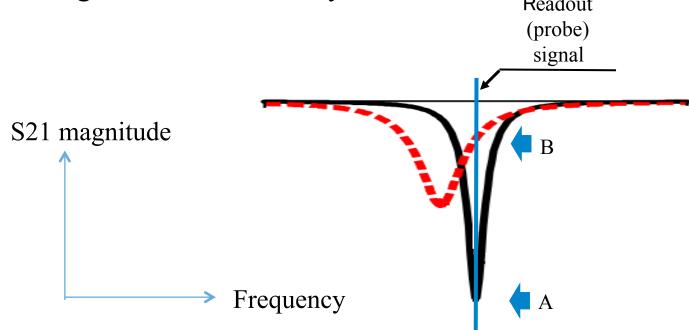
# Frequency sweeping scheme Readout system

### Hitoshi Kiuchi National Astronomical Observatory of Japan

### How to detect optical loading?

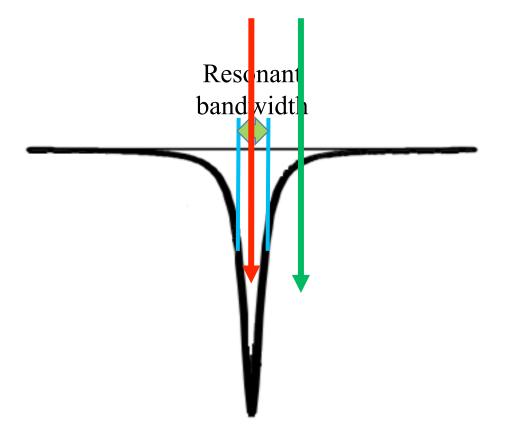
The MKID detects the optical loading as the change in the resonance frequency. The probe signal is modulated by the resonator.<sub>Readout</sub>



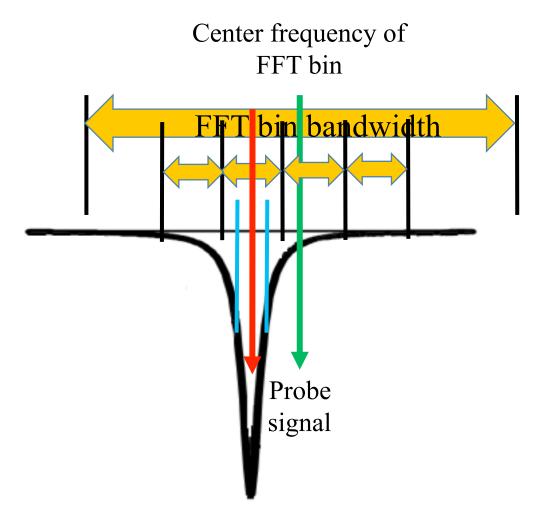
Initially, a microwave probing signal is tuned to the resonant frequency of the lowest optical loading state, and any optical loadings are estimated from the changed phase and amplitude at the probing signal frequency.

### Requirements on the readout (probe) signal?

Measured SNR is not sufficient if the probe signal is not within the resonance bandwidth. The probe signal frequency should be adjusted into the resonant bandwidth precisely or the detected signals have low SNR.



One possible solution to improve the frequency resolution is to increase the number of FFT points.

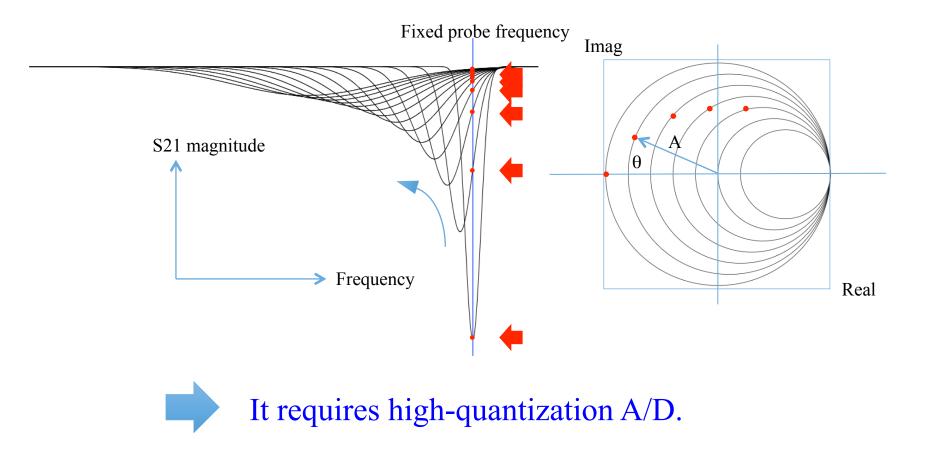


Increasing the number of FFT points makes the readout system larger.

An optical loading causes the resonance frequency to shift to lower frequency and the resonance dip to decrease in depth.

The shifted resonant frequency is estimated by the fixed probing signal.

The changed resonant frequency is measured indirectly.



(1) To maintain sufficient SNR even with the optical loading states, this method requires a high-quantization A/D converter.

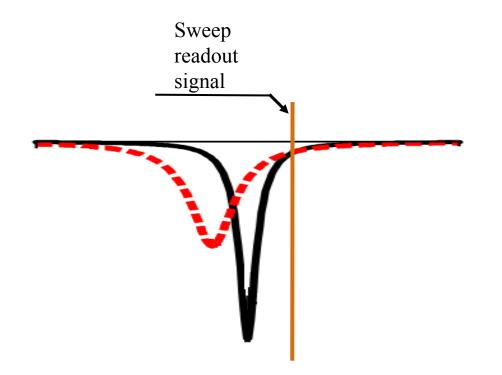
 (2) The frequency bandwidth of the probe comb signals increase with the array scale, frequency domain multiplexing array requires a high-rate-sampling A/D converter.



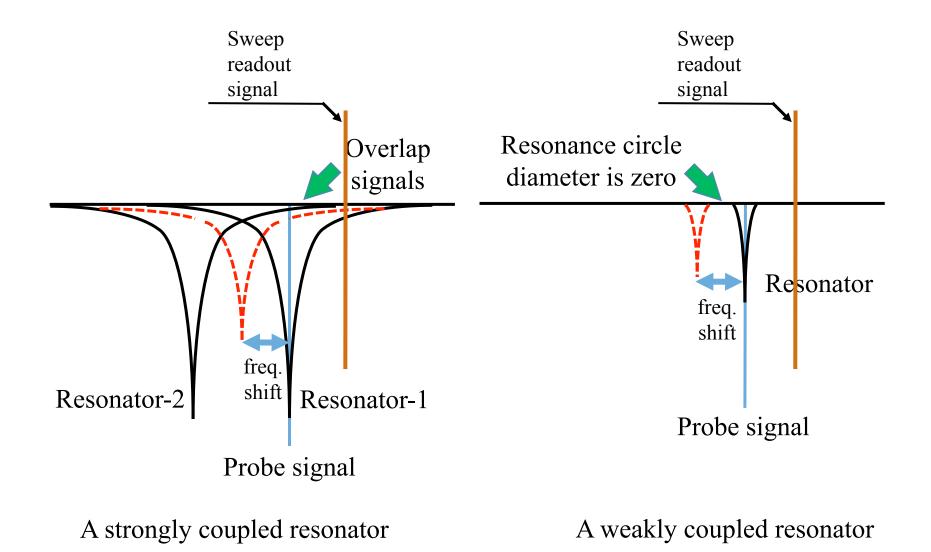
The high-quantization and the high-rate-sampling are incompatible functions.

## Local frequency sweeping scheme

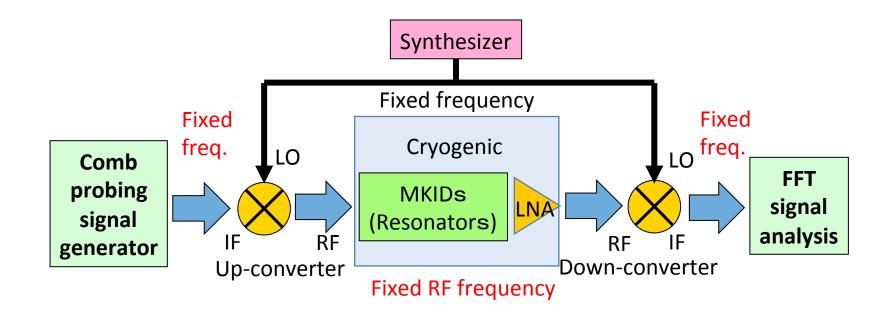
In this scheme, the resonance frequency is directly measured by a swept frequency probe signal, making it possible to measure the phase/amplitude after optical loading with a high-SNR.



The precise initial frequency tuning is unnecessary.

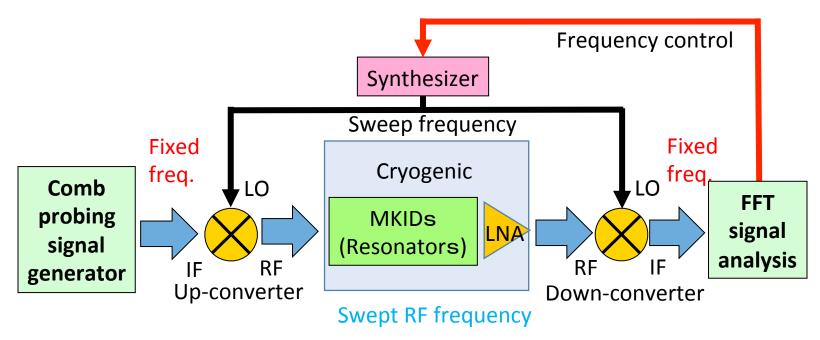


### Current scheme



A synthesizer provides the LO signal, forming a homodyne frequency conversion scheme.

# Local frequency sweeping scheme



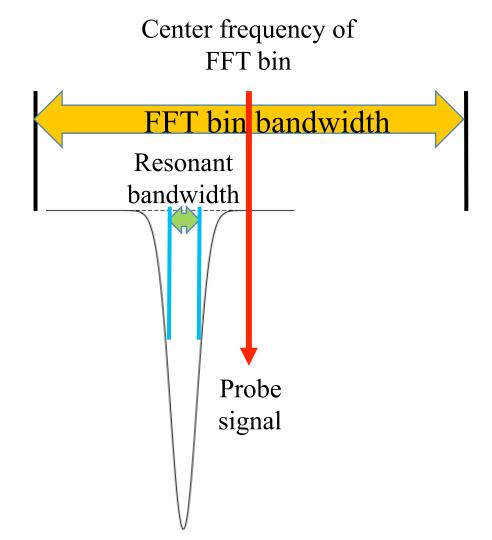
Forming a homodyne frequency conversion scheme.

From a view of the readout system, the digitally generated comb signal and the received comb signal are exactly same frequency in any case. Due to the LO signals of down-converter and up-converter being common, the phase instability of the frequency sweep of the LO signal is canceled out. Two applications:

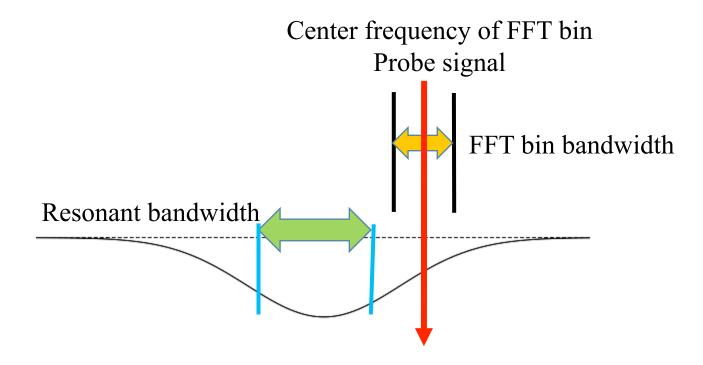
(1) Increase frequency resolutionSweep frequency range < FFT bin bandwidth</li>

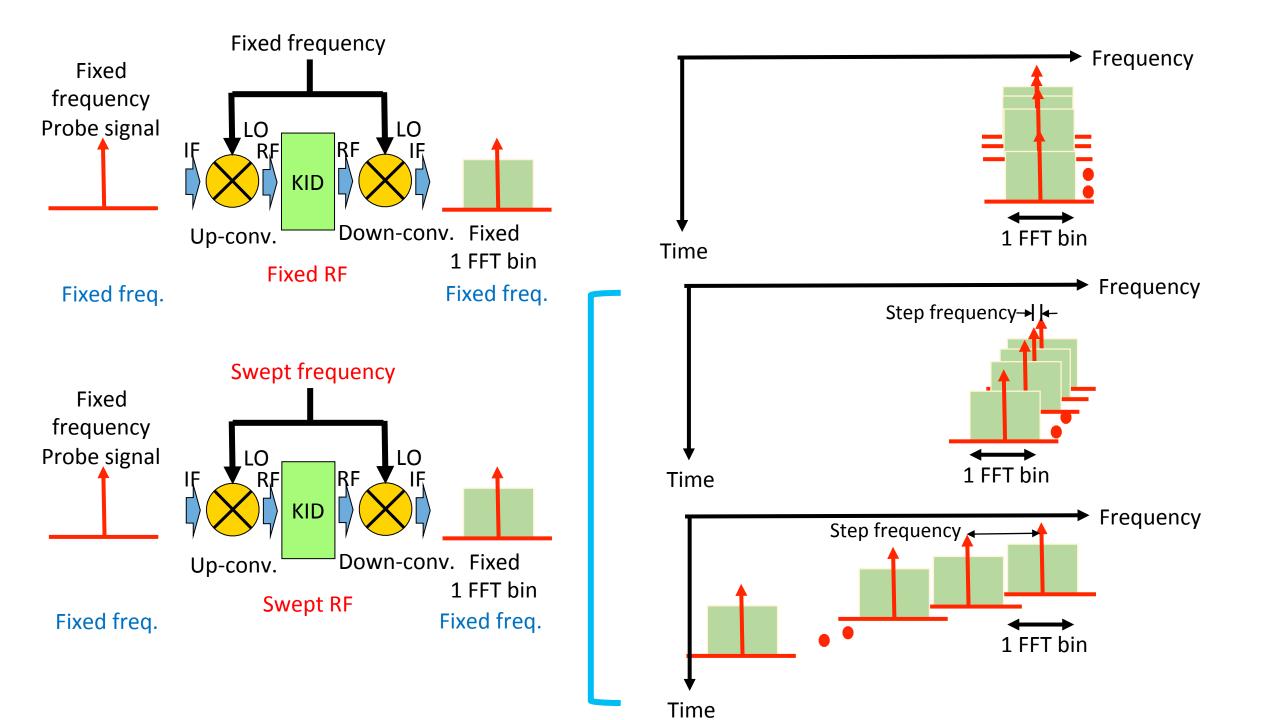
(2) Increase frequency dynamic rangeSweep frequency range > FFT bin bandwidth

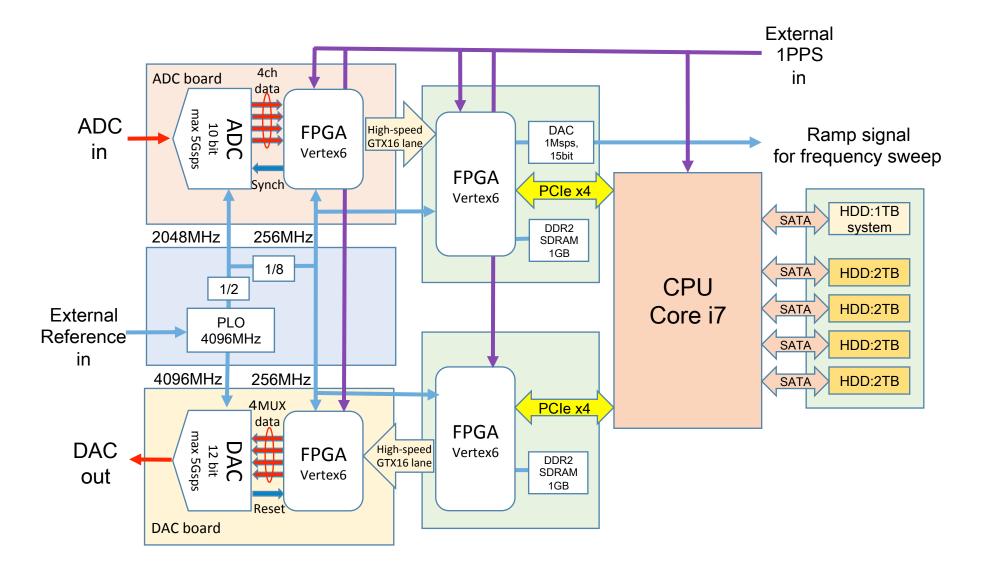
Increase frequency resolution Sweep range < FFT bin bandwidth



Increase the dynamic range Sweep range > FFT bin bandwidth



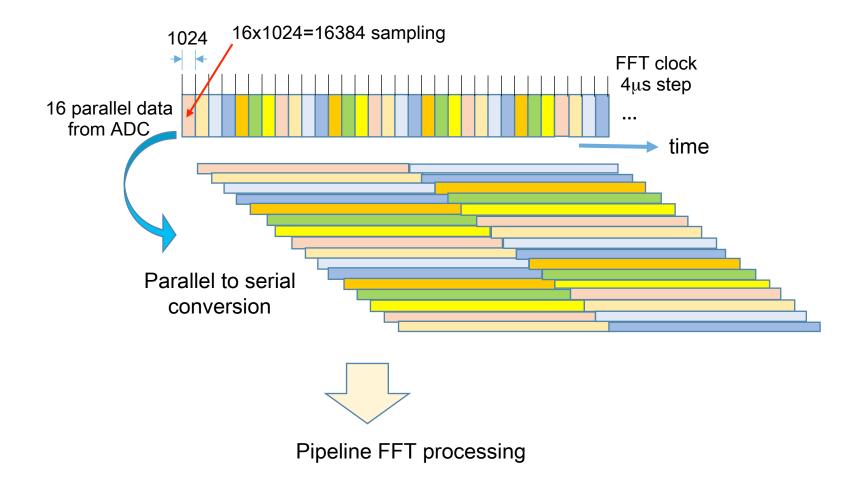


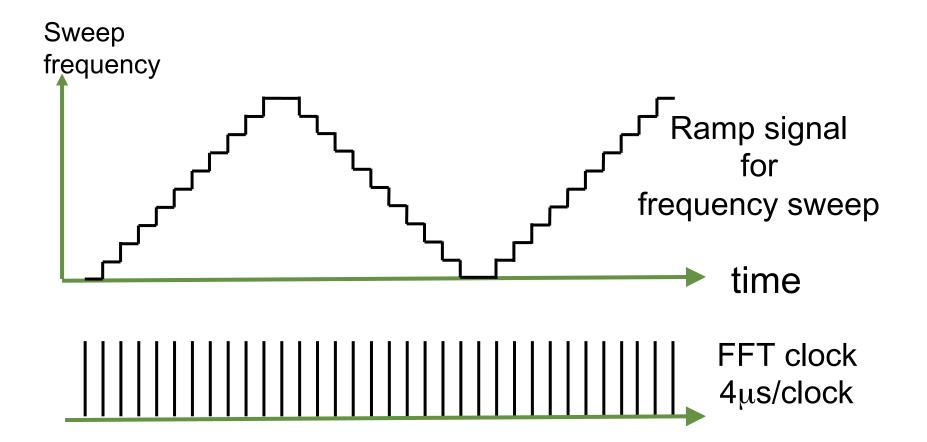


### An issue of consideration

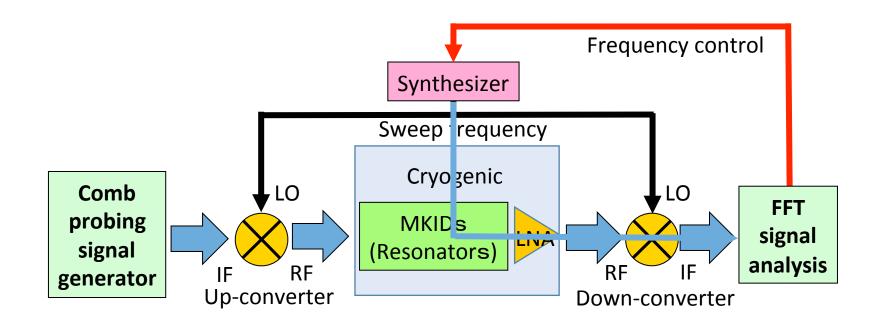
There is a possibility that the frequency sweep causes a measurement error during FFT processing.It is desirable that the probe frequency is unchanged during the FFT processing (data acquisition) period.

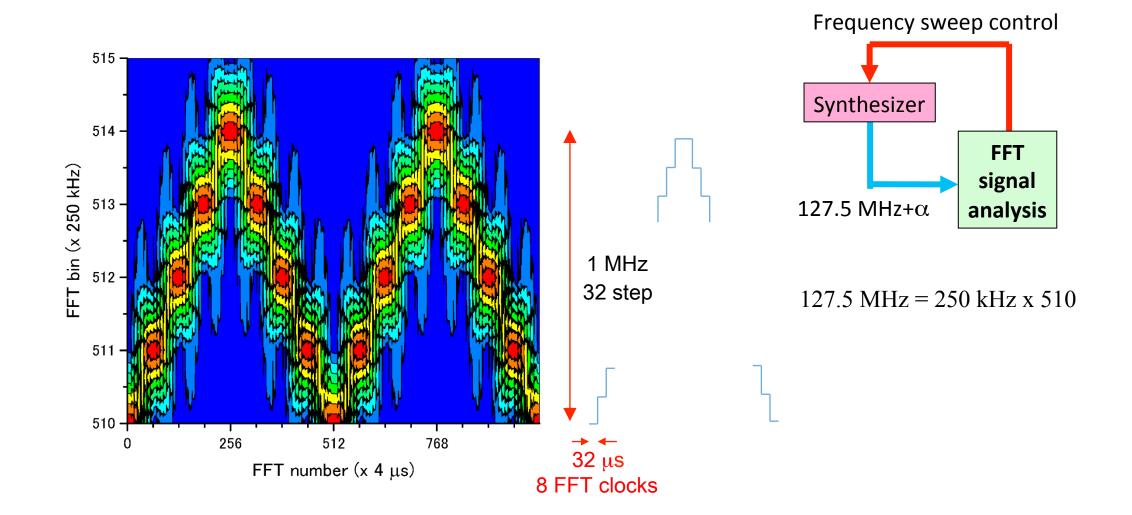
This means that the frequency sweep should be step wise and the timing of the frequency step should be synchronized with the FFT processing (data acquisition) timing.

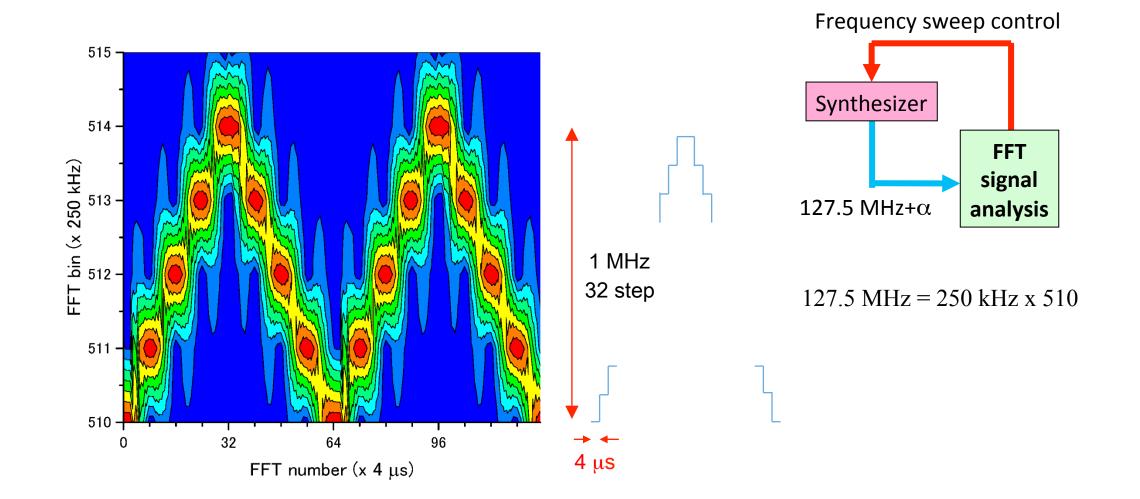




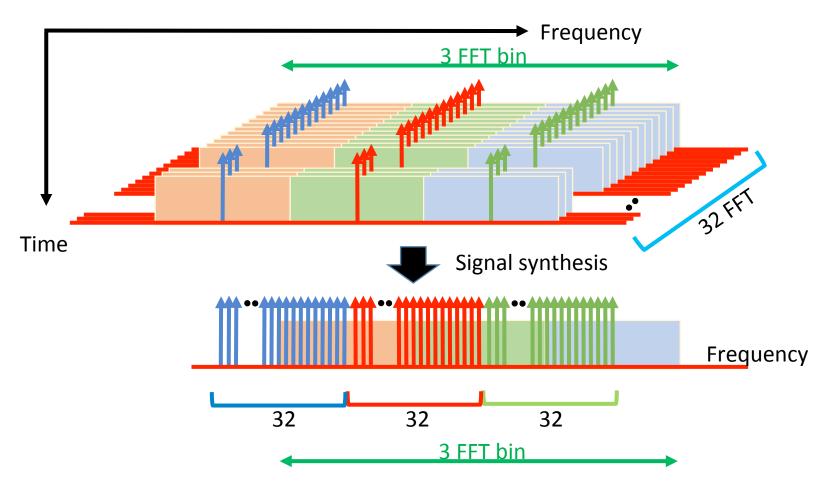
## Signal timing check



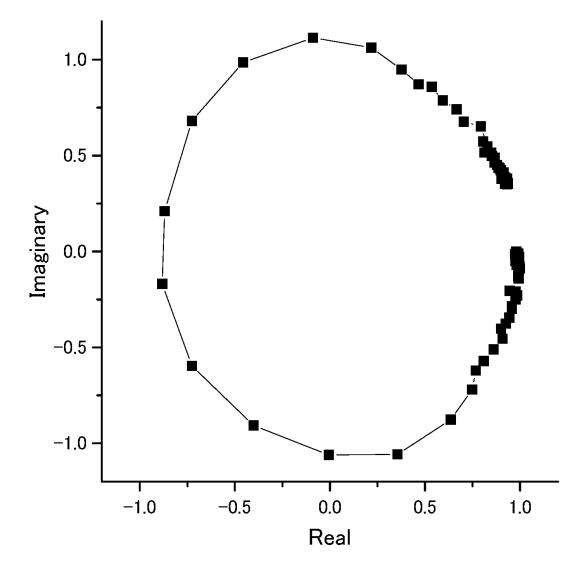




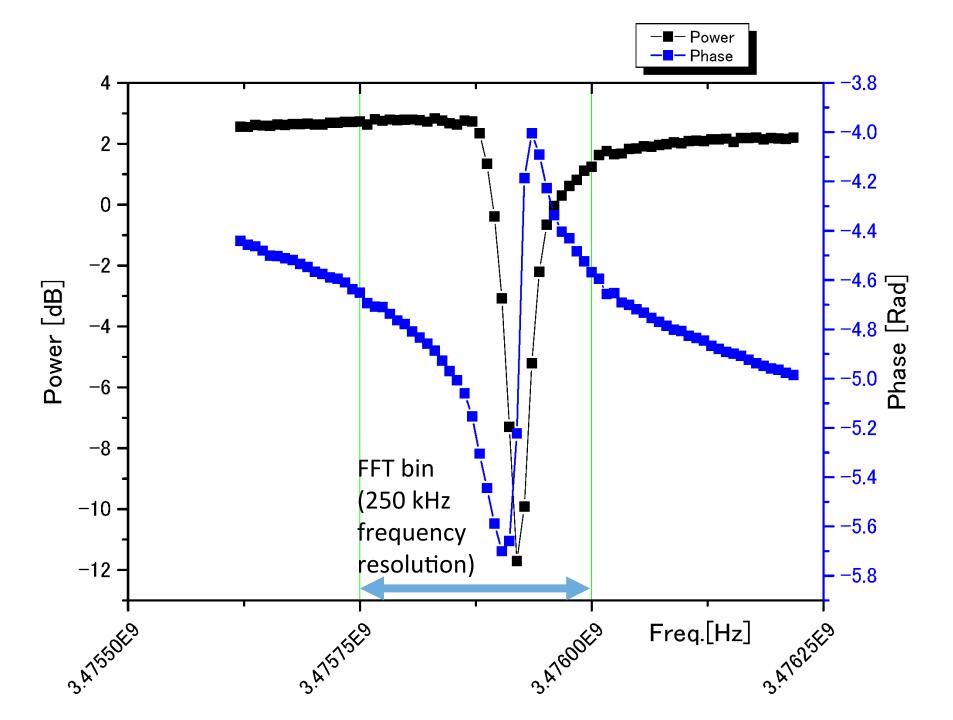
### Increase frequency resolution (32 times)



Frequency sweeping with 3 probe signals. In the normal usage, a probe signal corresponds to a resonator on one-to-one basis.



The trajectory in the complex plane of S21 parameter through a MKID with frequency sweep.



#### Features of the system

The LO frequency sweep scheme is intended to increase the frequency dynamic range: the frequency interval of the sweep frequency is greater than that of the FFT bin for the ground-based telescope,

and to increase frequency resolution without increasing the number of FFT points: the frequency interval of the sweep frequency is narrower than that of the FFT bin for the space telescope.

Neither the precise initial tuning (tune the probing signal to the resonant frequency of the lowest optical loading state) is necessary any more.

