is the major concern
 - 18% slope compared to 5.6% for CNGS

	Distance [m]	Depth [m]
target	-	0
hadron stop	300	-54.3
muon station	330	-59.9
1st position for near detector	500	-90.57
2nd position for near detector	800	-145.21



Starting the beam from the SPS level adds **~100m** to the depth of the installations

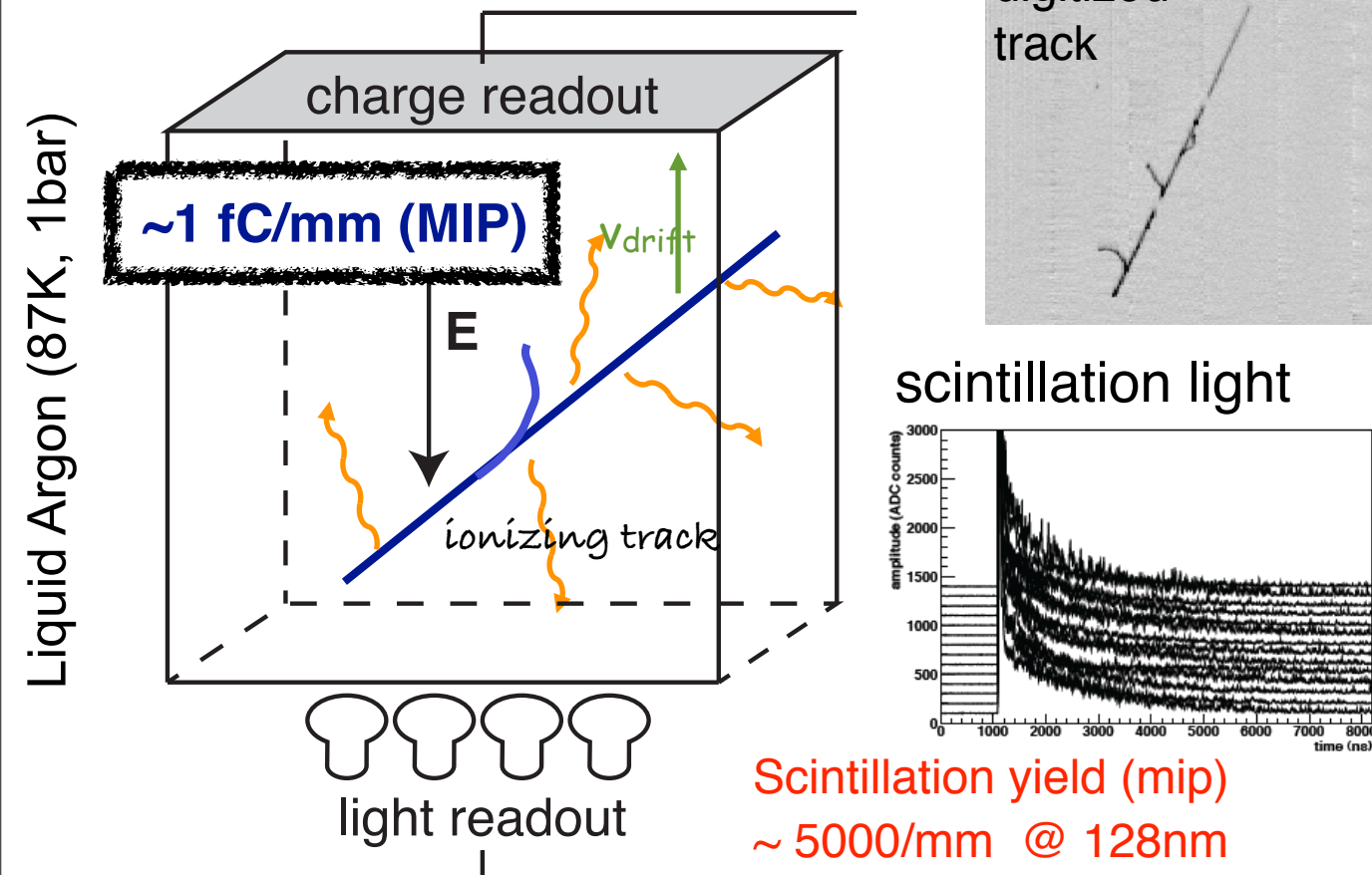
► Staying in the **molasse layer** has quite some advantages for the civil engineering and radiation to environment (underground water activation) issues

Physical parameters and challenges

Liquid Argon:

- + High density, cheap medium
- + Quasi free electrons from ionizing tracks are drifted in LAr (87K, 1bar) by E_{drift} .
- + Electron drift velocity $\approx 2\text{mm}/\mu\text{s}$ @ 1 kV/cm
- + Electron cloud diffusion is small
($\sigma \approx \sqrt{2Dx}/v_{\text{drift}} \approx \text{mm}$ after several meters of drift)
- + High scintillation yield (@ 128 nm) can be used for T_0 , trigger, ...

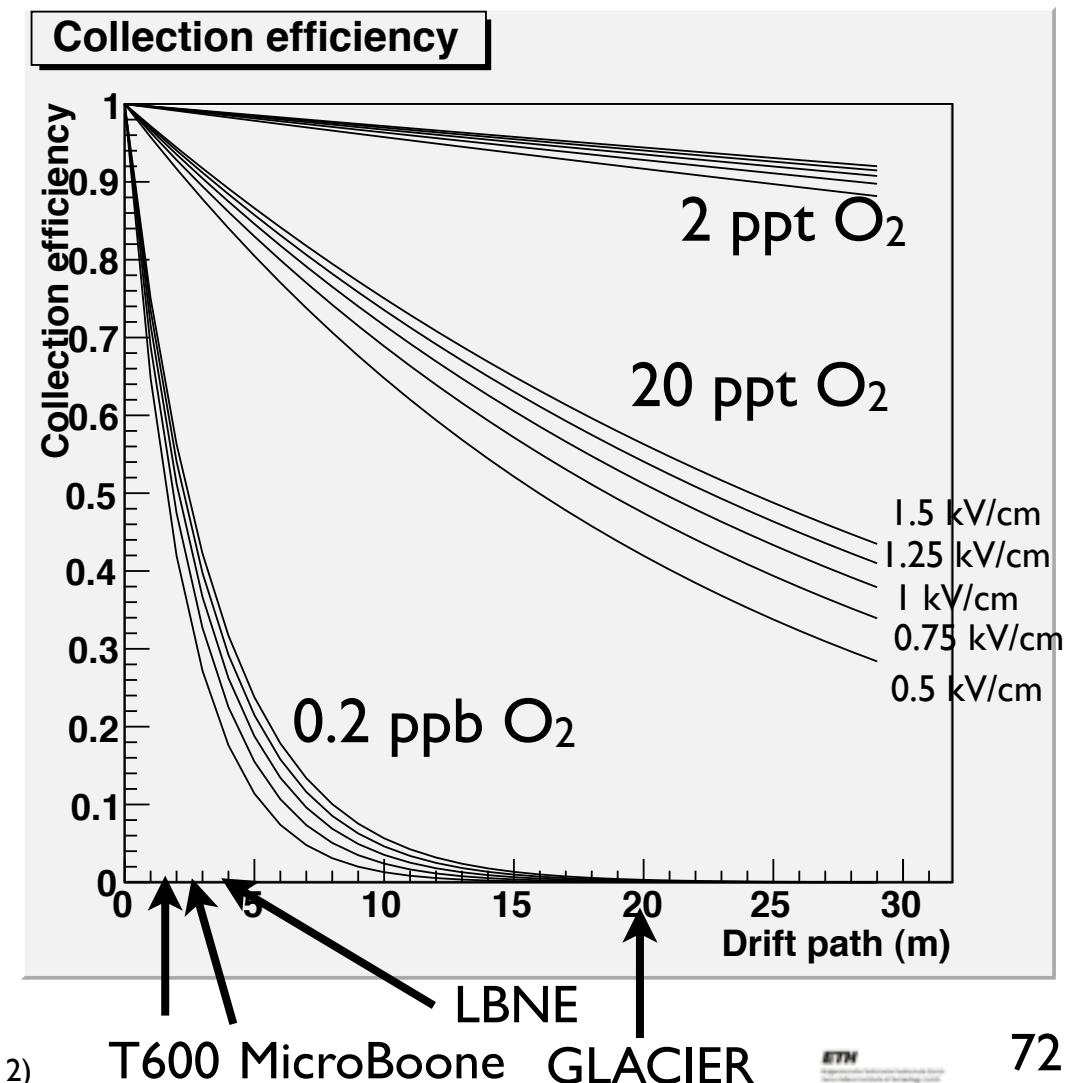
Charge yield after e-ion recombination (mip) $\sim 1\text{ fC/mm}$
(~ 6000 electrons/mm) before finite e-lifetime correction



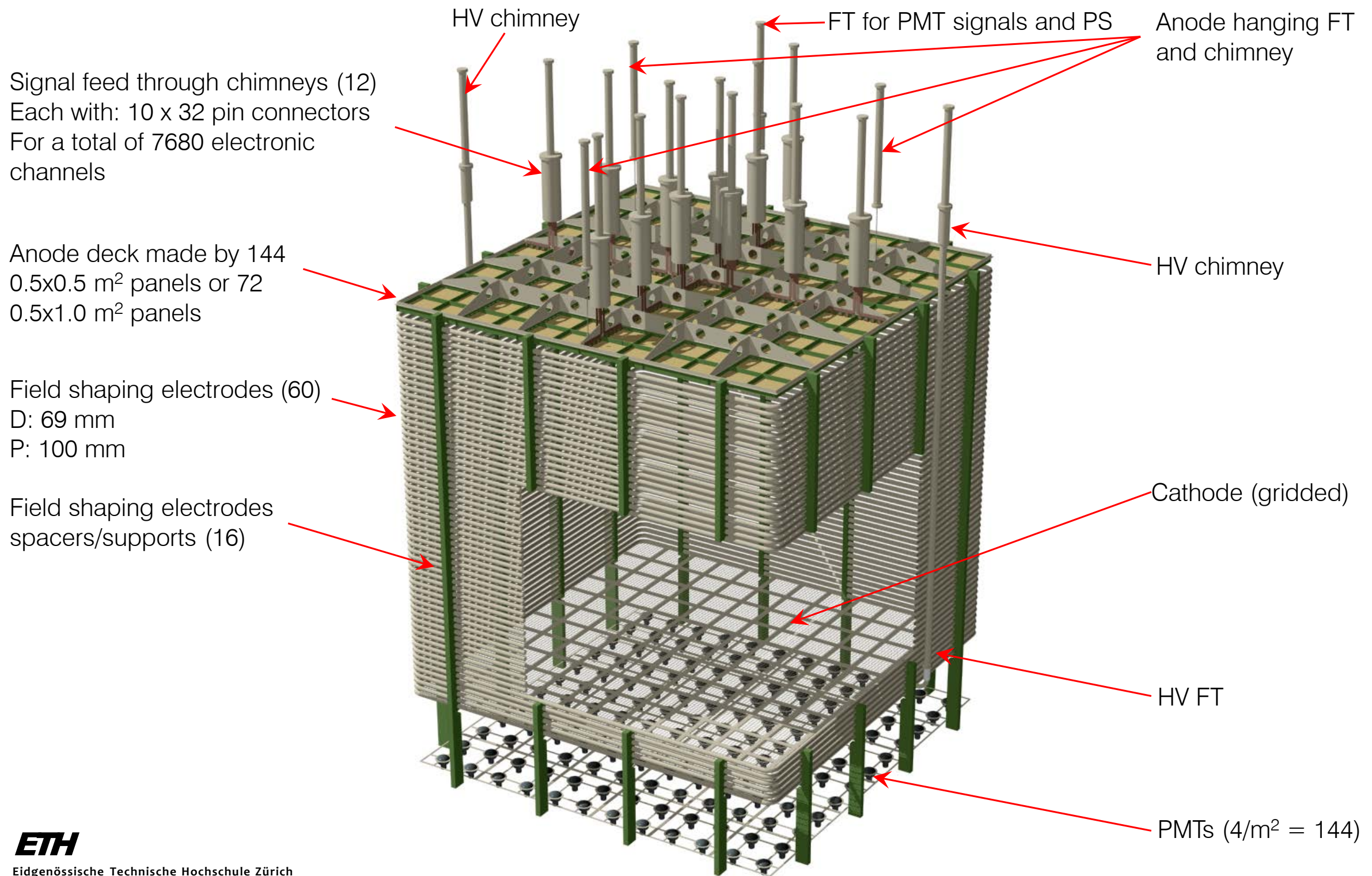
Technical challenges:

- Long drift requires ultra high purity
 - * free of electro-negative molecules (O_2 , H_2O , ...)
 - Goal $\ll 100\text{ ppt O}_2$ equivalent !!
 - * Drift field implies high voltage on the cathode
- Large wire chambers at cryogenic T
- No charge amplification in liquid: fC-level charge sensitive preamplifiers (can be partially solved by LAr LEM TPC – see later)
- Large #readout channels
- Large cryogenic systems

Charge attenuation because of attachment to impurities



LAGUNA LAr prototype @ CERN



ETH

Eidgenössische Technische Hochschule Zürich
Swiss Federal Institute of Technology

Franco Sergiampietri, 20 August 2012, 7

The LAr TPC features

★ The LAr TPC is the very successful marriage between the “gaseous TPC” and “the liquid argon calorimeter” to obtain a dense and very fine grained three dimensional tracking device with local dE/dx information and a homogenous full sampling calorimeter

★ Detector performance:

- 3D tracking, mm-scale spatial resolution with local dE/dx
- fully sensitive, $\approx 2\%X_0$ sampling rate and excellent energy resolution
- excellent particle identification (range vs dE/dx), e/π^0 separation with $\varepsilon=90\%$ for rejection factor >100 .
- continuously sensitive (“trigger-less” mode)

★ Technology achievement:

- large ultra high vacuum ($<1\text{e-}9$ mbar lt/s) and cryogenic systems ($T=87\text{K}$)
- ultra high purity liquid argon (<30 ppt O_2 equiv $\Rightarrow >10$ ms lifetime)
- large cryogenic wire chambers (up to 9m long wires)
- very high drift voltage (up to 150 kV)
- low noise fC charge sensitive readout electronics ($S/N > 10$ for m.i.p with $C_{\text{det}} \approx 400\text{pF}$)

★ Physics performance:

- Kinematical reconstruction of QE events (ICARUS 50L@CERN WNF)
- Inclusive cross-section measurement on Argon (ArgoNEUT)
- Many published studies on:

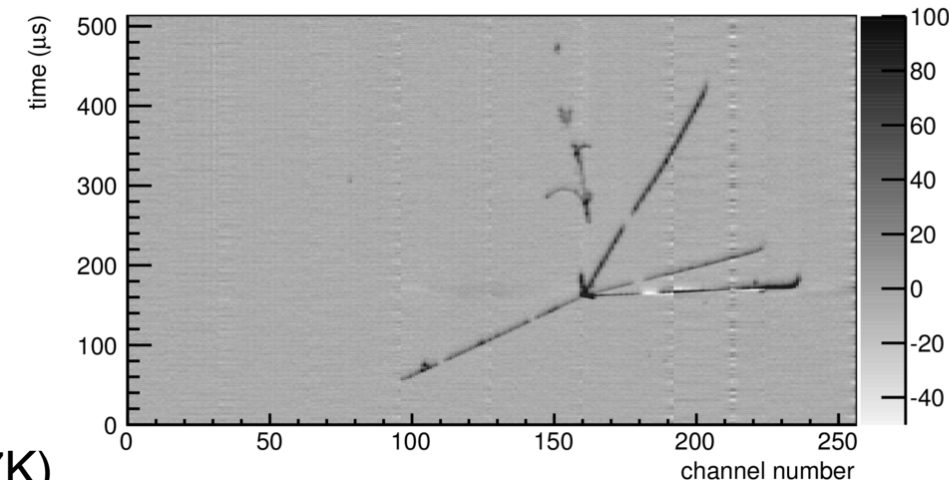
For example:

- Proton decay
- Atmospheric neutrinos: e.g. detection of $\nu\tau$
- Supernova core collapse neutrinos
- Diffuse supernova neutrino background
- Indirect DM detection
- Long baseline neutrino oscillation for CPV & MH

JHEP 0704 (2007) 041
 JCAP 0408 (2004) 001
 JCAP 0412 (2004) 002
 arXiv:1105.4077 [hep-ph]
 Nucl. Phys. B 589 (2000) 577

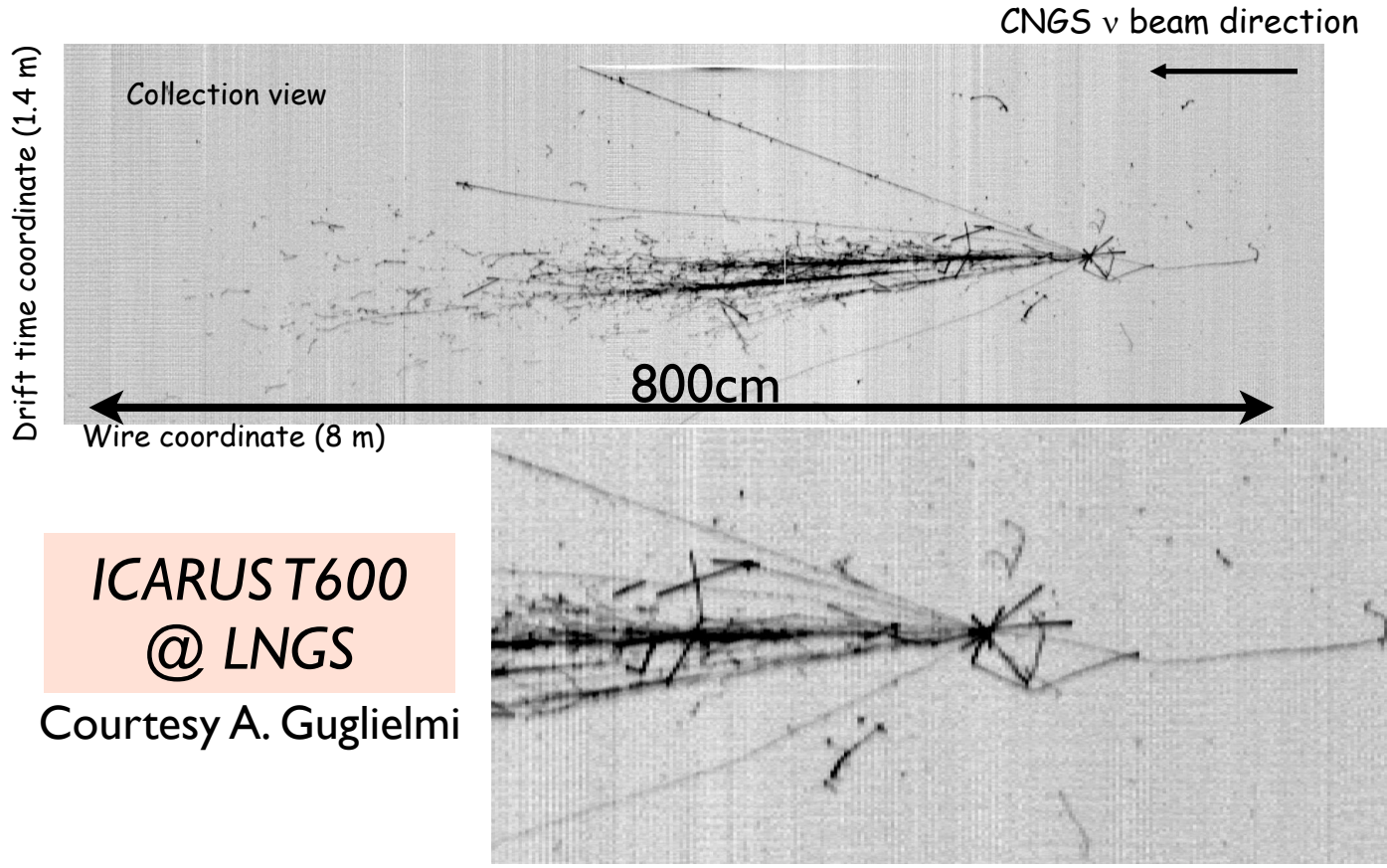
PRD82, 093012 (2010)
 Nucl. Phys. B. Proc. Suppl. 91, 223 (2001)
 arXiv:1003.1921 [hep-ph]
 arXiv:0804.2111 [hep-ph]
 arXiv:0801.4035 [hep-ph]
 JHEP 0611 (2006) 032
 + etc...

Cosmic interaction in double phase $40\times 80\text{cm}^2$ LAr-LEM TPC with adjustable gain @ CERN-ETHZ

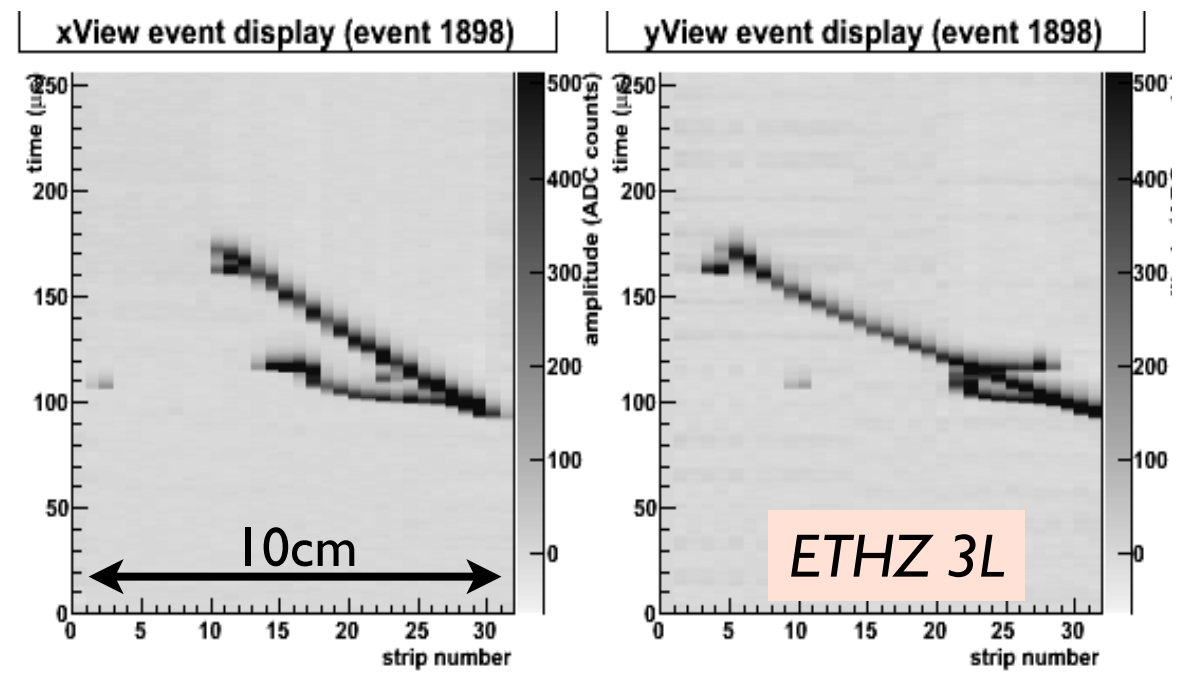


The “electronic bubble chamber”

Cosmic track in double phase 3L LAr-LEM TPC with adjustable gain @ CERN-ETHZ

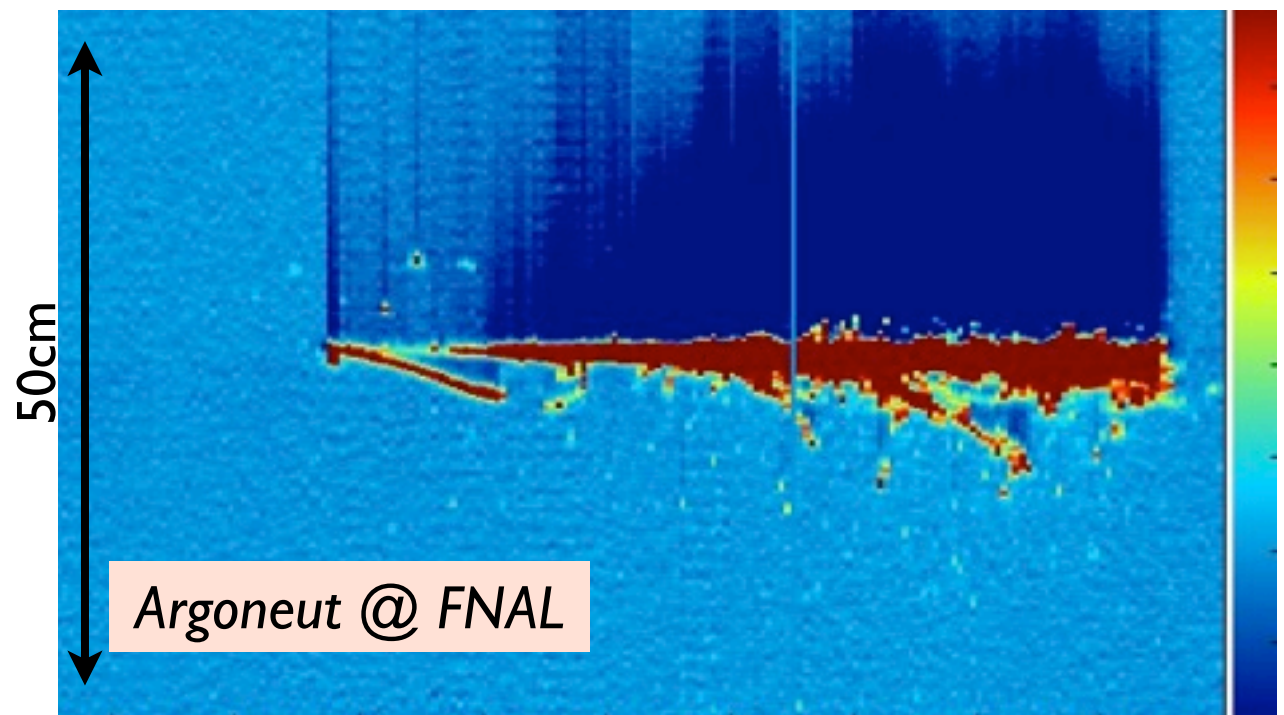


ICARUS T600
@ LNGS
Courtesy A. Guglielmi

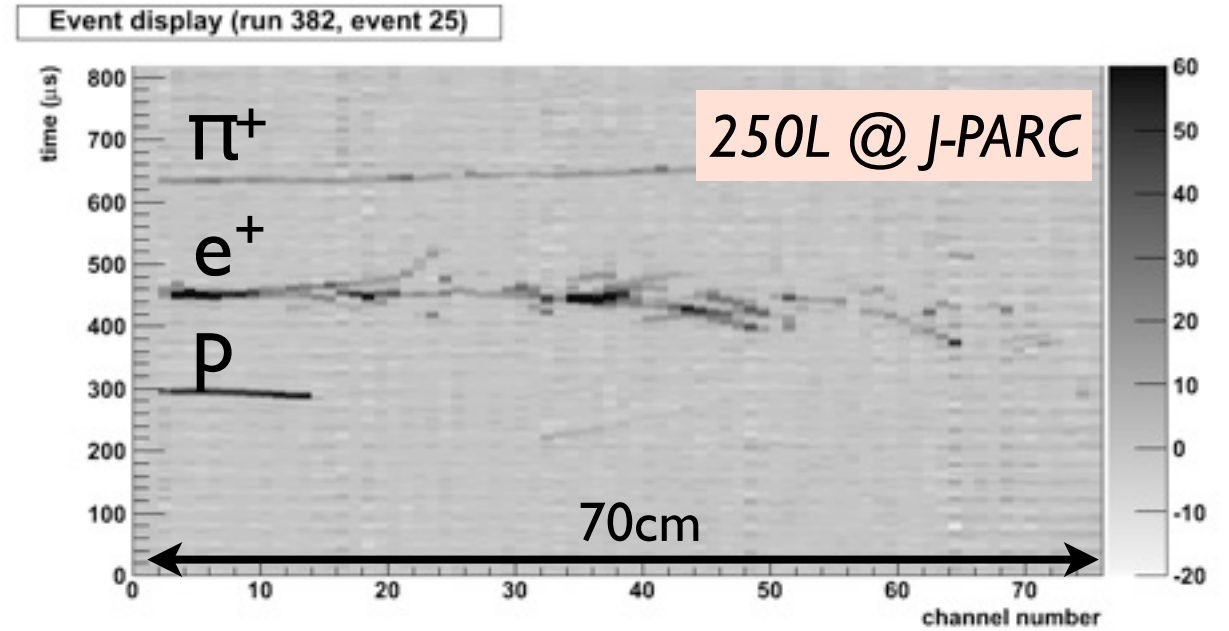


Much improved S/N (>100) compared to single-phase LAr operation (~15)

Charged particle beam ≈ 800 MeV/c exposure



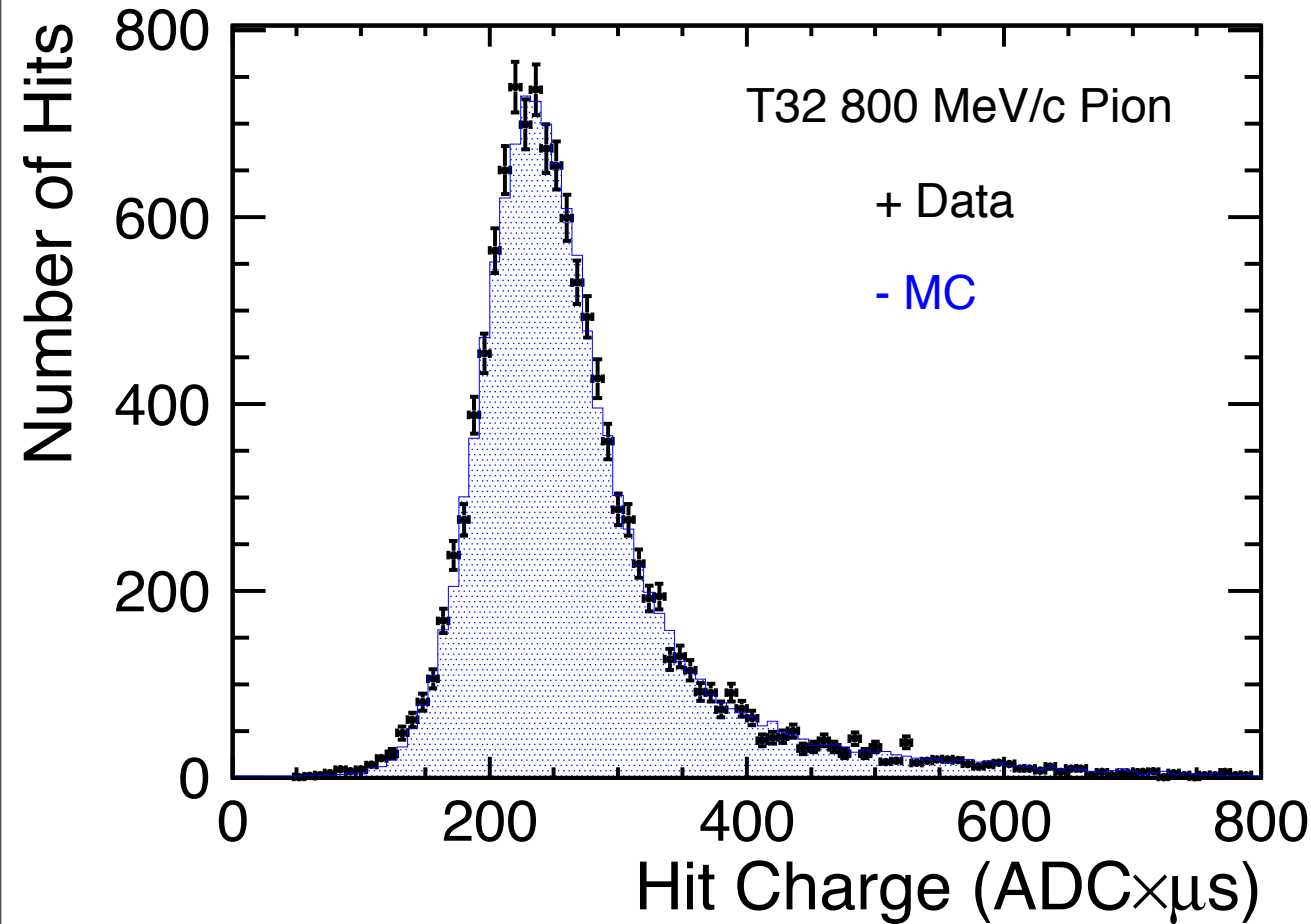
Courtesy B. Fleming



Tracking performance

JPARC T32 exposed to KI.IBR tagged beam

J.Phys.Conf.Ser. 308 (2011) 012008



Data well described by:

$$Q = A \frac{Q_0}{1 + (k / \epsilon) \times (dE/dx) \times (1/\rho)}$$

Observable charge

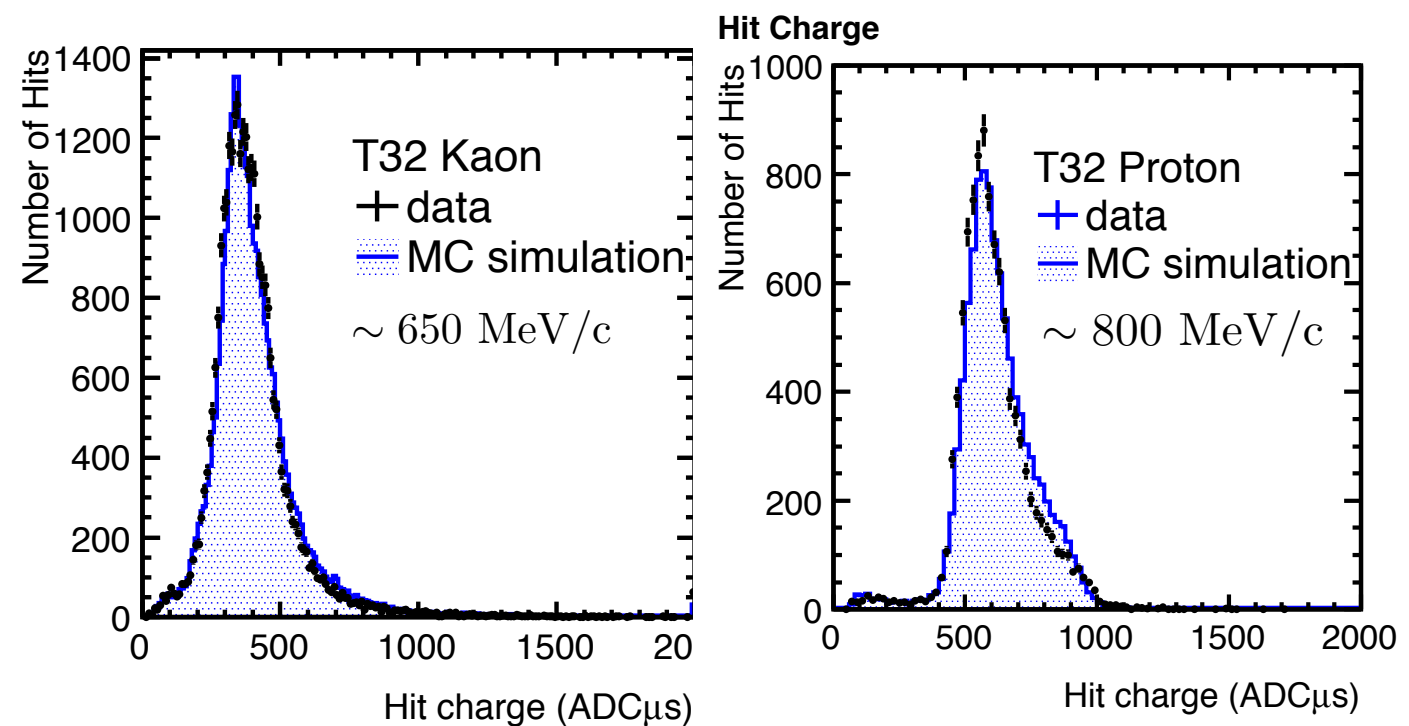
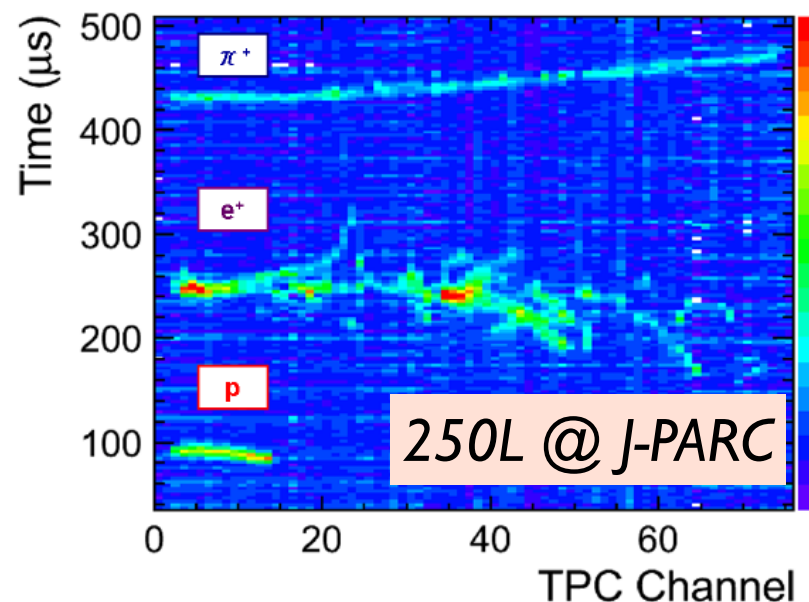
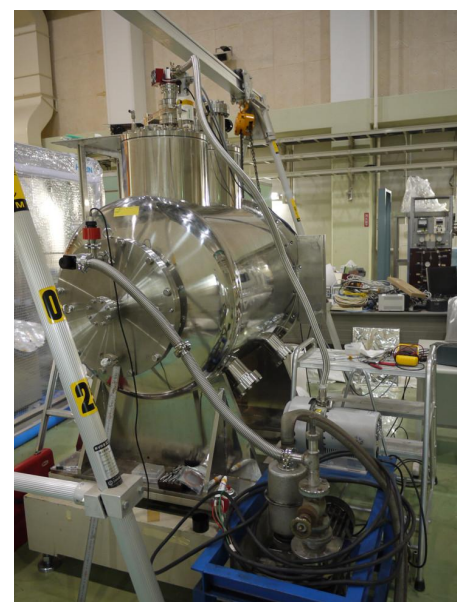
Raw charge

$$A = 0.8$$

$$k = 0.0486 \text{ kV/cm} \frac{\text{g/cm}^2}{\text{MeV}}$$

NIM A 523, 275 (2004)

J-PARC T32 chamber (ETHZ-KEK-Iwate-Waseda)



Good understand of tracking

Courtesy T. Maruyama

Calorimetric performance

Michel electrons form
stopping muon decay sample

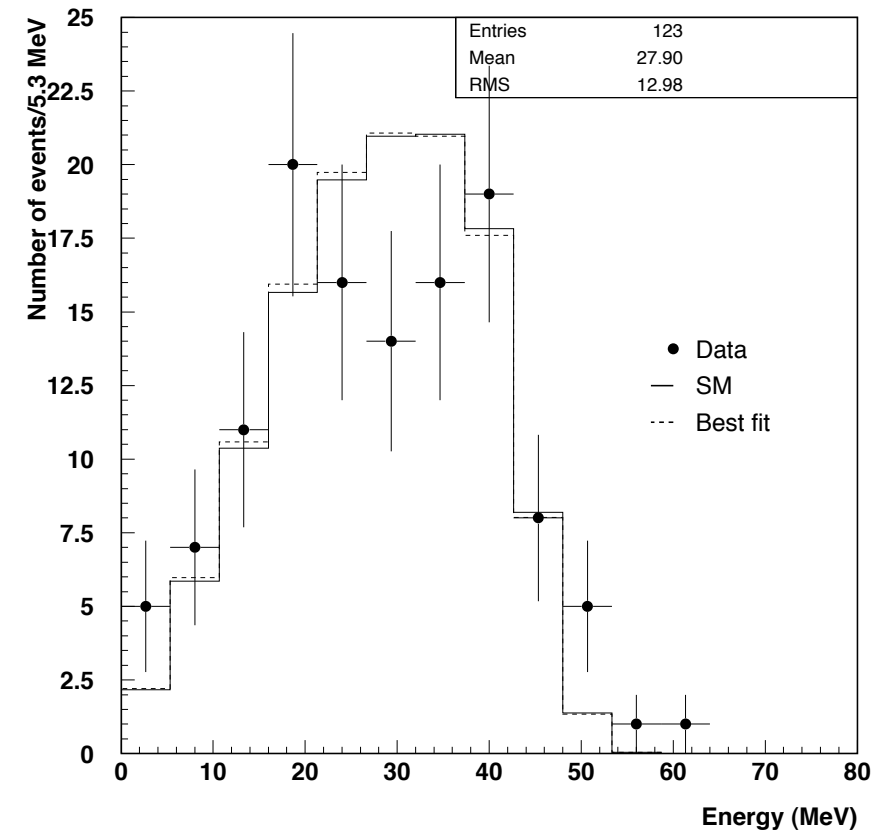
$$\frac{\sigma_e}{E} \simeq \frac{11\%}{\sqrt{E(\text{MeV})}} \oplus 4\%$$

MC simulations at
higher energies:

$$\frac{\sigma_{em}^{MC}}{E} \simeq \frac{3\%}{\sqrt{E}} \oplus 1\%$$

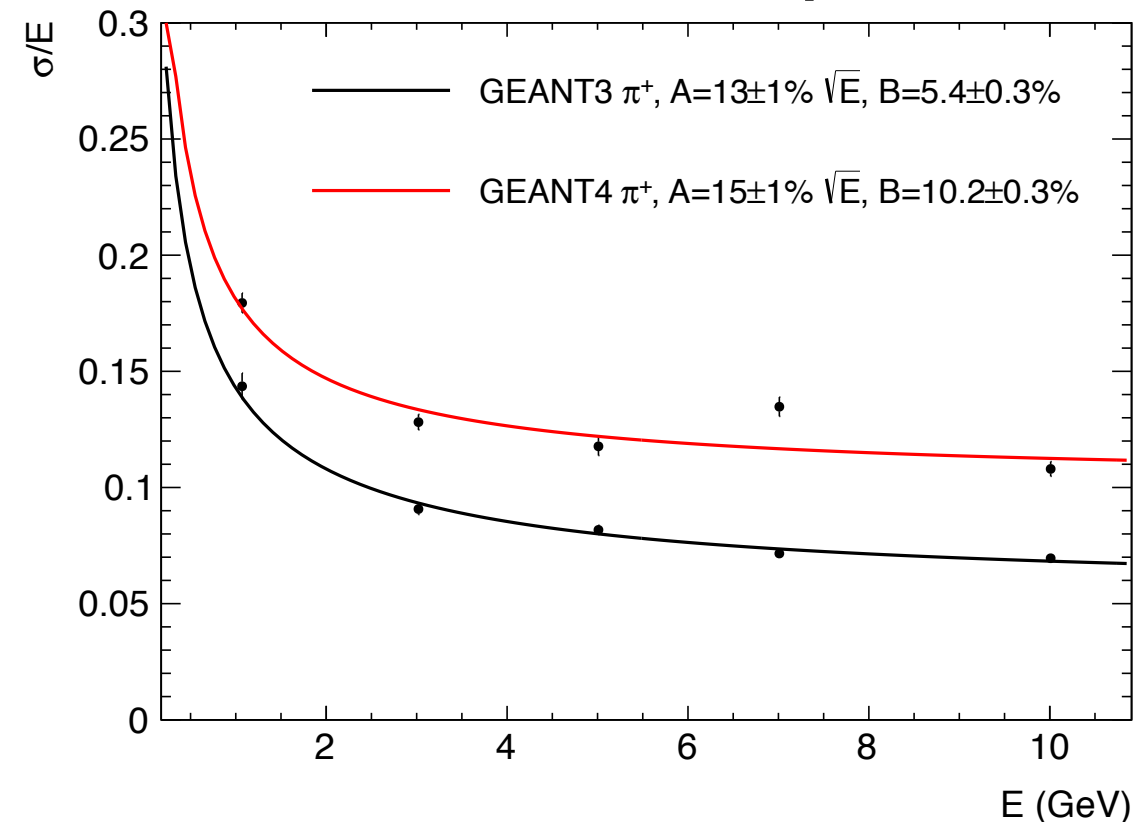
$$\frac{\sigma_{had}^{MC}}{E} \simeq \frac{15\%}{\sqrt{E}} \oplus 10\%$$

↑
needs to be confirmed
by experimental data



Eur. Phys. J. C33, 233 (2004)

G3 and G4 comparison



Purity and vessel evacuation

★ Several independent groups performed numerical simulations and concluded that the vacuum evacuation phase could be avoided for larger detectors:

- more favorable surface / volume ratio for large volume (also larger volumes are less sensitive to micro leaks !!)
- initial purity of argon when delivered is typ. $O(1)$ ppmv $O_2 \rightarrow$ purification from ppm to $\ll 1$ ppb anyhow needed
- outgassing of material from hot components, impurities “frozen” at low temperature

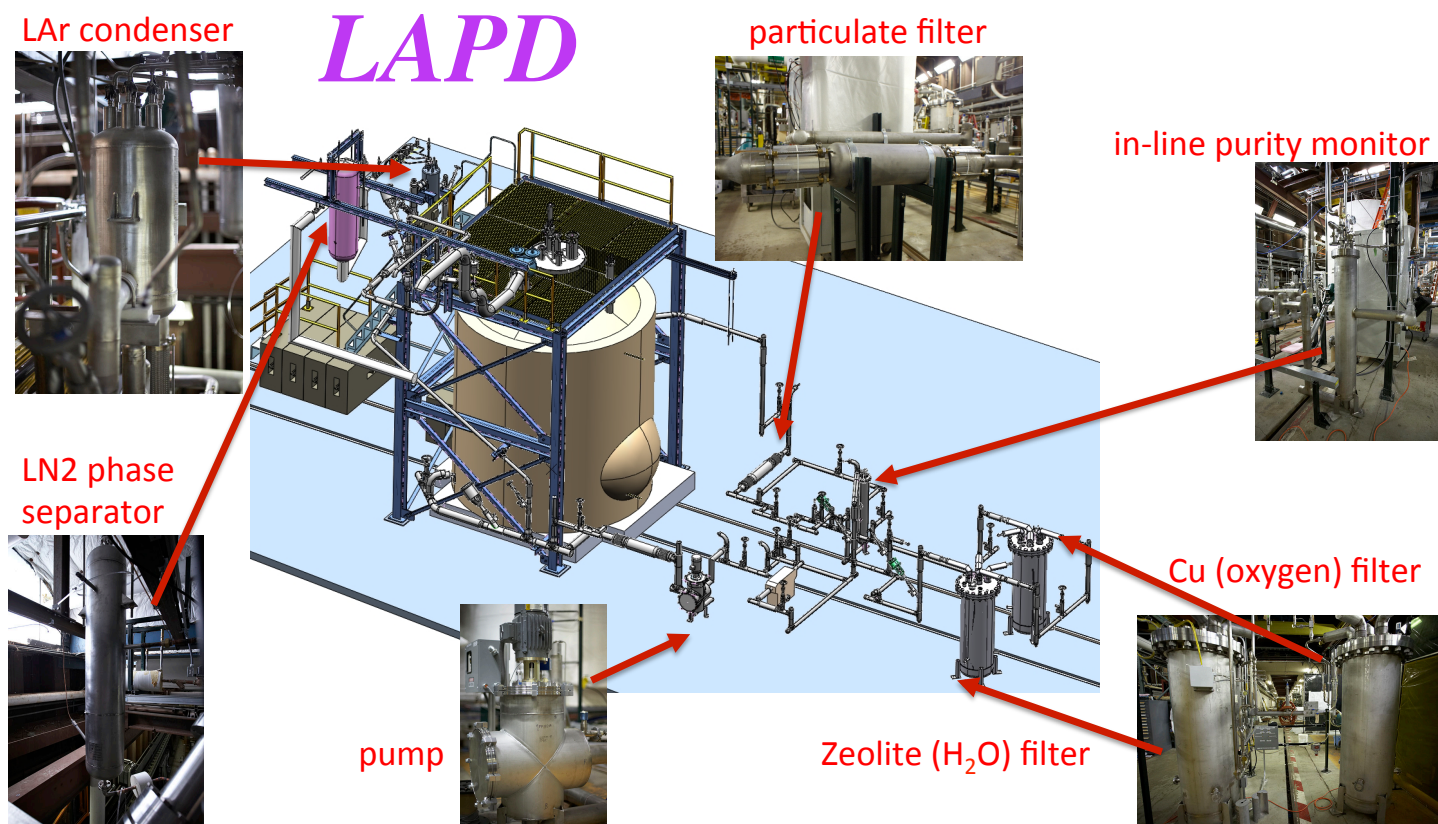
★ GAr flushing and purging are effective ways to remove air and impurities.

★ Purging on 6m3 volume (ETHZ-KEK-Liverpool @ CERN)

- Piston effect seen in gas and reached 3ppm O_2 after several volumes exchange (J.Phys.Conf.Ser. 308 (2011) 012024)

★ LAPD @ FNAL – Liquid Argon Purity Demonstrator – First test in Liquid Phase !

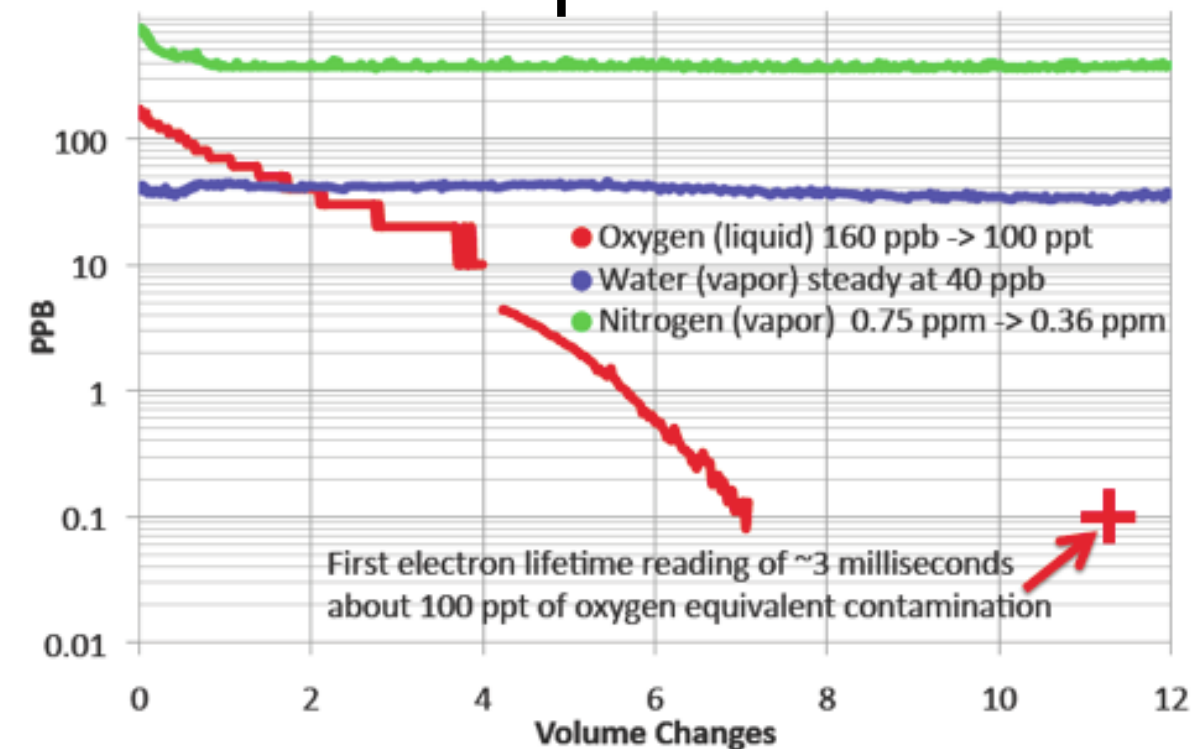
- Tank size: 30 ton LAr (25,000 liters)
- Milestone successfully reached!! it is possible to obtain a **better than 3 ms electron lifetime** in a large non-evacuated vessel !



A. Rubbia

Future liquid Argon detectors (Neutrino 2012)

Meas'd impurities in LAr



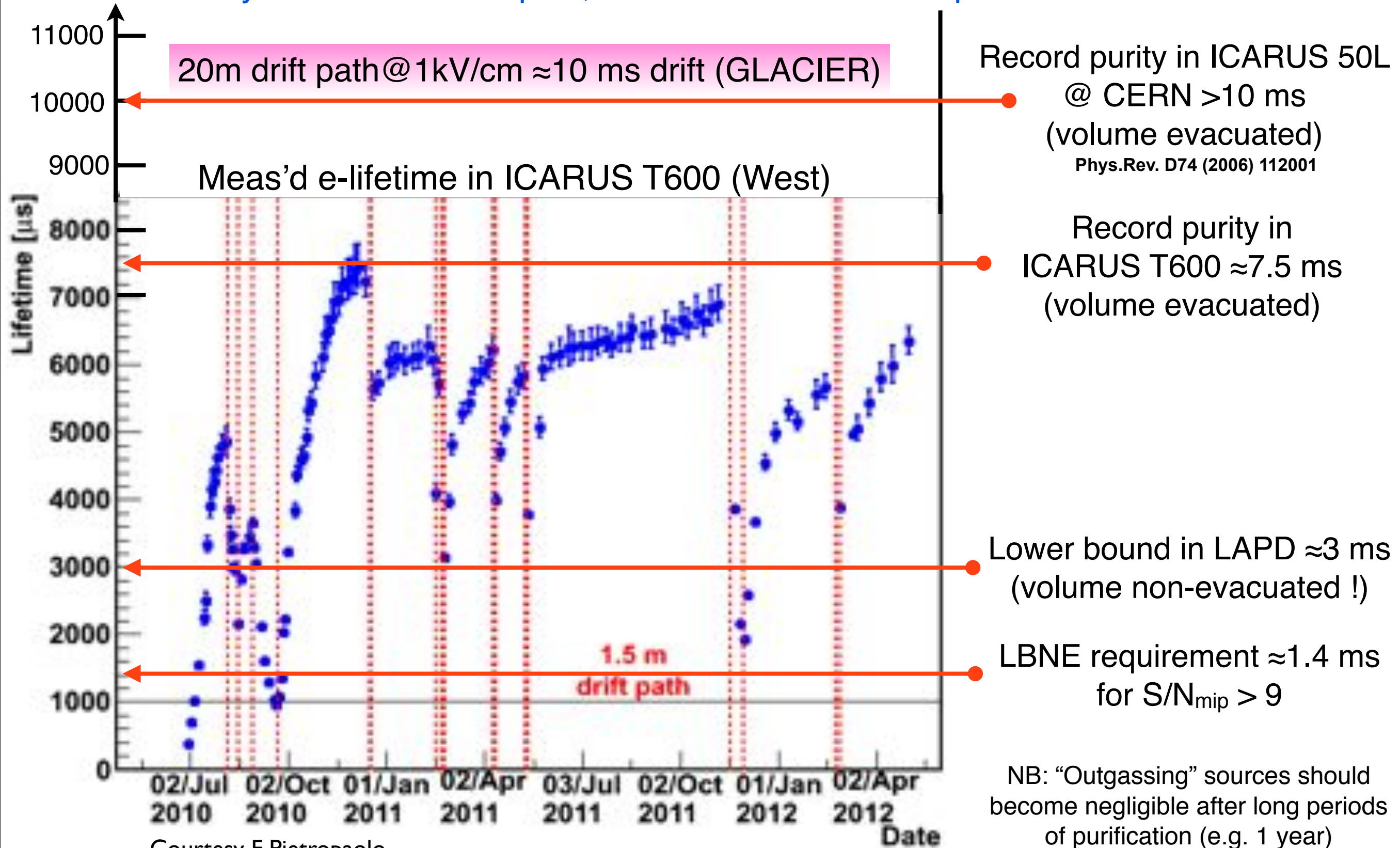
Courtesy B. Baller & B. Rebel

ETH

78

Purity and evacuation

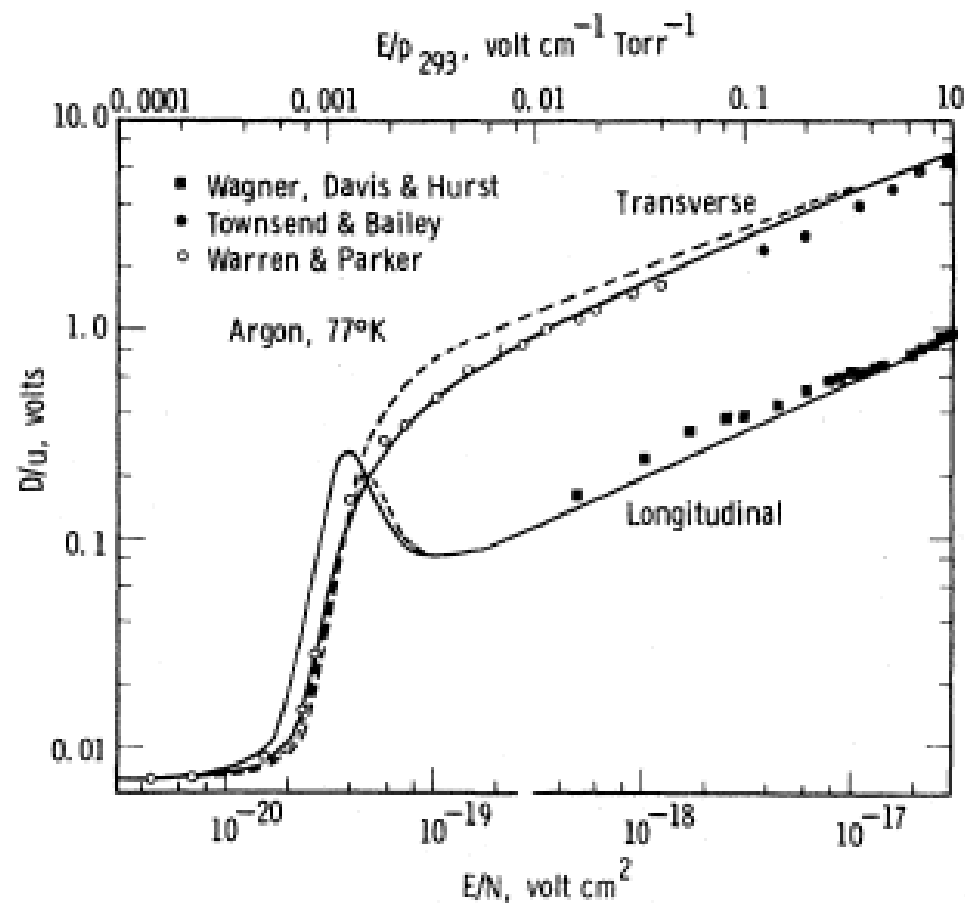
- ★ Excellent purity has been reproducibly achieved in various setups always relying on commercially available techniques, of various sizes and capacities.



Courtesy F. Pietropaolo

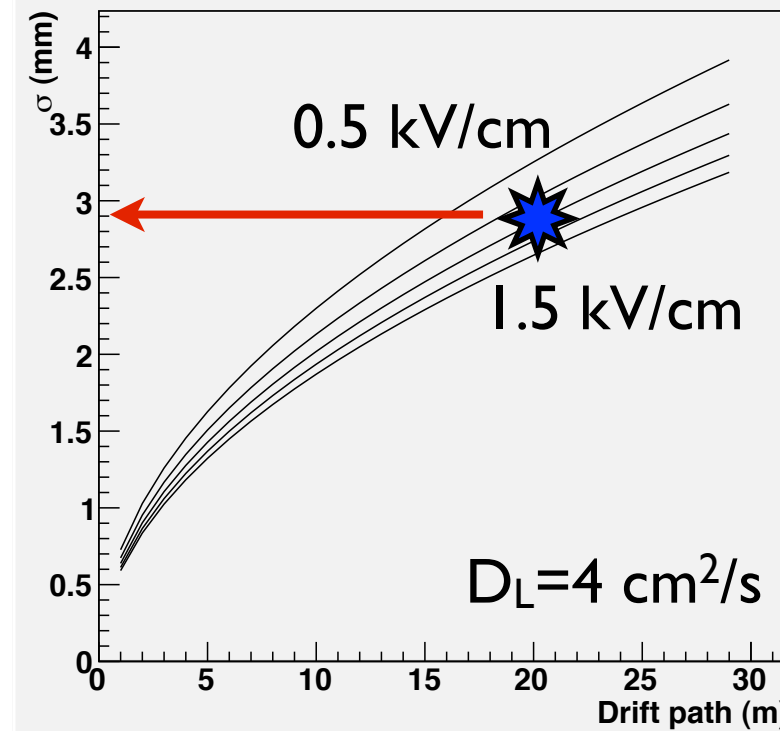
Electron cloud diffusion

★ The physical limit to long drifts is determined by diffusion → likely 20m !

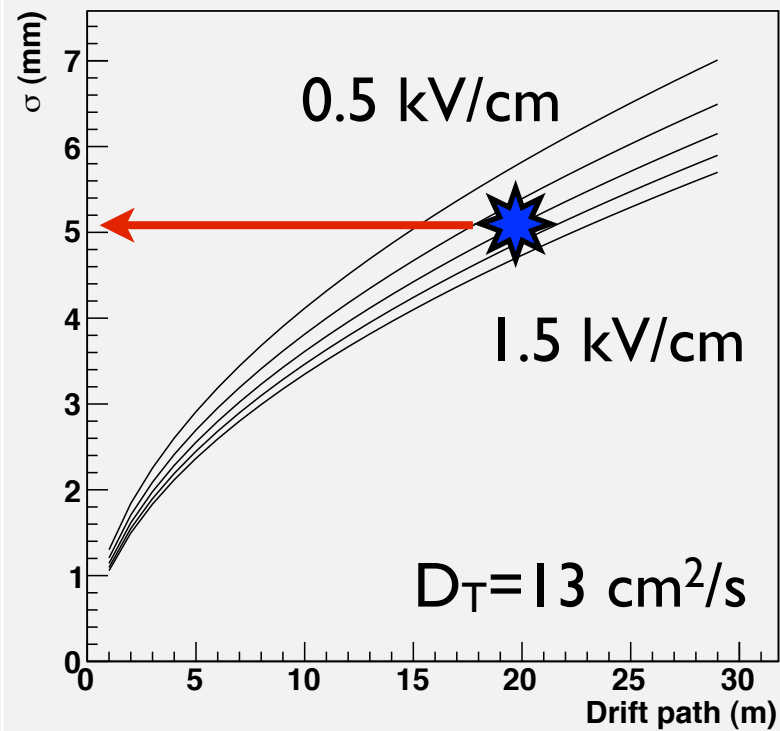


Drift fields $E=0.5, 0.75, 1, 1.25, 1.5$ kV/cm

Longitudinal Diffusion



Transverse Diffusion



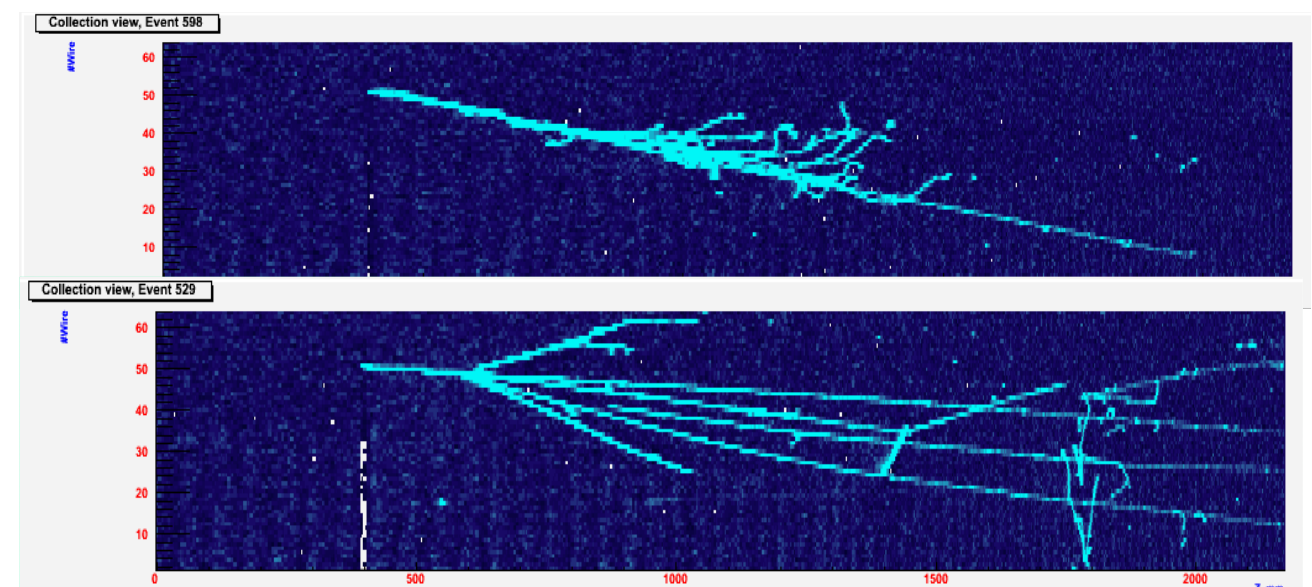
★ Diffusion coefficients not well known (in particular for transverse diff.):

- after 20 m drift: transverse diffusion \approx **5mm**, longitudinal diffusion \approx **3mm**

★ New measurements:

- ArgonTube (Bern University)
 - tracks >4 m length observed !
 - lifetime ≈ 2 ms after 24hrs
- 5m drift (UCLA)

Courtesy I. Kreslo



Future LAr TPC detectors

Project	LAr mass (tons)	Goal	Baseline (km)	Where	Status
MicroBOONE	170 (70 fid.)	short baseline	0.47	FNAL BNB	Under construction
LAr1	≈1'000	2 nd detector for short baseline	≈0.7	FNAL BNB	Proposal submitted
ICARUS-NESSIE	150 + 478	two-detectors short baseline	0.3 + 1.6	CERN + new SBL beam	Proposal submitted
MODULAR	5'000 unit	shallow depth far detector	730	Italy, new lab nearby LNGS	plan
GLADE	5000	surface	810	NUMI off-axis	Letter of Intent
LBNE LAr (*)	2x17'000(*)	underground(*) far detector	1300(*)	Homestake(*) + new FNAL beam(*)	CD-0
GLACIER LAGUNA-LBNO	initially 20'000 (incremental)	underground far detector	2300	Finland + new CERN LBL beam	Expression of interest in preparation
GLACIER Okinoshima	up to 100'000	underground far detector	665	Japan + JPARC neutrino beam	R&D proposal at JPARC

() LBNE reconfiguration for cost reduction / staging in progress (cf. Svoboda's talk)*

GLACIER detector design

★ Concept unchanged since 2003: Simple, scalable detector design, from one up to 100 kton (hep-ph/0402110)

★ Single module non-evacuatable cryo-tank based on industrial LNG technology

- industrial conceptual design (Technodyne, AAE, Ryhal engineering, TGE, GTT)
- two tank options: 9% Ni-steel or membrane (detailed comparison up to costing of assembly in underground cavern)
- three volumes: 20, 50 and 100 kton

★ Liquid filling, purification, and boiloff recondensation

- industrial conceptual design for liquid argon process (Sofregaz), 70kW total cooling power @ 87 K
- purity < 10 ppt O₂ equivalent

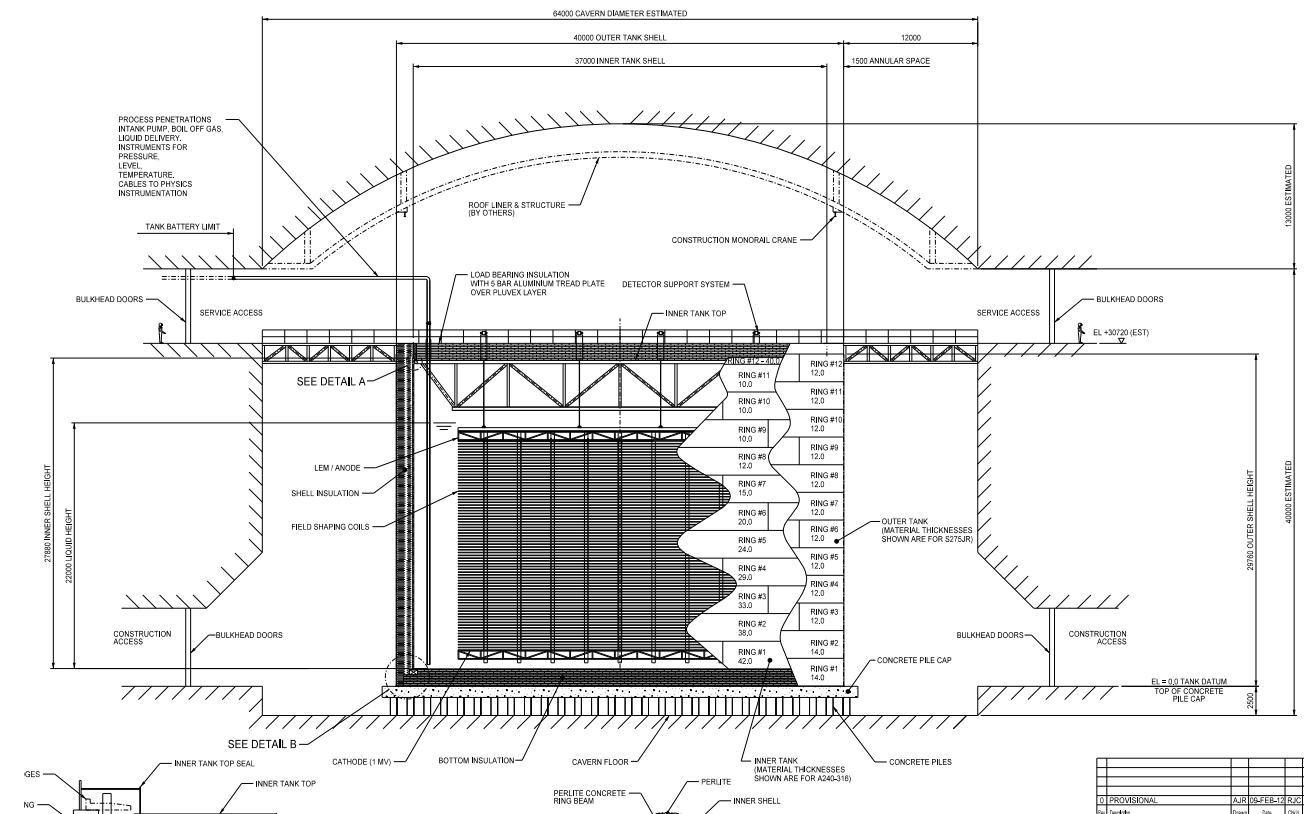
★ Charge readout (e.g. 20 kton fid.)

- 23'072 kton active, 824 m² active area
- 844 readout planes, 277'056 channels total
- 20 m drift

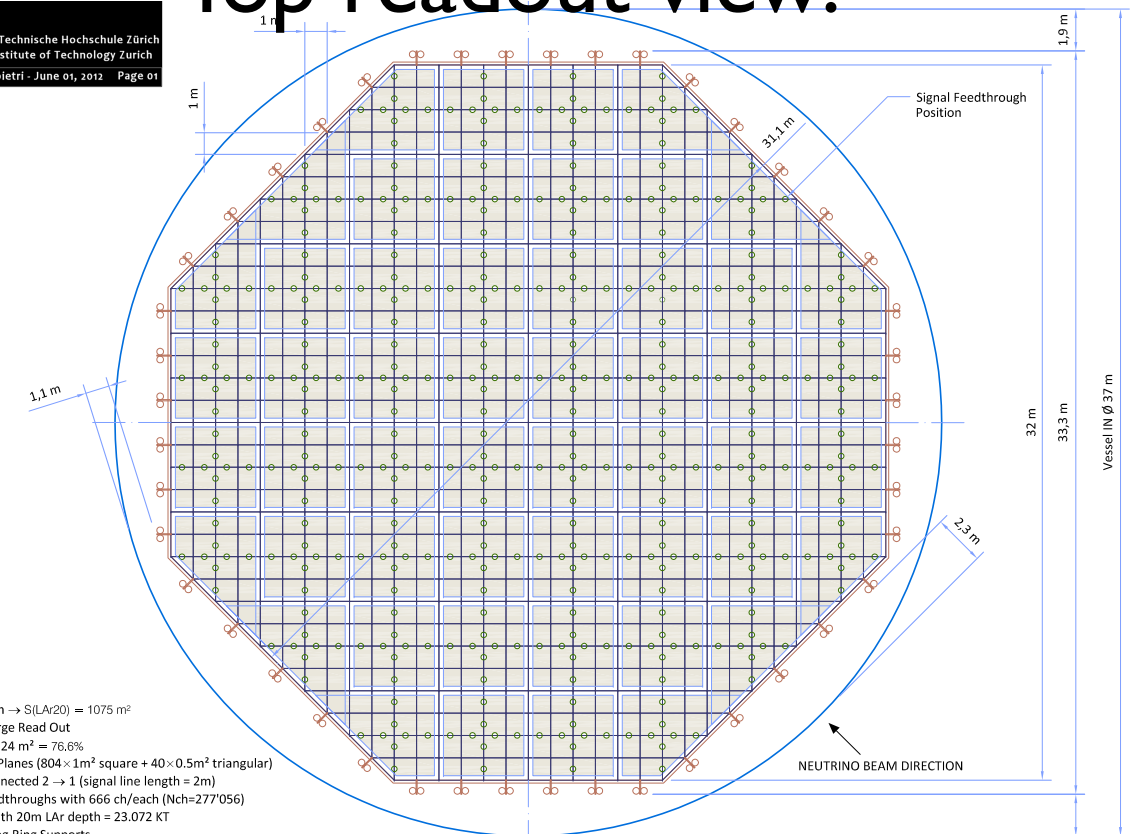
★ Light readout (trigger)

- 804 8" PMT (e.g. Hamamatsu R5912-02MOD) WLS coated placed below cathode

★ The concept and the designs are reaching the required level of maturity for submission to SPSC.

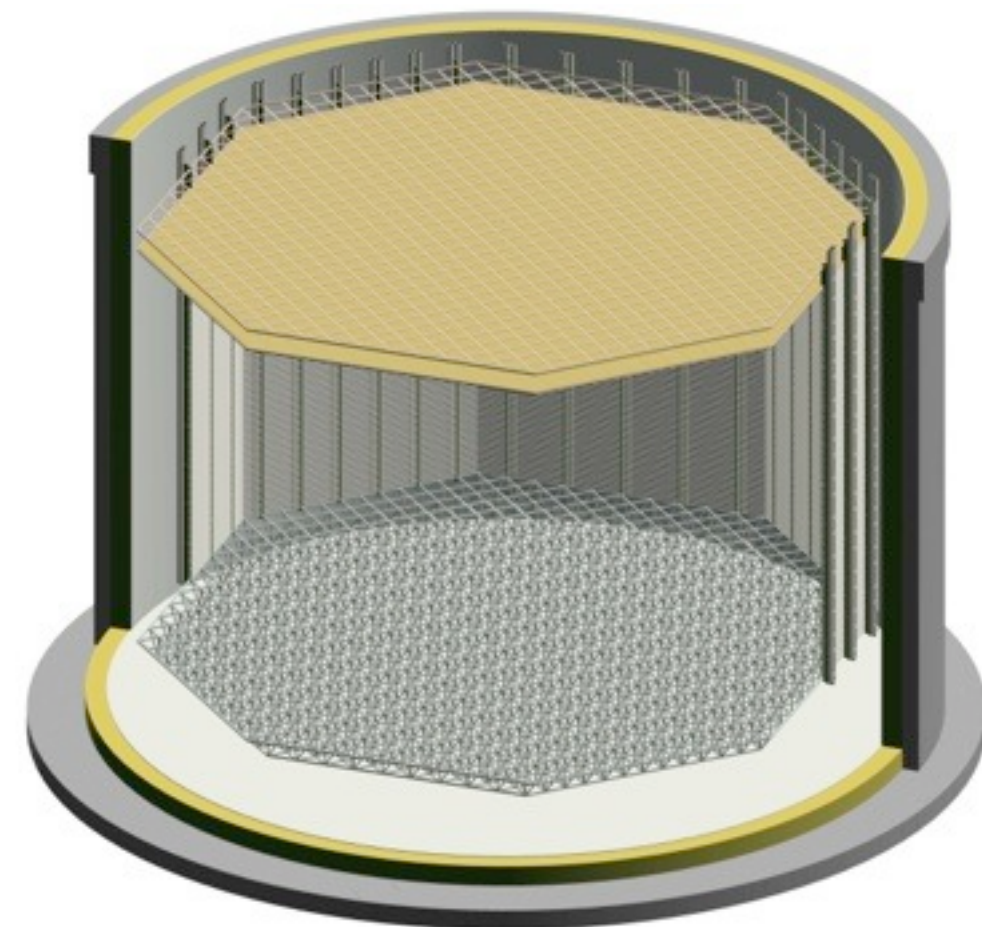
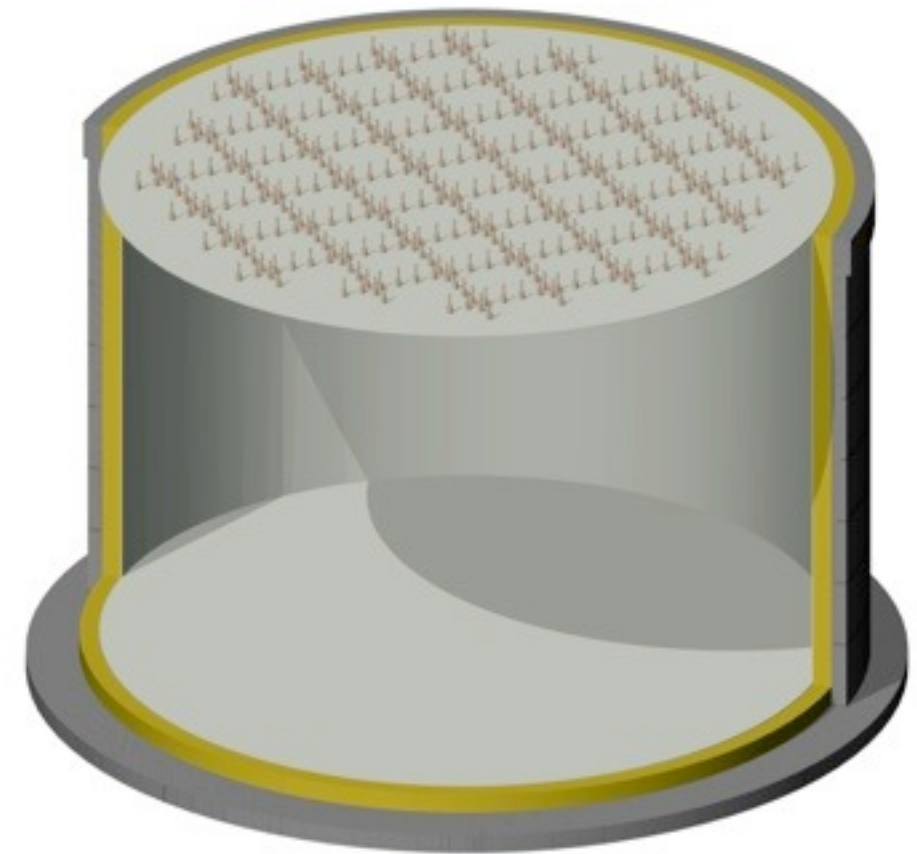


Top readout view:



GLACIER detector parameters

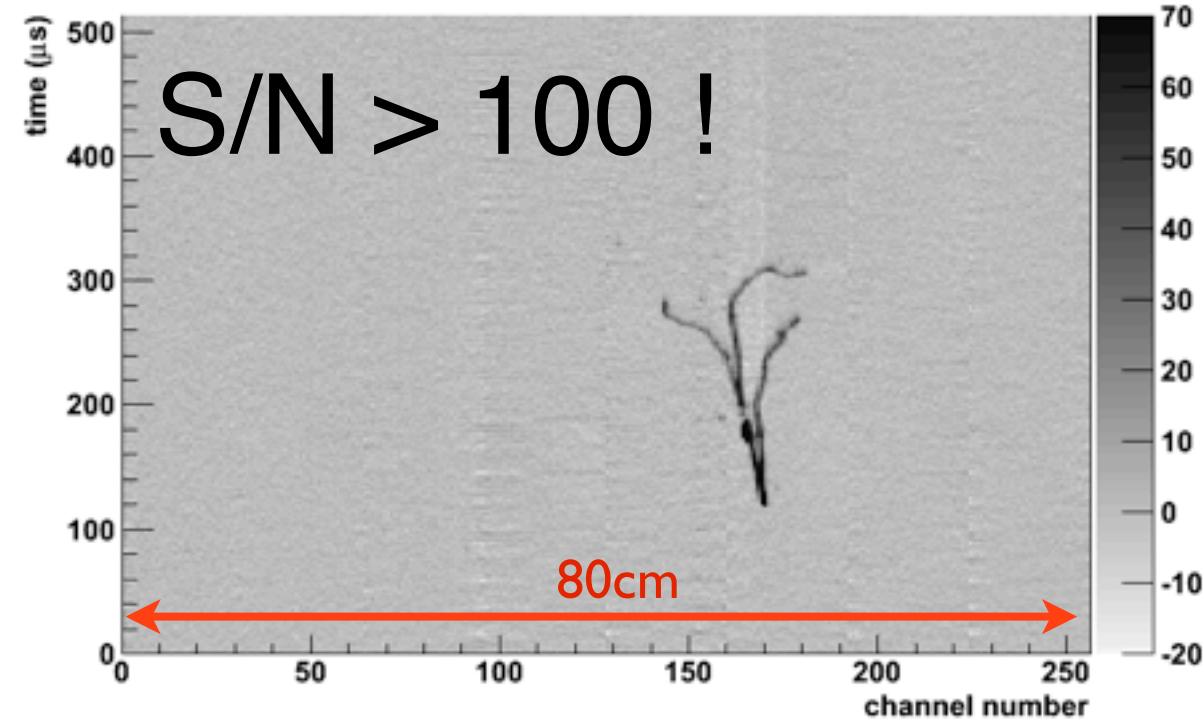
		20 KT	50 KT	100 KT
Liquid argon density at 1.2 bar	[T/ m ³]	1.38346		
Liquid argon volume height	[m]	22		
Active liquid argon height	[m]	20		
Pressure on the bottom due to LAr	[T/ m ²]	30.4 ($\equiv 0.3$ MPa $\equiv 3$ bar)		
Inner vessel diameter	[m]	37	55	76
Inner vessel base surface	[m ²]	1075.2	2375.8	4536.5
Liquid argon volume	[m ³]	23654.6	52268.2	99802.1
Total liquid argon mass	[T]	32525.6	71869.8	137229.9
Active LAr area (percentage)	[m ²]	824 (76.6%)	1854 (78%)	3634 (80.1%)
Active (instrumented) mass	[KT]	22.799	51.299	100.550
Charge readout square panels (1m×1m)		804	1824	3596
Charge readout triangular panels (1m×1m)		40	60	72
Number of signal feedthroughs (666 channels/FT)		416	1028	1872
Number of readout channels		277056	660672	1246752
Number of PMT (area for 1 PMT)		804 (1m×1m)	1288 (1.2m×1.2m)	909 (2m×2m)
Number of field shaping electrode supports (with suspension SS ropes linked to the outer deck)		44	64	92



GLACIER charge readout

- A. Badertscher, et al., NIM A 641 (2011) 48-57
- See also arXiv:1204.3530 [physics.ins-det]

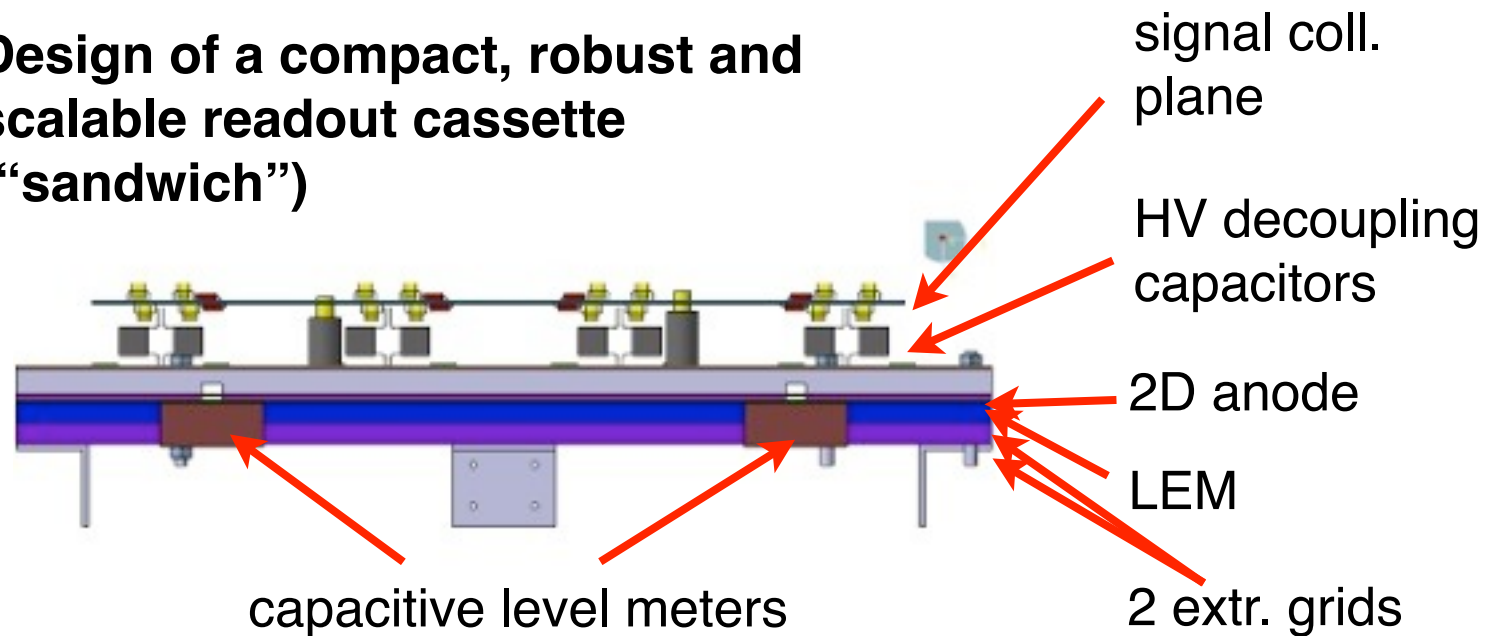
Cosmic Data from 40x80cm² LAr LEM TPC@CERN-ETHZ



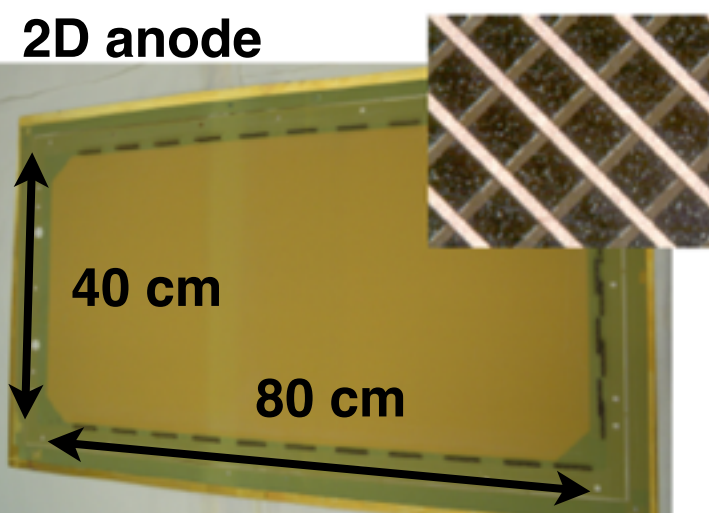
★ Novel double phase LAr LEM-TPC readout:

- ionization electrons are drifted to the liquid-gas interface
- if the E-field is high enough (≈ 3 kV/cm) they can efficiently be extracted to the gas phase
- in the holes of the LEM the E-field is high enough to trigger an electron avalanche
- the multiplied charge is collected on a 2D readout
- gain allows **sharing charge in collection mode for both views!!**

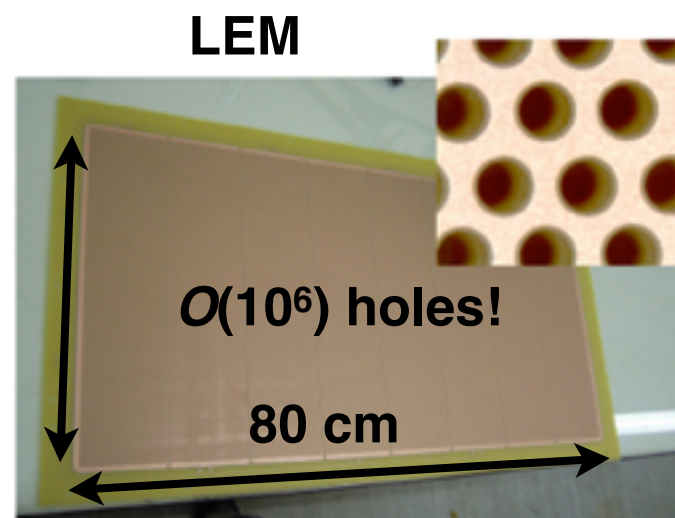
Design of a compact, robust and scalable readout cassette ("sandwich")



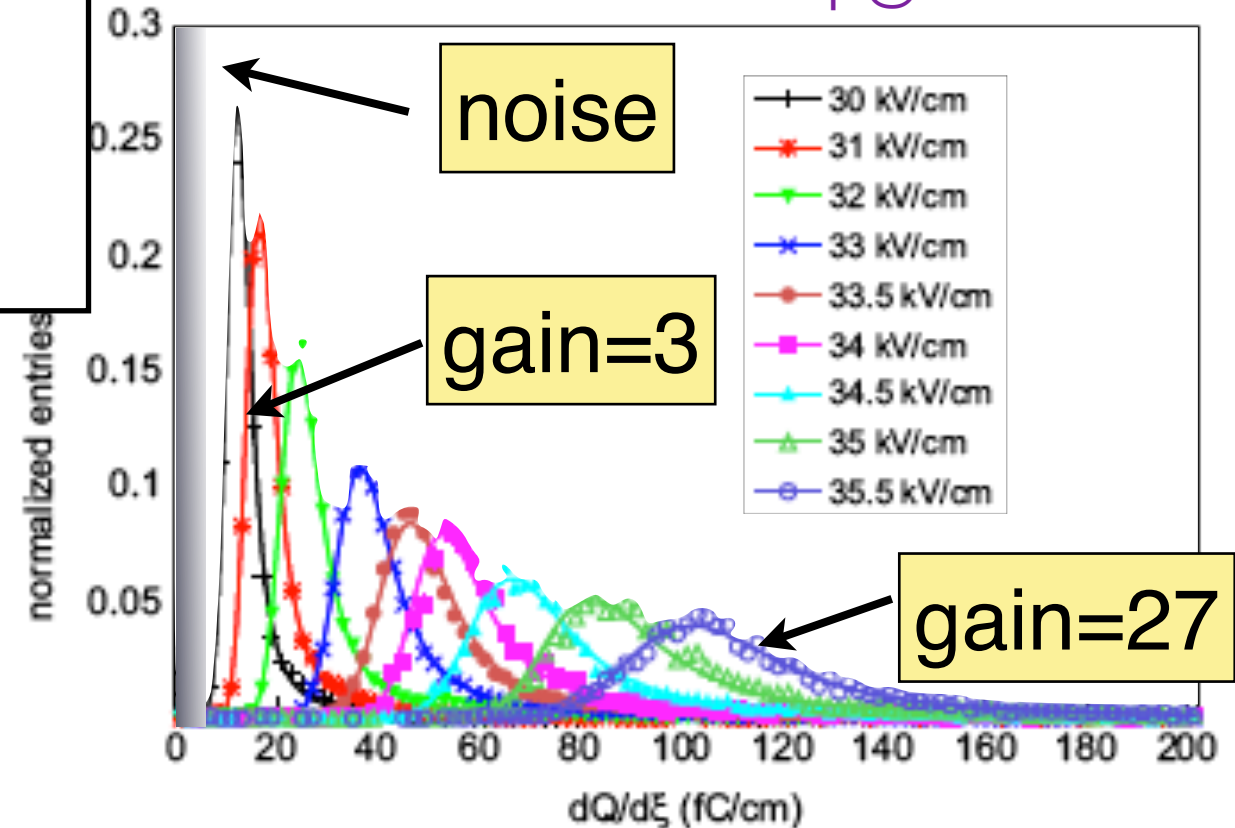
2D anode



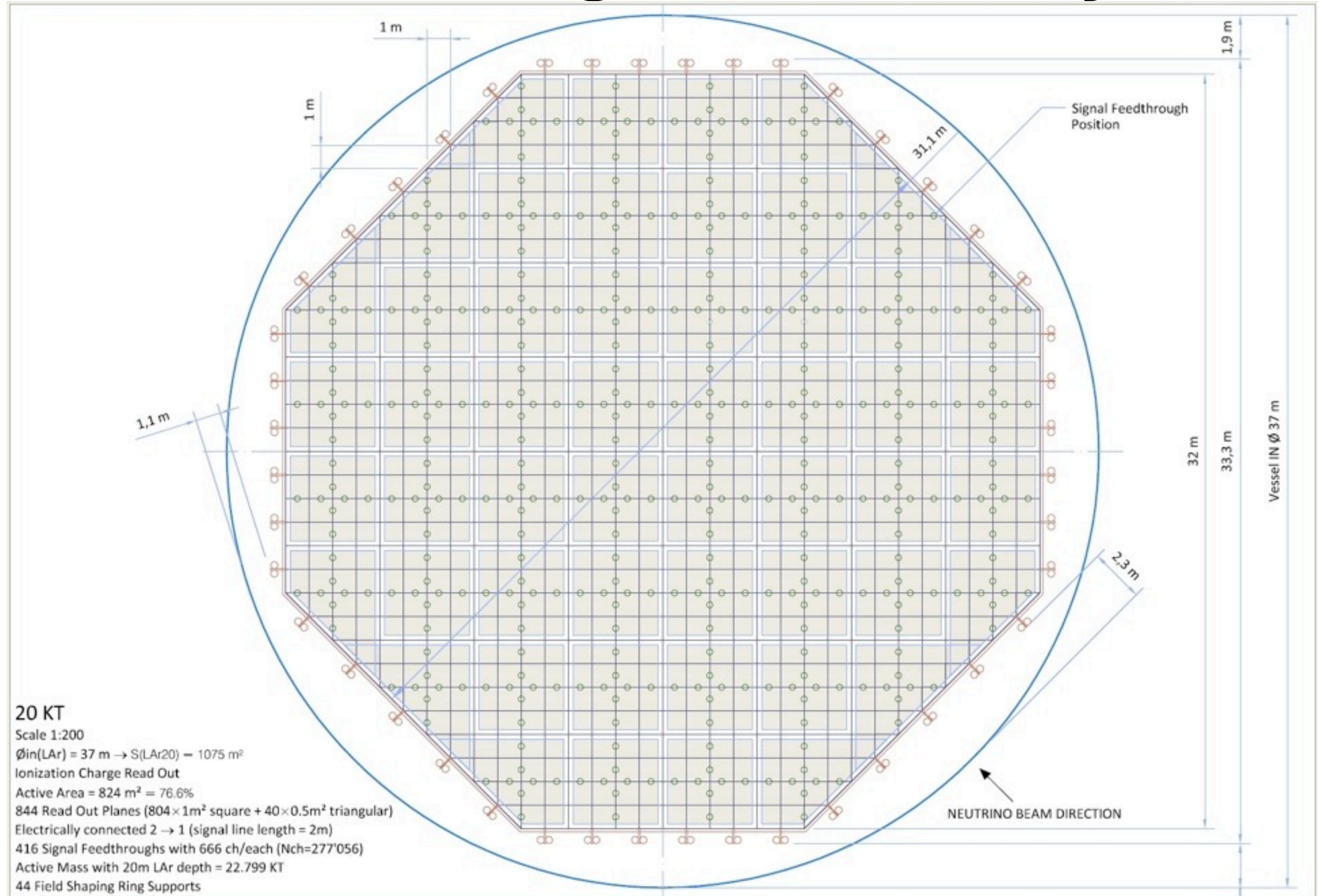
LEM



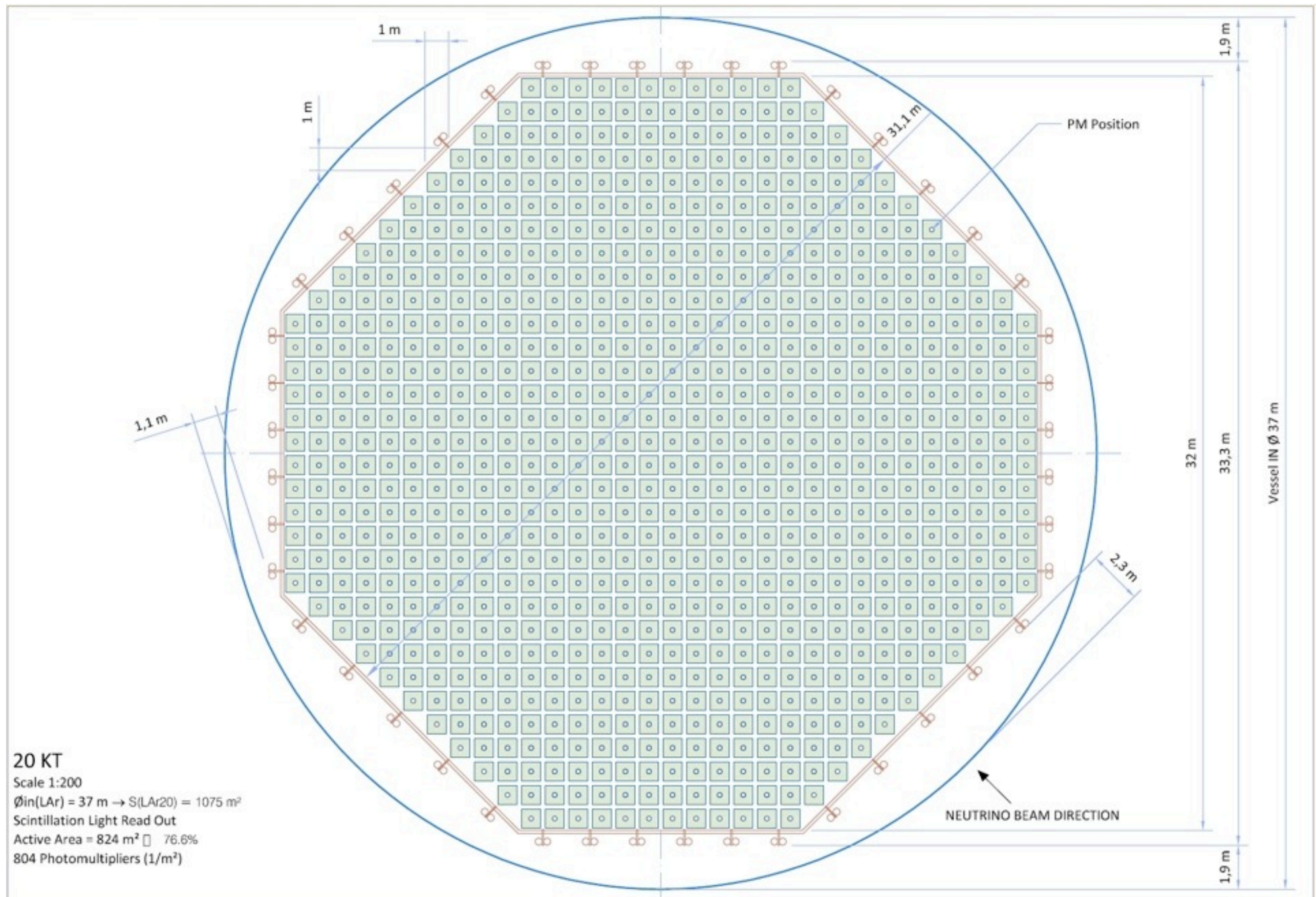
Landau distribution fitted to dE/dx distributions of muons on 3L LAr LEM-TPC setup @ CERN-ETHZ



GLACIER charge readout layout

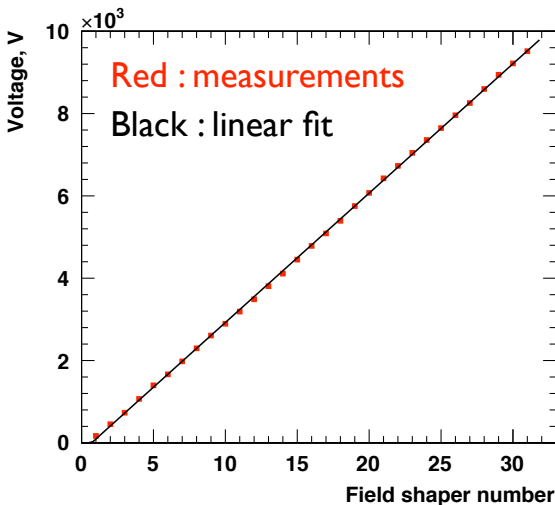
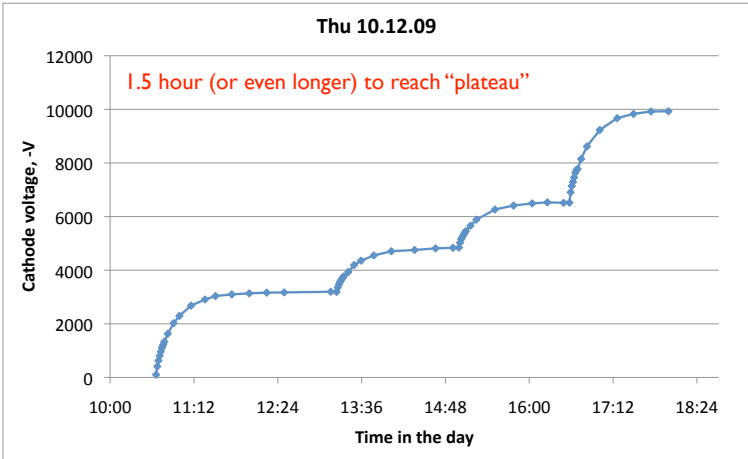


GLACIER light readout layout



Drift high voltage multiplier

J.Phys.Conf.Ser. 308 (2011) 012027
arXiv:1204.3530 [physics.ins-det]



Extrapolation to long drift

Extrapolation of the ArDM design

Changing Cs for fixed Cp = 2.35 pF and Vpp-in = 2E = 2.5 kV

ArDM

Drift length	m	1.24	5	10
Total output voltage for 1 kV/cm	V	124k	500k	1M
Input voltage Vpp-in = 2E	V	820	2.5k	2.5k
Shunt capacitance, Cp	F	2.35p	2.35p	2.35p
Capacitor	F	328/164n	475n	1.90μ
Number of stages, N	–	210	319	638
N per 10 cm	–	16.9	6.38	6.38
Total capacitance	F	125μ	303μ	2.43m
Capacitance per 10 cm	F	10.4μ	5.99μ	24.3μ
Total stored energy	J	21.7	948	7.58k

$\times \sqrt{2}$
 $\times 1/2$



20
2M
3.5k
1.18p
1.90μ
903
4.5l
3.43m
17.2μ
21.5k

Actual ArDM parameters are given just for comparison.

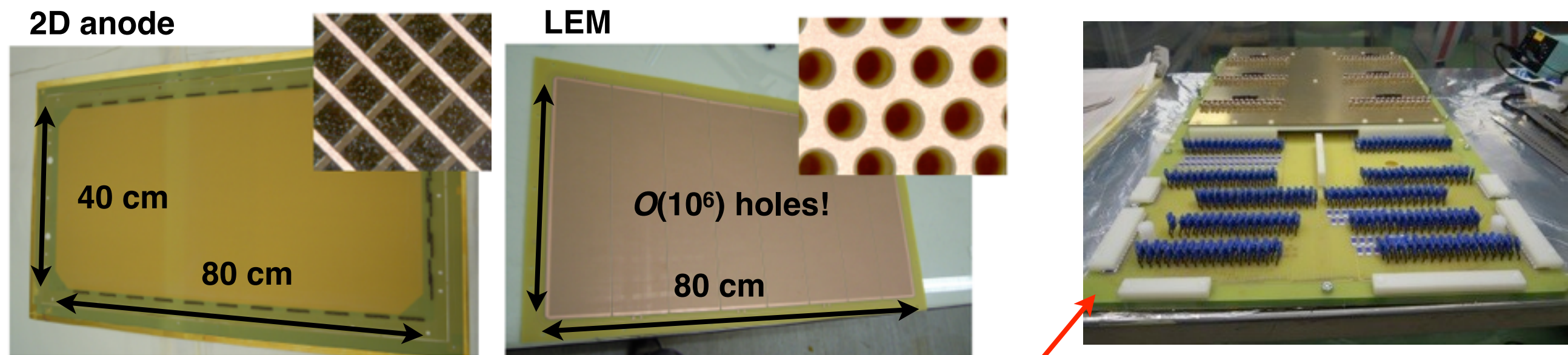
For extrapolation, $2\gamma N = 1.42$ is always assumed.

LAr vaporization heat 160 kJ/kg

$$V_{\max} = \frac{E}{\gamma}, \quad \gamma \approx \sqrt{\frac{C_p}{C_s}}$$

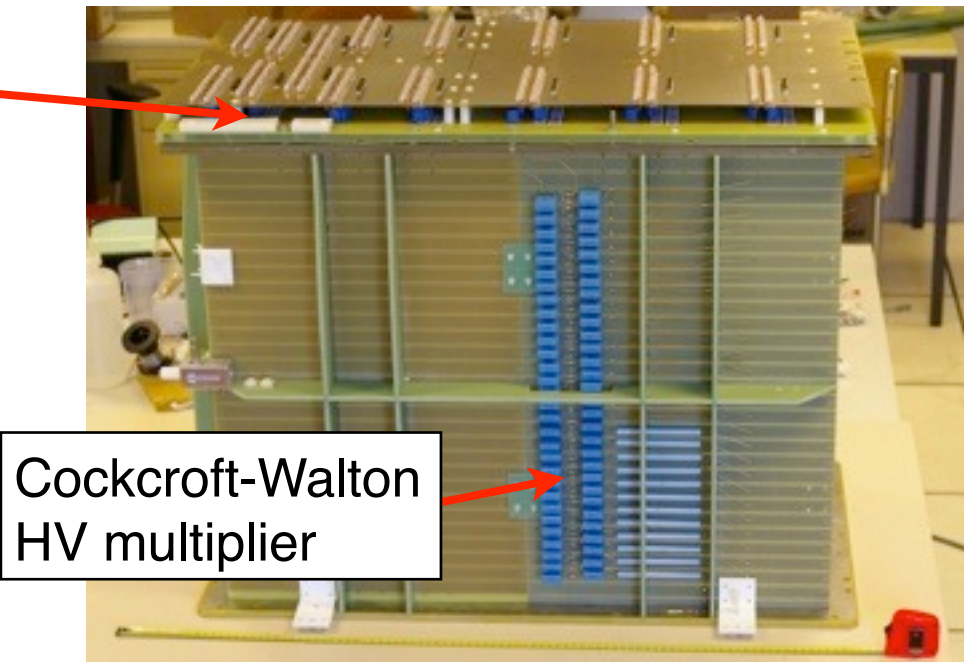
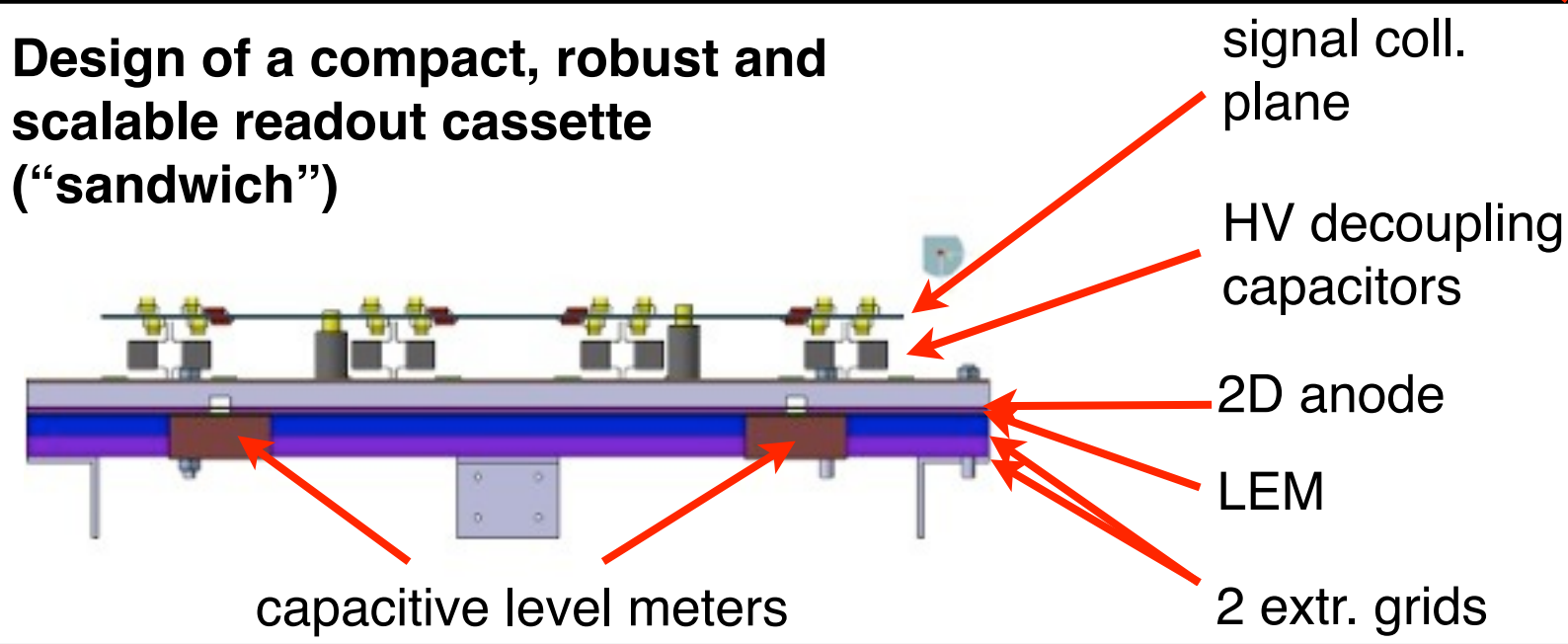
LAr-LEM TPC@CERN: Production of a 40x80 cm² charge readout sandwich

- ▶After successful test of LEM and 2D anode in the 3L setup we designed and produced a 40x80 cm² charge readout for a new 250L LAr LEM-TPC (production and assembling finished by summer 2011)
- ▶The ArDM cryostat @CERN was used for a first test of the new charge readout system

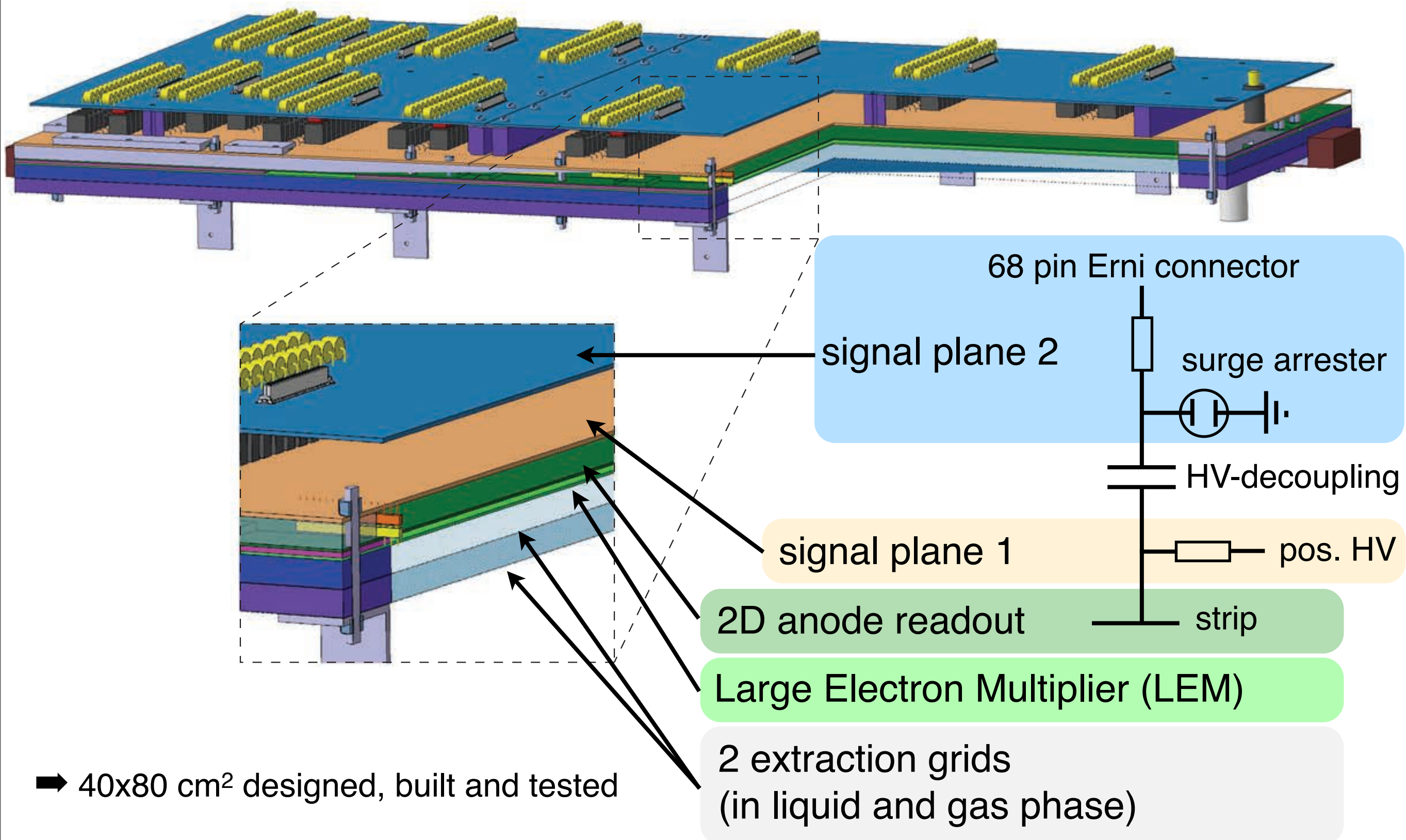


- Manufacturer: CERN TS/DEM group and ELTOS company (Italy)
- Largest LEM/THGEM and 2D readout ever produced!!!

Design of a compact, robust and scalable readout cassette (“sandwich”)



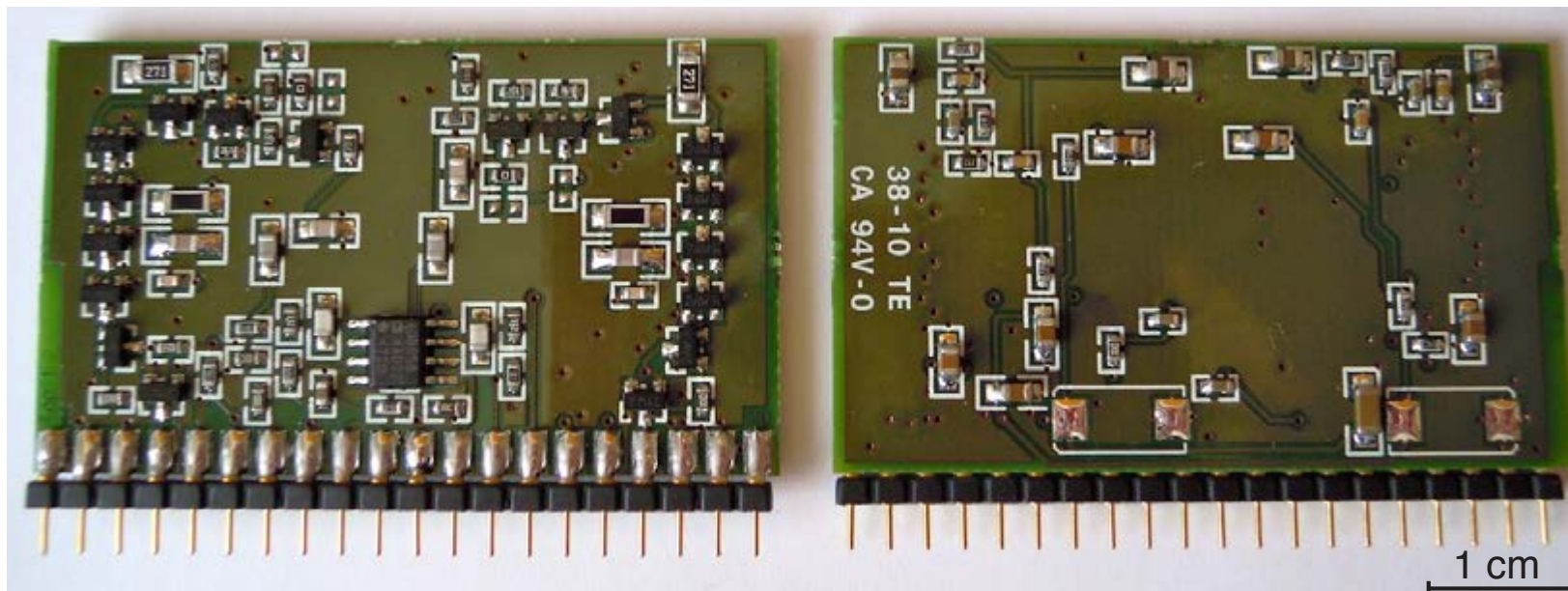
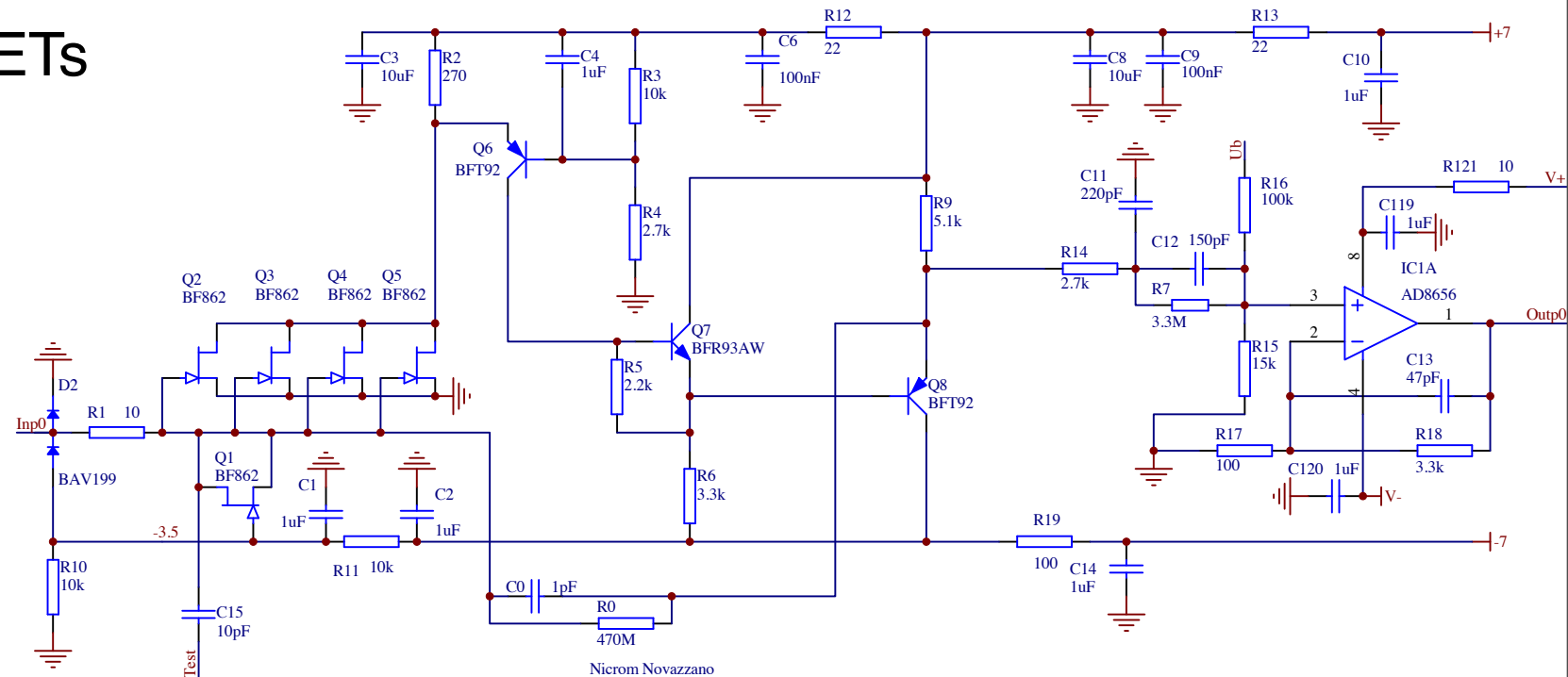
Charge readout sandwich



The ETHZ preamplifier

electric layout

- Cascode design with 4 parallel JFETs at the input (C. Boiano et al. IEEE Trans. Nucl. Sci. 52 (2004) 1931)
- $RC=470\ \mu s$ feedback ($C=1\text{pF}$)
- RC-CR shaper with zero-pole sub. mechanism (no undershoot)
- over-voltage protection at input

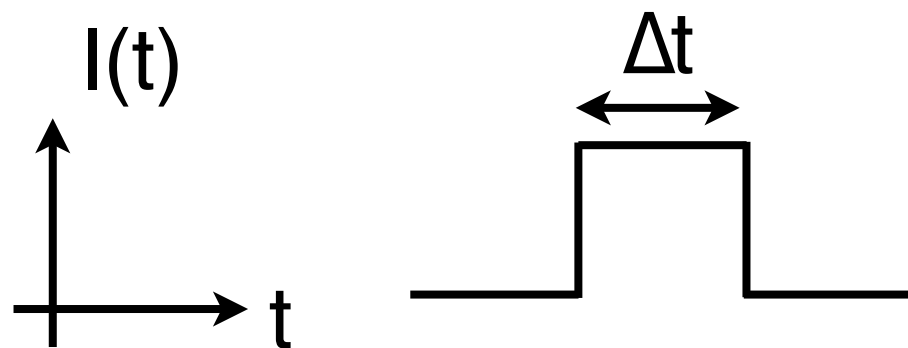


realization

- preamplifier is realized with discrete components
- two preamplifier circuits are implemented on a single 4-layer PCB

Performance of the ETHZ preamplifier

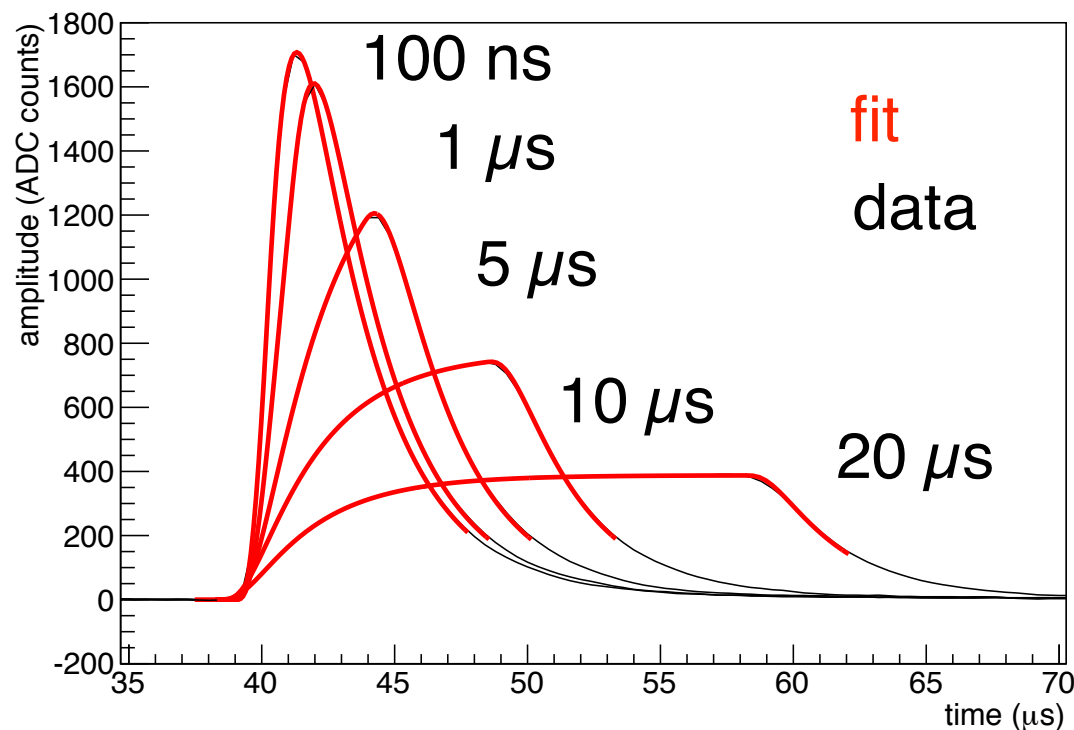
32 preamplifiers have been characterized with a well defined charge input:



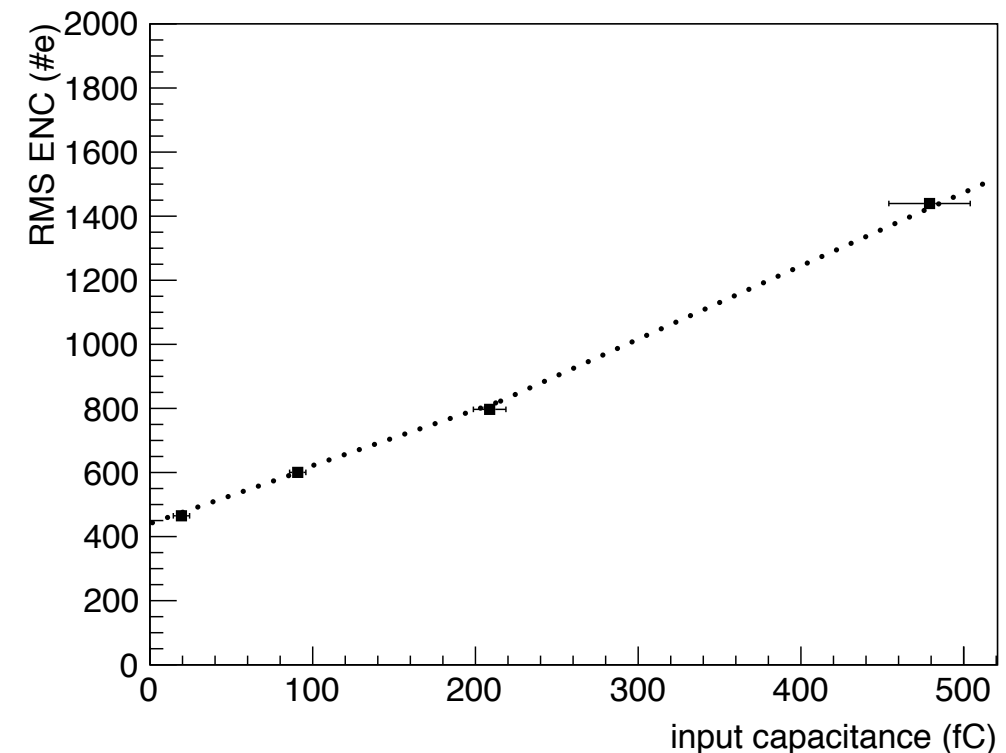
Summary

shaping time τ_D	$2.8 \pm 0.1 \mu\text{s}$
shaping time τ_I	$0.45 \pm 0.02 \mu\text{s}$
sensitivity	$13.8 \pm 0.4 \text{ mV/fC}$
open loop gain	$\approx 10^4$
linearity (0-180 fC)	$\pm 1\%$
ENC (RMS, $C \approx 200 \text{ pF}$)	$770 \pm 30 \text{ electrons}$
S/N (1 fC, $C \approx 200 \text{ pF}$)	8.1 ± 0.3

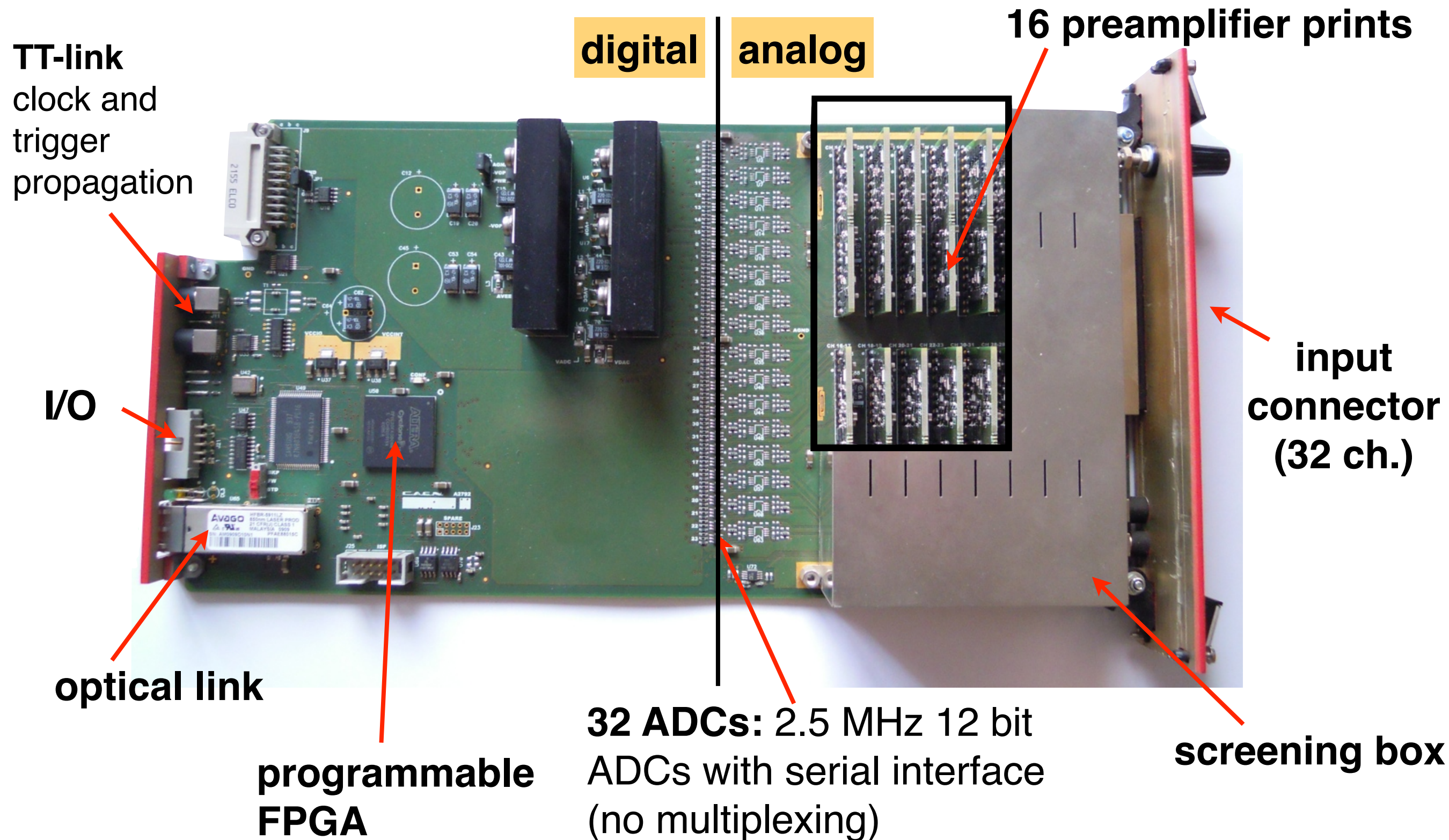
pulse shaping (varying Δt)



RMS ENC vs. input capacitance



The CAEN A2792 acquisition board

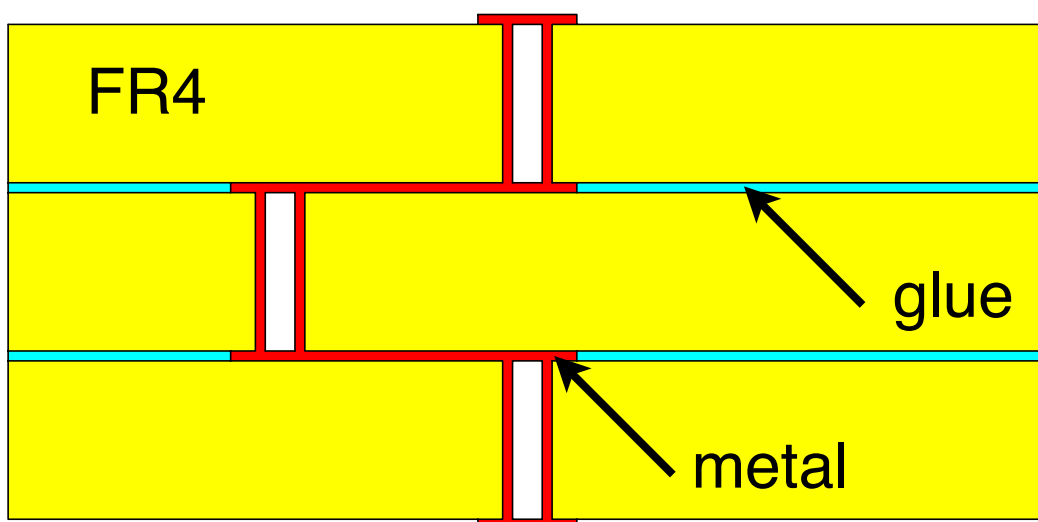


Signal feed-through technology

➡ vacuum tightness

- Four layer printed circuit board (PCB) with displaced vias (through-hole paths) guarantees tightness
- The PCB is attached to a metal flange
 - CF-standard for cryogenic temperatures
 - sealed with glue or gasket

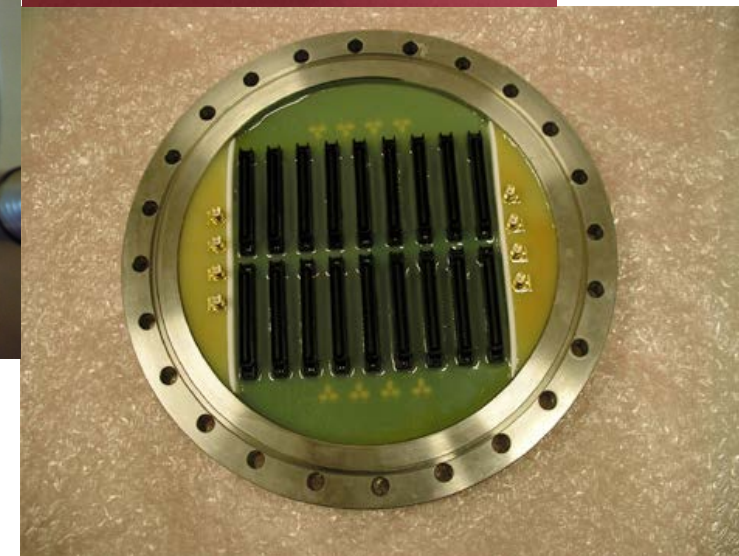
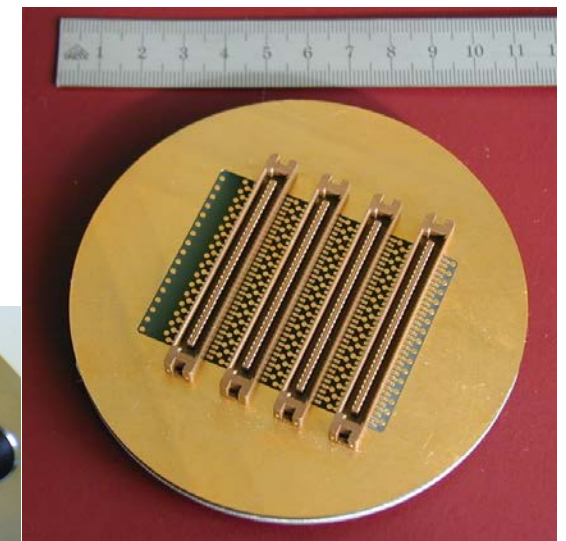
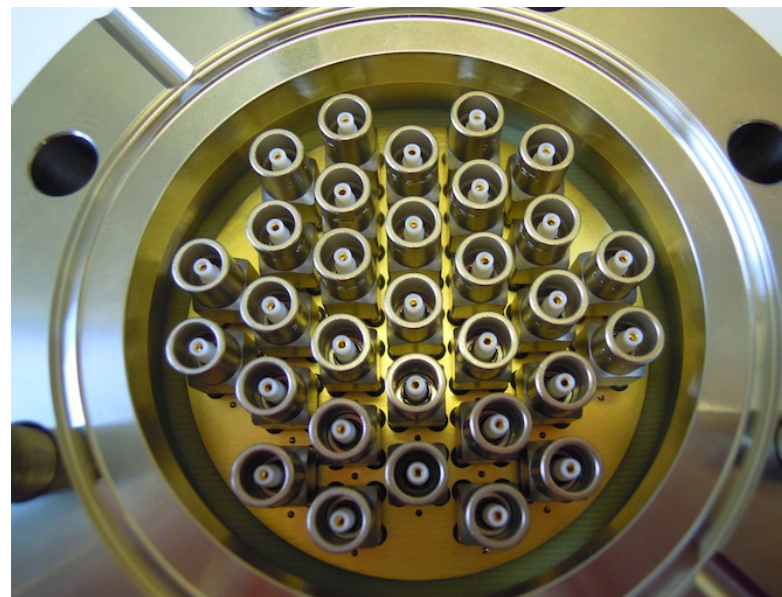
design



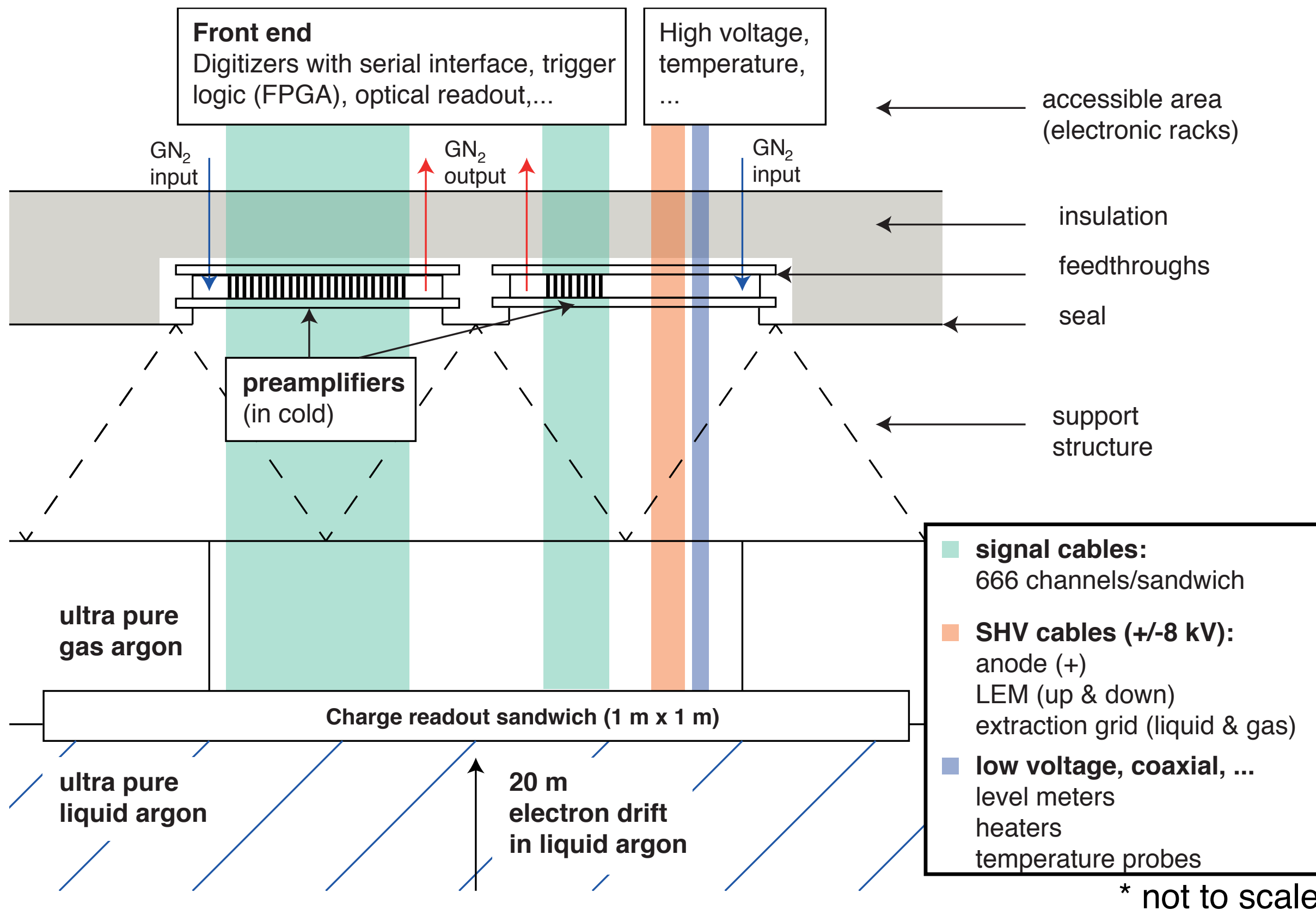
➡ High channel density

PCB technique allows to place 18 68-pin plugs on a single CF-250 lange

Examples from different experiments



Charge readout unit



Development of a front end ASIC CMOS preamplifier

working group

E.Bechetoille, S.Gardien, C.Girerd, H.Mathez (electronics), Bruno Carlus (software),
Dario Autiero, Yves Déclais, Jacques Marteau

CNRS / IN2P3 / UCBL - Institut de Physique Nucléaire de Lyon

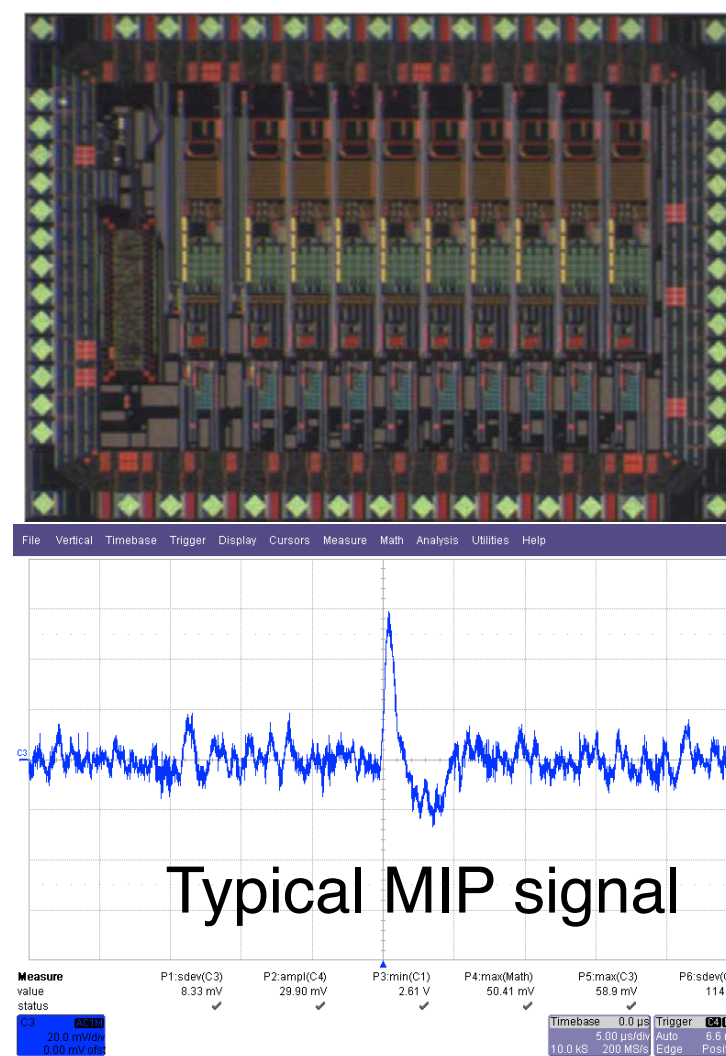
Advantages of ASIC cold electronics

- Exploit intrinsic noise reduction at low T (minimum around 110K)
- Large scale integration and costs reduction (1~1.5 eur/ch)

Summary

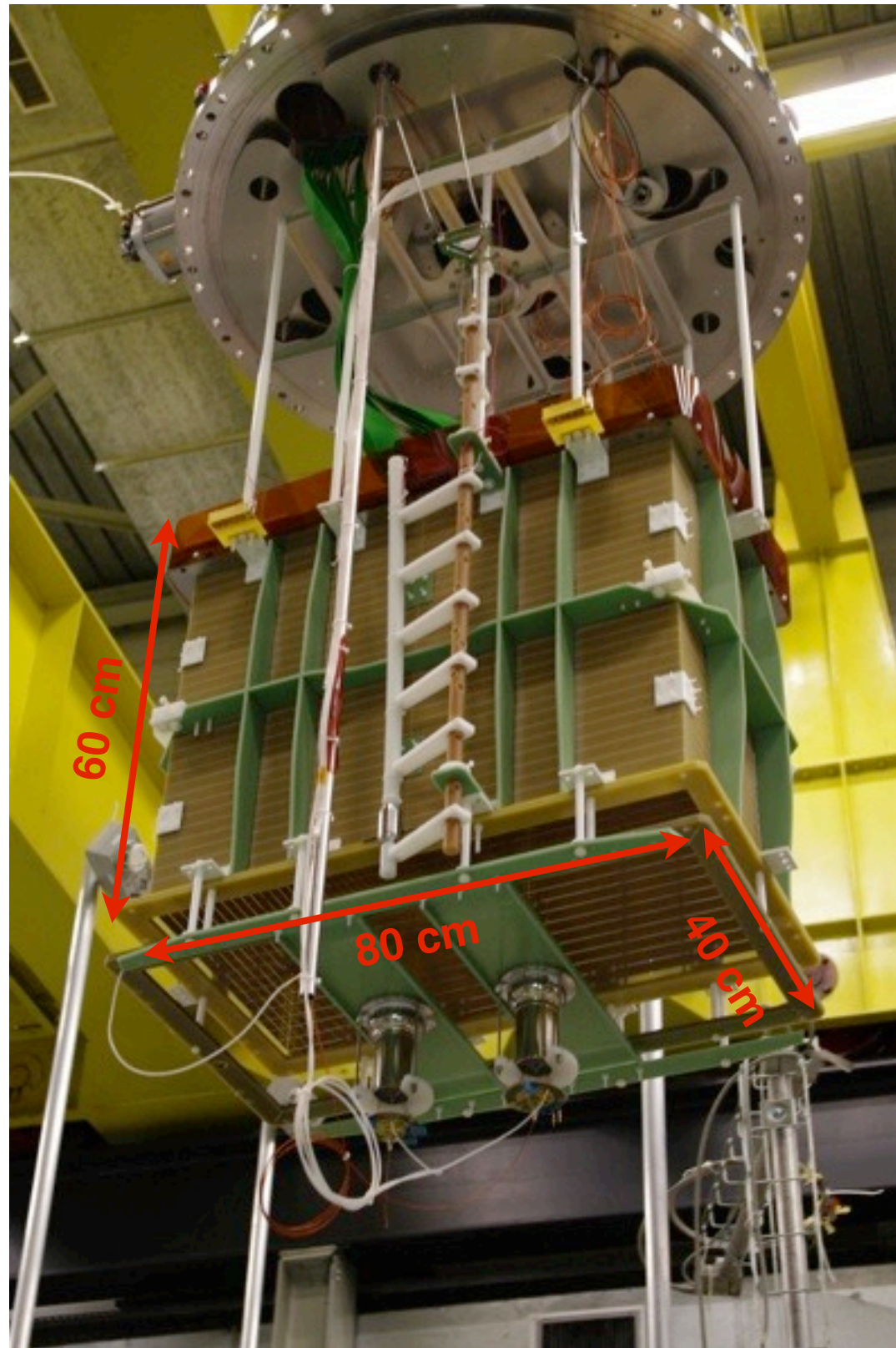
- Activity started in 2007
- R&D on a analog ASIC preamplifier working at cryogenic temperature for the charge readout of the LAr TPC
- Performance of the 4th generation:
 - 18 mV/fC sensitivity
 - 1500 ENC at the end of the full chain with 250 pF input capacitance @ 110K (-20%)
- R&D on the Gigabit Ethernet readout chain + network time distribution system PTP (IEEE1588)

ASIC V4



LAr-LEM TPC@CERN: The largest LEM-TPC ever

Detector fully assembled

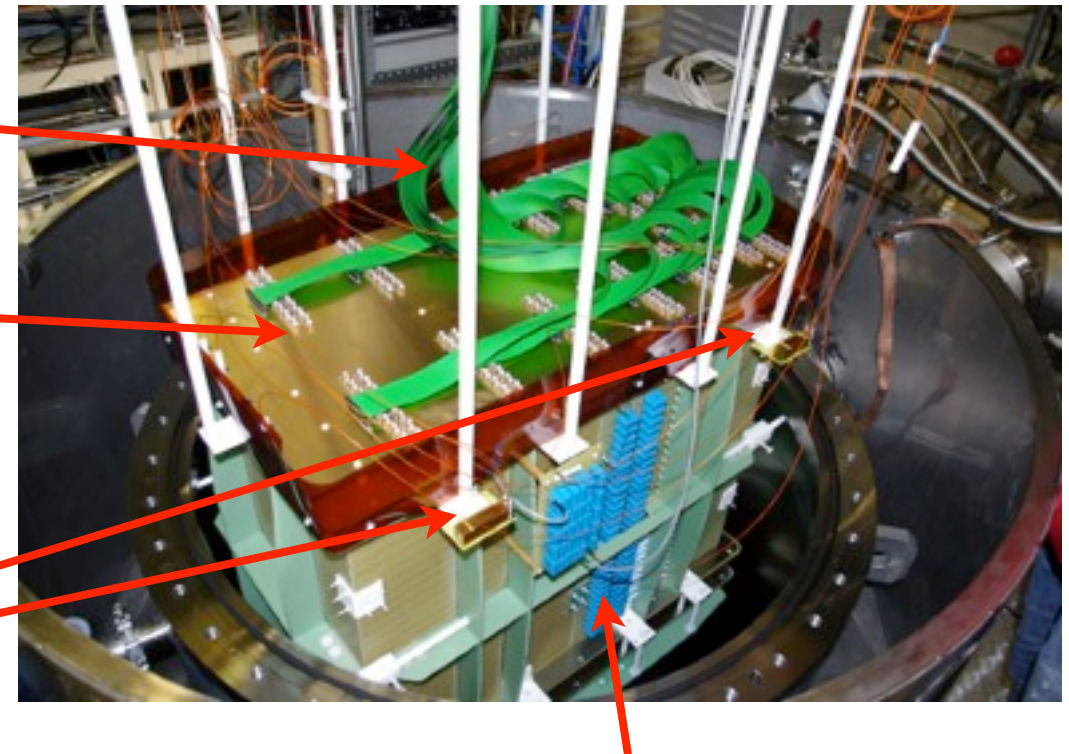


Chamber going into the ArDM cryostat

16 signal cables

charge readout sandwich

4 capacitive level meters



Cockcroft-Walton HV system

Final connection to the DAQ system

